

# Testing SUSY

G.G.Ross, Planck 2010



Low energy

# Testing SUSY

The Hierarchy problem:  $M_{Higgs}, M_{W,Z} \ll M_{Planck}, M_{GUT}, \dots$  ✓

$\ll M_{SUSY}$  ??

Fine tuning  $\Rightarrow$  SUSY accessible to the LHC

Low energy

# Testing SUSY

MSSM: 105 +(19) Parameters !

Low energy

# Testing SUSY

MSSM: 105 +(19) Parameters !

- CMSSM  $\mu_0, m_0, m_{1/2}, A_0, B_0$



Focus point...non-degenerate scalars more fine-tuned

Low energy

# Testing SUSY

MSSM: 105 +(19) Parameters !

- CMSSM  $\mu_0, m_0, m_{1/2}, A_0, B_0$
- Gauge mediation - no focus point

Low energy

# Testing SUSY

MSSM: 105 +(19) Parameters !

- CMSSM  $\mu_0, m_0, m_{1/2}, A_0, B_0$
- Gauge mediation - no focus point
- Non universal gaugino masses - new focus point
- New heavy states - higher dimension operators

Low energy

# Testing SUSY

MSSM: 105 +(19) Parameters !

- CMSSM  $\mu_0, m_0, m_{1/2}, A_0, B_0$
- Gauge mediation - no focus point
- Non universal gaugino masses - new focus point
- New heavy states - higher dimension operators
- LHC - discovery potential in 1<sup>st</sup> run

# ● Fine tuning in the CMSSM

$\mu_0, m_0, m_{1/2}, A_0, B_0$

$$M_{h^0}^2 = M_Z^2 \cos^2 2\beta + \frac{3M_t^2 h_t^2}{4\pi^2} \left( \ln\left(\frac{M_S^2}{M_t^2}\right) + \delta_t \right) + \dots \geq 114 \text{ GeV (SM)}$$

$$M_Z^2 = a_0 m_0^2 + a_{1/2} m_{1/2}^2 + a_\mu \mu^2 + \dots \ll \tilde{m}_{q_i}^2, M_i^2$$

$$0.6m_{q_3}^2(M_X) + 0.6m_{U_3}^2(M_X) + 3M_3^2(M_X) + 0.2A_t^2(M_X) - 2\mu^2(M_X) + \dots$$

SPS1a

Quantitative measure of fine tuning:

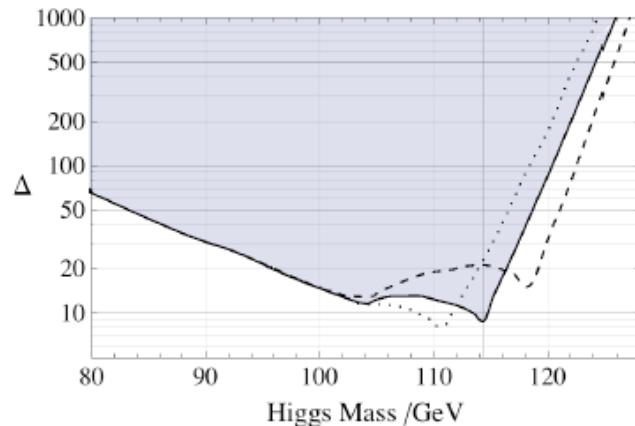
$$\Delta(a_i) = \left| \frac{(a_i \text{ or } \delta a_i)}{M_Z} \frac{\partial M_Z}{\partial a_i} \right|$$

$$\Delta_{\max} = \text{Max}_{a_i} \Delta(a_i)$$

Ellis, Enquist, Nanopoulos, Zwirner  
 Barbieri, Giudice  
 Ciafaloni, Strumia, Romanino

# ● Fine tuning in the CMSSM

## Constraints



$$\Delta \equiv \max \left| \Delta_p \right|_{p=\{\mu_0^2, m_0^2, m_{1/2}^2, A_0^2, B_0^2, h_t \dots\}}, \quad \Delta_p \equiv \frac{\partial \ln v^2}{\partial \ln p}$$

Sample technique: Mathematica code search, then SoftSUSY <sup>†</sup>

Couplings and masses evaluated to two loop (leading log) order  
...enhanced sensitivity due to small tree-level quartic coupling

$$\lambda = \frac{1}{8}(g_1^2 + g_2^2) \cos^2 2\beta$$

Cassel, Ghilencea, GGR

c.f. earlier work : Dimopoulos, Giudice  
Chankowski, Ellis, Olechowski, Pokorski

$$\Delta_{Min} = 9, \quad m_h = 114 \pm 2 \text{ GeV}$$

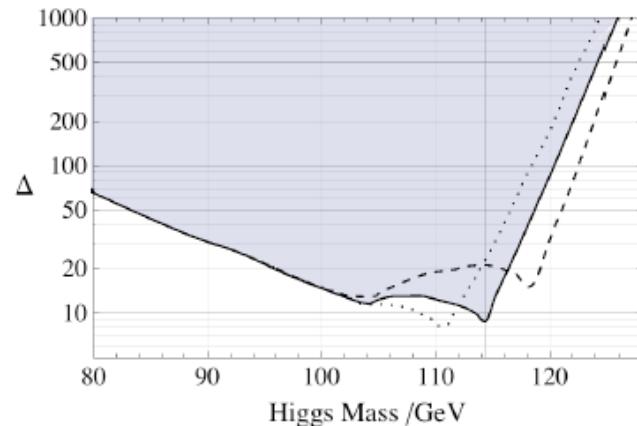
(No Higgs bound applied)

5.5 (improved minimisation of potential)

Horton

# ● Fine tuning in the CMSSM

## Constraints



SUSY particle masses

$$3.20 < 10^4 \text{ Br}(b \rightarrow s\gamma) < 3.84$$

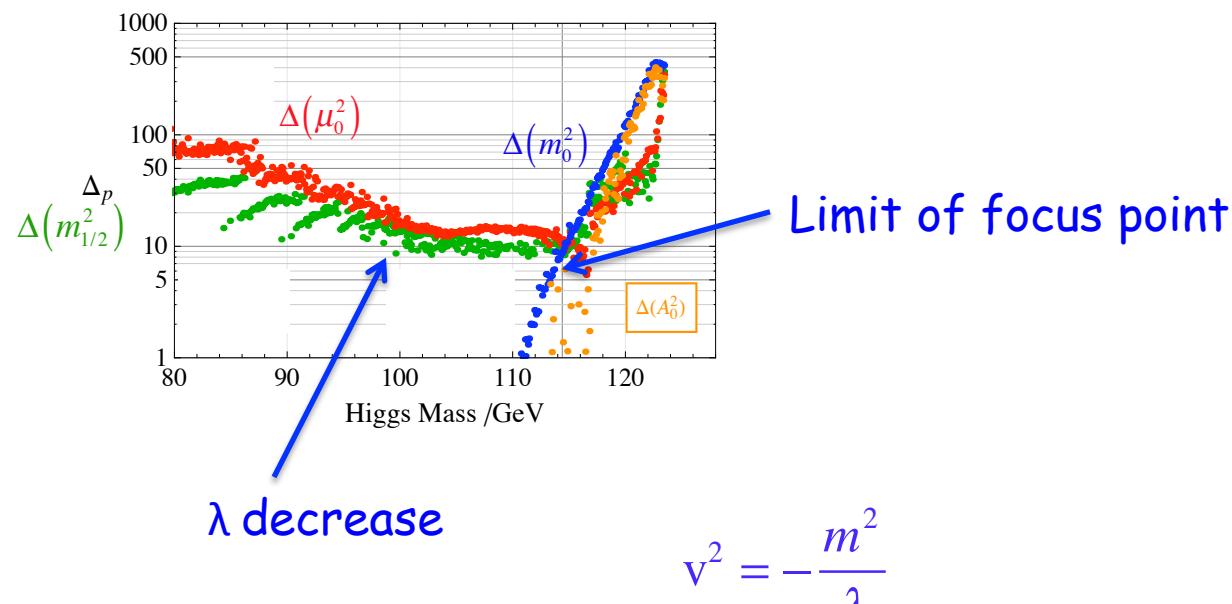
$$\text{Br}(b \rightarrow \mu\mu) < 1.8 \times 10^{-8}$$

$$\delta a_\mu < 292 \times 10^{-11}$$

$$-0.0007 < \delta\rho < 0.0012$$

Radiative EW breaking

Relic density unrestricted



$$v^2 = -\frac{m^2}{\lambda}$$

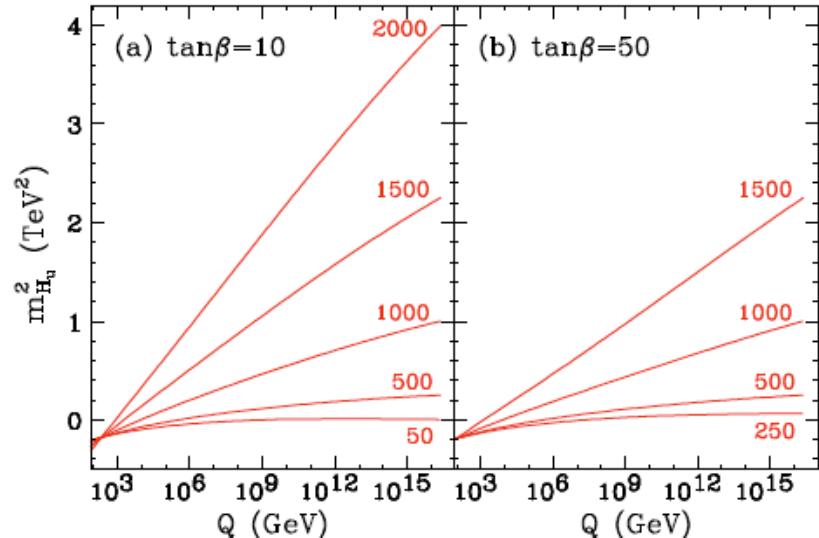
# Focus Point

$$2|y_t|^2(m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2|a_t|^2$$

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3X_t - 6g_2^2 |M_2|^2 - \frac{6}{5}g_1^2 |M_1|^2$$

$$16\pi^2 \frac{d}{dt} m_{Q_3}^2 = X_t + X_b - \frac{32}{3}g_3^2 |M_3|^2 - 6g_2^2 |M_2|^2 - \frac{2}{15}g_1^2 |M_1|^2$$

$$16\pi^2 \frac{d}{dt} m_{u_3}^2 = 2X_t - \frac{32}{3}g_3^2 |M_3|^2 - \frac{32}{15}g_1^2 |M_1|^2$$



$$m_{H_u}^2(Q^2) = m_{H_u}^2(M_P^2) + \frac{1}{2} \left( m_{H_u}^2(M_P^2) + m_{Q_3}^2(M_P^2) + m_{u_3}^2(M_P^2) \right) \left[ \left( \frac{Q^2}{M_P^2} \right)^{\frac{3y_t^2}{4\pi^2}} - 1 \right]$$

$m_0^2$        $3m_0^2$        $\simeq -\frac{2}{3}, Q^2 \simeq M_Z^2$

**“Focus point”:**  $m_{H_u}^2(0) = m_{Q_3}^2(0) = m_{u_3}^2(0) \equiv m^2$

$$m_{H_u}^2(t_0) = a_0 m^2 + \dots, a_0 \leq 0.1$$

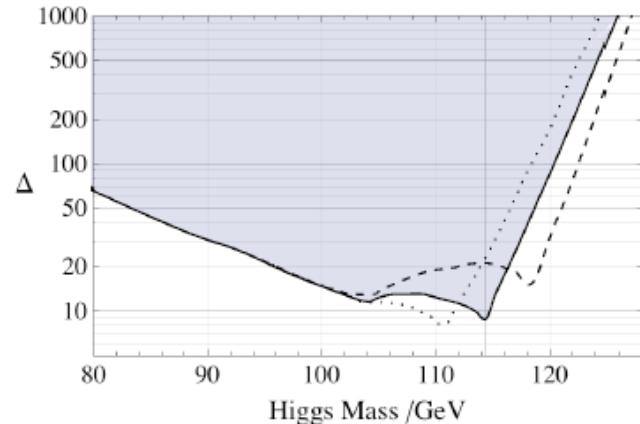
i.e.  $m_{Q_3}^3, m_{u_3}^2 \gg M_Z^2$  possible

“Natural” choice

Feng, Matchev, Moroi  
Chan, Chattopadhyay, Nath  
Barbieri, Giudice

# ● Fine tuning in the CMSSM

## Constraints



SUSY particle masses

$$3.20 < 10^4 \text{ Br}(b \rightarrow s\gamma) < 3.84$$

$$\text{Br}(b \rightarrow \mu\mu) < 1.8 \times 10^{-8}$$

$$\delta a_\mu < 292 \times 10^{-11}$$

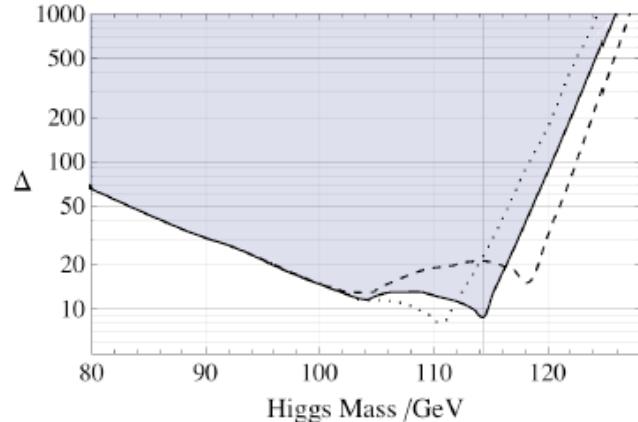
$$-0.0007 < \delta\rho < 0.0012$$

Relic density unrestricted

$$\Delta_{Min} = 9(5.5), \quad m_h = 114 \pm 2 \text{ GeV}$$

# ● Fine tuning in the CMSSM

## Constraints



SUSY particle masses

$$3.20 < 10^4 \text{ Br}(b \rightarrow s\gamma) < 3.84$$

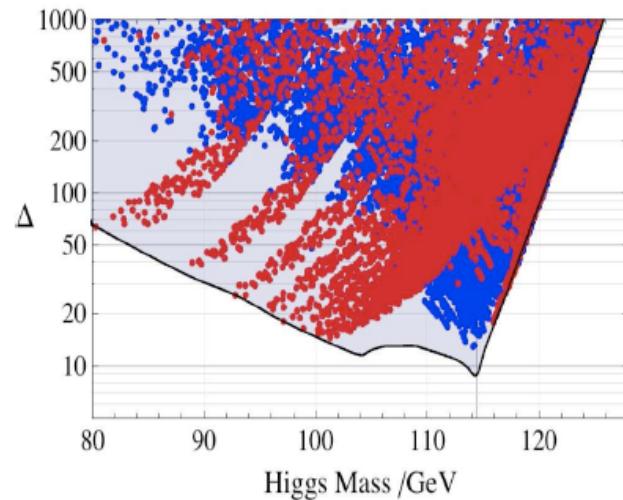
$$\text{Br}(b \rightarrow \mu\mu) < 1.8 \times 10^{-8}$$

$$\delta a_\mu < 292 \times 10^{-11}$$

$$-0.0007 < \delta\rho < 0.0012$$

Relic density unrestricted

$$\Delta_{Min} = 9(5.5), \quad m_h = 114 \pm 2 \text{ GeV}$$



Relic density restricted

■ within  $3\sigma$  WMAP

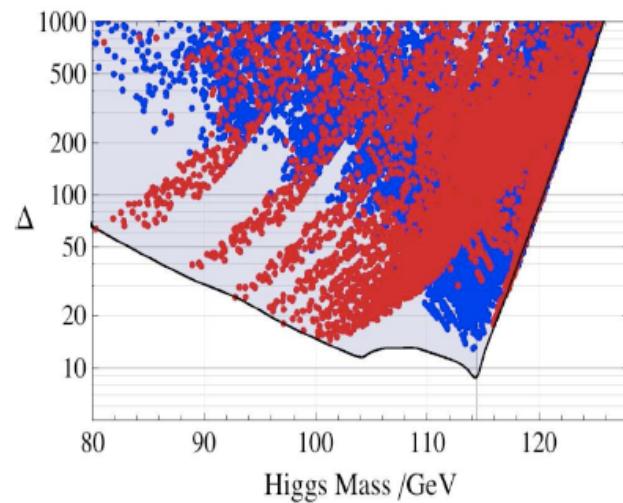
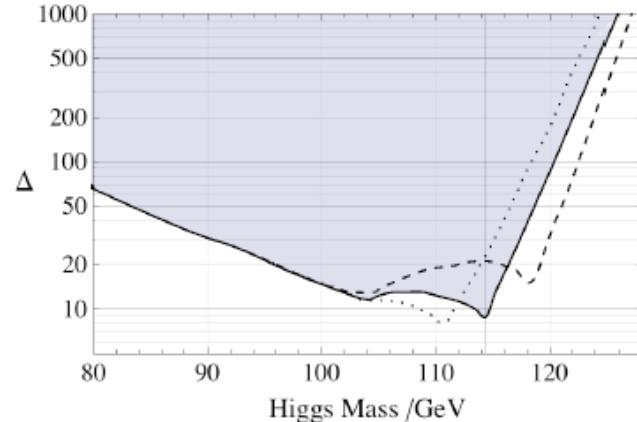
■  $< 3\sigma$  WMAP

$$\Delta_{Min} = 15, \quad m_h = 114.7 \pm 2 \text{ GeV}$$

$$\Delta_{Min} = 18, \quad m_h = 115.9 \pm 2 \text{ GeV}$$

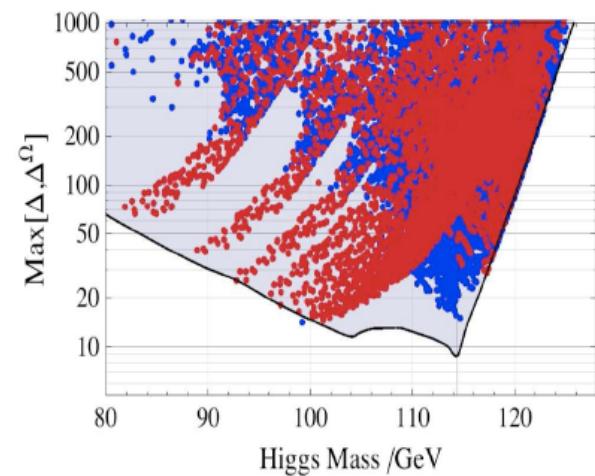
# ● Fine tuning in the CMSSM

## Constraints



$$\Delta^\Omega = \max_{q=m_0, m_{1/2}, A_0, B_0} \left| \frac{\partial \ln \Omega h^2}{\partial \ln q} \right|$$

Ellis, Olive



$$\Delta_{Min} = 29, \quad m_h = 117 \pm 2 \text{ GeV}$$

# Direct dark matter searches: CDMS (spin independent)

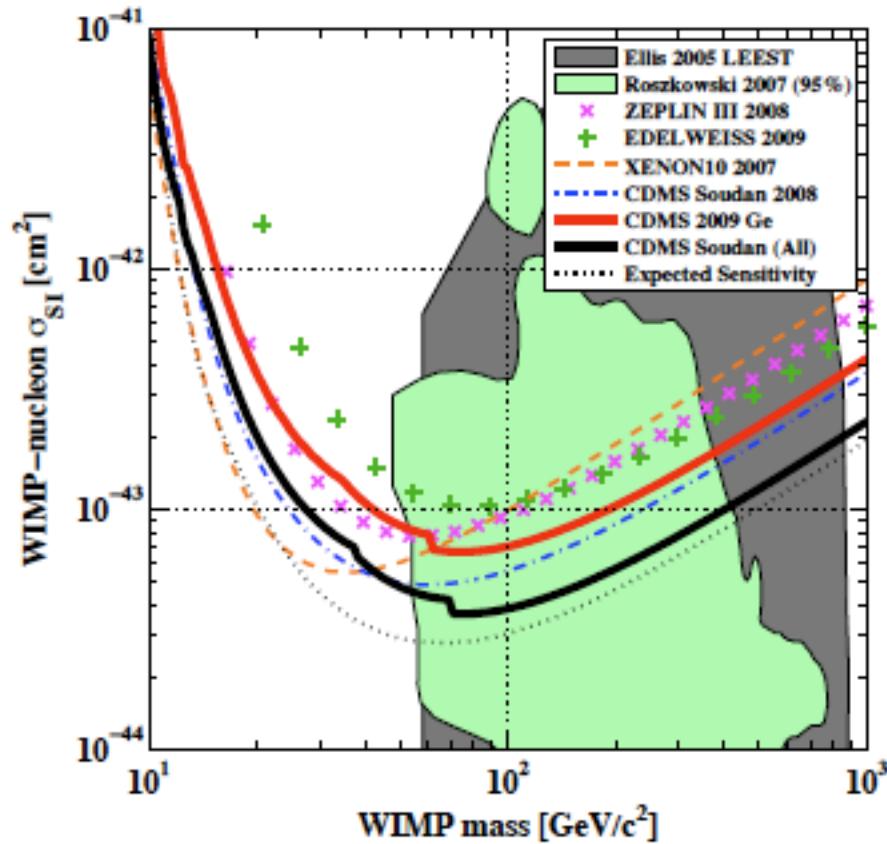
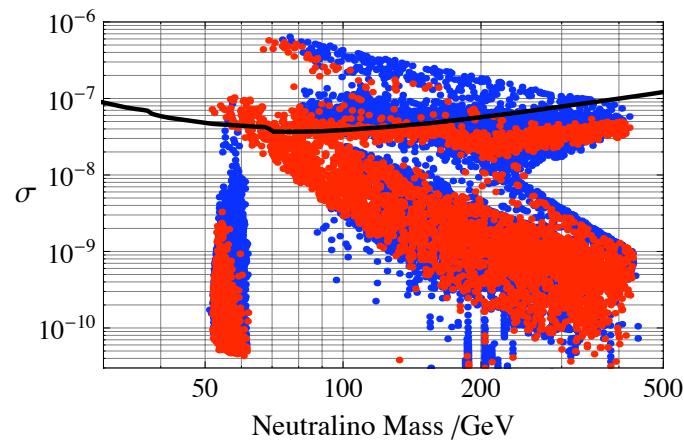
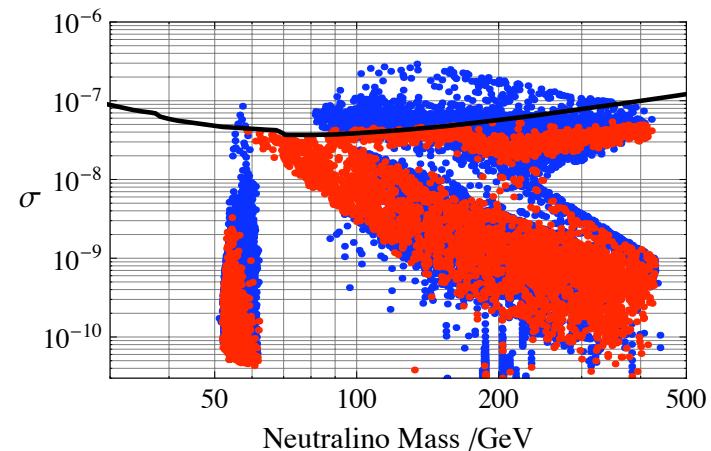


FIG. 4: 90% C.L. upper limits on the WIMP-nucleon spin-independent cross section as a function of WIMP mass. The red (upper) solid line shows the limit obtained from the exposure analyzed in this work. The solid black line shows the combined limit for the full data set recorded at Soudan.

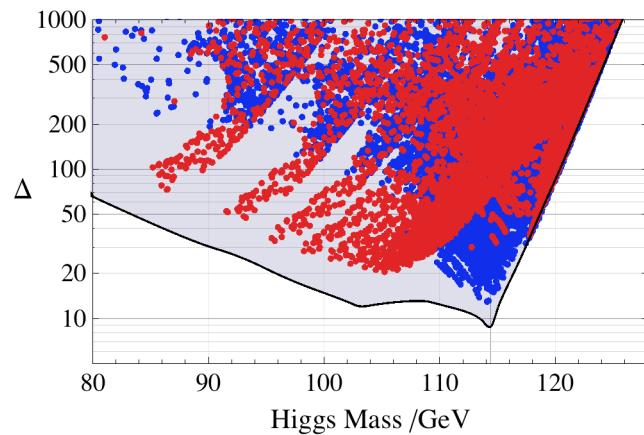
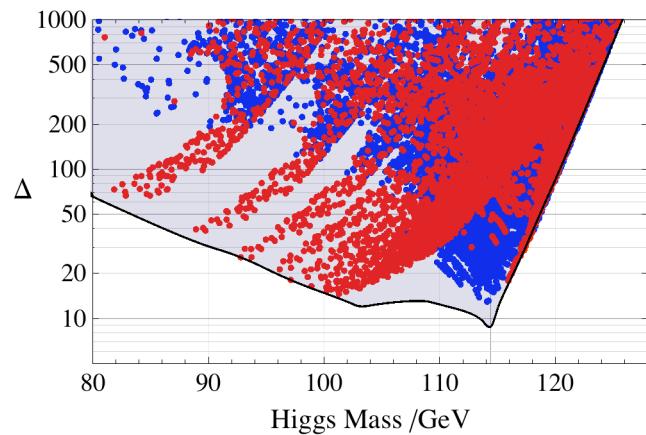
# CDMS sensitivity to CMSSM



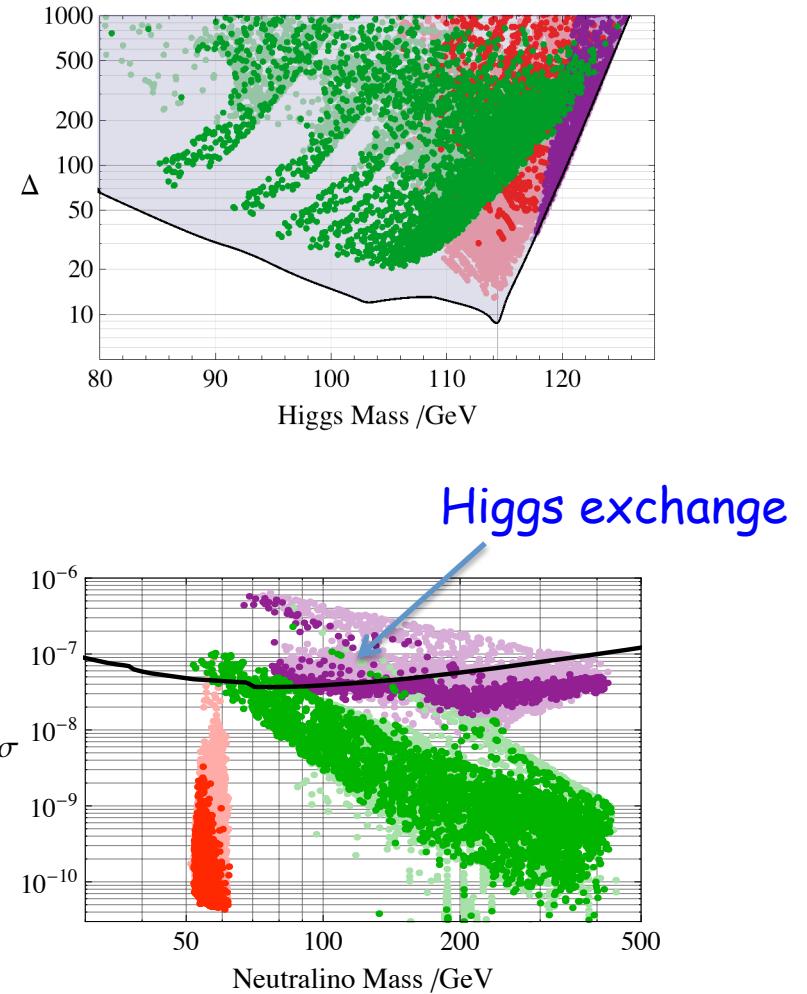
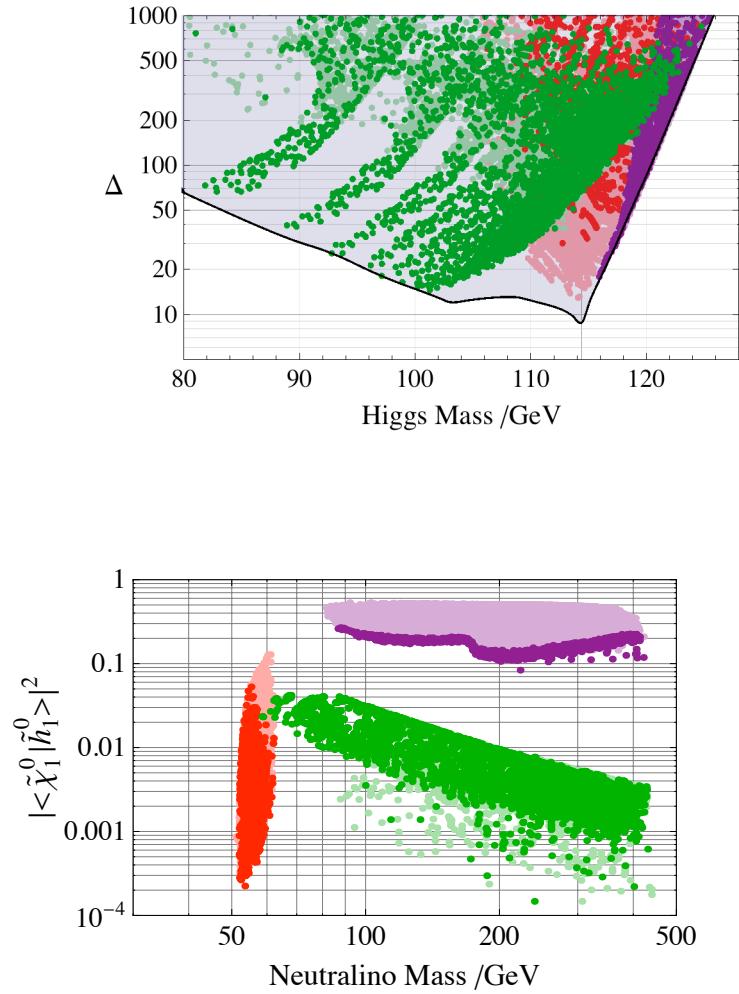
Unconstrained



... with CDMS constraint



# LSP composition



# SUSY parameters

$\Delta < 100$

$$m_h < 121 \text{ GeV}$$

$$\mu < 680 \text{ GeV}$$

$$m_0 < 3.2 \text{ TeV}$$

$$5.5 < \tan \beta < 55$$

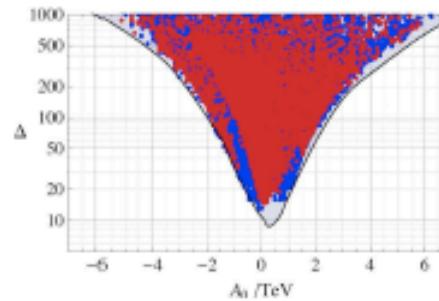
$$120 \text{ GeV} < m_{1/2} < 720 \text{ GeV}$$

$$-2.0 \text{ TeV} < A_0 < 2.5 \text{ TeV}$$

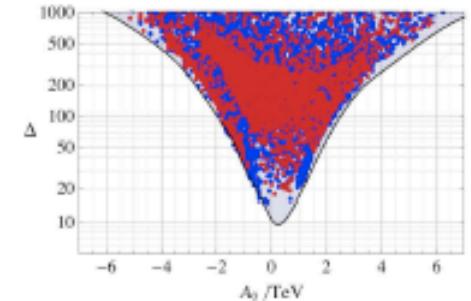
## Sparticle upper bounds

$\tilde{g}$	$\chi_1^0$	$\chi_2^0$	$\chi_3^0$	$\chi_4^0$
1720	305	550	660	665

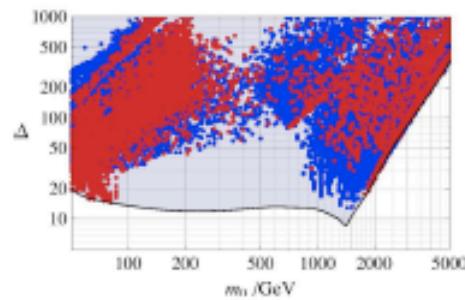
$\chi_1^\pm$	$\chi_2^\pm$	$\tilde{t}_1$	$\tilde{t}_2$	$\tilde{b}_1$	$\tilde{b}_2$
550	670	2080	2660	2660	3140



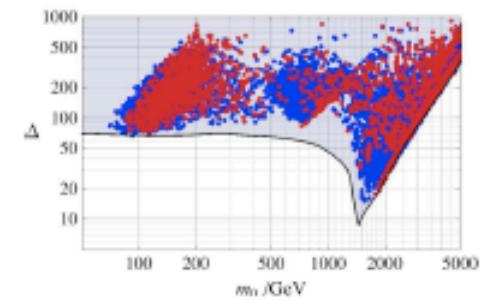
(a) Fine tuning vs  $A_0$



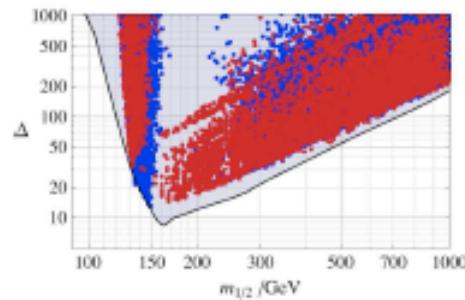
(b) Fine tuning vs  $A_0$ ,  $m_h > 114.4 \text{ GeV}$



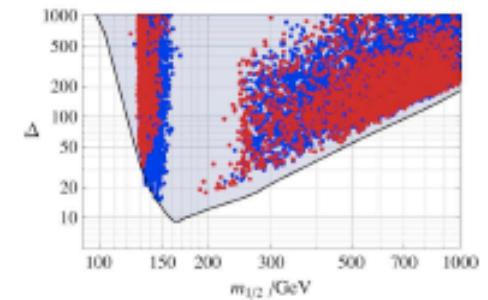
(c) Fine tuning vs  $m_0$



(d) Fine tuning vs  $m_0$ ,  $m_h > 114.4 \text{ GeV}$



(e) Fine tuning vs  $m_{1/2}$



(f) Fine tuning vs  $m_{1/2}$ ,  $m_h > 114.4 \text{ GeV}$

- (General) Gauge mediation in the MSSM

$$M_{\tilde{\lambda}_i}(M_{mess}) = k_i \frac{\alpha_i(M_{mess})}{4\pi} \Lambda_G$$

$$m_f^2(M_{mess}) = 2 \sum_{i=1}^3 C_i k_i \frac{\alpha_i^2(M_{mess})}{(4\pi)^2} \Lambda_S^2 \dagger$$

$$k_i = (\frac{5}{3}, 1, 1)$$

$$k_i \alpha_i(M_{GUT}) = 1, \quad i = 1, 2, 3$$

(Ordinary gauge mediation  $\Lambda_G = \Lambda_S$  )

Meade, Seiberg, Shih

# • (General) Gauge mediation in the MSSM

$$M_{\tilde{\lambda}_i}(M_{mess}) = k_i \frac{\alpha_i(M_{mess})}{4\pi} \Lambda_G$$

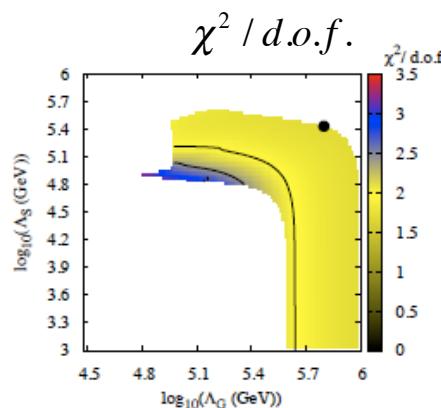
$$m_f^2(M_{mess}) = 2 \sum_{i=1}^3 C_i k_i \frac{\alpha_i^2(M_{mess})}{(4\pi)^2} \Lambda_S^2 \dagger$$

$$k_i = \left(\frac{5}{3}, 1, 1\right)$$

$$k_i \alpha_i(M_{GUT}) = 1, \quad i = 1, 2, 3$$

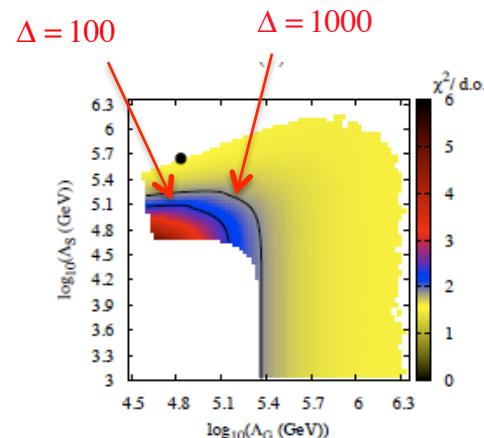
(Ordinary gauge mediation  $\Lambda_G = \Lambda_S$ )

Meade, Seiberg, Shih



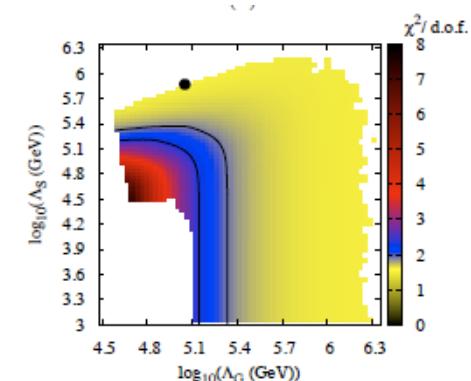
$M_{Messenger}$

$10^6 \text{ GeV}$



$10^{10} \text{ GeV}$

$B \rightarrow X_s \gamma, B \rightarrow \tau \mu, B \rightarrow \mu^+ \mu^-, B \rightarrow D \tau \mu,$   
 $D_s \rightarrow \mu \nu, D_s \rightarrow \tau \nu, K \rightarrow \mu \nu / \pi \rightarrow \mu \nu, \Delta_0 -$



$10^{14} \text{ GeV}$

$\Delta > 100$

(no focus point) $\dagger$

Abel, Dolan, Jaeckel, Khoze  
(Giusti, Romanino, Strumia)

# ● Nonuniversal gaugino masses

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3 \left( 2 |y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2 |a_t|^2 \right) - 6g_2^2 |M_2|^2 - \frac{6}{5} g_1^2 |M_1|^2$$



New focus point: cancellation between  $M_3$  and  $M_2$  contributions if  $|M_2|^2 \simeq |M_3|^2$  at  $M_{SUSY}$

Natural ratios? e.g.:

GUT:  $SU(5)$ :  $\Phi^N \subset (24 \times 24)_{symm} = 1 + 24 + 75 + 200$ ;  $SO(10)$ :  $(45 \times 45)_{symm} = 1 + 54 + 210 + 770$

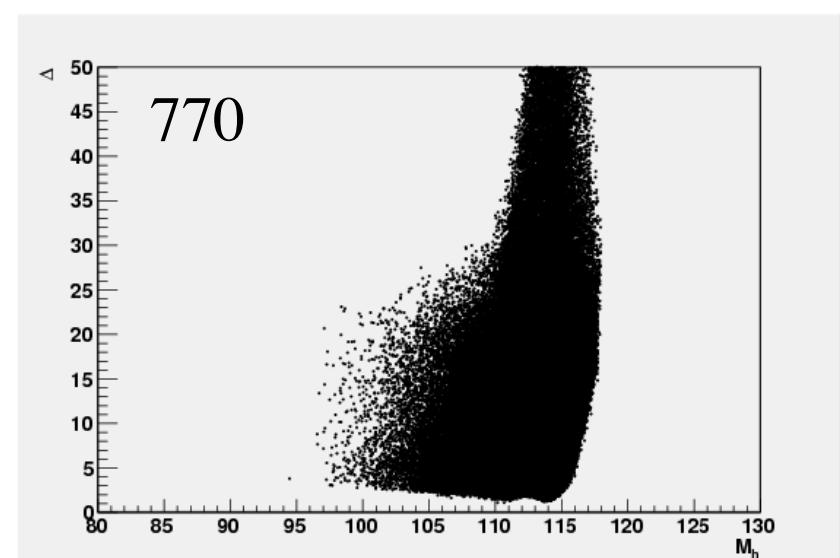
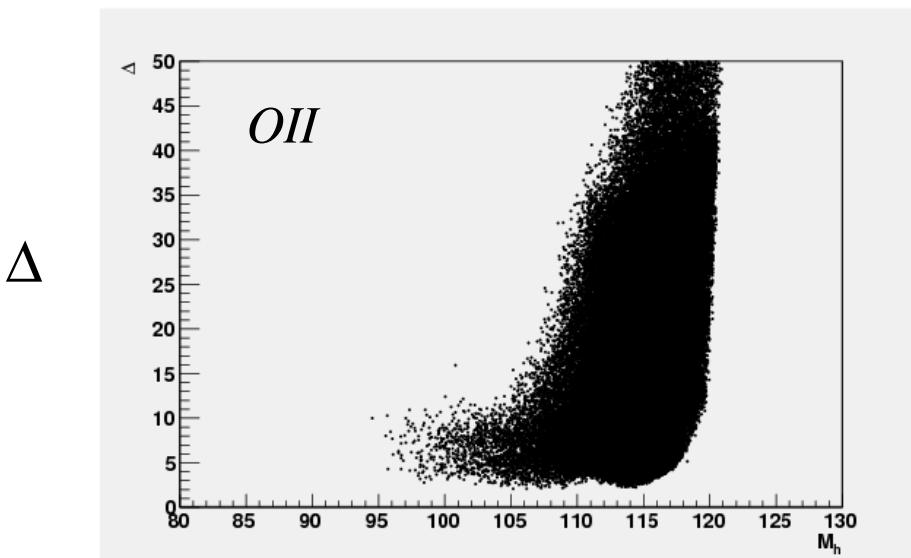
$$\eta_3 : 1 : \eta_1 \quad 2.7\eta_3 : 1 : 0.5\eta_1$$

Representation	$M_3 : M_2 : M_1$ at $M_{GUT}$	$M_3 : M_2 : M_1$ at $M_{EWSB}$
1	1:1:1	5:2:1
24	2:(-3):(-1)	11:(-6):(-1)
75	1:3:(-5)	5:6:(-5)
210	5:15:(-1)	27:30:(-1)
770	(-10):15:101	(-54):30:101

String:  $(3 + \delta_{GS}) : (-1 + \delta_{GS}) : \left( -\frac{33}{5} + \delta_{GS} \right)$   $(-11) : (-12) : 12$  (OII, also mixed moduli anomaly)

Horton, GGR  
Choi, Jeong, Kobayashi, Okumura  
Lebedev, Nilles, Ratz  
Abe, Kobayashi, Okumura

# ● Nonuniversal gaugino masses



$M_H$

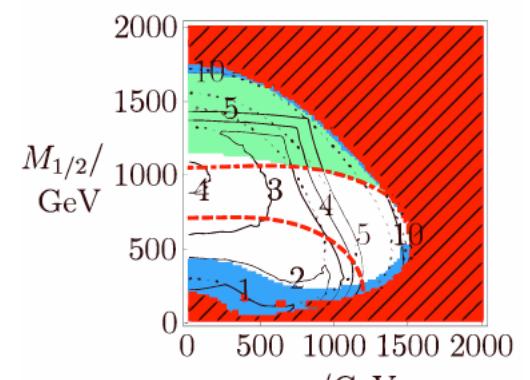
$M_H$

D.Horton, GGR

Characteristic signals:  $M_1 : M_2 : M_3 = 0.5\eta_1 : 1 : 2.7\eta_3$ , light Higgsino  $|\mu| \leq 200 \text{GeV} \ll M_{1/2}$

...gauginos can be very heavy

(Higgsino LSP subdominant cont to DM)



(d) O-II

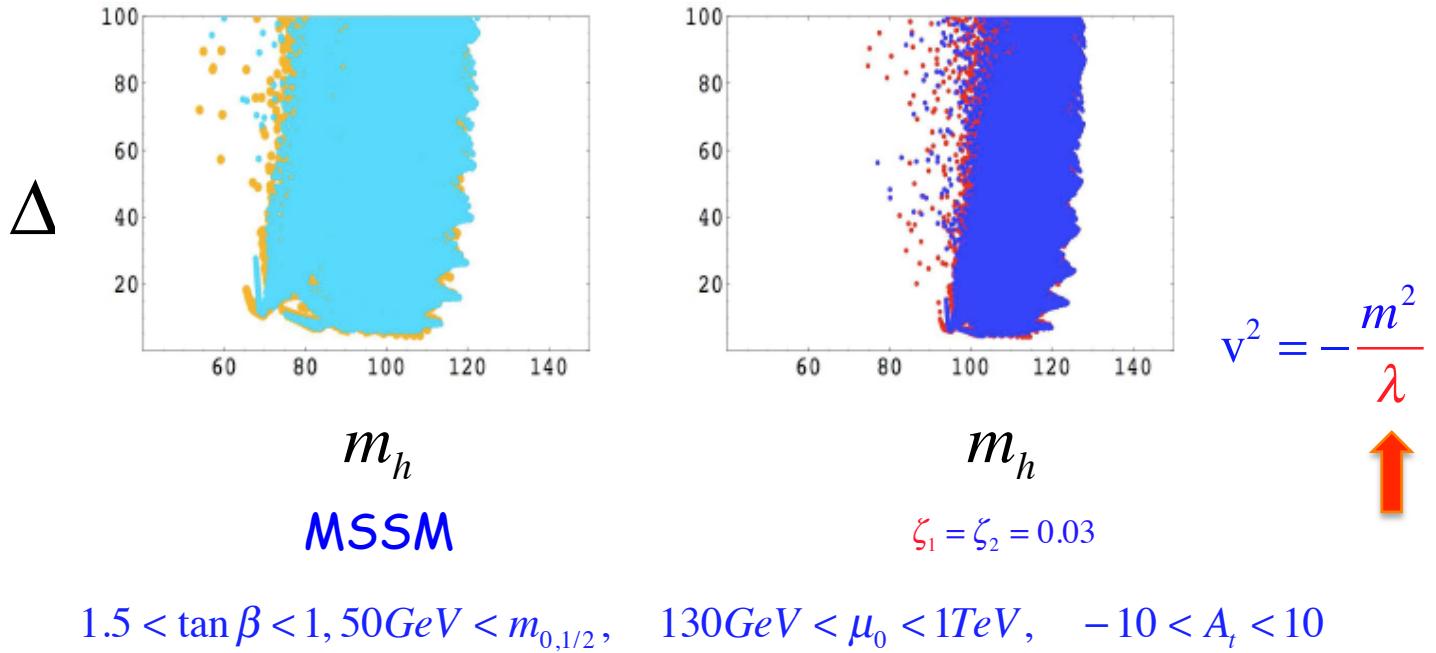
- New heavy states - higher dimension operators

e.g. Massive gauge singlet (or triplet) coupled to Higgs fields:

$$\delta L = \int d^2\theta \frac{1}{M_*} (1 + c_0 P) (H_1 H_2)^2, \quad P = m_0 \theta \theta \quad \text{Dimension 5}$$

$$\delta V = \zeta_1 2 \frac{\mu_0}{M_*} (|h_1|^2 + |h_2|^2) h_1 h_2 + \frac{1}{2} \zeta_2 (h_1 h_2)^2; \quad \zeta_1 = 2 \frac{\mu_0}{M_*}, \quad \zeta_2 = -2 \frac{c_0 m_0}{M_*}$$

## ● New heavy states - higher dimension operators



$$1.5 < \tan \beta < 1, 50 \text{GeV} < m_{0,1/2}, \quad 130 \text{GeV} < \mu_0 < 1 \text{TeV}, \quad -10 < A_t < 10$$

Cassel, Ghilencea, GGR

Even for  $M_* = 65\mu_0$  a significant shift of  $m_h$  for constant  $\Delta$   
but ... no lose ... SUSY spectrum doesn't change much!

$$\delta V = \zeta_1 2 \frac{\mu_0}{M_*} \left( |h_1|^2 + |h_2|^2 \right) h_1 h_2 + \frac{1}{2} \zeta_2 (h_1 h_2)^2$$

0 in NMSSM (only small change in fine tuning)

≠ 0 with  $M_* S^2$  term

(For an explicit model see Delgado, Kolda, Olson, la Puente)

# Summary

- Hierarchy problem  $\Rightarrow$  Low-energy SUSY

$$\Delta = \text{Max}_{a_i} \left| \frac{a_i \text{ or } \delta a_i}{M_Z} \frac{\partial M_Z}{\partial a_i} \right|, \text{ measure of residual fine-tuning}$$

- CMSSM  $\Delta_{EW} = 9(18), m_h = 114(116) \pm 2 GeV$   
 $\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29), m_h = 114(116) \pm 2 GeV$   
+ CDMS  $\rightarrow 15(32)$

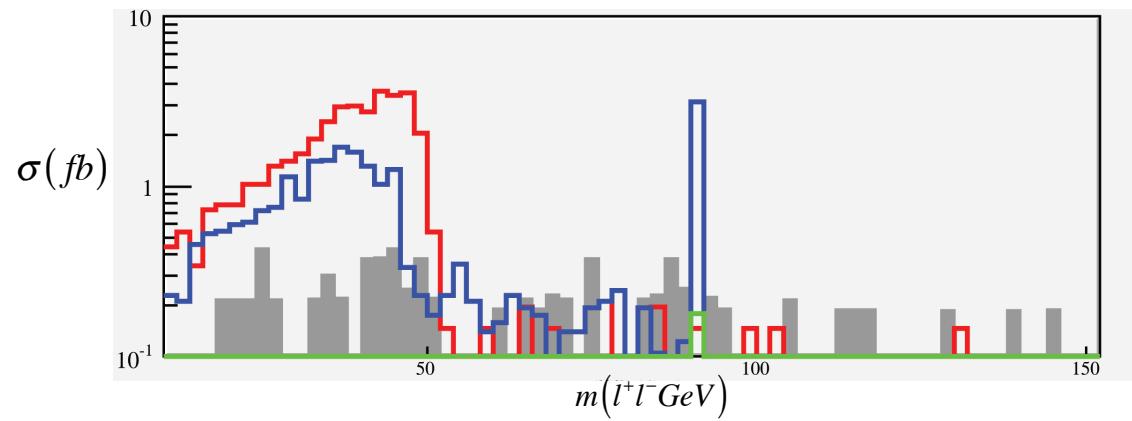
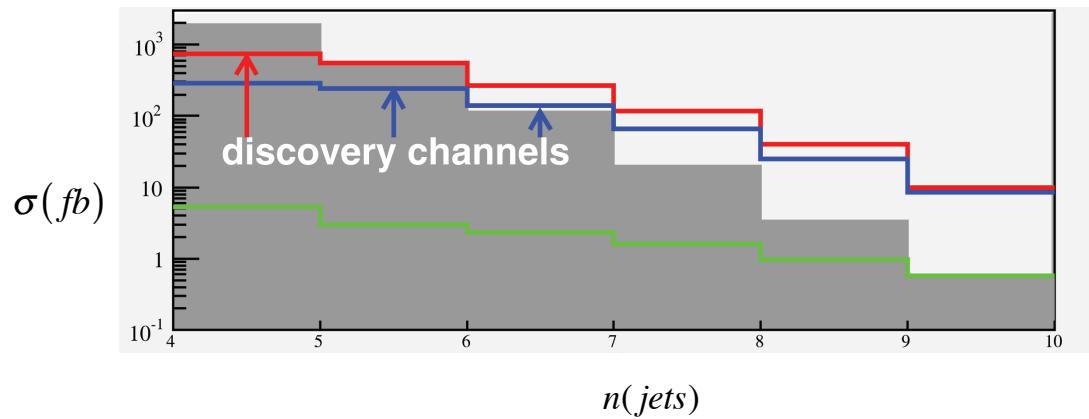
LHC  $\sqrt{s} = 10(14) TeV$   $L = 10 fb^{-1}$   $m_{\tilde{g}} \leq 1.9(2.4) TeV$   $\Delta_{EW} = 120(180)$

Baer et al

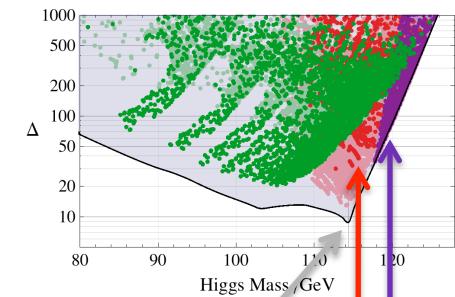
$\sqrt{s} = 7 TeV$  ?

## LHC - discovery potential in 1<sup>st</sup> run

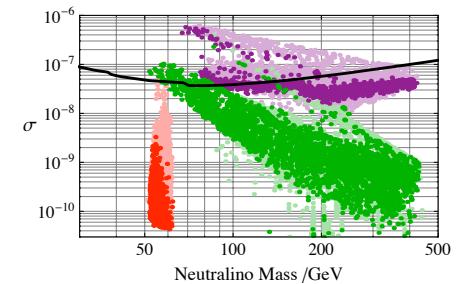
Cuts:  $E_T^{miss} > 100\text{GeV}$ ,  $n(j) \geq 4$ ,  $p_T(j) > 50\text{GeV}$ ,  $p_T(l) > 10\text{GeV}$



SM background



DM search complementarity:



# Summary

- Hierarchy problem  $\Rightarrow$  Low-energy SUSY

$$\Delta = \text{Max}_{a_i} \left| \frac{a_i \text{ or } \delta a_i}{M_Z} \frac{\partial M_Z}{\partial a_i} \right|, \text{ measure of residual fine-tuning}$$

- CMSSM  $\Delta_{EW} = 9(18), m_h = 114(116) \pm 2 GeV$   
 $\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29), m_h = 114(116) \pm 2 GeV$   
+ CDMS  $\rightarrow 15(32)$

LHC  $\sqrt{s} = 10(14) TeV$   $L = 10 fb^{-1}$   $m_{\tilde{g}} \leq 1.9(2.4) TeV$   $\Delta_{EW} = 120(180)$

$\sqrt{s} = 7 TeV$   $L = 1 fb^{-1}$   $m_{\tilde{g}} \leq 0.45 - 0.65 TeV$   $\Delta_{EW} = 18(23)$   $\Delta_{TOT} = 35(40)$

# Summary

- Hierarchy problem  $\Rightarrow$  Low-energy SUSY

$$\Delta = \text{Max}_{a_i} \left| \frac{a_i \text{ or } \delta a_i}{M_Z} \frac{\partial M_Z}{\partial a_i} \right|, \text{ measure of residual fine-tuning}$$

- CMSSM  $\Delta_{EW} = 9(18), m_h = 114(116) \pm 2 GeV$   
 $\text{Max}[\Delta_{EW}, \Delta_\Omega] = 15(29), m_h = 114(116) \pm 2 GeV$   
+ CDMS  $\rightarrow 15(32)$

LHC  $\sqrt{s} = 10(14) TeV$   $L = 10 fb^{-1}$   $m_{\tilde{g}} \leq 1.9(2.4) TeV$   $\Delta_{EW} = 120(180)$

$\sqrt{s} = 7 TeV$   $L = 1 fb^{-1}$   $m_{\tilde{g}} \leq 0.45 - 0.65 TeV$   $\Delta_{EW} = 18(23)$   $\Delta_{TOT} = 35(40)$

- MSSM...
  - Gauge mediation:  $\Delta > 100$
  - Nondegenerate gauginos:  $\Delta > 3, m_{\tilde{g}} \leq 4 TeV$
  - + D=5 operators:  $m_h \rightarrow 130 GeV$

...Higgs search ultimately the most sensitive

