

Conformal viscous hydrodynamics

and

finite temperature gauge/gravity duality

Rudolf Baier

Department of Physics, University of Bielefeld

mainly based on following work - and references therein - by:

R. Baier, P. Romatschke, D. T. Son, A. O. Starinets and
M. A. Stephanov

“Relativistic viscous hydrodynamics, conformal invariance, and holography”

JHEP 0804 (2008) 100 (arXiv: 0712.2451 [hep-th])

R. Loganayagam

“Entropy current in conformal hydrodynamics”

JHEP 0805 (2008) 087 (arXiv: 0801.3701 [hep-th])

P. Romatschke

“New developments in relativistic viscous hydrodynamics”

Int.J.Mod.Phys E19 (2010) 1 (arXiv: 0902.3663 [hep/ph])

CONTENT

- motivation: viscous effects in RHIC data
viscosity η - relaxation time τ_π, \dots
- conformal invariance,
relativistic hydrodynamics -
up to second-order gradients
- gauge/gravity duality -
(AdS/CFT) correspondence

A fluid which has no shear stresses, viscosity or heat conduction is called a

PERFECT FLUID

Analysis of measurements at RHIC indicate a strongly interacting quark-gluon plasma

- called **sQGP** -

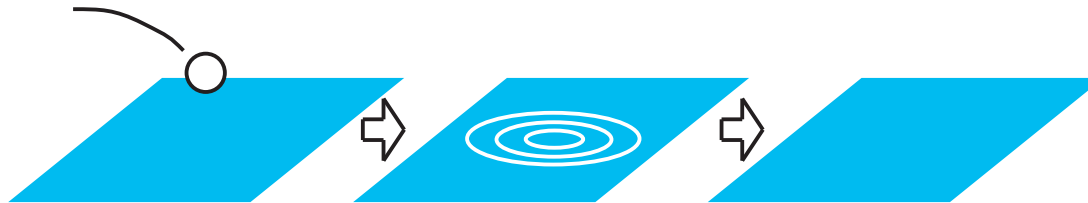
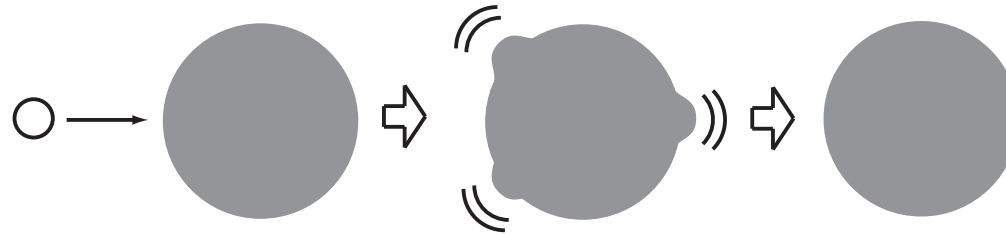
which behaves almost like a perfect fluid, with very low

viscosity Is it correct? It requires to calculate corrections in a systematic way!

remarkable approach: **black hole theory**

is now used to explain properties of colliding nuclei

dissipation: consequence of **black hole absorption**



dissipation: consequence of **viscosity**

[from M. Natsuume]

(damped) quasinormal modes - excited classical oscillations:
gravitational perturbation to a black hole ("ringing BH")
and to a hydrodynamic system ("ringing plasma")

hydrodynamics

energy momentum tensor: ϵ energy density and p pressure

$$T^{\mu\nu} = (\epsilon + p)u^\mu u^\nu + pg^{\mu\nu} + \Pi^{\mu\nu}$$

u^μ fluid velocity, $u_\mu u^\mu = -1$, and $\Pi^{\mu\nu}$ shear stress tensor

with

$$u_\mu \Pi^{\mu\nu} = 0 \quad (u_\mu T^{\mu\nu} = -\epsilon u^\nu),$$

momentum density is due to energy flow, and

$$\Pi^\mu{}_\mu = 0$$

required from conformal invariance

local conservation law (covariant derivative ∇_μ): $\nabla_\mu T^{\mu\nu} = 0$

(assume: no net charge in the system)

a few definitions

conformal fluid in $d = 4$ (curved) dimensions -

line element $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$

metric $g_{\mu\nu}$ with signature $(- + + +)$, $\mu, \nu..$ space-time indices 0, 1, 2, 3

(geometric) covariant derivative: ∇_μ

$$\nabla_\mu \Phi(x) = \partial_\mu \Phi(x) ,$$

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\rho}{}^\nu V^\rho \text{ etc., e.g. } \nabla_\mu g^{\nu\rho} = 0$$

Christoffel symbols:

$$\Gamma_{\mu\rho}{}^\nu = \frac{1}{2} g^{\nu\sigma} (\partial_\mu g_{\rho\sigma} + \partial_\rho g_{\mu\sigma} - \partial_\sigma g_{\mu\rho})$$

Riemann tensor: $[\nabla_\mu, \nabla_\nu] V^\rho = R_{\mu\nu\sigma}{}^\rho V^\sigma$

approximation to $\Pi^{\mu\nu}$

only retaining shear viscosity terms - NO bulk viscosity

- Navier - Stokes = first-order theory in gradients
with constitutive relation:

$$\Pi^{\mu\nu} = -2\eta \langle \nabla^\mu u^\nu \rangle \equiv -2\eta \sigma^{\mu\nu}$$

with the traceless and symmetric shear strain tensor

$$\sigma^{\mu\nu} = \sigma^{\nu\mu} \equiv \frac{1}{2} (\Delta^{\mu\rho} \nabla_\rho u^\nu + \Delta^{\nu\rho} \nabla_\rho u^\mu) - \frac{1}{3} \Delta^{\mu\nu} \Delta^{\rho\sigma} \nabla_\rho u_\sigma$$

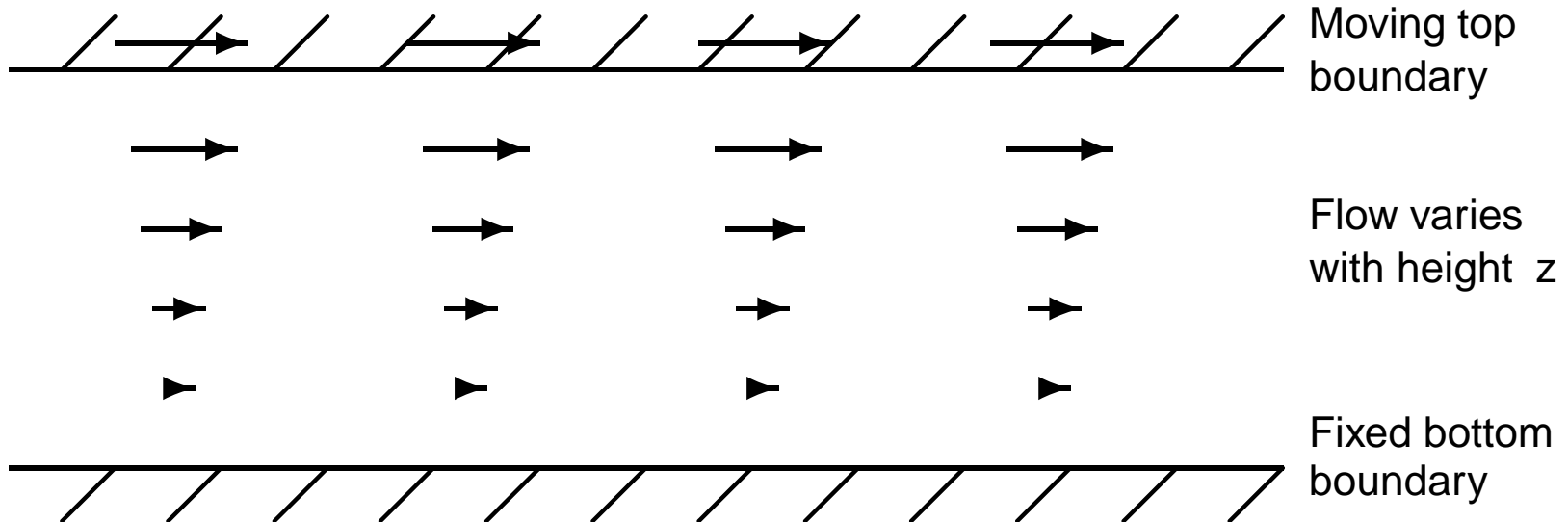
and transverse projection

$$\Delta^{\mu\nu} = g^{\mu\nu} + u^\mu u^\nu, \quad \Delta^{\mu\nu} u_\nu = 0$$

shear viscosity η

interacting field theories:

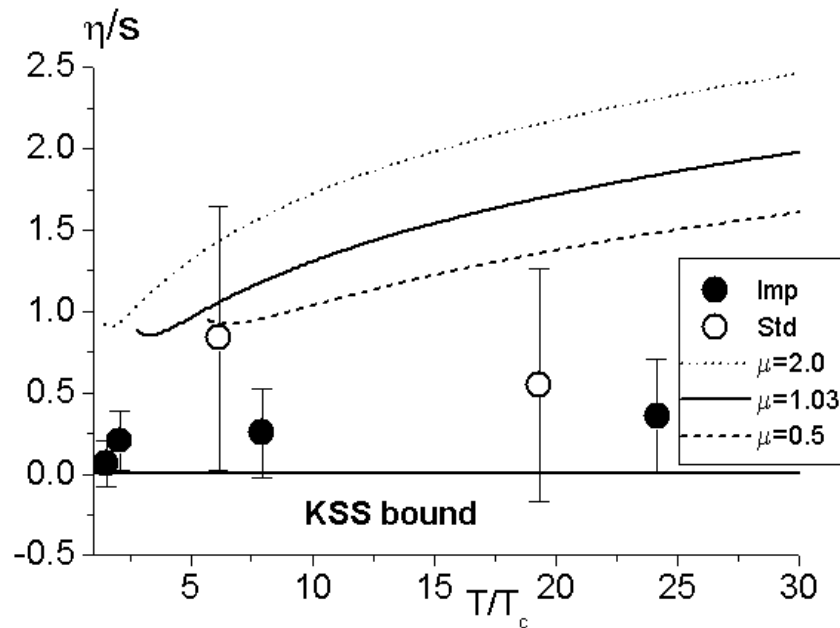
non-vanishing transport coefficients



e.g. shear flow in x - direction \rightarrow force per unit area/momentum transfer

$$\frac{F_x}{A} = -\Pi_{xz} = \eta \frac{\partial u_x}{\partial z}$$

viscosity/entropy density ratio



η/s from quenched QCD **lattice simulations** [Nakamura and Sakai 05],
 also $\eta/s = 0.134(33)$ at $T = 1.65T_c$ [Meyer 07]

compared with the **perturbative result** $\sim \frac{1}{g^4 \ln(\mu/gT)} \sim \mathcal{O}(1)$

for strong coupling $\lambda = g_{YM}^2 N_c$, $\lambda \rightarrow \infty$ [Kovtun, Son, Starinets 05]

$$\frac{\eta}{s} \geq \frac{1}{4\pi} \sim 0.08 : \text{AdS/CFT universal (?) lower bound}$$

Navier-Stokes

transverse projection of EoM $\Delta_{\alpha}^{\mu} \nabla_{\beta} T^{\alpha\beta} = 0 \rightarrow$
relativistic Navier-Stokes equation in first-order theory

$$(\epsilon + p)u^{\alpha} \nabla_{\alpha} u^{\mu} + \Delta^{\mu\alpha} \nabla_{\alpha} p + \Delta_{\alpha}^{\mu} \nabla_{\beta} [-2\eta < \nabla^{\alpha} u^{\beta} >] = 0$$

i.e. **parabolic equation:**

time derivative is of first-order ($u^{\alpha} \nabla_{\alpha} \rightarrow \partial/\partial t$)

while

space derivative is of second-order ($\nabla_j \sigma^{ij} \rightarrow \partial_j \partial_i u^j$)

**“relativistic first-order dissipative theory is highly pathological,
and therefore should be discarded
in favor of the second-order one”**

[Hiscock and Lindblom 1983-1985]

transverse mode

small linear perturbation and first-order \rightarrow

diffusion equation in shear channel:

$$\delta u_{\perp} = \sqrt{\frac{(\epsilon + p)}{4\pi\eta t}} \exp\left[-\frac{(\epsilon + p)x^2}{4\eta t}\right]$$

i.e. propagates outside the light-cone (Gaussian)
starting from $\delta u_{\perp}(x, t = 0) = \delta(x)$!!

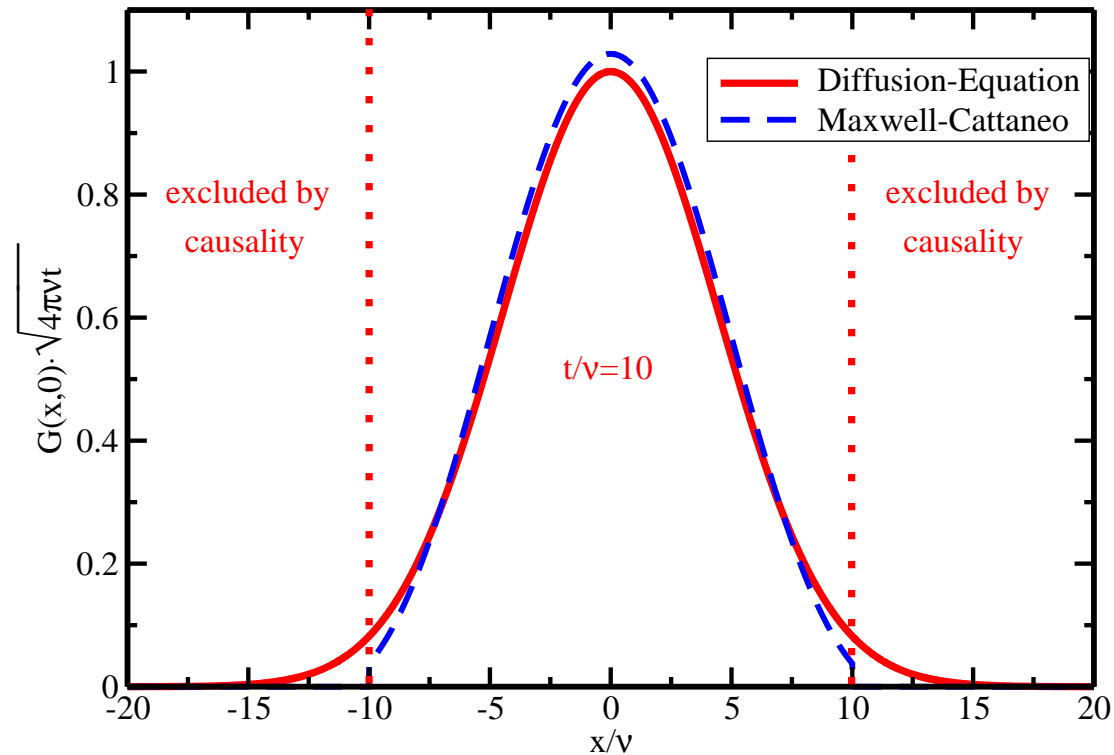
minimal modification \rightarrow second-order:

relaxation time $\tau_{\pi} > 0$ \rightarrow hyperbolic “telegraph” equation

$$\left[\tau_{\pi}\partial_t^2 + \partial_t - \frac{\eta}{(\epsilon + p)}\partial_x^2\right] \delta u_{\perp} = 0$$

diffusion vs. Cattaneo

Diffusion Eq. vs. Maxwell-Cattaneo



$$\nu = \eta / (\epsilon + p) = \text{const}; \quad \tau_\pi = \nu$$

Green's function for diffusion and Maxwell-Cattaneo law

[Romatschke (2009)]

a systematic result

all second-order terms classified by conformal symmetry,
by space-time dependent Weyl transformation:

$$g^{\mu\nu} = e^{-2\phi(x)} \tilde{g}^{\mu\nu}, \quad T^{\mu\nu} = e^{-6\phi} \tilde{T}^{\mu\nu}, \quad T^\mu{}_\mu = 0, \quad \dots$$

constitutive relation (d= 4) with five new terms:

$$\begin{aligned} \Pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} + 2\eta\tau_\pi u^\lambda \mathcal{D}_\lambda \sigma^{\mu\nu} \\ & + 4\lambda_1 \sigma^{<\mu}{}_\lambda \sigma^{\nu>\lambda} + 2\lambda_2 \sigma^{<\mu}{}_\lambda \Omega^{\nu>\lambda} + \lambda_3 \Omega^{<\mu}{}_\lambda \Omega^{\nu>\lambda} \\ & + 2\kappa u_\alpha C^{\alpha\mu\nu\beta} u_\beta \end{aligned}$$

$C^{\alpha\beta\gamma\delta}$... Weyl tensor, $\Omega^{\alpha\beta}$... antisymmetric vorticity tensor, \mathcal{D}_λ ... Weyl derivative

[Baier, Romatschke, Son, Starinets, Stephanov 07]

[Bhattacharyya, Hubeny, Minwalla, Rangamani 07, Loganayagam 08]

relaxation phenomena/**strong coupling limit**

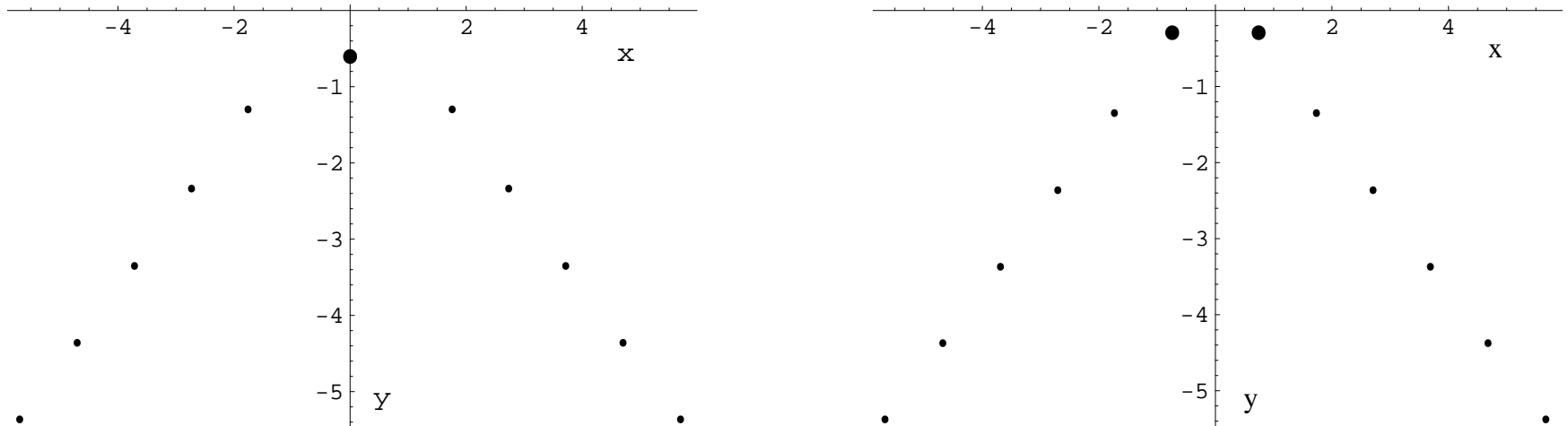
hydrodynamic transport coefficients by gauge/gravity duality:
compare with quite involved AdS/CFT-gravity calculations
at Hawking temperature $T = T_H$ and for momentum
 $\omega, k \ll T_H$ i.e. long wave-length limit:

e.g. from sound channel dispersion up to $O(k^3)$, etc :

$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T} = 2\lambda_1,$$

$$\lambda_2 = -\frac{\ln 2}{2\pi T}\eta, \quad \lambda_3 = 0$$

hydrodynamic modes



quasinormal spectrum of black brane gravitational fluctuations in the shear and sound channels at fixed spatial momentum as a function of complex frequency

[Kovtun and Starinets (2005)]

hydrodynamic poles are marked by full dots, e.g. shear

$$\omega = -i \frac{\eta}{T_s} q^2 + O(q^3) = -i \frac{1}{4\pi T} q^2 + O(q^3)$$

approach the real frequency axis for $q \rightarrow 0$

[Kovtun, Son and Starinets (2005)]

quasinormal mode (QNM)

Wikipedia:

Quasinormal modes are the modes of energy dissipation of a perturbed object or field (example: perturbation of a wine glass with a knife - the glass begins to ring, but not forever!)

Quasinormal ringing is approximated by the amplitude

$$\psi(t) \approx e^{-\omega_I t} \cos(\omega_R t) \approx e^{-i\omega t}$$

complex quasinormal frequency, $Im\omega = -\omega_I < 0$

in gravity: characteristic excited classical oscillations of (AdS-) black holes (BH) and black branes responding to generic perturbations - approach to thermal equilibrium:

“ringing BH”

transport coefficients

kinetic QCD theory (LO weak coupling: $\alpha_s < 0.7$)

$$\frac{\eta/s}{\tau_\pi T} \simeq 0.17 - 0.2, \quad \lambda_1 \simeq (5.2 - 4.1) \frac{\eta^2}{sT}, \quad \lambda_2 = -2\tau_\pi \eta, \quad \lambda_3 = \kappa = 0$$

[York and Moore 08]

finite 't Hooft coupling $\lambda \equiv g_{YM}^2 N_c$, $\lambda \gg 1$ corrections to coefficients by gauge/gravity duality, e.g.:

$$\frac{\eta/s}{\tau_\pi T} = 0.383 (1 - 3.52 \lambda^{-3/2} + \dots), \quad \lambda_1 = \frac{2\eta^2}{sT},$$

$$\lambda_2 = -0.530\tau_\pi \eta, \quad \kappa = \frac{\eta}{\pi T} \left(1 - \frac{145\zeta(3)}{8} \lambda^{-3/2} + \dots\right)$$

[Buchel and Paulos 08]

non-conformal transport coefficients

The shear viscosity η and the second order coefficients τ_π (“shear” relaxation time), κ , λ_1 , λ_2 , λ_3 are the only ones defined in conformal fluids, as the one of $\mathcal{N} = 4$ SYM. All the others coefficients, i.e. the bulk viscosity ζ and the second order coefficients κ^* , τ_π^* , λ_4 , τ_Π (“bulk” relaxation time), ξ_1 , ξ_2 , ξ_3 , ξ_4 , ξ_5 , ξ_6 , are only defined in non-conformal plasmas.

Defining:

$$\delta \equiv (1 - 3c_s^2)$$

transport coefficients at leading order in the deformation parameter δ :

[Bigazzi and Cotrone (2010)]

$\frac{\eta}{s}$	$\frac{1}{4\pi}$	$T\tau_\pi$	$\frac{2-\log 2}{2\pi} + \frac{3(16-\pi^2)}{64\pi}\delta$	$\frac{T\kappa}{s}$	$\frac{1}{4\pi^2} \left(1 - \frac{3}{4}\delta\right)$
$\frac{T\lambda_1}{s}$	$\frac{1}{8\pi^2} \left(1 + \frac{3}{4}\delta\right)$	$\frac{T\lambda_2}{s}$	$-\frac{1}{4\pi^2} \left(\log 2 + \frac{3\pi^2}{32}\delta\right)$	$\frac{T\lambda_3}{s}$	0
$\frac{T\kappa^*}{s}$	$-\frac{3}{8\pi^2}\delta$	$T\tau_\pi^*$	$-\frac{2-\log 2}{2\pi}\delta$	$\frac{T\lambda_4}{s}$	0
$\frac{\zeta}{\eta}$	$\frac{2}{3}\delta$	$T\tau_\Pi$	$\frac{2-\log 2}{2\pi}$	$\frac{T\xi_1}{s}$	$\frac{1}{24\pi^2}\delta$
$\frac{T\xi_2}{s}$	$\frac{2-\log 2}{36\pi^2}\delta$	$\frac{T\xi_3}{s}$	0	$\frac{T\xi_4}{s}$	0
$\frac{T\xi_5}{s}$	$\frac{1}{12\pi^2}\delta$	$\frac{T\xi_6}{s}$	$\frac{1}{4\pi^2}\delta$		

$\frac{\eta}{s}$	$\frac{1}{4\pi}$	$T\tau_{\pi}$	0.222	$\frac{T\kappa}{s}$	0.022
$\frac{T\lambda_1}{s}$	0.014	$\frac{T\lambda_2}{s}$	-0.021	$\frac{T\lambda_3}{s}$	0
$\frac{T\kappa^*}{s}$	-0.006	$T\tau_{\pi}^*$	-0.031	$\frac{T\lambda_4}{s}$	0
$\frac{\zeta}{\eta}$	0.101	$T\tau_{\Pi}$	0.208	$\frac{T\xi_1}{s}$	0.001
$\frac{T\xi_2}{s}$	0.001	$\frac{T\xi_3}{s}$	0	$\frac{T\xi_4}{s}$	0
$\frac{T\xi_5}{s}$	0.001	$\frac{T\xi_6}{s}$	0.004		

transport coefficients at $T \sim 1.5T_c$ and $c_s^2 \sim 0.283$

conformal hydrodynamics

Weyl transformation

scaling transformation:

$$g_{\mu\nu} = e^{2\phi} \tilde{g}_{\mu\nu}; \quad g^{\mu\nu} = e^{-2\phi} \tilde{g}^{\mu\nu}, \quad \phi = \phi(x)$$

leads to:

$$\Gamma_{\lambda\mu}{}^{\nu} = \tilde{\Gamma}_{\lambda\mu}{}^{\nu} + \delta_{\lambda}^{\nu} \partial_{\mu} \phi + \delta_{\mu}^{\nu} \partial_{\lambda} \phi - \tilde{g}_{\lambda\mu} \tilde{g}^{\nu\sigma} \partial_{\sigma} \phi$$

$$u^{\mu} = e^{-\phi} \tilde{u}^{\mu}, \quad \Delta^{\mu\nu} = g^{\mu\nu} + u^{\mu} u^{\nu} = e^{-2\phi} \tilde{\Delta}^{\mu\nu}$$

usefull quantities (d dimensions):

$$'divergence' : \vartheta \equiv \nabla_{\mu} u^{\mu} = e^{-\phi} \left[\tilde{\vartheta} + (d-1) \tilde{u}^{\sigma} \partial_{\sigma} \phi \right],$$

$$'acceleration' : a^{\nu} \equiv u^{\mu} \nabla_{\mu} u^{\nu} = e^{-2\phi} \left[\tilde{a}^{\nu} + \tilde{\Delta}^{\nu\sigma} \partial_{\sigma} \phi \right],$$

$$'gauge field' : \mathcal{A}_{\nu} \equiv a_{\nu} - \frac{\vartheta}{d-1} u_{\nu} = \tilde{\mathcal{A}}_{\nu} + \partial_{\nu} \phi$$

Weyl derivative

cf. (abelian) gauge transform and derivative:

$$\begin{aligned}\tilde{\psi} &= e^{i\Lambda(x)}\psi, \quad \tilde{A}^\mu = A^\mu - \partial^\mu\Lambda, \\ \tilde{D}^\mu\tilde{\psi} &= e^{i\Lambda}(\partial^\mu + iA^\mu)\psi = e^{i\Lambda}D^\mu\psi\end{aligned}$$

elegant way: define covariant Weyl derivative [Loganayagam 07]

$$Q_{\nu\dots}^{\mu\dots} = e^{-w\phi}\tilde{Q}_{\nu\dots}^{\mu\dots}, \quad \mathcal{D}_\lambda Q_{\nu\dots}^{\mu\dots} = e^{-w\phi}\tilde{\mathcal{D}}_\lambda\tilde{Q}_{\nu\dots}^{\mu\dots}, \quad w..integer$$

where Weyl derivative - linear in \mathcal{A}_λ -

$$\begin{aligned}\mathcal{D}_\lambda Q_{\nu\dots}^{\mu\dots} &\equiv \nabla_\lambda Q_{\nu\dots}^{\mu\dots} + w\mathcal{A}_\lambda Q_{\nu\dots}^{\mu\dots} \\ &+ [g_{\lambda\alpha}\mathcal{A}^\mu - \delta_\lambda^\mu\mathcal{A}_\alpha - \delta_\alpha^\mu\mathcal{A}_\lambda] Q_{\nu\dots}^{\alpha\dots} + \dots \\ &- [g_{\lambda\nu}\mathcal{A}^\alpha - \delta_\lambda^\alpha\mathcal{A}_\nu - \delta_\nu^\alpha\mathcal{A}_\lambda] Q_{\alpha\dots}^{\mu\dots} - \dots\end{aligned}$$

examples

$$\mathcal{D}_\lambda g_{\mu\nu} = 0, \quad \mathcal{D}_\mu u^\mu = 0, \quad u^\mu \mathcal{D}_\mu u^\nu = 0$$

and

$$\mathcal{D}_\mu u^\nu = \nabla_\mu u^\nu + u_\mu a^\nu - \frac{\vartheta}{d-1} \Delta_\mu{}^\nu = \sigma_\mu{}^\nu + \Omega_\mu{}^\nu = e^{-\phi} \tilde{\mathcal{D}}_\mu \tilde{u}^\nu$$

$$\begin{aligned} \sigma^{\mu\nu} &\equiv \frac{1}{2} (\Delta^{\mu\lambda} \nabla_\lambda u^\nu + \Delta^{\nu\lambda} \nabla_\lambda u^\mu) - \frac{1}{d-1} \vartheta \Delta^{\mu\nu} = \frac{1}{2} (\mathcal{D}^\mu u^\nu + \mathcal{D}^\nu u^\mu) \\ &= e^{-3\phi} \tilde{\sigma}^{\mu\nu} \end{aligned}$$

antisymmetric vorticity tensor :

$$\Omega^{\mu\nu} \equiv \frac{1}{2} (\Delta^{\mu\lambda} \nabla_\lambda u^\nu - \Delta^{\nu\lambda} \nabla_\lambda u^\mu) = \frac{1}{2} (\mathcal{D}^\mu u^\nu - \mathcal{D}^\nu u^\mu) = e^{-3\phi} \tilde{\Omega}^{\mu\nu}$$

- partition function $Z = \tilde{Z}$ is conformal invariant
($d = 4$ dimensions)

$$T^{\mu\nu} \propto \frac{1}{\sqrt{-g}} \frac{\delta \ln Z}{\delta g_{\mu\nu}} \rightarrow$$

$$T^{\mu\nu} = e^{-6\phi(x)} \tilde{T}^{\mu\nu} \quad \text{with } g_{\mu\nu} = e^{2\phi} \tilde{g}_{\mu\nu}$$



$$\text{conservation : } \nabla_{\mu} T^{\mu\nu} = 0 \rightarrow \tilde{\nabla}_{\mu} \tilde{T}^{\mu\nu} = 0 \rightarrow$$

$$\tilde{\nabla}_{\mu} \tilde{T}^{\mu\nu} = e^{6\phi} [\nabla_{\mu} T^{\mu\nu} + T_{\mu}^{\mu} \partial^{\nu} \phi] \rightarrow T_{\mu}^{\mu} = 0$$

- only scale is temperature T :

$$T_{\text{perfect}}^{\mu\nu} = (g^{\mu\nu} + 4u^{\mu}u^{\nu})p, \quad \epsilon = 3p$$

$$\rightarrow p = e^{-4\phi} \tilde{p}, \quad p \propto T^4, \quad T = e^{-\phi} \tilde{T}, \quad \mathcal{D}_{\mu} T = e^{-\phi} \tilde{\mathcal{D}}_{\mu} \tilde{T}$$

invariant Weyl tensor

$$\begin{aligned}
 (d-2)C_{\mu\alpha\nu\beta}u^\alpha u^\beta = & \\
 & \Delta^{\mu\lambda}\Delta^{\nu\sigma}R_{\lambda\sigma} + (d-2)\Delta^{\mu\lambda}\Delta^{\nu\sigma}R_{\lambda\alpha\sigma\beta}u^\alpha u^\beta \\
 & - \frac{\Delta^{\mu\nu}}{d-1}(\Delta^{\lambda\sigma}R_{\lambda\sigma} + (d-2)\Delta^{\lambda\sigma}R_{\lambda\alpha\sigma\beta}u^\alpha u^\beta)
 \end{aligned}$$

$R_{\mu\nu}\dots$ Ricci tensor

symmetry properties (same as for the Riemann tensor):

$$C_{\mu\nu\lambda\sigma} = -C_{\nu\mu\lambda\sigma} = -C_{\mu\nu\sigma\lambda} = C_{\lambda\sigma\mu\nu}$$

$$\text{and } C_{\mu\alpha\lambda}{}^\alpha = 0$$

$$C_{\alpha\mu\nu}{}^\beta = \tilde{C}_{\alpha\mu\nu}{}^\beta \implies u_\alpha C^{\alpha\mu\nu\beta} u_\beta = e^{-4\phi(x)} \tilde{u}_\alpha \tilde{C}^{\alpha\mu\nu\beta} \tilde{u}_\beta$$

$$u_\mu u_\alpha C^{\alpha\mu\nu\beta} u_\beta = 0$$

$\Pi^{\mu\nu}$ has in $d = 4$ dimensions 5 independent tensor structures
first-order

$$\Pi^{\mu\nu} = -2\eta \sigma^{\mu\nu}$$

$$\Pi^{\mu\nu} = e^{-6\phi} \tilde{\Pi}^{\mu\nu}, \quad \sigma^{\mu\nu} = e^{-3\phi} \tilde{\sigma}^{\mu\nu} \rightarrow$$

$$\eta = e^{-3\phi} \tilde{\eta}, \quad \text{i.e. } \eta \propto T^3$$

e.g. $\eta = (\pi T)^3$, $p = (\pi T)^4$,

$$s = \frac{\partial p}{\partial T} = 4\pi(\pi T)^3 \rightarrow \eta/s = 1/4\pi \quad \text{i.e. only a number!}$$

second-order

$$u^\lambda \mathcal{D}_\lambda \sigma^{\mu\nu} = e^{-4\phi} \tilde{u}^\lambda \tilde{\mathcal{D}}_\lambda \tilde{\sigma}^{\mu\nu}, \quad \text{etc.}$$

$$\eta \tau_\pi = e^{-2\phi} \tilde{\eta} \tilde{\tau}_\pi, \quad \tau_\pi \propto \frac{1}{T}, \quad \kappa \propto T^2, \quad \lambda_{1,2,3} \propto T^2$$

conformal $\Pi^{\mu\nu}$

all 5 second-order terms classified by conformal symmetry

constitutive relation (d= 4):

$$\begin{aligned}\Pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} + 2\eta\tau_\pi u^\lambda \mathcal{D}_\lambda \sigma^{\mu\nu} \\ & + 4\lambda_1 \sigma^{\langle\mu}{}_\lambda \sigma^{\nu\rangle\lambda} + 2\lambda_2 \sigma^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} + \lambda_3 \Omega^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} \\ & + 2\kappa u_\alpha C^{\alpha\mu\nu\beta} u_\beta\end{aligned}$$

$$\eta \propto T^3, \quad \tau_\pi \propto \frac{1}{T}, \quad \kappa \propto \lambda_{1,2,3} \propto T^2$$

consistent evaluation of transport coefficients in strong coupling limit via

AdS/CFT correspondence

duality \sim **equality**

is between

Quantum Field Theory (a special one) in $d = 4$

and

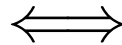
Classical Gravity in $d = 5$ (for $N_c \gg 1$, $g_{YM}^2 N_c \gg 1$)

(at finite temperature)

AdS/CFT [Maldacena 98, Gubser et al. 98, Witten 98]

strongly coupled quantized conformal gauge theory
in $d = 4$ dimensions ($\mathcal{N} = 4$ SYM with $8N_c^2$ (1 gauge and 6
scalar) bosons and $(4N_c^2)$ Weyl fermions)

[[NOT QCD !]]



weakly coupled classical supergravity (type IIB)
in $d = 10$ dimensions (on $AdS_5 \otimes S^5$)

via **holographic property**: radial coordinate $r_0 \leq r < \infty$ with
gauge theory on the boundary at ∞

AdS/CFT cont.

crucial limit of interest:

't Hooft coupling $\lambda = g_{YM}^2 N_c$ is large, $N_c \rightarrow \infty$, $g_{YM}^2 \ll 1$

i.e. string coupling $g_s = \frac{g_{YM}^2}{4\pi} \ll 1$

– NO LOOPS

and

small curvature $\frac{l_s^4}{L^4} = \frac{1}{\lambda} \ll 1$ – RADIUS L of CURVATURE is

LARGE compared to the STRING SCALE $l_s = \sqrt{\alpha'}$

– CLASSICAL GRAVITY

AdS_5 metric at finite T

generalization of the Einstein-Hilbert-Maxwell equation
- strong analogy to Reissner-Nordström black hole

near extremal black $D3$ -brane metric with horizon $r = r_0$ for
 $r_0 < r \ll L$, i.e. factorized metric for $AdS_5 \otimes S^5$ for:

$$ds_{(5)}^2 = \frac{r^2}{L^2} (-f(r) dt^2 + d\vec{x}^2) + \frac{L^2}{r^2 f(r)} dr^2$$

with $f(r) = 1 - \frac{r_0^4}{r^4}$

from AdS_5 action - negative cosmological constant $\Lambda = -\frac{6}{L^2}$:

$$S = \frac{1}{16\pi G_{(5)}} \int d^5x \sqrt{-g} \left[R + \frac{12}{L^2} \right],$$

with Newton constant $G_{(5)} = \frac{\pi L^3}{2N_c^2}$

master formula for AdS/CFT

schematically in terms of coinciding partition functions:

$$\int e^{iS_{4d}^{gauge} + i\Phi_0 O} = \int e^{iS_{5d}[\Phi]} \simeq e^{iS^{classical}[\Phi_0]}$$

S_{5d} is computed with non-trivial boundary condition
(holography)

$$\Phi(t, \vec{x}, r) \stackrel{r \rightarrow \infty}{=} \Phi_0(t, \vec{x})$$

\implies

quantum correlation = classical two-point function

$$G(x, y) = -i \langle T O(x) O(y) \rangle = - \frac{\delta^2 S^{classical}}{\delta\Phi_0(x) \delta\Phi_0(y)} \Big|_{r=\infty}$$

gauge: $O = T_{\mu\nu}$.. energy-momentum tensor

gravity: $\phi = g_{\mu\nu}$.. graviton

correlators from gravity, $d = 4$

response of the fluid to small metric perturbations

e.g.: $g_{xy} = h_{xy}(t, z, r) = -h^{xy}(t, z, r) \neq 0$, $u^0 = 1$, $T = \text{const}$

it leads from Christoffel symbols in the covariant derivatives to linear approximation:

$$\sigma^{xy} = \sigma_{xy} \approx \frac{1}{2}(\Gamma_{x0}^y + \Gamma_{y0}^x) \approx \frac{1}{2}\partial_t h_{xy}$$

and for the energy-momentum tensor

$$\delta T^{xy} \approx -p h_{xy} - \eta \partial_t h_{xy} + \eta \tau_\pi \partial_t^2 h_{xy} - \frac{\kappa}{2} [\partial_t^2 h_{xy} + \partial_z^2 h_{xy}]$$

Fourier transform $h(t, z) = \exp(-i\omega t + ikz) h(\omega, k)$, etc.

$$\delta T^{xy}(\omega, k) = -G_R^{xy,xy}(\omega, k) h_{xy}(\omega, k)$$

via **linear response**

correlators, cont.

retarded Green function (in momentum space):

$$G_R^{xy,xy}(\omega, k) = p - i\eta\omega + \eta\tau_\pi\omega^2 - \frac{\kappa}{2}(\omega^2 + k^2) + \dots$$

i.e. **Kubo formula**

$$\eta = \lim_{\omega \rightarrow 0} \frac{1}{\omega} iG_R^{xy,xy}(\omega, \vec{0})|_{\omega=0}$$

$$= \lim_{\omega \rightarrow 0} \frac{1}{\omega} \int dt d^3x \exp(i\omega t) \Theta(t) \langle [T_{xy}(t, \vec{x}), T_{xy}(0, \vec{0})] \rangle |_{\omega=0}$$

classical action $S^{classical}[\Phi_0]$ from the solution Φ_0 for $h_{xy}(\omega, k, r)$ at the boundary $r \rightarrow \infty$ and keeping only the surface contribution:

\Rightarrow

$$G_R^{xy,xy}(\omega, k) = \frac{\pi N_c^2 T^3}{8} \left[\pi T - i\omega + \frac{1 - 2 \ln 2}{2\pi T} \omega^2 - \frac{1}{2\pi T} k^2 \pm \dots \right],$$

$$p = \frac{\pi^2 N_c^2 T^4}{8}, \quad \eta = \frac{\pi N_c^2 T^3}{8}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T}$$

sound mode

consistency requirement from sound channel:

$$\omega_{1,2} = \pm c_s k - i\Gamma k^2 \pm \frac{\Gamma}{c_s} (c_s^2 \tau_\pi - \frac{\Gamma}{2}) k^3 + \mathcal{O}(k^4)$$

with $\Gamma = \frac{2}{3} \frac{\eta}{\epsilon + p}$ and

gravity perturbation $g_{tz} = h_{tz} \ll 1$ [i.e. poles in $G_R^{tz,tz}(\omega, k)$]

$$\omega_{1,2} = \pm \frac{1}{\sqrt{3}} k - i \frac{1}{6\pi T} k^2 \pm \frac{3 - 2 \ln 2}{24\sqrt{3}\pi^2 T^2} k^3$$

gives sound velocity $c_s = \frac{1}{\sqrt{3}}$ and consistent values for

$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}$$

RHIC physics

hydro + RHIC

heavy-ion collisions require **beyond hydrodynamics**:

- hydrodynamics = differential equations
initial conditions !
- initial = equilibration time
- distribution of energy density (Glauber? CGC ?)
- QCD equation of state
- hadronisation prescription (Cooper-Frye?)
-

main results [Luzum and Romatschke 08]

- viscous hydrodynamics simulation give a good description of RHIC data for

$$\frac{\eta}{s} = 0.1 \pm 0.1(\text{theory}) \pm 0.08(\text{experiment})$$

- modest estimate:

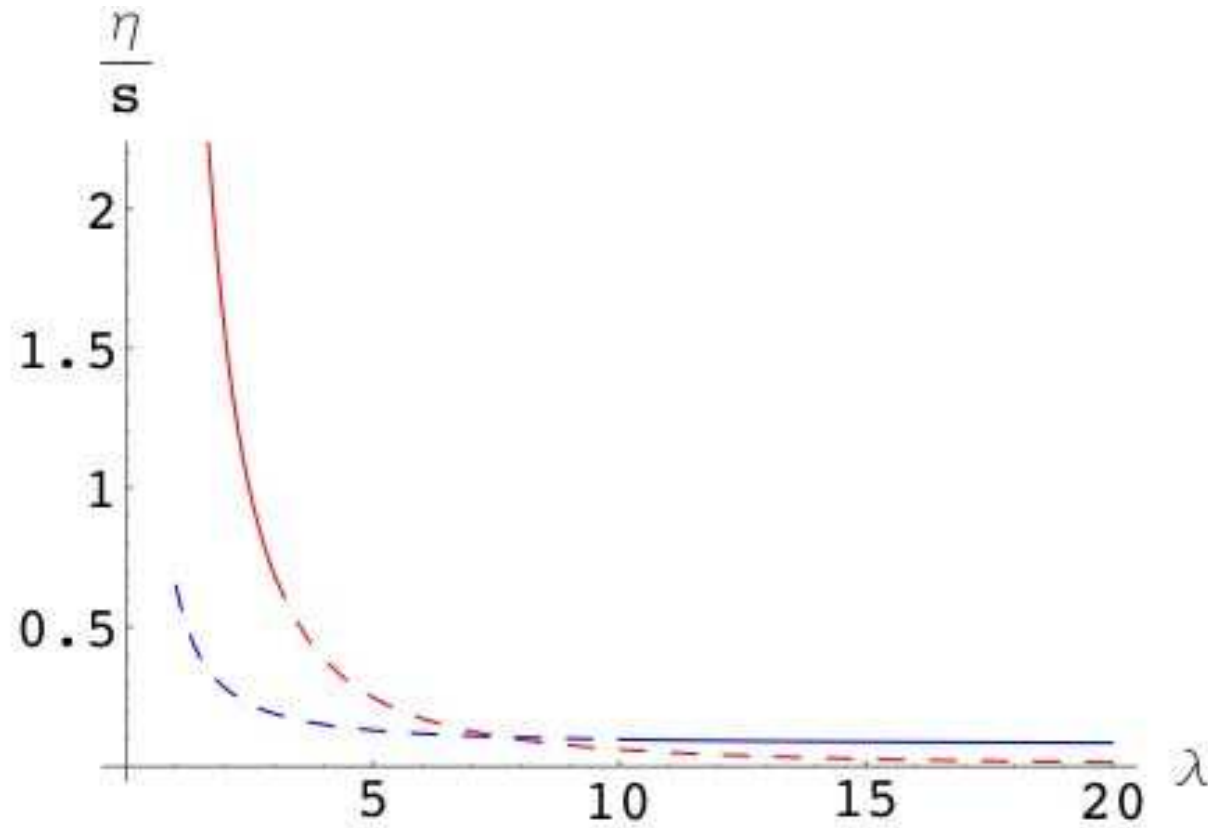
$$\frac{\eta}{s} < 0.5$$

- weak dependence on the values of the second-order parameters $\tau_\pi, \lambda_1, \dots$

- early thermalisation time is questioned, but

$$\tau_0 < 2 \text{ fm}$$

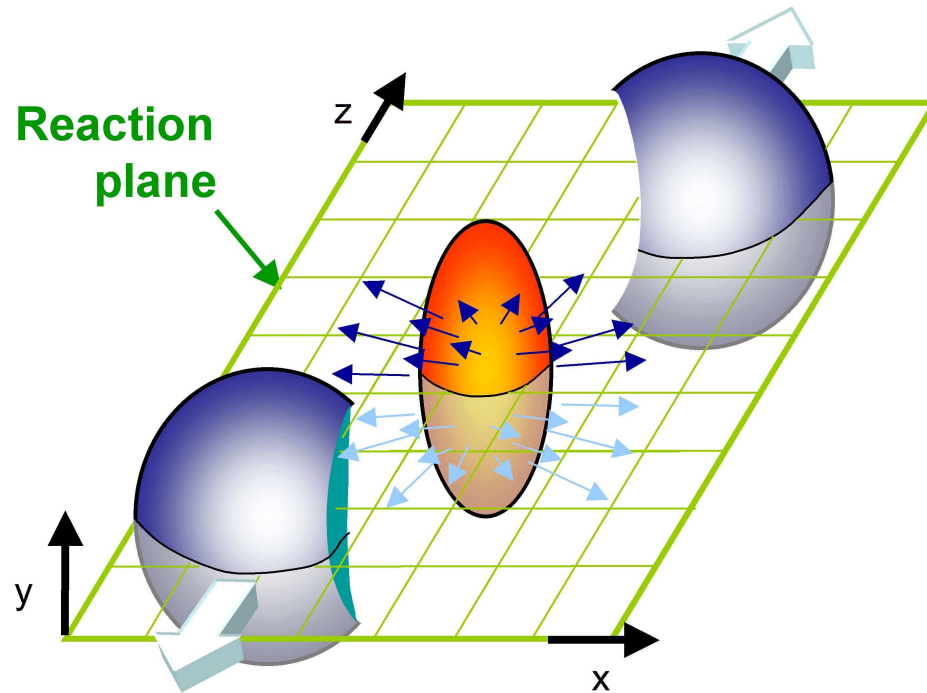
$$\eta/s$$



behaviour of η/s as a function of the 't Hooft coupling $\lambda = gYM^2N_c$

$\frac{\eta}{s} \approx 0.5$: border-value between sQGP and pQGP

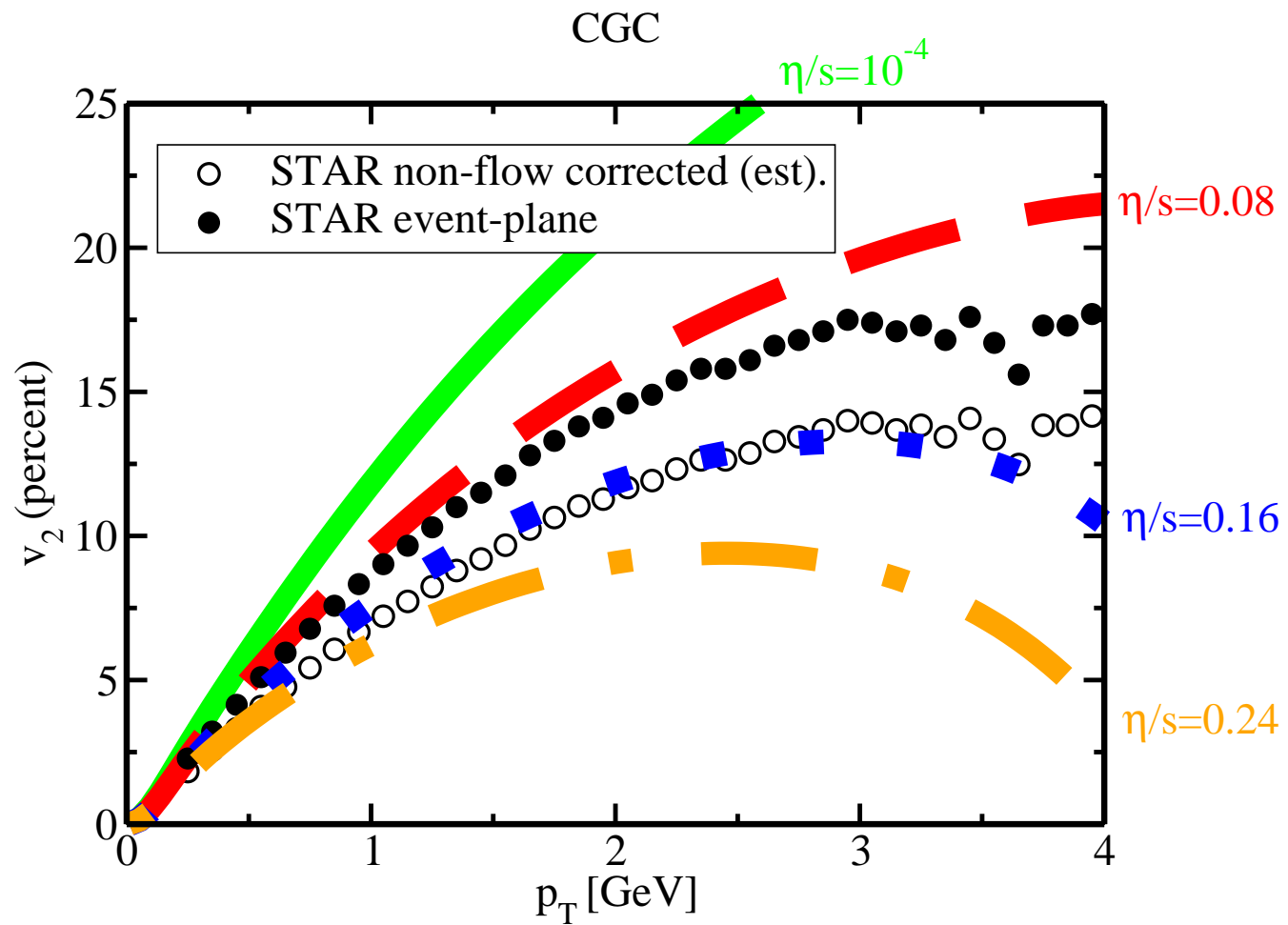
as important example



spatial asymmetry

$$\frac{dN}{dy dp_{\perp} d\phi} = \left\langle \frac{dN}{dy dp_{\perp} d\phi} \right\rangle_{\phi} (1 + 2v_2(p_{\perp}) \cos(2\phi) + \dots)$$

elliptic flow: $v_2(p_{\perp})$



elliptic flow: $\eta/s > 0$ reduces v_2

[Luzum and Romatschke 08]

EXTRAS

some classical gravity

Einstein-Hilbert with nonzero cosmological constant
($d = 5$ dimensions)



$$R_{\mu\nu} - \frac{R}{2}g_{\mu\nu} = -\Lambda_{(5)}g_{\mu\nu} , \quad \mu, \nu = 0, 1, \dots, 4$$

• ansatz:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + \dots , \quad g_{tt} = -\frac{1}{g_{rr}} = -\frac{r^2}{L^2}f(r)$$



$$R\left(1 - \frac{5}{2}\right) = -5\Lambda_{(5)}, \quad R_{\mu\nu} = \frac{2\Lambda_{(5)}}{3}g_{\mu\nu} = -\frac{4}{L^2}g_{\mu\nu}$$

solution

$$R_t^t = -\frac{1}{2} \frac{[r^5 f' + 2r^4 f]'}{r^3 L^2} = -\frac{4}{L^2}$$

$$r f' + 2f = 2 + \frac{const}{r^4}, \quad const = 2r_0^4$$

$$f(r) = 1 - \left(\frac{r_0}{r}\right)^4$$

i.e.

$$ds^2 = -\frac{r^2}{L^2} f(r) dt^2 + \frac{L^2}{r^2 f(r)} dr^2 + \dots$$

Hawking temperature

black hole metric:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + \dots, \quad g_{tt} = -\frac{1}{g_{rr}} = -\left[1 - \left(\frac{r_0}{r}\right)^{(d-3)}\right]$$

expand around the horizon $r \approx r_0$:

$$g_{tt} \approx -\frac{d-3}{r_0}(r - r_0) = -\gamma_t(r - r_0), \quad g_{rr} \approx \frac{\gamma_r}{r - r_0}, \quad \gamma_r = \frac{r_0}{d-3}$$

claim:

$$T_H = \frac{1}{4\pi} \sqrt{\frac{\gamma_t}{\gamma_r}} = \frac{d-3}{4\pi r_0}$$

and AdS_5 : $\gamma_t = 1/\gamma_r = 4r_0/R^2$

$$T_H = \frac{r_0}{\pi R^2}$$

proof: with $\rho^2 = \frac{4r_0(r-r_0)}{d-3}$

$$ds^2 = \rho^2 \left(\frac{d-3}{2r_0} \right)^2 d\tau^2 + d\rho^2 + \dots, \quad \text{euclidean } d\tau^2 = -dt^2$$

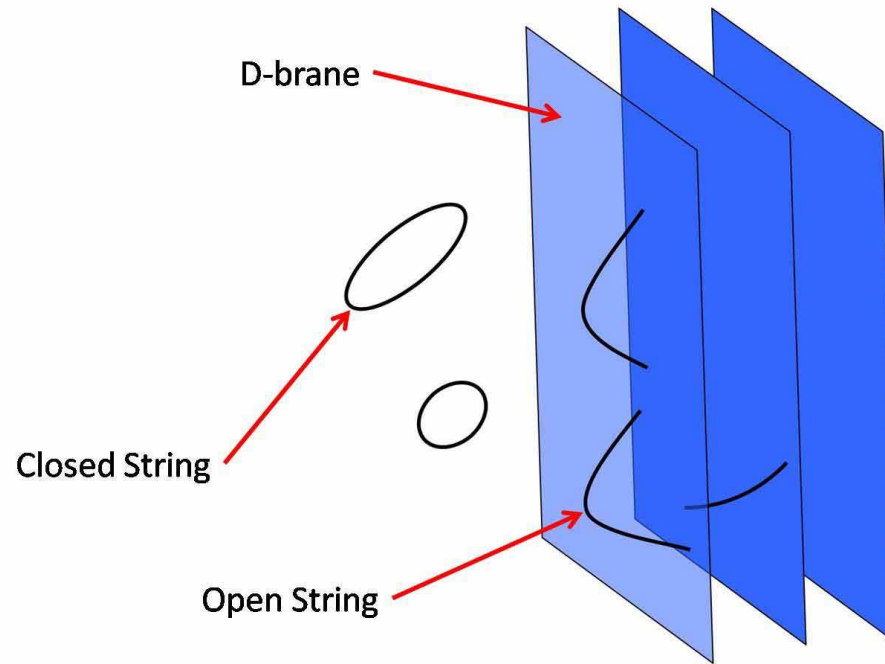
→

$$ds^2 = \rho^2 d\phi^2 + d\rho^2, \quad \phi = \frac{d-3}{2r_0} \tau$$

requirement: periodicity $\phi \rightarrow \phi + 2\pi$, i.e. **NO** conical singularity

$$2\pi = \frac{(d-3)}{2r_0} \beta, \quad \beta = \frac{1}{T_H}$$

D -branes



[from Myers and Vazquez 08]

dynamical walls on which strings can end:

theory of open strings living on $N_c - D3$ -branes ($\mathcal{N} = 4$

SYM, $d = 4$) \iff

gravity theory of fields living in the space curved by the
branes (AdS_5 , $d = 5$)

symmetries ($T = 0$)

duality maps the operators of QFT to the boundary values of the SUGRA fields

● QFT ($\mathcal{N} = 4, d = 4$ CFT)

$$S = -\frac{1}{g_{YM}^2} \int d^4x \left[\frac{1}{4} F_{\mu\nu}^a F^{a\mu\nu} + \frac{1}{2} D_\mu \Phi_i^a D^\mu \Phi_i^a + \dots \right]$$

.... massless Weyl fermions + interactions , $i = 1, \dots, 6, a = 1, \dots, N_c^2$

has Poincaré and conformal symmetry $SO(2, 4)$:

$x^\mu \rightarrow \lambda x^\mu$, $\Phi_i^a(x) \rightarrow \lambda \Phi_i^a(\lambda x)$, $F_{\mu\nu}(x) \rightarrow \lambda^2 F_{\mu\nu}(\lambda x)$ **etc,**

$\beta(\mu) = \mathbf{0}$, and $SU(\mathcal{N} = 4)$ symmetry and $SU(N_c)$

● $AdS_5 \otimes S^5$:

$$SO(2, 4) \otimes SO(6) = SO(2, 4) \otimes SU(4)$$

besides $SU(N_c)$ due to N_c parallel $D - 3$ branes

Bekenstein- Hawking entropy

thermodynamics: $dS = \frac{dE}{T}$

• Schwarzschild BH ($G = G_{(4)}, T_H = \frac{1}{4\pi r_0}$)

$$E = M = \frac{r_0}{2G}, \quad dE = \frac{dr_0}{2G}, \quad dS = \frac{4\pi}{2G} r_0 dr_0,$$

entropy

$$S = \frac{\pi r_0^2}{G} = \frac{A}{4G}$$

a universal relation with $A = 4\pi r_0^2$... area of the horizon

Bekenstein-Hawking entropy, cont.

● AdS_5 : $T_H = \frac{r_0}{\pi L^2}$, $S = \frac{A_{(3)}}{4G_{(5)}}$

$$A_{(3)}(r = r_0) = \left(\frac{r_0}{L}\right)^3 V_3 = \pi^3 T_H^3 L^3 V_3 , \quad G_{(5)} = \frac{\pi L^3}{2N_c^2}$$

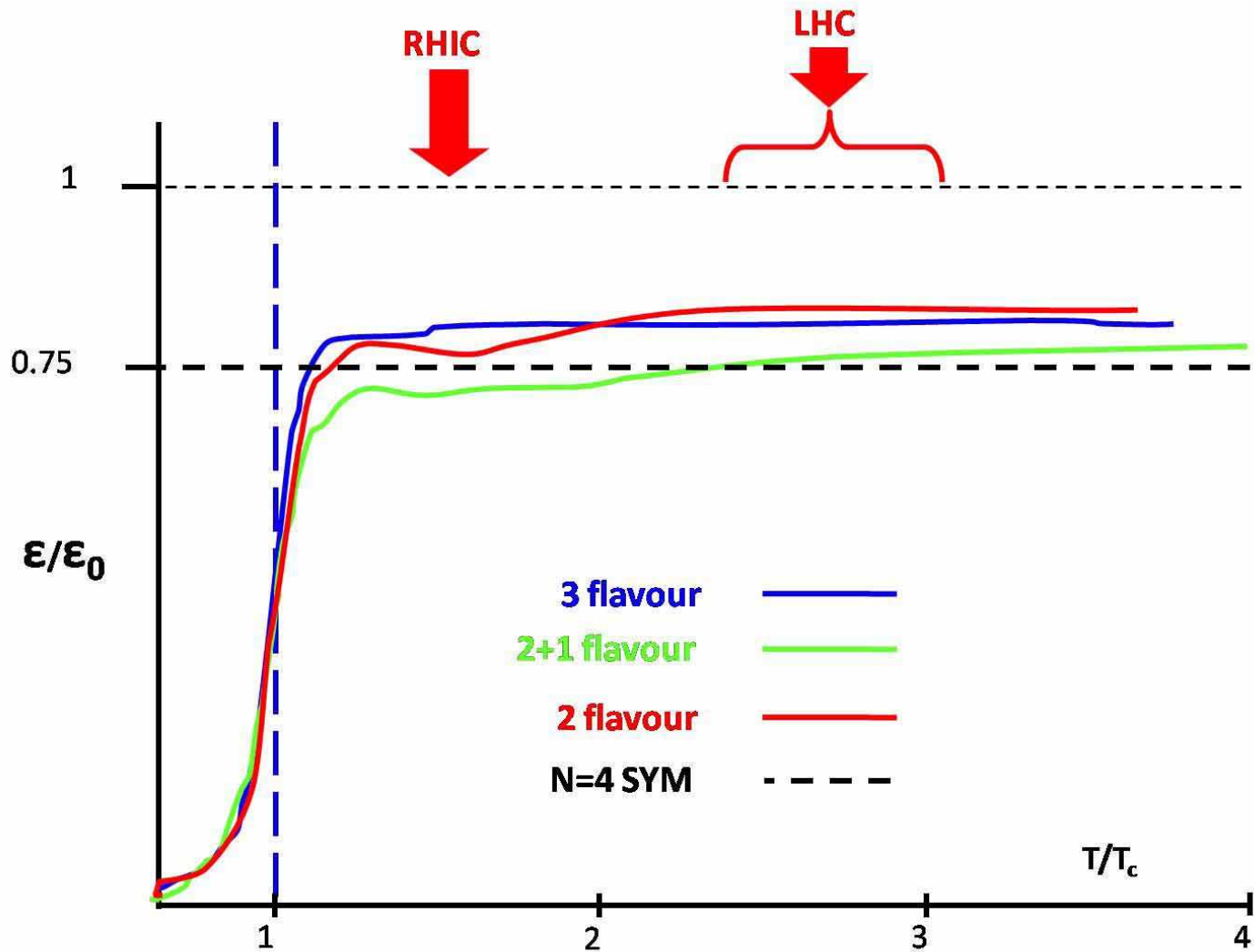
gives for the entropy density

$$s_{BH} = \frac{S}{V_3} = \frac{\pi^2 N_c^2}{2} T_H^3$$

[Gubser, Klebanov and Peet 96]

NOTE: $s_{BH} = \frac{3}{4} s_{Boltzmann}$

$$s_{Boltzmann} = \frac{4p}{T} = \frac{4\pi^2 T^3}{90} \left[\frac{7}{4} \cdot 4 + 2 + 6 \right] N_c^2 = \frac{2\pi^2 N_c^2}{3} T^3$$



energy density of QCD and SYM - via BH entropy

[from Myers and Vazquez 08]

quasinormal mode: some details

G_R from gravity action with AdS_5 metric:

$$S_{(5d)} = \frac{N_c^2}{8\pi^2 L^3} \int d^5x [\sqrt{-g}R_{(5d)} + \dots],$$

weak field limit: scalar mode and EoM (of Heun's type)

$$\sqrt{-g}R_{(5d)} \rightarrow -\frac{1}{2}\sqrt{-g}g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi, \quad \Phi \equiv h_{xy} :$$

$$\Phi(t, z, u = \frac{r_0^2}{r^2}) = \exp(-i\omega t + ikz) \Phi_k(u), \quad \Phi'_k = \frac{d}{du} \Phi_k,$$

$$\Phi_k'' - \frac{1+u^2}{u(1-u^2)} \Phi_k' + \frac{1}{4\pi^2 T^2} \frac{\omega^2 - k^2(1-u^2)}{u(1-u^2)^2} \Phi_k(u) = 0$$

quasinormal mode, cont.

solution with ansatz $\Phi_k(u) = f_k(u)\Phi_0(k)$, $f_k(0) = 1$ and
boundary condition: only the incoming wave, the one which
moves toward the horizon at $r \rightarrow r_0 \equiv u = 1$, i.e. nothing
comes out the horizon

$$f_k(u) = (1 - u^2)^{-i\omega/(4\pi T)} F_k(u) \approx 1 + i\frac{\omega}{4\pi T}u^2 + \dots$$

note, near $u \sim 1$:

$$\exp(-i\omega t) f_k(u) \rightarrow \exp[-i\omega(t + r^*)], \quad r^* = \frac{\ln(1 - u)}{4\pi T}$$

moves from $u = 0$ at $t = 0$ to $u = 1$ at $t \rightarrow \infty$

classical gravity action

inserting the solution into $S_{(5d)}$

$$S_{(5d)} \approx \frac{N_c^2}{8\pi^2 L^3} \int d^4x \int_0^1 du \sqrt{-g} \left(-\frac{1}{2}\right) g^{uu} \partial_u \Phi \partial_u \Phi$$

and keeping only the surface contribution at $u = 0 \Rightarrow$

$$S^{classical} = -\frac{\pi^2 N_c^2 T^4}{8} \int \frac{d^4k}{(2\pi)^4} \Phi_0(-k) \left[\frac{(1-u^2)}{u} f'_k(u) f_k(u) \right] \Big|_{u=0} \Phi_0(k)$$

and finally (including a counter term) \Rightarrow

$$G_R^{xy,xy}(\omega, k)$$

Bjorken flow

boost-invariant (irrotational) 1 + 1 flow [Bjorken 83]

second-order equations (proper time τ , Φ ... viscous flow):

$$\partial_\tau \epsilon = -\frac{4\epsilon}{3\tau} + \frac{\Phi}{\tau}$$

$$\tau_\pi \partial_\tau \Phi = \frac{4\eta}{3\tau} - \Phi - \frac{4\tau_\pi}{3\tau} \Phi - \frac{\lambda_1}{2\eta^2} \Phi^2$$

non-linear term NOT in MIS theory!

[BRSSS 07]

compare with AdS/CFT calculation:

$$\frac{\lambda_1 T}{\eta} = \frac{1}{2\pi} \left[1 + \frac{215 \zeta(3)}{8} \lambda^{-3/2} + \dots \right]$$

[Janik, Peschanski, Heller 06; Buchel, Paulos 08]

COMPARISON

	QCD	$\mathcal{N}=4$ SYM
$T=0$	$N_c=3=N_f$, confinement, discrete spectrum, scattering,	N_c large, N_f/N_c small, deconfined, conformal, supersymmetric,
	very different !!	
$T>T_c$	strongly-coupled plasma of gluons & fundamental matter deconfined, screening, finite corr. lengths, . . .	strongly-coupled plasma of gluons & adjoint and fundamental matter deconfined, screening, finite corr. lengths, . . .
	very similar !!	
$T \gg T_c$	runs to weak coupling	remains strongly-coupled
	very different !!	

QCD and $\mathcal{N} = 4$ SYM as a function of temperature

[from Myers and Vazquez 08]

parametric pQCD estimate

for thermalisation in an expanding gluonic medium

near equilibrium at $T(\tau)$: **Knudsen number Kn**

$$\frac{1}{Kn} = \frac{\text{longitudinal expansion time}}{\text{mean free path}} = \frac{\tau}{\lambda_f} \left(\sim 1/\frac{\eta}{s} \right) \gg 1$$

from many gluon interactions (including saturation):

Arnold et al.: $\tau_{eq} Q_s \geq \alpha_s^{-7/3}$

‘bottom-up’ [Baier, Son, Mueller and Schiff 01]

$$\tau_{eq} Q_s \geq \alpha_s^{-13/5}$$

$$\text{RHIC: } \tau_{eq} \geq 2 - 3 \text{ fm}$$