



## Measurement of Directed Flow of Identified Particles in Au+Au √s<sub>NN</sub>=4.5 GeV Fixed-target Collisions at STAR

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- ✓ Recent flow analysis in the STAR fixed-target program
- ✓ Motivation
- ✓ Analysis method (EP method)
- ✓ p<sub>T</sub> dependence of directed flow
- Reprive the second s
- ✓ v₁ slope
- ✓ Summary and Outlook

# **STAR STAR fixed-target program (FXT)**



1.3M events from half hour test run top 30% central trigger. Au+Au  $\sqrt{s_{NN}}$ =4.5 GeV





## **STAR Previous flow results from FXT**

E895 PRL 84(2000) 5488 STAR PRL 112(2014) 162301



1.2

1.6

1.8 p<sub>\_</sub> (GeV/c)

0.02

STAR, Quark Matter 2018

# **STAR** Directed flow

Directed flow is ...  $v_1 = \langle \cos(\phi - \Psi) \rangle$ 

- ✓ Evaluated by the coefficients of the 1st harmonic in the Fourier expansion.
- ✓ Generated by the interaction between participants and spectators.
- Observable signature that was suggested to be sensitive to the first-order phase transition. (softest point)
- ✓ Possible probe of search for the QGP signature. (anti-flow)



# **STAR Directed flow analysis**



STAR PRL 120 (2018) 062301

- ✓ Opposite sign of dv₁/dy at midrapidity is observed for baryons and mesons at low energies.
- ✓ Minimum at √s<sub>NN</sub>=10-20 GeV for net baryon is observed.

This is called "softest point" and may be a possible sign of the first-order phase transition.

- Models cannot explain energy dependence of the directed flow
  - experiment:minimum at 10-20 GeV model:minimum at 4 GeV

H. Stocker NPA750, 121-147(2005)

# **STAR** Motivation

- The directed flow is an observable probe which suggested to be sensitive to the first-order phase transition.
- $\checkmark$  The fixed-target program extends the RHIC BES to higher  $\mu_B.$

To clarify the structure of the first-order phase transition, study the characteristics of directed flow at low energies is important.



# The Solenoidal Tracker At RHIC star

### Beam-Beam Counter

Time-Of-Flight detector

### Time Projection Chamber

Gold Target was installed inside the vacuum pipe at z = 211 cm





### **Data Set**

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- ✓ Au+Au : √s<sub>NN</sub>=4.5 GeV
- ✓ Test run conducted in 2015
- ✓ 1.3 central million events
- ✓ Midrapidity : -1.52

## **Event Selection**

- ✓ Vertex X = -1.5 to 1.0 cm
- ✓ Vertex Y = -2.5 to -1.0 cm
- ✓ Vertex Z = 210 to 212 cm

## Track Selection

- ✓ nHitsFit > 20
- ✓ nHitFit/nHitsPoss > 0.52



## **STAR Particle identification**

#### ۰п

- ✓ Inσ(π)I < 2</li>
  ✓ 0.2 < p<sub>T</sub> GeV/c
  ✓ p < 1.6 GeV/c</li>
  ✓ -0.15 < m<sup>2</sup> < 0.14 (GeV/c<sup>2</sup>)<sup>2</sup> (If TOF available)
   K
  ✓ Inσ(K)I < 2</li>
  ✓ p<sub>T</sub> < 2.0 GeV/c</li>
  - ✓  $0.14 < m^2 < 0.4$  (GeV/c<sup>2</sup>)<sup>2</sup> (If TOF available)

## Ρ

- ✓ Inσ(p)I < 2</p>
- ✓ 0.4 < p<sub>T</sub> < 2.0 GeV/c
- ✓  $0.4 < m^2 < 1.4$  (GeV/c<sup>2</sup>)<sup>2</sup> (If TOF available)
- ✓ If no hit in TOF, but satisfying other selection criteria, track is kept as a proton candidate.



# **STAR Deuteron identification**





1.Define event plane from the direction in which the generated particles are emitted.

$$\Psi_1 = \tan^{-1} \left( \frac{\Sigma w_i \sin \phi_i}{\Sigma w_i \cos \phi_i} \right)$$

2.Event plane distortion from detector non-uniformity and/or beam offset are corrected using the recentering and flattening methods.

3.Perform the Fourier expansion of the angular distribution of particles.

$$\frac{dN}{d(\phi-\Psi_n)} \propto 1 + \sum_{n=1}^{\infty} 2v_n \cos\left[n(\phi-\Psi_n)\right] \qquad \Longrightarrow \qquad v_1 = \langle \cos(\phi-\Psi) \rangle$$

4.Measured v<sub>n</sub> include the effect of the finite detector resolution that can be corrected using so-called event plane resolution correction.

I. Selyuzhenkov and S. Voloshin, PRC 77 (2008), 034904

A.M.Poskanzer, S.A.Voloshin, PRC 58 (1998), 1671-1678

## **STAR Event plane method : 2 or 3 subevents**

## 2 subevents

Divide into 2 groups using random number (group A and B)

Because A and B are essentially the same, we calculate resolution using 2 subevents.

### 3 subevents



Assuming a true event plane, calculate the resolution by taking correlations for each two of three regions.

$$Res_A = \sqrt{\frac{\sigma_{AB} \cdot \sigma_{AC}}{\sigma_{BC}}}$$

## **STAR EP correlation and resolution**

<cos(Ψ<sub>A</sub> - Ψ<sub>B</sub>)>

### **EP Correlation**

<sin(Ψ<sub>A</sub> - Ψ<sub>B</sub>)>



## **STAR Systematic uncertainty**

- 1. 2 TPC subevent planes divided at y = -0.8, and BBC east.
- 2. Randomly assigning particles to subevent A or B. Use charged particles.
- 3. Randomly assigning particles to subevent A or B. Use not protons.
- 4. Randomly assigning particles to subevent A or B. Use protons only.

### **3 subevents**

$$Res_A = \sqrt{\frac{\sigma_{AB} \cdot \sigma_{AC}}{\sigma_{BC}}}$$

Systematic uncertainty from EP definition difference.

### 2 subevents

$$Res_A = Res_B = \sqrt{\langle cos(\Psi_A - \Psi_B) \rangle}$$

### Systematic uncertainty

$$Err_{sys}^2 = \frac{\sum_{n=1}^{N} (x_i - x_{ave})^2}{N}$$

# **STAR** *p*<sub>T</sub> dependence of v<sub>1</sub> (π, K)



# **STAR** *p***T** *dependence of v*<sup>1</sup> (*p*, *d*)



 $v_1$  increases with increasing  $p_T$  at 1.0< $y_{cm}$ <1.5

### Detector acceptance effects have been taken into account





## **STAR Rapidity dependence of V<sub>1</sub> (π, K)**



## **STAR Rapidity dependence of v1 (p, d)**



**V**cm

reflected marker shown stat. err only



### п, К

✓  $v_1$  sign is negative at low  $p_T$ , and is positive at high  $p_T$ . ✓  $p_T$  dependence is consistent for  $\pi$  + , $\pi$  - and K.

### p, d

 $\checkmark v_1$  increases with increasing  $p_T$ .

 $\checkmark v_1$  of p and d become closer after performing the m<sub>T</sub>-scaling.

## **STAR <b>PT** dependence of **V**1 slope

Centrality : 5-30%



## **STAR <b>PT** dependence of V1 slope







- ✓ We presented the results of the identified particle directed flow at  $\sqrt{s_{NN}} = 4.5$  GeV in Au+Au fixed-target collisions as a function of p<sub>T</sub> and rapidity.
- ✓ Baryon and meson dv₁/dy slopes indicate opposite sign due to the shadowing effect.
- ✓ v<sub>1</sub> of baryons and mesons have the same sign at high-p<sub>T</sub> region (> 0.8 GeV/c), which is averaged out in the previous measurement.
- ✓ The STAR FXT physics program is now ongoing (>300 million event at 3.0 and 7.2 GeV in 2018, will acquire > 100 million events at series of energies from 3.2 to 7.7 GeV from 2019-2021). Higher statistics will allow a more definitive physics message.







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## **STAR EP resolution for each region**

![](_page_26_Figure_1.jpeg)

![](_page_27_Picture_0.jpeg)

- No correction
- Re-centering
- Re-centering+Flattening

![](_page_27_Figure_4.jpeg)