

# Detector technologies

A brief overview

Many thanks to Erik Butz, [Simon Spannagel](#), [Freya Blekman](#), [Peter Schleper](#), Erika Garutti

[jory.sonneveld@cern.ch](mailto:jory.sonneveld@cern.ch)

# Particles from outer space

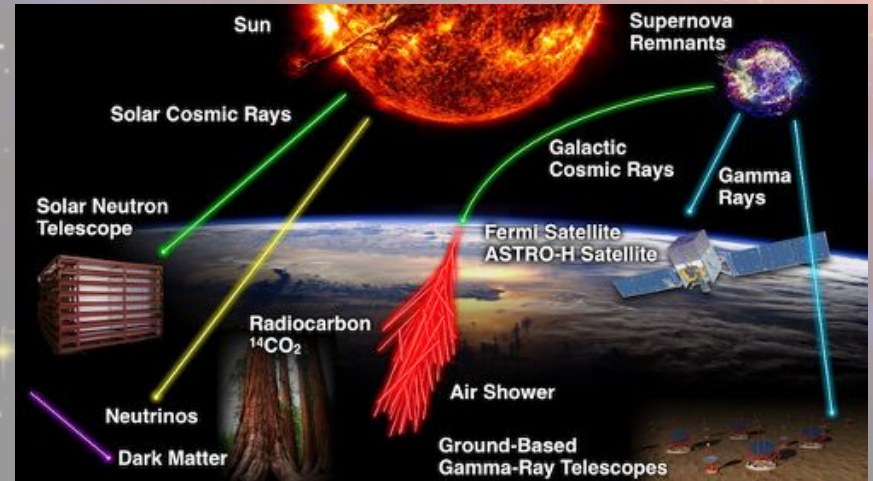
10000 times a second you have particles from cosmic rays passing through you



**HESS:** high energy stereoscopic system, in Namibia, can detect gamma rays

[http://www2.cors.fr/sites/en/image/hess\\_new\\_large\\_hd.jpg](http://www2.cors.fr/sites/en/image/hess_new_large_hd.jpg)

What are these particles and how do they behave?  
What are we and what is the universe made of?



# How to detect a particle?

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# $\gamma$ rays

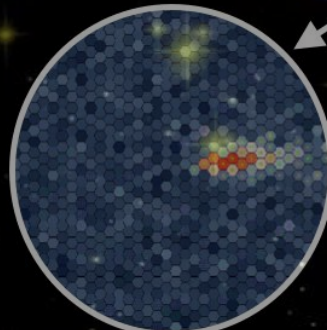
$\gamma$ -ray enters the atmosphere

Electromagnetic cascade



## Cherenkov telescope:

- light is 0.03 % slower in air
- ultra-high energy particles can travel faster than light in air
- then a blue flash of “Cherenkov light” is created
- similar to the sonic boom created by an aircraft exceeding the speed of sound



10 nanosecond snapshot



0.1 km<sup>2</sup> “light pool”, a few photons per m<sup>2</sup>.

# Discovery of antimatter

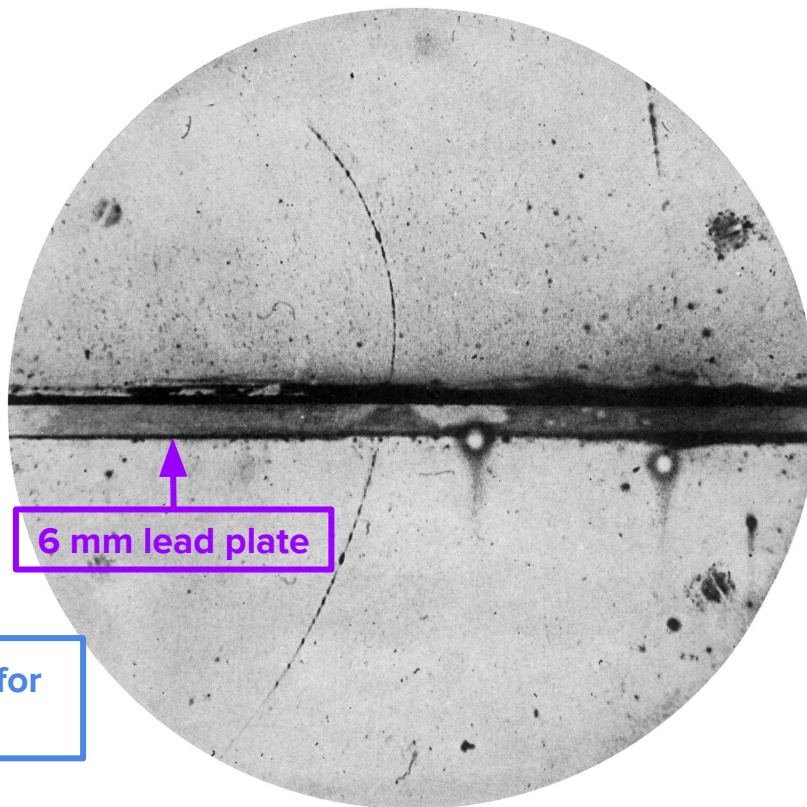
<https://upload.wikimedia.org/wikipedia/commons/6/69/PositronDiscovery.jpg>

C.D. Anderson <https://journals.aps.org/pr/pdf/10.1103/PhysRev.43.491>

The first positron ever observed!

**Wilson cloud chamber:** gaseous mixture of supersaturated water or alcohol. Energetic particle ionizes gas and ions form condensation centers visible as a 'cloud'.

15000 Gauss = 1.5T magnetic field Wilson chamber for detecting cosmic rays

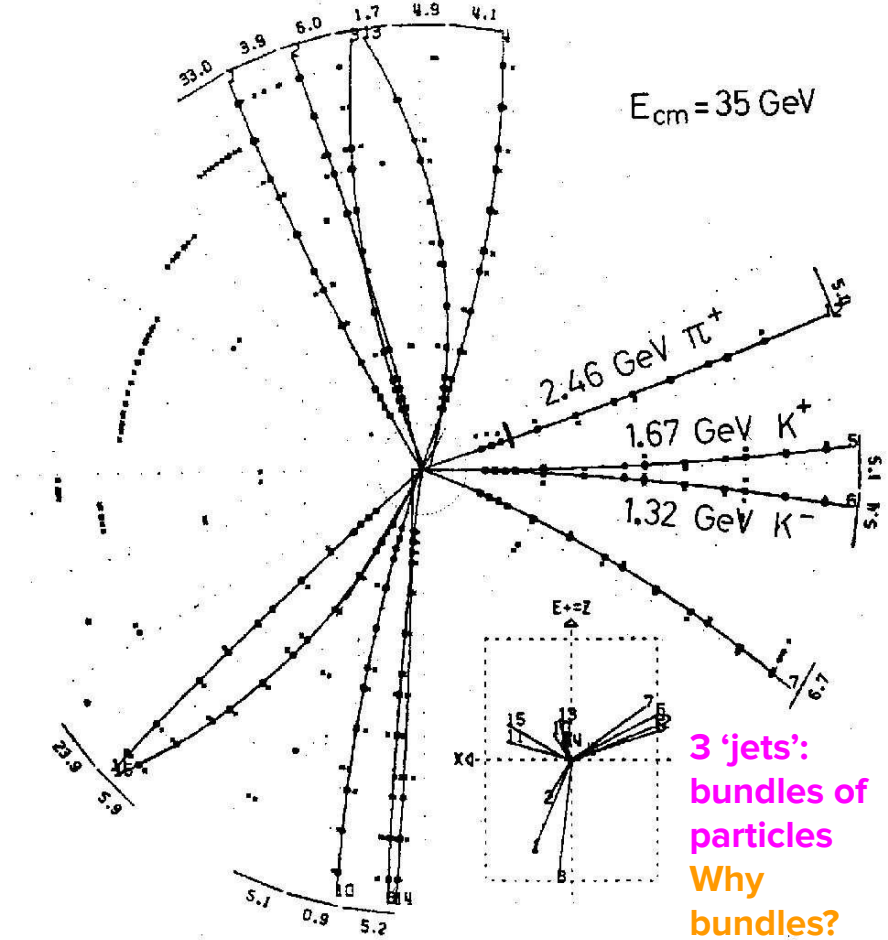
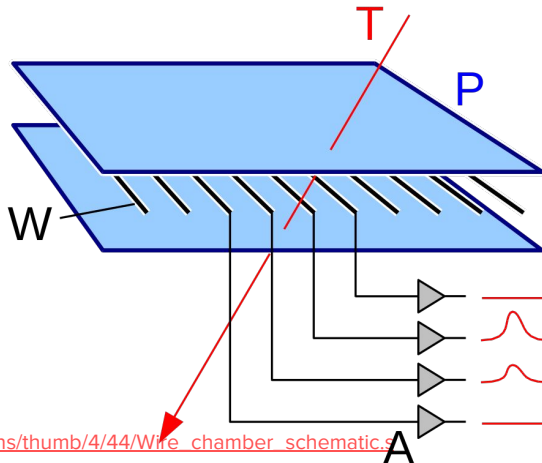


# Gluon discovery

Event in *drift chamber* of JADE experiment at PETRA collider at DESY.

Such events were used to prove the existence of gluons:  $e^+e^- \rightarrow qq\bar{g}$ .

**Wire chamber:** particle T passes through grounded plate P and ionizes gas in chamber. Charge drifts in electric field to high voltage wires W and is collected at an amplifier A.



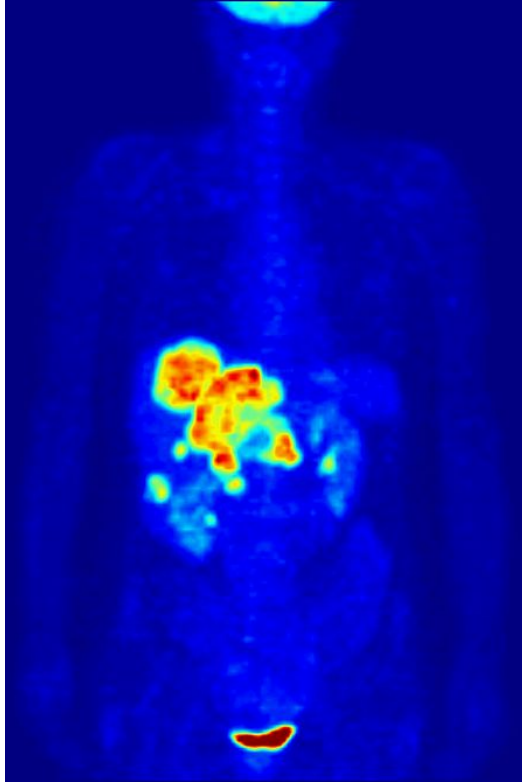
From

[http://www.desy.de/sites2009/site\\_www-desy/content/e409/e287332/e287337/e287345/1980-09-22\\_TASSO-Event\\_Gluon\\_Entdeckung\\_sw\\_ger.jpg](http://www.desy.de/sites2009/site_www-desy/content/e409/e287332/e287337/e287345/1980-09-22_TASSO-Event_Gluon_Entdeckung_sw_ger.jpg)<sup>6</sup>

22.9.80

[https://www.desy.de/~schlepèr/lehre/physik5/WS\\_2018\\_19/Physik\\_5\\_72-95.pdf](https://www.desy.de/~schlepèr/lehre/physik5/WS_2018_19/Physik_5_72-95.pdf)

# More detectors



Magnetic resonance imaging

(b)

Positron emission tomography



# CERN and the Large Hadron Collider

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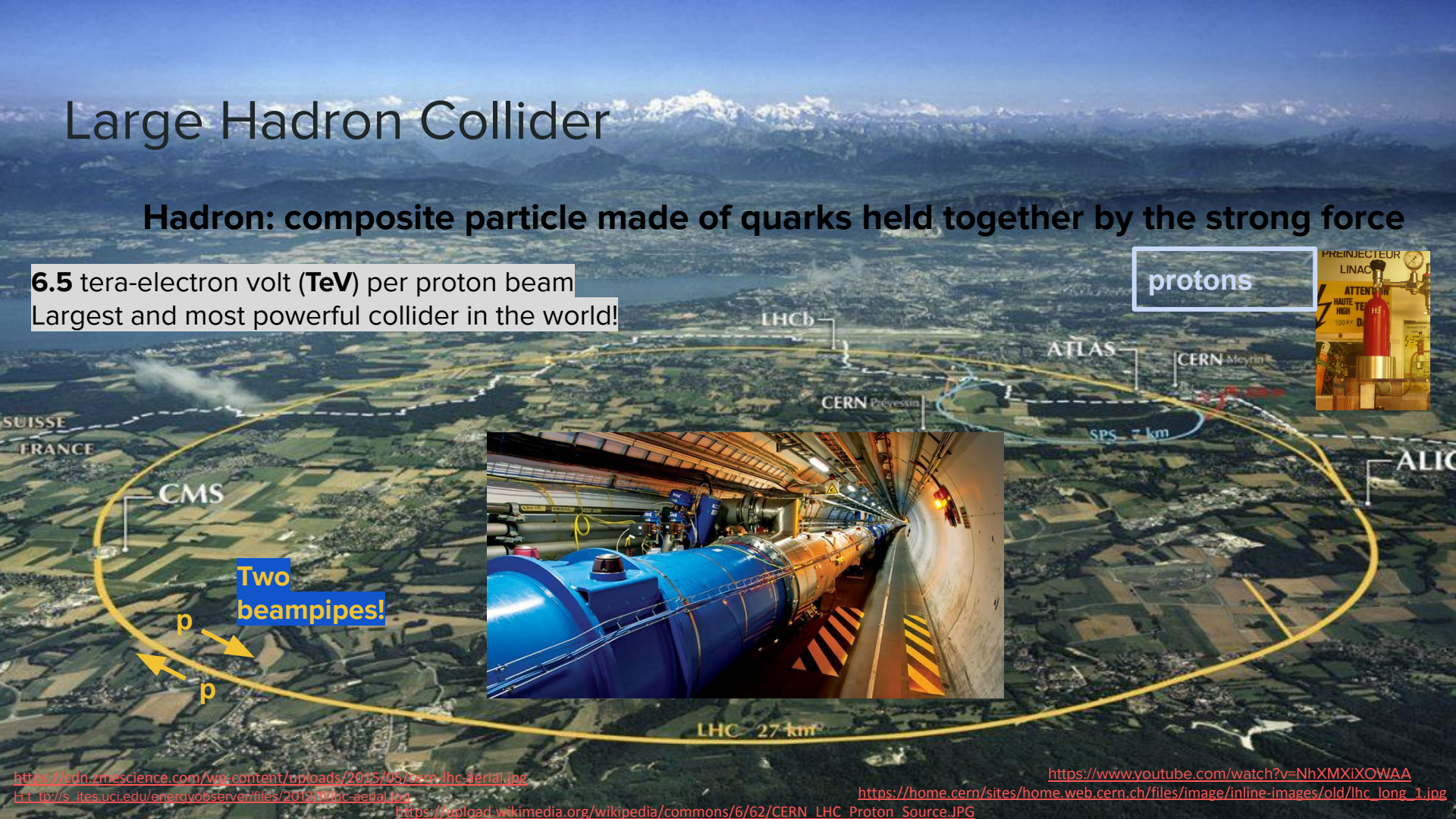


# Large Hadron Collider

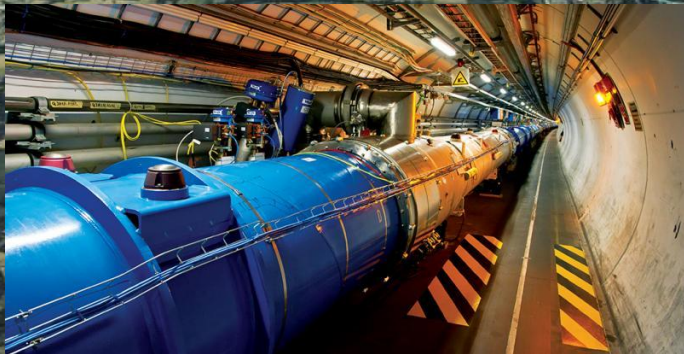
**Hadron: composite particle made of quarks held together by the strong force**

**6.5** tera-electron volt (**TeV**) per proton beam  
Largest and most powerful collider in the world!

protons



Two beampipes!



[https://cdn.zmescience.com/wp-content/uploads/2015/05/cern\\_lhc-aerial.jpg](https://cdn.zmescience.com/wp-content/uploads/2015/05/cern_lhc-aerial.jpg)

<http://sites.uci.edu/energyobserver/files/2012/10/lhc-aerial.jpg>

[https://upload.wikimedia.org/wikipedia/commons/6/62/CERN\\_LHC\\_Proton\\_Source.JPG](https://upload.wikimedia.org/wikipedia/commons/6/62/CERN_LHC_Proton_Source.JPG)

<https://www.youtube.com/watch?v=NhXMXiXOWAA>

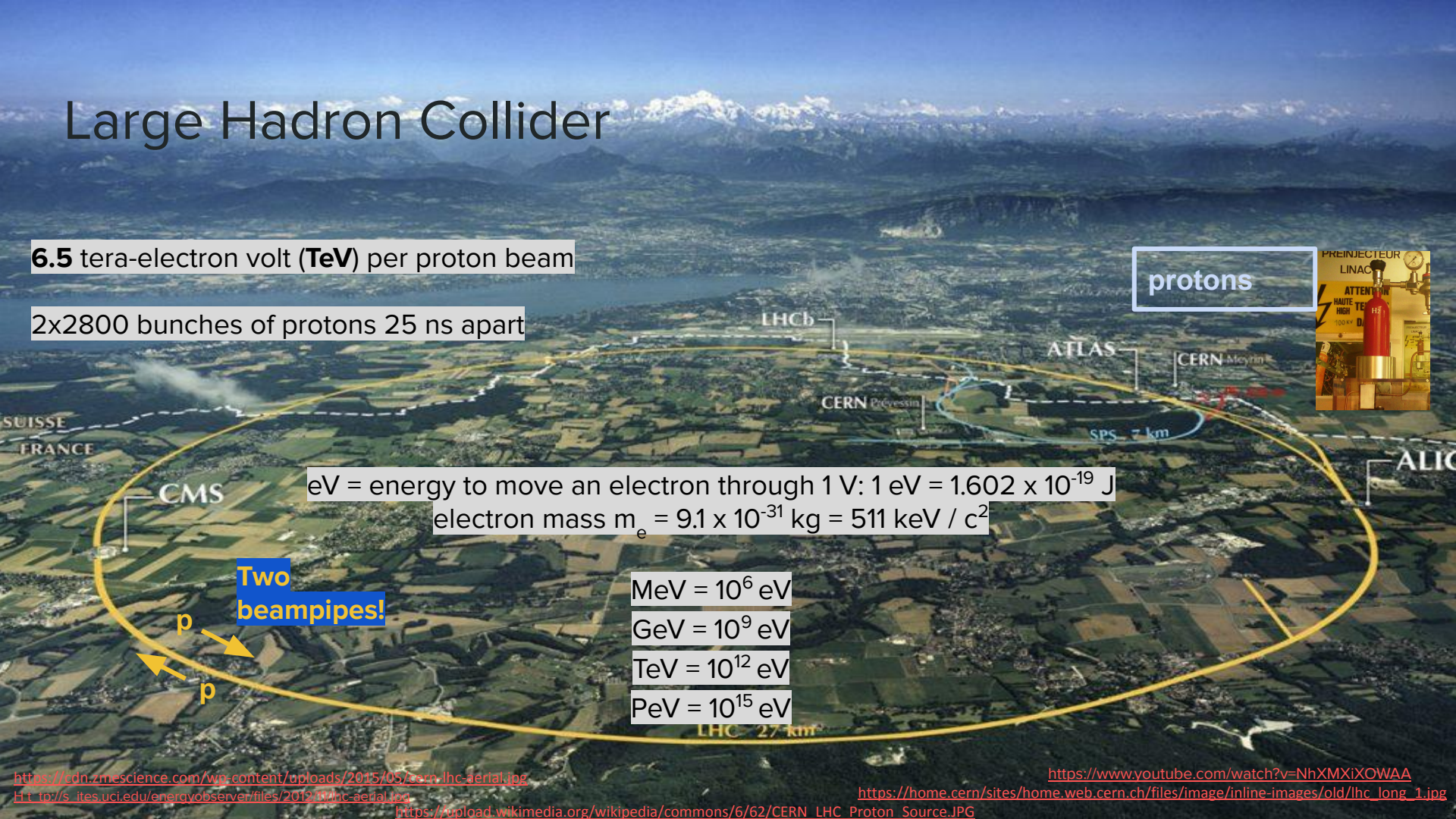
[https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc\\_long\\_1.jpg](https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc_long_1.jpg)

# Large Hadron Collider

6.5 tera-electron volt (TeV) per proton beam

2x2800 bunches of protons 25 ns apart

protons



eV = energy to move an electron through 1 V:  $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$   
electron mass  $m_e = 9.1 \times 10^{-31} \text{ kg} = 511 \text{ keV} / c^2$

Two beampipes!



- MeV =  $10^6 \text{ eV}$
- GeV =  $10^9 \text{ eV}$
- TeV =  $10^{12} \text{ eV}$
- PeV =  $10^{15} \text{ eV}$

[https://cdn.zmescience.com/wp-content/uploads/2015/05/CERN\\_LHC-aerial.jpg](https://cdn.zmescience.com/wp-content/uploads/2015/05/CERN_LHC-aerial.jpg)

<http://sites.uci.edu/energyobserver/files/2012/11/lhc-aerial.jpg>

[https://upload.wikimedia.org/wikipedia/commons/6/62/CERN\\_LHC\\_Proton\\_Source.JPG](https://upload.wikimedia.org/wikipedia/commons/6/62/CERN_LHC_Proton_Source.JPG)

<https://www.youtube.com/watch?v=NhXMXiXOWAA>

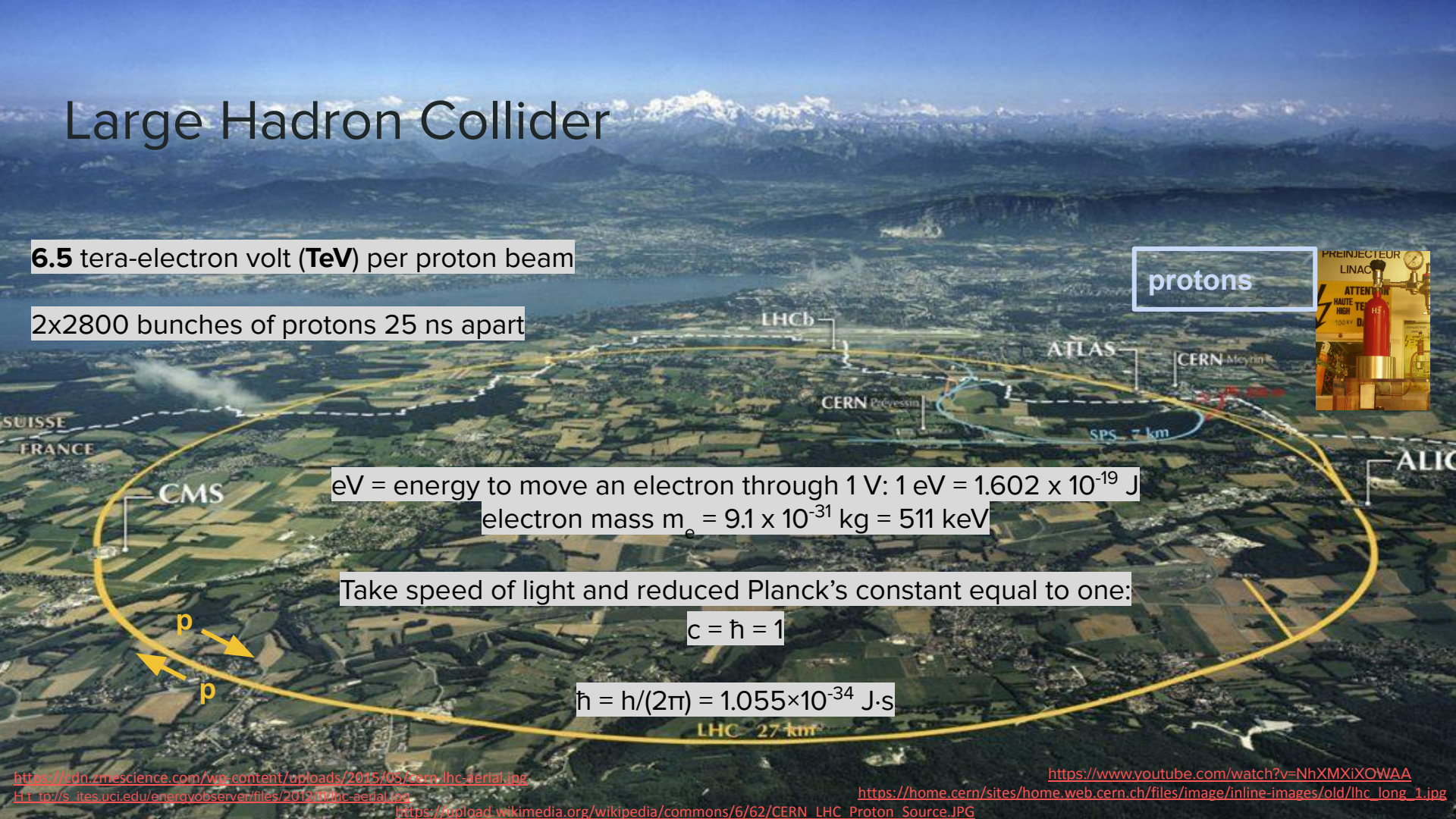
[https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc\\_long\\_1.jpg](https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc_long_1.jpg)

# Large Hadron Collider

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Take speed of light and reduced Planck's constant equal to one:

$$c = \hbar = 1$$

$$\hbar = h/(2\pi) = 1.055 \times 10^{-34} \text{ J}\cdot\text{s}$$

# Detectors at the LHC



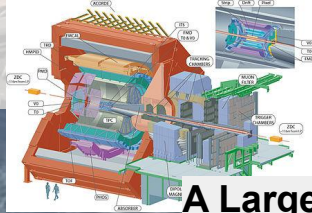
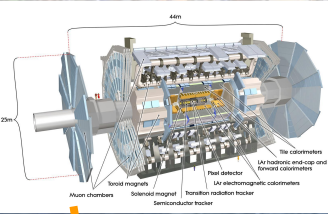
ALICE



## Compact Muon Solenoid

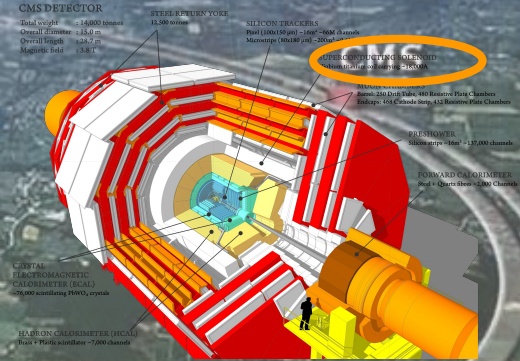
14000 tons: 1.5\* Eiffel tower weight, half the size of ATLAS: 15 m x 15 m x 21 m very compact!

Largest superconducting solenoid magnet ever made



## A Large Ion Collider

**Experiment:** specialized in heavy ion collisions and quark-gluon plasma: fraction of second after big bang!



## LHC beauty:

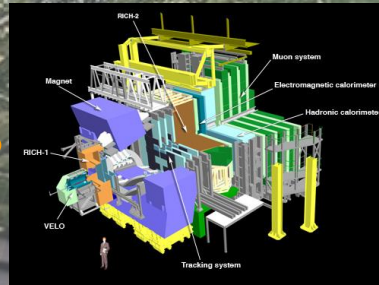
A single-arm forward spectrometer designed for the study of particles containing b or c quarks.

**Other detectors:** MoEDAL, TOTEM, LHCforward

## A Toroidal LHC

### Apparatus:

25 m x 25 m x 46m  
The inner detector has 3 air core toroidal magnets and one solenoidal magnet.

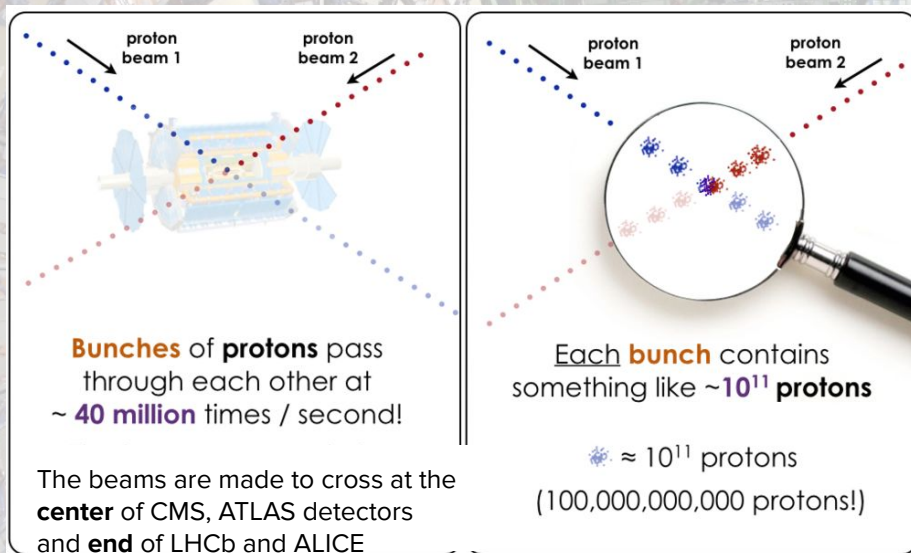


ALICE

ATLAS

LHCb

# Proton-proton collisions

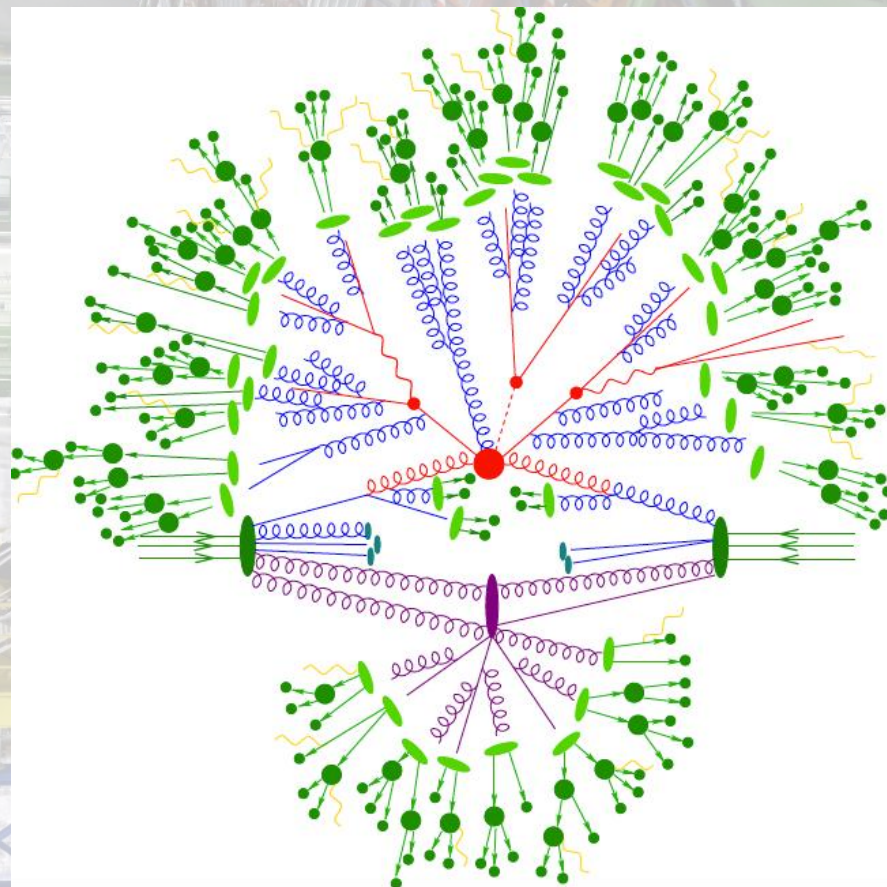


## Trigger system:

Choose what events are interesting

**How do we choose?**

**Jet:** quarks and gluons hadronized to kaons, protons, pions in a collimated stream



<https://sciencenode.org/feature/sherpa-and-open-science-grid-predicting-emergence-jets.php>

<http://wlcg-public.web.cern.ch/sites/wlcg-public.web.cern.ch/files/WLCG-snapshot-28112013.jpg>

What do we detect?

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# Not all known elementary particles

Directly detect:

Decay products

jets

Indirectly detect:

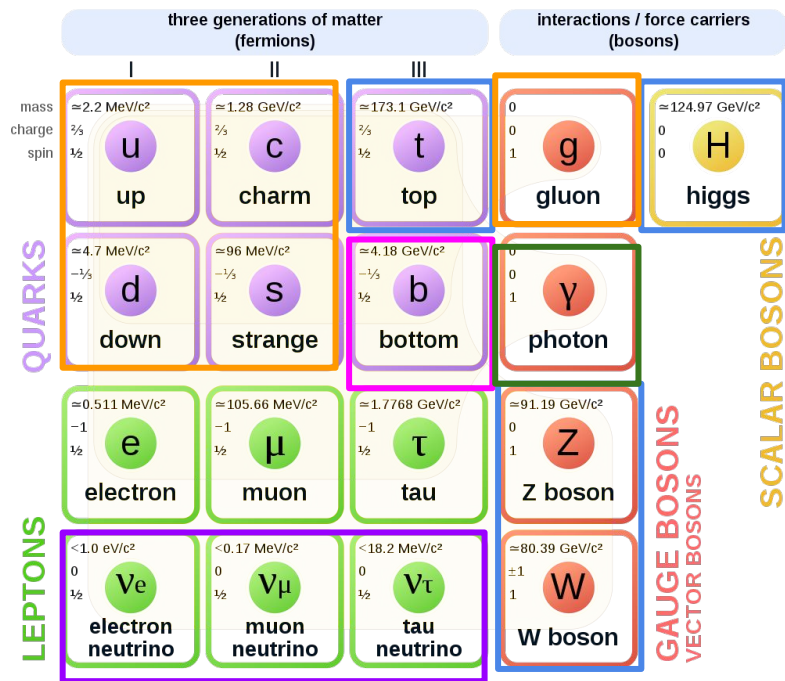
Missing energy

Secondary vertex + jets

Neutral particles

Should be able to detect and identify:  
 $e^\pm, \mu^\pm, \gamma, \pi^\pm, K^\pm, p^\pm, K^0, n$   
 using mass, charge, interaction

## Standard Model of Elementary Particles



From [https://upload.wikimedia.org/wikipedia/commons/0/00/Standard\\_Model\\_of\\_Elementary\\_Particles.svg](https://upload.wikimedia.org/wikipedia/commons/0/00/Standard_Model_of_Elementary_Particles.svg)

# What do we measure and how?

Observable	Measurable quantity
Momentum (p)	Bending radius in magnetic field
Speed (v)	Time of flight, Cherenkov radiation
Charge (Q)	Bending in magnetic field
Lifetime ( $\tau$ )	Distance traveled before decay
Energy (E)	Absorption in calorimeters
Mass (m)	Indirectly from momentum
Spin	<u><a href="#">Angular distributions</a></u>

- $d = c\tau\gamma$
- $\gamma = 1/\sqrt{1-\beta^2}$
- $\beta = v/c$
- $E^2 = m^2c^4 + p^2c^2$
- $p = \gamma mv = mv/\sqrt{1-v^2/c^2}$

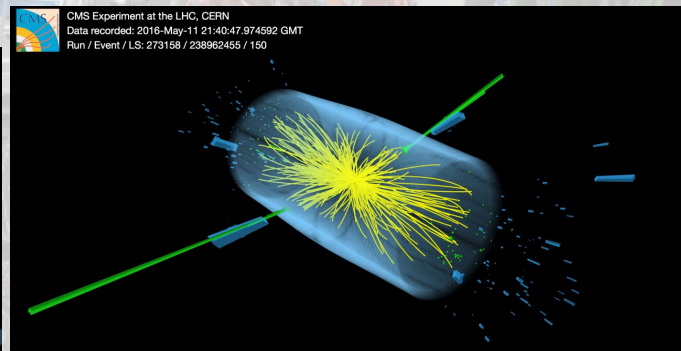
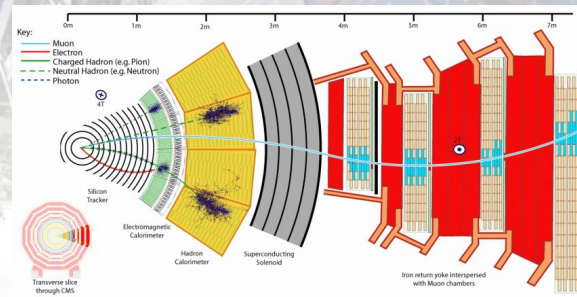
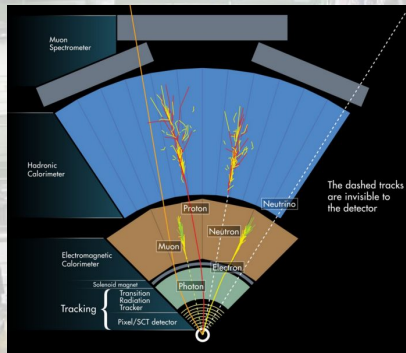
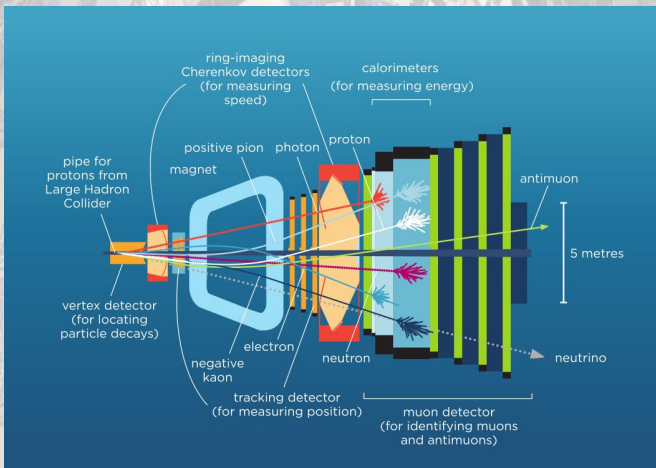
For some examples of measuring spin see

<https://arxiv.org/pdf/1202.6660.pdf> and  
<http://moriond.in2p3.fr/QCD/2013/proceedings/Muehleleitner.pdf>

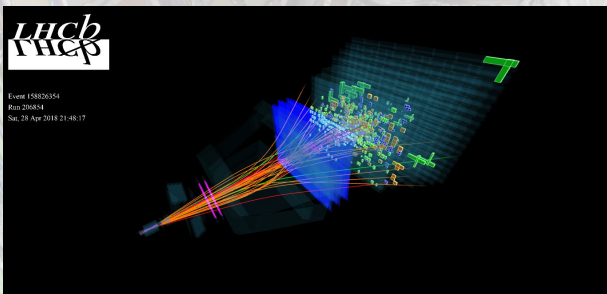
**Need 1) a magnetic field and 2) interaction with material**



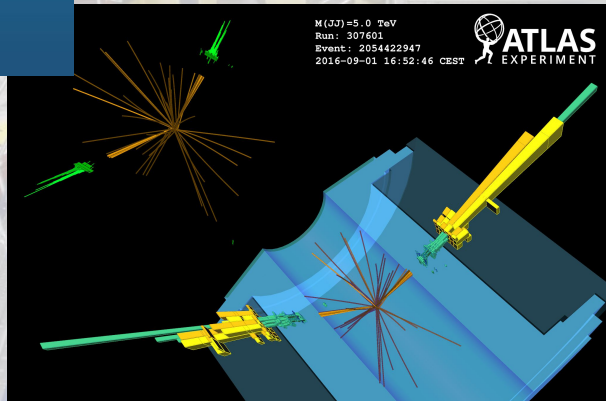
# Detecting particles at the LHC



CMS dijet event

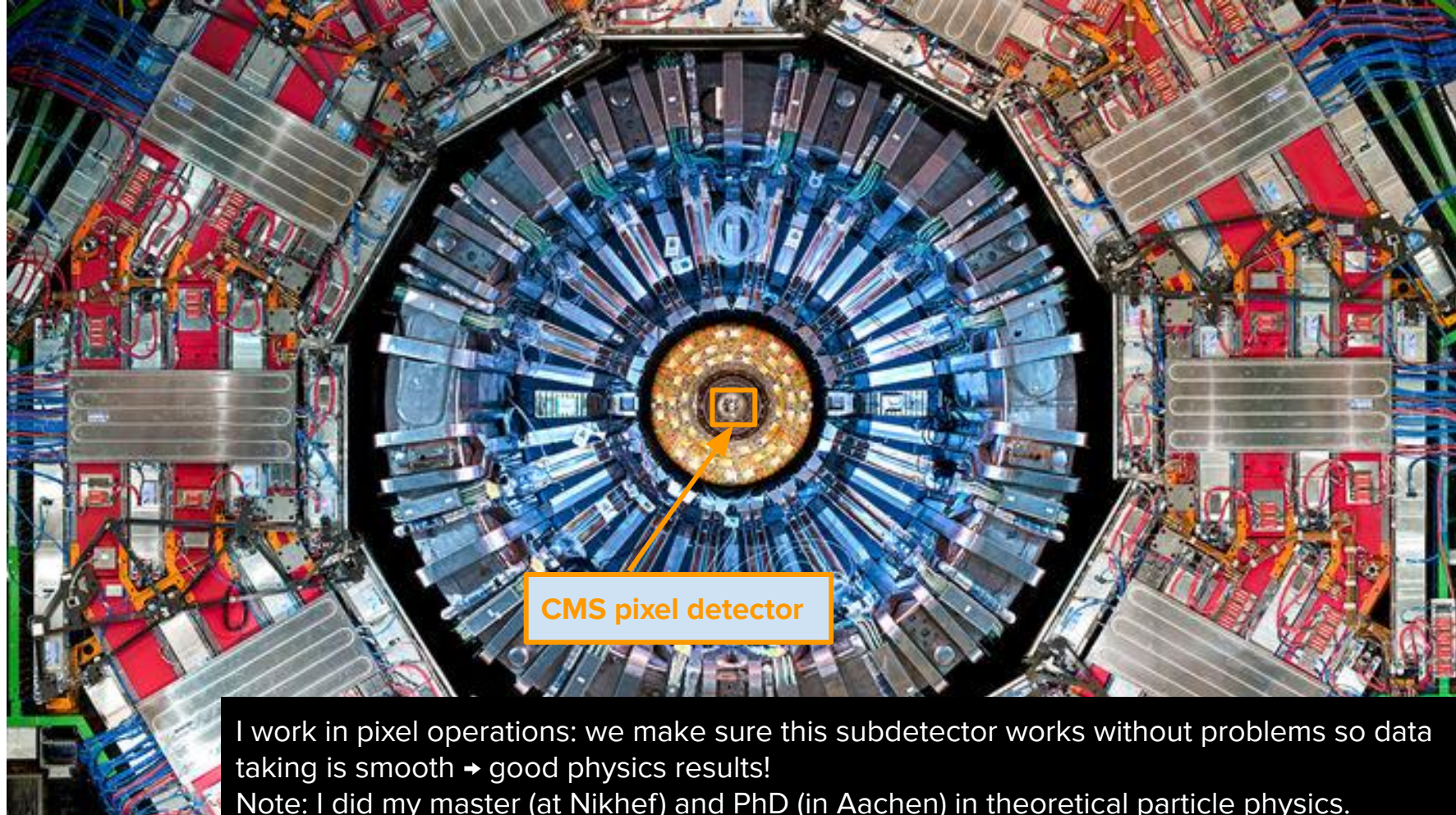


LHCb b-jet event



ATLAS dijet event

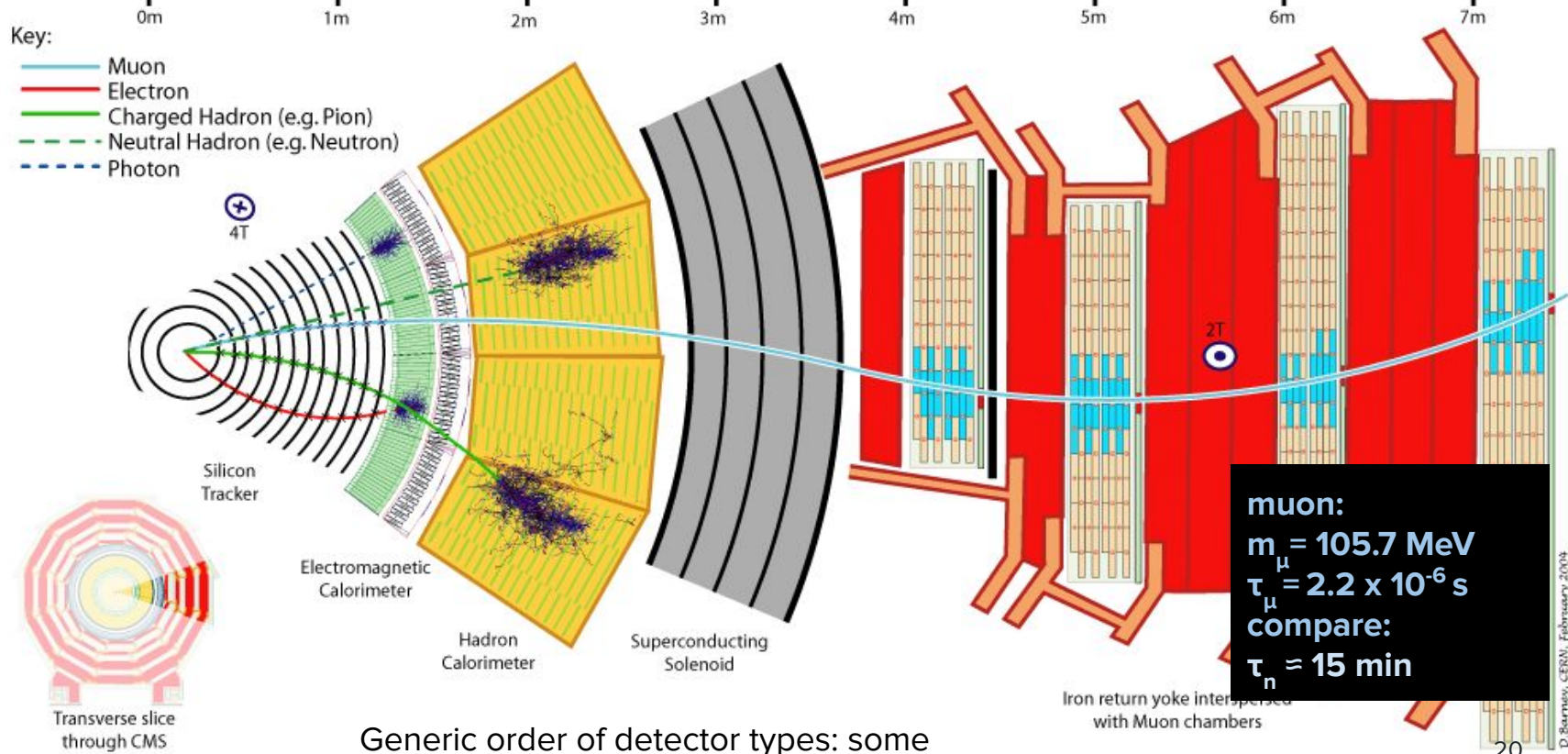




CMS pixel detector

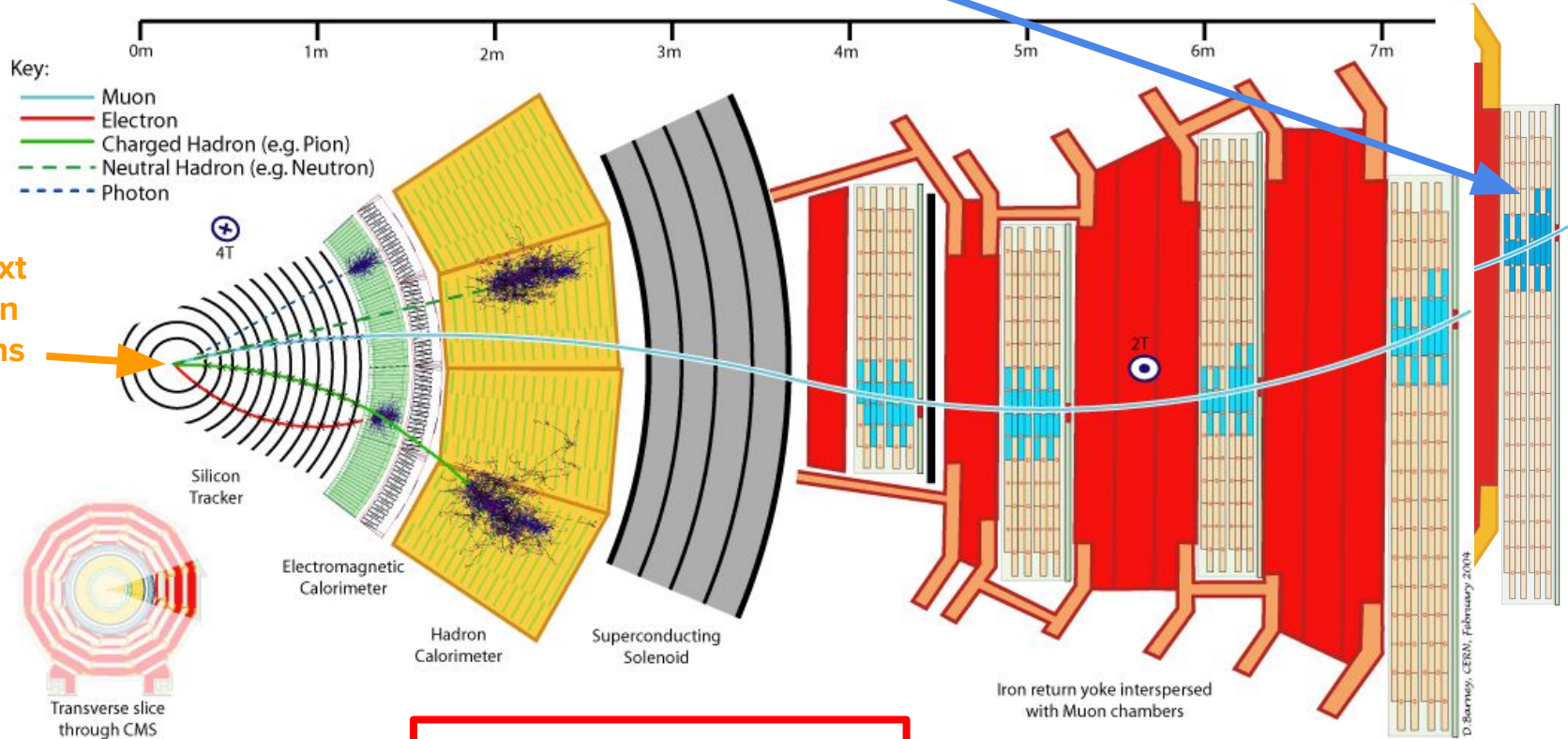
I work in pixel operations: we make sure this subdetector works without problems so data taking is smooth → good physics results!  
Note: I did my master (at Nikhef) and PhD (in Aachen) in theoretical particle physics.

# Detectors at the large hadron collider: onion-like



Generic order of detector types: some measurements destructive!

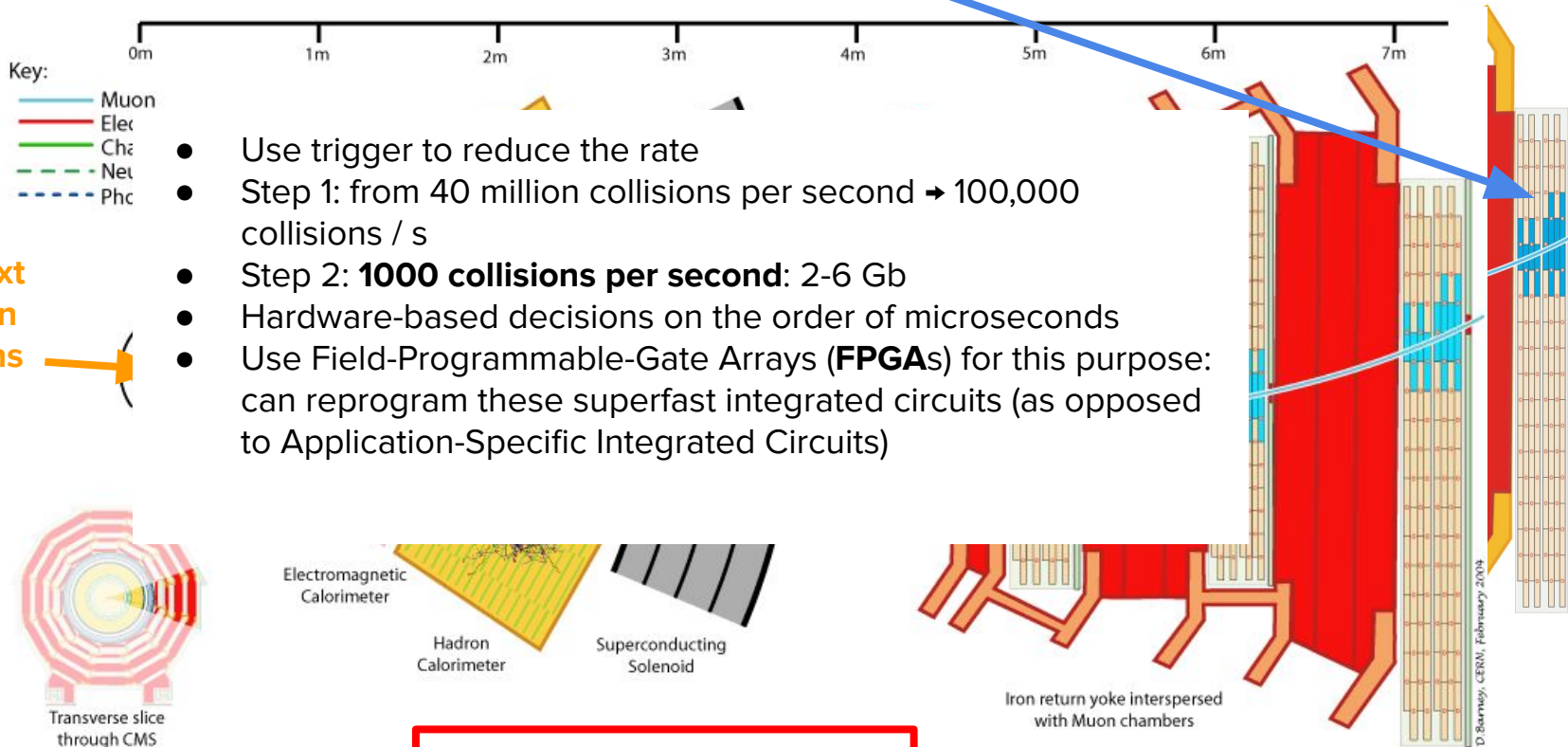
# Note when the muon arrives here



The next collision happens here:

$$25 \text{ ns} \cdot c \approx 7.5 \text{ m}$$

# Note when the muon arrives here

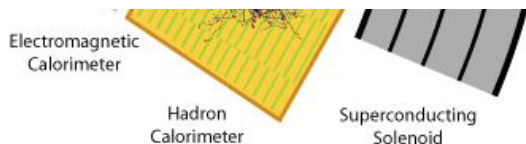


Key:

- Muon
- Elec
- Ch̄
- - - Net
- - - Phc

- Use trigger to reduce the rate
- Step 1: from 40 million collisions per second → 100,000 collisions / s
- Step 2: **1000 collisions per second**: 2-6 Gb
- Hardware-based decisions on the order of microseconds
- Use Field-Programmable-Gate Arrays (**FPGAs**) for this purpose: can reprogram these superfast integrated circuits (as opposed to Application-Specific Integrated Circuits)

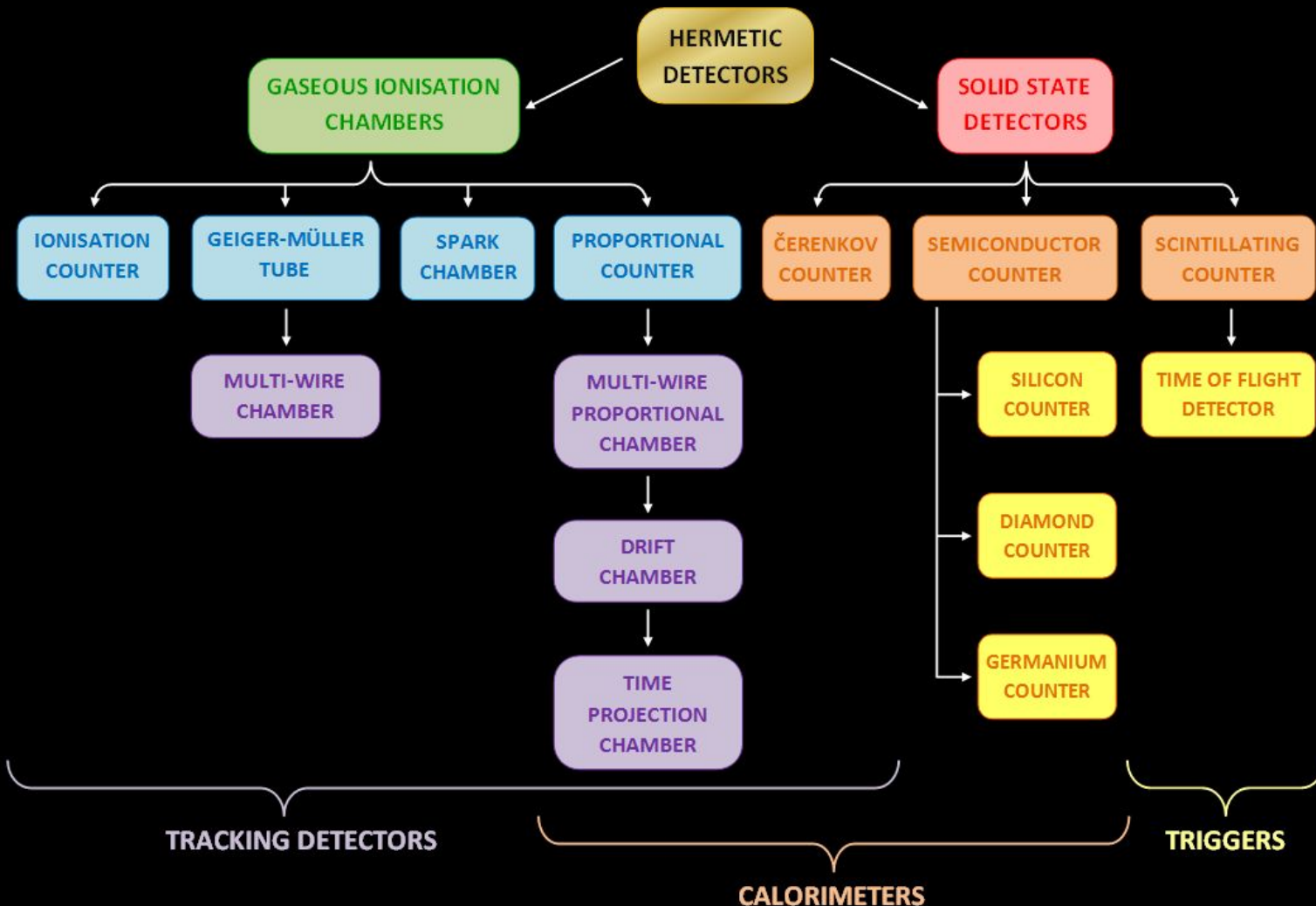
The next collision happens here:



$$25 \text{ ns} \cdot c \approx 7.5 \text{ m}$$

Iron return yoke interspersed with Muon chambers

D. Barney, CERN, February 2013

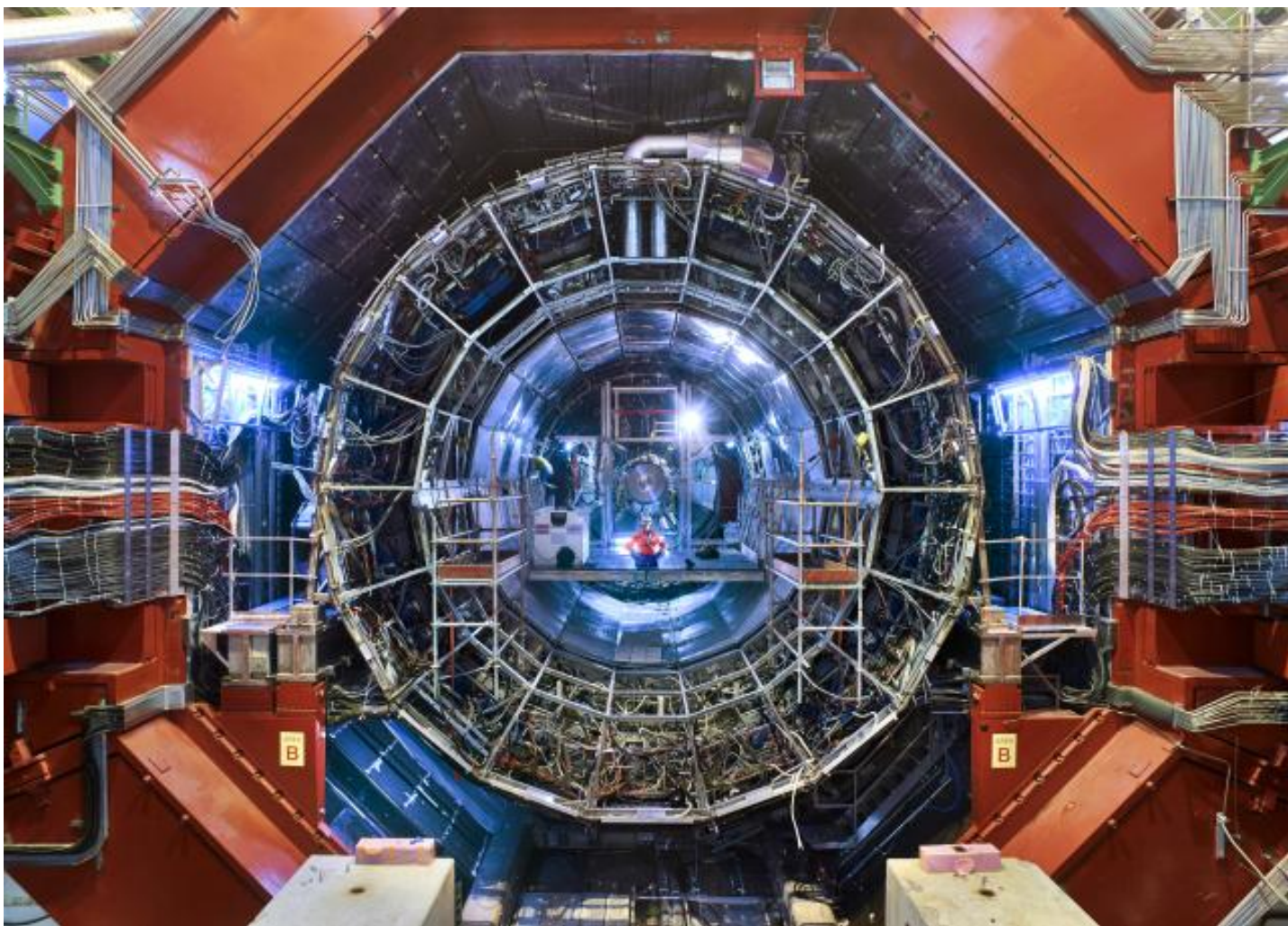


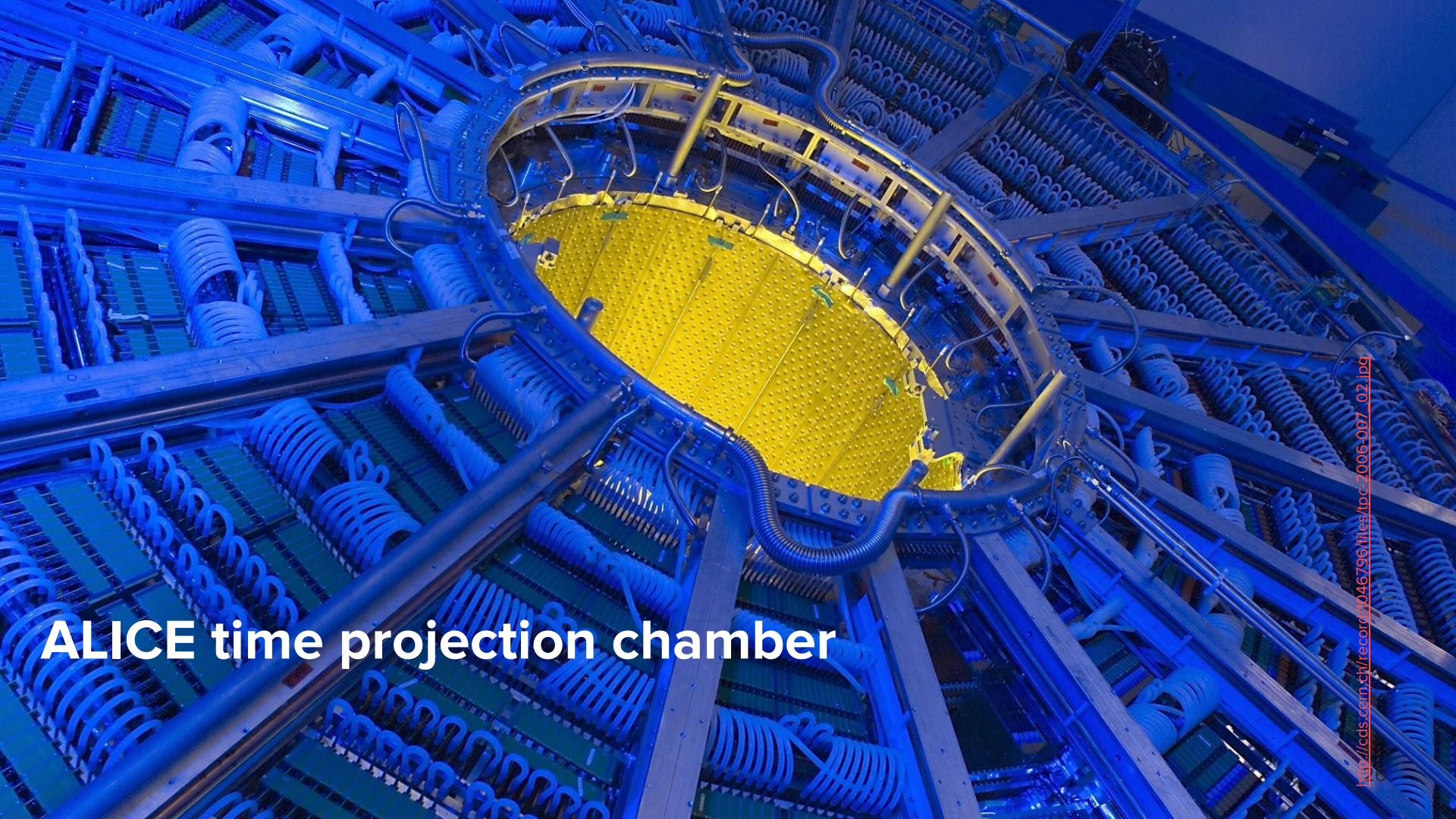
[https://upload.wikimedia.org/wikipedia/commons/c/c0/Detectors\\_summary\\_3.png](https://upload.wikimedia.org/wikipedia/commons/c/c0/Detectors_summary_3.png)

# Gaseous detectors

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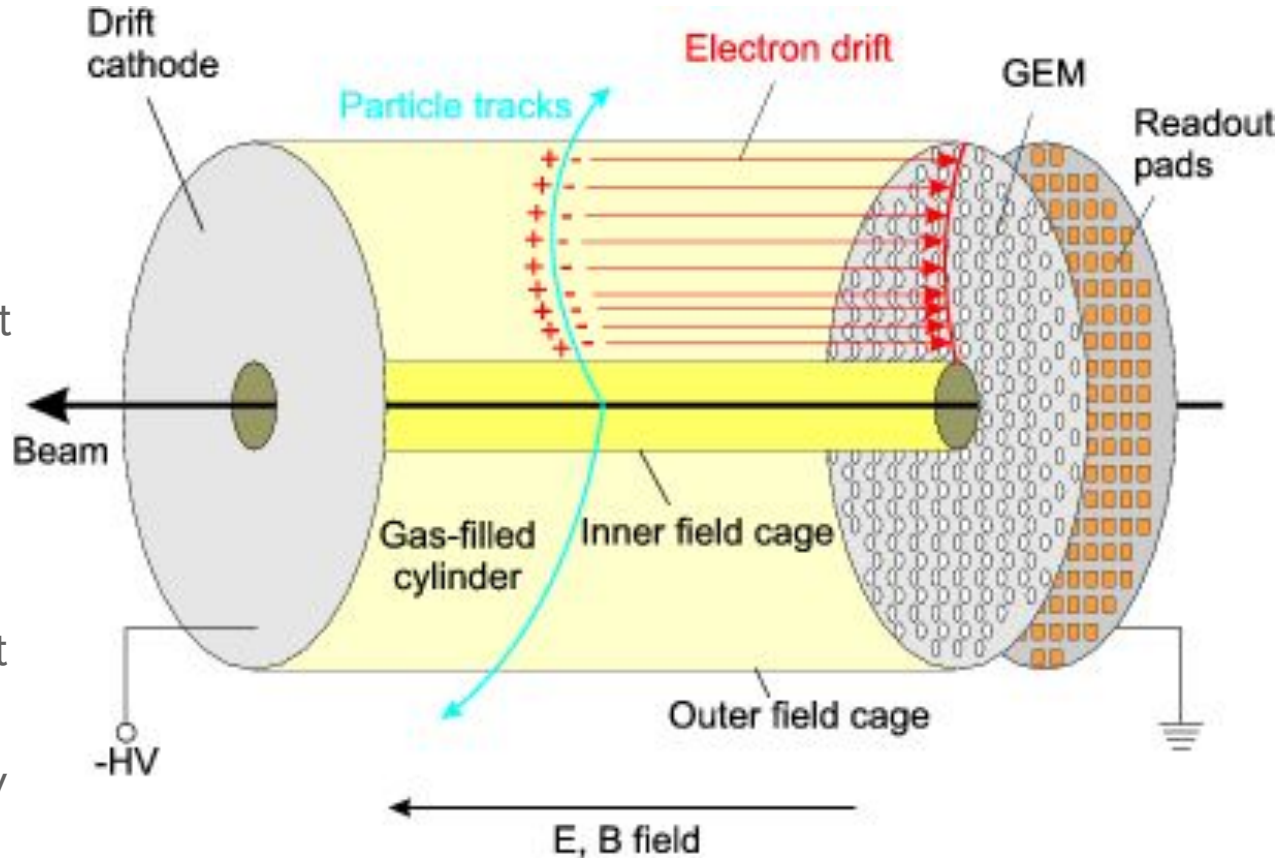




**ALICE time projection chamber**

# TPC

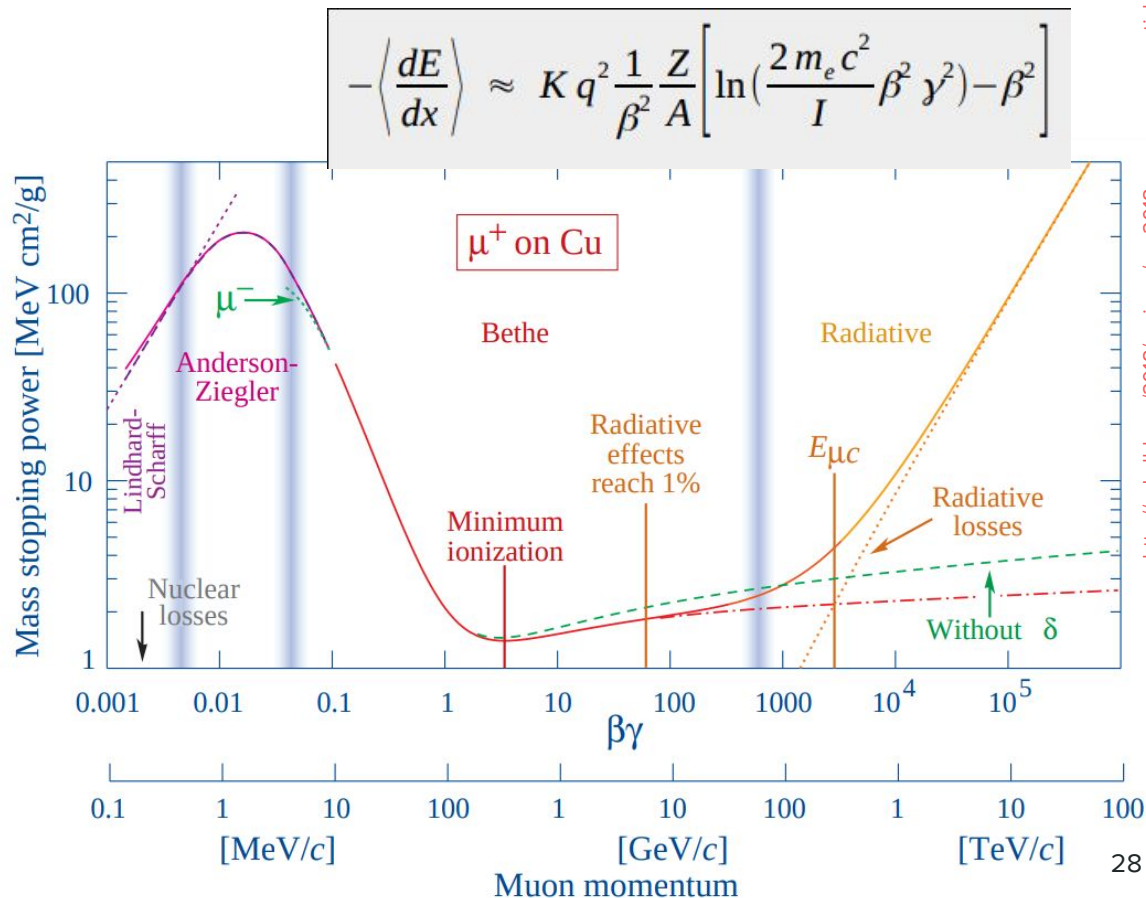
1. Ionization of gas in **chamber** with electric field causes electron drift
2. Signal gets amplified, in this case by gas electron multipliers  $\rightarrow$  electron avalanche
3. Readout pads can detect signal that can be **projected** onto trajectory
4. z (along beam) information from **timing**



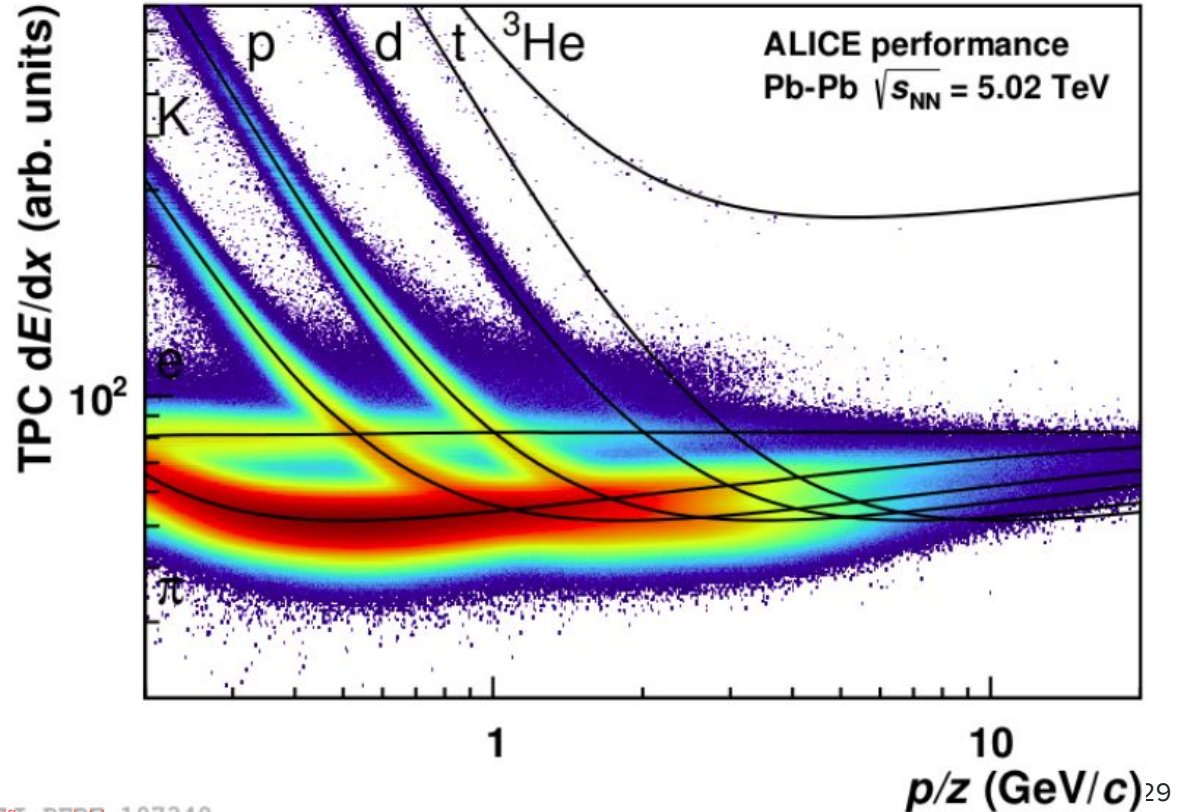
# Ionization loss

- Can measure ionization loss  $dE/dx$
- $K$  is a coefficient:  
 $K = .307 \text{ MeV mol}^{-1}\text{cm}^{-2}$
- $I$  is the mean excitation energy

Depends on charge, atom number, ionization energy, density



- Every point is one measurement!
- Can identify particles for low momenta
- For higher momenta, all particles behave like a minimum ionizing particle (MIP)



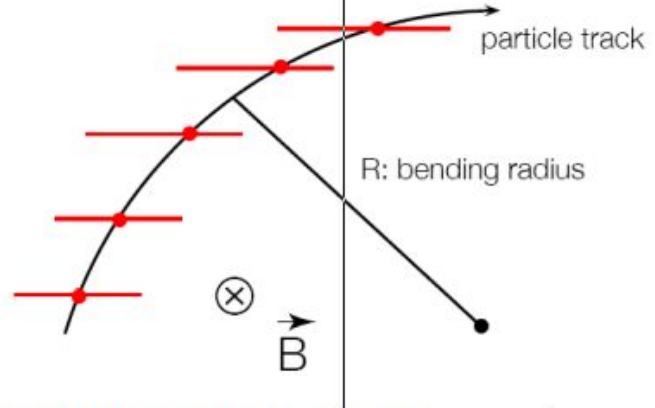
# Tracking detectors

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Momentum determination  
in a cylindrical drift chamber ...

$$\frac{mv^2}{R} = evB \quad \rightarrow \quad p = eB \cdot R$$

$$p \left[ \frac{\text{GeV}}{c} \right] = 0.3 \text{ B[T]} R[\text{m}]$$



momentum component perpendicular to the B-field  
transverse momentum  $p_t$

For Sagitta  $s$ :

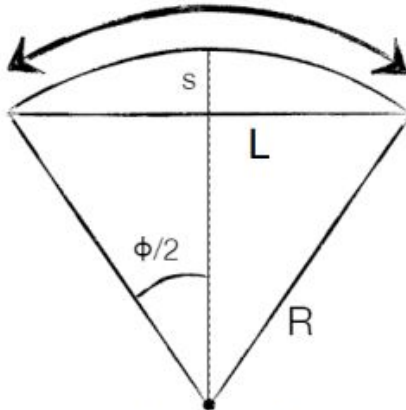
$$s = R - R \cos \frac{\phi}{2} \approx R \frac{\phi^2}{8}$$

$$s = R \frac{L^2}{8R^2} = \frac{L^2}{8R} \quad \text{and} \quad R = \frac{L^2}{8s}$$

$$\rightarrow \frac{\Delta p}{p} = \frac{\Delta R}{R} = \frac{L^2}{8Rs} \cdot \frac{\Delta s}{s}$$

$$\text{with } \phi = \frac{L}{R}$$

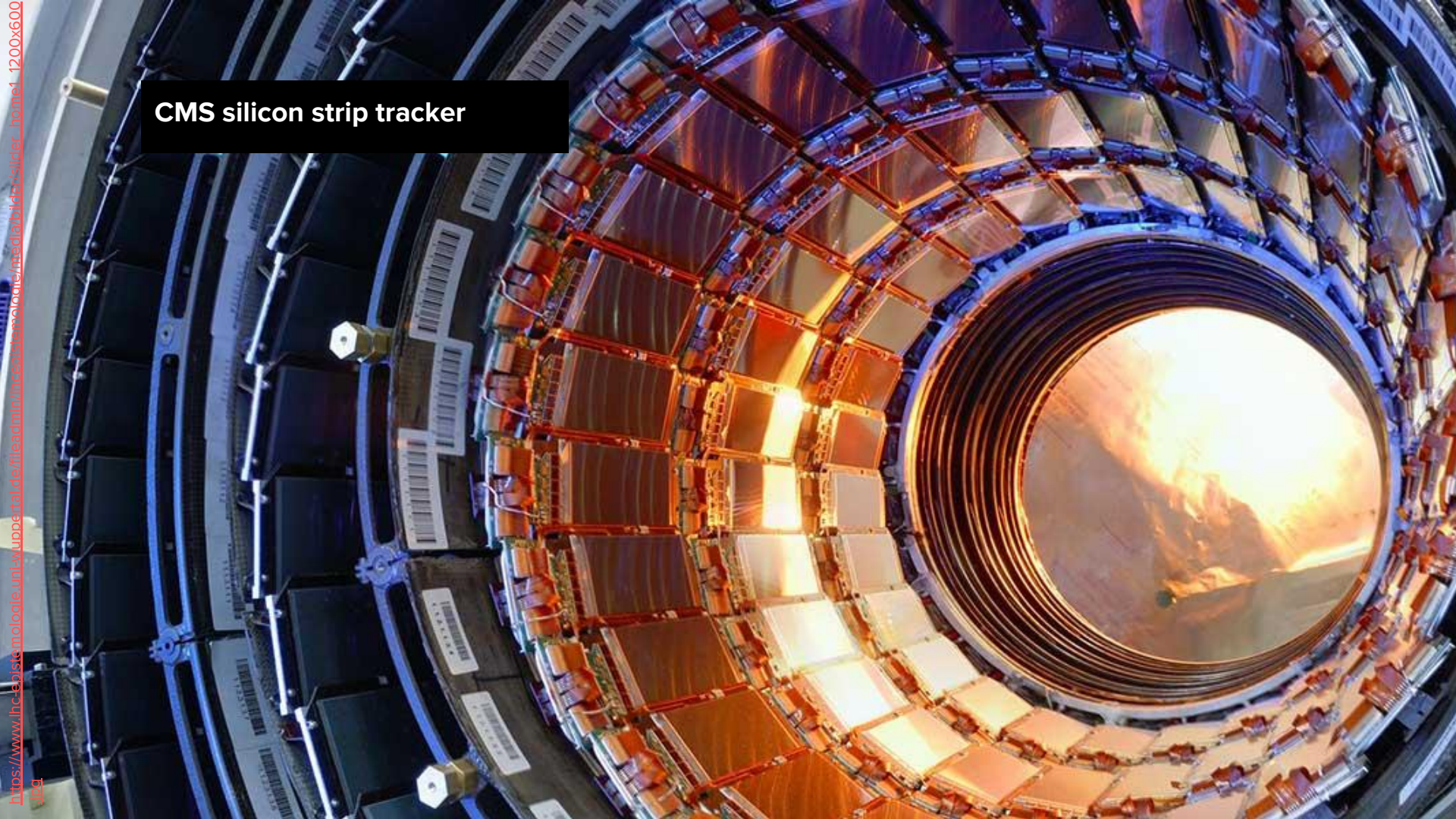
→ radius is obtained by:  
circle fit through  
measurement points along  
the track with point  
resolution  $\sigma_{\text{fp}}$



$$\begin{aligned} 1\text{T} &= 1\text{kg/C/s}^2 \\ 1\text{eV}/c &= 0.535 \cdot 10^{-27} \text{ kg m/s} \\ 1\text{eV} &= 1.6 \cdot 10^{-19} \text{ C} \end{aligned}$$

Magnet is  
important!

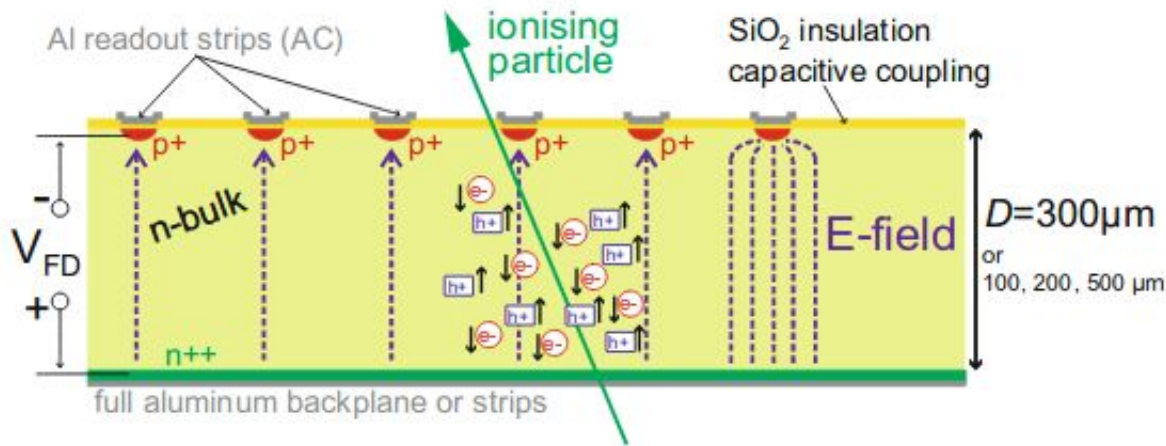
CMS silicon strip tracker





# Ideal signal detection with silicon sensors

- A minimum ionizing particle (MIP) traveling through a fully depleted region ( $V_{FD}$ ) creates electron hole pairs
- The charges drift to opposite directions under the electric field
- Within nanoseconds, charges are collected at the readout



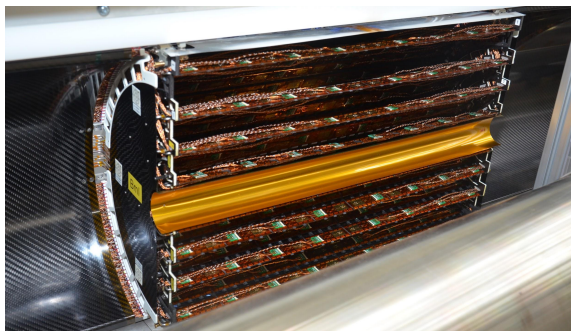
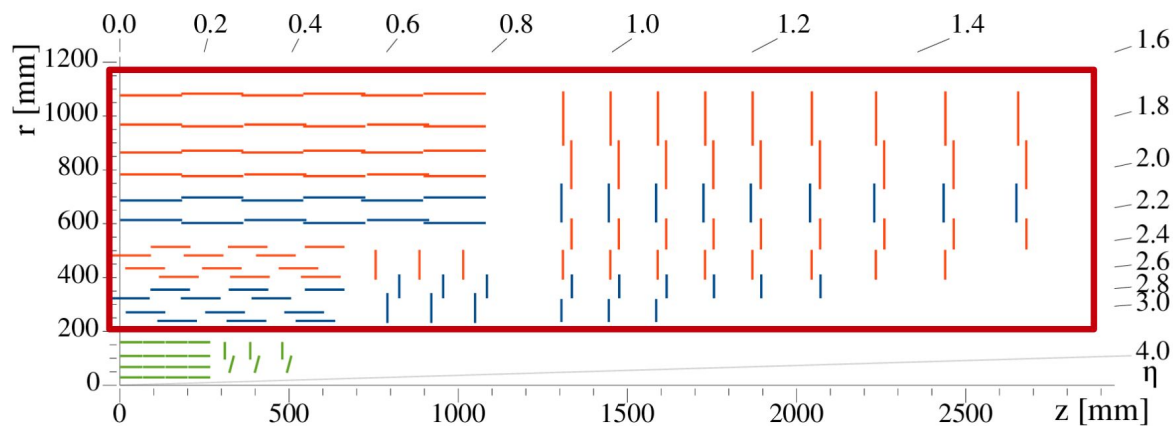
p-in-n silicon sensor

# CMS silicon tracker



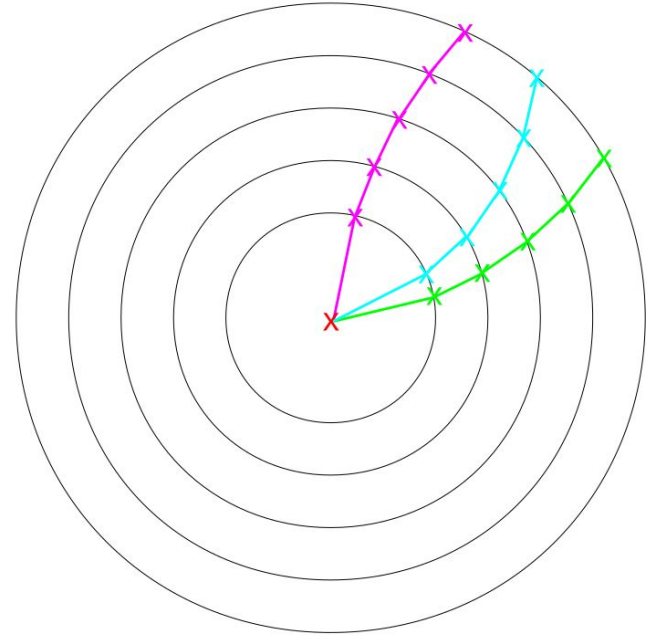
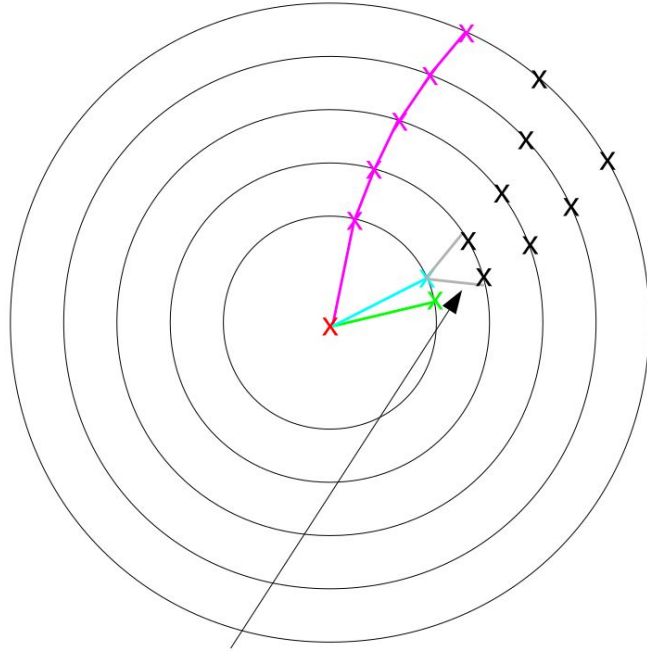
Strips vs pixels: how to determine location with strips?

- 10-12 layers of silicon sensors
- 15148 modules
- 9.3 million electronic channels
- Operated at  $-20^{\circ}\text{C}$  and  $< 20\%$  humidity
- In over 10 years of beam more than a billion particles fly through detector!

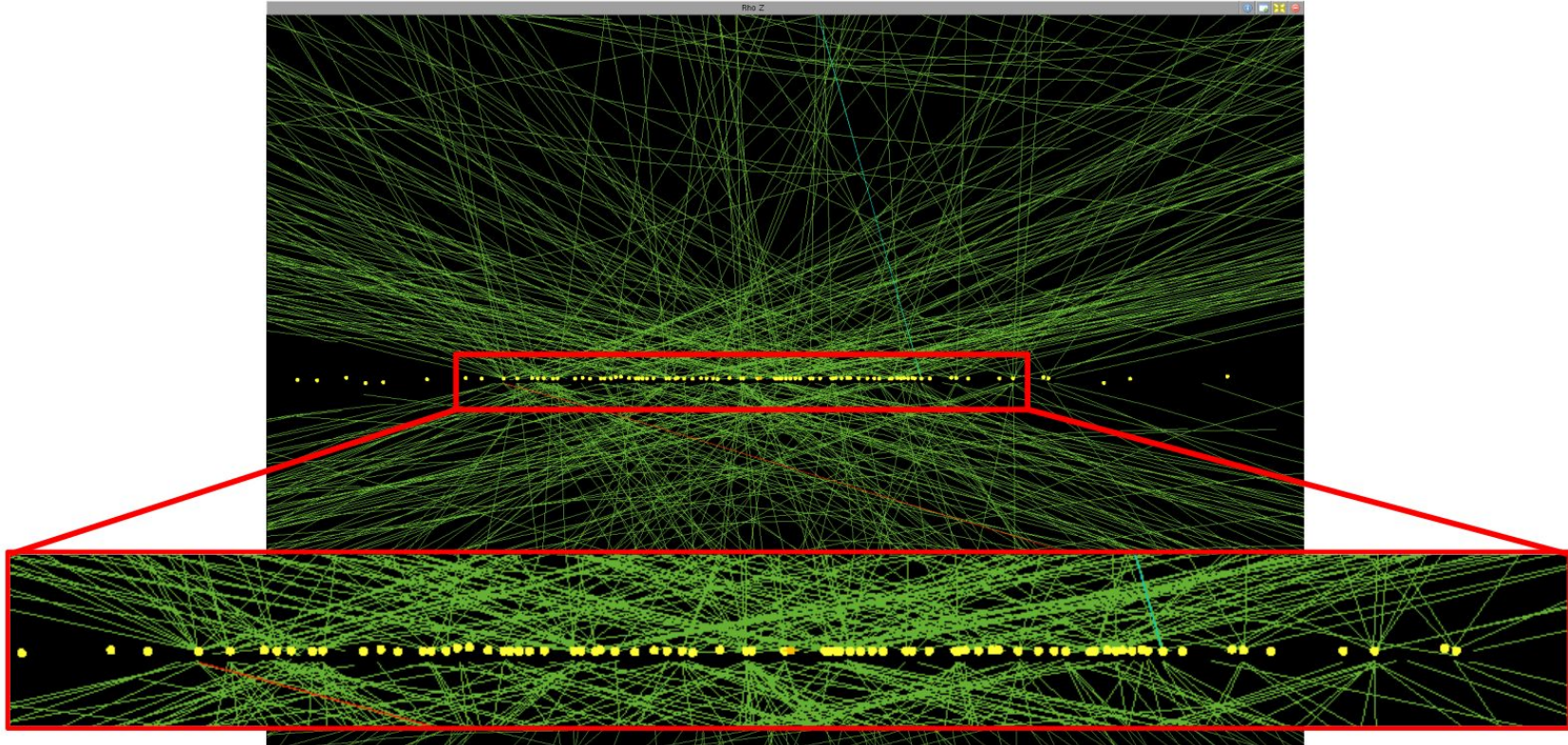


The pixel (or vertex) detector is so close to the beam pipe, it cannot survive this radiation: replaced in 2017, now inner layer will be replaced

# Track reconstruction: find hits that belong to track

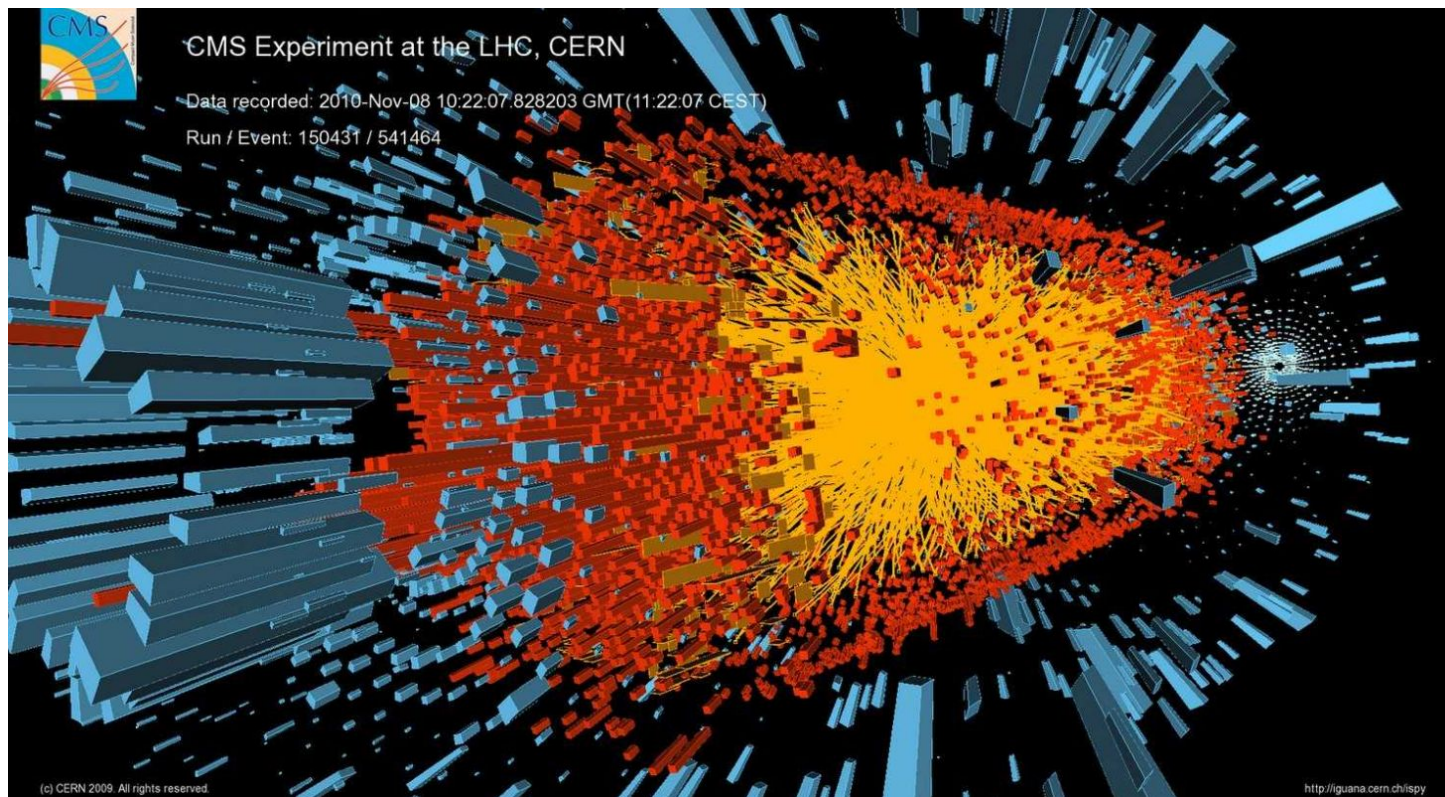


What if 78 interactions happen simultaneously?



# Or a collision of 2 lead nuclei?

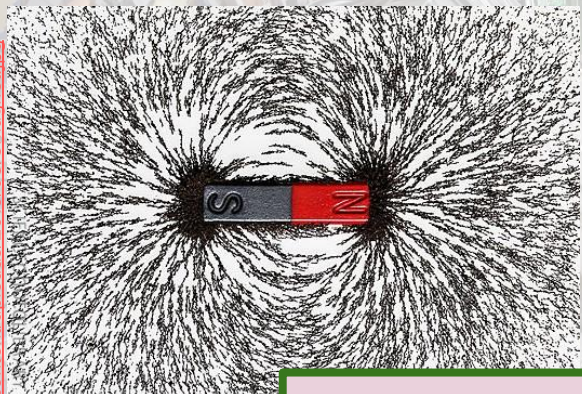
10000  
charged  
tracks!



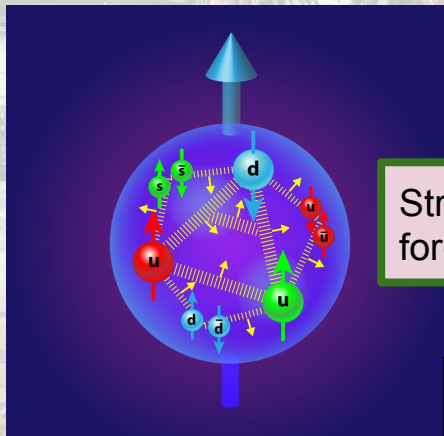
# Calorimeters

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# Interactions: four known forces



electromagnetism

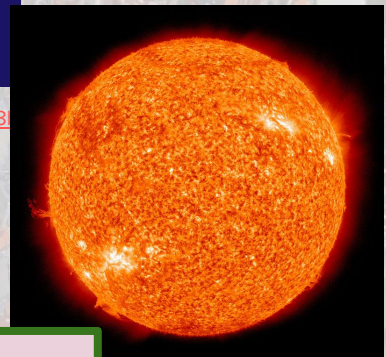
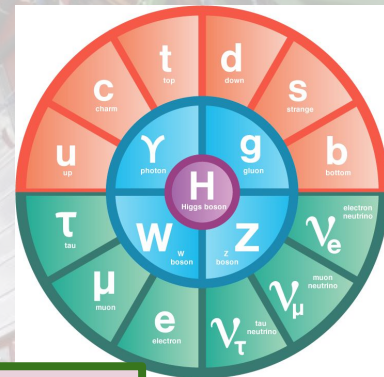


Strong nuclear force

[https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23\\_2.png](https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23_2.png)

What are we made of?  
How do particles get mass?  
**Gravity is not described by the Standard Model!**

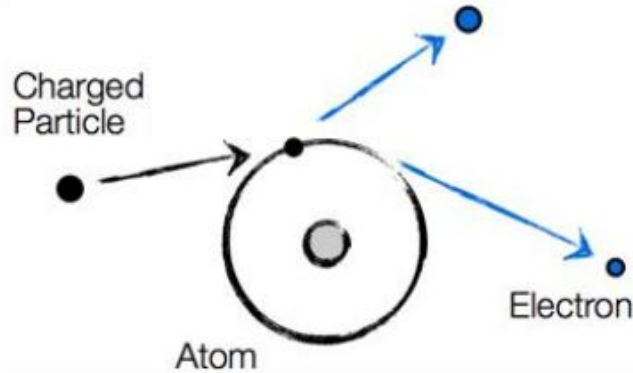
Weak nuclear force



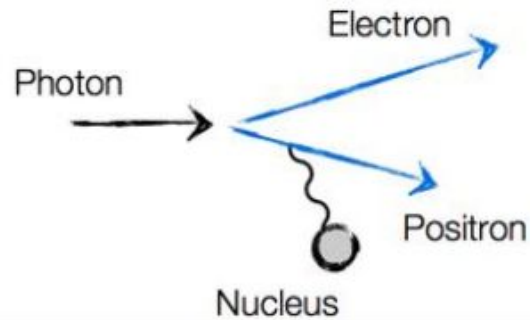
[https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/In\\_e\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg/800px-The\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg](https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/In_e_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg/800px-The_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg)

# Interaction with matter: destructive measurement

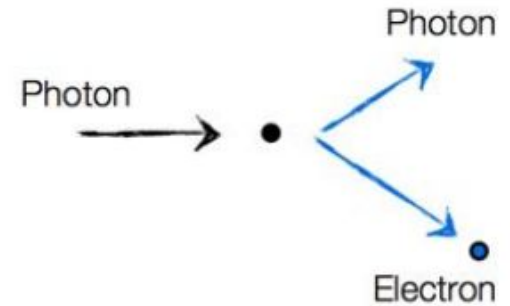
ionisation



Electron-positron  
pair production



Compton  
scattering

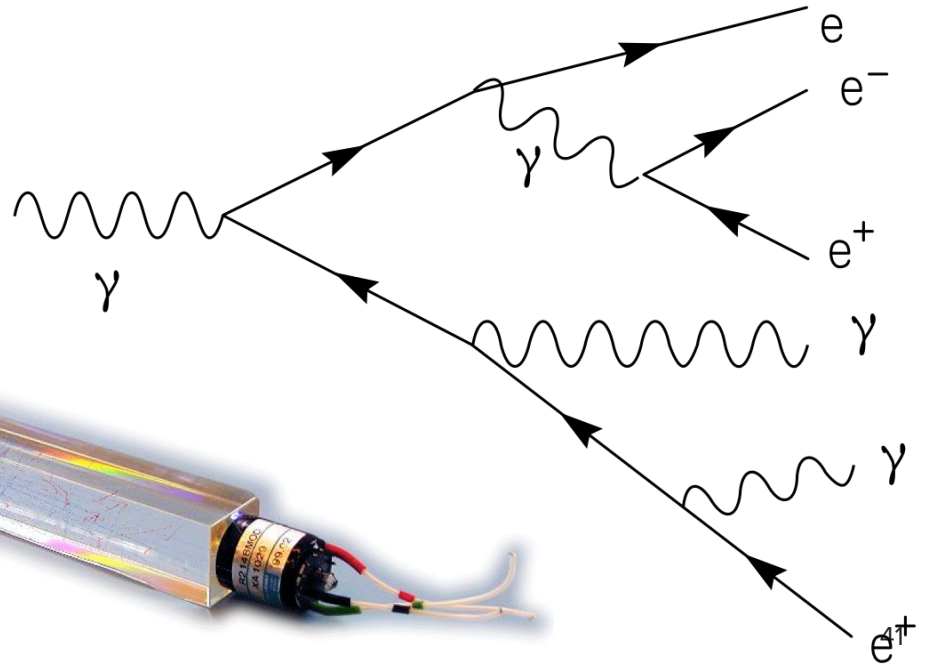
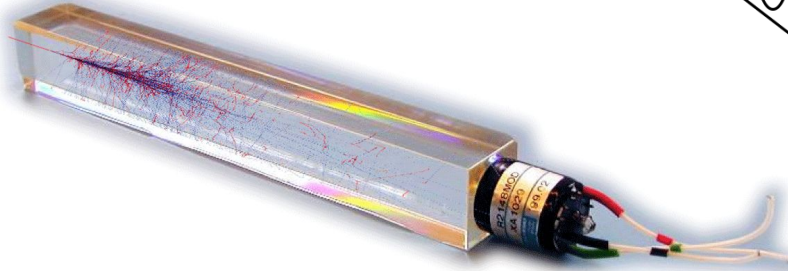




# Electromagnetic calorimeter

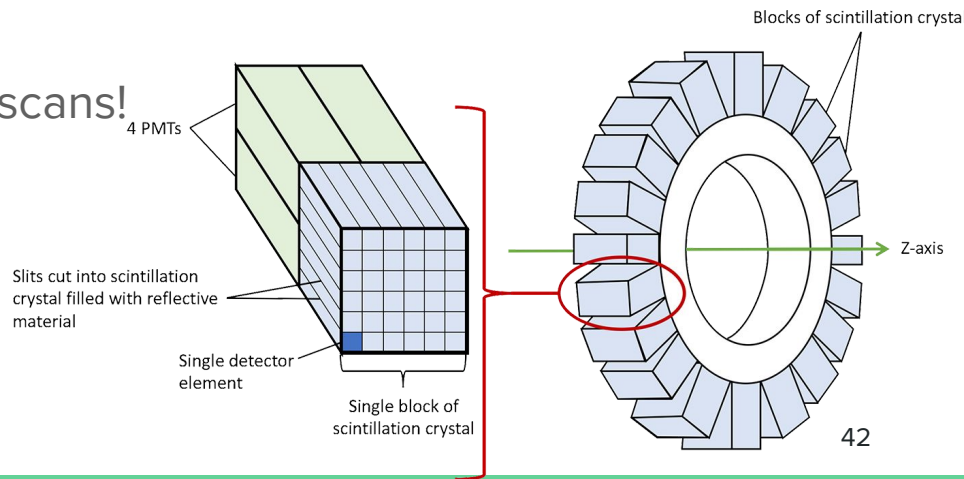
- Electromagnetic shower by interaction with material
- Depth of shower in a material is determined by
  - Energy
  - Critical energy where Bremsstrahlung rate = ionization rate
  - Radiation length of material

$$X = X_0 \frac{\ln(E_0/E_c)}{\ln 2}$$



# CMS electromagnetic calorimeter

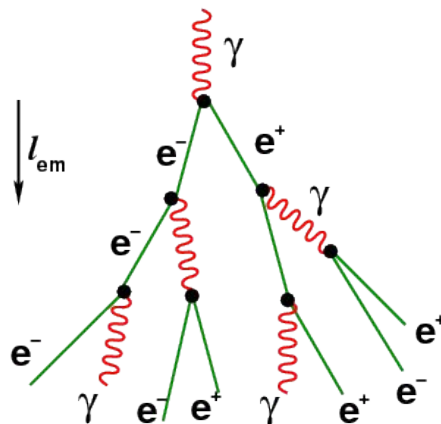
- Crystals made of lead tungstate, weighing 1.5 kg each
- 80,000 crystals each of which took 2 days to grow
- A crystal scintillates when an electron or photon passes through: produces light
- Light is read out with photomultiplier tubes: vacuum tubes that convert light into an electric signal
- Same scintillating crystals as for PET scans!



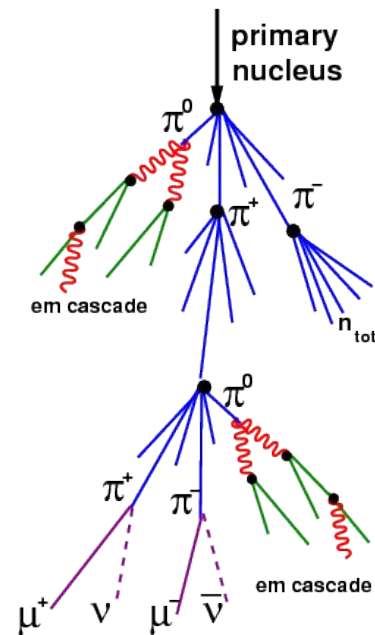
# Hadronic calorimeter

	$\lambda_{\text{int}}$ [cm]	$X_0$ [cm]
Szint.	79.4	42.2
LAr	83.7	14.0
Fe	16.8	1.76
Pb	17.1	0.56
U	10.5	0.32
C	38.1	18.8

## em cascade



## hadronic cascade



$$X_0 \sim \frac{A}{Z^2}$$

$$\lambda_{\text{int}} \sim A^{1/3}$$

$$\lambda_{\text{int}} \gg X_0$$

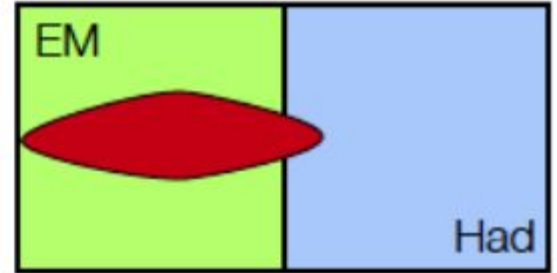
$$\rightarrow \frac{\lambda_{\text{int}}}{X_0} \sim A^{4/3}$$

Hadronic calorimeters are much thicker: larger shower depth!

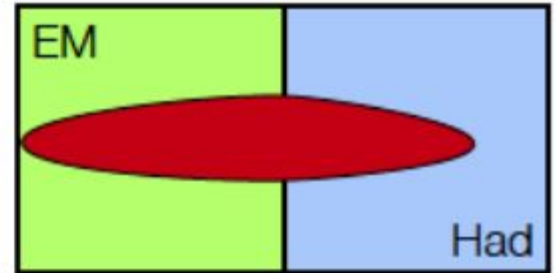
# Hadronic calorimeter

Need to consider fraction of electromagnetic energy

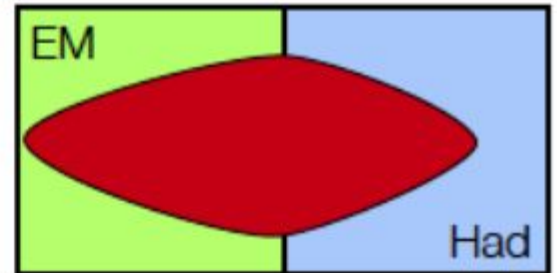
Electrons  
Photons



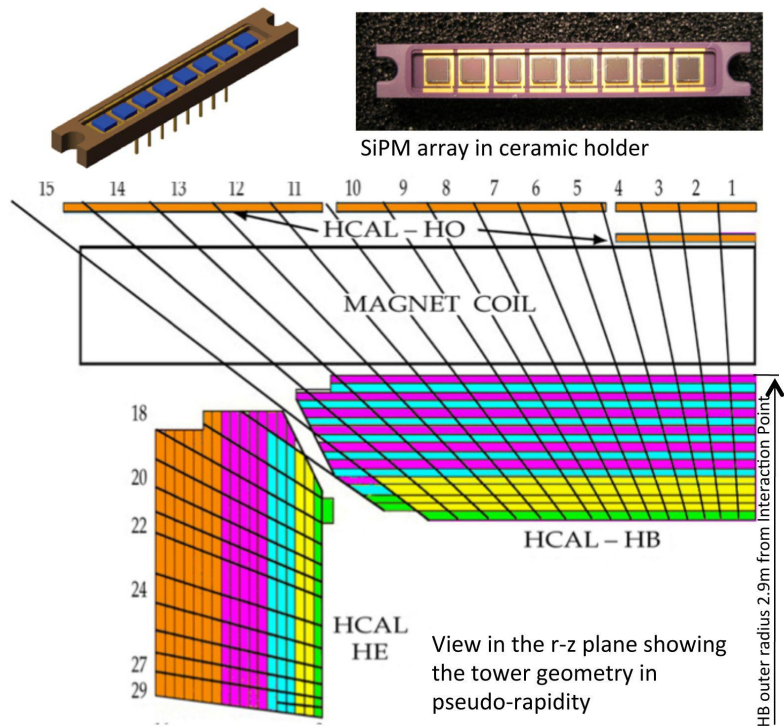
Taus  
Hadrons



Jets



# Photomultipliers → silicon photomultipliers



From hit information for 4 cells (1 tower) → hit information per cell

# How to measure neutrinos?

---

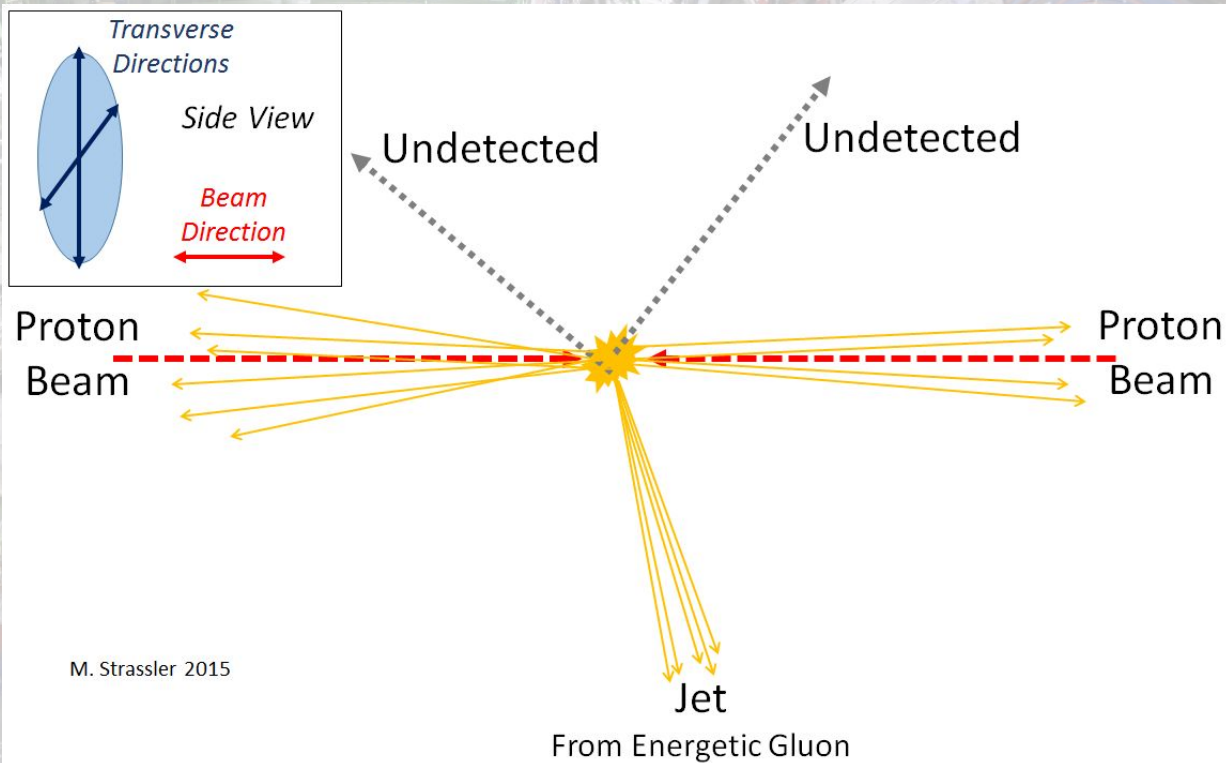
# Missing transverse momentum

Neutrinos?

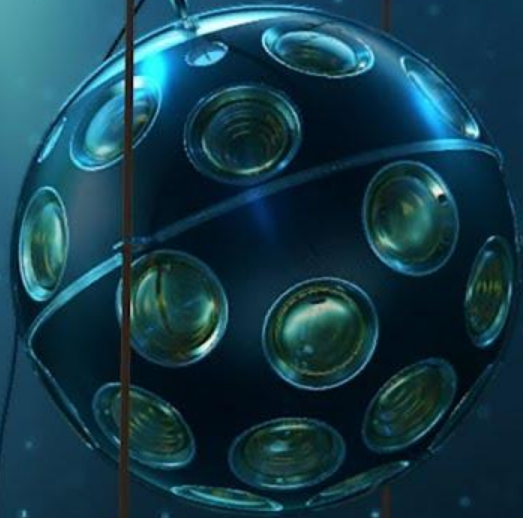
Mismeasurement?

Detector effect?

Dark matter?



- $10^9$  neutrinos /  $\text{cm}^2/\text{s}$
- Most from sun and atmosphere
- Rare events from black holes, supernovae...



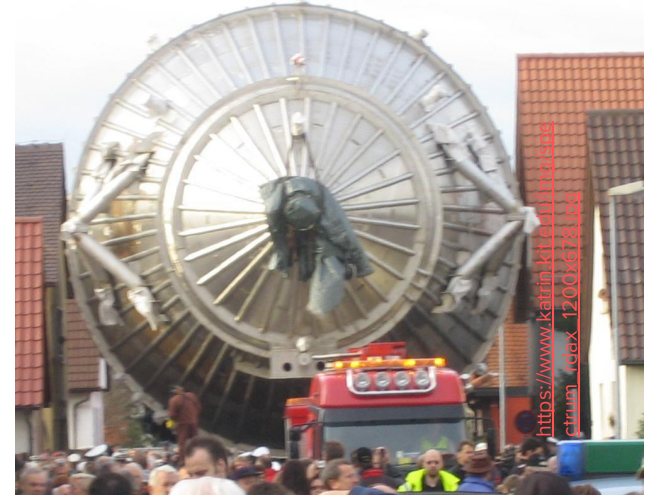
## **KM3NeT: cubic kilometer neutrino telescope**

- **Between 2 and 4 km deep in Mediterranean (FR-IT-GR)**
- **12000 digital optical modules (DOMs) on 600 strings**
- **Cherenkov detection with photomultipliers**
- **GeV, TeV, and PeV neutrinos**

**Netherlands plays a large role in construction**



# KATRIN: neutrino mass measurement

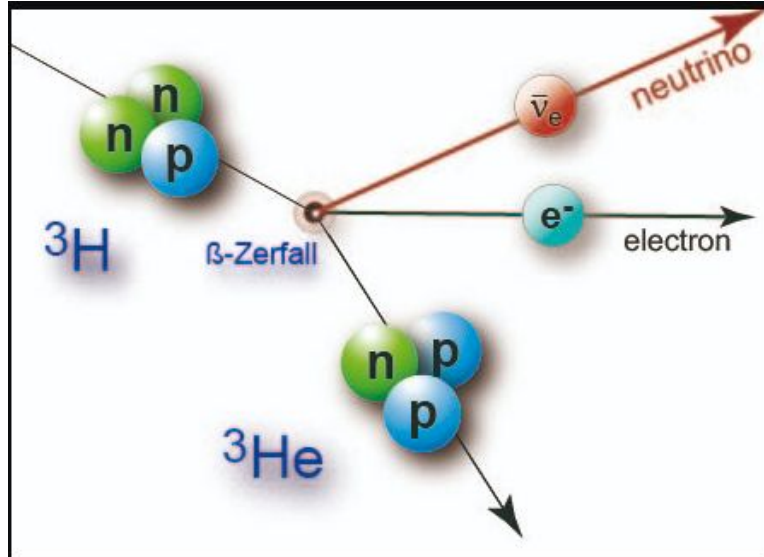


$m_\nu < 1.1$  eV (90% confidence level): twice as precise as previous measurements!

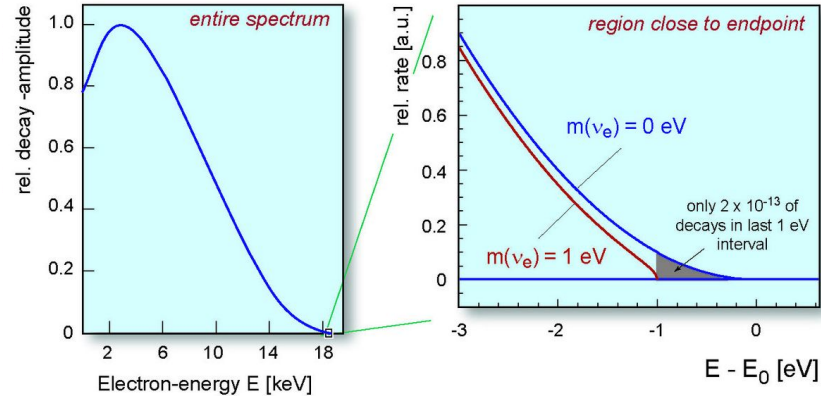
Recently published! <https://arxiv.org/abs/1909.06048>

# KATRIN: neutrino mass measurement

Karlsruhe tritium neutrino experiment



$m_\nu < 1.1 \text{ eV}$  (90% confidence level)



Derive  
neutrino mass  
information  
from electron  
energy

Recently published!

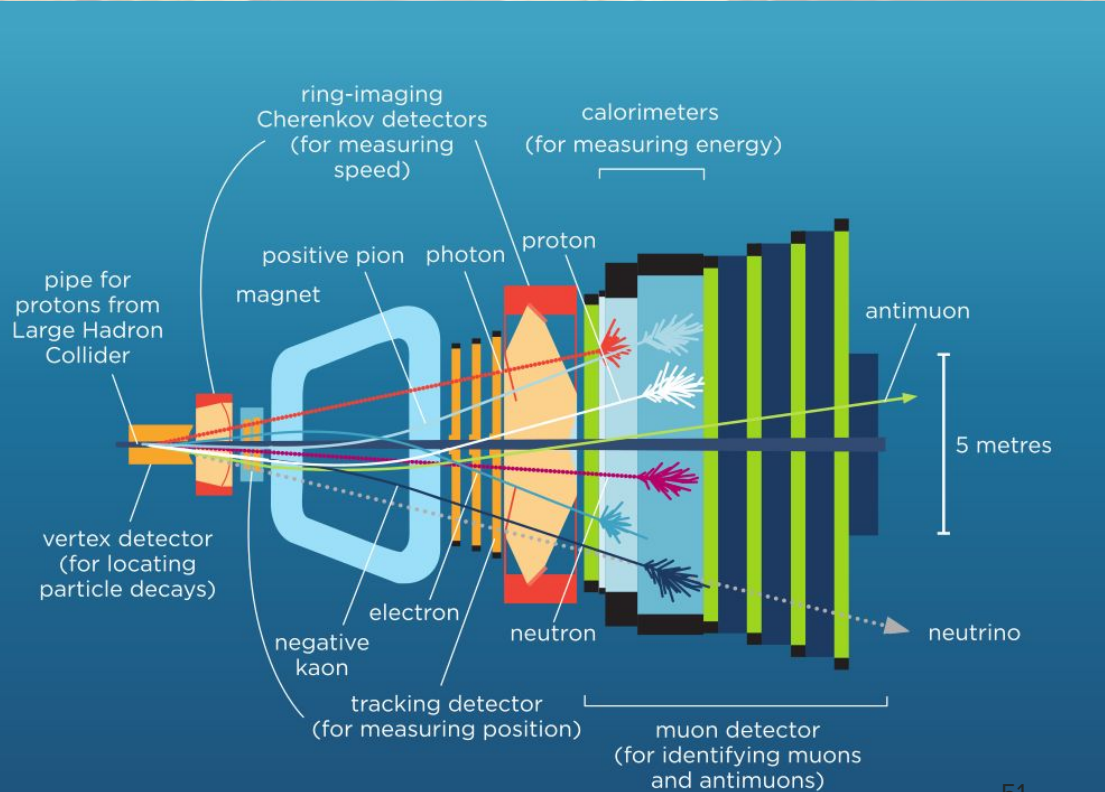
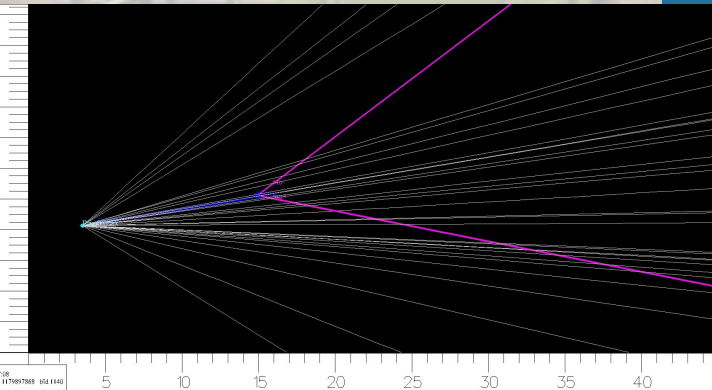
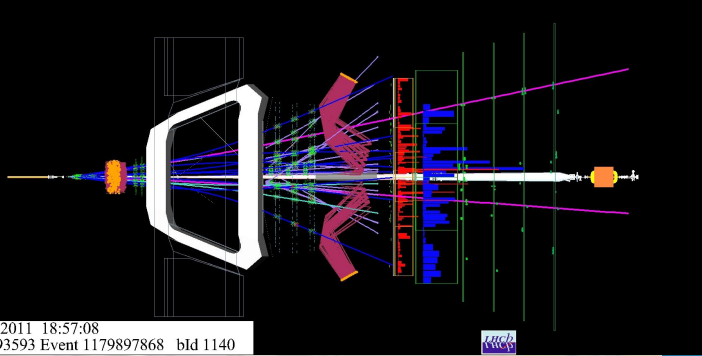
<https://arxiv.org/abs/1909.06048>

[https://www.katrin.kit.edu/img/spdctrum\\_rdx\\_1200x678.jpg](https://www.katrin.kit.edu/img/spdctrum_rdx_1200x678.jpg)

# Detecting particles with LHCb

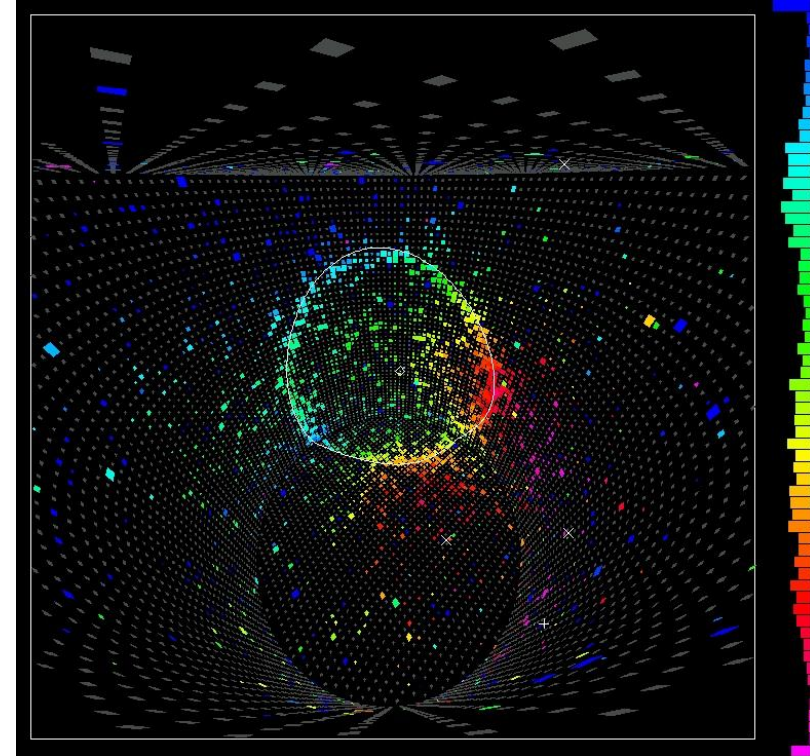
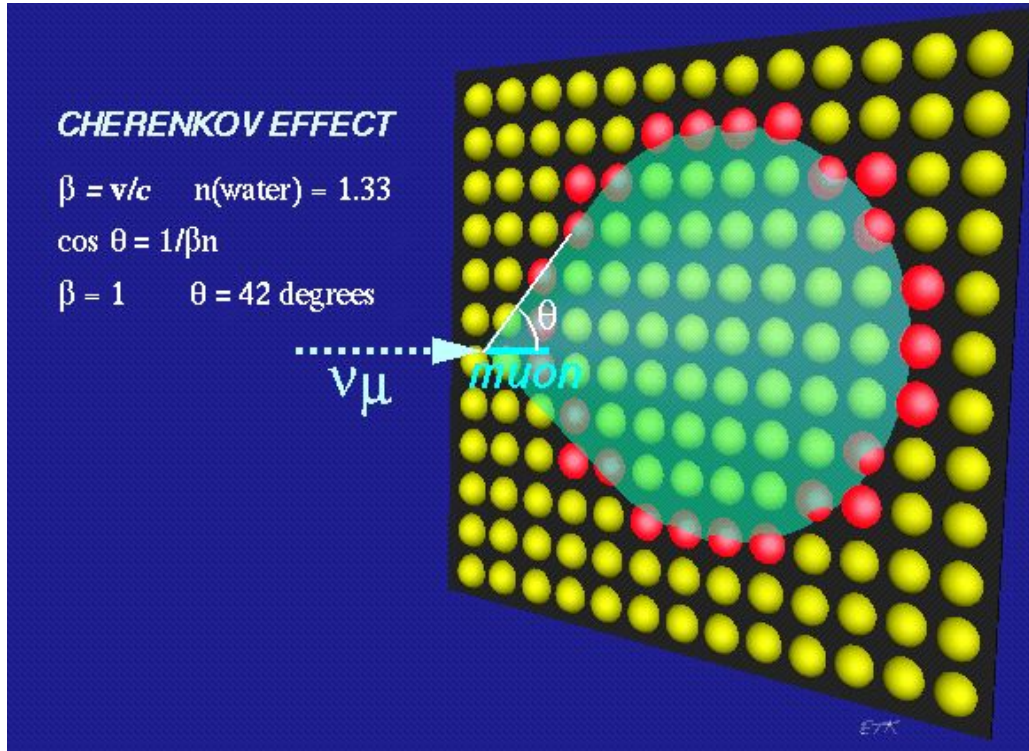
Cerenkov radiation like in HESS

LHCb Event Display



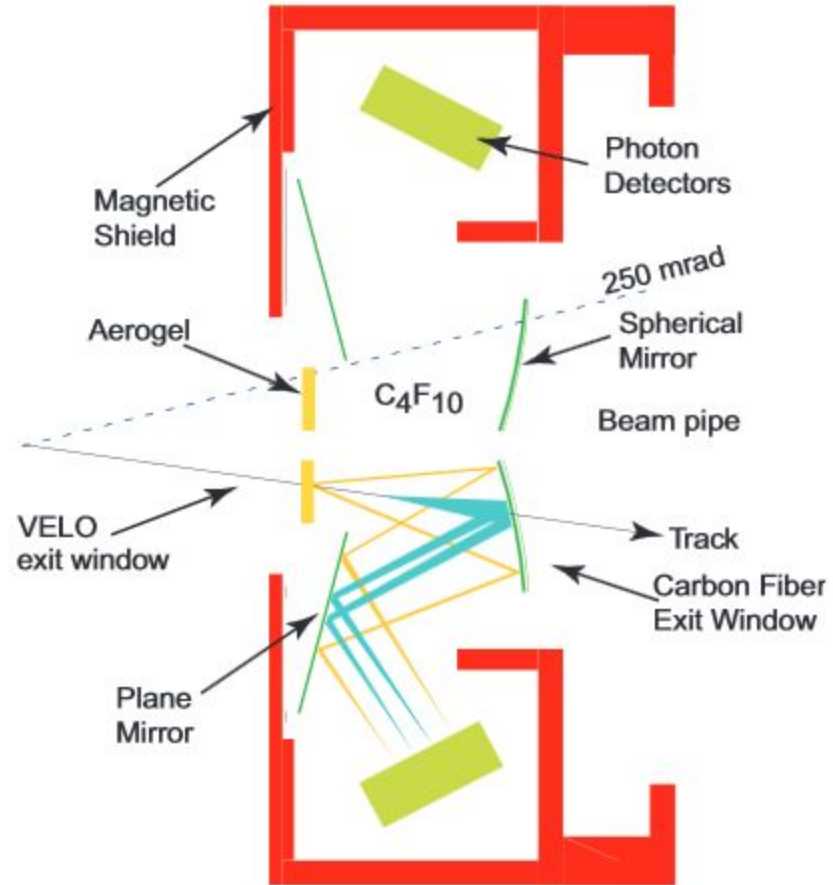
[http://lhcb-public.web.cern.ch/lhcb-public/Images/2011/BsMuMuPt\\_b.jpg](http://lhcb-public.web.cern.ch/lhcb-public/Images/2011/BsMuMuPt_b.jpg)

# Ring imaging Cherenkov detector



Cherenkov ring from a charged particle from a neutrino interaction in Kamiokande

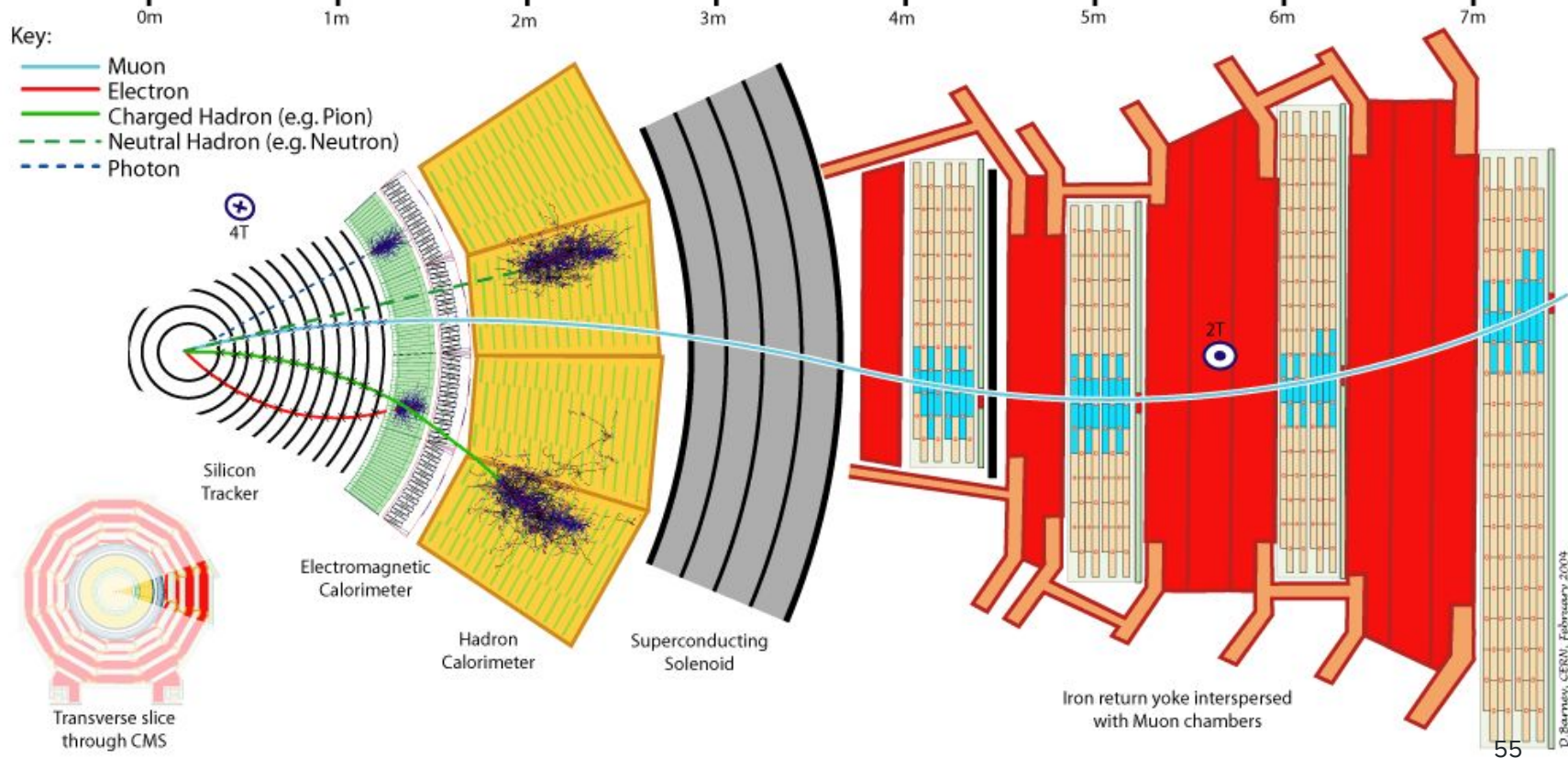
# LHCb RICH



# How to identify a particle?

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# Detectors at the large hadron collider: onion-like



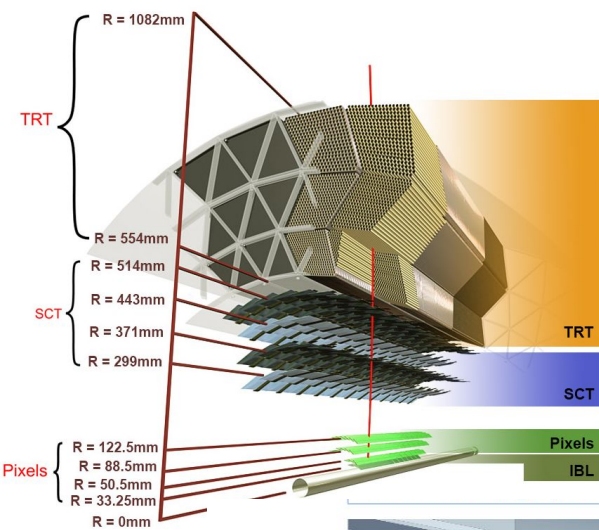
# Future detectors

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# Plan for the LHC

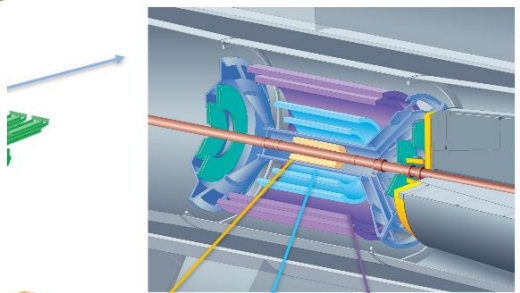




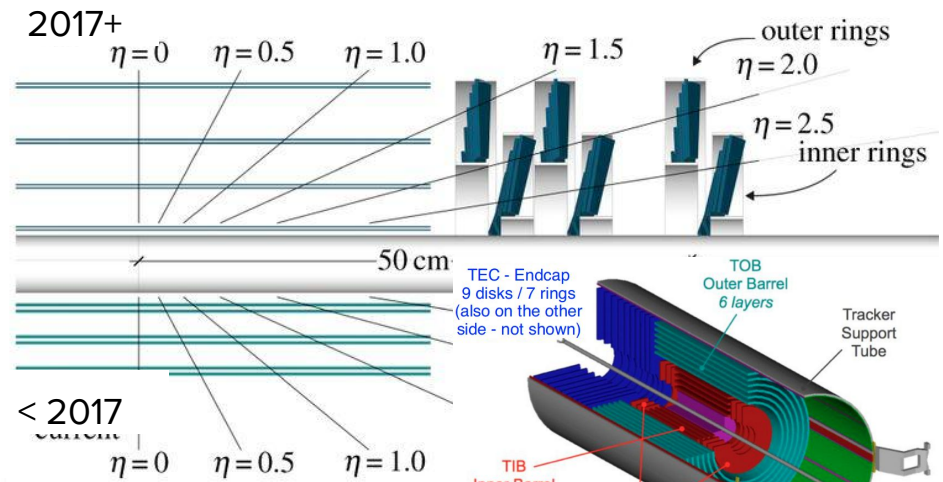
**ATLAS:** 3.325 cm from beam line

**LHC Run 2:** inner detector systems December 2018

**ALICE:** 3.3 cm from beam line

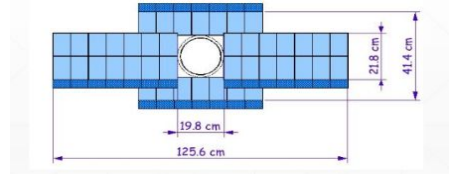
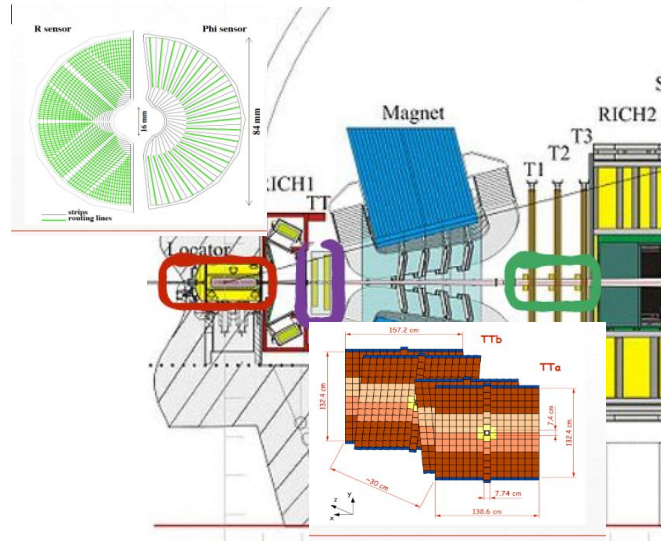


**Inner tracker upgrade in LS2**

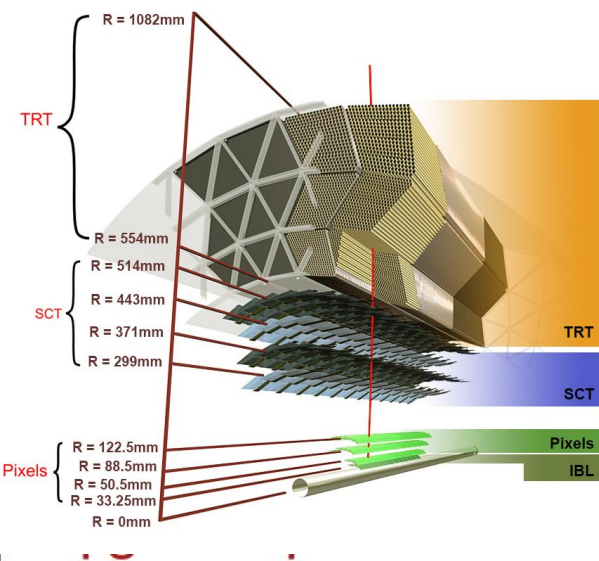


**CMS:** 2.9 cm from beam line

**Velo upgrade in LS2**

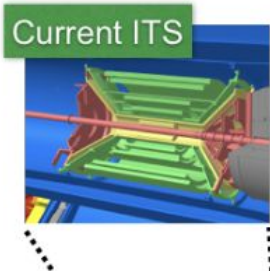
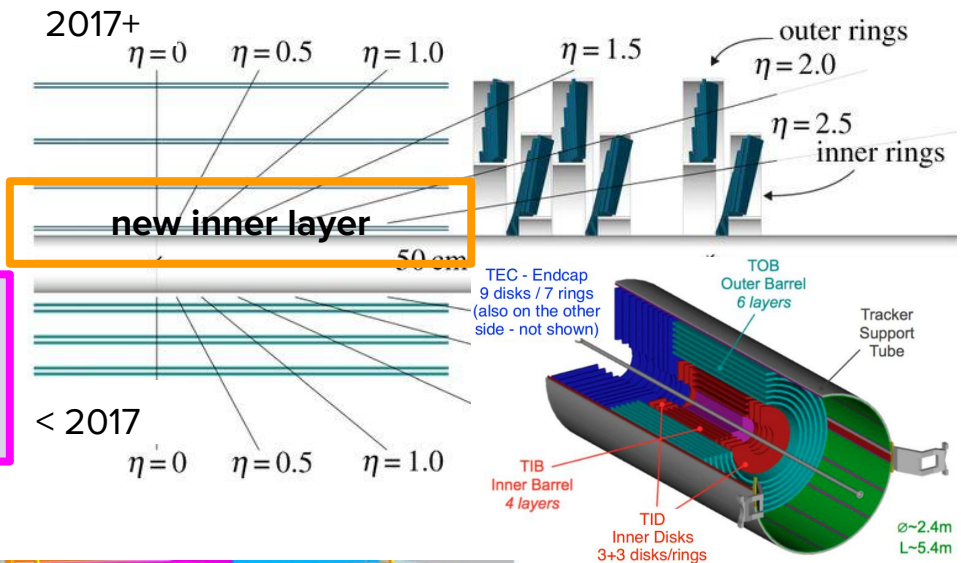


**LHCb:** 0.7 cm from beam line

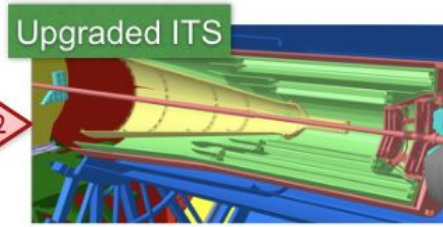


**ATLAS:** 3.325cm from beam line

**LHC Run 2:** inner detector systems early 2021

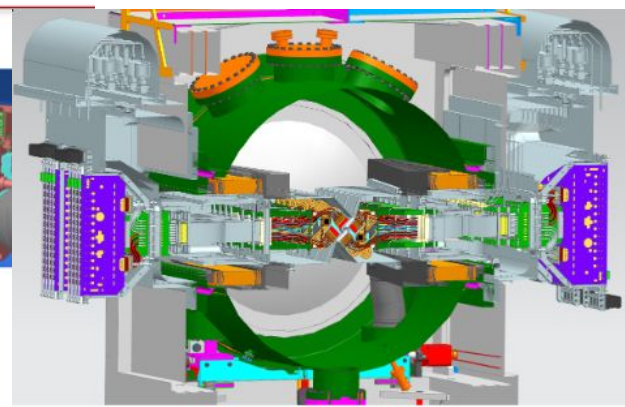


LHC LS2



**Inner tracker upgrade in LS2**

**ALICE:** 2.2 cm from beam line, monolithic active pixel sensors  $33\ \mu\text{m} \times 33\ \mu\text{m}$



**CMS:** 2.9 cm from beam line

**LHCb:** 0.51 cm from beam line, pixel sensors  $55\ \mu\text{m} \times 55\ \mu\text{m}$  in VELO

**VELO and trackers upgrade in LS2**

