# Linear Theory

$$
a = \exp(A \cdot \nabla)I. \tag{3.28}
$$

The map  $I$  is the identity map of phase space: it is made of two trivial projection functions in 1-d-f, in other words

$$
I(\mathbf{z}) = \begin{pmatrix} z_1 \\ z_2 \end{pmatrix} . \tag{3.29}
$$

If the vector field A is symplectic (or Hamiltonian), it means that a Hamiltonian  $H_a$ can achieve the same result:

$$
A \cdot \nabla = \frac{\partial H_a}{\partial z_1} \frac{\partial}{\partial z_2} - \frac{\partial H_a}{\partial z_2} \frac{\partial}{\partial z_1} =: H_a : .
$$
 (3.30)

It can be shown easily, in the Hamiltonian case, that Eq.  $(3.26)$  can be written as:

$$
K^{new} = K \circ a = \exp\left(\frac{H_a}{H_a}\right)K. \tag{3.31}
$$



$$
M = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \implies M = ARA^{-1}
$$
  
1 = ad - bc



A is also assumed to have  $det(A)=1$ .

$$
\text{Invariant (nonlinear OK)}\\
\begin{pmatrix} x_n \\ p_n \end{pmatrix} = A^{-1} \begin{pmatrix} x \\ p \end{pmatrix} = \begin{pmatrix} A_{22} & -A_{12} \\ -A_{21} & A_{11} \end{pmatrix} \begin{pmatrix} x \\ p \end{pmatrix}
$$

$$
M = A R A^{-1}
$$

$$
\text{Invariant} = x_n^2 + p_n^2 = (A_{22}x - A_{12}p)^2 + (-A_{21}x + A_{11}p)^2
$$
  
=  $\underbrace{(A_{22}^2 + A_{21}^2)x^2 + 2(-A_{12}A_{22} - A_{21}A_{11})xp + (A_{11}^2 + A_{12}^2)p^2}_\alpha$ 

**Lattice Functions (linear only)**  
\n
$$
\underbrace{\begin{pmatrix} a & b \\ c & d \end{pmatrix}}_{M} = \underbrace{\begin{pmatrix} A_{11} & A_{12} \\ A_{21} & A_{22} \end{pmatrix}}_{A} \begin{pmatrix} \cos(\mu) & \sin(\mu) \\ -\sin(\mu) & \cos(\mu) \end{pmatrix} \underbrace{\begin{pmatrix} A_{22} & -A_{12} \\ -A_{21} & A_{11} \end{pmatrix}}_{A^{-1}}
$$
\n
$$
M = \cos(\mu)I + \sin(\mu) \begin{pmatrix} \underbrace{-A_{11}A_{21} - A_{12}A_{22}}_{\alpha} & \underbrace{A_{11}^2 + A_{12}^2}_{\beta} \\ -\underbrace{A_{21}^2 - A_{22}^2}_{\gamma} & \underbrace{A_{11}A_{21} + A_{12}A_{22}}_{\alpha} \end{pmatrix}
$$

$$
\Rightarrow M = \begin{pmatrix} \cos(\mu) + \alpha \sin(\mu) & \beta \sin(\mu) \\ -\gamma \sin(\mu) & \cos(\mu) - \alpha \sin(\mu) \end{pmatrix}
$$

 $1+\alpha^2=\beta\gamma$ Consequence of det(A)=1

#### 2.2.4 Lattice functions as coefficients of the moments  $\langle z_iz_j \rangle$

Here let us use the matrix A directly and compute the three quadratic moments of the 1-d-f linear theory.

$$
\langle z_1^2 \rangle = \langle (A_{11} z_1^{new} + A_{12} z_2^{new})^2 \rangle = A_{11}^2 \langle z_1^{new2} \rangle + A_{12}^2 \langle z_2^{new2} \rangle + 2A_{11} A_{12} \langle z_1^{new} z_2^{new} \rangle
$$
  
= 
$$
\underbrace{(A_{11}^2 + A_{12}^2)}_{\beta} \frac{\langle r^2 \rangle}{2}.
$$
 (2.34)

In Eq.  $(2.34)$ , if we are dealing with a single particle and performing a time average, then the average can be removed in the final expression. In Eq.  $(2.34)$  I use the fact that that  $z_1^{new}$  and  $z_2^{new}$  move on circles and thus the value  $z_1^{new2} + z_2^{new2}$  is constant on a trajectory. Other averages can be computed as well:

$$
\langle z_2^2 \rangle = \langle (A_{21} z_1^{new} + A_{22} z_2^{new})^2 \rangle = A_{21}^2 \langle z_1^{new2} \rangle + A_{22}^2 \langle z_2^{new2} \rangle + \langle A_{21} A_{22} \langle z_1^{new} z_2^{new} \rangle
$$
  
= 
$$
\underbrace{(A_{21}^2 + A_{22}^2)}_{\gamma} \underbrace{\langle r^2 \rangle}_{\gamma}
$$
 (2.35)

and more interestingly,

$$
\langle z_1 z_2 \rangle = \langle (A_{11} z_1^{new} + A_{12} z_2^{new}) (A_{21} z_1^{new} + A_{22} z_2^{new}) \rangle
$$
  
=  $A_{11} A_{21} \langle z_1^{new2} \rangle + A_{12} A_{22} \langle z_2^{new2} \rangle + \{A_{11} A_{22} + A_{12} A_{21} \} \langle z_1^{new} z_2^{new} \rangle$   
=  $\underbrace{(A_{11} A_{21} + A_{12} A_{22})}_{-\alpha} \frac{\langle r^2 \rangle}{2}$  (2.36)

$$
M = \cos(\mu)H + \sin(\mu)B\tag{3.1}
$$

where

$$
H = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \text{ and } B = \begin{pmatrix} \alpha & \beta \\ -\gamma & -\alpha \end{pmatrix} \text{ with } B^2 = -H
$$

Using this formula, which is a representation of the unit complex numbers, we immediately see that

 $\sim$ 

$$
M^N = \cos(N\mu)H + \sin(N\mu)B\tag{3.2}
$$

Eq.  $(B.2)$  is the famous de Moivre formula known to accelerator physicists. The notation  $H$  (as in capital  $\eta$ ), for what appears to be the identity, anticipates the results in more than 1-d-f. We will see that in 3-d-f, the matrix  $H$  is not the identity but corresponds to generalised dispersions of reference [4].

The matrix  $B$  (as in beta functions) contains the regular lattice functions as defined by Ripken. In fact we have the following results:

$$
B = SK \tag{3.3}
$$

$$
B = ES \tag{3.4}
$$

$$
K = SES^{\mathrm{T}} \tag{3.5}
$$

where 
$$
S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}
$$

The matrix K of Eq.  $(B.3)$  defines the Courant-Snyder invariant

$$
K = \begin{pmatrix} \gamma & \alpha \\ \alpha & \beta \end{pmatrix} \implies \varepsilon(z) = z^{\mathrm{T}} K z = \gamma z_1^2 + 2\alpha z_1 z_2 + \beta z_2^2 \tag{3.6}
$$

The matrix E of Eq.  $(3.4)$  gives us the quadratic moments:

$$
E = \begin{pmatrix} \beta & -\alpha \\ -\alpha & \gamma \end{pmatrix} \implies \langle z_i z_j \rangle = E_{ij} \langle \frac{z_1^2 + z_2^2}{2} \rangle = E_{ij} \langle J \rangle
$$
  
for example  $\langle z_1^2 \rangle = \beta \langle J \rangle$  (3.7)

Finally, the one-turn map  $M$  can be written with a single exponent and we will see that all the results of this section can be extended to many dimensions.

$$
M = \exp(\mu B) = \exp(\mu SK) \tag{3.8}
$$

From Eq.  $(6.8)$  one can derive two interesting corollaries. Firstly the invariant of Eq.  $(6.6)$ , the so called Courant-Snyder invariant, is a pseudo-Hamiltonian for the matrix M. The Hamiltonian  $\mathcal{H}$ ,

$$
\mathcal{H} = \frac{\mu}{2} z^{\mathrm{T}} K z \tag{3.9}
$$

#### 4. De Moivre's formula in many degrees of freedom

The derivation of this formula only requires the existence of a normal form. We might as well do it in the non-symplectic case. In 3-d-f, when classical radiation is present, the map  $M$  can be factorised as

$$
M = A\Lambda R A^{-1}
$$
\n
$$
R = \begin{pmatrix} r_1 & 0 & 0 \\ 0 & r_2 & 0 \\ 0 & 0 & r_3 \end{pmatrix} \qquad r_i = \begin{pmatrix} \cos \mu_i & \sin \mu_i \\ -\sin \mu_i & \cos \mu_i \end{pmatrix}
$$
\n
$$
\Lambda = \begin{pmatrix} \Lambda_1 & 0 & 0 \\ 0 & \Lambda_2 & 0 \\ 0 & 0 & \Lambda_3 \end{pmatrix} \qquad \Lambda_i = \begin{pmatrix} \exp(-\alpha_i) & 0 \\ 0 & \exp(-\alpha_i) \end{pmatrix}
$$
\n(4.2)

<sup>&</sup>lt;sup>3</sup>It was reported to me by Frank Schmidt of CERN, that Dr. Ripken referred to the dispersion-like lattice functions of reference [4] as "the only thing in linear Hamiltonian perturbation theory" he had been unaware. I surmise that he would have like the cute result of this paper- that his theory contained an exact form of Sands' approximate methods.

Applying Eq. (4.2) to Eq. (4.1) we get

$$
M = A\left(\sum_{i} e^{-\alpha_i} \left\{\cos\left(\mu_i\right)I^i + \sin\left(\mu_i\right)SI^i\right\}\right) A^{-1} \tag{4.3}
$$

$$
=\sum_{i}e^{-\alpha_{i}}\left\{\cos\left(\mu_{i}\right)\underbrace{AI^{i}A^{-1}}_{H^{i}}+sin\left(\mu_{i}\right)\underbrace{ASI^{i}A^{-1}}_{B^{i}}\right\}\tag{4.4}
$$

where  $I^i$  is the identity only in the  $i^{th}$  plane. For example,

$$
I^{2} = \begin{pmatrix} 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \end{pmatrix}
$$
(4.5)

It is easy to show that the matrices  $H^i$  and  $B^i$  form three independent representations of the complex numbers 1 and  $i$ :

$$
H^i H^j = \delta_{ij} H^j \tag{4.6}
$$

$$
H^i B^j = B^j H^i = \delta_{ij} B^j \tag{4.7}
$$

$$
B^i B^j = -\delta_{ij} H^j \tag{4.8}
$$

De Moivre's general formula follows from the three previous equations.

$$
M^N = \sum_i e^{-N\alpha_i} \left\{ \cos(N\mu_i) H^i + \sin(N\mu_i) B^i \right\}
$$
 (4.9)

Moreover, we can again rewrite the map  $M$  in terms of a single exponent:

$$
M = \exp\left(\sum_{i} -\alpha_{i}H^{i} + \mu_{i}B^{i}\right) = \exp\left(\sum_{i} -\alpha_{i}H^{i} + \mu_{i}SK^{i}\right)
$$
(4.10)

There are a few comments we need to make concerning the results of this section.

- 1. As shown by Eq. (4.10), the lattice functions  $H^i$  cannot be removed easily from any treatment of a non-symplectic system. Of course, using Eq.  $(\underline{4.8})$ , we can replace them by quadratic polynomials in the  $B^{i}$ 's.
- 2. Forest, in [4], has shown that analytical tune shift formulas for an arbitrary force depend on the  $B^{i}$ 's while the damping shifts depend on the  $H^{i}$ 's.
- 3. In a symplectic system, if perturbation theory is done on the Hamiltonian, the lattice functions  $H^i$  will not naturally appear since the exponent in Eq. (4.10) only depends of the  $B^i$ 's (when the damping is zero). This is why they never appear in Ripken's extensive literature on the subject.

#### 5.2 Other Useful Properties

Properties equivalent those of Eqs.  $(6.3)$ ,  $(6.4)$  and  $(6.5)$  can be proved in the multidimensional case:

$$
B^i = SK^i \tag{5.11}
$$

$$
B^i = E^i S \tag{5.12}
$$

$$
K^i = SE^i S^T \tag{5.13}
$$

For example, the pseudo-Hamiltonian which generates the matrix  $M$  is in the sum of the three invariants weighted by the tunes:

$$
\mathcal{H} = \sum_{i=1,3} \frac{\mu_i}{2} z^{\mathrm{T}} K^i z = \sum_{i=1,3} \frac{\mu_i}{2} \varepsilon_i(z)
$$
\n(5.14)

$$
\varepsilon_i(z) = z^T K^i z \quad \leftarrow \quad \text{Coupled Courant} - \text{Snyder Invariants} \tag{5.15}
$$

As for the quadratic moments, they are given by  $E$ :

$$
\langle z_a z_b \rangle = \sum_{i=1,3} E_{ab}^i \langle J_i \rangle \tag{5.16}
$$

There are many more interesting corollaries but that would take our story into an unwanted tangent. Let us return to our main discovery: the strange and unexpected equivalence between Sands and Chao.

## Possible A: select  $A_{11}$  > 0 and  $A_{12}$  = 0

$$
A_{11} = \left| \frac{M_{12}}{\sin{(\mu)}} \right|^{1/2}
$$

$$
A_{22} = A_{11}^{-1}
$$

$$
A_{21} = \frac{(M_{21} + \sin{(\mu)}A_{11}^{-2})A_{11}}{\cos{(\mu)} - M_{22}}
$$

### A in terms of lattice functions

$$
A_{11} = A_{22}^{-1} = \left| \frac{M_{12}}{\sin(\mu)} \right|^{1/2} = \left| \frac{\beta \sin(\mu)}{\sin(\mu)} \right|^{1/2} = \sqrt{\beta}
$$

$$
A_{21} = \frac{\left(M_{21} + \sin\left(\mu\right)A_{11}^{-2}\right)A_{11}}{\cos\left(\mu\right) - M_{22}} = \frac{\left(-\gamma\sin\left(\mu\right) + \frac{\sin\left(\mu\right)}{\beta}\right)\sqrt{\beta}}{\cos\left(\mu\right) - \left(\cos\left(\mu\right) - \alpha\sin\left(\mu\right)\right)}
$$

$$
= \frac{\left(-\gamma + \frac{1}{\beta}\right)\sqrt{\beta}}{\alpha} = -\frac{\alpha}{\sqrt{\beta}}
$$

$$
\Longrightarrow A = \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix}
$$

### Evolution of invariant (nonlinear OK)

$$
H = \frac{H_{22}(s)}{2}p^2 + \frac{H_{11}(s)}{2}x^2 + H_{12}(s)xp
$$

$$
H_{11} = k(s) \equiv \text{Focusing quadrupole strength}
$$
  

$$
H_{22} = 1
$$
  

$$
H_{12} = 0.
$$

$$
\mathcal{M}_{s+ds} = \mathcal{M}_{s \to s+ds}^{-1} \mathcal{M}_s \mathcal{M}_{s \to s+ds}
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.89)} \qquad \exp\left(ds : H : \right) \mathcal{M}_s \exp\left(ds : -H : \right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left(ds : H : \right) \exp\left(: -\frac{\mu}{2} \varepsilon_s : \right) \exp\left(ds : -H : \right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left(ds : H : \right) \exp\left(: -\frac{\mu}{2} e^{ds : H : \varepsilon_s : \right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left(: -\frac{\mu}{2} e^{ds : H : \varepsilon_s : \cdot} \right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left((ds : H : \cdot) \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \right) \exp\left((ds : -H : \cdot) \right)\right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left((ds : H : \cdot) \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \right) \exp\left((ds : -H : \cdot) \right)\right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \right) \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \right)\right)
$$
  
\n
$$
\mathbb{E}_{\mathbf{q}^s. (2.43, 2.44, 2.45)} \exp\left((s + \frac{\mu}{2} \varepsilon_s : \cdot) \exp\left((s + \frac{\mu}{2} \varepsilon_s : \
$$

$$
\frac{d\varepsilon_s}{ds} = [H_2, \varepsilon_s]
$$
  
= 2(\alpha H\_{11} - \gamma H\_{12})x^2 + 2(\beta H\_{11} - \gamma H\_{22})xp  
+2(\beta H\_{12} - \alpha H\_{22})p^2  

$$
\Rightarrow \begin{cases} \frac{d\gamma}{ds} = 2(\alpha H_{11} - \gamma H_{12}) \\ \frac{d\alpha}{ds} = \beta H_{11} - \gamma H_{22} \\ \frac{d\beta}{ds} = 2(\beta H_{12} - \alpha H_{22}) \end{cases}.
$$

If the Hamiltonian is the usual one:

$$
\frac{d\gamma}{ds} = 2\alpha \ k(s)
$$

$$
\frac{d\alpha}{ds} = \beta k(s) - \gamma
$$

$$
\frac{d\beta}{ds} = -2\alpha.
$$

### Nonlinear works

$$
\frac{dI_s}{ds} = [H, I_s].\tag{2.94}
$$

With Maps: In general, consider a quantity  $I_s$  which is invariant under  $\mathcal{M}_s$ :

$$
I_s = \mathcal{M}_s I_s
$$
  
\n
$$
\mathcal{M}_{s \to s+\Delta}^{-1} I_s = \underbrace{\mathcal{M}_{s \to s+\Delta}^{-1} \mathcal{M}_s \mathcal{M}_{s \to s+\Delta}}_{\mathcal{M}_{s+ds}} \mathcal{M}_{s \to s+\Delta}^{-1} I_s
$$
  
\n
$$
\downarrow
$$
  
\n
$$
I_{s+\Delta} = \mathcal{M}_{s \to s+\Delta}^{-1} I_s.
$$
\n(2.95)

In Equation (2.95), the time shift  $\Delta$  can be finite. The result of Equation (2.94) is gotten by replacing  $\Delta$  by ds.

Writing  $\mathcal{R}_{s\to s+ds}$  as  $\exp(-d\mu : J :)$ , we then compute the map  $\mathcal{A}_{s+ds}$ to first order in  $ds$ :

$$
\exp(-d\mu : J: ) = A_s e^{ds: -H:} A_{s+ds}^{-1}
$$
  

$$
\downarrow \qquad \downarrow
$$
  

$$
A_{s+ds} = A_s - ds A_s : H: +d\mu : J: A_s. \qquad (2.99)
$$

Up to this point, we have not used the fact that we interested in the Courant-Snyder transformation. This transformation is characterized by the vanishing  $A_{12}$  entry of its associated transfer map. To get this entry, we act on the function "x" with  $A_{s+ds}$  and we require the coefficient of the function " $p$ " to be zero:

$$
\mathcal{A}_{s+ds}x = \mathcal{A}_{s}x - ds\mathcal{A}_{s}[H, x] + d\mu : J : \mathcal{A}_{s}x
$$
  
\n
$$
= \sqrt{\beta_{s}}x - ds\mathcal{A}_{s}(-H_{22}p - H_{12}x)] + d\mu\sqrt{\beta_{s}}[J, x]
$$
  
\n
$$
= \sqrt{\beta_{s}}x + dsH_{22}(\frac{p}{\sqrt{\beta_{s}}} - \frac{\alpha_{s}x}{\sqrt{\beta_{s}}}) + dsH_{12}\sqrt{\beta_{s}}x - d\mu\sqrt{\beta_{s}}p
$$
  
\n
$$
= \left(\sqrt{\beta_{s}} + ds\frac{-H_{22}\alpha_{s} + H_{12}\beta_{s}}{\sqrt{\beta_{s}}}\right)x + \underbrace{\left(\frac{dsH_{22}}{\sqrt{\beta_{s}}} - d\mu\sqrt{\beta_{s}}\right)p}_{\text{Must has more for}}
$$

Must be zero for the Courant-Snyder Functional

$$
\frac{d\mu}{ds} = \frac{H_{22}}{\beta_s}.
$$

 $a_{1}$ 

 $\overline{a_2^1}$ 

 $r_{12}$ 

 $m_{12}$ 

 $\mathbf{A}$   $\mathbf{P}$ 

 $(2.100)$ 

#### Why Courant-Snyder choice?

$$
z^{1}(n) = M_{01}M_{0}^{n}z^{0}(0)
$$
  
=  $A_{1}R_{01}A_{0}^{-1}A_{0}R_{0}^{n}A_{0}^{-1}z^{0}(0)$   
=  $A_{1}R_{01}R_{0}^{n}A_{0}^{-1}z^{0}(0)$   
=  $A_{1}A_{0}^{-1}\underbrace{A_{0}R_{01}R_{0}^{n}A_{0}^{-1}}_{\text{De Moivre applicable}}z^{0}(0)$  (4.14)

We can apply the result of Sec.  $(3.4.1)$  to re-express Eq.  $(4.14)$  as a function of the lattice functions at  $s = 0$ :

$$
z^{1}(n) = A_{1}A_{0}^{-1} \begin{pmatrix} \cos\left(\Phi_{n}^{c}\right) + \alpha_{0} \sin\left(\Phi_{n}^{c}\right) & \beta_{0} \sin\left(\Phi_{n}^{c}\right) \\ -\gamma_{0} \sin\left(\Phi_{n}^{c}\right) & \cos\left(\Phi_{n}^{c}\right) - \alpha_{0} \sin\left(\Phi_{n}^{c}\right) \end{pmatrix} z^{0}
$$
\n
$$
= \begin{pmatrix} \sqrt{\frac{\beta_{1}}{\beta_{0}}} & 0 \\ -\frac{(\alpha_{1}-\alpha_{2})}{\sqrt{\beta_{1}\beta_{0}}} & \sqrt{\frac{\beta_{0}}{\beta_{1}}} \end{pmatrix} \begin{pmatrix} \cos\left(\Phi_{n}^{c}\right) + \alpha_{0} \sin\left(\Phi_{n}^{c}\right) & \beta_{0} \sin\left(\Phi_{n}^{c}\right) \\ -\gamma_{0} \sin\left(\Phi_{n}^{c}\right) & \cos\left(\Phi_{n}^{c}\right) - \alpha_{0} \sin\left(\Phi_{n}^{c}\right) \end{pmatrix} z^{0}
$$
\nwhere  $\Phi_{n}^{c} = n\mu + \mu_{01}^{c}$  here  $\mu_{01}^{c}$  results from Eq. (4.12)

The final is step is simply to write the position q at  $s = 1$ :

$$
q^{1}(n) = \sqrt{\frac{\beta_{1}}{\beta_{0}}} \left( \left( \cos \left( \Phi_{n}^{c} \right) + \alpha_{0} \sin \left( \Phi_{n}^{c} \right) \right) q_{0} + \beta_{0} \sin \left( \Phi_{n}^{c} \right) p_{0} \right) (4.16)
$$

Extends to coupled system if  $A_{12}=0$  and  $A_{34}=0$