Pseudo Goldstone Dark Matter and Inflation

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- Inflation and Dark matter can't be explained within SM.
- The two unknown problems may be connected.
- We look for a solution by considering a different sector that can address these two.
- Inflation sector can be embedded into a hidden susy breaking sector where inflationary energy scale can be dynamically generated.
- The susy breaking sector could be a SQCD sector in the form of supersymmetric QCD.
- SQCD-embedded inflation model has existence of a UV complete theory.

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Smooth Hybrid Inflation in SQCD

• Inflationary sector represented by a strongly coupled supersymmetric SU(N) gauge group with $N_f = N_C$ flavours of quark superfields Q_i and \bar{Q}_i .

Global symmetry: $SU(N_f) \times SU(N_f) \times U(1)_B \times U(1)_R$. Below the strong coupling regime (Λ_0) they form mesons and baryons [hep-th/0602239]

$$T_{ij} = \frac{Q_i \bar{Q}_i}{\Lambda_0}, \ B = \frac{\epsilon_{abcd} Q_1^a Q_2^b Q_3^c Q_4^d}{\Lambda_0^3}, \ \bar{B} = \frac{\epsilon_{1234} \bar{Q}_1^a \bar{Q}_2^b \bar{Q}_3^c \bar{Q}_4^d}{\Lambda_0^3} \quad (1)$$

Superpotential: $W_{N_f=N_c} = S(\frac{\det T}{\Lambda_0^2} - B\bar{B} - \Lambda_{eff}^2)$

- $N_f = N_C$ theory can be realized from $N_f = N_c + 1$ version of SQCD by making the N_f th quark heavy $(W_m = m_Q Q_{N_f+1} \overline{Q}_{N_f+1})$ below it.
- The heavy quark can be identified with the S-field and $\Lambda_{\text{eff}} = m_Q \Lambda_0$.

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contd.

• Inflation can be realized with $N_f = N_c = 4$ along $B = \overline{B} = 0$ and $T_{ij} = \chi \delta_{ij} [arXiv:0902.0972]$.

So inflationary superpotential [hep-ph/9606297]

$$W_{\rm Inf} = S\left(\frac{\chi^4}{\Lambda_0^2} - \Lambda_{\rm eff^2}\right)$$
 (2)

At global minimum $T = \sqrt{\Lambda_0 \Lambda_{\text{eff}}}$, thus breaks the global symmetry $SU(4)_L \times SU(4)_R \times U(1)_B \times U(1)_R \rightarrow SU(4)_V \times U(1)_B \times U(1)_R.$

Inflationary predictions: $r \simeq 10^{-7}$, $n_s = 0.967$. Allowed by Planck 2016???[arXiv:1502.01589]

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NGB as DM

• Associated flavor symmetry breaks down after inflation, 15 NGB's will appear.

$$T_{N_f \times N_f} = \chi \exp\left(\frac{iG_S^2 \lambda^a}{\langle \chi \rangle}\right),\tag{3}$$

Can those (t^a) be DM??

• Masses of NGB's can be generated through Dashen formulla by explicit breaking of broken SU(4) in the superpotential $(m = \text{diag}\{m_1, m_2, m_3, m_4\}).$

$$\langle \chi \rangle^2 (m_{\mathcal{G}_S}^a)^2 = \langle 0 | [\tilde{\mathcal{Q}}_a, [\tilde{\mathcal{Q}}_a, H]] | 0 \rangle \tag{4}$$

$$= \bar{\psi} \left[\frac{\lambda_a}{2}, \left[\frac{\lambda_a}{2}, m_{\text{diag}} \right]_+ \right]_+ \psi$$
 (5)

We modify the inflationary superpotential to generate the interactions of NGBs with SM

$$W_{\rm Inf} = S\left(\frac{\det T}{\Lambda_0^2} - \Lambda_{\rm eff}^2\right) + \kappa_1 S\left\{{\rm Tr}(T^2) - \frac{({\rm Tr}\ T)^2}{N_f}\right\} + \kappa_2 SH_u H_d,$$

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contd.

- If $m_1 = m_2 = m_3 = m_4$ we get 15 no's of degenerate DM.
- For a different choice $m_1 = m_2 = m_3 = m_\gamma$ and $m_4 \gg m_\gamma$, three different sets of NGB (i) $m_A = m_\gamma \frac{\Lambda_{\text{eff}}^3}{\langle \chi \rangle^2}$, (ii) $m_B = \left(m_4 + m_\gamma\right) \frac{\Lambda_{\text{eff}}^3}{\langle \chi \rangle^2}$,

(iii)
$$m_C = \left(\frac{3}{2}m_4 + \frac{m_\gamma}{2}\right)\frac{\Lambda_{\text{eff}}^3}{\langle\chi\rangle^2}.$$

Degeneracies 8,6 and 1.

Interaction Lagrangian with visible sector:

$$V \supset -\left(\lambda' h^2 + \lambda'' h v_d\right) \sum_{a}' (G_S^a)^2, \qquad (7)$$

where

$$\lambda = \frac{\kappa_1 \kappa_2}{2}, \lambda' = \frac{1}{2} \lambda \sin \alpha \cos \alpha, \quad \lambda'' = \frac{1}{2} \lambda \cos \alpha (\tan \alpha - \tan \beta)$$
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Important formulas

$$\tan \beta = \frac{v_u}{v_d}, v = \sqrt{v_u^2 + v_d^2} \simeq 246 \text{GeV}$$
(9)

The mixing angle α can be expressed in terms of β and the pseudoscalar A mass as

$$\tan 2\alpha = \tan 2\beta \, \frac{M_A^2 + M_Z^2}{M_A^2 - M_Z^2},\tag{10}$$

In large M_A limit $tan2\alpha = tan 2\beta$. Solutions: $\alpha = \beta$ and $\alpha = \beta + \frac{\pi}{2}$. We work with $\alpha = \beta + \frac{\pi}{2}$ otherwise $\lambda'' \to 0$. WHEPP 2019

Relic density and direct search

Three parameters m_{G_s} , λ , tan β . **Case I**: Boltzman equation:

$$\dot{n}_{G_S} + 3Hn_{G_S} = -\langle \sigma v \rangle_{G_S G_S \to SM} (n_{G_S}^2 - n_{eq}^2)$$
(12)

Relic density $\Omega_T = 15\Omega_{G_s}$.



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Case II: Three sets of DM ($m_A < m_B < m_C$).

Assume A-type having mass $m_h/2$.

 $CC \rightarrow BB, AA$ and $BB \rightarrow AA$ possible.

$$\Omega_T = 6\Omega_B + \Omega_C \tag{13}$$

Parameters: $\{\kappa_1, \lambda, m_B, m_C\}$



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Boltzman equations:

$$\begin{split} \frac{dn_{C}}{dt} + 3Hn_{C} &= -\langle \sigma v \rangle_{G_{C}G_{C} \to \mathrm{SM}} (n_{C}^{2} - n_{C}^{\mathrm{eq2}}) - 6\langle \sigma v \rangle_{G_{C}G_{C} \to G_{B}G_{B}} (n_{C}^{2} - \frac{n_{C}^{\mathrm{eq2}}}{n_{B}^{\mathrm{eq2}}} n_{B}^{2}) \\ &- 8\langle \sigma v \rangle_{G_{C}G_{C} \to G_{A}G_{A}} (n_{C}^{2} - \frac{n_{C}^{\mathrm{eq2}}}{n_{A}^{\mathrm{eq2}}} n_{A}^{2}) \\ \frac{dn_{B_{i}}}{dt} + 3Hn_{B_{i}} &= -\langle \sigma v \rangle_{G_{B_{i}}G_{B_{i}} \to \mathrm{SM}} (n_{B_{i}}^{2} - n_{B_{i}}^{\mathrm{eq2}}) + \langle \sigma v \rangle_{G_{C}G_{C} \to G_{B_{i}}G_{B_{i}} \sigma_{B_{i}}} n_{B_{i}}^{2} n_{B_{i}}^{2}) \\ &+ 8\langle \sigma v \rangle_{G_{B_{i}}G_{B_{i}} \to G_{A_{i}}G_{A_{i}}} (n_{B_{i}}^{2} - \frac{n_{B_{i}}^{\mathrm{eq2}}}{n_{A_{i}}^{2}} n_{A_{i}}^{2}) \end{split}$$

Results:



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- Chiral symmetry broken down spontaneously at the end of inflation, NGB appears in the set-up.
- Depending on the explicit chiral symmetry breaking term, there could be different degree of degenracy among the masses of these pNGBs.
- 15 dengenerate DM is almost ruled out from direct detection limit
- However degerenrate DM scenario is still valid thanks to the interactions among them.

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