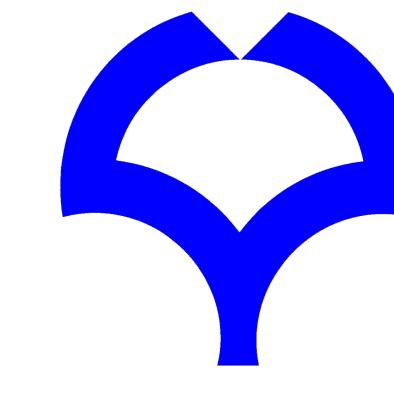


# LFV decays of the Higgs boson and dark matter in a new model for Dirac neutrino masses

Based on the work in progress by KE, S. Kanemura, K. Sakurai and H. Sugiyama



Osaka U.  
Kazuki Enomoto

## Summary

In [7], it was shown that if the signal of  $h \rightarrow \mu\tau$  is observed in future collider experiments without detecting the signal of  $\tau \rightarrow \mu\gamma$ , almost of all models generating neutrino masses are excluded. We construct a new model for neutrino masses and dark matter which can be survive in such a case.

## 1. Models generating neutrino masses

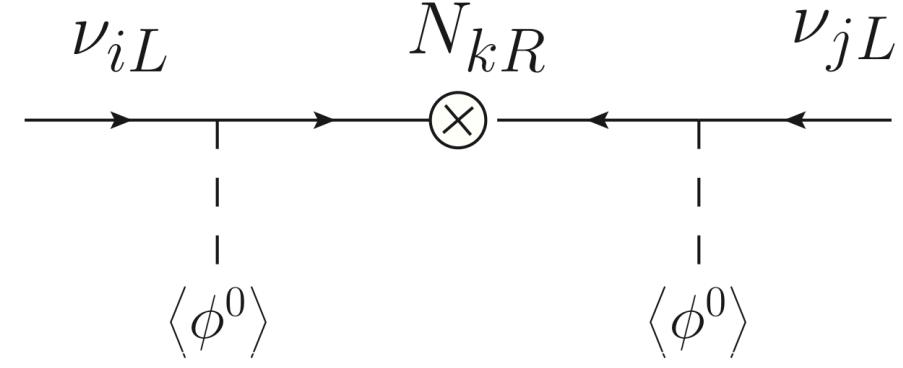
### Seesaw mechanisms

	Type-I	Type-II	Type-III
New field	$N_{iR}$	$\Delta$	$\Sigma_{iR}$
Spin	1/2	0	1/2
SU(2) <sub>L</sub>	<b>1</b>	<b>3</b>	<b>3</b>
U(1) <sub>Y</sub>	0	1	0

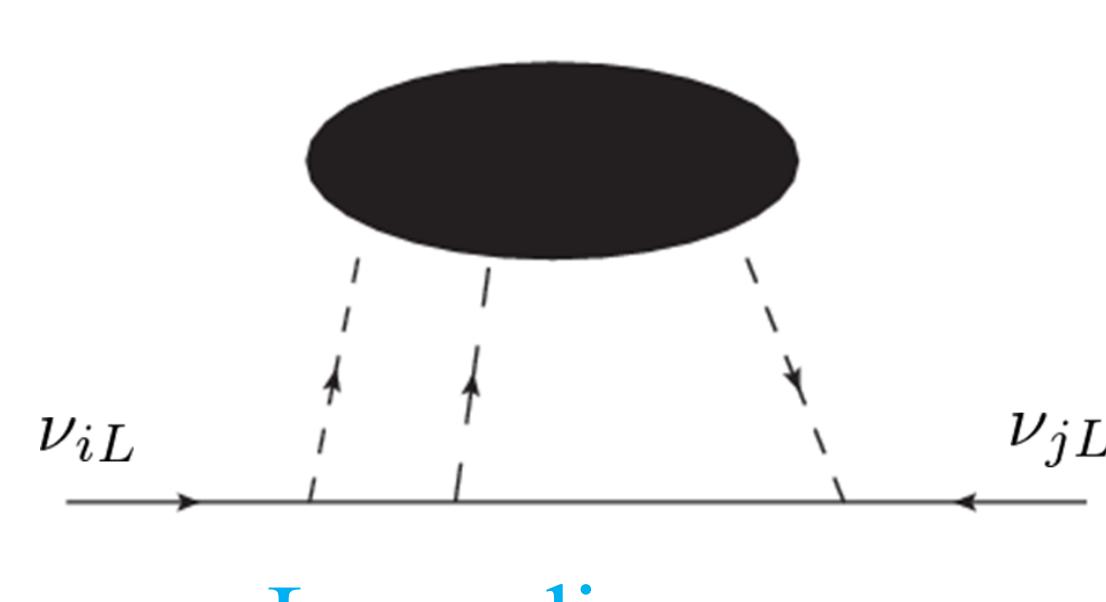
Table1. Seesaw mechanisms ( $i = 1 \sim 3$ )

Majorana masses are generated by the [tree diagram](#)

e.g., Type-I



### Models which generate neutrino masses radiatively [1]



Loop diagram

Some models can also explain DM by a new unbroken  $Z_2$  symmetry [2].

Some models generate Dirac type neutrino masses [3,4]

### Classification of simple models [5,6]

In the case of Majorana neutrinos [5]

#### Assume

1. w/o new sym.

No new fermions.

2. w/ an unbroken  $Z_2$  sym.

	SU(2) <sub>L</sub>	U(1) <sub>Y</sub>	$Z_2$
$\psi_{\alpha R}$	1	0	Odd

( $\alpha = 1 \sim 3$ )

	$\Delta$	$\phi_2$	$s^{++}$	$s_1^+$	$\eta$	$s_2^+$		$e_i \rightarrow e_j \gamma$
SU(2) <sub>L</sub>	<b>3</b>	<b>2</b>	<b>1</b>	<b>1</b>	<b>2</b>	<b>1</b>		
U(1) <sub>Y</sub>	+1	+1/2	+2	+1	+1/2	+1		
$Z_2$			Even			Odd		
M1			○	○				1 1
M2		○	○					1
M3			○					1
M4	○							1
M5				○		○		1 1
M6		○				○		1
M7					○			1
M8				○				1

Table1. Classification of models which generate Majorana neutrino masses

In the case of Dirac neutrinos [6]

	SU(2) <sub>L</sub>	U(1) <sub>Y</sub>	$Z_2$	$Z'_2$
$\nu_{iR}$	<b>1</b>	0	Even	Odd

$Z'_2$  : Softly broken  $Z_2$  sym.

	$\Delta$	$\phi_2$	$\phi_\nu$	$s^{++}$	$s_L^+$	$s_R^+$	$s^0$	$\eta$	$s_2^+$	$s_2^0$	$e_i \rightarrow e_j \gamma$
SU(2) <sub>L</sub>	<b>3</b>	<b>2</b>	<b>2</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>1</b>	<b>2</b>	<b>1</b>	<b>1</b>	
U(1) <sub>Y</sub>	+1	+1/2	+1/2	+2	+1	+1	0	+1/2	+1	0	
$Z_2$				Even							
$Z'_2$	Even	Even	Odd	Even	Even	Odd	Even	Even	Odd		
L : Lepton #	-2	0	0	-2	-2	-2	-2	-1	-1	-1	
D1				○	○						1 1
D2	○										1 1
D3		○		○		○					2
D4			○		○						2
D5	○				○	○					1
D6					○	○					1
D7		○									1
D8				○				○	○		1 1
D9	○							○	○		1 1
D10					○			○			1 1
D11		○				○					2
D12					○			○			2
D13	○							○			1
D14								○			1
D15	○							○			1
D16								○	○		1
D17				○		○		○			1 2
D18								○			1

Table2. Classification of Models which generate Dirac neutrino masses

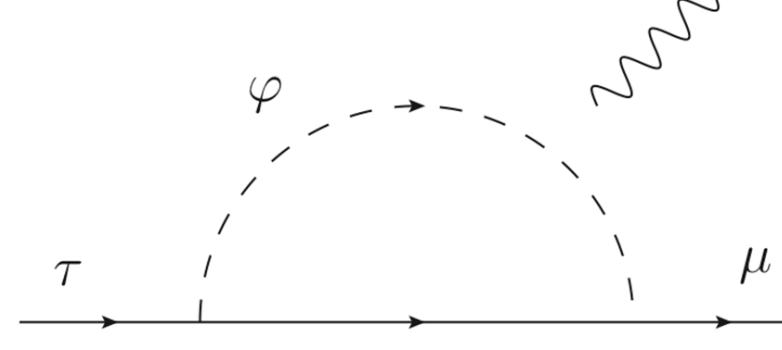
## References

- [1] A. Zee, Phys. Lett. B 93, (1980) 389. [2] L. M. Krauss, S. Nasri and M. Trodden, Phys. Rev. D 67, 085002 (2003). [3] S. Nasri and S. Moussa, Mod. Phys. Lett. A 17 (2002) 771. [4] S. Kanemura, K. Sakurai and H. Sugiyama, Phys. Rev. D 96 (2017) no.9, 095024. [5] S. Kanemura and H. Sugiyama, Phys. Lett. B 753 (2016) 161. [6] S. Kanemura, K. Sakurai and H. Sugiyama, Phys. Lett. B 758 (2016) 465. [7] M. Aoki, S. Kanemura, K. Sakurai, H. Sugiyama, Phys. Lett. B 763 (2016) 352.

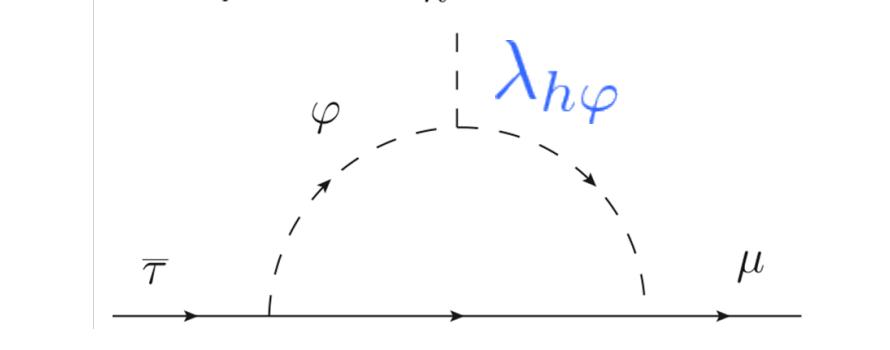
## 2. LFV decays of the Higgs boson [7]

LFV decays by Yukawa interaction with new scalar  $\varphi$

$\tau \rightarrow \mu\gamma$



$h \rightarrow \mu\tau$



When models have only one new Yukawa interaction,

$$\text{Br}(h \rightarrow \mu\tau) \sim 10^{-2} \text{ Br}(\tau \rightarrow \mu\gamma)$$

If  $h \rightarrow \mu\tau$  is observed w/o  $\tau \rightarrow \mu\gamma$ ,

M2  $\sim$  4, M6  $\sim$  8 and D5, D6, D13  $\sim$  16, D18 and Type-I, III seesaw mechanisms are excluded.

When two kinds of scalars couple with left (right)-handed lepton,

$$\text{Br}(\tau \rightarrow \mu\gamma) \propto \left| \frac{(Y_1^\dagger Y_1)_{\tau\mu}}{m_{\varphi_1}^2} + \frac{(Y_2^\dagger Y_2)_{\tau\mu}}{m_{\varphi_2}^2} \right|^2 \ll \left| \frac{(Y_1^\dagger Y_1)_{\tau\mu}}{m_{\varphi_1}^2} \right|^2 \quad \text{Can be cancelled}$$

$$\text{Br}(h \rightarrow \mu\tau) \propto \left| \lambda_{h\varphi_1} \frac{(Y_1^\dagger Y_1)_{\tau\mu}}{m_{\varphi_1}^2} + \lambda_{h\varphi_2} \frac{(Y_2^\dagger Y_2)_{\tau\mu}}{m_{\varphi_2}^2} \right|^2 \simeq \left| \lambda_{h\varphi_1} \frac{(Y_1^\dagger Y_1)_{\tau\mu}}{m_{\varphi_1}^2} \right|^2$$

Signs of  $\lambda_{h\varphi_1}$  and  $\lambda_{h\varphi_2}$  are different, the cancellation occur in only  $\tau \rightarrow \mu\gamma$ .

$\rightarrow$  D3, D4, D11, D12, D17

$$(m_\nu)_{ij} = m_{e_i} (Y \dots)_{ij} \quad Y_{ek} \gg Y_{\mu k} > Y_{\tau k}$$

If  $h \rightarrow \mu\tau$  is observed w/o  $\tau \rightarrow \mu\gamma$ , only D17 may survive in simple models generating neutrino masses.

## 3. Our model

	$\nu_{iR}$	$\psi_{\alpha R}$	$\Phi$	$s_1^+$	$\eta$	$s_2^+$	
Spin J	1/2	1/2	1	1	1	1	
SU(2) <sub>L</sub>	1	1	2	1	2	1	
U(1) <sub>Y</sub>	0	0	3/2	1	1/2	1	
$Z'_2$	-	+	(+)	-	+	+	
L	1	0	-2	-2	-1	-1	-1
$Z_2$	+	-					