

WIMP dark matter in the parity solution to strong CP problem

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Based on the work (arXiv:1812.07004)
with J. Kawamura (*Ohio State U.*), S. Okawa (*U. of Victoria*), Y. Tang (*U. of Tokyo*).

Introduction

SM succeeds in many experiments,
but still we cannot be satisfied because of
a lot of “why”, e.g.

Why is Higgs mass tachyonic and $O(100)$ GeV?

Why are there three generations?

Why is the gauge symmetry like that?

Why is parity broken?

Why is θ -term so small? (strong CP problem)

cosmological observation

dark matter, Baryogenesis, etc.

Introduction

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My proposal

parity symmetric
model

cosmological observation

dark matter, Baryogenesis, etc.

WIMP DM

What is strong CP problem?

CP is explicitly broken in the SM, so that namely θ -term is allowed.

$$\theta \frac{g_s^2}{32\pi^2} F_{\mu\nu}^a \tilde{F}^{a\mu\nu}$$

To avoid the experimental bound, θ should be tiny:

$$|\theta| \lesssim 10^{-10}$$

Why is it so small?

well-known solutions

$$\theta \frac{g_s^2}{32\pi^2} F_{\mu\nu}^a \tilde{F}^{a\mu\nu}$$



global $U(1)_{PQ}$ symmetry

$$a(x) \frac{g_s^2}{32\pi^2} F_{\mu\nu}^a \tilde{F}^{a\mu\nu}$$

axion

Peccei, Quinn '77; Weinberg '77; Wilczek '77

parity is respected

θ -term is forbidden

Babu, Mohapatra '90; etc.

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My talk

Point of the model building

- The SM gauge symmetry is

$$SU(3)_c \times SU(2)_L \times U(1)_Y$$

$$Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix} \quad \begin{matrix} u_R \\ d_R \end{matrix}$$

We have to extend or introduce extra symmetries and matters, in order to respect parity.

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We have to extend or introduce extra symmetries and matters, in order to respect parity.

- We may need more fields to avoid conflict with experimental results.

There are so many fields, but some of extra fields could be good DM candidates.

Content

- Setup
- Introduction of DM
- Phenomenology
- Summary

Setup

Parity transformation

For Dirac fermion (q),

$$P q(t, x) P = \gamma_0 q(t, -x) \iff \begin{pmatrix} q_R(t, x) \\ q_L(t, x) \end{pmatrix} \rightarrow \begin{pmatrix} q_L(t, -x) \\ q_R(t, -x) \end{pmatrix}$$

LR exchanging

Parity transformation

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LR exchanging

In the SM

SU(2)_L doublet

$$Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}$$

SU(2)_L singlet

$$u_R \quad d_R$$

need extra something to make parity symmetric

Our extension to realize LR exchanging

SU(2)_L doublet

$$Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}$$

SU(2)_L singlet

$$u_R \quad d_R$$



parity transformation



extra something

SU(2)_R doublet

$$Q'_R = \begin{pmatrix} u'_R \\ d'_R \end{pmatrix}$$

SU(2)_R singlet

$$u'_L \quad d'_L$$

gauge symmetry

$$SU(3)_c \times SU(2)_L \times U(1)_L \quad \times SU(2)_R \times U(1)_R$$

$$A_L^\mu$$



$$A_{R\mu}$$

parity transformation

gauge symmetry

$$SU(3)_c \times SU(2)_L \times U(1)_L \quad \times SU(2)_R \times U(1)_R$$

$$A_L^\mu$$



$$A_{R\mu}$$

parity transformation

Higgs fields

$$H_L$$



$$H_R$$

parity transformation

matter content

“SM side”

Fields	spin	$SU(3)_c$	$SU(2)_L$	$SU(2)_R$	$U(1)_R$	$U(1)_L$
Q_L^i	1/2	3	2	1	0	1/6
u_R^i	1/2	3	1	1	0	2/3
d_R^i	1/2	3	1	1	0	-1/3
l_L^i	1/2	1	2	1	0	-1/2
e_R^i	1/2	1	1	1	0	-1
H_L	0	1	2	1	0	1/2

“Mirror side”



parity transformation

Fields	spin	$SU(3)_c$	$SU(2)_L$	$SU(2)_R$	$U(1)_R$	$U(1)_L$
Q_R^i	1/2	3	1	2	1/6	0
u_L^i	1/2	3	1	1	2/3	0
d_L^i	1/2	3	1	1	-1/3	0
l_R^i	1/2	1	1	2	-1/2	0
e_L^i	1/2	1	1	1	-1	0
H_R	0	1	1	2	1/2	0

Yukawa couplings

$$\mathcal{L}_Y = -Y_d^{ij} \overline{Q_L^i} H_L d_R^j - Y_u^{ij} \overline{Q_L^i} \tilde{H}_L u_R^j - Y_e^{ij} \overline{l_L^i} H_L e_R^j + h.c.$$
$$-Y_d^{ij} \overline{Q_R^i} H_R d_L^j - Y_u^{ij} \overline{Q_R^i} \tilde{H}_R u_L^j - Y_e^{ij} \overline{l_R^i} H_R e_L^j + h.c.$$

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$$-Y_d^{ij} \overline{Q_R'^i} H_R d_L'^j - Y_u^{ij} \overline{Q_R'^i} \tilde{H}_R u_L'^j - Y_e^{ij} \overline{l_R'^i} H_R e_L'^j + h.c.$$

problem 1

Mirror fermions should be heavy $\rightarrow \langle H_L \rangle \ll \langle H_R \rangle$

Yukawa couplings

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problem 1

Mirror fermions should be heavy $\rightarrow \langle H_L \rangle \ll \langle H_R \rangle$

problem 2

Stable colored particles appear

\rightarrow introduce extra scalars, X_b and X_l with

$$\lambda_u^{ij} X_b \overline{u_R^i} u_L'^j + \lambda_e^{ij} X_l \overline{e_R^i} e_L'^j$$

The detail of X_b and X_l

charge assignment

Fields	spin	$U(1)_R$	$U(1)_L$
X_b	0	$-2/3$	$2/3$
X_l	0	1	-1

They are not charged under the other symmetries.

Coupling with quarks and electrons

$$\lambda_u^{ij} X_b \overline{u_R^i} u_L^{\prime j} + \lambda_e^{ij} X_l \overline{e_R^i} e_L^{\prime j}$$

Mirror quarks decay through these couplings:

$$u_j' \rightarrow u_i X_b, \quad e_j' \rightarrow e_i X_l$$

We can consider two cases:

$$(I) \langle X_b \rangle = 0 \text{ and } \langle X_l \rangle \neq 0$$

$$(II) \langle X_b \rangle \neq 0 \text{ and } \langle X_l \rangle = 0$$

For instance

$$(I) \langle X_b \rangle = 0 \text{ and } \langle X_l \rangle \neq 0$$

$$\cdot \text{SU}(3)_c \times \text{SU}(2)_L \times \text{U}(1)_L \times \text{SU}(2)_R \times \text{U}(1)_R$$

$$\rightarrow \text{SU}(3)_c \times \text{SU}(2)_L \times \text{U}(1)_L \times \text{U}(1)' \rightarrow \text{SU}(3)_c \times \text{SU}(2)_L \times \text{U}(1)_Y$$

$$\langle H_R \rangle \neq 0$$

$$\langle X_l \rangle \neq 0$$

For instance

$$(I) \langle X_b \rangle = 0 \text{ and } \langle X_l \rangle \neq 0$$

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$$\langle H_R \rangle \neq 0$$

$$\langle X_l \rangle \neq 0$$

- X_b is neutral under the SM gauge symmetry and stable because of the remnant discrete symmetry.

X_b is a good DM candidate that couples to up-type quarks

“baryonic DM”

For instance

$$(II) \langle X_b \rangle \neq 0 \text{ and } \langle X_l \rangle = 0$$

$$\cdot SU(3)_c \times SU(2)_L \times U(1)_L \times SU(2)_R \times U(1)_R$$

$$\rightarrow SU(3)_c \times SU(2)_L \times U(1)_L \times U(1)' \rightarrow SU(3)_c \times SU(2)_L \times U(1)_Y$$

$$\langle H_R \rangle \neq 0$$

$$\langle X_b \rangle \neq 0$$

- X_l is neutral under the SM gauge symmetry and stable because of the remnant discrete symmetry.

X_l is a good DM candidate that couples to leptons

“leptonic DM”

Phenomenology

We can explicitly calculate most parts.

The relevant free parameters are the Yukawa couplings:

$$\lambda_u^{ij} X_b \overline{u_R^i} u_L'^j + \lambda_e^{ij} X_l \overline{e_R^i} e_L'^j$$

These couplings contribute to

flavor physics (D - \bar{D} mixing, $t \rightarrow qV$, $\mu \rightarrow e \gamma$, etc.)

LHC physics ($pp \rightarrow jj + \text{missing}$, $ll + \text{missing}$, etc.)

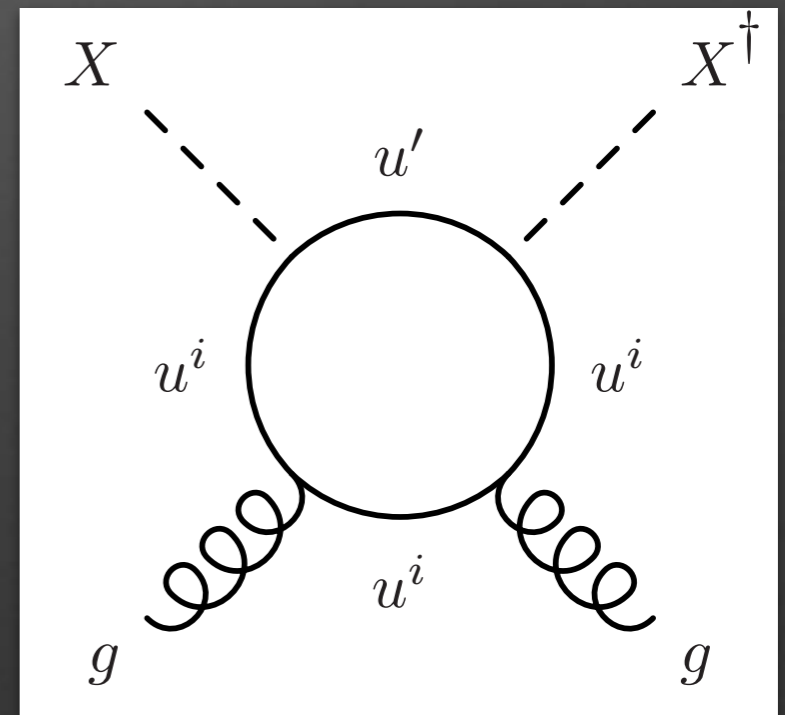
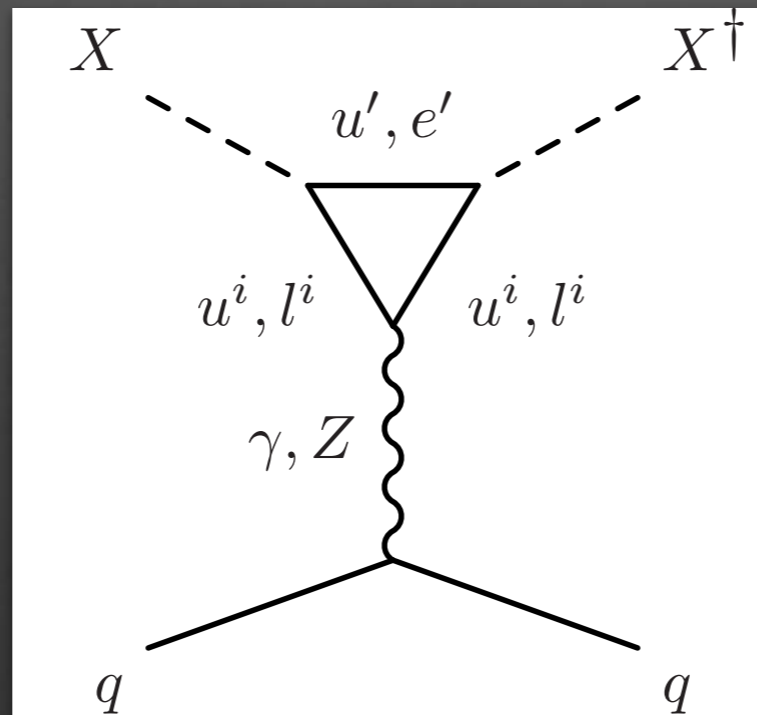
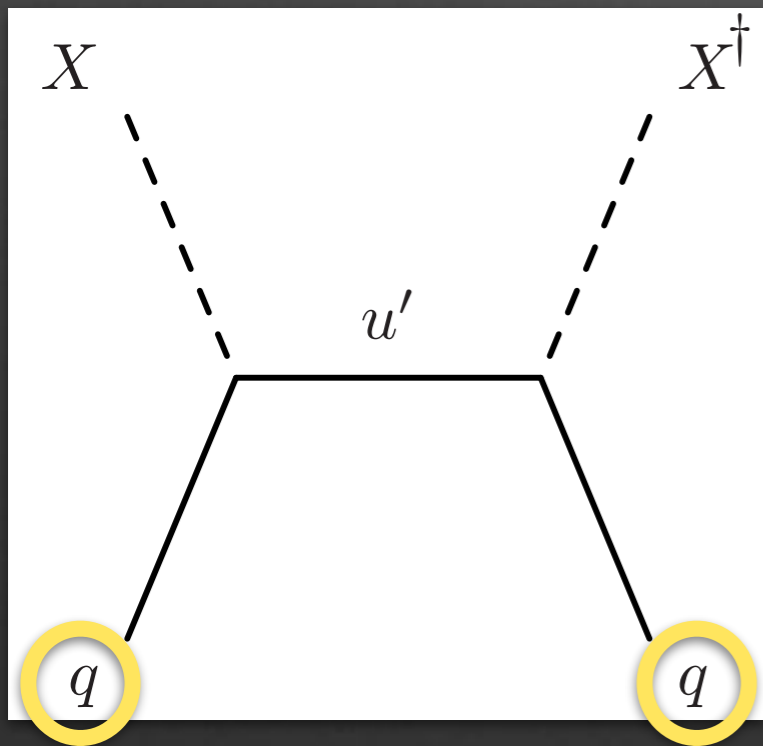
dark matter physics

In particular, DM physics strongly constrains this setup.

Dark Matter Physics

$$\lambda_u^{ij} X_b \overline{u_R^i} u_L'^j + \lambda_e^{ij} X_l \overline{e_R^i} e_L'^j$$

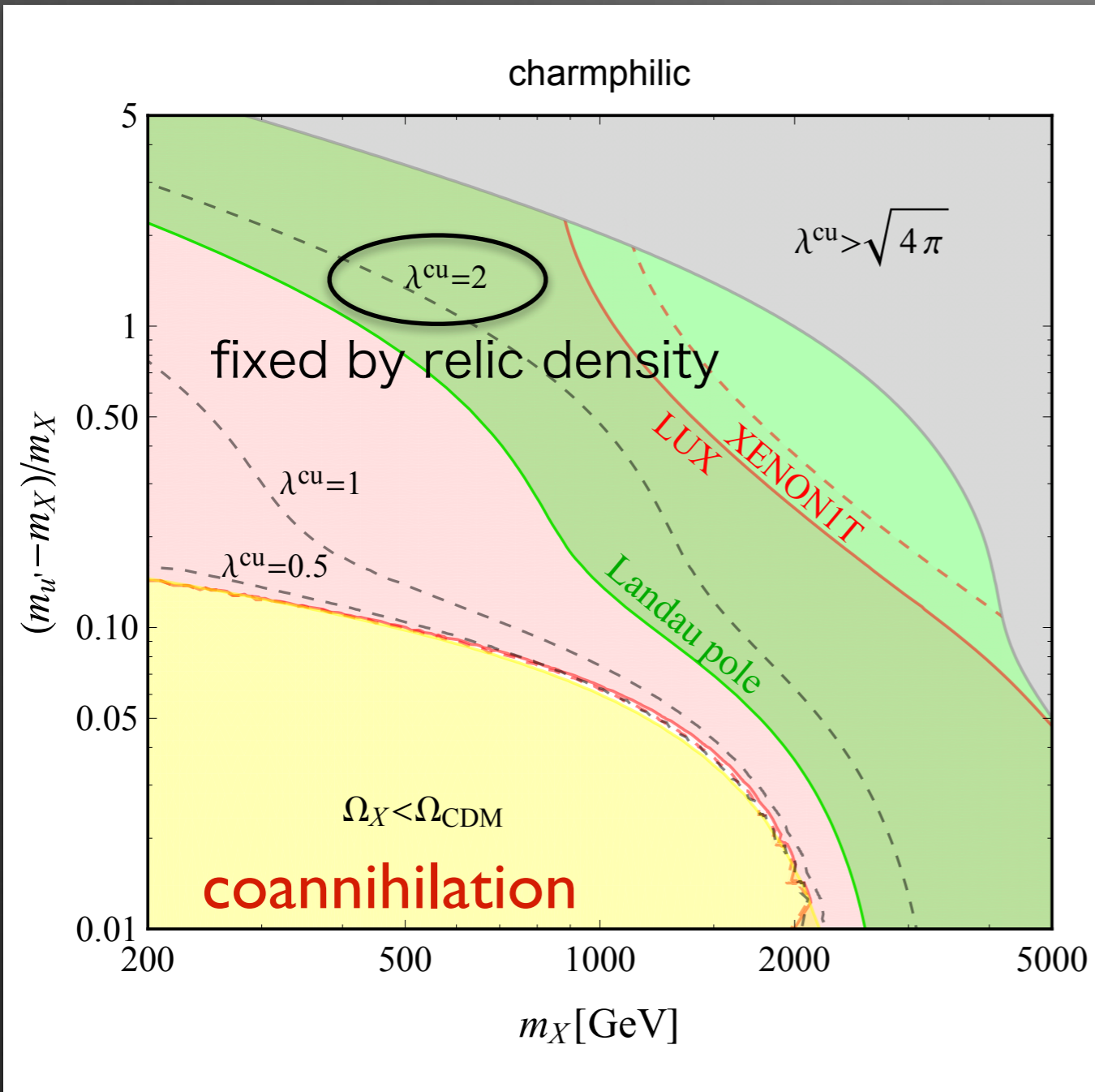
contribute to annihilation and direct detection:



flavor depends on the Yukawa alignment, λ_u

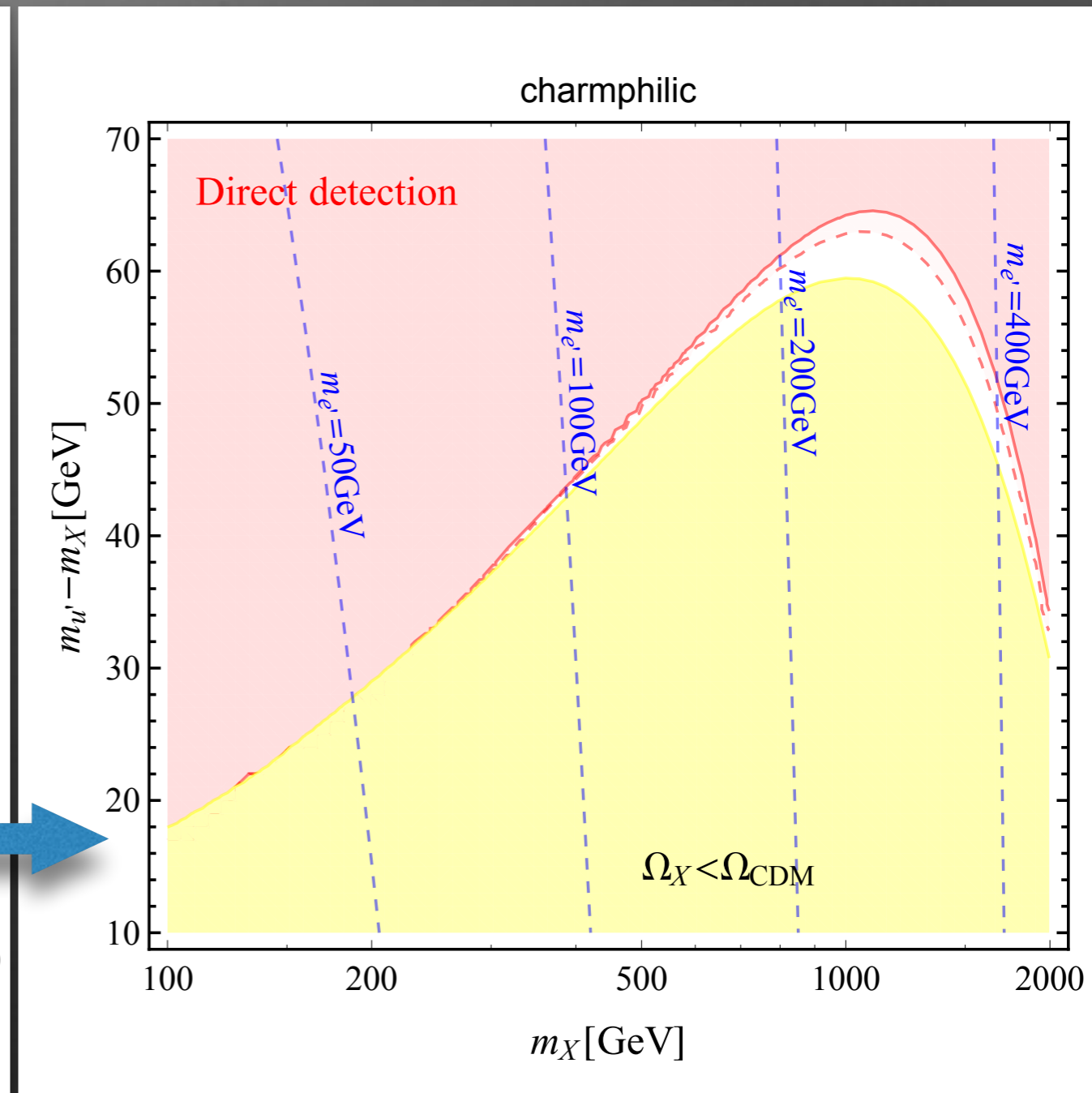
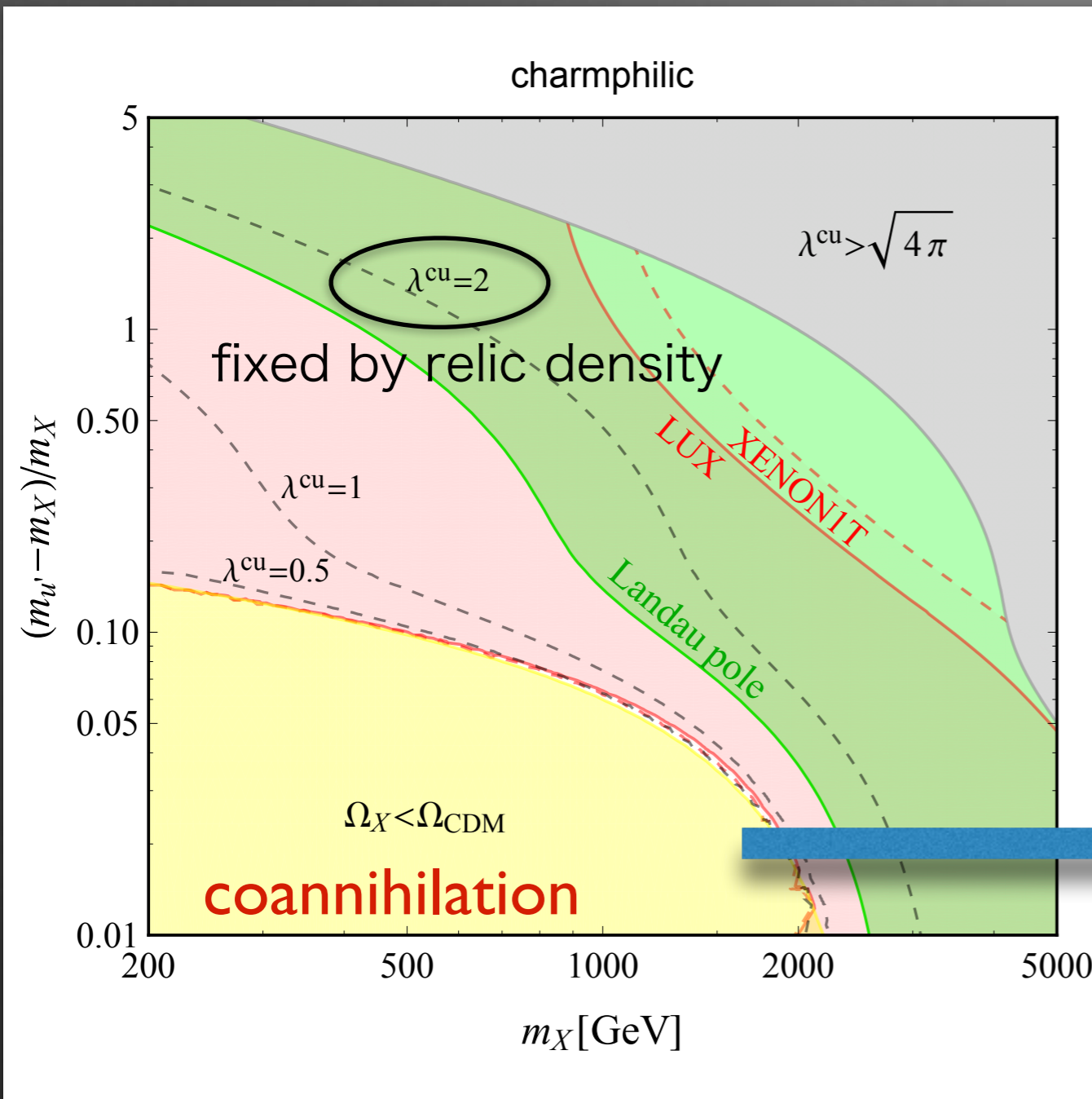
Allowed region in baryonic DM (X_b)

(I) Assume DM dominantly couples to **charm quark** and mirror up quark
($u' \rightarrow c X_b$)



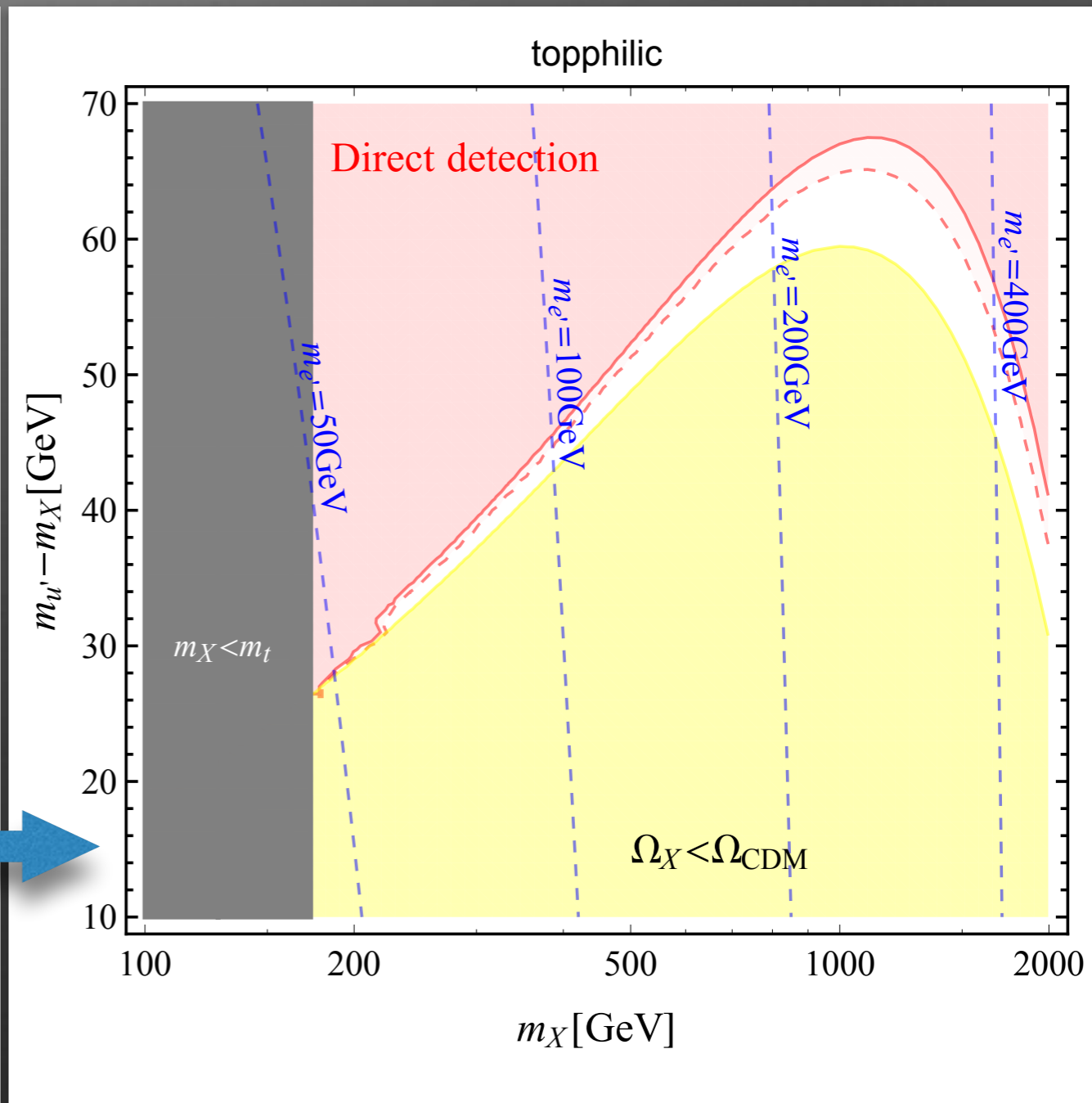
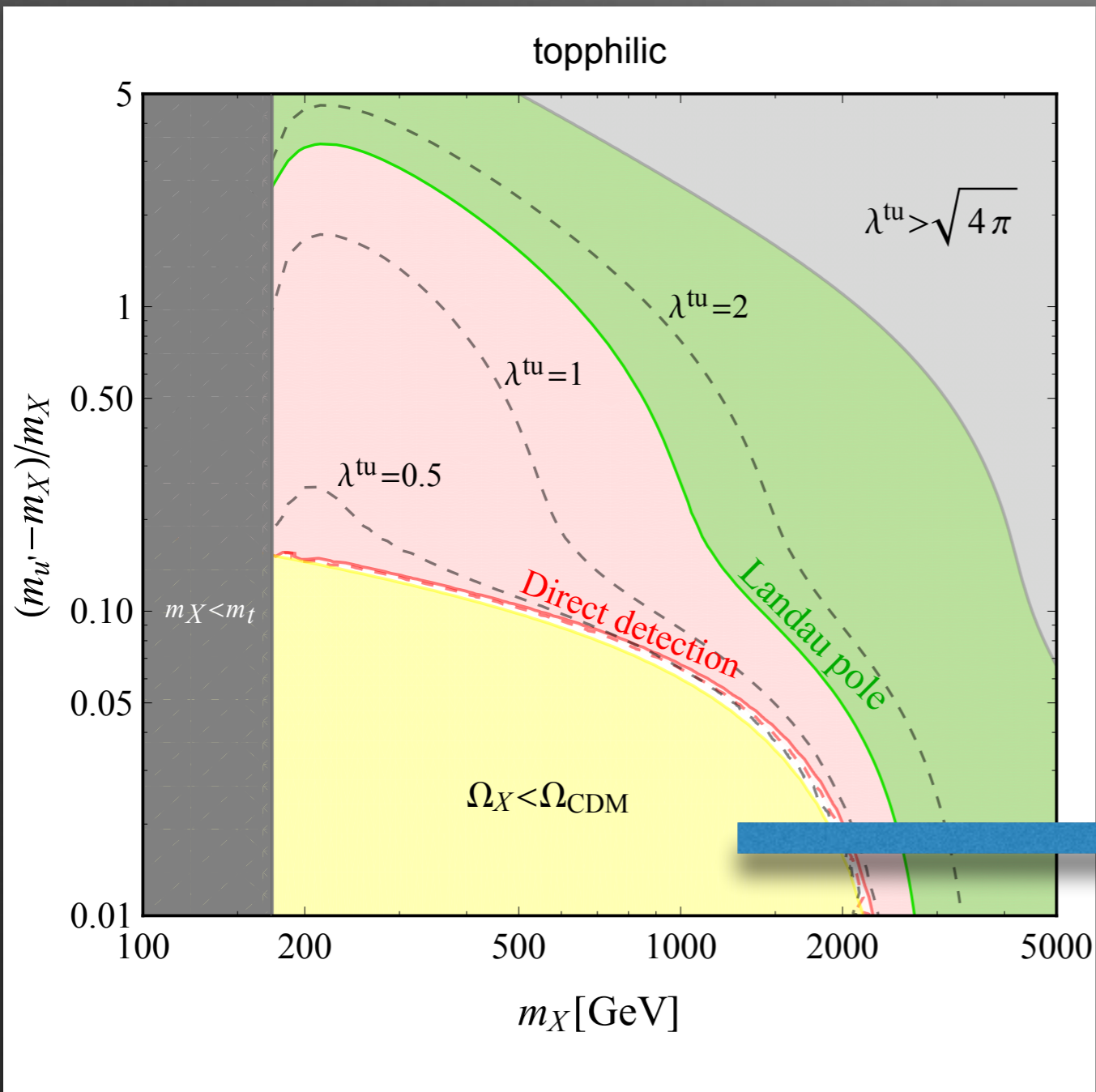
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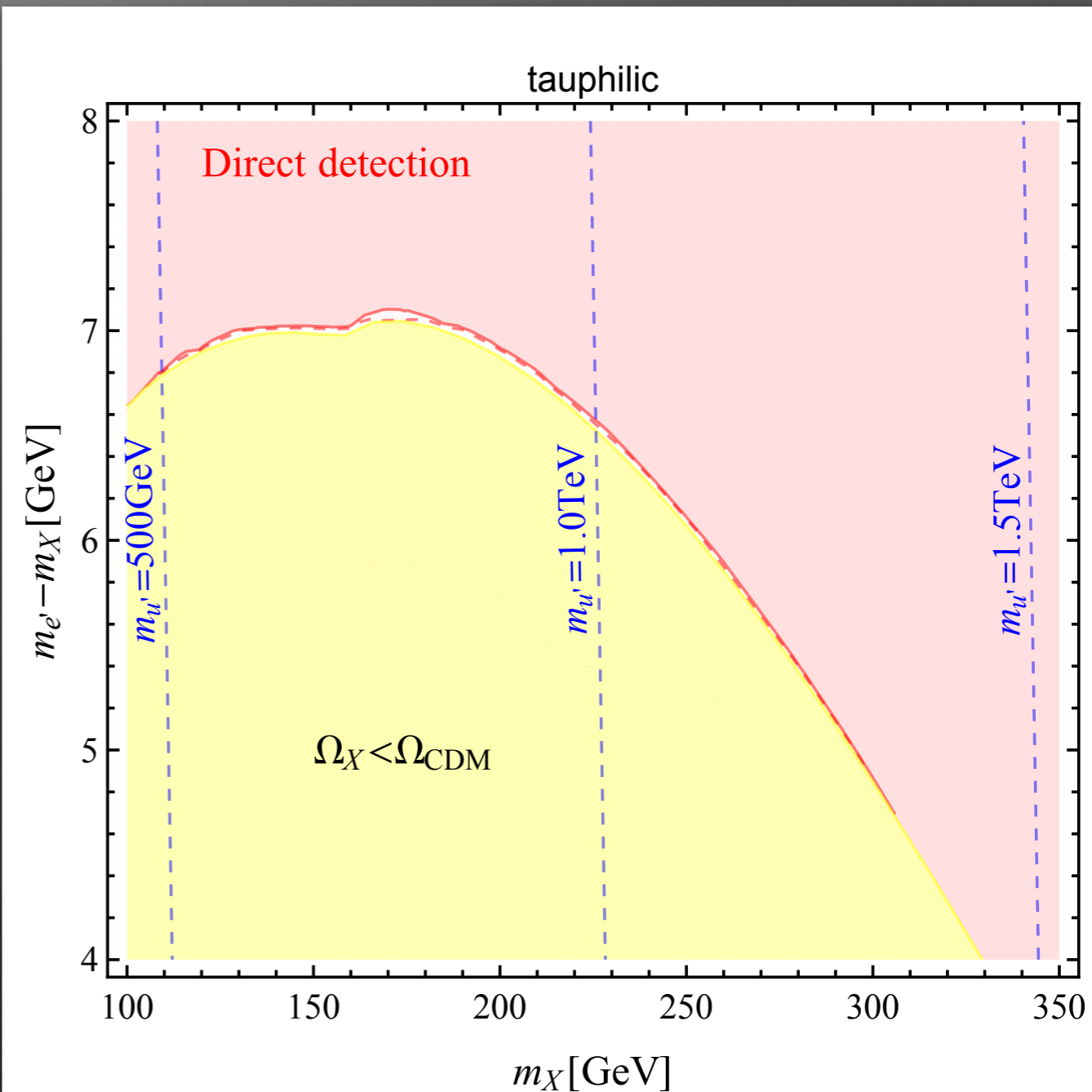
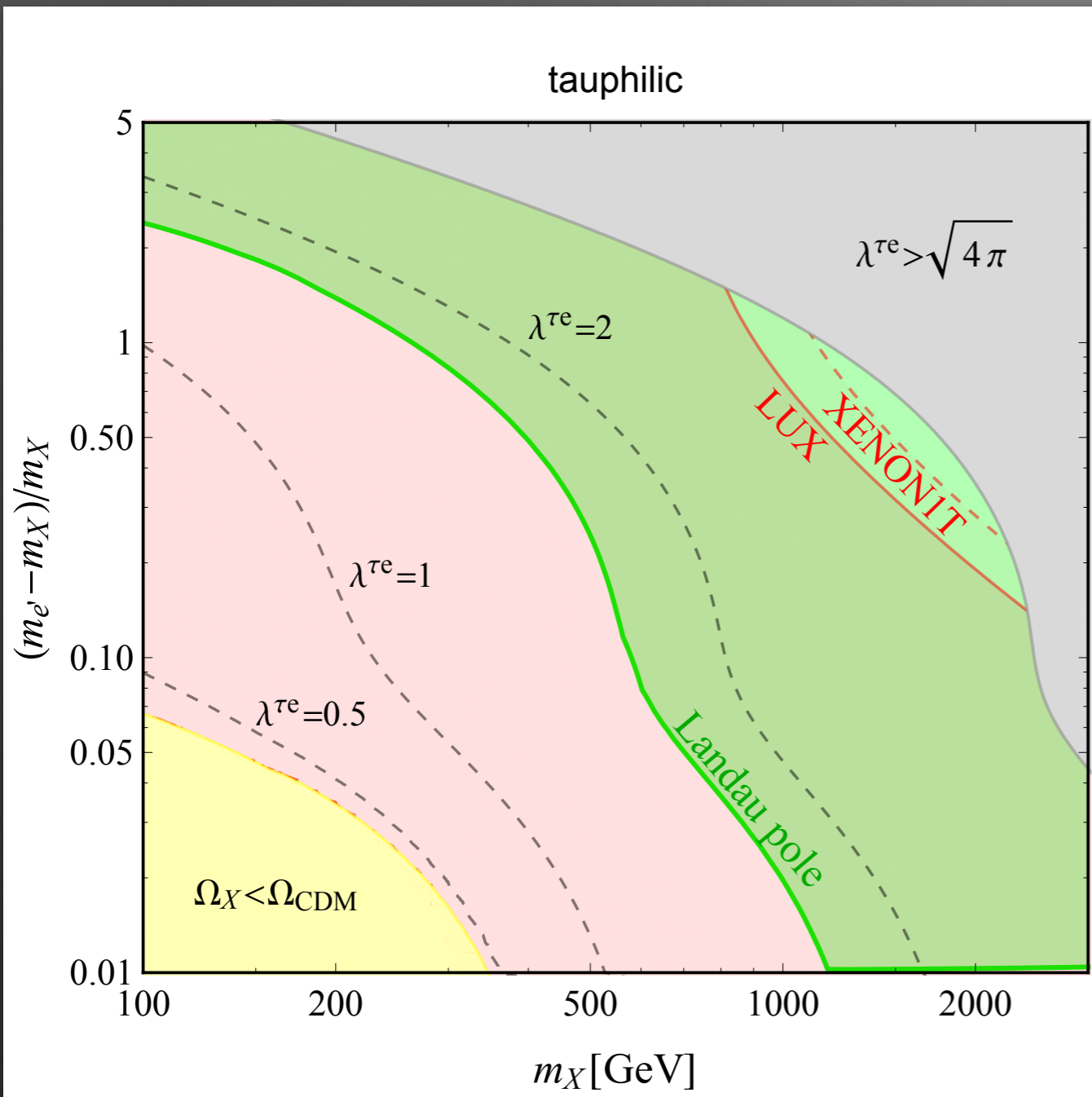
Allowed region in baryonic DM (X_b)

(2) DM dominantly couples to **top quark** and mirror up quark
 ($u' \rightarrow t X_b$)



Allowed region in leptonic DM (X_1)

(3) DM dominantly couples to τ and mirror electron



Summary

- We propose one model with parity symmetry, motivated by *the strong CP problem*.

One well-known simple setup

$$SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix} \quad Q_R = \begin{pmatrix} u_R \\ d_R \end{pmatrix}$$

Babu, Mohapatra, '89

We have to discuss how to construct the realistic fermion mass spectrum and how to avoid avoid the tree-level FCNC.

- We propose one model with parity symmetry **introducing the mirror sector**, motivated by *the strong CP problem*.

1812.07004, Kawamura, Okawa, YO, Tang

- Yukawa couplings are the same as the SM.
- Some mirror (colored) charged particles become stable, so we introduce extra scalars, X_b and X_l .
- X_b/X_l is a baryonic/leptonic DM candidate.
- DM physics strongly constrains this model, as far as DM is thermally produced.
- LHC physics is also important ($pp \rightarrow e'_+ e'_- \rightarrow l+l-$ missing etc.).
- Neutrino physics would have rich phenomenology.

1812.07004, Kawamura, Okawa, YO, Tang
and work in progress

END