



20th Annual RDMS CMS Collaboration Conference



Tashkent-Samarkand

Uzbekistan, 12-15 September 2018

Conference dedicated to new physics results, Future Physics beyond the LHC era, and the CMS Detector Upgrade, including Endcap Calorimetry

Registration: <http://rdms2018.jinr.ru/>

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WHAT MAKES US THINK THAT THERE IS PHYSICS BEYOND THE STANDARD MODEL?

Dmitry Kazakov

BLTP JINR

The Standard Model of Fundamental Interactions

Higgs Sector

Neutrino Sector

Flavour Sector

Dark Matter

New particles and Interactions

THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{Yukawa} + \mathcal{L}_{Higgs},$$

$$\mathcal{L}_{gauge} = -\frac{1}{4}G_{\mu\nu}^a G_{\mu\nu}^a - \frac{1}{4}W_{\mu\nu}^i W_{\mu\nu}^i - \frac{1}{4}B_{\mu\nu} B_{\mu\nu}$$

$$+ i\bar{L}_\alpha \gamma^\mu D_\mu L_\alpha + i\bar{Q}_\alpha \gamma^\mu D_\mu Q_\alpha + i\bar{E}_\alpha \gamma^\mu D_\mu E_\alpha$$

$$+ i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha + i\bar{D}_\alpha \gamma^\mu D_\mu D_\alpha + (D_\mu H)^\dagger (D_\mu H),$$

$$+ i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha \quad \leftarrow \text{possible right handed neutrino ?}$$

$$\mathcal{L}_{Yukawa} = y_{\alpha\beta}^L \bar{L}_\alpha E_\beta H + y_{\alpha\beta}^D \bar{Q}_\alpha D_\beta H + y_{\alpha\beta}^U \bar{Q}_\alpha U_\beta \tilde{H} + h.c.,$$

$$+ y_{\alpha\beta}^N \bar{L}_\alpha N_\beta \tilde{H} \quad \leftarrow$$

$$\mathcal{L}_{Higgs} = -V = m^2 H^\dagger H - \frac{\lambda}{2} (H^\dagger H)^2$$

THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}$$

$$\mathcal{L}_{gauge} = -\frac{1}{4}G_{\mu\nu}^c G_{\mu\nu}^c$$

$$+ i\bar{L}_\alpha \gamma^\mu D_\mu L_c$$

$$+ i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha +$$

$$+ i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha \quad \leftarrow$$

$$\mathcal{L}_{Yukawa} = y_{\alpha\beta}^L \bar{L}_\alpha E_\beta H$$

$$+ y_{\alpha\beta}^N \bar{L}_\alpha \Lambda$$

$$\mathcal{L}_{Higgs} = -V =$$

$$\begin{aligned} \mathcal{L}_{SM} = & \frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- \\ & M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - igc_w (\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\ & W_\mu^- W_\nu^+) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+)) - \\ & ig s_w (\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\mu^- W_\nu^+) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + A_\nu (W_\mu^+ \partial_\nu W_\mu^- - \\ & W_\mu^- \partial_\nu W_\mu^+)) - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \frac{1}{2}g^2 W_\mu^- W_\mu^+ W_\nu^- W_\nu^+ + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - \\ & Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w (A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\ & W_\mu^- W_\nu^+) - 2A_\mu Z_\nu^0 W_\mu^+ W_\nu^-) - \frac{1}{2}\partial_\mu H \partial_\mu H - 2M^2 \alpha_h H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \\ & \beta_h \left(\frac{2M^2}{g^2} + \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right) + \frac{2M^4}{g^2} \alpha_h - \\ & g\alpha_h M (H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-) - \\ & \frac{1}{8}g^2 \alpha_h (H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2) - \\ & gM W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \\ & \frac{1}{2}ig (W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)) + \\ & \frac{1}{2}g (W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) + W_\mu^- (H \partial_\mu \phi^+ - \phi^+ \partial_\mu H)) + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) + \\ & M (\frac{1}{c_w} Z_\mu^0 \partial_\mu \phi^0 + W_\mu^+ \partial_\mu \phi^- + W_\mu^- \partial_\mu \phi^+)) - ig \frac{e^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + ig s_w M A_\mu (W_\mu^+ \phi^- - \\ & W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + ig s_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \\ & \frac{1}{4}g^2 W_\mu^+ W_\mu^- (H^2 + (\phi^0)^2 + 2\phi^+ \phi^-) - \frac{1}{8}g^2 \frac{1}{c_w^2} Z_\mu^0 Z_\mu^0 (H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-) - \\ & \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\ & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\ & g^2 s_w^2 A_\mu A_\mu \phi^+ \phi^- + \frac{1}{2}ig_s \lambda_{ij}^2 (\bar{q}_i^\alpha \gamma^\mu q_j^\alpha) g_\mu^a - \bar{e}^\lambda (\gamma^\partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda (\gamma^\partial + m_\nu^\lambda) \nu^\lambda - \bar{u}_j^\lambda (\gamma^\partial + \\ & m_u^\lambda) u_j^\lambda - \bar{d}_j^\lambda (\gamma^\partial + m_d^\lambda) d_j^\lambda + ig s_w A_\mu (-\bar{e}^\lambda \gamma^\mu e^\lambda + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)) + \\ & \frac{ig}{4c_w} Z_\mu^0 \{ (\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - 1 - \gamma^5) d_j^\lambda) + \\ & (\bar{u}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 + \gamma^5) u_j^\lambda) \} + \frac{ig}{2\sqrt{2}} W_\mu^+ ((\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) U^{lep}{}_{\lambda e} e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda e} d_j^\lambda)) + \\ & \frac{ig}{2\sqrt{2}} W_\mu^- ((\bar{e}^\lambda U^{lep}{}_{\lambda e} \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\lambda C_{\lambda e}^\dagger \gamma^\mu (1 + \gamma^5) u_j^\lambda)) + \\ & \frac{ig}{2M\sqrt{2}} \phi^+ (-m_e^\lambda (\bar{\nu}^\lambda U^{lep}{}_{\lambda e} (1 - \gamma^5) e^\lambda) + m_\nu^\lambda (\bar{\nu}^\lambda U^{lep}{}_{\lambda e} (1 + \gamma^5) e^\lambda) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_e^\lambda (\bar{e}^\lambda U^{lep}{}_{\lambda e} (1 + \gamma^5) \nu^\lambda) - m_\nu^\lambda (\bar{e}^\lambda U^{lep}{}_{\lambda e} (1 - \gamma^5) \nu^\lambda) - \frac{g}{2} \frac{m_\nu^\lambda}{M} H (\bar{\nu}^\lambda \nu^\lambda) - \\ & \frac{g}{2} \frac{m_e^\lambda}{M} H (\bar{e}^\lambda e^\lambda) + \frac{ig}{2} \frac{m_\nu^\lambda}{M} \phi^0 (\bar{\nu}^\lambda \gamma^5 \nu^\lambda) - \frac{ig}{2} \frac{m_e^\lambda}{M} \phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda) - \frac{1}{4} \bar{\nu}_\lambda M_{\lambda e}^R (1 - \gamma_5) \hat{\nu}_\lambda - \\ & \frac{1}{4} \bar{\nu}_\lambda M_{\lambda e}^R (1 - \gamma_5) \hat{\nu}_\lambda + \frac{ig}{2M\sqrt{2}} \phi^+ (-m_d^\lambda (\bar{u}_j^\lambda C_{\lambda e} (1 - \gamma^5) d_j^\lambda) + m_u^\lambda (\bar{u}_j^\lambda C_{\lambda e} (1 + \gamma^5) d_j^\lambda) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_d^\lambda (\bar{d}_j^\lambda C_{\lambda e}^\dagger (1 + \gamma^5) u_j^\lambda) - m_u^\lambda (\bar{d}_j^\lambda C_{\lambda e}^\dagger (1 - \gamma^5) u_j^\lambda) - \frac{g}{2} \frac{m_\nu^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \\ & \frac{g}{2} \frac{m_d^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_\nu^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \frac{ig}{2} \frac{m_d^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c + \\ & \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \\ & \partial_\mu \bar{X}^+ X^0) + ig s_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \partial_\mu \bar{X}^+ Y) + igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \\ & \partial_\mu \bar{X}^0 X^+) + ig s_w W_\mu^- (\partial_\mu \bar{X}^- Y - \partial_\mu \bar{Y} X^+) + igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^- - \\ & \partial_\mu \bar{X}^- X^+) + ig s_w A_\mu (\partial_\mu \bar{X}^+ X^- - \\ & \partial_\mu \bar{X}^- X^+) - \frac{1}{2}gM (\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H) + \frac{1-2c_w^2}{2c_w} igM (\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-) + \\ & \frac{1}{2c_w} igM (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + igM s_w (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + \\ & \frac{1}{2}igM (\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0) . \end{aligned}$$

THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{Yukawa} + \mathcal{L}_{Higgs},$$

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Three gauge couplings

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Three gauge couplings

Three or four Yukawa matrices

THE LAGRANGIAN

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Three gauge couplings

Three or four Yukawa matrices

Two parameters

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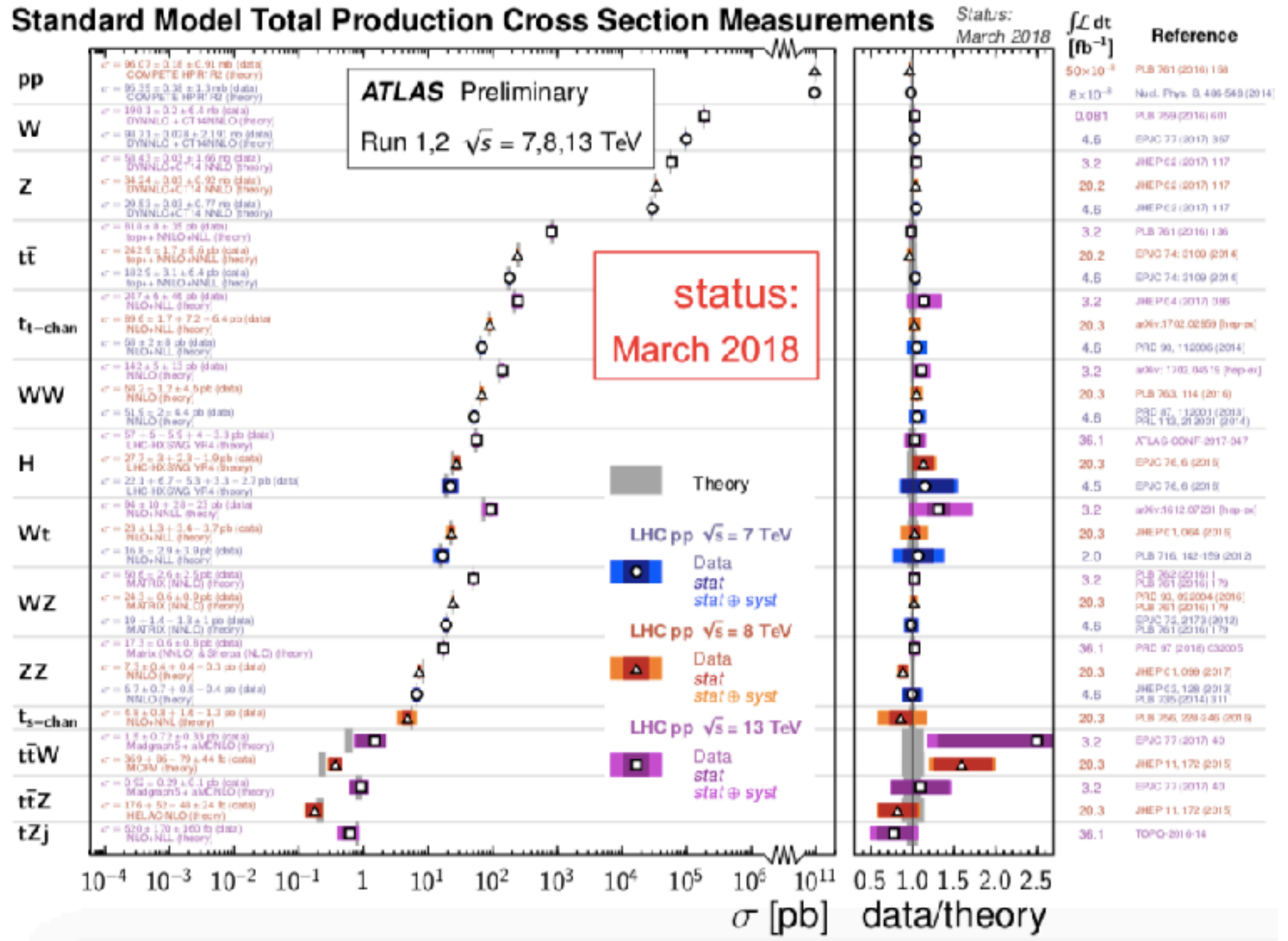
All these parameters are not predicted by the SM and determined experimentally

Three gauge couplings

Three or four Yukawa matrices

Two parameters

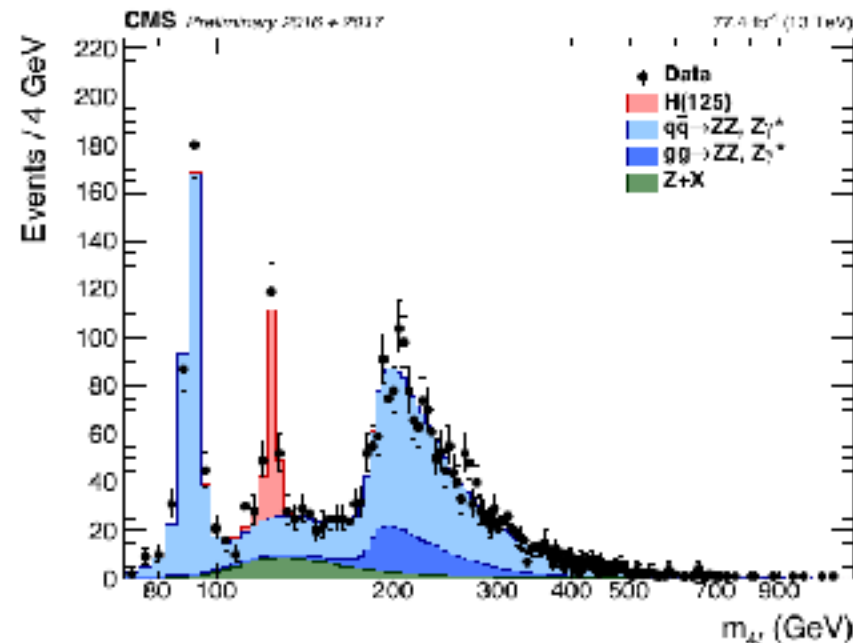
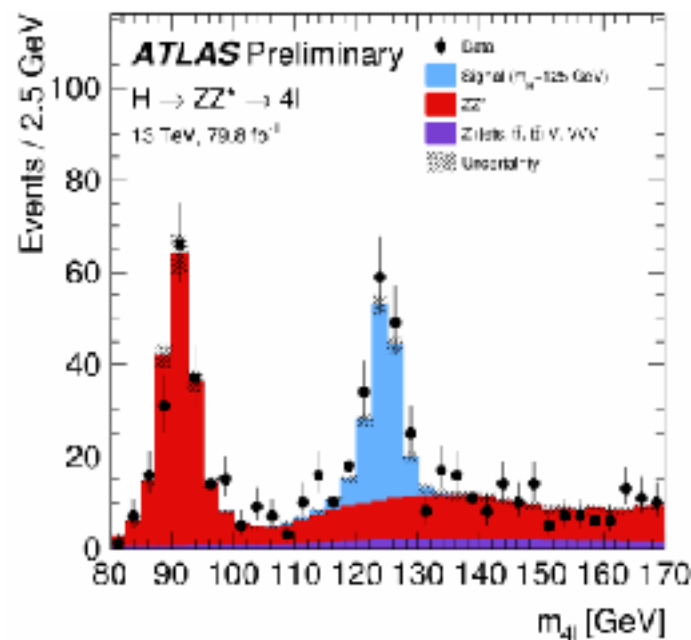
C. Gwenlan
ICHEP2018



Extraordinary agreement between measurements and SM predictions

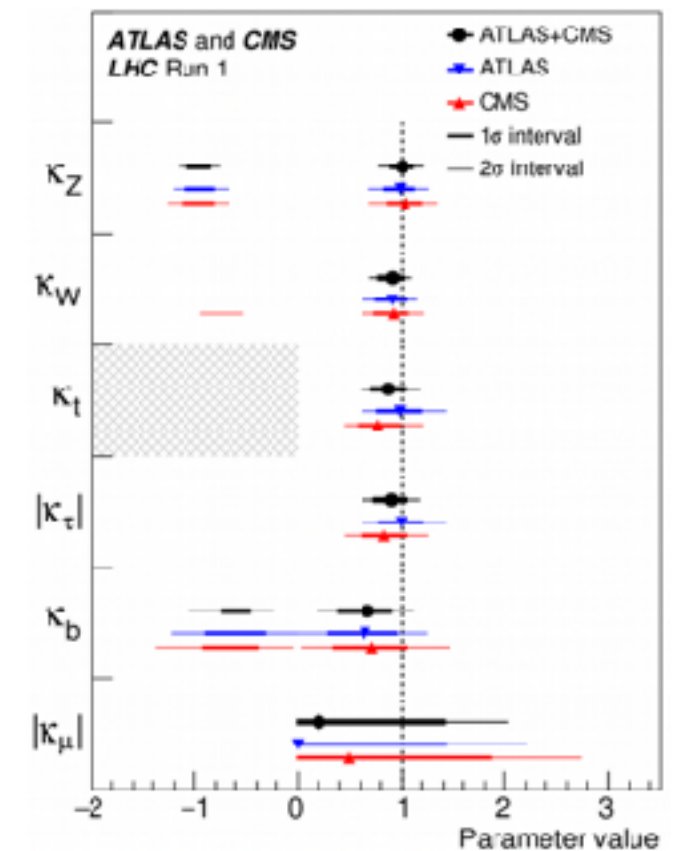
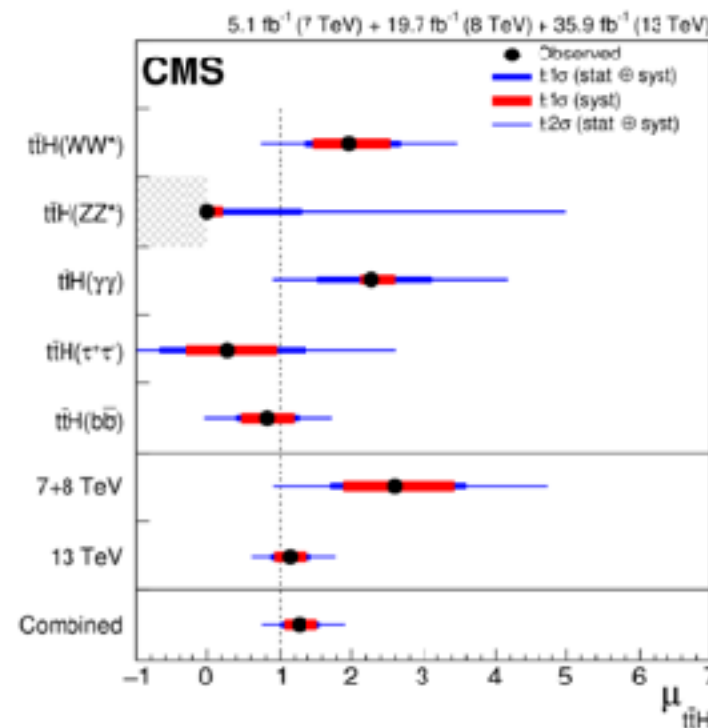
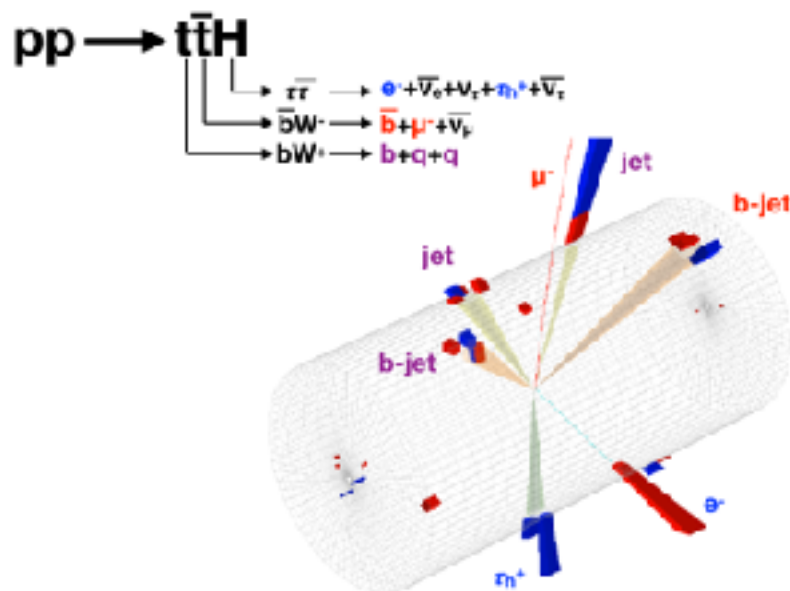
Higgs bosons - entering precision era

Run-2 analyses with 80 fb^{-1} for the first time – higher precision is coming!



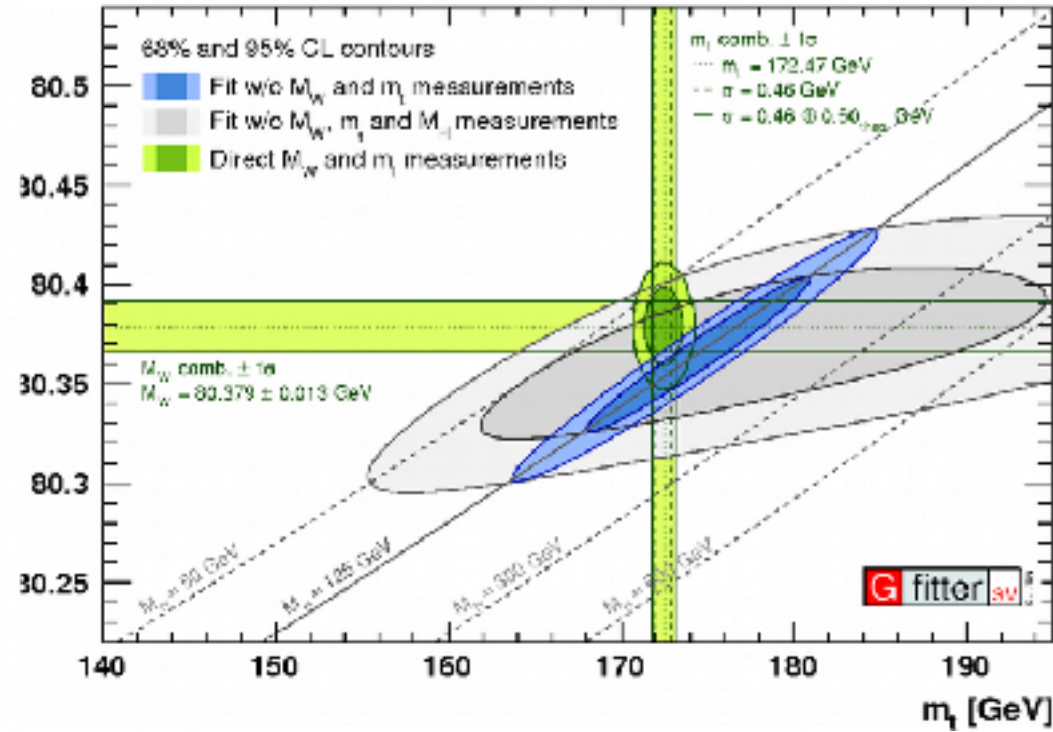
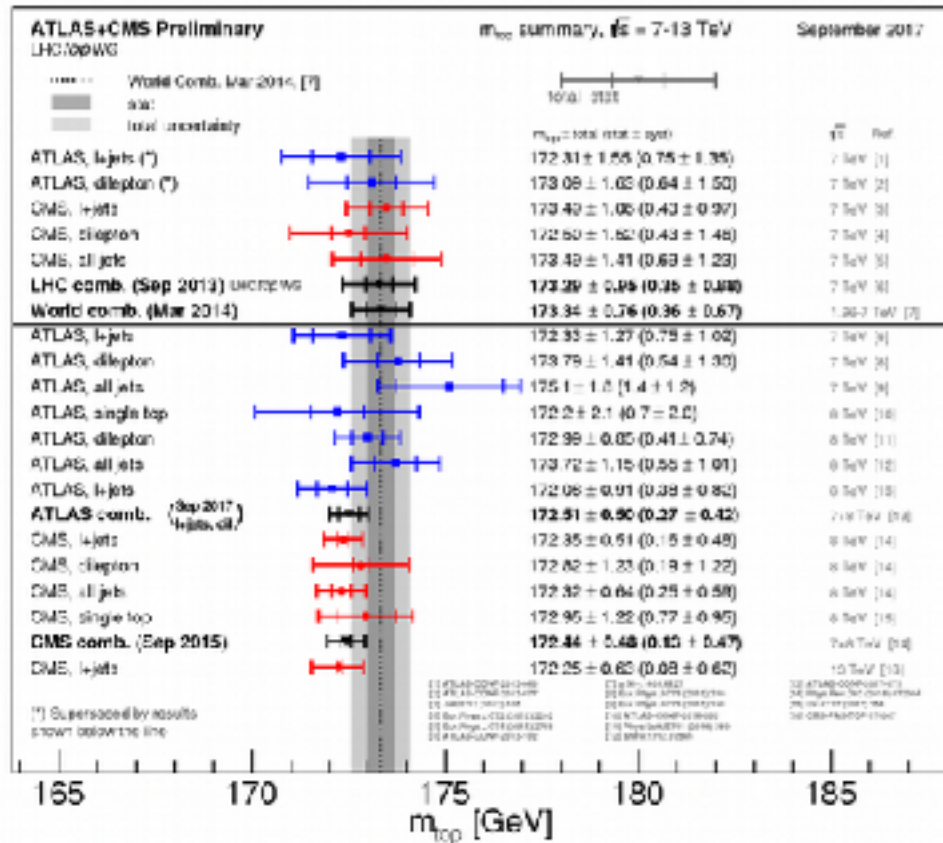
D. Charlton
LHCp2018

ttH observation



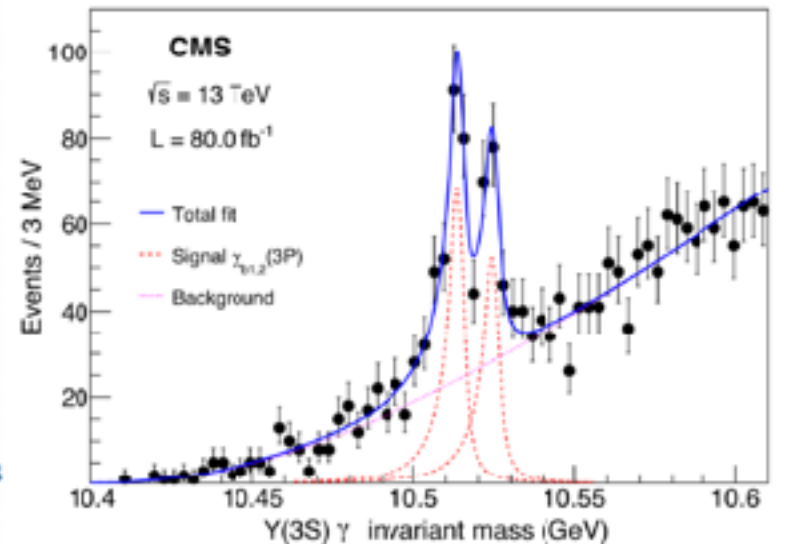
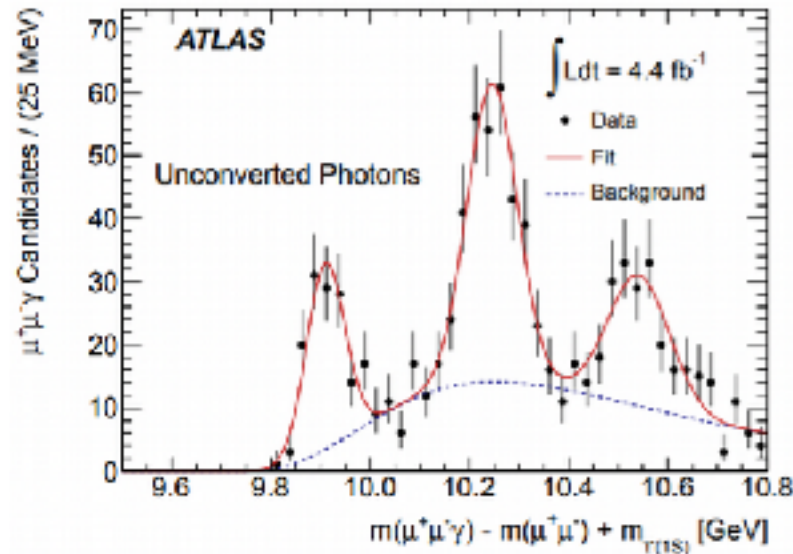
Precision EW mass measurements

D. Charlton
LHCp2018



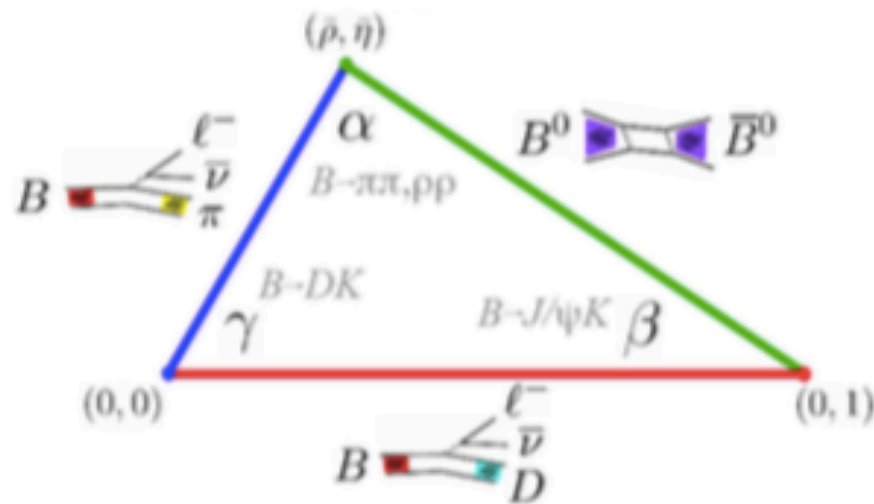
Precision spectroscopy!

$$m(\chi_{b2}(3P)) - m(\chi_{b2}(3P)) = 10.60 \pm 0.64(\text{stat}) \pm 0.17(\text{syst}) \text{ MeV}$$



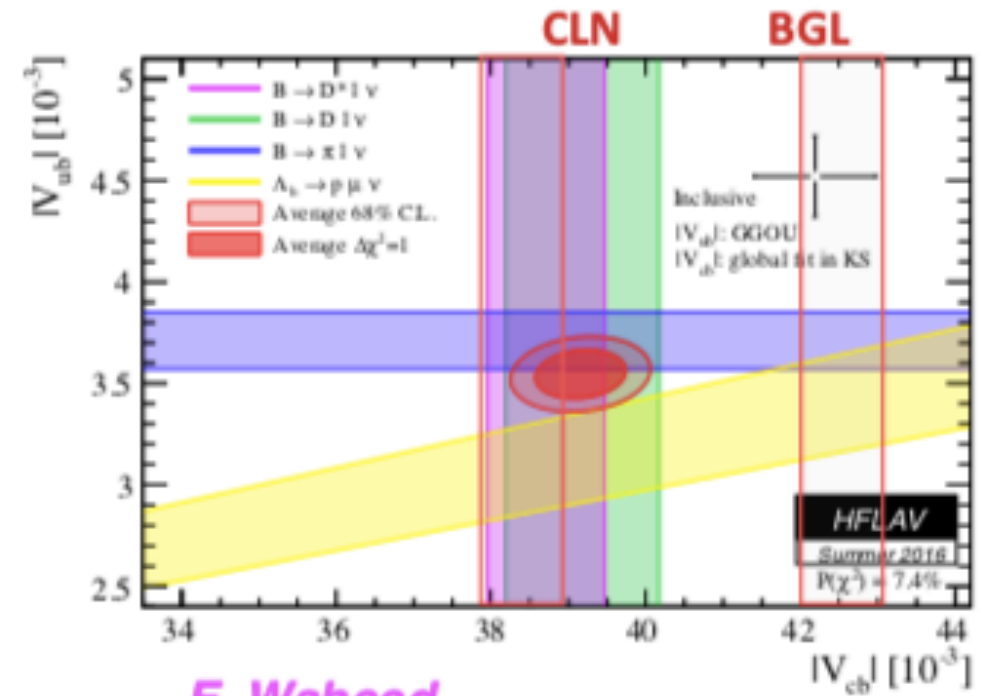
Flavor Physics

- CKM and CPV
 - $|V_{cb}|$ puzzle resolved (not $|V_{ub}|$)
 - new γ (φ_3) from LHCb

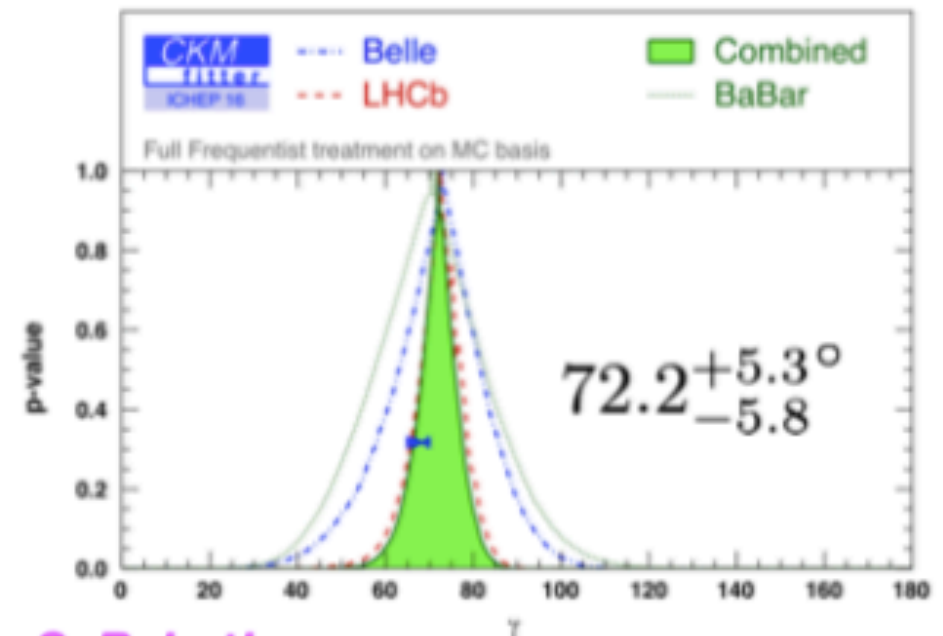


P. Urquijo

- | | |
|--|-------------------|
| $B \rightarrow \pi\pi, \rho\rho$ | α / Φ_2 |
| $B \rightarrow D^{(*)} K^{(*)}$ | γ / Φ_3 |
| $B \rightarrow J/\psi K_S$ | β / Φ_1 |
| $B_s \rightarrow J/\psi \Phi$ | β_s |
| $K \rightarrow \pi \nu \text{ anti-}\nu$ | ρ, η |



E. Waheed



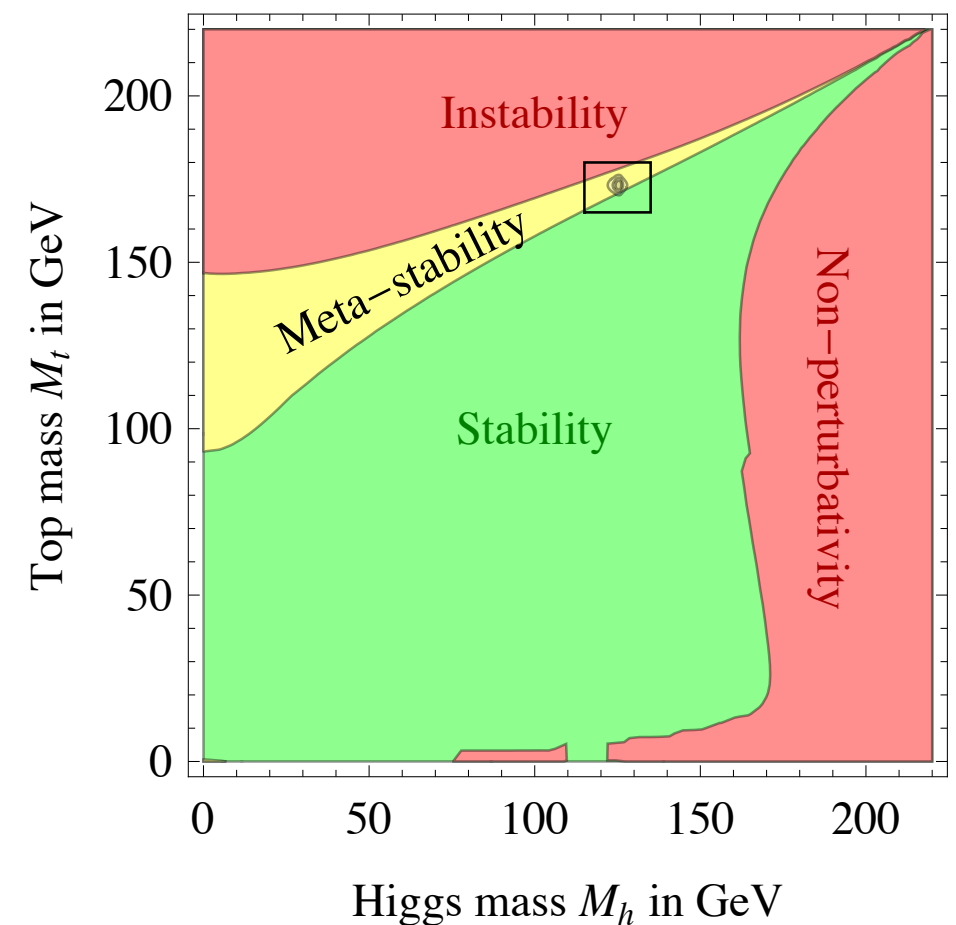
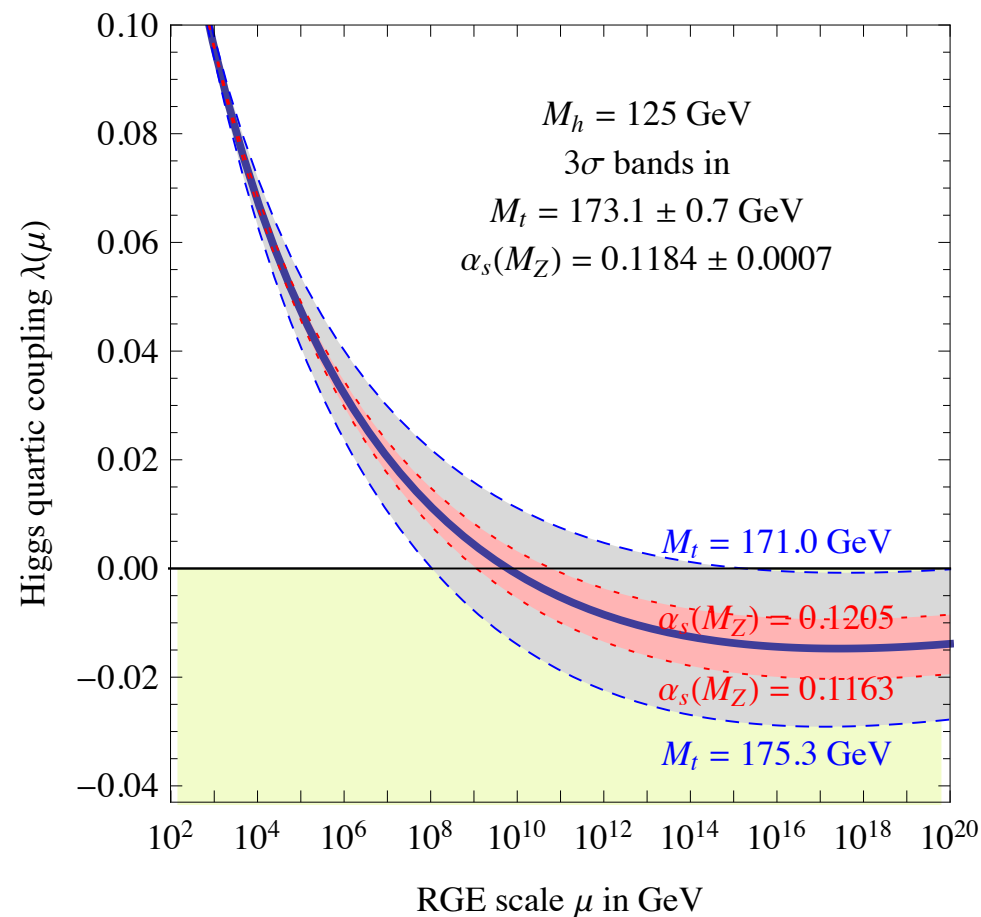
S. Rahatlou

Paul Langacker (IAS)

THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

The electroweak vacuum is unstable under radiative corrections

The whole construction of the SM may be in trouble being metastable or even unstable



the situation crucially depends on the top and Higgs mass values and requires severe fine-tuning and high accuracy of calculations (3 loops)

Muon anomalous magnetic moment

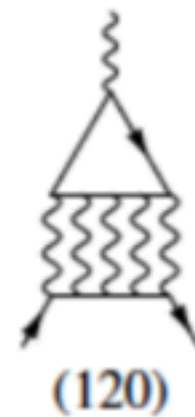
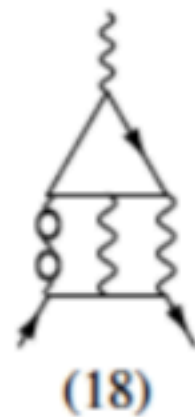
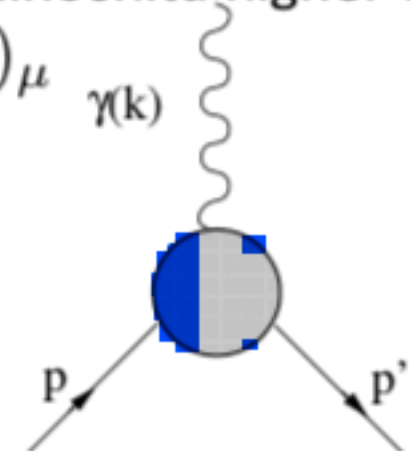
$$ie\bar{u}_\ell(p') \left[\gamma^\mu - \frac{a_\ell}{2m_\ell} i\sigma^{\mu\nu} q_\nu \right] u_\ell(p) \epsilon_\mu^*$$

(Schwinger α/π ,
Kinoshita higher orders in α)

$$q_\mu = (p - p')_\mu$$

Dirac equation predicts $g=2$ $a = (g - 2)/2$

For electron a_e theory and experiment agrees!



$$a_\mu^{th} - a_\mu^{exp} = -(3.06 \pm 0.76) \times 10^{-8} \quad 4\sigma$$

Theory: uncertainty in hadronic contributions to the muon $g - 2$, (Jägerlehner, 1802.08019).
Lattice QCD great progress light-by-light study (RBC & UKQCD, 1801.07224).

Fermilab and J-Park experiments are expected to clarify existing discrepancy!

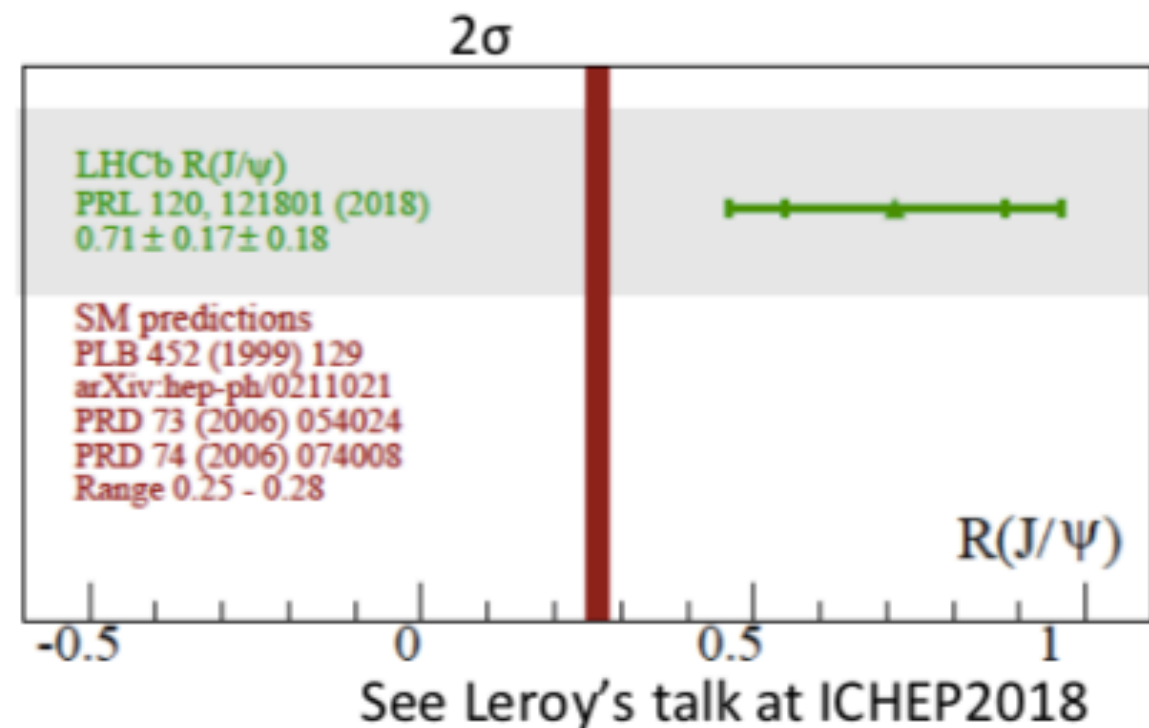
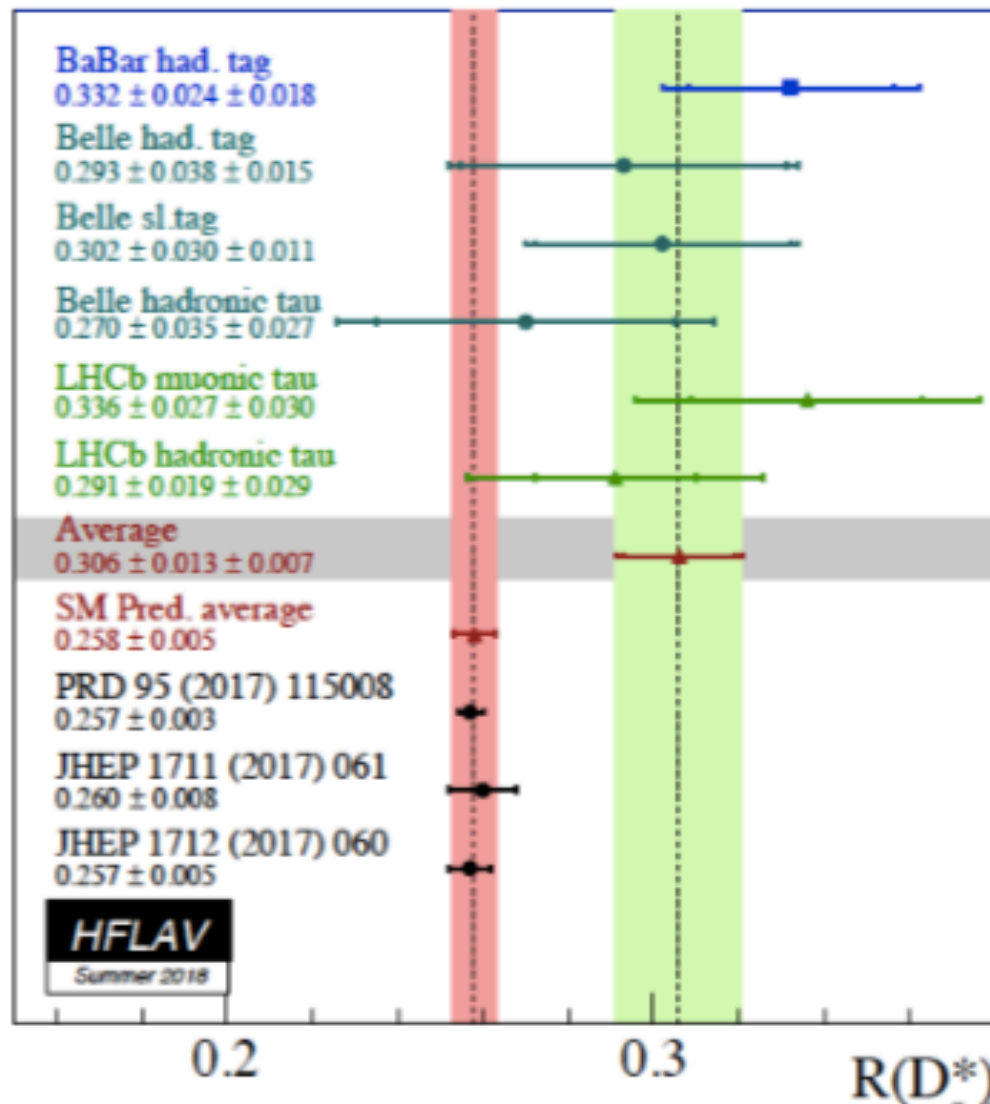
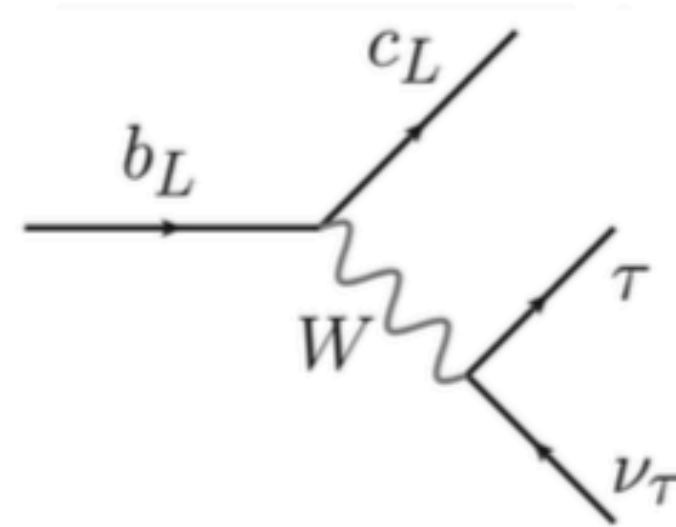
THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

B physics anomalies: experimental results \neq SM predictions!

charged current (SM tree level)

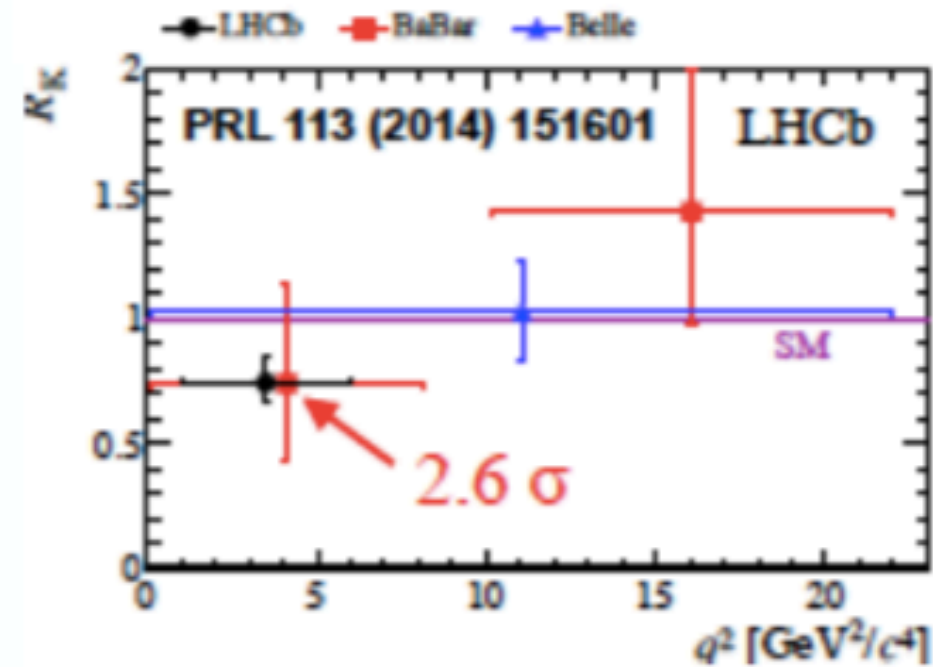
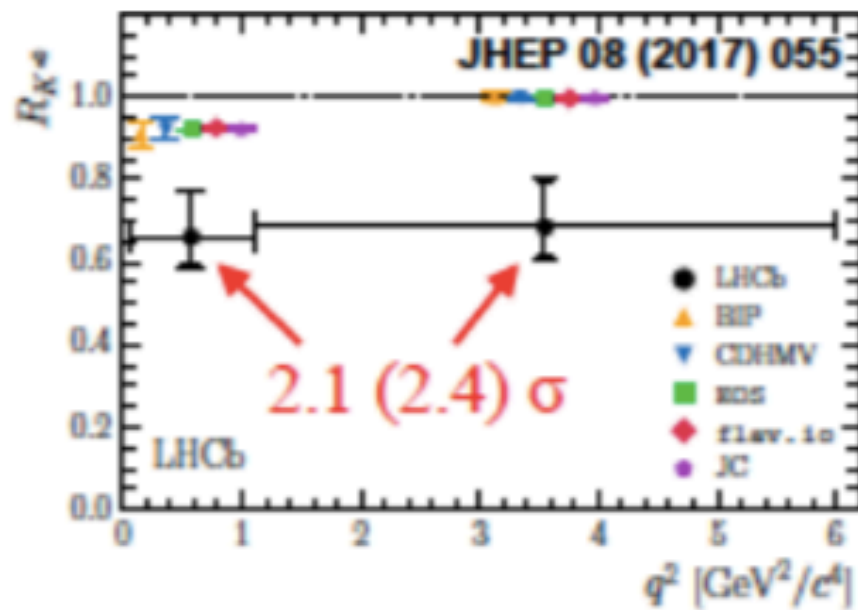
$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)}$$

3.8σ



FCNC - SM loop process: $R_{K^{(*)}}$ anomaly

$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)} \mu\mu)}{BR(B \rightarrow K^{(*)} ee)} \Bigg|_{q^2 \in [q_{min}^2, q_{max}^2]}$$

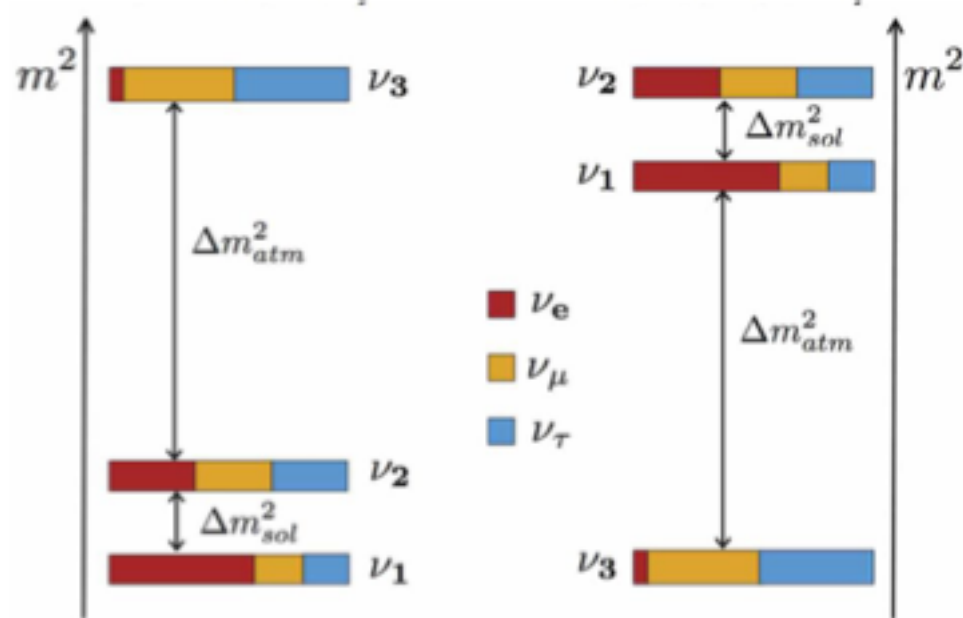


P_5' in $B \rightarrow K^* \mu^+ \mu^-$ (angular distribution functions) 3σ

LHCb: the discrepancy present in $B_s \rightarrow \phi \mu\mu$ and $\Lambda_b \rightarrow \Lambda \mu\mu$

(see Capriotti talk at ICHEP2018)

Neutrino Physics



- Absolute value of neutrino masses ?
- Mass hierarchy?
- Dirac or Majorana?
- Fourth sterile neutrino?
- Neutrino dark matter?

$$0.06 \text{ eV} < \sum m_\nu < 0.12 \text{ eV}$$

↑
↑
ν-OSC
CMB

PMNS-matrix parameters are measured with high accuracy of few %

- Normal hierarchy favoured at 3.1 σ
- Nonzero CP phase favoured
- Upper octant favoured

parameter	best fit $\pm 1\sigma$	3 σ range
Δm_{21}^2 [10^{-5}eV^2]	$7.55^{+0.20}_{-0.16}$	7.05–8.14
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (NO)	2.50 ± 0.03	2.41–2.60
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (IO)	$2.42^{+0.03}_{-0.04}$	2.31–2.51
$\sin^2 \theta_{12}/10^{-1}$	$3.20^{+0.20}_{-0.16}$	2.73–3.79
$\sin^2 \theta_{23}/10^{-1}$ (NO)	$5.47^{+0.20}_{-0.30}$	4.45–5.99
$\sin^2 \theta_{23}/10^{-1}$ (IO)	$5.51^{+0.18}_{-0.30}$	4.53–5.98
$\sin^2 \theta_{13}/10^{-2}$ (NO)	$2.160^{+0.083}_{-0.069}$	1.96–2.41
$\sin^2 \theta_{13}/10^{-2}$ (IO)	$2.220^{+0.074}_{-0.076}$	1.99–2.44
δ/π (NO)	$1.32^{+0.21}_{-0.15}$	0.87–1.94
δ/π (IO)	$1.56^{+0.13}_{-0.15}$	1.12–1.94

Is it just the SM or requires New physics?

Three Types of Seesaw Mechanisms

Require the existence of new degrees of freedom (particles) beyond those present in the SM

Type I seesaw mechanism: ν_{IR} - RH ν s' (heavy).

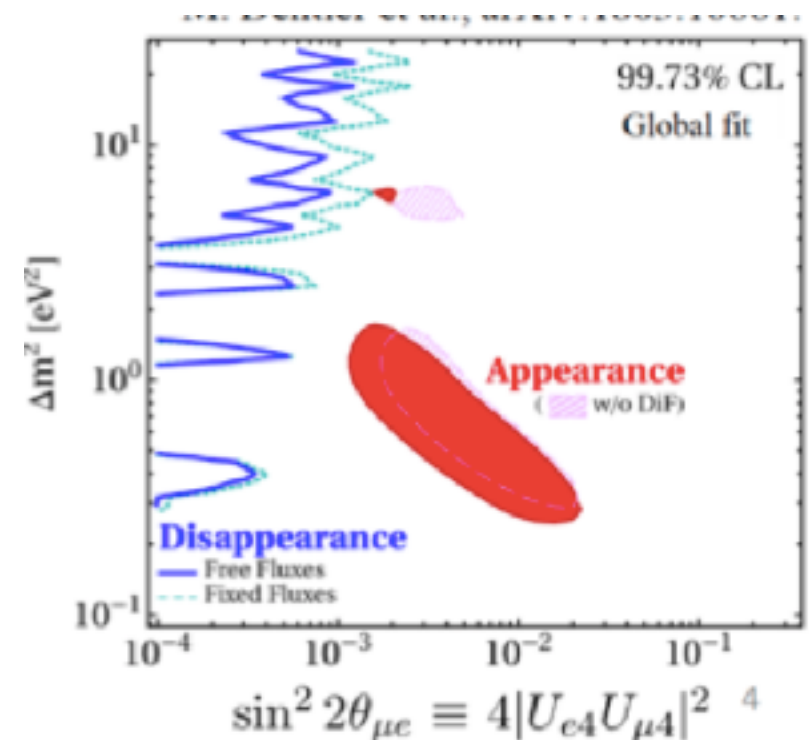
Type II seesaw mechanism: $H(x)$ - a triplet of H^0, H^-, H^{--} Higgs fields.

Type III seesaw mechanism: $T(x)$ - a triplet of fermion fields.

M. Weber ICHEP2018

• Possible Sterile Neutrino?

- **New MiniBooNE consistent with LSND (but low energy excess?)**
- **Reactor anomaly questioned by Daya Bay/RENO time dependence**
- **New SBL and source experiments**
- **Conflict with ν_μ disappearance**



Major problem: 85% of matter is dark and remains invisible!

Is this compatible with the SM?

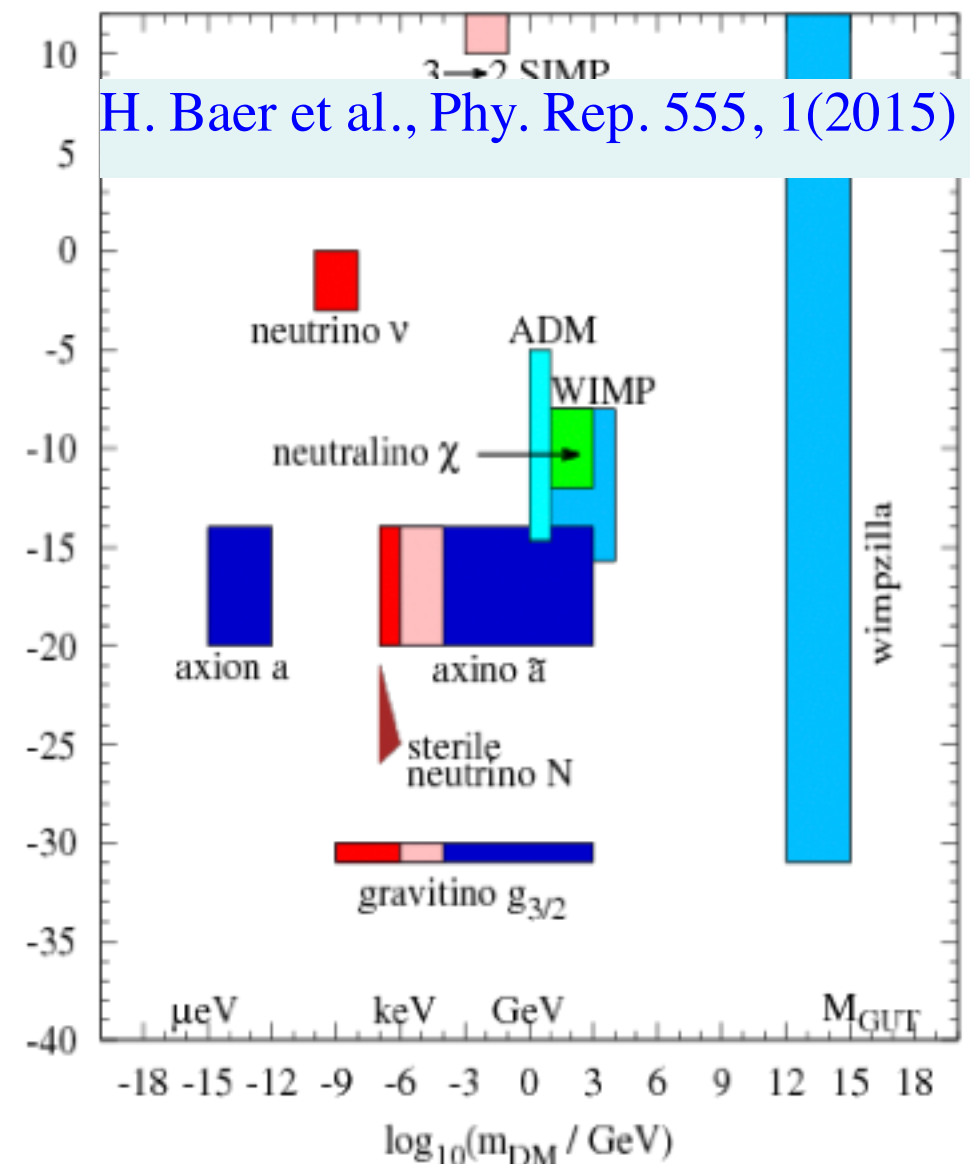
Does it requires modification of the SM or addition of gravity?

• Many candidates in many orders of magnitude of mass:

- **MOND** (Problems: large scales, Bullet cluster)
- **Primordial black holes** (LIGO, but constraints)
- **Fuzzy** (very light bosons)
- **Warm** (KeV sterile)
- **WIMP**
- **Axions/ALPs**
- **Dark sector**
- **Gravitinos**
- **Moduli**
- **Wimpzillas**



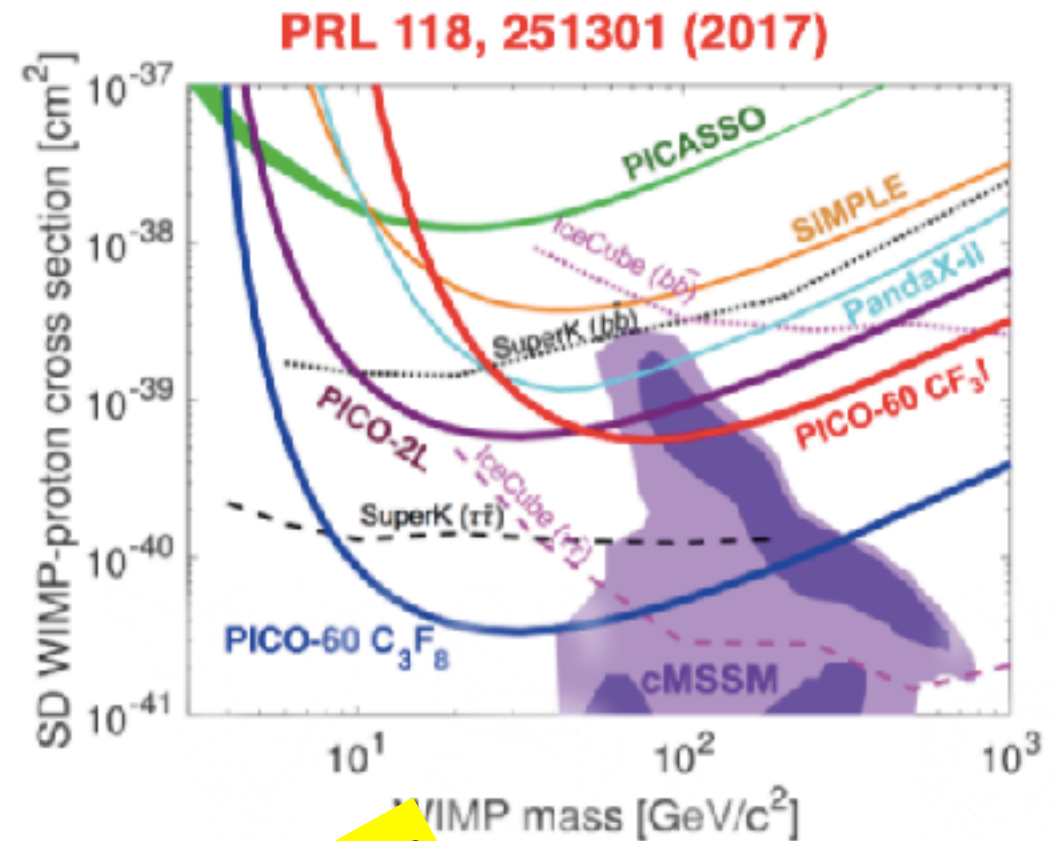
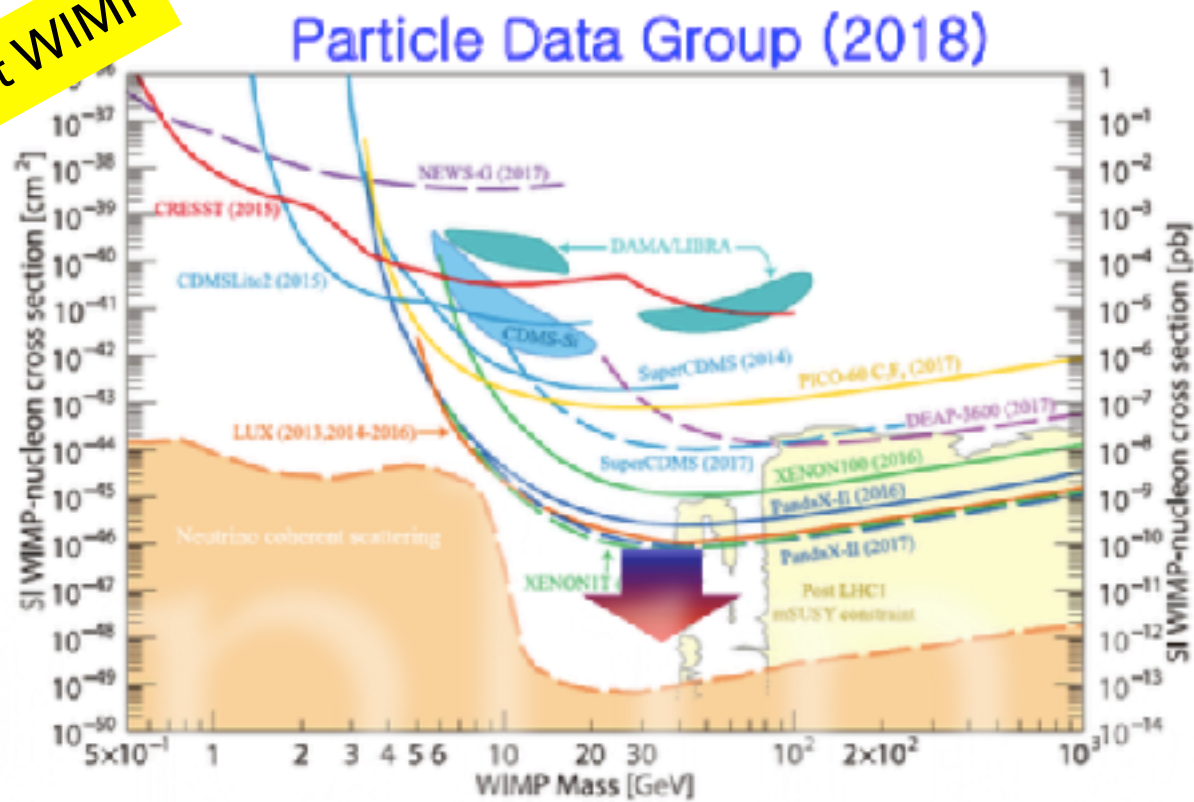
M. Drees



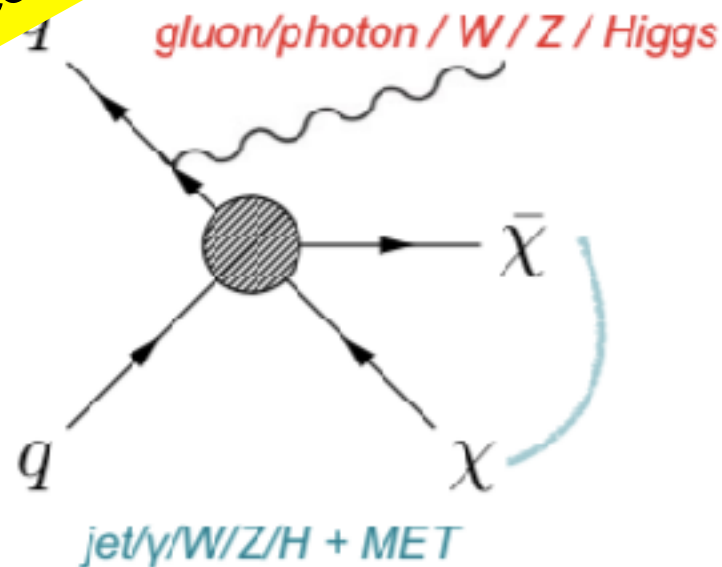
• Direct, indirect, collider

BEYOND THE STANDARD MODEL: DARK MATTER SEARCHES

Direct WIMP



Colliders WIMP

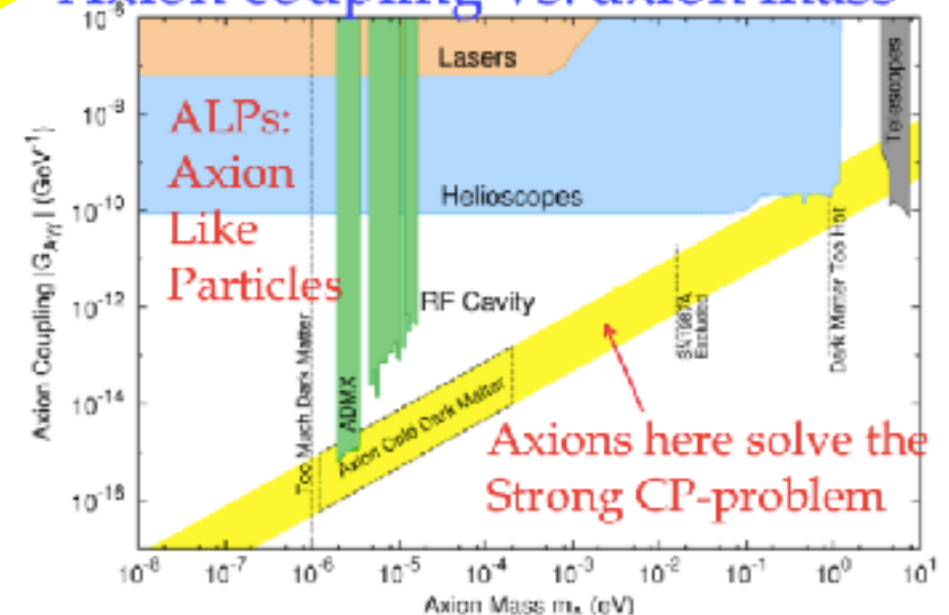


- **mono-jet**
 - most general signature, constraints on many models
- **mono-photon**
 - more challenging for background estimation
 - less powerful: EW vs. strong interaction
- **mono-W/Z leptonic**
 - clean signature and simple trigger
 - penalized by W/Z branching fraction
- **mono-W/Z hadronic**
 - larger statistics with larger background
- **tt+MET/bb+MET and mono-top**
 - more complicated experimentally
 - powerful in some scenarios
- **mono-Higgs**
 - powerful in some scenarios

D. del Re

Axion-likes

Axion coupling vs. axion mass



Y. Semertzidis

BEYOND THE STANDARD MODEL: THE MASS SPECTRUM AND MIXINGS

- Mass spectrum?

$$m_{quark} = y_{quark} \cdot v$$

$$m_{lepton} = y_{lepton} \cdot v$$

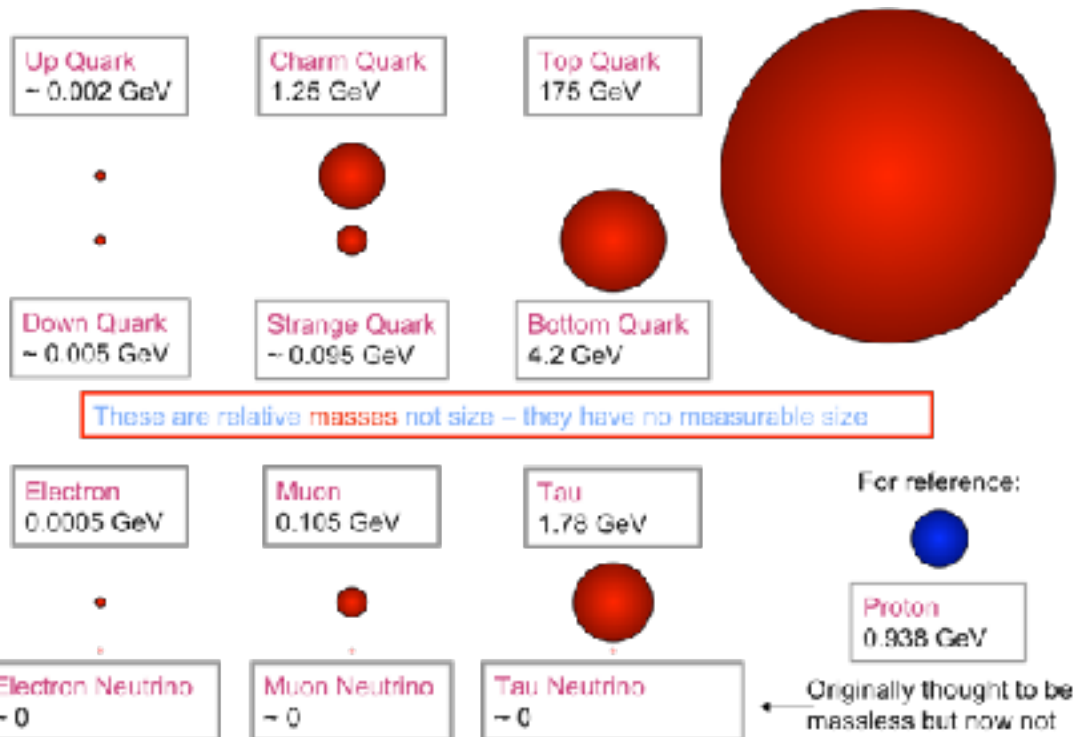
$$m_W = g/\sqrt{2} \cdot v$$

$$m_Z = \sqrt{g^2 + g'^2}/\sqrt{2} \cdot v$$

$$m_H = \sqrt{\lambda} \cdot v$$

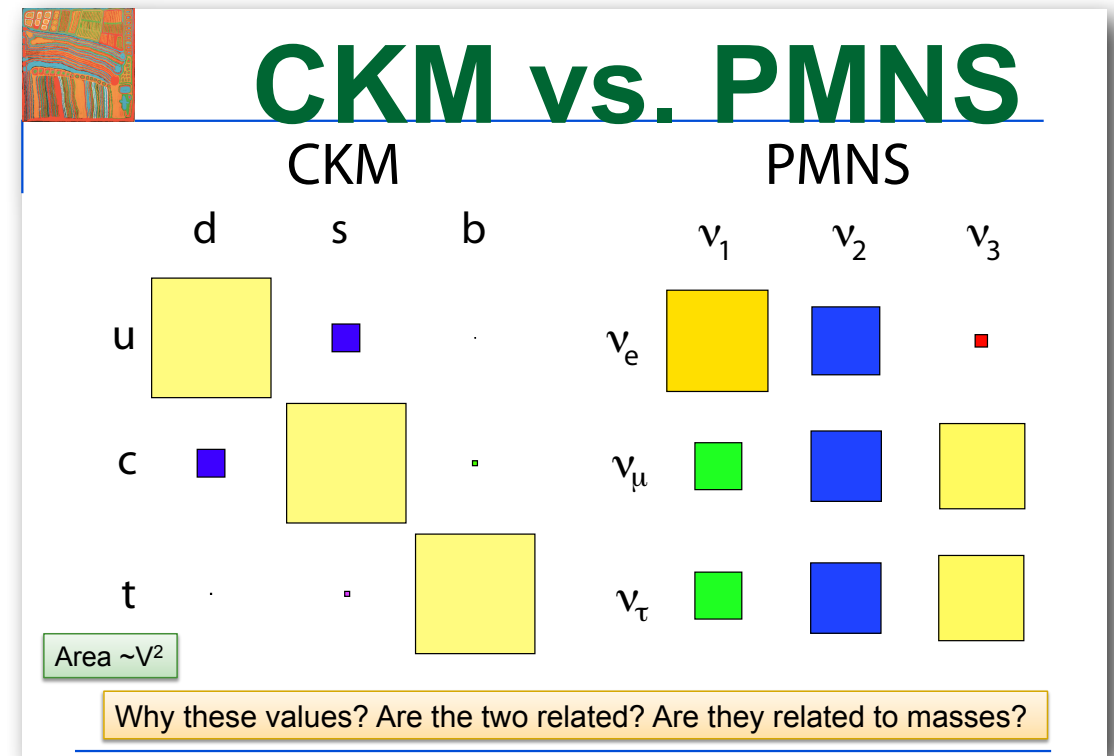
SM $m_\gamma = 0$

$$m_{gluon} = 0$$



- Mixing Matrices?

- Quark-Lepton Symmetry
- Strong difference in parameters



- What are the CKM and PMNS phases?
- Where lies the source of CP violation: in quark or lepton sector?

$$J_{CP} = \frac{1}{8} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta$$

Neutrino Mixing: New Symmetry?

- $\theta_{12} = \theta_{\odot} \simeq \frac{\pi}{5.4}$, $\theta_{23} = \theta_{\text{atm}} \simeq \frac{\pi}{4}(\?)$, $\theta_{13} \simeq \frac{\pi}{20}$

$$U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}; \quad U_{\text{BM}} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \pm \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \pm \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & \mp \frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix}. \quad U_{\text{HGM}} = \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & 0 \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

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Follows the attempts in quark sector with 30 years delay: so far unsuccessful

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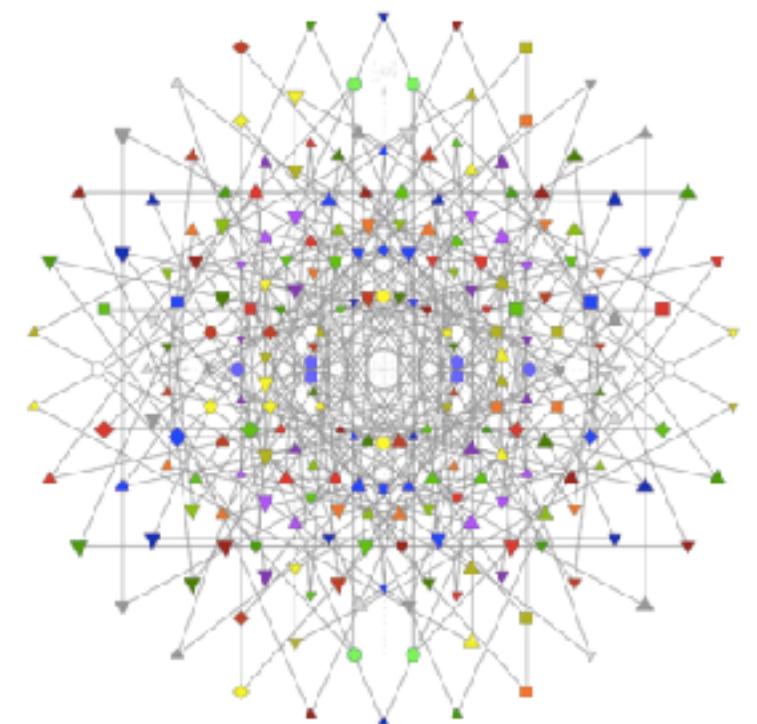


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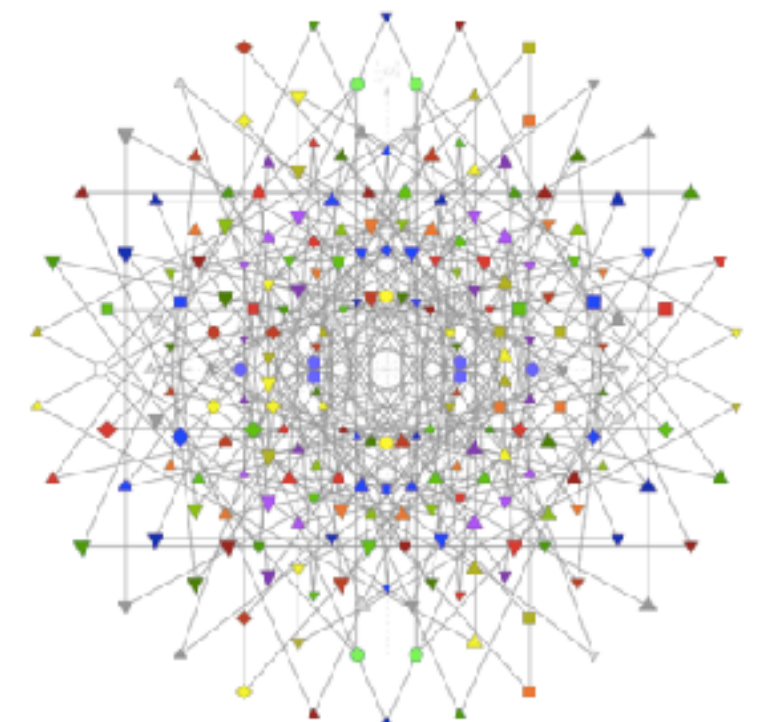
E8 roots

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Symmetry might be tricky

E8 roots

THE STANDARD MODEL: CONCEPTUAL PROBLEMS

- Baryon asymmetry of the Universe

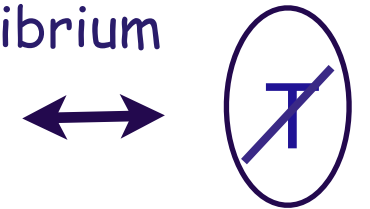
$$\frac{N(B) - N(\bar{B})}{N_\gamma} \sim (6.19 \pm 0.14) \times 10^{-10}$$

- still not explained
- three conditions (A.D.Sakharov)

- Violation of a thermal equilibrium in early Universe

A possible scenario in the early Universe when particles drop from thermal equilibrium violations T invariance

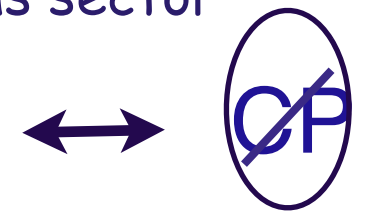
2. Violation of baryon number \longleftrightarrow B $B = \frac{N_q - N_{\bar{q}}}{3}$



Baryon number is conserved in the SM with exponential accuracy

Violation of baryon number occurs in Grand Unified Theories and in Lepton=fourth color models (Pati-Salam model) } New particles = Leptoquarks, Extended Higgs sector

3. Violation of CP invariance (requires larger CP than in the SM)



In the SM achieved via phase factors in the CKM and PMNS mixing matrices

The presence of new phase factors in extended models (2HDM, SUSY, etc)

CPT is exact symmetry of Nature

BEYOND THE STANDARD MODEL: CONCLUSIONS

WHAT MAKES US THINK THAT THERE IS PHYSICS BEYOND THE STANDARD MODEL?

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- Baryon asymmetry of the Universe is a fundamental problem (Baryon and Lepton genesis might require new ingredients)
- Lack of understanding of flavor structure of the SM calls for explanation at higher level
- New era in gravity due to discovery of gravitational waves and black holes might change the landscape

Physics with a single generation

Back to the middle of the XX century



Physics with a single generation

Back to the middle of the XX century



All the world around us is made of the 1st generation

Physics with a single generation

Back to the middle of the XX century



All the world around us is made of the 1st generation

p
1919

n
1932

e
1897

Physics with a single generation

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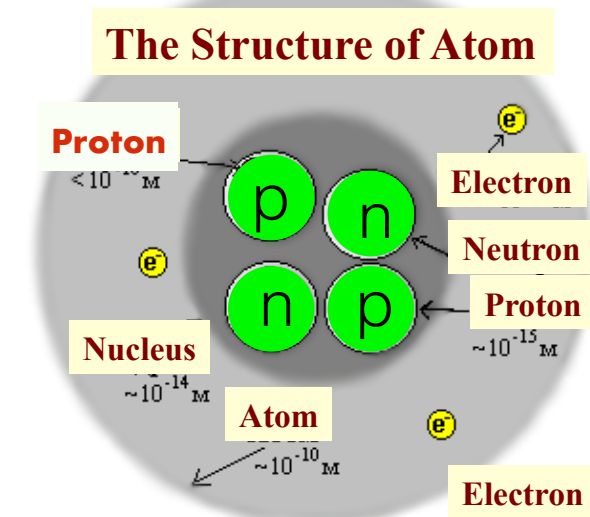
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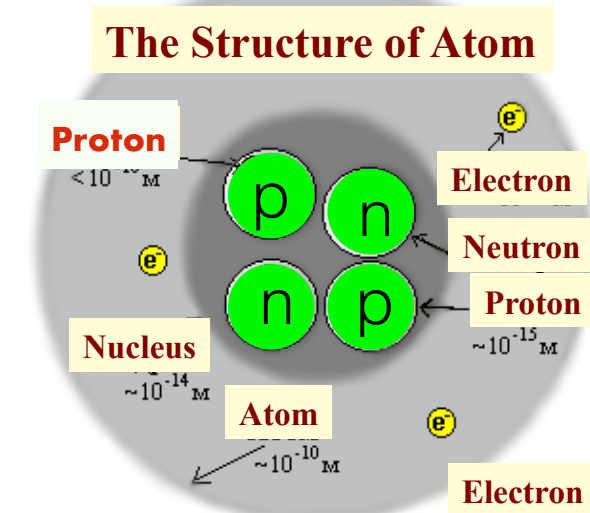
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- Who expected new physics to come?

Physics with a single generation

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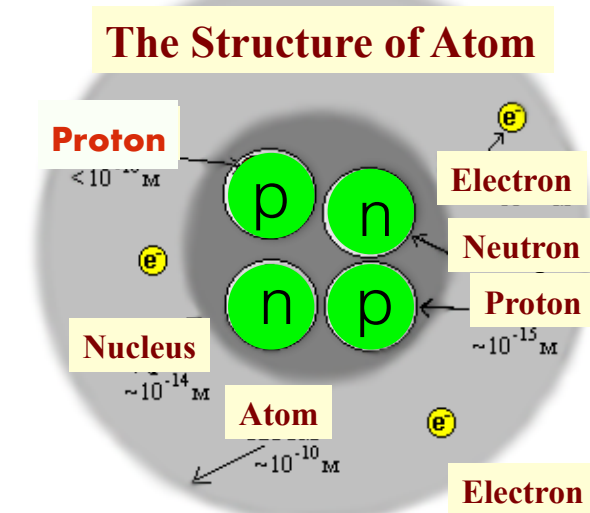
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- Who expected new physics to come?
- What scale of NP?

Physics with a single generation

Back to the middle of the XX century



All the world around us is made of the 1st generation

p
1919

ν_e
1956

ν_μ
1963

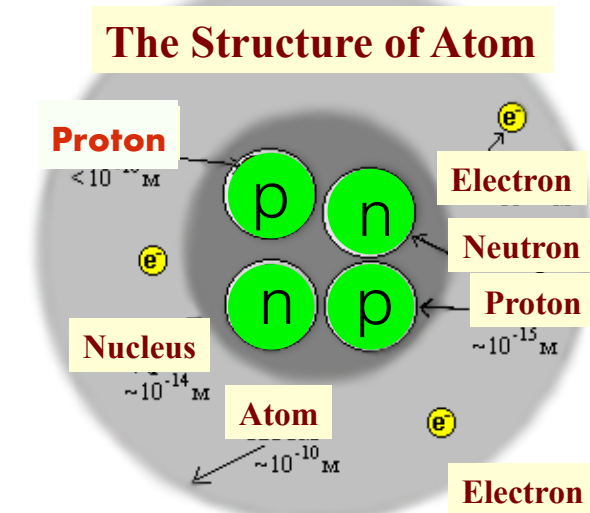
n
1932

e
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- Muon - heavy electron - 2nd generation ?



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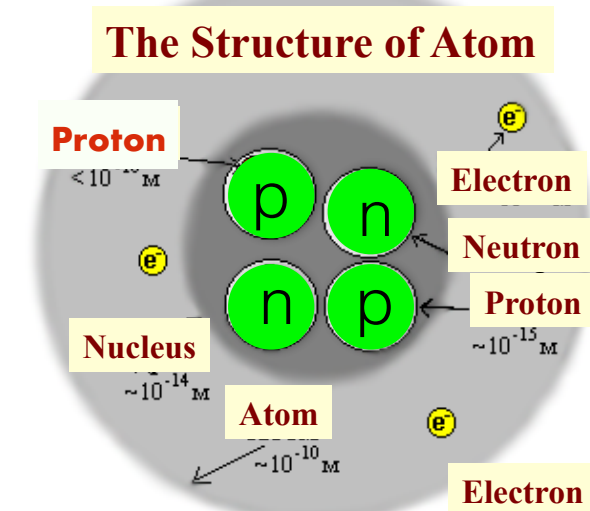
n
1932

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γ

- Muon - heavy electron - 2nd generation ?
- K-meson - strangeness ?



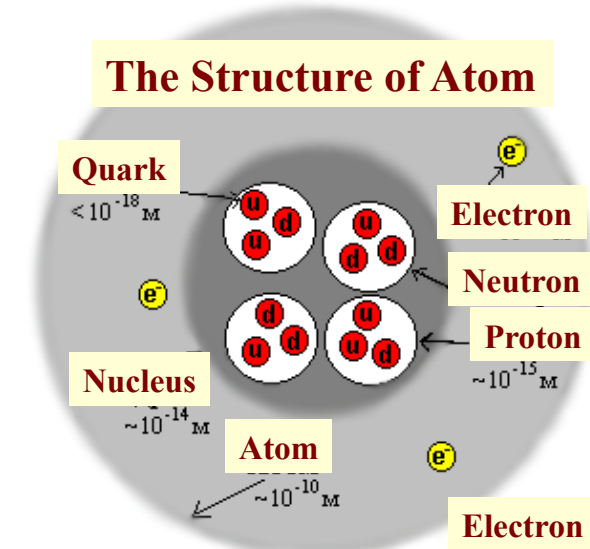
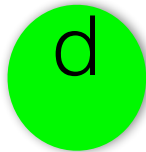
- Who expected new physics to come?
- What scale of NP?

Physics with a single generation

Back to the middle of the XX century



All the world around us is made of the 1st generation



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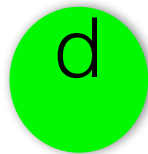
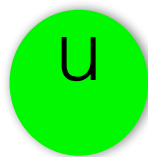
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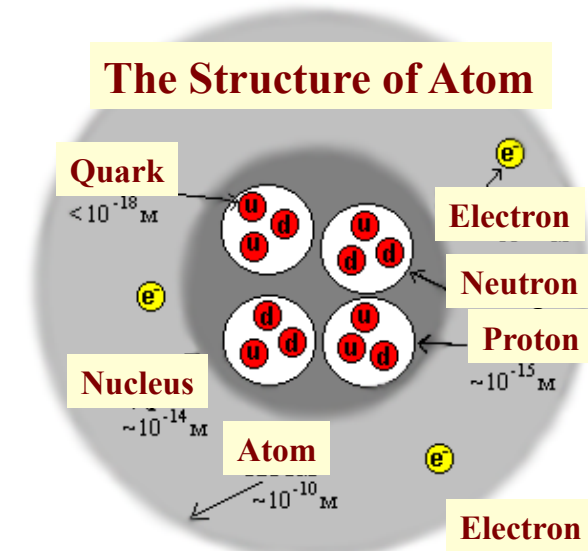
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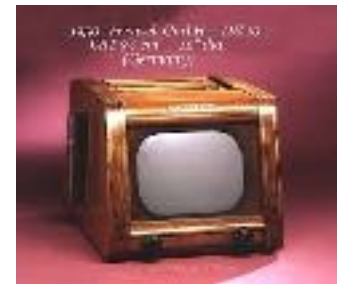
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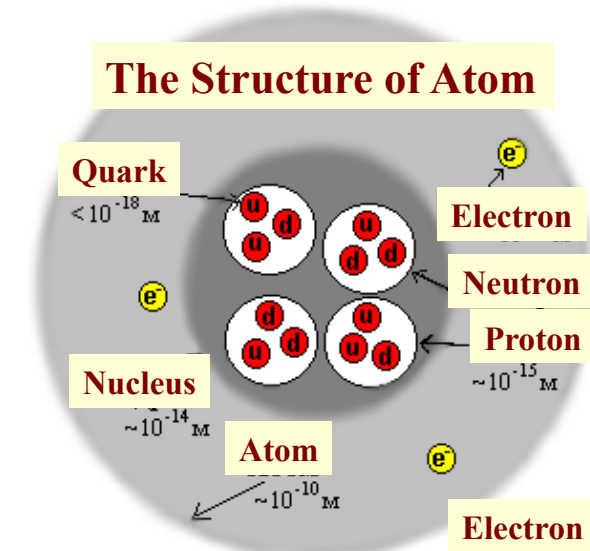
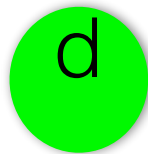
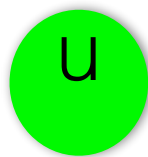
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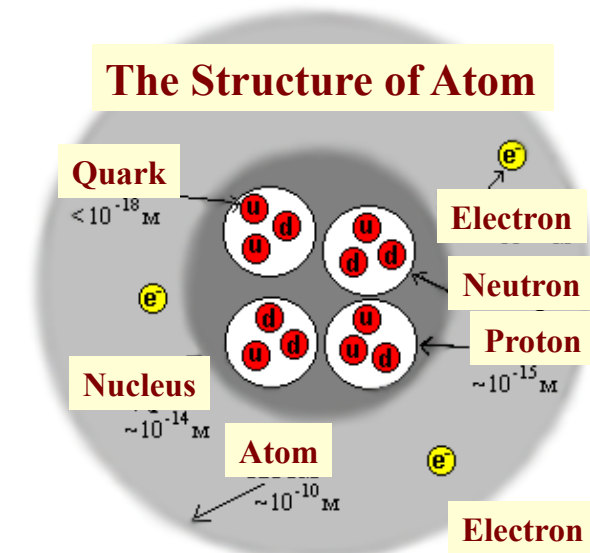
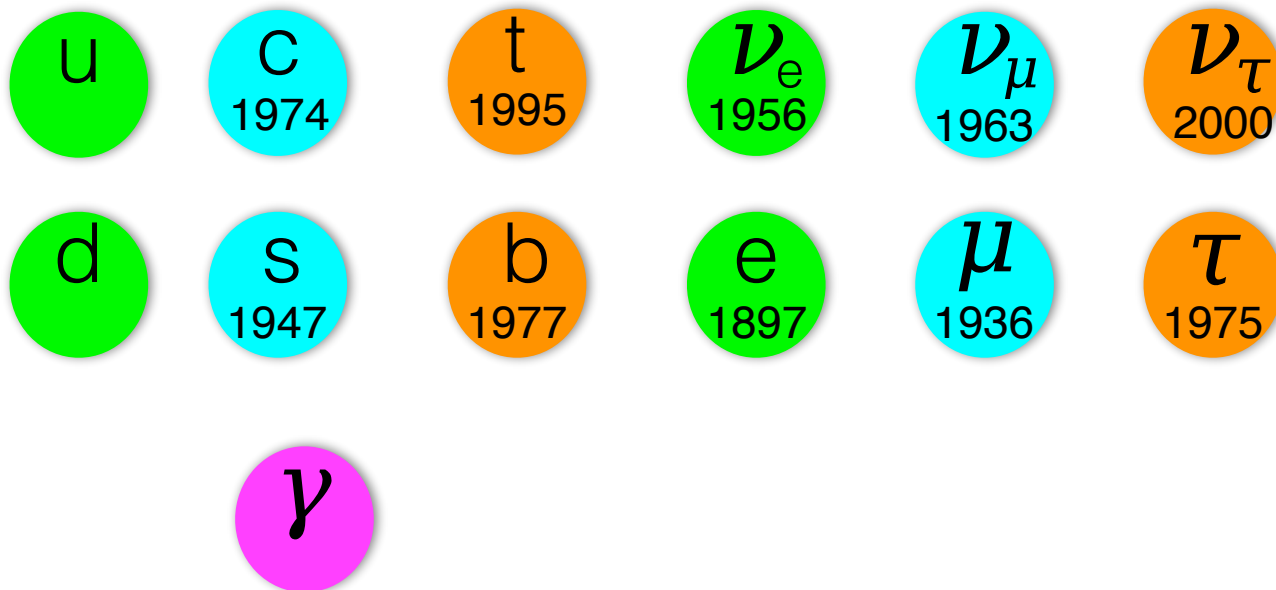
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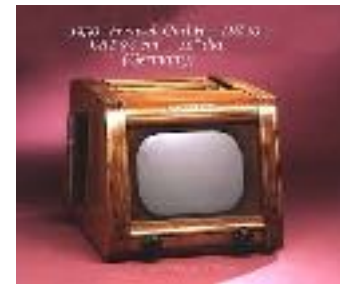


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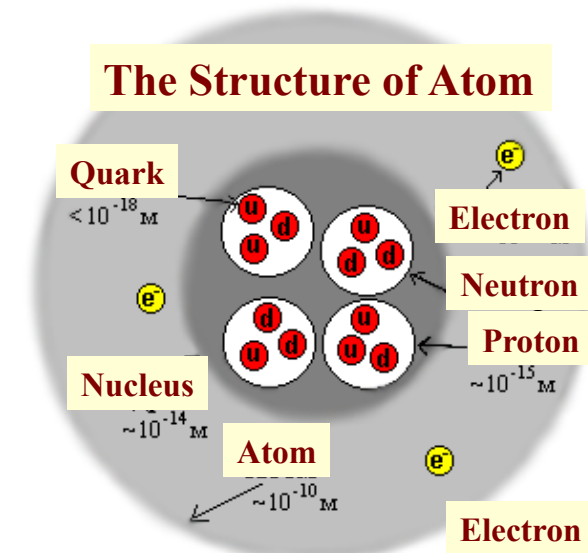
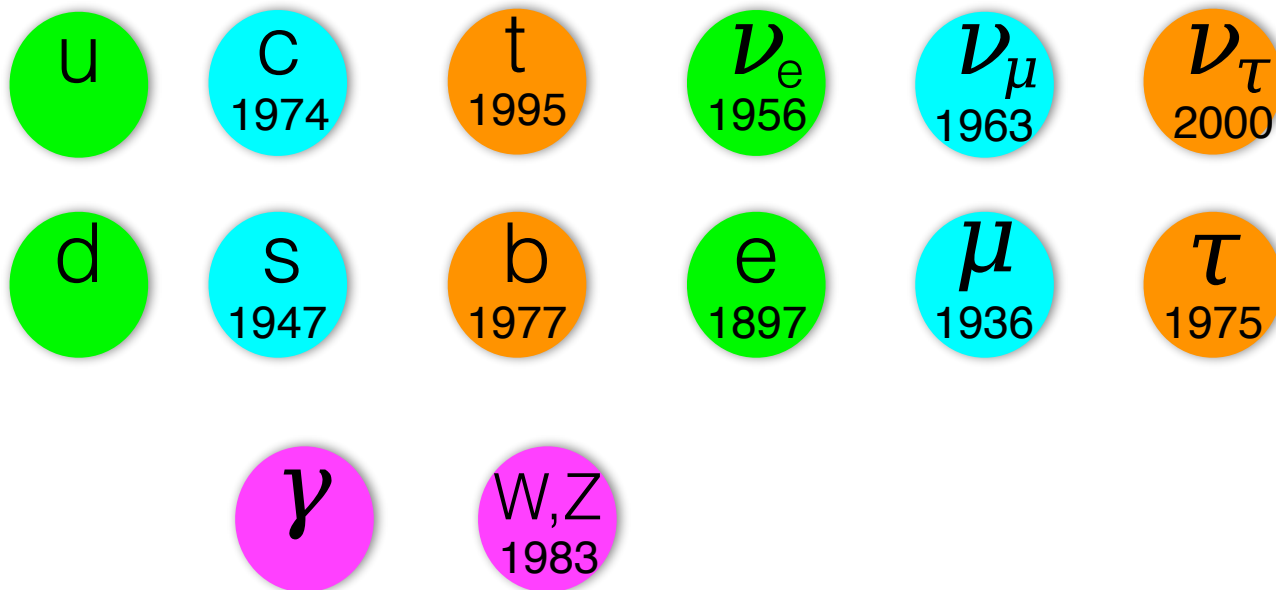
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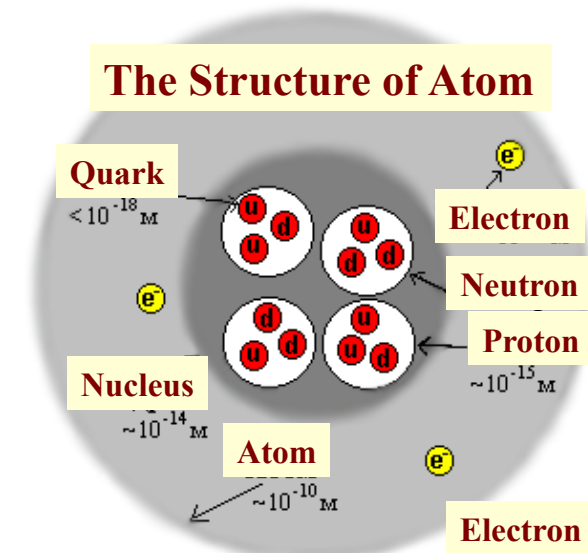
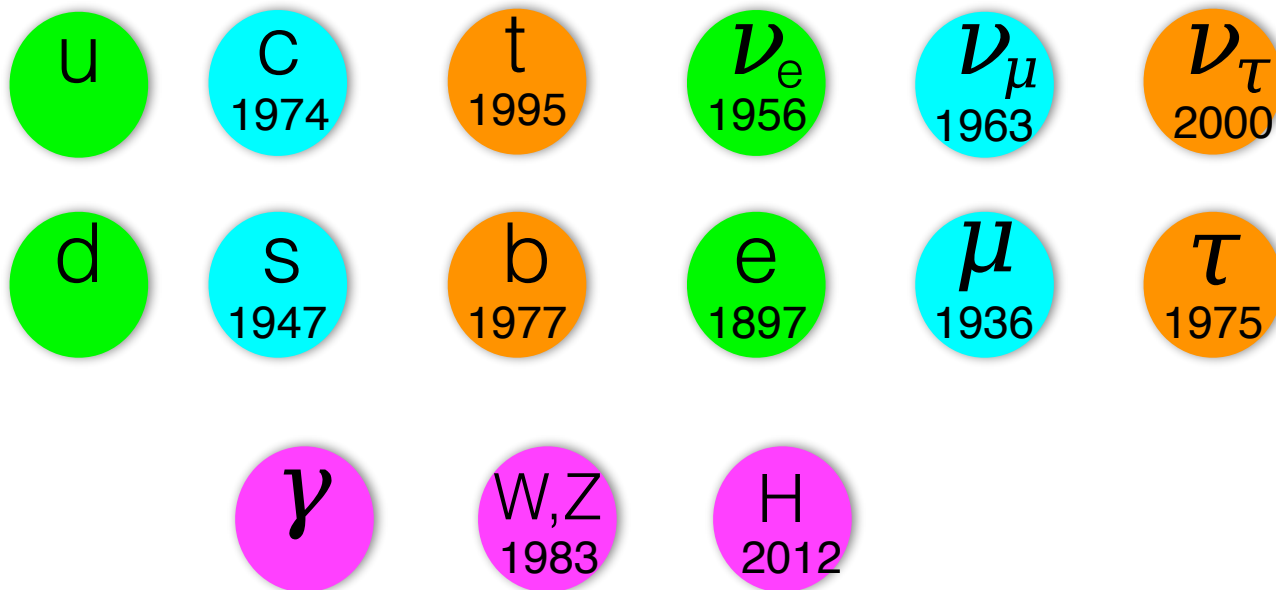
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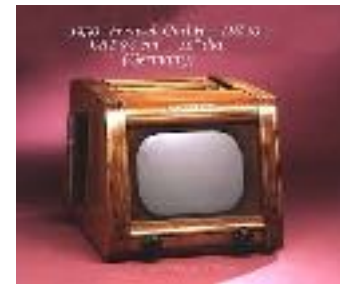


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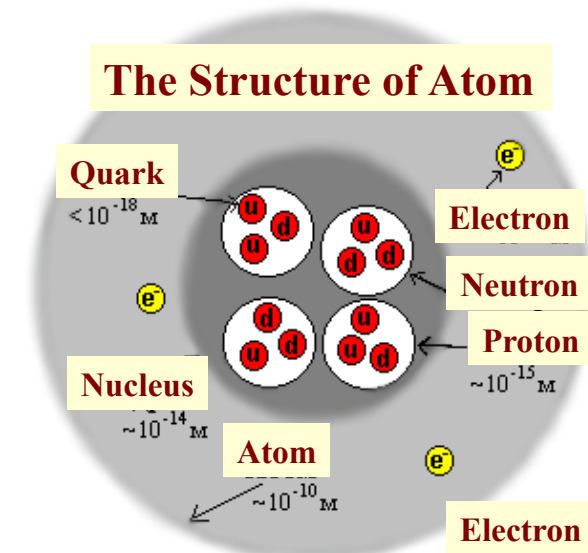
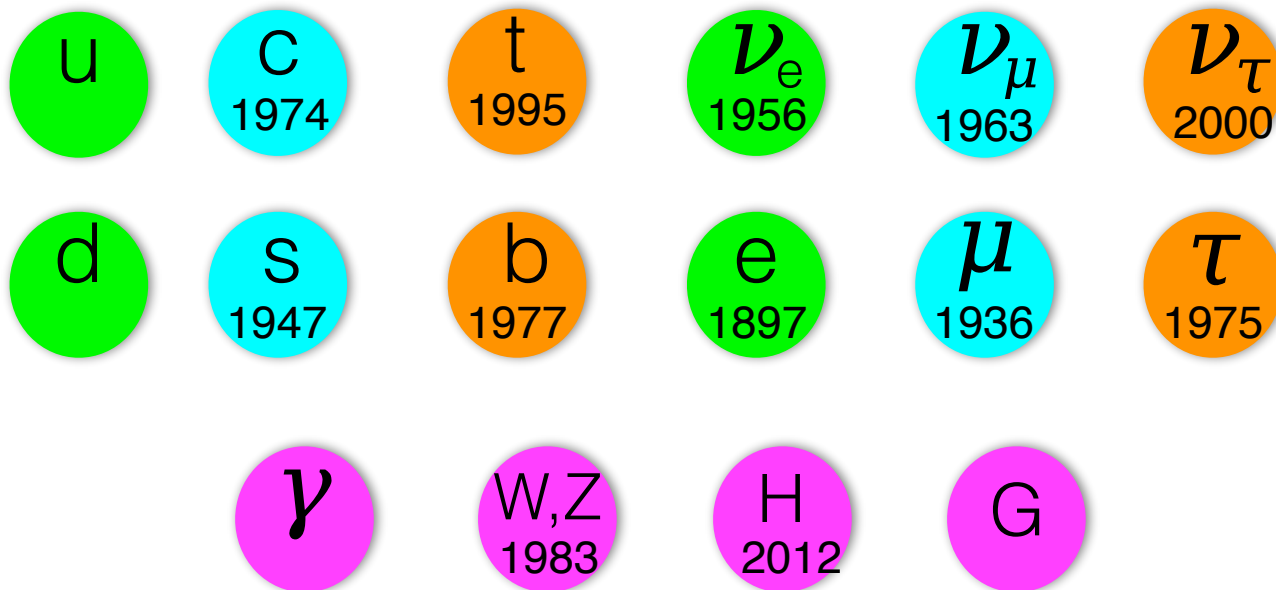
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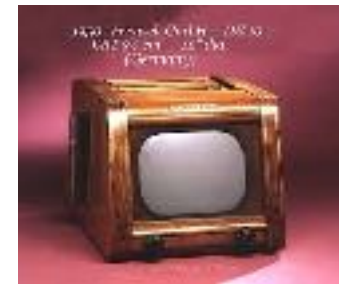


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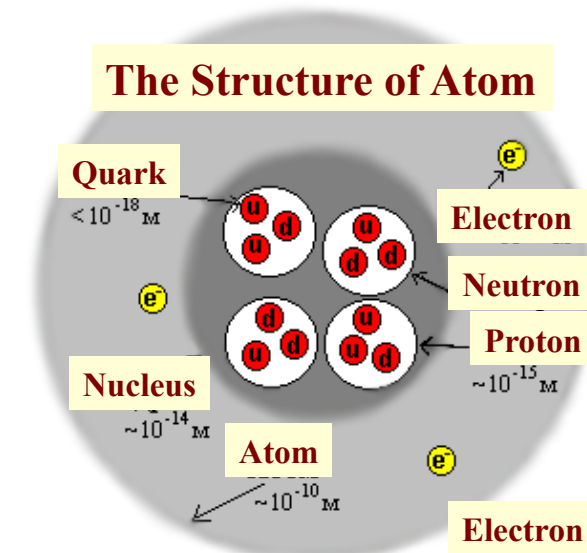
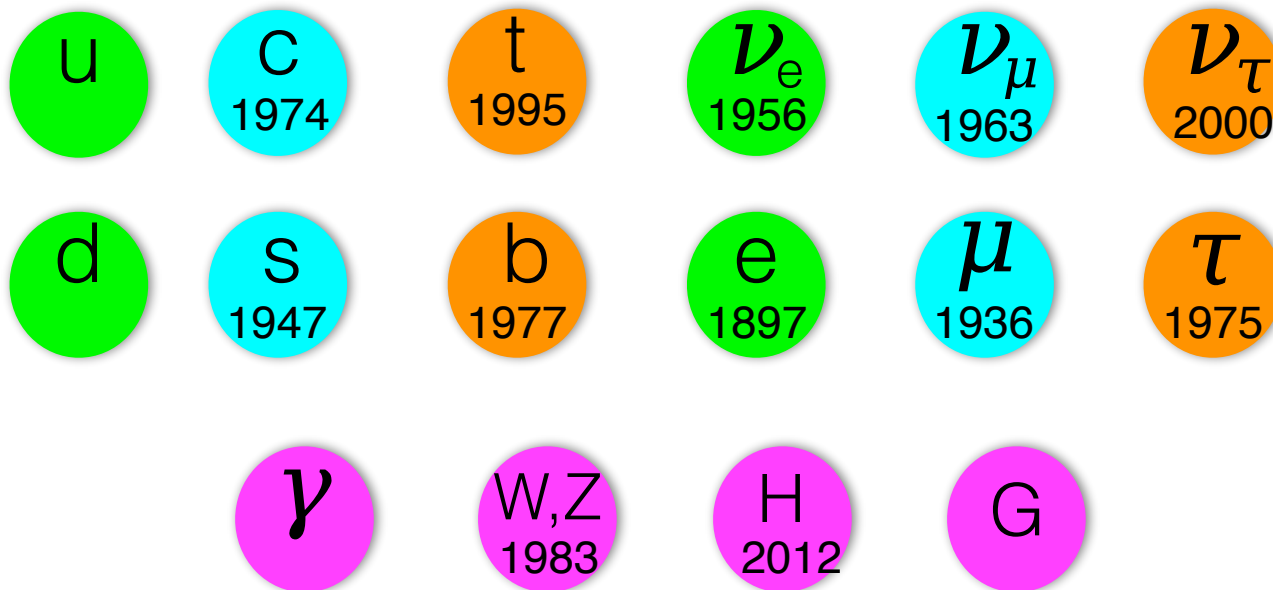
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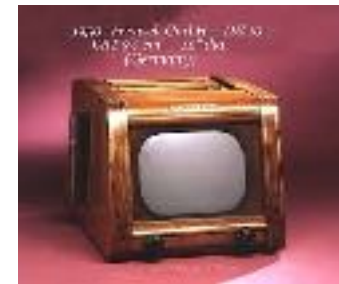
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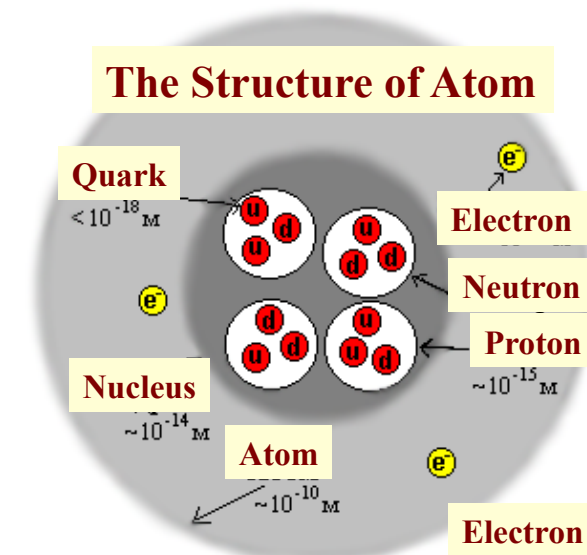
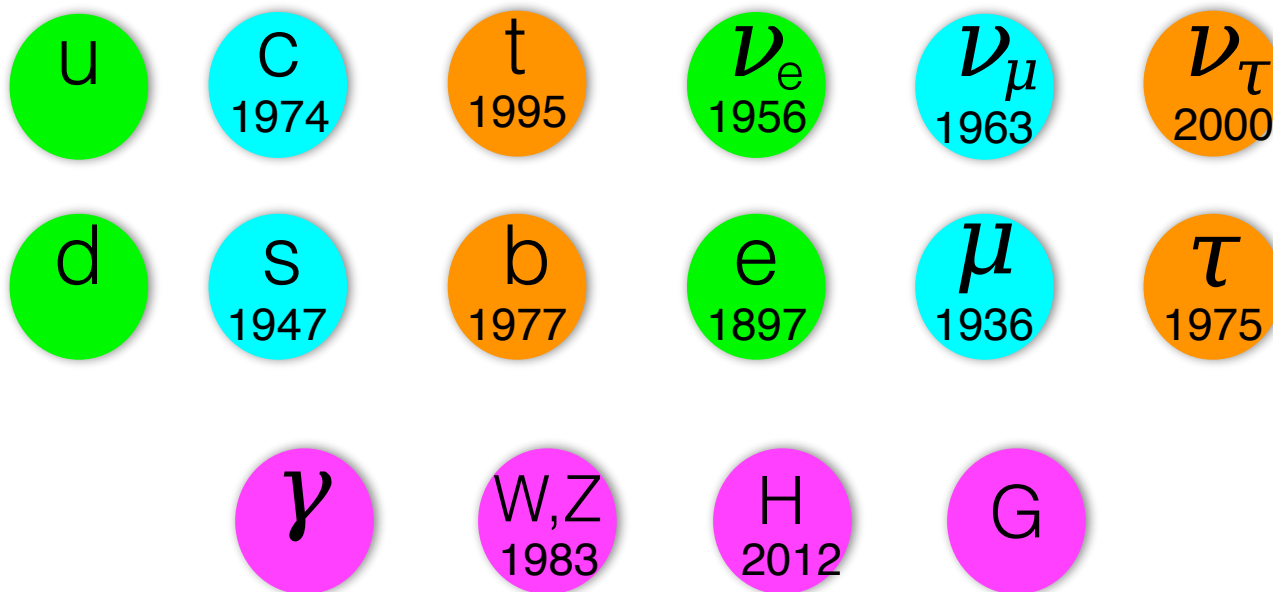
Astrophysics & Cosmology challenge

Physics with a single generation

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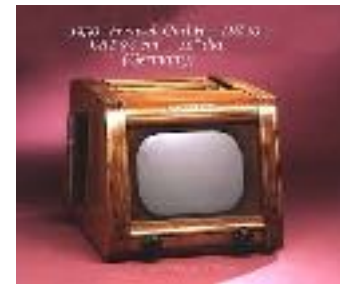
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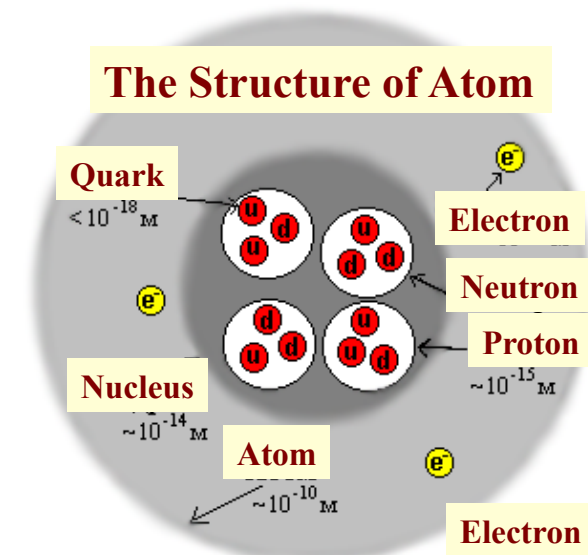
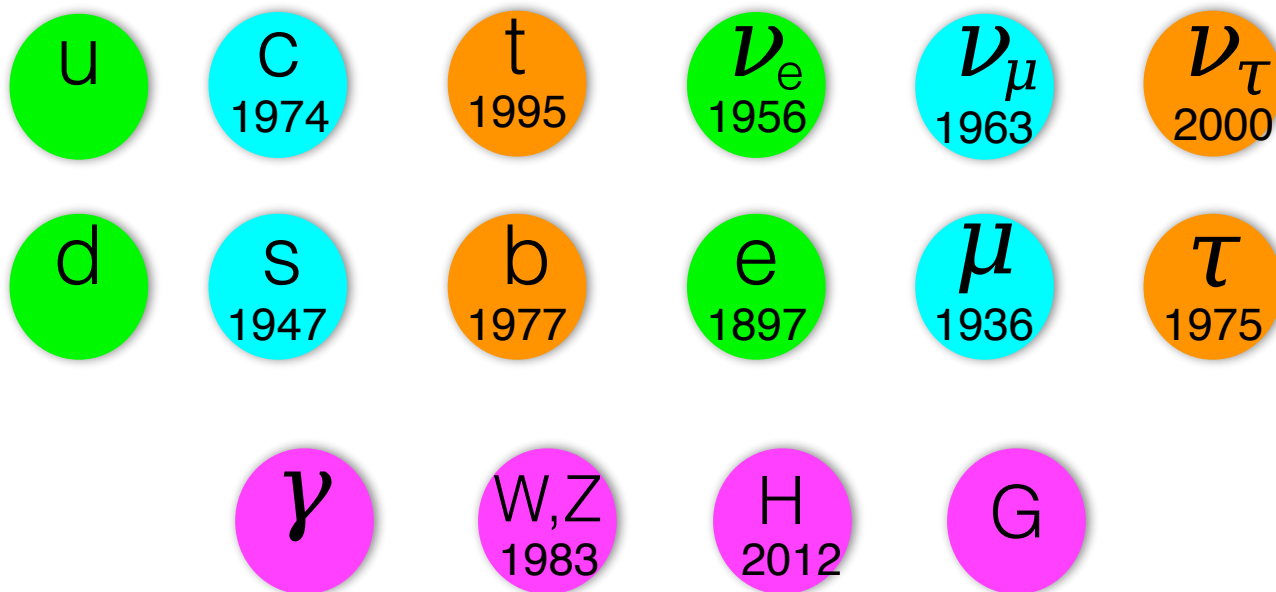
- Baryon asymmetry of the Universe
- Description of the Dark Matter

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Astrophysics & Cosmology challenge

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Ideas (conventional and not)

- **Symmetries**
 - Supersymmetry, family, ...
- **Compositeness**
 - Higgs, fermions, ...
- **Extra dimensions**
 - large, warped, ...
- **Dark or hidden sectors**
 - Dark, SUSY-breaking, random, ...
- **Unification**
 - GUT, string, ...
- **New dynamical ideas**
 - Relaxion, unnaturalness, clockwork, string instantons, ...
- **Random or environmental**
 - multiverse
- **String remnants**
 - (need not solve SM problem)
 - Z' , vector fermions, extended Higgs, dark, moduli, axions, ...

BEYOND THE STANDARD MODEL: CONCLUSIONS



BEYOND THE STANDARD MODEL: CONCLUSIONS



BEYOND THE STANDARD MODEL: CONCLUSIONS



How Will We Make Progress?

• • •

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• • •

The energy frontier

The precision frontier and neutrinos

Cosmology and astrophysics