

Next-to-minimal dark matter

Felix Brümmer

1703.00370, with A. Bharucha and R. Ruffault
1804.02357, with A. Bharucha and N. Desai
and work in progress



The summary to start with

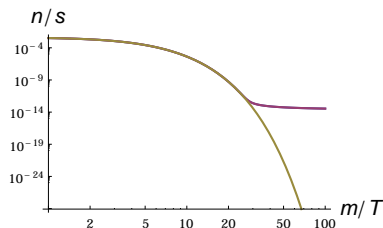
What I want to show in this talk:

- Sub-TeV WIMPs are still alive. . .
- . . . even the Original WIMPTM = the MSSM neutralino
- The LHC will further constrain them using long-lived particle searches
- More outlandish models give more outlandish signatures

What is a WIMP?

This talk's definition:

- A WIMP is a **thermal relic**.
- At large temperatures it is in **thermal equilibrium**. At temperatures $T \lesssim m$ its number density **decreases exponentially** due to scattering processes.
- At temperatures $T \lesssim T_{\text{freeze-out}}$ these processes become inefficient and its number density becomes constant.



- WI = “weakly interacting”. We take as part of the definition: DM abundance generated by **electroweak processes**.

WIMPs

First possibility: DM **itself** carries electroweak quantum numbers.

E.g. generic MSSM neutralino:

$$\chi_1^0 = \alpha \tilde{H}_u^0 + \beta \tilde{H}_d^0 + \gamma \tilde{B} + \delta \tilde{W}^0$$

This is **largely ruled out below a TeV** by direct detection → see e.g. Krall&Reece '17

Some remaining options:

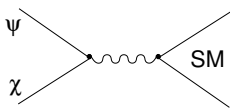
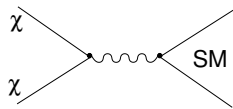
- mostly-bino χ_1^0 coannihilating with sleptons, or annihilating via Z - or Higgs funnel → e.g. GAMBIT collaboration '18
- pseudo-Dirac higgsino, $m_{\chi_1^0} = 1.1$ TeV → A. Delgado's talk
- pure wino, $m = 2.5$ TeV (but under pressure from indirect detection)
- SU(2) 5-plet fermion, $m_5 \approx 10$ TeV
- SU(2) 7-plet scalar, $m_7 \approx 25$ TeV
- ...

} SUSY
} MDM

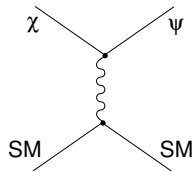
WIMPs

Second possibility: DM χ does not interact much with the EW sector but **remains in equilibrium** by scattering with some particle ψ which does.

This will be the case for most of this talk's WIMPs. Specifically:

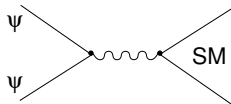


can be small



provided that χ remains in equilibrium via conversion

and that



is efficient.

Annihilation of ψ efficiently depletes χ , even if $\psi - \chi$ mixing angle is $\ll 1$.

The case of marginally efficient conversion processes is especially interesting

→ D'Agnolo/Pappadopulo/Ruderman '17, Garmy/Heisig/Lülf/Vogl '17

Next-to-minimal dark matter

For the purposes of this talk:

- DM χ is mostly a Majorana singlet
- Stabilized by \mathbb{Z}_2
- Coannihilation partner ψ is a \mathbb{Z}_2 -odd fermionic n -plet of SU(2)
- Sub-TeV particle content is minimal: there is only ψ and χ
- We don't consider SU(2) doublets (mostly ruled out by direct detection anyway) → extensive literature on well-tempered bino-higgsino and its non-SUSY version

Dark matter is a singlet fermion χ mixing with an n -plet fermion ψ ($n \geq 3$) through higher-dimensional operators.

States inducing the mixing live at scales \gtrsim TeV \Rightarrow irrelevant for LHC if carrying only EW charges.

A familiar example

Split SUSY with somewhat heavy higgsinos and $M_1 < M_2$.

DM is **mostly bino** (singlet), mixing with **wino** (3-plet) through dimension-5 operator

$$\mathcal{L}_{\text{mix}} = \frac{\tilde{g}_u \tilde{g}'_d + \tilde{g}_d \tilde{g}'_u}{\mu} \phi^\dagger \tau^a \phi \widetilde{W}^a \widetilde{B}$$

where ϕ = SM-like Higgs doublet.

Example of a “**well-tempered neutralino**” → e.g. Arkani-Hamed/Delgado/Giudice '06

Pheno details: e.g. → Rolbiecki/Sakurai '15, Nagata/Otono/Shirai '15

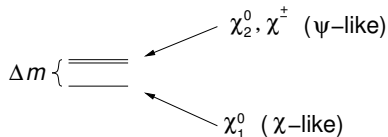
Features:

- Rather small direct detection cross section (small mixing angle in MSSM, no Z coupling to either bino or wino)
- Tiny indirect detection cross section (DM mostly singlet)

First case study: SU(2) triplets

→ See also talk by A. Filimonova

Particle content: one charged and two neutral fermions



Interactions of χ_1^0 with SM mainly through two operators:

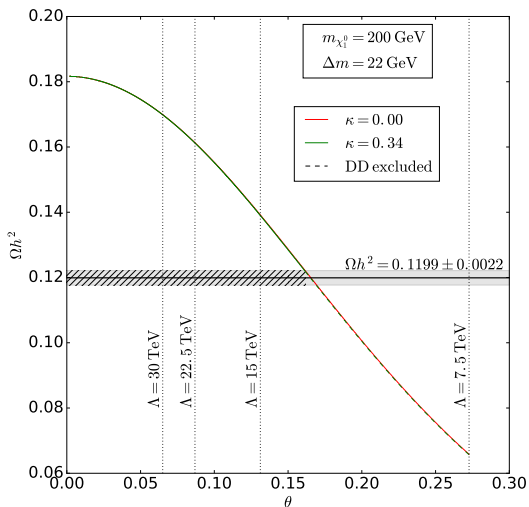
$$\mathcal{L} = \frac{1}{2} \frac{\kappa}{\Lambda} \phi^\dagger \phi \chi \chi + \frac{\lambda}{\Lambda} \phi^\dagger \tau^a \phi \psi^a \chi + \text{h.c.}$$

ϕ = SM Higgs; Λ = cutoff scale.

Wilson coefficients κ, λ both contribute to **DM annihilation** (\Rightarrow thermal relic density) and **DM-nucleus scattering** (\Rightarrow direct detection).

Trade λ for **mixing angle** $\theta = \frac{\lambda}{\Lambda} \frac{v^2}{\Delta m}$

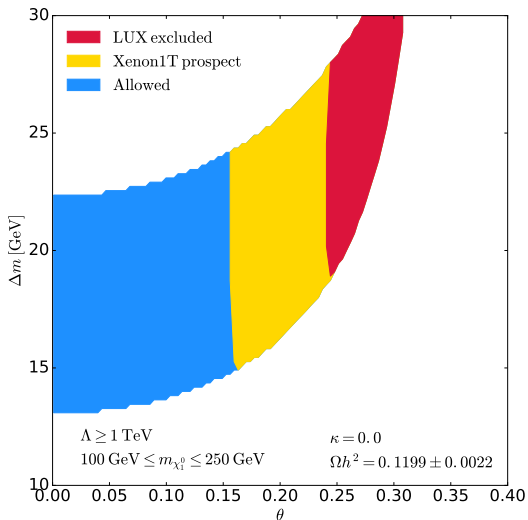
First case study: SU(2) triplets



Any κ large enough to significantly influence relic density is ruled out by DD.

First case study: SU(2) triplets

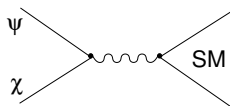
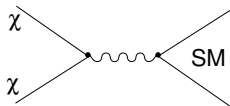
For $\kappa = 0$:



SU(2) triplets at (really) small mixing angle

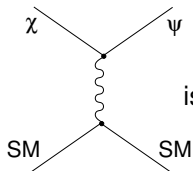
How far away can the cutoff scale Λ be?

Remember: we don't need



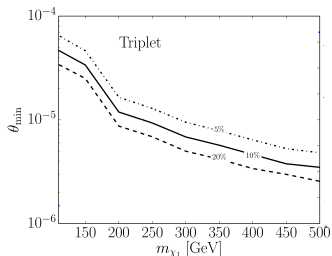
($\propto \theta^2$, doubly Boltzmann-suppressed)

provided that



is effective ($\propto \theta^2$, **singly** Boltzmann-suppressed)

Still the case for surprisingly small θ :

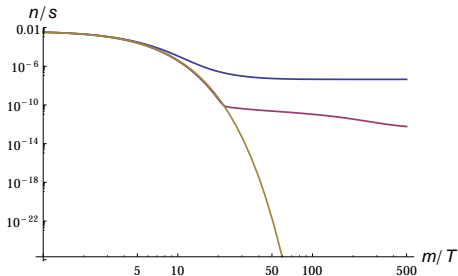


SU(2) triplets at (really) small mixing angle

At even smaller mixing: Relic density no longer determined by ψ annihilations freezing out but by χ SM \rightarrow ψ SM conversion rate dropping below Hubble rate

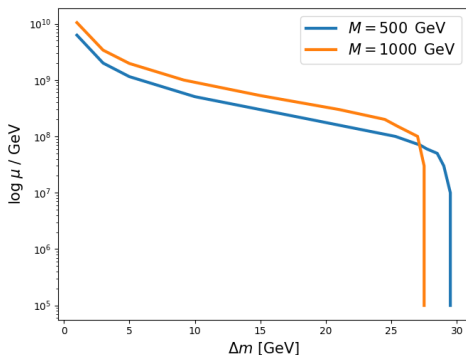
\rightarrow D'Agnolo/Pappadopulo/Ruderman '17, Garny/Heisig/Lülf/Vogl '17

- “Effective number density” formalism used by standard codes no longer applicable: need to solve full Boltzmann equations, including conversion terms
- DM number density departs from equilibrium earlier \Rightarrow increased Ωh^2
- Can compensate (to some extent) by varying Δm
- “Conversion-driven freeze-out”, “Coscattering”



χ abundance, ψ abundance, equilibrium

Special case: Split SUSY at (really) large μ

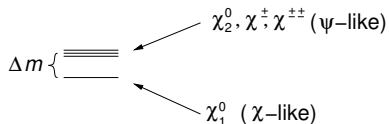


$$\tan \beta = 1 \quad \Omega h^2 = 0.12 \quad \Delta m \equiv m_{\chi_1^\pm} - m_{\chi_1^0}$$

- Coannihilation phase: $\Delta m \approx \text{const.}$
- Coscattering phase at large $\mu \gtrsim 10^7$ GeV

Second case study: SU(2) quintuplets

Particle content: one doubly charged, one singly charged and two neutral fermions

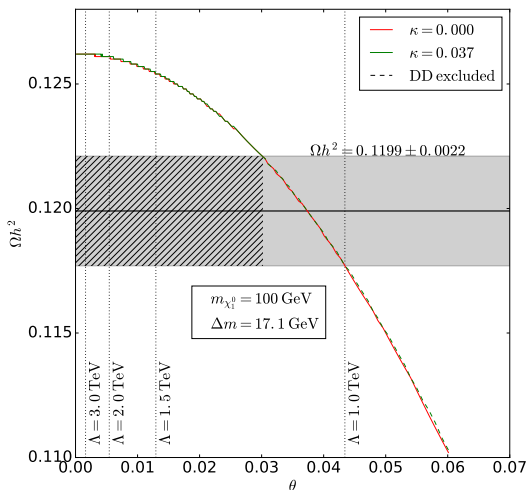


- Doubly charged state \Rightarrow potentially characteristic signatures at LHC (long-lived)
- Mixing operator is now **dimension 7**:

$$\frac{\lambda}{\Lambda^3} \phi^{\dagger i} \phi_j \phi^{\dagger k} \phi_\ell C_{A ik}^{j\ell} \psi^A \chi + \text{h.c.}$$

\Rightarrow mixing angles **guaranteed to be small** for cutoff scale $\Lambda \gtrsim \text{TeV}$

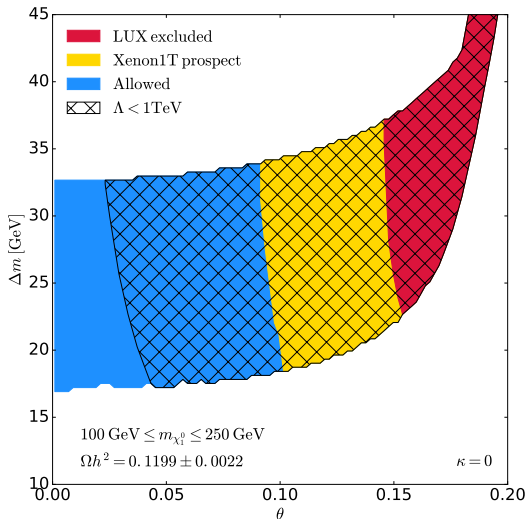
Second case study: SU(2) quintuplets



Similar to triplet case. Mass parameters need some tuning. Wilson coefficient κ for dimension-5 operator $\frac{\kappa}{\Lambda} \phi^\dagger \phi \chi \chi$ already **tightly constrained** by direct detection.

First case study: SU(2) quintuplets

For $\kappa = 0$:



On quadruplets

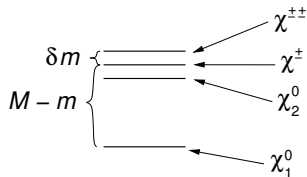
The quadruplet (isospin 3/2) case is more complicated because

- we now need a Dirac fermion $(\psi, \bar{\psi})$
- one doubly charged, two singly charged mass eigenstates
- **three** neutral mass eigenstates \Rightarrow **two** relevant neutral mixing angles
- spectrum depends on additional Wilson coefficients inducing non-universal mass splittings in the ψ sector
- **proliferation of parameters**

Result of our analysis: Qualitatively similar conclusions as for triplet/quintuplet case \rightarrow [FB/Bharucha/Ruffault '17](#)

Collider phenomenology: Displaced leptons

Zoom in on the quintuplet mass spectrum:



where

$$\delta m = \delta m^{(\text{tree})} + \delta m^{(1\text{-loop})}$$

Source for $\delta m^{(\text{tree})}$ is dimension-7 operator:

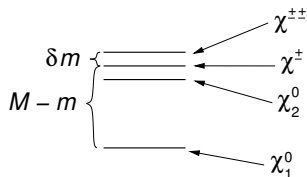
$$\frac{1}{\Lambda^3} \underbrace{C^{ABC}}_{\text{tot. symmetric}} \psi_A \psi_B (\phi^\dagger \phi \phi^\dagger \phi)_C \quad \Rightarrow \quad \left| \delta m^{(\text{tree})} \right| \lesssim \text{few } 100 \text{ MeV}$$

Source for $\delta m^{1\text{-loop}}$ are electroweak loops:

$$\chi^{\pm(\pm)} \text{---} \text{---} \chi^{\pm(\pm)} \quad \Rightarrow \quad \delta m^{(\text{one-loop})} \approx 500 \text{ MeV}$$

Collider phenomenology: Displaced leptons

Zoom in on the quintuplet mass spectrum:



Numerically: $M - m \approx 15 - 50$ GeV (coannihilation), $\delta m \approx$ few 100 MeV.

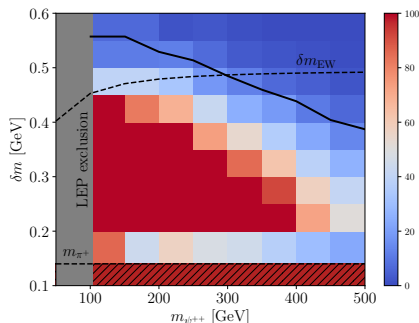
Typical $\chi^{\pm\pm}$ decay:

$$\chi^{\pm\pm} \rightarrow \chi^\pm (\rightarrow \chi^0_1 \ell^\pm \nu_\ell) \pi^\pm$$

- Decay into singly-charged χ^\pm and pion via off-shell W is **only open 2-body mode**. Pion too soft to be seen.
- Small mass splitting $\delta m \Rightarrow$ small phase space \Rightarrow macroscopic decay length ≈ 0.5 mm
- Lepton from subsequent χ^\pm decay will be **displaced**.

Collider phenomenology: Displaced leptons

- CMS has published searches for displaced OS leptons at 8 TeV \rightarrow PRL 114 (2015) 6 and 13 TeV \rightarrow CMS-PAS-EXO-16-022
- 8 TeV analysis gives **better constraints** because of looser lepton p_T cuts (At 13 TeV, hard cuts to completely remove heavy flavour backgrounds)
- Exclusion (black curve) as a function of M and δm :



Collider phenomenology: Soft dileptons

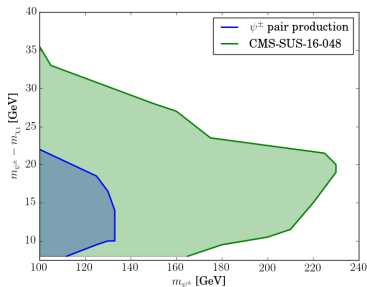
Triplet model doesn't have long-lived states except at very small θ

Instead: **Low-momentum OS lepton pairs**

Search for **compressed SUSY neutralino-chargino pair production**

→ CMS-PAS-SUS-16-048 allows to **constrain triplet model**. Exclusion depends on preferred χ_2 decay mode:

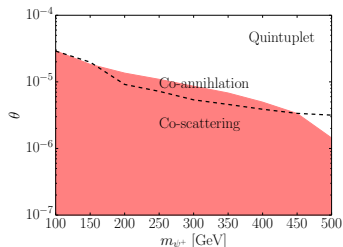
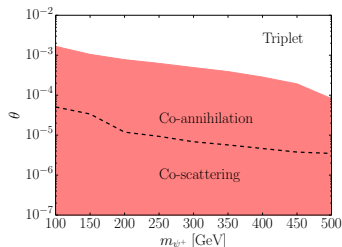
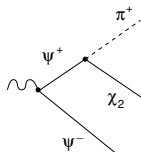
- **green curve**: CMS analysis, assuming $\chi_2 \psi^\pm$ production with $\chi_2 \rightarrow \chi_1 Z^*$
- **blue curve**: recast for $\psi^+ \psi^-$ production if $\chi_2 \rightarrow \chi_1 h^*$ (no leptons)



Collider phenomenology: Disappearing tracks

For **small mixing angles** $\theta \lesssim 10^{-3}$: singly-charged ψ^+ decays into χ_2 (ψ -like) rather than directly into χ_1 (χ -like).

Mass degeneracy in n -plet sector \Rightarrow long-lived $\psi^+ \Rightarrow$ **disappearing track**



Conclusions

- Dark matter could be a mixed singlet – n -plet with an EW-scale mass.
- Effective theory. Mixing induced by higher-dimensional operators.
- Simplest example: Well-tempered bino-wino in split SUSY. “Phase diagram” for neutralino dark matter.
- LUX already very constraining, Xenon1T even more. But at small mixing, DD constraints go away.
- Collider signatures:
 - Displaced leptons in quintuplet model
 - Soft dileptons in triplet model (weaker constraints)
 - Disappearing tracks in both (if mixing angle small)
 - A 500 GeV bino-like neutralino is not excluded if it coannihilates/coscatters with a wino-like chargino and neutralino. If it coscatters, the chargino will soon be found in disappearing track searches.