

Implications for New Physics in $b \rightarrow s\mu\mu$ transitions after recent measurements by Belle and LHCb

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based on work with D.Kumar and E.M.Sessolo

arXiv: 1903.10932

*22nd International Conference From the Planck Scale to the Electroweak Scale
PLANCK 2019*



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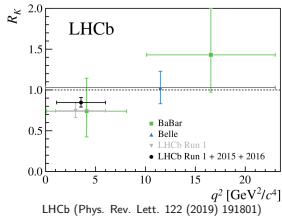
Granada, 04 June 2019



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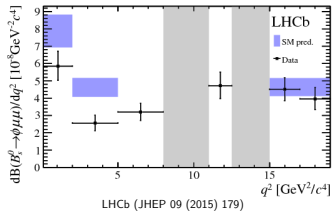
Deviations from the SM in $b \rightarrow s$ transitions

$$R_K = \frac{\text{BR}(B^+ \rightarrow K^+ \mu^+ \mu^-)}{\text{BR}(B^+ \rightarrow K^+ e^+ e^-)}$$



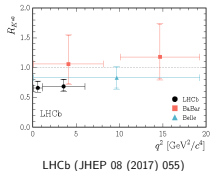
$\sim 2.5\sigma$ **NEW!**

$$\frac{d\text{BR}(B_s^0 \rightarrow \phi \mu^+ \mu^-)}{dq^2}$$

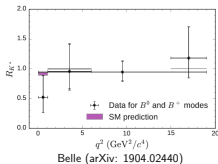


$\sim 3.7\sigma$

$$R_{K^*} = \frac{\text{BR}(B \rightarrow K^* \mu^+ \mu^-)}{\text{BR}(B \rightarrow K^* e^+ e^-)}$$

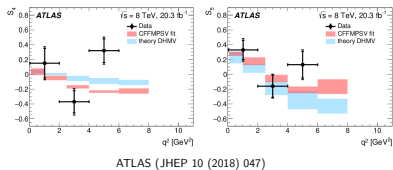


$\sim 2.5\sigma$



$\sim 0.5\sigma$ **NEW!**

angular observables in $B^0 \rightarrow K^{*0} \mu^+ \mu^-$



$\sim 3.0\sigma$

- The effective Hamiltonian for $b \rightarrow sll$ transitions

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \frac{e^2}{16\pi^2} \sum_{i,l} (C_i^l O_i^l + C_i^{\prime l} O_i^{\prime l}) + \text{h.c.}$$

- We assume the presence of NP in **semileptonic operators**:

$$\begin{aligned} O_9^l &= (\bar{s}_L \gamma^\mu b_L)(\bar{l} \gamma_\mu l), & O_9^{\prime l} &= (\bar{s}_R \gamma^\mu b_R)(\bar{l} \gamma_\mu l), \\ O_{10}^l &= (\bar{s}_L \gamma^\mu b_L)(\bar{l} \gamma_\mu \gamma_5 l), & O_{10}^{\prime l} &= (\bar{s}_R \gamma^\mu b_R)(\bar{l} \gamma_\mu \gamma_5 l) \end{aligned}$$

- NP in scalar and pseudoscalar operators $O_S^{(l)}$ and $O_P^{(l)}$ are severely constrained by the $B_s \rightarrow \mu^+ \mu^-$ measurements.

R. Alonso et al. PRL 113(2014) 241802, W. Altmannshofer et al. JHEP 05 (2017) 076.

- NP in electromagnetic dipole operator $O_7^{(l)}$ is tightly constrained by radiative decays.

A. Paul et al. JHEP 04 (2017) 027

- **other recent fits**: J. Aebischer et al. arXiv 1903.10434, M. Alguero et al., arXiv:1903.09578, A. K. Alok et al., arXiv:1903.09617, M. Ciuchini et al., arXiv:1903.09632, A. Datta et al. arXiv:1903.10086, A. Arbey et al.

arXiv:1904.08339

- Bayes' theorem

$$p(\mathbf{m}|\mathbf{d}) = \frac{p(\mathbf{d}|\xi(\mathbf{m}))\pi(\mathbf{m})}{p(\mathbf{d})}$$

- **posterior pdf** (*probability density function*)

$$p(\psi_{i=1,\dots,r}|\mathbf{d}) = \int p(\mathbf{m}|\mathbf{d})d^{n-r}m$$

- **likelihood function**

$$p(\mathbf{d}|\xi(\mathbf{m})) \equiv \mathcal{L}(\mathbf{m}) = \exp \left\{ -\frac{1}{2} [\mathcal{O}_{\text{th}}(\mathbf{m}) - \mathcal{O}_{\text{exp}}]^T (\mathcal{C}^{\text{exp}} + \mathcal{C}^{\text{th}})^{-1} [\mathcal{O}_{\text{th}}(\mathbf{m}) - \mathcal{O}_{\text{exp}}] \right\}$$

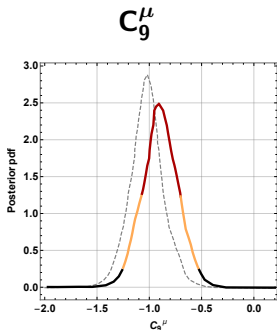
- **evidence** → comparison of different models through the *Bayes factor*

$$p(\mathbf{d})_{\mathcal{M}_1} / p(\mathbf{d})_{\mathcal{M}_2}$$

140 observables included in the likelihood function ($B^0 \rightarrow K^{*0} l^+ l^-$, $B^+ \rightarrow K^+ l^+ l^-$, $B^+ \rightarrow K^{*+} \mu^+ \mu^-$, $B^0 \rightarrow K^0 \mu^+ \mu^-$, $B_s^0 \rightarrow \phi \mu^+ \mu^-$, $\Lambda_b^0 \rightarrow \Lambda \mu^+ \mu^-$, $\text{BR}(B_s^0 \rightarrow X_s l^+ l^-)$, $\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-)$)

Parameter	Range	Prior
C_9^μ	$(-3, 3)$	Flat
$C_9^\mu = -C_{10}^\mu$	$(-3, 3)$	Flat
C_9^μ, C_{10}^μ	$(-3, 3)$	Flat
$C_9^\mu, C_9^{\prime\mu}$	$(-3, 3)$	Flat
$C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}$	$(-3, 3)$	Flat
$C_9^\mu, C_{10}^\mu, C_9^e, C_{10}^e$	$(-3, 3)$	Flat
$C_9^\mu, C_9^{\prime\mu}, C_9^e, C_9^{\prime e}$	$(-3, 3)$	Flat
$C_9^\mu, C_9^{\prime\mu}, C_{10}^\mu, C_{10}^{\prime\mu}$ $C_9^e, C_9^{\prime e}, C_{10}^e, C_{10}^{\prime e}$	$(-3, 3)$	Flat
CKM matrix element V_{cb}	$(4.22, 0.08) \times 10^{-2}$	Gaussian

1 Wilson coefficient



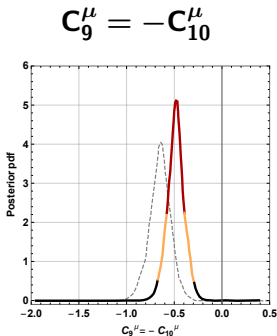
$$-\ln \mathcal{Z} = \mathbf{77.3}, \text{ pull } \mathbf{4.7 \sigma} \text{ (} 5.0 \sigma \text{)}$$

$$\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-) = 3.5 \times 10^{-9}$$

$$\chi_{R_K}^2 = 0.3, \chi_{R_{K^*}}^2 = 8.9$$

both predict $R_K \approx R_{K^*}$

$$\text{Bayes factor } \frac{\mathcal{Z}_{C_9^\mu}}{\mathcal{Z}_{C_9^\mu = -C_{10}^\mu}} = \mathbf{1.2} \text{ (1/4)}$$



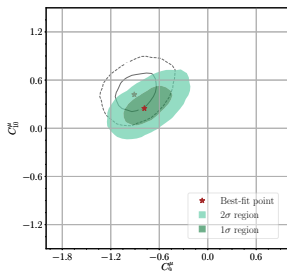
$$-\ln \mathcal{Z} = \mathbf{77.5}, \text{ pull } \mathbf{4.8 \sigma} \text{ (} 5.3 \sigma \text{)}$$

$$\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-) = 2.8 \times 10^{-9}$$

$$\chi_{R_K}^2 = 1.6, \chi_{R_{K^*}}^2 = 7.0$$

2 Wilson coefficients

C_9^μ, C_{10}^μ



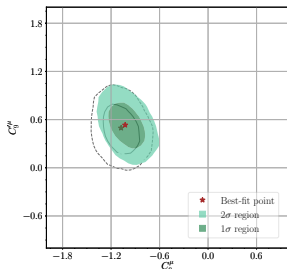
$$-\ln \mathcal{Z} = \mathbf{77.6}, \text{ pull } \mathbf{4.7} \sigma \text{ (} 5.3 \sigma \text{)}$$

$$\chi_{R_K}^2 = 1.5, \chi_{R_{K^*}}^2 = 7.6$$

$R_K \approx R_{K^*}$ (tension arises)

$$\text{Bayes factor } \frac{\mathcal{Z}_{C_9^\mu, C_9^{\prime\mu}}}{\mathcal{Z}_{C_9^\mu, C_{10}^\mu}} = 6 \quad (\text{positive}) \quad (\text{old: } 1/2)$$

$C_9^\mu, C_9^{\prime\mu}$



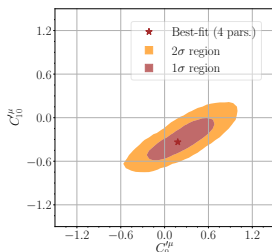
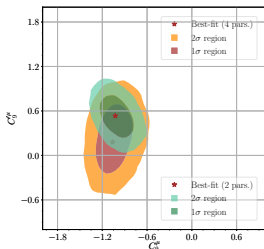
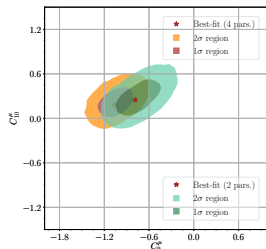
$$-\ln \mathcal{Z} = \mathbf{75.8}, \text{ pull } \mathbf{5.0} \sigma \text{ (} 5.2 \sigma \text{)}$$

$$\chi_{R_K}^2 = 0.5, \chi_{R_{K^*}}^2 = 7.3$$

$R_K \not\approx R_{K^*}$

4 vs. 2 Wilson coefficients

$$C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}$$



- perfect fit in R_K and $R_{K^*}^{[1.1,6]}$ ($\chi_{R_K}^2 = 0.0$, $\chi_{R_{K^*}}^2 = 6.8$)
- $\overline{\text{BR}}(B_s^0 \rightarrow \mu^+ \mu^-) = 2.8 \times 10^{-9}$
- larger negative C_9^μ w.r.t. (C_9^μ, C_{10}^μ)
- $C_9^{\prime\mu} \leq 0$ allowed w.r.t. $(C_9^\mu, C_9^{\prime\mu})$

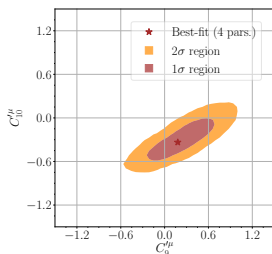
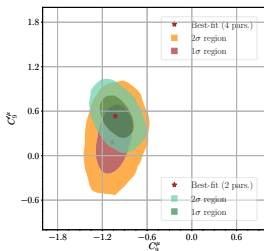
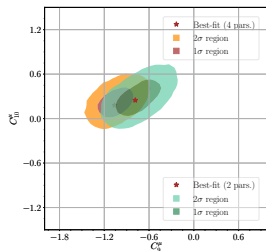
G.Hiller, M.Schmaltz JHEP 02 (2015) 055

$$R_K \approx 1 + 0.24 (C_9^\mu - C_{10}^\mu) + C_9^{\prime\mu} - C_{10}^{\prime\mu}$$

$$R_{K^*} \approx 1 + 0.24 (C_9^\mu - C_{10}^\mu) - 0.17 (C_9^{\prime\mu} - C_{10}^{\prime\mu})$$

4 vs. 2 Wilson coefficients

$C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}$



$-\ln \mathcal{Z} = 76.0$, pull 5.1σ

$$\frac{\mathcal{Z}_{C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}}}{\mathcal{Z}_{C_9^\mu, C_{10}^\mu}} = 5.0 \quad (\text{positive})$$

$$\frac{\mathcal{Z}_{C_9^\mu, C_9^{\prime\mu}}}{\mathcal{Z}_{C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}}} = 1.2 \quad (\text{barely worth mentioning})$$

Input parameters	$-\ln \mathcal{Z}$	Pull	χ^2_{TOT}	$\frac{\chi^2_{\text{TOT}}}{d.o.f}$	χ^2_{μ}	$\chi^2_{R_K}$	$\chi^2_{R_{K^*}}$
C_9^μ, C_{10}^μ	77.6	4.7σ	146.1	1.05	130.3	1.5	7.6
$C_9^\mu, C_9^{\prime\mu}$	75.8	5.0σ	142.3	1.02	127.6	0.5	7.3
$C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}$	76.0	5.1σ	136.8	1.00	123.2	0.0	6.8
$C_9^\mu, C_{10}^\mu, C_9^e, C_{10}^e$	78.0	4.5σ	142.7	1.04	129.8	0.1	5.8
$C_9^\mu, C_9^e, C_9^{\prime\mu}, C_9^{\prime e}$	77.7	4.6σ	141.6	1.03	127.2	0.2	6.7
$C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu}$ $C_9^e, C_{10}^e, C_9^{\prime e}, C_{10}^{\prime e}$	78.3	4.4σ	135.4	1.02	123.3	0.2	5.4

No improvement w.r.t fit with the muon-only Wilson coefficients

Heavy Z'

- the most generic Lagrangian, parametrizing couplings of Z' to b, s, μ

$$\mathcal{L} \supset Z'_\alpha \left(\Delta_L^{sb} \bar{s}_L \gamma^\alpha b_L + \Delta_R^{sb} \bar{s}_R \gamma^\alpha b_R + \text{H.c.} \right) + Z'_\alpha \left(\Delta_L^{\mu\mu} \bar{\mu}_L \gamma^\alpha \mu_L + \Delta_R^{\mu\mu} \bar{\mu}_R \gamma^\alpha \mu_R \right)$$

- the corresponding Wilson coefficients

$$C_{9,\text{NP}}^\mu = -\frac{\Delta_L^{sb}(\Delta_L^{\mu\mu} + \Delta_R^{\mu\mu})}{V_{tb}V_{ts}^*} \left(\frac{\Lambda_v}{m_{Z'}}\right)^2 \quad C_{9,\text{NP}}^{\prime\mu} = -\frac{\Delta_R^{sb}(\Delta_L^{\mu\mu} + \Delta_R^{\mu\mu})}{V_{tb}V_{ts}^*} \left(\frac{\Lambda_v}{m_{Z'}}\right)^2$$

$$C_{10,\text{NP}}^\mu = -\frac{\Delta_L^{sb}(\Delta_R^{\mu\mu} - \Delta_L^{\mu\mu})}{V_{tb}V_{ts}^*} \left(\frac{\Lambda_v}{m_{Z'}}\right)^2 \quad C_{10,\text{NP}}^{\prime\mu} = -\frac{\Delta_R^{sb}(\Delta_R^{\mu\mu} - \Delta_L^{\mu\mu})}{V_{tb}V_{ts}^*} \left(\frac{\Lambda_v}{m_{Z'}}\right)^2$$

- Z' is the gauge boson of $U(1)_X \rightarrow$ its couplings to the gauge eigenstates must be flavor-conserving $\rightarrow \Delta_L^{sb}$ and Δ_R^{sb} from VLF



Model 1 P.Fox et al. Phys. Rev. D84 (2011) 115006, C. Bobeth et al. JHEP 04 (2017) 079

- VL quarks $Q^{(\prime)} : (\mathbf{3}, \mathbf{2}, 1/6, -1)$, $D^{(\prime)} : (\bar{\mathbf{3}}, \mathbf{1}, 1/3, -1)$
- scalar singlet $S : (\mathbf{1}, \mathbf{1}, 0, -1)$
- SM leptons: $Q(e) = 0$, $Q(\mu_{R/L}) = -1/1$, $Q(\tau_{L/R}) = -1/1$

SM and BSM mixing given by

$$\mathcal{L} \supset (-\lambda_{Q,i} S Q' q_i - \lambda_{D,i} S D' d_{R,i} + \text{H.c.}) - M_Q Q' Q - M_D D' D$$

effective Wilson coefficients ($v_S = m_{Z'}/g_X$)

$$C_{9,\text{NP}}^\mu = \frac{2\Lambda_v^2}{V_{tb} V_{ts}^*} \frac{\lambda_{Q,2} \lambda_{Q,3}}{2M_Q^2 + (\lambda_{Q,2}^2 + \lambda_{Q,3}^2) v_S^2}, \quad C_{10,\text{NP}}^\mu = 0$$

$$C'_{9,\text{NP}}{}^\mu = -\frac{2\Lambda_v^2}{V_{tb} V_{ts}^*} \frac{\lambda_{D,2} \lambda_{D,3}}{2M_D^2 + (\lambda_{D,2}^2 + \lambda_{D,3}^2) v_S^2}, \quad C'_{10,\text{NP}}{}^\mu = 0$$

Heavy Z' with $L_\mu - L_\tau$ symmetry

Model 2 W. Altmannshofer et al. Phys. Rev. D94 no. 9, (2016) 095026

- VL quarks $Q^{(\prime)} : (\mathbf{3}, \mathbf{2}, 1/6, -1)$, $E^{(\prime)} : (\mathbf{1}, \mathbf{1}, 1, 0)$
- scalar singlet $S : (\mathbf{1}, \mathbf{1}, 0, -1)$
- SM leptons: $Q(e) = 0$, $Q(\mu_{R/L}) = -1/1$, $Q(\tau_{L/R}) = -1/1$

SM and BSM mixing given by

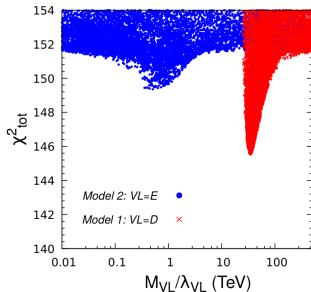
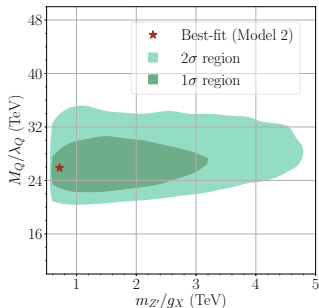
$$\mathcal{L} \supset \left(-\lambda_{E,2} S^* E' \mu_R - \lambda_{E,3} S E' \tau_R - \tilde{Y}_E \phi^\dagger I_1 E + \text{H.c.} \right) - M_E E' E$$

effective Wilson coefficients ($v_S = m_{Z'}/g_X$)

$$C_{9,\text{NP}}^\mu = \frac{\Lambda_v^2}{V_{tb} V_{ts}^*} \left(\frac{\lambda_{Q,2} \lambda_{Q,3}}{2M_Q^2 + (\lambda_{Q,2}^2 + \lambda_{Q,3}^2) v_S^2} \right) \left(1 + \frac{2M_E^2}{2M_E^2 + \lambda_{E,2}^2 v_S^2} \right), \quad C_{9,\text{NP}}^{\prime\mu} = 0$$

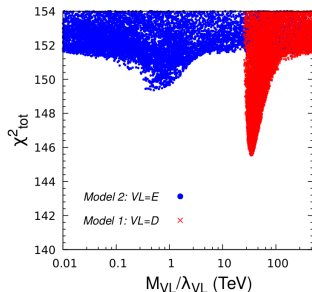
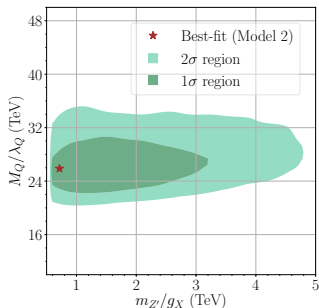
$$C_{10,\text{NP}}^\mu = \frac{\Lambda_v^2}{V_{tb} V_{ts}^*} \left(\frac{\lambda_{Q,2} \lambda_{Q,3}}{2M_Q^2 + (\lambda_{Q,2}^2 + \lambda_{Q,3}^2) v_S^2} \right) \left(-1 + \frac{2M_E^2}{2M_E^2 + \lambda_{E,2}^2 v_S^2} \right), \quad C_{10,\text{NP}}^{\prime\mu} = 0$$

Heavy Z' with $L_\mu - L_\tau$ symmetry



- VLQ mass range is determined by the 2σ range in $C_{9,\text{NP}}^\mu$
- VLD and VLE mass is unbounded from the above at the 2σ level ($C_{9,\text{NP}}^{\prime\mu}$ (model 1) and $C_{10,\text{NP}}^\mu$ (model 2) are consistent with the zero at the 2σ)
- $m_{Z'}/g_X$ is limited to values below 5 TeV as a result of the B_s mixing constraint

Heavy Z' with $L_\mu - L_\tau$ symmetry



$Z' + VL$	$-\ln \mathcal{Z}$	Pull	$m_{Z'}/g_X$	M_Q/λ_Q	M_{VL}/λ_{VL}
Model 1	78.4	4.5σ	0.7 TeV	24.4 TeV	34.2 TeV
Model 2	80.0	4.1σ	0.7 TeV	24.7 TeV	0.5 TeV

$$\frac{\mathcal{Z}_{\text{Model 1}}}{\mathcal{Z}_{\text{Model 2}}} = 5.0 \quad (\text{positive})$$

- Global Bayesian analysis of NP effects with new measurements of R_K and R_{K^*} in Morionod 2019.
- R_K shifts closer to SM predictions hence (C_9^μ, C_{10}^μ) shifts slightly towards zero.
- Pull w.r.t. SM still at 5σ (!).
- $(C_9^\mu, C_9^{\prime\mu})$ and $(C_9^\mu, C_{10}^\mu, C_9^{\prime\mu}, C_{10}^{\prime\mu})$ are favored by the data.