Conformal invariance versus Weyl invariance

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Física

In the UV masses should be unimportant \rightarrow CFT

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Intuitive idea of scale invariance, no preferred length

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Three realizations of this symmetry

Scale invariance (in flat spacetime)

$$
x^{\mu} \to \lambda x^{\mu}
$$

Conformal invariance (in flat spacetime)

Conformal invariance in curved space, Weyl Symmetry

Conformal group

$$
g_{\mu\nu} \to \Omega(x)^2 \, g_{\mu\nu}
$$

Scale invariance Conformal invariance

Weyl invariance

Scale invariance

Conformal invariance

Rosenfeld energy momentum tensor

 $\eta^{\mu\nu}T_{\mu\nu}=\partial_{\mu}V^{\mu}_{\lambda}$ Virial Weyl
invariance (on-shell) $\eta^{\mu\nu} I_{\mu\nu} = \theta_\mu V^\mu$ invariance

Conserved scale current

$$
j^{\mu} = x^{\lambda} T^{\mu}_{\lambda} - V^{\mu}
$$

[Callan, Coleman & Jackiw]

Scale invariance

Conformal invariance

The virial is a total derivative

Weyl invariance

$$
V^\mu = \partial_\nu \sigma^{\mu\nu}
$$

Conserved current

$$
K^{\mu\nu} = \left(2x^{\nu}x_{\rho} - x^2\delta^{\nu}_{\rho}\right)T^{\rho\mu} - 2x^{\nu}V^{\mu} + 2\sigma^{\mu\nu}
$$

The EM tensor can be improved

$$
\Theta^{\mu\nu} = T^{\mu\nu} + \frac{1}{2} \partial_{\lambda} \partial_{\rho} X^{\lambda \rho \mu \nu} \times X = X(\sigma)
$$

Conditions for scale inv. to become conformal invariants d=2 scale inv. implies conformal invariance for local theories [Zamolodchikov] [Nakayama] [Polchinski] Counterexample 2d theory of elasticity [Riva & Cardy]

d=4 not known example of scale inv. but not conformal inv. [Coleman & Jackiw]

> Weyl invariance

Conformal invariance

These symmetries only make sense in flat spacetimes What happens when gravity is present?

> Weyl invariance

Scale invariance Conformal invariance

Weyl invariance

We can study this symmetry in the linear limit $g_{\mu\nu} = \eta_{\mu\nu} + \kappa\,h_{\mu\nu}$

$$
g_{\mu\nu} = \Omega(x)^2 g_{\mu\nu} \quad \rightarrow \quad \delta h_{\mu\nu} = \frac{2\omega(x)}{\kappa} \eta_{\mu\nu}
$$

Linearized symmetry LWeyl

Scale invariance **Conformal invariance Weyl invariance** ?

The scalar field conformal coupling to gravity is an easy example

Improvement of the EM tensor Conformal and Weyl invariant coupling of scalar field to gravity

$$
\int d^4x \left\{ \frac{1}{6} \phi^2 R + (\partial_\mu \phi)^2 \right\}
$$

Our aim is to explore the difference between this symmetries for the spin 2 field

We study operators that come from the weak field expansion of linear and quadratic theories of gravity

We study the gravitational field as a fluctuation around flat spacetime

$$
g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}
$$

Symmetric tensor representing the graviton

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Up to quadratic order in the fluctuations

Dimension 4 operators **Dimension 6 operators**

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Symmetric tensor representing the graviton

Up to quadratic order in the fluctuations

Dimension 4 operators **Dimension 6 operators**

$$
\begin{aligned}\n\mathcal{D}_1 &= \frac{1}{4} \partial_{\mu} h_{\nu \rho} \partial^{\mu} h^{\nu \rho} \\
\mathcal{D}_2 &= -\frac{1}{2} \partial^{\lambda} h_{\lambda \rho} \partial_{\sigma} h^{\rho \sigma} \\
\mathcal{D}_3 &= \frac{1}{2} \partial^{\lambda} h \partial^{\sigma} h_{\lambda \sigma} \\
\mathcal{D}_4 &= -\frac{1}{4} \partial_{\mu} h \partial^{\mu} h\n\end{aligned}
$$
\n
$$
\begin{aligned}\nS_{lin.} &= \int d^4 x \sum_{i=1}^4 \alpha_i \mathcal{D}_i \\
S_{lin.} &= \int d^4 x \sum_{i=1}^4 \alpha_i \mathcal{D}_i\n\end{aligned}
$$

$$
\begin{aligned}\n\mathcal{O}_1 &= h_{\alpha\beta} \left(\partial^{\alpha} \partial^{\beta} \partial^{\gamma} \partial^{\delta} \right) h_{\gamma\delta} \\
\mathcal{O}_2 &= h_{\alpha\beta} \left(\partial^{\alpha} \partial^{\beta} \eta^{\gamma\delta} \Box \right) h_{\gamma\delta} \\
\mathcal{O}_3 &= h_{\alpha\beta} \left(\partial^{\alpha} \partial^{\gamma} \eta^{\beta\delta} \Box \right) h_{\gamma\delta} \\
\mathcal{O}_4 &= h_{\alpha\beta} \left(\eta^{\alpha\gamma} \eta^{\beta\delta} \Box^2 \right) h_{\gamma\beta} \\
\mathcal{O}_5 &= h_{\alpha\beta} \left(\eta^{\alpha\beta} \eta^{\gamma\delta} \Box^2 \right) h_{\gamma\delta}\n\end{aligned}
$$

[Alvarez, Blas, Garriga & Verdaguer]

We study the gravitational field as a fluctuation around flat spacetime

$$
g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}
$$

Symmetric tensor representing the graviton

Up to quadratic order in the fluctuations

Dimension 4 operators **Dimension 6 operators**

Come from linearization of theories linear in curvature

$$
\int d^4x \left(\sqrt{g}R\right)^{O(\kappa^2)}
$$

Come from linearization of theories quadratic in curvature

$$
\int d^4x \left[\sqrt{g} (\alpha R^2 + \beta R_{\mu\nu}^2 + \gamma R_{\mu\nu\rho\sigma}^2) \right]^{O(\kappa^2)}
$$

Invariance under *δhμν* = ∂*μξν* + ∂*νξμ*

Conditions $\alpha_1 = \alpha_2, \quad \alpha_2 = \alpha_3, \quad \alpha_3 = \alpha_4$

Conditions $\alpha_1 + \alpha_3 = n\alpha_4$, $2\alpha_2 = n\alpha_3$

Only theory with LDiff and linear in curvature

$$
S_{EH} = \int d^4x \left(\sqrt{g}R\right)^{O(\kappa^2)}
$$

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LDiff and LWeyl symmetries are incompatible in $n=4$

LDiff **LTDiff LWeyl** Conformal

Only theory with LDiff and linear in curvature

$$
S_{EH} = \int d^4x \left(\sqrt{g}R\right)^{O(\kappa^2)}
$$

LDiff and LWeyl symmetries are incompatible in $n=4$

LTDiff and LWeyl compatible giving

$$
S_{WTDiff} = \int d^n x \left[g^{1/n} \left(R + \frac{(n-1)(n-2)}{4n^2} \frac{(\nabla g)^2}{g^2} \right) \right]^{O(\kappa^2)}
$$

Description of Unimodular Gravity S_{EH} [$g^{-1/n}g_{\mu\nu}$]

We can infer the theories that are invariant under this symmetries with the lowest order

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In this case

$$
\sigma^{\mu\nu} = \left(3 - \frac{n}{2}\right) \left\{ \frac{\alpha_1}{8} \eta^{\mu\nu} h^{\alpha\beta} h_{\alpha\beta} - \frac{\alpha_2}{4} h^{\rho\mu} h^{\nu}_{\rho} + \frac{\alpha_4}{4} h^{\mu\nu} h - \frac{\alpha_5}{8} \eta^{\mu\nu} h^2 \right\}
$$

No constraints on the coupling constants

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At this order, LWeyl is more restrictive than Conformal invariance

Invariance under *δhμν* = ∂*μξν* + ∂*νξμ*

Conditions $2g_1 + g_2 + g_3 = 0$, $g_2 + 2g_5 = 0$, $g_3 + 2g_4 = 0$

Theories LDiff and quadratic in curvature

$$
\int d^4x \left[\sqrt{g} (\alpha R^2 + \beta R_{\mu\nu}^2 + \gamma R_{\mu\nu\rho\sigma}^2) \right]^{O(\kappa^2)}
$$

Theories LDiff and quadratic in curvature

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\int d^4x \left[\sqrt{g} (\alpha R^2 + \beta R_{\mu\nu}^2 + \gamma R_{\mu\nu\rho\sigma}^2) \right]^{O(\kappa^2)}
$$

LDiff and LWeyl symmetries can be compatible

$$
\int d^n x \sqrt{g} \left\{ \left(\alpha - \frac{(n-1)(n-2)}{2(n-3)} \right) E_4 + g_2 \frac{(n-1)(n-2)}{2(n-3)} W_n \right\}
$$

Theories LDiff and quadratic in curvature

$$
\int d^4x \left[\sqrt{g} (\alpha R^2 + \beta R_{\mu\nu}^2 + \gamma R_{\mu\nu\rho\sigma}^2) \right]^{O(\kappa^2)}
$$

LDiff and LWeyl symmetries can be compatible

$$
\int d^n x \sqrt{g} \left\{ \left(\alpha - g_2 \frac{(n-1)(n-2)}{2(n-3)} \right) \frac{E_4}{2} + g_2 \frac{(n-1)(n-2)}{2(n-3)} \frac{W_n}{2}
$$
\nGauss-Bonet

\nWeyl tensor

\ndensity

\nSquared

\n
$$
E_4 \equiv R_{\alpha\beta\gamma\delta}^2 - 4R_{\alpha\beta}^2 + R^2 \qquad W_n \equiv R_{\alpha\beta\gamma\delta}^2 - \frac{4}{n-2} R_{\alpha\beta}^2 + \frac{2}{(n-1)(n-2)} R^2
$$
\n(total derivative)

LTDiff and LWeyl compatible giving

$$
S_{WTDiff} = \int d^n x \sqrt{\tilde{g}} \left\{ \alpha \tilde{R}_{\mu\nu\rho\sigma} \tilde{R}^{\mu\nu\rho\sigma} + \beta \tilde{R}_{\mu\nu} \tilde{R}^{\mu\nu} + \gamma \tilde{R}^2 \right\}
$$

where $\tilde{g}_{\mu\nu} = g^{-1/n} g_{\mu\nu}$

LTDiff and LWeyl compatible giving

$$
S_{WTDiff} = \int d^{n}x \sqrt{\tilde{g}} \left\{ \alpha \tilde{R}_{\mu\nu\rho\sigma} \tilde{R}^{\mu\nu\rho\sigma} + \beta \tilde{R}_{\mu\nu} \tilde{R}^{\mu\nu} + \gamma \tilde{R}^{2} \right\}
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where $\tilde{g}_{\mu\nu} = g^{-1/n} g_{\mu\nu}$

We need to see if we can write the EM tensor as $T = \partial_{\mu} \partial_{\nu} \sigma^{\mu\nu}$

In this case

$$
\sigma^{\mu\nu} = \left(4 - \frac{n}{2}\right) \left\{ \frac{g_1}{2} \left(h^{\mu\alpha} \partial_\alpha \partial^\beta h^\nu_\beta + h^{\nu\alpha} \partial_\alpha \partial^\beta h^\mu_\beta \right) + \frac{g_2}{2} \left(h^{\mu\nu} \Box h + h \Box h^{\mu\nu} \right) + \frac{g_3}{2} \left(h^{\mu\lambda} \Box h^\nu_\lambda + h^{\nu\lambda} \Box h^\mu_\lambda \right) \right\}
$$

+ $g_4 \eta^{\mu\nu} h_{\alpha\beta} \Box h^{\alpha\beta} + g_5 \eta^{\mu\nu} h \Box h \right\}$

We cannot construct an scale invariant theory with dimension 4 and dimension 6 operators altogether

LDiff LTDiff LWeyl **Conformal**

We cannot construct an scale invariant theory with dimension 4 and dimension 6 operators altogether

For example the following Lagrangian is not even scale invariant

$$
L' = \alpha \phi \Box \phi + \frac{\beta}{M^2} \phi \Box^2 \phi = \alpha L_1 + \beta L_2
$$

As we have

As we have
\n
$$
T = -\alpha \left(1 - \frac{n}{2}\right) L_1 - \frac{\beta}{M^2} \left(2 - \frac{n}{2}\right) L_2 \neq \partial_\mu V^\mu
$$
\nDifferent factors

$$
\alpha \Box \phi + \frac{\beta}{M^2} \Box^2 \phi = 0
$$

Conformal invariance and Weyl invariance are different symmetries

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Local Weyl invariant theories of gravity are very restricted

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What about global Weyl invariance?

The lowest order is not enough

Can only be analyzed order by order (as conformal)

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Different Ward identities for Conformal and Weyl. Also, sensitive to the dynamical character of gravity?

LDiff LWeyl **Conformal** LTDiff

In this case

$$
\sigma^{\mu\nu} = \left(3 - \frac{n}{2}\right) \left\{ \frac{\alpha_1}{8} \eta^{\mu\nu} h^{\alpha\beta} h_{\alpha\beta} - \frac{\alpha_2}{4} h^{\rho\mu} h^{\nu}_{\rho} + \frac{\alpha_4}{4} h^{\mu\nu} h - \frac{\alpha_5}{8} \eta^{\mu\nu} h^2 \right\}
$$

We can find an improvement for any dimension

$$
\Theta^{\mu\nu} = T^{\mu\nu} + \frac{1}{2} \partial_{\lambda} \partial_{\rho} X^{\lambda \rho \mu \nu} \longrightarrow \Theta = 0
$$

$$
X^{\lambda \rho \mu \nu} = g^{\lambda \rho} \sigma^{\mu \nu} + g^{\lambda \mu} \sigma^{\rho \nu} + g^{\lambda \nu} \sigma^{\mu \rho} + g^{\mu \nu} \sigma^{\lambda \rho} + \frac{1}{3} g^{\lambda \rho} g^{\mu \nu} \sigma^{\alpha}_{+\alpha} + \frac{1}{3} g^{\lambda \mu} g^{\rho \nu} \sigma^{\alpha}_{+\alpha}
$$

Backup

$$
0 = \delta Z = \int \mathcal{D}g_{\mu\nu} \mathcal{D}\phi \int d(vv) \,\omega(x) \left[-2g^{\mu\nu} \frac{\delta S}{\delta g^{\mu\nu}} - \frac{n-2}{2} \phi \frac{\delta S}{\delta \phi} \right]
$$

Unimodular Gravity (UG)

WTDiff reduces to UG in the gauge $g = 1$

^{*} Non invertible

Renormalization procedure — Introduction of an energy scale? Flat spacetime

Breaking of scale invariance summarised in β-functions

Some known theories with vanishing β -functions ($\mathcal{N}=2$ Supersymmetry)

Curved spacetime (background)

Anomaly in the trace of the energy (Conformal anomaly momentum tensor

Duff, Capper, Christensen, Fulling, Desser, Isham, Englert, Truffin, Gastman, Berends, Vilkovisky, Fradkin, Tseytlin,…

$$
T^{\mu}_{\mu} = a\left(W_4^2 + \frac{2}{3}\,\square R\right) + b\,E_4 + c\,F^a_{\mu\nu}F^{\mu\nu}_a
$$

Backup

Weyl operator in the Ward identity

$$
0 = \delta Z = \int \mathcal{D}g_{\mu\nu} \mathcal{D}\phi \int d(vv) \,\omega(x) \left[-2g^{\mu\nu} \frac{\delta S}{\delta g^{\mu\nu}} - \frac{n-2}{2} \phi \frac{\delta S}{\delta \phi} \right]
$$