Detectability of pseudo-Goldstone dark matter

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Dark matter

Dark matter

There is a lot of evidence of dark matter.

- Rotation curves of spiral galaxies
- CMB observations
- Gravitational lensing
- Large scale structure of the universe
- Collision of bullet cluster
- DM existence is crucial.



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Albada et al.



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DM existence is crucial.

WIMPs are good DM candidate.

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WIMP search status

XENON1T, PRL (2018)

Fermi-LAT, PRL (2015)

- Experimental bounds for WIMPs are stronger and stronger.
- Interactions between DM and SM are very weak?
 - \rightarrow possibilities other than WIMPs are discussed.

Model of pseudo-Goldstone DM

Model

Model of pseudo-Goldstone DM

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

- Introduce complex scalar field $S = (s + i\chi)/\sqrt{2}$
- Global U(1) symmetry is assumed (invariant under $S \rightarrow e^{i\alpha}S$)

$$\begin{split} \mathcal{V} = & -\frac{\mu_H^2}{2} |H|^2 - \frac{\mu_S^2}{2} |S|^2 + \frac{\lambda_H}{2} |H|^4 + \lambda_{HS} |H|^2 |S|^2 + \frac{\lambda_S}{2} |S|^4 \\ & - \left(\frac{\mu_S'^2}{4} S^2 + \text{H.c.}\right) \quad \leftarrow \text{ soft breaking mass term} \end{split}$$

• After H and S get VEVs, ϕ and s mix

$$H = \begin{pmatrix} 0 \\ (v+\phi)/\sqrt{2} \end{pmatrix}, \qquad S = \frac{v_s + s + i\chi}{\sqrt{2}}$$
$$\begin{pmatrix} \phi \\ s \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

• $\sin \theta \leq 0.3$ \leftarrow Constrained by EWPT, h_2 direct search at LHC

Model of pseudo-Goldstone DM

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

• χ is mass eigenstate itself $m_{\chi}^2 = \mu_S'^2$ Invariant under $S \to S^{\dagger}$, $\Rightarrow \chi$ can be a DM candidate Higgs portal DM

Scalar potential $\mathcal{V} = \mu_{h_1\chi\chi}h_1\chi^2 + \mu_{h_2\chi\chi}h_2\chi^2 + \cdots$

$$\mu_{h_1\chi\chi} = -\frac{m_{h_1}^2 \sin \theta}{v_s}, \quad \mu_{h_2\chi\chi} = \frac{m_{h_2}^2 \cos \theta}{v_s},$$

SM Yukawa int. $\mathcal{L} \supset y_q \Big(\cos \theta h_1 + \sin \theta h_2 \Big) \overline{q} q$

$$\lambda_{H} = \frac{\cos^{2} \theta m_{h_{1}}^{2} + \sin^{2} \theta m_{h_{2}}^{2}}{v^{2}}, \quad \lambda_{S} = \frac{\sin^{2} \theta m_{h_{1}}^{2} + \cos^{2} \theta m_{h_{2}}^{2}}{v_{s}^{2}},$$
$$\lambda_{HS} = \frac{\sin \theta \cos \theta (m_{h_{2}}^{2} - m_{h_{1}}^{2})}{v v_{s}}$$

2

Δ

Direct detection (tree level) C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

Scattering amplitude is suppressed by transfer momentum q^2

$$i\mathcal{M} \sim i\left(\frac{m_{h_1}^2}{q^2 - m_{h_1}^2} - \frac{m_{h_2}^2}{q^2 - m_{h_2}^2}\right) \sim i\frac{q^2(m_{h_1}^2 - m_{h_2}^2)}{m_{h_1}^2 m_{h_2}^2}$$

→ All the interactions are derivative couplings $\mathcal{L}_{int}(\partial_{\mu}\chi)$ ■ 4 independent parameters $(m_{\chi}, m_{h_2}, \sin\theta, v_s (\lambda_S))$

Direct detection (1-loop level)

K. Ishiwata, TT, JHEP (2018) D. Azevedo, M. Duch, B. Grzadkowski, D. Huang M. Iglicki, R. Santos, JHEP (2019)

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Compute Feynman diagrams at 1-loop level

Numerical analysis (1-loop level) K. Ishiwata, TT, JHEP [arXiv:1810.08139]

- Invisible Higgs decay $Br(h_1 \rightarrow inv) \lesssim 10\%$ at LHC
- Two resonances at (h_1,h_2) , $v/v_s\sim \sqrt{\lambda_S}v/m_{h_2}$
- Perturbative unitarity $\lambda_S \leq 8\pi/3$ Chen et al., arXiv:1410.5488
- Partial region can be tested by DARWIN. $\sigma_{\rm SI}^p = \mathcal{O}(10^{-48})~{\rm cm}^2$

Numerical analysis (1-loop level) K. Ishiwata, TT, JHEP [arXiv:1810.08139]

 $\bullet \sin \theta = 0.2$

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Indirect detection

Huitu, Koivunen, Lebedev, Mondal, TT, arXiv:1812.05952

Small parameter space is excluded by Fermi-LAT gamma-ray observation $(\chi \chi \to WW, bb \to \gamma)$

- Thermal WIMP scenarios can be tested only when $m_{\chi} = \mathcal{O}(100)$ GeV
- CTA is sensitive in heavy DM mass region (DM profile dependent) but $\chi\chi \rightarrow h_2h_2$ is dominant in this mass range.

Collider search

Huitu, Koivunen, Lebedev, Mondal, TT, arXiv:1812.05952

• Constraint on h_2 production cross section at LHC

 $\sigma_{\text{prod}} = \sigma(pp \to h_2) \operatorname{Br}(h_2 \to SM) \propto \sin^2 \theta \operatorname{Br}(h_2 \to SM) \lesssim \#_{\exp}$

■ $pp \rightarrow h_2 \rightarrow ZZ$ mode When $\sin \theta \gtrsim 0.2$ and $m_{h_2} \lesssim 2m_{h_1}$, parameters are constrained.

Summary

- **1** Thermal WIMP scenarios are strongly constrained by direct DM search.
- Pseudo-Goldstone DM can naturally avoid the strong constraint.
- 3 Elastic cross section with nucleaon is $\sigma_{SI}^N = \mathcal{O}(10^{-48}) \text{ cm}^2$ at most. (1-loop)
- The model is testable by indirect detection and collider search.