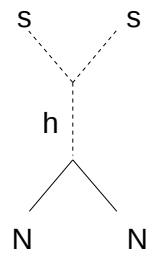
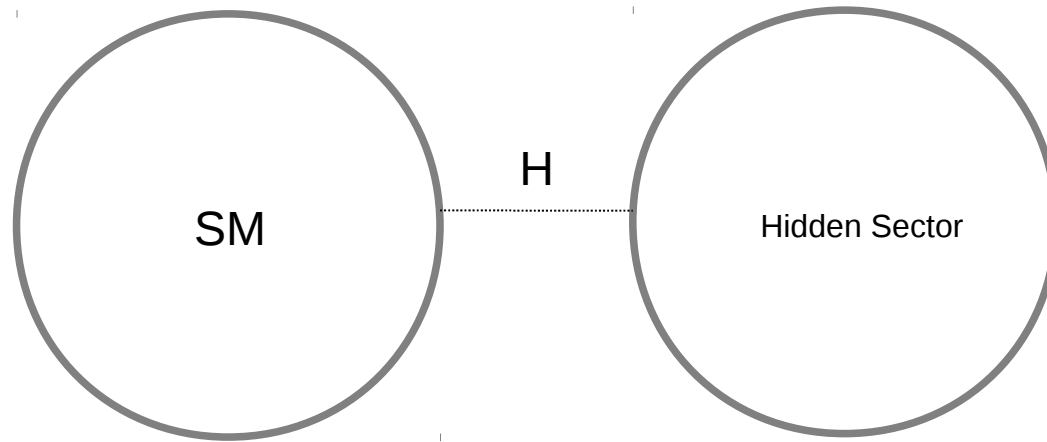

Pseudo-Goldstone Dark Matter

through the Higgs portal

Oleg Lebedev



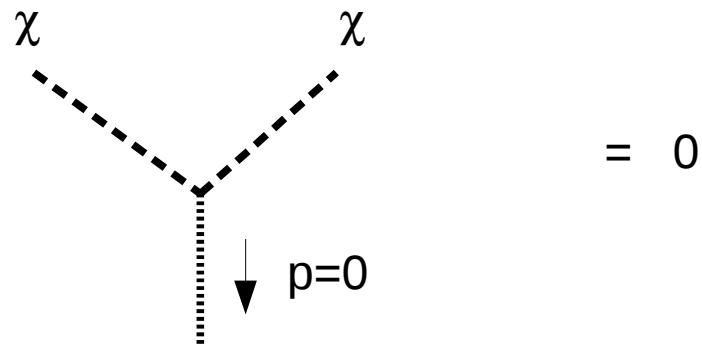
University of Helsinki



strong constraints on a WIMP

Main Goldstone feature:

derivative couplings $\partial\chi \dots$



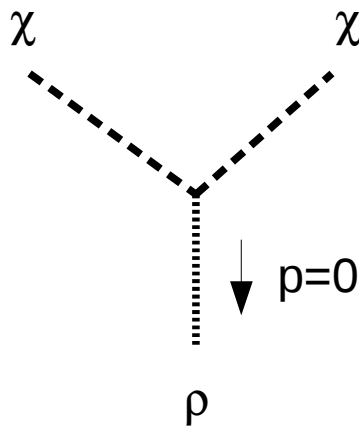
Perfect for χ – dark matter:

- no direct detection
- efficient annihilation ($p \neq 0$)

Massive "Goldstone" case:

$$L = \partial_\mu S \partial^\mu S^* - m^2 S S^* + (\mu^2 S^2 + \text{h.c.}) - V(SS^*)$$

$$S = \rho e^{i\chi} \quad (\chi = \text{pseudo-Goldstone boson with mass } \sim \mu)$$



$$\sim \chi \rho (\square + 4\mu^2) \chi = 0 \quad (\text{EOM})$$

Still works! (NB: only for an S^2 soft term)

Pseudo-Goldstone dark matter

SM + a complex scalar S , require softly broken U(1) symmetry:

$$V = V_0 + V_{\text{soft}} ,$$

$$V_0 = -\frac{\mu_H^2}{2} |H|^2 - \frac{\mu_S^2}{2} |S|^2 + \frac{\lambda_H}{2} |H|^4 + \lambda_{HS} |H|^2 |S|^2 + \frac{\lambda_S}{2} |S|^4 ,$$

$$V_{\text{soft}} = -\frac{\mu_S'^2}{4} S^2 + \text{h.c.}$$

All parameters are real $\Rightarrow \langle S \rangle = \text{real}$, $S \rightarrow S^*$ symmetry

Im S = Dark Matter

States:

$$S = \frac{1}{\sqrt{2}}(v_s + s + i\chi)$$

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v + h \end{pmatrix}$$

CP-even mass eigenstates = h_1, h_2 (mixtures of s and h):

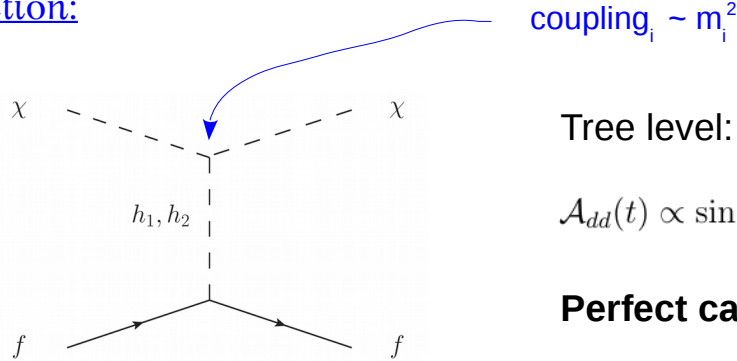
$$O^T \mathcal{M}^2 O = \begin{pmatrix} m_{h_1}^2 & 0 \\ 0 & m_{h_2}^2 \end{pmatrix} \quad O = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

$$m_{h_1, h_2}^2 = \frac{1}{2} \left(\lambda_H v^2 + \lambda_S v_s^2 \mp \frac{\lambda_s v_s^2 - \lambda_H v^2}{\cos 2\theta} \right)$$

CP-odd mass eigenstate = dark matter χ :

$$m_\chi^2 = \mu'_S{}^2$$

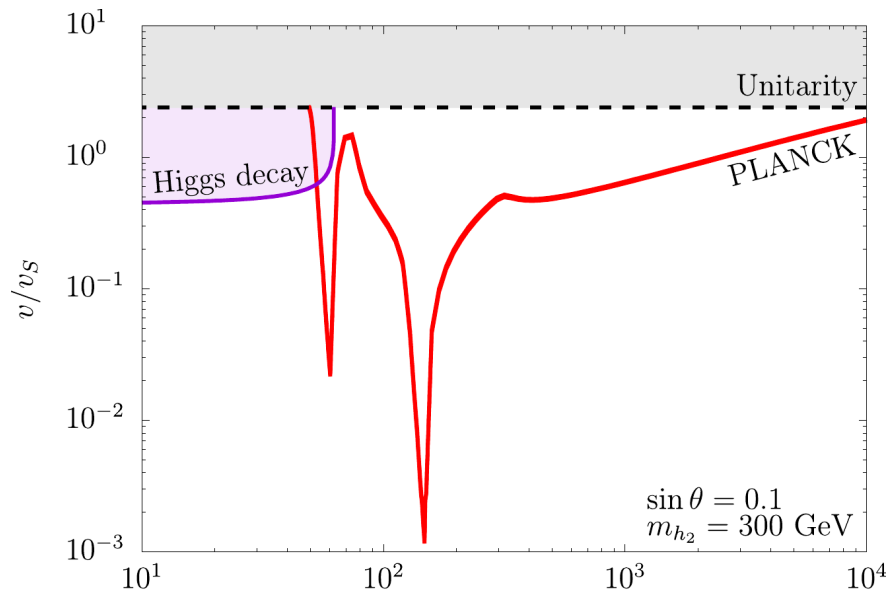
Direct detection:



Tree level:

$$\mathcal{A}_{dd}(t) \propto \sin \theta \cos \theta \left(\frac{m_2^2}{t - m_2^2} - \frac{m_1^2}{t - m_1^2} \right) \rightarrow 0$$

Perfect cancellation for any parameter choice !



Direct detection = loop-suppressed

Annihilation = unsuppressed



Excellent WIMP

(from 60 GeV to 10 TeV)

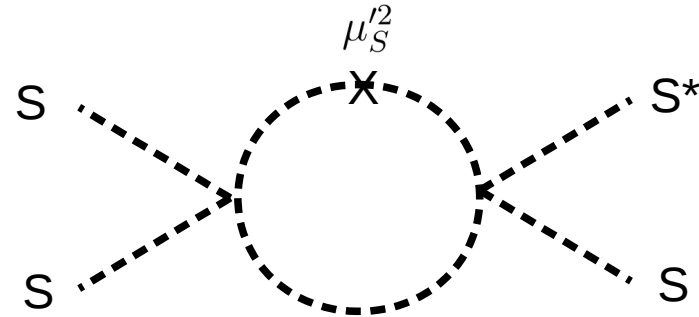
Loop corrections to the potential:

$$V_1 = \frac{\lambda'_{HS}}{2} |H|^2 S^2 + \frac{\lambda''_S}{4} |S|^2 S^2 + \frac{\lambda'_S}{4} S^4 + \text{h.c.}$$

$$\lambda'_{HS} = \frac{\lambda_{HS} \lambda_S}{32\pi^2} \ln \frac{\mu_S^2 + \mu_S'^2}{\mu_S^2 - \mu_S'^2},$$

$$\lambda''_S = \frac{\lambda_S^2}{8\pi^2} \ln \frac{\mu_S^2 + \mu_S'^2}{\mu_S^2 - \mu_S'^2},$$

$$\lambda'_S = \frac{\lambda_S^2}{64\pi^2} \left(\frac{\mu_S^2}{\mu_S'^2} \ln \frac{\mu_S^2 - \mu_S'^2}{\mu_S^2 + \mu_S'^2} + 2 \right)$$



$$\sigma \sim 10^{-49} \text{ cm}^2$$

Refined by

Azevedo, Duch, Grzadkowski, Huang, Iglicki, Santos '18

Ishiwata, Toma '18

Couplings as spurions:

$$\text{gauged U(1) + S + } \Phi : \left\{ \begin{array}{l} q_S = \text{even} \\ q_\Phi = \text{odd} \end{array} \right.$$

$$\Delta V \sim a S^2 \Phi^n + b S^4 \Phi^{2n} + \dots \quad (\text{no odd powers of S})$$

Here

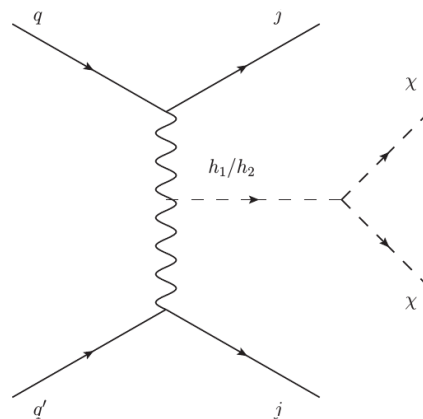
$$n \equiv -2 \frac{q_S}{q_\Phi}, \quad \epsilon \equiv \frac{\langle \Phi \rangle}{\Lambda}$$

$$\underbrace{\mu_S'^2 \sim \langle \Phi \rangle^2 \epsilon^{n-2}}_{\text{significant}}, \quad \underbrace{\lambda'_{HS} \sim \lambda''_S \sim \epsilon^n, \quad \lambda'_S \sim \epsilon^{2n}}_{\text{small}}$$

LHC search

Main signature: missing energy (via Higgs production)

VBF:

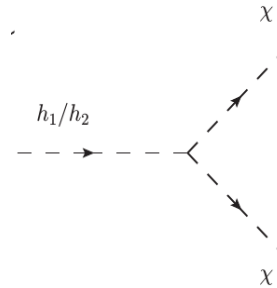


BSM states: $h_2, \chi + (h_1)$

The signal is significant only for

$$m_{h_2} \text{ or } m_{h_1} > 2 m_\chi$$

invisible decay



Constraints:

- h_1 Higgs signal strength

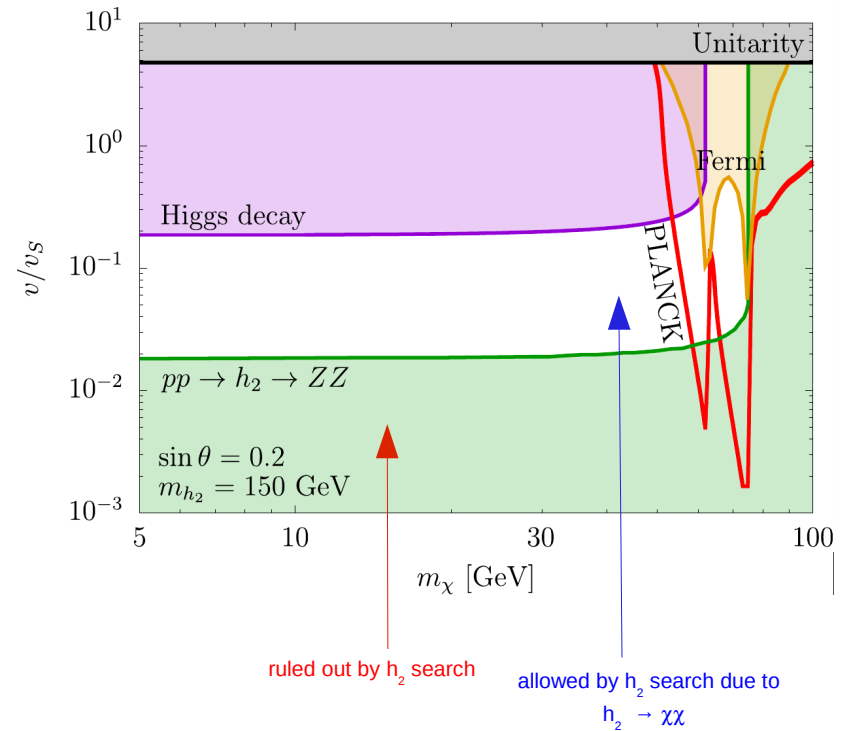
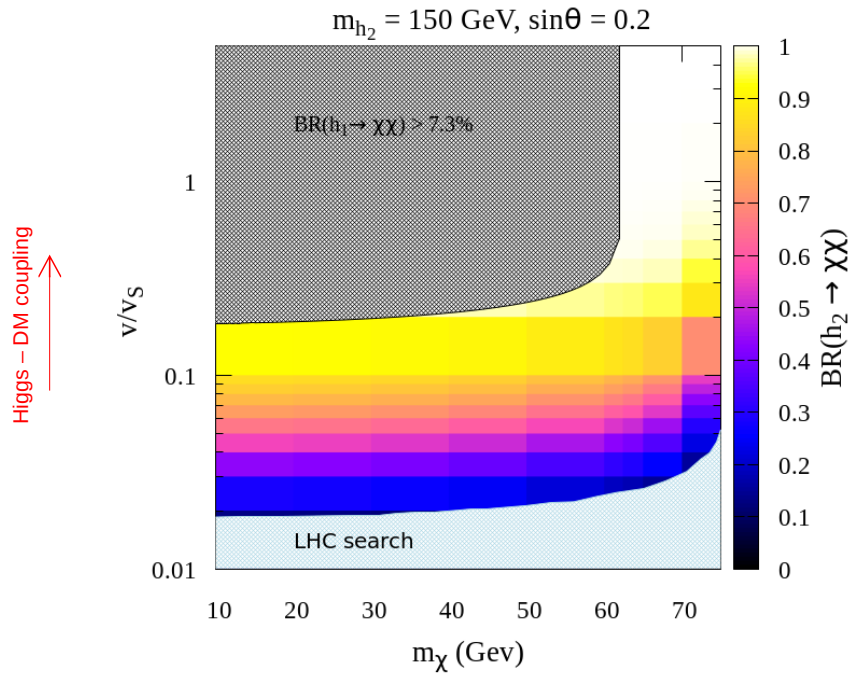
$$\mu = \cos^2 \theta (1 - BR_{inv})$$

└─> θ - dependent bound on BR_{inv}

- h_2 search

Example:

Use $\mu = 1.09^{+0.11}_{-0.10}$



- light h_2 allowed (e.g. ~ 150 GeV)
- h_2 invisible decay very efficient

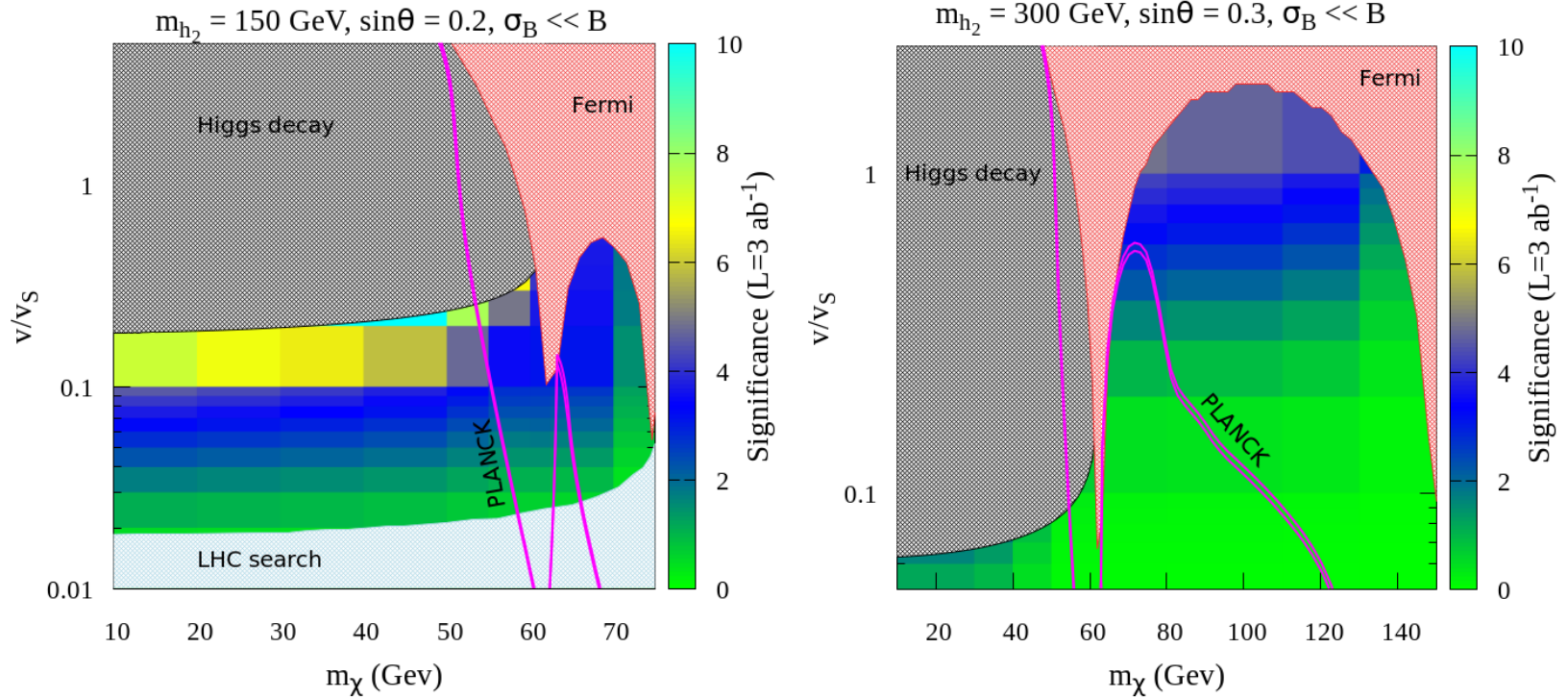
Signal selection (CMS, 13 TeV):

Observables	Requirements
Leading (trailing) jet	$p_T > 80(40) \text{ GeV}, \eta < 4.7$
\cancel{E}_T	$> 250 \text{ GeV}$
$\Delta\phi(\vec{\cancel{E}}_T, \vec{p}_T^{\text{jet}})$	> 0.5
$ \Delta\phi_{jj} $	< 1.5
$\eta_{j1} \cdot \eta_{j2}$	< 0
$ \Delta\eta_{jj} $	> 4
$ m_{jj} $	$> 1300 \text{ GeV}$
Leptons	$N_{\mu,e} = 0, p_T > 10 \text{ GeV}, \eta < 2.4(2.5)$
τ leptons	$N_{\tau_h} = 0, p_T > 10 \text{ GeV}, \eta < 2.3$
Photons	$N_\gamma = 0, p_T > 15 \text{ GeV}, \eta < 2.5$
b -jets	$N_b = 0, p_T > 20 \text{ GeV}, \eta < 2.4$

Background: Z + jets, W + jets, ...

$$S = \frac{S}{\sqrt{S+B+\sigma_B^2}}$$

Discovery potential at 3 ab⁻¹:



good prospects for non-thermal dark matter

Further studies:

- non–Abelian generalization Karamitros '19
 - gravitational waves Kannike, Raidal '19
-

Conclusion

- very simple DM model: *SM + complex singlet*
- natural direct detection suppression: *pseudo-Goldstone nature*
- light DM $< O(100)$ GeV can be probed at HL-LHC in *VBF + \cancel{E}_T*