

Higgs inflation in the NMSSM: impact at the electroweak scale

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in collaboration with

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Higgs inflation

- inflation required by cosmology
- instead of new field use Higgs boson as inflaton
- SM becomes unnatural [Barbon, Espinosa, [arXiv:0903.0355](#)]
- scale-free extension of the SM in canonical superconformal supergravity
[Einhorn, Jones, [arXiv:0912.2718](#)] [Ferrera, Kallosh, Linde, Marrani, Van Proeyen, [arXiv:1004.0712](#), [arXiv:1008.2942](#)]
inflation triggered by non-minimal coupling to Einstein gravity:

$$\mathcal{L}_\chi = -6 \int d^2\theta \mathcal{E} \left[R + X(\hat{\phi}) R - \frac{1}{4} (\bar{D}^2 - 8R) \hat{\phi}^\dagger \hat{\phi} + \mathcal{W}(\hat{\phi}) \right] + \text{h. c.} + \dots$$

only possible choice: $X = \chi \hat{H}_u \cdot \hat{H}_d$

- MSSM no viable model for inflation [Einhorn, Jones, [arXiv:0912.2718](#)]
additional scalar singlet + stabilisator term at high energy works
[Ferrera, Kallosh, Linde, Marrani, Van Proeyen, [arXiv:1008.2942](#)] [Lee, [arXiv:1005.2735](#)]



The standard NMSSM

\mathbb{Z}_3 invariant NMSSM:

- two Higgs doublets, one Higgs singlet:

$$H_u = \begin{pmatrix} \eta_u^+ \\ v_u + \frac{1}{\sqrt{2}} (\sigma_u + i \phi_u) \end{pmatrix}, \quad H_d = \begin{pmatrix} v_d + \frac{1}{\sqrt{2}} (\sigma_d + i \phi_d) \\ \eta_d^- \end{pmatrix}, \quad S = v_s + \frac{1}{\sqrt{2}} (\sigma_s + i \phi_s)$$

- superpotential

$$\mathcal{W} = \lambda S H_u \cdot H_d + \frac{1}{3} \kappa S^3 + \text{Yukawa}$$

dynamically generated term $\mu_{\text{eff}} = \lambda v_s$ solves μ -problem of MSSM



The NMSSM with inflation

- term $X = \chi H_u \cdot H_d$ breaks \mathbb{Z}_3 symmetry,
appears in Kähler potential

$$\mathcal{K} = -3 \log \left[1 - \frac{1}{3} \left(|S|^2 + |H_d|^2 + |H_d|^2 \right) - \frac{1}{2} \chi (H_u \cdot H_d + \text{h. c.}) \right]$$

- superpotential changed: [Ferrera, Kallosh, Linde, Marrani, Van Proeyen, [arXiv:1008.2942](https://arxiv.org/abs/1008.2942)]

$$\mathcal{W} \rightarrow \mathcal{W} \exp \left(\frac{X}{M_P^2} \right) \approx \mathcal{W} + \frac{1}{M_P^2} \langle \mathcal{W}_{\text{hidden}} \rangle X \approx \mathcal{W} + m_{3/2} X$$

- can be accommodated by more general NMSSM with superpotential

$$\mathcal{W} = \lambda S H_u \cdot H_d + \frac{1}{3} \kappa S^3 + \frac{3}{2} m_{3/2} \chi H_u \cdot H_d + \text{Yukawa}$$



New parameters

- additional term appears as an MSSM-like μ term with $\mu = \frac{3}{2} m_{3/2} \chi$
- approximate value of $\chi \approx 10^5 \lambda$
(last 60 e-folds of inflation, COBE normalization of scalar perturbations)
[Ferrera, Kallosh, Linde, Marrani, Van Proeyen, [arXiv:1008.2942](#)] [Lee, [arXiv:1005.2735](#)]
- additional soft-breaking term

$$-\mathcal{L}_{\text{soft}} = \left[A_\lambda \lambda S H_u \cdot H_d + \frac{1}{3} A_\kappa \kappa S^3 + \color{red} B_\mu \mu H_u \cdot H_d + \text{h. c.} \right] + m_{H_d}^2 |H_d|^2 + m_{H_u}^2 |H_u|^2 + m_s^2 |S|^2$$

- additional \mathbb{Z}_3 -breaking parameters possible,
in the following: equal to zero at tree level
 - superpotential parameters zero at all orders
 - running of soft-breaking parameters in general small
- further studies in extended NMSSM or GNMSSM:

[Ellwanger, Hugonie, Teixeira, [arXiv:0910.1785](#)] [Ross, Schmidt-Hoberg, [arXiv:1108.1284](#)]

[Ross, Schmidt-Hoberg, Staub, [arXiv:1205.1509](#)] [Kaminska, Ross, Schmidt-Hoberg, [arXiv:1308.4168](#)] [Badziak, Wagner, [arXiv:1611.02353](#)]

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The setup

- NMSSM with \mathbb{Z}_3 -breaking μ, B_μ terms,
all parameters considered to be real
- compare phenomenology with \mathbb{Z}_3 -preserving NMSSM
→ focus set on μ (and B_μ)
- take into account constraints from:
 - vacuum stability
 - SM-like Higgs at 125 GeV
 - observed Higgs data with HiggsSignals
 - limits on extended Higgs sector with HiggsBounds
 - neutralino and chargino masses (higgsinos)
 - sfermion mixing (charge-, color-breaking minima)



Vacuum stability

- minimization conditions of Higgs potential

$$\left. \frac{\partial V}{\partial h} \right|_{\text{vev}} = 0$$

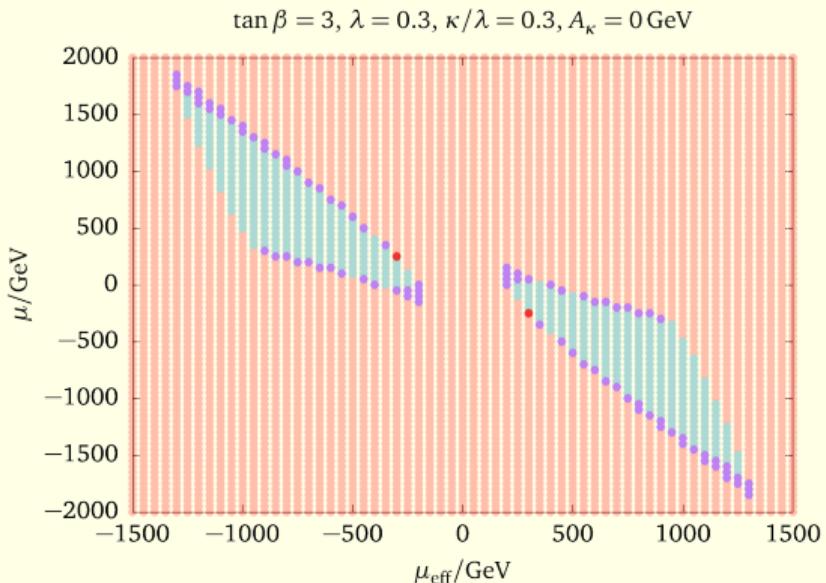
may be misleading

- scenarios with different minima at large field values possible
- check for global minimum = electroweak minimum
- other minima may be viable as well (meta-stable with long life time)
- minima typically hard to find analytically
→ numerical minimization of tree-level Higgs potential
- μ term in general increases allowed parameter region
- substitution of A_λ by charged Higgs mass m_{H^\pm} :

$$A_\lambda = \frac{c_\beta s_\beta}{\mu_{\text{eff}}} \left(m_{H^\pm}^2 - m_W^2 + v^2 \lambda^2 \right) - \frac{B_\mu \mu}{\mu_{\text{eff}}} - \frac{\kappa}{\lambda} \mu_{\text{eff}}.$$



Vacuum stability – example

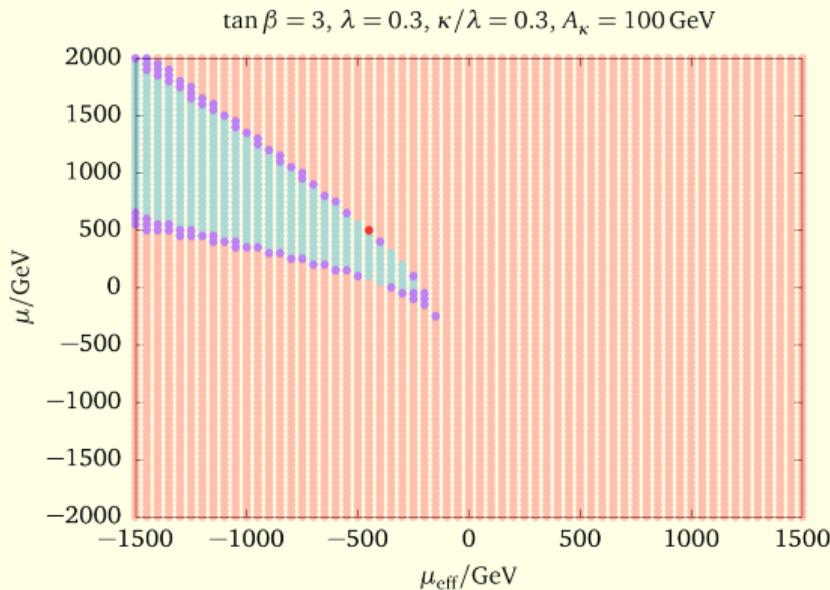


light blue: stable vacuum,

purple: long-lived metastable vacuum,

red: short-lived metastable vacuum,

rose: unstable vacuum (tachyons)



value of A_κ has severe impact (singlet states),

region around $\mu_{\text{eff}} = 0$ not accessible,

in most scenarios: sign $A_\kappa = -$ sign μ_{eff} ,

non-zero μ : allow scenarios impossible in NMSSM



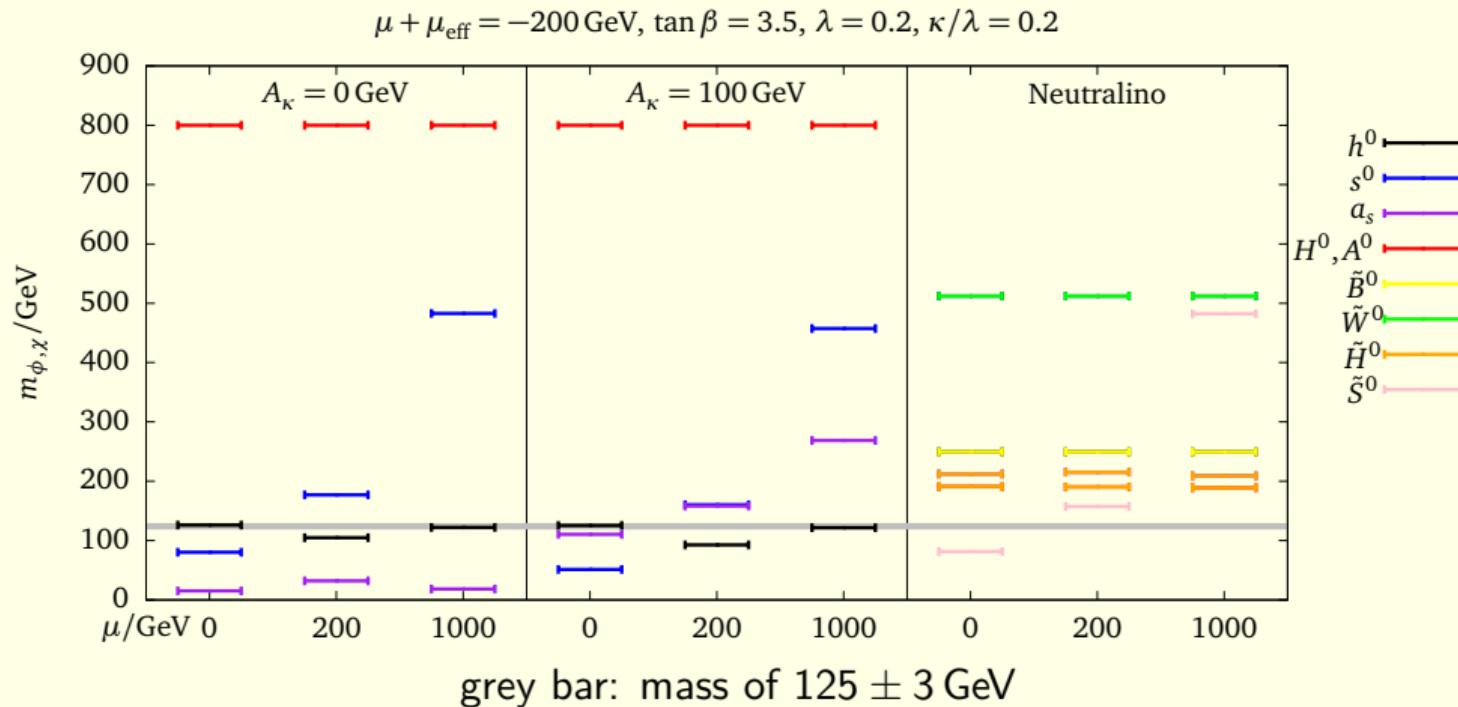
Higgs masses

- at tree level:
SM-like Higgs mass shifted upwards compared to MSSM (same as in NMSSM),
 $v^2(\mu + \mu_{\text{eff}})$, $v A_\lambda$ appear in singlet–doublet mixing,
 $v^2 \mu/\mu_{\text{eff}}$, $v^2 A_\lambda/\mu_{\text{eff}}$, $\mu_{\text{eff}} A_\kappa$, μ_{eff}^2 appear in diagonal elements of singlets
- full one-loop corrections:
 $\overline{\text{DR}}$ scheme for μ , B_μ (and also λ , κ , A_κ , μ_{eff})
- additional two-loop corrections in the MSSM-limit with FeynHiggs:
important mass shifts of $\mathcal{O}(\alpha_t \alpha_s, \alpha_t^2)$ to SM-like state
- masses determined from poles of

$$\hat{\Delta}(k^2) = -i \left[k^2 \mathbf{1} - M_{\text{tree}}^2 + \hat{\Sigma}^{(1L)}(k^2) + \hat{\Sigma}_{\text{MSSM}}^{(\alpha_t \alpha_s, \alpha_t^2)}(0) \right]^{-1}$$



Higgs masses – example



fixed sum $\mu + \mu_{\text{eff}}$: large positive $\mu \rightarrow$ large negative μ_{eff} ,
singlets and singlinos sensitive to μ_{eff}



Neutralino masses

$$\mathcal{M}_\chi = \begin{pmatrix} M_1 & 0 & -M_Z s_w c_\beta & M_Z s_w s_\beta & 0 \\ \cdot & M_2 & M_Z c_w c_\beta & -M_Z c_w s_\beta & 0 \\ \cdot & \cdot & 0 & -(\mu + \mu_{\text{eff}}) & -\lambda v s_\beta \\ \cdot & \cdot & \cdot & 0 & -\lambda v c_\beta \\ \cdot & \cdot & \cdot & \cdot & 2 \frac{\kappa}{\lambda} \mu_{\text{eff}} \end{pmatrix}$$

sum $\mu + \mu_{\text{eff}}$ in MSSM-like higgsino mass terms (analogous for charged),
term $\frac{\kappa}{\lambda} \mu_{\text{eff}}$ in singlino mass term,

for $\mu \neq 0$ all neutralino/chargino masses constant by shifting and rescaling

$$\mu'_{\text{eff}} = \mu_{\text{eff}} + \mu,$$

$$\kappa' = \kappa \frac{\mu_{\text{eff}} + \mu}{\mu_{\text{eff}}}$$



Higgs production and decays

- production:
in general: SM-normalized effective couplings of a Higgs boson to gluons,
light-singlet scenario: NMSSM version of SusHi [Liebler, [arXiv:1502.07972](#)]
- decays: SM-normalized effective couplings
- mixing: employ Z matrix in the algorithm of Ref. [Domingo, Drechsel, SP, [arXiv:1706.00437](#)]
- relevant couplings λ_{ijk} of $\phi_{i,j,k} \in (\sigma_d, \sigma_u, \sigma_s, A, \phi_s)$ containing μ and/or μ_{eff} :

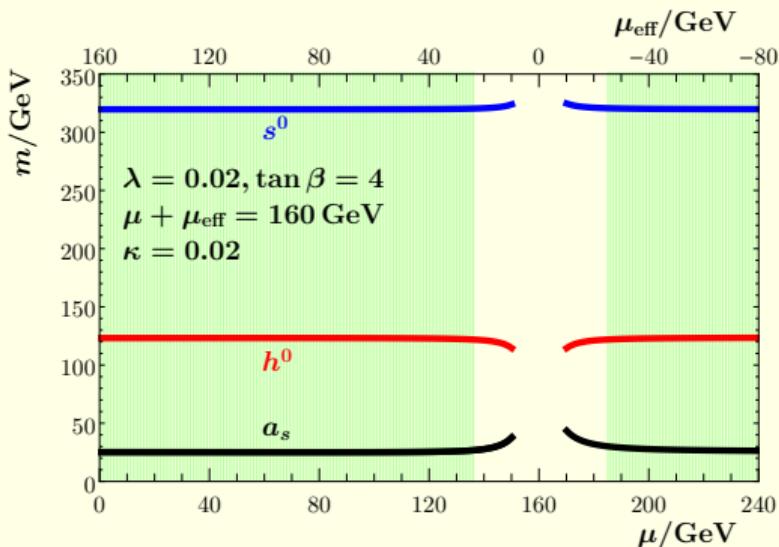
$$\begin{aligned}\lambda_{113} = \lambda_{223} = \lambda_{344} = \lambda_{355} &= -2\lambda(\mu + \mu_{\text{eff}}) \\ \lambda_{123} = -\lambda_{345} &= \lambda A_\lambda + 2\kappa\mu_{\text{eff}}\end{aligned}$$

(in addition: couplings to charged Higgs)

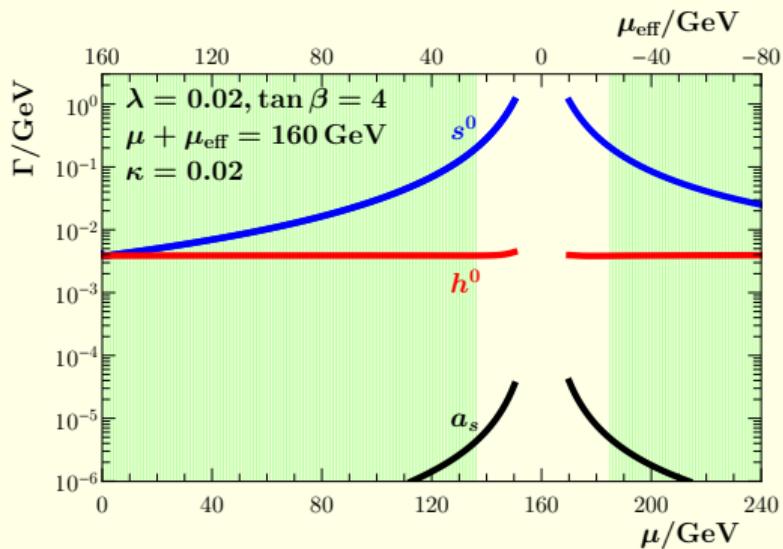
- sensitive decays: $s^0 \rightarrow h^0 h^0$, $H^0 \rightarrow s^0 h^0$, $A^0 \rightarrow s^0 a_s$
(h^0 , H^0 , s^0 , A^0 , a_s denote states with most contribution of this type,
mixing matrices appear and are relevant)



Higgs decays – example



masses nearly constant via rescaling



total width

scenario with large κ' and very small λ and μ_{eff} , but large mixing of h^0 and s^0
(not possible in NMSSM due to constraints on higgsino mass)

① From cosmology to the model

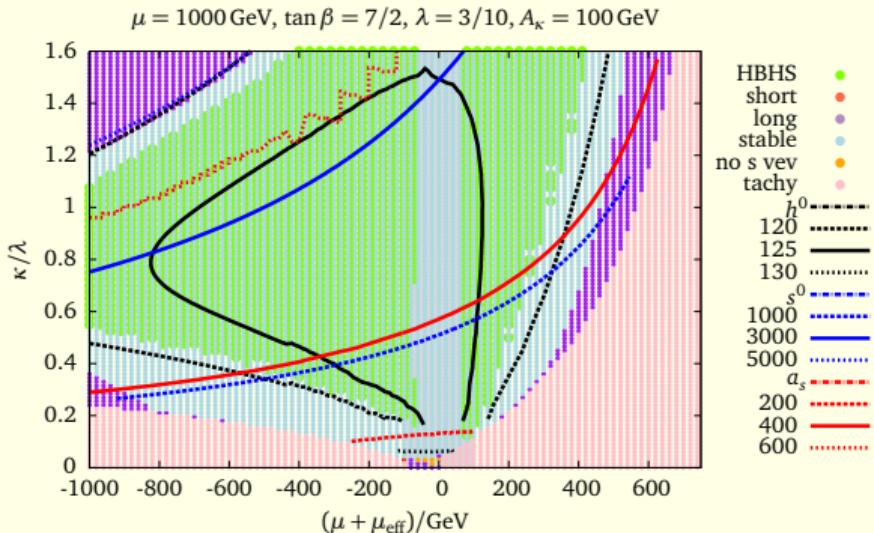
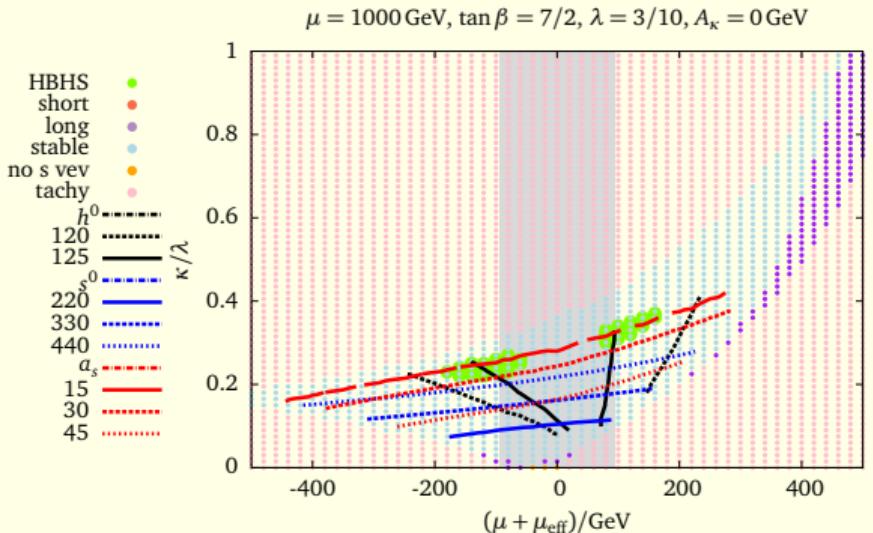
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Dependence on κ/λ over $\mu + \mu_{\text{eff}}$



grey bar: LEP-limit for chargino masses,

green background: allowed by HiggsSignals and HiggsBounds

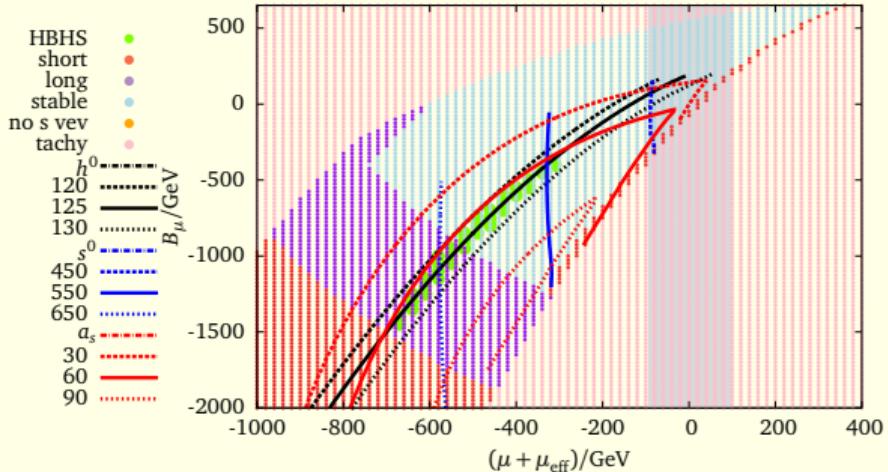
singlet masses very sensitive to A_κ (like NMSSM),

SM-like Higgs with 125 GeV only in region with stable vacuum (light blue)

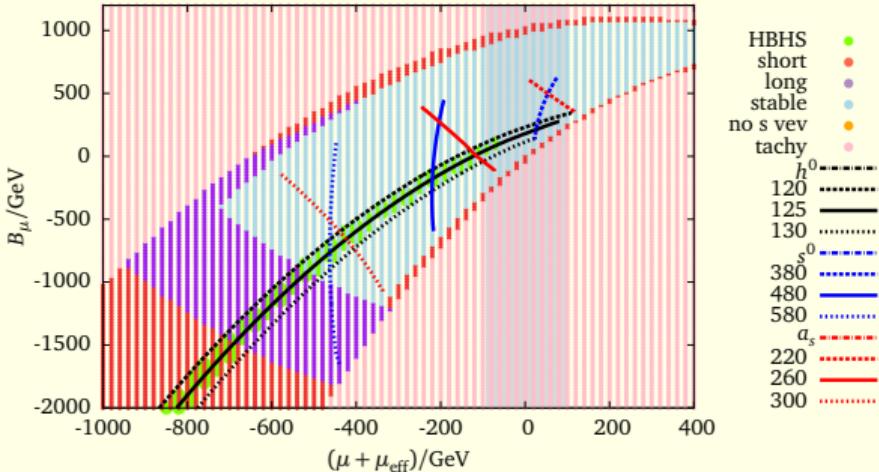


Dependence on B_μ over $\mu + \mu_{\text{eff}}$

$\mu = 1000 \text{ GeV}, \tan \beta = 3/2, \kappa = 1/10, A_\kappa = 0 \text{ GeV}$



$\mu = 1000 \text{ GeV}, \tan \beta = 3/2, \kappa = 1/10, A_\kappa = 100 \text{ GeV}$



non-zero B_μ may cause instable vacua,
differences between tachyons at tree level and loop level can be large
→ large loop corrections,
for some scenarios: more two-loop (or higher-order) corrections necessary



Conclusions and outlook

- Higgs inflation in extension of NMSSM by \mathbb{Z}_3 -breaking μ term
- vacuum stability and SM-like Higgs at 125 GeV significantly constrain parameter space
- parameter rescaling to keep impact of μ on Higgs masses small, but: decay widths of singlets strongly affected
- attributes of singlet-like Higgs bosons required to distinguish inflation-inspired model from standard NMSSM
- light singlet scenarios which do not exist in standard NMSSM