





D. Mirarchi

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MANCHESTER 1824 The University of Manchester

9th Evian Workshop, 31th January 2019





### Outline

# I. UFO II. ULO III. 16L2 IV. 10Hz V. Conclusions









# I. UFO II. ULO III.16L2 IV.10Hz V. Conclusions





### Introduction









#### Very fast failure! Few turns!

28/09/2018, 11h03, Quench of MB.B18R7



- Unidentified Falling Object:
  - **1.** Dust particles falling into the beam (tens of  $\mu$ m)
  - 2. Inelastic proton-UFO collisions
  - 3. Hadronic showers and energy deposition on magnets coil
  - 4. Quench





## **UFO detection**





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#### Tight connection between **quench level**, **BLM threshold** and **energy deposition simulations**



- Main lessons learnt:
  - ✓ **Similar losses** in all quenches (within a factor of 2)
  - Inelastic collision in the range of 6x10<sup>7</sup>-1.2x10<sup>8</sup>  $\checkmark$
  - $\checkmark$  UFO radius in the range of 40-100  $\mu$ m

of quench level models!



### **UFO rate evolution**



#### UFO rates in the arc (cells $\geq$ 12) in stable beams



Clear de-conditioning during LS1 and re-conditioning during Run 2

Similar UFO rate as end of Run 1 achieved after about 1.5-2 years



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# I. UFO II. ULO III.16L2 IV.10Hz V. Conclusions





### Introduction



- Significant UFO activity in cell 15R8 during commissioning in 2015 (14 dump, 3 quench)
- First actions taken:
  - Energy deposition simulations

Located vertex of hadronic showers at MB.C15R8.B2 centre ±1m



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### **ULO** measurements



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#### 4 correctors bump in V plane 2.0Orbit center 014 016 Vertical Position [mm] 1.5B15 1.0015 0.0 23900 23950 2405024100 23800 238502400024150Longitudinal Position [m]

#### **Measurement procedure:**

- **Beam shaped with IR7-TCPs**:  $2/4\sigma$  in H/V 1.
- Local aperture probed systematically: 2.
  - H/V bump steps of 3/0.5mm
- 3. Max bump excursion:

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in  $H \sim \pm 14$ mm (losses at Q15) in  $V \sim \pm 8$ mm (losses at Q14 and Q16)



![](_page_10_Figure_11.jpeg)

#### **3 correctors bump in H plane**

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![](_page_11_Picture_0.jpeg)

## **ULO evolution and mitigation**

![](_page_11_Picture_2.jpeg)

![](_page_11_Figure_3.jpeg)

#### • Mitigation strategy:

 $\Rightarrow$  Orbit bump to bypass it, while ensuring 10 $\sigma$  at injection

Constant monitoring during Run 2

ULO scan during commissioning to deploy optimal bump

#### Limitation successfully removed! Access planned mid-April, let's see what it is!

![](_page_11_Picture_9.jpeg)

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![](_page_11_Figure_11.jpeg)

![](_page_12_Picture_0.jpeg)

![](_page_12_Picture_1.jpeg)

![](_page_12_Picture_2.jpeg)

# I. UFO II. ULO III.16L2 IV.10Hz V. Conclusions

![](_page_12_Picture_4.jpeg)

![](_page_13_Picture_0.jpeg)

## Introduction

![](_page_13_Picture_2.jpeg)

- Sudden increase of losses in the half-cell **16L2** represented **"The" machine limitation in 2017** 
  - **67 dumps** induced: only two at injection, one quench induced at high energy
- <u>Three stages of loss rate:</u>
  - 1. Steady loss in 16L2 arising during the ramp along the entire fill (~ few 1e-6 Gy/s)
  - 2. Sharp rise of losses in 16L2 "UFO-like" (1e-3 ÷ 1e-2 Gy/s)
  - 3. Very fast rise of losses at collimators triggering a dump (few ms)-

![](_page_13_Figure_9.jpeg)

Scattered particles reach primary collimators

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![](_page_14_Picture_0.jpeg)

![](_page_14_Picture_2.jpeg)

#### • Energy deposition simulations to identify longitudinal location of hadronic showers vertex

![](_page_14_Figure_4.jpeg)

![](_page_15_Picture_0.jpeg)

![](_page_15_Picture_2.jpeg)

![](_page_15_Figure_3.jpeg)

#### Mitigation strategy:

- ✓ Dipolar field of MCBs
- ✓ 8b4e filling pattern

Prevent "UFO-like" events suppressing multipacting

✓ Solenoid in the field free region of the interconnection

10% increase of ppb in 2017 Mitigation at inj. in 2018

✓ BCM production scheme → Brighter beams

![](_page_15_Picture_11.jpeg)

![](_page_16_Picture_0.jpeg)

![](_page_16_Picture_2.jpeg)

- Team work made the difference:
  - Many people involved and actions also discussed in a dedicated Task Force
  - ✓ Additional hardware installed with impressive schedule
- Most probable scenario: Accidental air venting during the removal of the mobile group for the final pinch-off (EYETS 2016-2017)
  - Upgraded pumping ports design
  - ✓ New pumping scheme and procedure integrated into a dedicated QA and checklist
- Temperature up to 80-90 K and analysis of gas released (YETS 2017-2018)
  - Compromise between risks associated to bring a sector to room temperature and vacuum quality reached
- Few dump events in 2018 with gradual conditioning along the year
  - Mitigated with few hours in SB with 900 bunches

![](_page_17_Picture_0.jpeg)

![](_page_17_Picture_1.jpeg)

![](_page_17_Picture_2.jpeg)

# I. UFO II. ULO III.16L2 IV.10Hz V. Conclusions

![](_page_17_Picture_4.jpeg)

![](_page_18_Picture_0.jpeg)

### Introduction

![](_page_18_Picture_2.jpeg)

- 8 dumps with similar signature in 2018:
  - 1 during the p-p run
  - 7 during the Pb-Pb run (out of 48 physics fills)

![](_page_18_Figure_6.jpeg)

![](_page_19_Picture_0.jpeg)

0.0

-0.2

-0.4

-0.6

-0.8

-1.0 └── 18

20

22

![](_page_19_Picture_2.jpeg)

![](_page_19_Figure_3.jpeg)

![](_page_19_Figure_4.jpeg)

Kink and orbit displacement of both beams compatible with horizontal movement of MQ

At which point of the cycle do they start? ADT and 100Hz BLM at TCPs show that are always present

![](_page_19_Figure_7.jpeg)

No ITR2 cryostat movement measured by survey team

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Q21L4

Phase advance [2<sub>π</sub>]

Internal vibration? Origin at Q4-Q5?

First time that IP2 squeezed to 50 cm during Pb-Pb run: induced displacement amplified Why 21L4 during p-p run ( $\beta_x$ =180m)?

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![](_page_20_Picture_0.jpeg)

![](_page_20_Picture_2.jpeg)

- Crucial better understanding of the events:
  - What is the trigger of sudden amplitude increase?
  - ✓ Will the amplitude keep increasing?
- Present studies based on dump events and PM data:
  - ✓ More observations needed and study events not leading to dump
  - ✓ T-b-T BPM acquisition triggered when these oscillation are detected?
  - ✓ Accelerometer on cryostat of suspected MQs?
- **Dump** triggered by **losses** due to **beam halo** intercepted by TCPs: **possible mitigations**?

![](_page_20_Figure_11.jpeg)

![](_page_20_Picture_12.jpeg)

![](_page_21_Picture_0.jpeg)

![](_page_21_Picture_1.jpeg)

![](_page_21_Picture_2.jpeg)

# I. UFO II. ULO III. 16L2 IV. 10Hz V. Conclusions

![](_page_21_Picture_4.jpeg)

![](_page_22_Picture_0.jpeg)

### **Conclusions**

![](_page_22_Picture_2.jpeg)

#### Unidentified Falling Object:

Experimental validation of quench level models and improved BLM threshold strategy Hard to predict UFO rate right after LS2: in the range 2012-2015, if equivalent to LS1

#### Unidentified Lying Object:

Successfully mitigated during Run 2 by means of local orbit bump

Access mid-April: if something found, remove it. If nothing found, replace magnet?

#### • 16L2:

Significant limitation in 2017, successfully mitigated and target integrated lumi reached Most probable cause identified and actions taken to avoid repetition Should literally "evaporate" during LS2 with warm up to room temperature

#### • 10Hz:

Sound explanation found for events during 2018 Pb run

Still more to understand on the trigger and dump during p run (improved diagnostic?) Possible loss mitigation: crystal collimation and TCLD (Run 3?), hollow e-lens (Run 4?)

#### Team work: many people involved, many ideas, many actions taken and mitigations found!

![](_page_22_Picture_14.jpeg)

![](_page_23_Picture_0.jpeg)

![](_page_23_Picture_1.jpeg)

![](_page_23_Picture_2.jpeg)

# Thanks for your attention!

![](_page_23_Picture_4.jpeg)

![](_page_24_Picture_0.jpeg)

![](_page_24_Picture_1.jpeg)

![](_page_24_Picture_2.jpeg)

# BACKUP

![](_page_24_Picture_4.jpeg)

![](_page_25_Picture_0.jpeg)

![](_page_25_Picture_1.jpeg)

![](_page_25_Picture_2.jpeg)

## *UFO...*

![](_page_25_Picture_4.jpeg)

![](_page_26_Picture_0.jpeg)

![](_page_26_Figure_2.jpeg)

![](_page_26_Figure_3.jpeg)

![](_page_27_Picture_0.jpeg)

![](_page_27_Picture_2.jpeg)

#### $\Rightarrow$ more relevant for the likelihood of a quench $\rightarrow$ # of UFO events accummulated over time

![](_page_27_Figure_4.jpeg)

Only fills with >100b per beam, includes triggers on all Run 2 standard BLMs (RS04 detection threshold:  $2 \times 10^{-4}$  Gy/sec).

![](_page_28_Picture_0.jpeg)

![](_page_28_Picture_2.jpeg)

![](_page_28_Figure_3.jpeg)

Only fills with  $\geq$  1 h in STABLE and with > 100b per beam.

A. Lechner

![](_page_29_Picture_0.jpeg)

![](_page_29_Picture_2.jpeg)

#### UFO rate (cells $\geq$ 12) in stable beams, normalized to beam intensity:

![](_page_29_Figure_4.jpeg)

Only fills with  $\geq$ 1 h in STABLE and with >100b per beam, only BLMs common to Run 1 and Run 2.

![](_page_30_Picture_0.jpeg)

![](_page_30_Picture_2.jpeg)

A. Lechner

#### Yes (at least to a good approximation), if ...

- ... we use different detection thresholds to account for the different beam energy
- ... we focus on a subset of BLMs common to both runs

#### Run 1 BLM layout:

![](_page_30_Figure_8.jpeg)

![](_page_31_Picture_0.jpeg)

![](_page_31_Picture_1.jpeg)

![](_page_31_Picture_2.jpeg)

## *ULO...*

![](_page_31_Picture_4.jpeg)

# Parasitic monitoring of beam losses

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LHC Collimation

CERN

• Clear loss spikes (i.e. exp. decay and peak > 1e-6 Gy/s) looking at 1.3s BLM running sum

![](_page_32_Figure_2.jpeg)

![](_page_33_Picture_0.jpeg)

### **UFO at the ULO**

![](_page_33_Picture_2.jpeg)

![](_page_33_Figure_3.jpeg)

![](_page_33_Figure_4.jpeg)

![](_page_34_Picture_0.jpeg)

![](_page_34_Picture_1.jpeg)

![](_page_34_Picture_2.jpeg)

## *16L2...*

![](_page_34_Picture_4.jpeg)

![](_page_35_Picture_0.jpeg)

![](_page_35_Figure_1.jpeg)

![](_page_35_Figure_2.jpeg)

LHC Collimation

Project

CERN

![](_page_36_Picture_0.jpeg)

![](_page_36_Picture_2.jpeg)

• First dump event: First fill with 72b trains during Scrubbing run (1236b)

![](_page_36_Figure_4.jpeg)

#### Fill 5725, 29<sup>th</sup> of May

- <u>First action taken right after this event:</u>
  - ✓ Local aperture measurements → No evident aperture restriction found
- *First hints on the source of this event:*

![](_page_37_Picture_0.jpeg)

# **Mitigation using MCB field**

![](_page_37_Picture_2.jpeg)

- Action taken the 20<sup>th</sup> of July:
  - ✓ Operational bump to set I<sub>MCB</sub> > 2.5 A
  - ✓ MCB removed from orbit feedback
    - Smooth 25 ns operations restored!

![](_page_37_Figure_7.jpeg)

![](_page_37_Figure_8.jpeg)

![](_page_38_Picture_0.jpeg)

![](_page_38_Picture_2.jpeg)

- Two warm up of the beam screen only, performed between 10<sup>th</sup> and 13<sup>th</sup> of August
  - Abnormal increase of pressure observed by additional gauges installed (up to 2e-3 mbar!)
- Main effects:
- ✓ Change of I<sub>MCB</sub> no longer sufficient
- ✓ Steady losses reduced but significant spikes
- Clear correlation between spikes and dumps
- Reduced intensity to get to Stable Beams
- Possible explanation:
- 1. Accidental air venting during at the end of EYETS
- 2. Part of the evaporated gas condensed in cold spots in the interconnection area during beam screen thermal regeneration

![](_page_38_Figure_13.jpeg)

![](_page_38_Figure_14.jpeg)

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![](_page_39_Figure_0.jpeg)

![](_page_40_Picture_0.jpeg)

### Latest events

![](_page_40_Picture_2.jpeg)

Operational bump to change of I<sub>MCB</sub> not incorporated in the 2.5 TeV cycle **Confirmation impossible** I<sub>MCB</sub> of Beam 2 at ~ 0 A Dumps occurred in Beam 2 to run with  $I_{\text{MCB}} \sim 0 \text{ A}$ Day with  $\geq 1$  dump by 16L2 ⊐2800 జె 1.5 Del. Int. Lumi [fb<sup>-1</sup>] 06 . . 2600 und 2400 # 22001.3 2000 1.2 1800 30 1600 1.11400 20 1200 1000 10 0.9 800 Special physics. MD4 600 0.8 13/11 04/12 20/11 27/11 time [day/month] **Limited time during MD4** to perform systematic tests on the effect of the solenoid: Single fill with physics beams but solenoid OFF (8b4e BCS, 1868b, 1.25e11 ppb)

Dump during ramp confirmed beneficial effect of solenoid on intensity threshold

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![](_page_41_Picture_0.jpeg)

# **YETS plans**

![](_page_41_Picture_2.jpeg)

- Initial plan: bring all sector 12 to room temperature
- First action: temperature increased up to 80-90 K (including cold mass) and analysis of the gas released

Confirmed contamination by atmospheric air

![](_page_41_Figure_6.jpeg)

- Estimated air entered: 7 bar | per beam pipe > 99.96% pumped > Amount of gas reduced by 10<sup>5</sup>
- Estimated amount of water vapor: 0.1 g per beam pipe Condensed on the pumping lines
- Warm up stopped:

<u>Compromise between risks associated to bring a sector to room temperature</u> <u>and vacuum quality reached</u>

![](_page_42_Picture_0.jpeg)

# **Avoid repetition**

Most probable scenario:

<u>Accidental air venting,</u> <u>during the removal of the mobile group for the final pinch-off</u>

• New pumping scheme: All the pump groups installed in positions equipped with vacuum gauges

Detection of air entering during mechanical activities

Pumping port upgrade:

![](_page_42_Picture_7.jpeg)

![](_page_42_Picture_8.jpeg)

#### Removal of pumping groups before cool down:

<u>Quick re-pumping in case of air leak,</u> <u>no need of magnet warm up</u> LHC Collimation

Project

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![](_page_43_Picture_0.jpeg)

### **Intensity Ramp up 2018**

![](_page_43_Picture_2.jpeg)

![](_page_43_Figure_3.jpeg)

![](_page_43_Picture_4.jpeg)

# Intensity Ramp up 2018 (normalized)

![](_page_44_Picture_1.jpeg)

![](_page_44_Figure_2.jpeg)

#### Very early activity w.r.t. 2017 but clear conditioning after scrubbing

![](_page_44_Picture_4.jpeg)

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![](_page_45_Picture_0.jpeg)

## Full Machine until TS1 (2018)

![](_page_45_Picture_2.jpeg)

10 9 ppb (× 1e11) Spikes B1 Spikes B2 8 Dump B1 Dump B2 7 6 5 4 3 2 1.12 1.11 1.17 1.11 I:12 F. 19 19 19 Ξ. 1.11 .... 60.1 1 0 6660 6680 6700 6720 6740 6760 Fill number 3 dump events in B1 followed by 4 "thunderstorm" Cured after fill 6746 with 900b recovery fill MANCHESTER D. Mirarchi

Average bunch intensity taken at Injection.

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![](_page_46_Picture_0.jpeg)

![](_page_46_Picture_2.jpeg)

![](_page_46_Figure_3.jpeg)

![](_page_46_Picture_4.jpeg)

![](_page_47_Picture_0.jpeg)

![](_page_47_Picture_2.jpeg)

![](_page_47_Figure_3.jpeg)

![](_page_47_Picture_4.jpeg)

![](_page_48_Picture_0.jpeg)

## From TS2 to end of Run2 p-p run (2018)

![](_page_48_Picture_2.jpeg)

![](_page_48_Figure_3.jpeg)

#### No dump events, only one single spike in B1

![](_page_48_Picture_5.jpeg)

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![](_page_49_Picture_0.jpeg)

![](_page_49_Picture_1.jpeg)

![](_page_49_Picture_2.jpeg)

- Very early activity w.r.t. 2017 obseved during intensity ramp up (first observation in 2017 during scrubbing run, in 2018 already visible with 75b)
- Clear conditioning after scrubbing and intensity ramp up
- Still 7 dumps caused between end on intensity ramp up and TS1 (4 dumps caused by following "thunderstorm")
- **Mitigation strategy** of the "thunderstorm" **of recovery fill with 900b** worked from fill 6746
- Two dump events in B1 between TS1 and MD2 without following dumps due to "thunderstorm"
- One dump events in B1 between MD2 and TS2 without following dumps due to "thunderstorm"
- No dump and only one single spike in B1 from TS2 to end of Run 2 p-p run

![](_page_49_Picture_10.jpeg)

![](_page_50_Picture_0.jpeg)

![](_page_50_Picture_1.jpeg)

![](_page_50_Picture_2.jpeg)

## 10Hz...

![](_page_50_Picture_4.jpeg)

![](_page_51_Picture_0.jpeg)

# Fill 7235 / pp

![](_page_51_Picture_2.jpeg)

- Extreme peak-to-peak trajectory on B1 and B2, ratio B1/B2 ~2.4.
  - A kink is visible @ 21L4 for B1.
  - MICADO points to cell 21L4 for both B1 and B2, kick ~ 1 μrad @ MQ (on p2p traj excursion), <u>consistent between B1 and B2</u> (also sign !).

![](_page_51_Figure_6.jpeg)

J. Wenninger

![](_page_51_Figure_8.jpeg)

![](_page_51_Picture_9.jpeg)

![](_page_52_Picture_0.jpeg)

# Fill 7235 / pp kink

LHC Collimation Project

• Kink @ 21L4 clearly visible on normalized position of B1 and B2.

![](_page_52_Figure_4.jpeg)

![](_page_52_Picture_5.jpeg)

![](_page_53_Picture_0.jpeg)

![](_page_53_Picture_2.jpeg)

- Time evolution: estimate period ~1300 turns ~ 9 Hz.
- Ratio B1/B2 ~ 2.44, matches the ratio of  $\sqrt{\beta_{x1}}/\sqrt{\beta_{x2}}$  for the H plane of cell 21L4.
- Consistent with a horizontal movement of the MQ in cell 21L4.
- □ Kick of 1  $\mu$ rad  $\rightarrow$  37  $\mu$ m quad displacement (p2p).
  - ~20  $\mu$ m oscillation amplitude.

![](_page_53_Figure_8.jpeg)

J. Wenninger

![](_page_53_Picture_10.jpeg)

![](_page_54_Picture_0.jpeg)

![](_page_54_Picture_2.jpeg)

- Triplet modes at ~8, 11 and 12 Hz in horizontal plane.
- Fits rather well the observed frequencies...

![](_page_54_Figure_5.jpeg)

J. Wenninger

![](_page_54_Picture_7.jpeg)

![](_page_55_Picture_0.jpeg)

### **Raw Oscillation**

![](_page_55_Picture_2.jpeg)

![](_page_55_Figure_3.jpeg)

(b) Ion fill 7447

Figure: Horizontal beam displacement observed in all cases

X. Buffat and A. Oeftiger

![](_page_55_Picture_7.jpeg)

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![](_page_56_Picture_0.jpeg)

### **BPM Spectrum**

![](_page_56_Picture_2.jpeg)

#### lons 50cm

![](_page_56_Figure_4.jpeg)

Figure: Beam 1 Frequency Spectrum of horizontal BPMs.

Arrange δριβ 8794.101.1.82 8794.101.2.82 8794.101.3.82 8794.101.4.82 8794.101.4.82 8794.101.4.82 8794.101.4.82 8794.101.8.82 8794.101.0.8.82 899719.54.82 899719.54.82 899719.548.82 899719.548.82 89971.64.82 8997.4.792.82 8997.4.792.82 8997.4.792.82 8997.4.792.82

Figure: Beam 2 Frequency Spectrum of horizontal BPMs.

- Lines around 10Hz are visible in the horizontal spectrum
- no lines can be found in the vertical spectrum
- often two lines at around 10Hz and 12Hz
- the **12Hz line** has usually **higher amplitude**
- sometimes only one line around 11Hz
- these lines are visible for both, lons and Protons
- they are more distinct for lon optics, especially in Beam 2
- Injection does not show clear lines.

![](_page_56_Picture_16.jpeg)

![](_page_57_Picture_0.jpeg)

### **100Hz BLM FFT**

![](_page_57_Figure_2.jpeg)

![](_page_57_Figure_3.jpeg)

![](_page_57_Figure_4.jpeg)

![](_page_57_Picture_5.jpeg)

![](_page_58_Picture_0.jpeg)

![](_page_58_Picture_1.jpeg)

![](_page_58_Picture_2.jpeg)

# Heat Load VS Losses...

![](_page_58_Picture_4.jpeg)

![](_page_59_Picture_0.jpeg)

![](_page_59_Picture_1.jpeg)

![](_page_59_Picture_2.jpeg)

![](_page_59_Figure_3.jpeg)

![](_page_59_Picture_4.jpeg)

### **50 ns**

![](_page_60_Picture_1.jpeg)

![](_page_60_Picture_2.jpeg)

![](_page_60_Figure_3.jpeg)

![](_page_60_Picture_4.jpeg)