(In)dependence of various LFV observables in the non-minimal SUSY

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- additional symmetry of the SUSY algebra allowed by the Haag Łopuszański Sohnius theorem
- for N=1 it is a global $U_R(1)$ symmetry under which the SUSY generators are charged
- implies that the spinorial coordinates are also charged $Q_R(\theta) = 1, \ \theta \to e^{i\alpha}\theta$
- superpotential example

$$\mathcal{L} \ni \int d^2\theta \, W$$

Superpotential is polynomial in fields. For W to transform homogeneously superfields must have definite R-charges

$$e^{i\alpha Q_R}$$
 $e^{i\alpha Q_R}$ $e^{i\alpha (Q_R-1)}$
 $\Phi = \phi(y) + \sqrt{2}\theta\psi(y) + \theta\theta F(y)$

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(we want it to be) R-invariant $\longrightarrow \mathcal{L} \quad \ni \quad \int d^2\theta \quad W$

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Low-energy R-symmetry realization

- Different possible models that one can construct
- * "Natural" choice $e^{i\alpha Q_R}$ $e^{i\alpha Q_R}$ $e^{i\alpha (Q_R-1)}$ $\Phi = \phi(y) + \sqrt{2}\theta\psi(y) + \theta\theta F(y)$ leptons and quarks $Q_R = 1$ $Q_R = 1$ $Q_R = 0$ Higgs $Q_R = 0$ $Q_R = 0$ $Q_R = -1$
 - **G**ood: no barion and lepton number violating terms
 - Bad: No Majorana masses for higgsinos and gauginos

: <u>Dirac mas</u>	<u>ses</u>				
netric Super	symn	netric Star	ndardmod	el (MRS	SM)
		$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	U(1)
Singlet	Ŝ	1	1	0	
Triplet	Ť	1	3	0	
Octet	Ô	8	1	0	
R-Higgses	Â _u	1	2	-1/2	
	Â _d	1	2	1/2	
	Singlet Triplet Octet R-Higgses	Singlet Ŝ Triplet Ĵ Octet Ô R-Higgses \hat{R}_u \hat{R}_d	$\begin{array}{c c} \hline \text{Dirac masses}\\ \hline \text{etric Supersymmetric Star}\\ \hline \\ \hline \\ Singlet & \hat{S} & 1\\ \hline \\ \text{Triplet} & \hat{T} & 1\\ \hline \\ \text{Octet} & \hat{O} & 8\\ \hline \\ \text{R-Higgses} & \hat{R}_u & 1\\ \hline \\ & \hat{R}_d & 1 \end{array}$	$\begin{array}{c c} \hline \text{Dirac masses}\\ \hline \text{etric Supersymmetric Standardmod}\\ \hline SU(3)_C & SU(2)_L\\ \hline \\ \hline \\ \text{Singlet} & \hat{S} & 1 & 1\\ \hline \\ \text{Triplet} & \hat{T} & 1 & 3\\ \hline \\ \\ \text{Octet} & \hat{O} & 8 & 1\\ \hline \\ \\ \text{R-Higgses} & \hat{R}_u & 1 & 2\\ \hline \\ \\ \hat{R}_d & 1 & 2 \end{array}$	$\begin{array}{c c} \hline \text{Dirac masses}\\ \hline \text{etric Supersymmetric Standardmodel (MRS)}\\ \hline & SU(3)_C & SU(2)_L & U(1)_Y\\ \hline & \text{Singlet} & \hat{S} & 1 & 1 & 0\\ \hline & \text{Triplet} & \hat{T} & 1 & 3 & 0\\ \hline & \text{Octet} & \hat{O} & 8 & 1 & 0\\ \hline & \text{R-Higgses} & \hat{R}_u & 1 & 2 & -1/2\\ & & \hat{R}_d & 1 & 2 & 1/2 \end{array}$

$$W = \mu_d \hat{R}_d \hat{H}_d + \mu_u \hat{R}_u \hat{H}_u$$

$$+ \Lambda_d \hat{R}_d \hat{T} \hat{H}_d + \Lambda_u \hat{R}_u \hat{T} \hat{H}_u + \lambda_d \hat{S} \hat{R}_d \hat{H}_d + \lambda_u \hat{S} \hat{R}_u \hat{H}_u$$

$$- Y_d \hat{d} \hat{q} \hat{H}_d - Y_e \hat{e} \hat{l} \hat{H}_d + Y_u \hat{u} \hat{q} \hat{H}_u$$

MSSM vs. MRSSM

- - soft-SUSY breaking terms
 - $\Box = B_{\mu}$ term
 - □ soft scalar masses
 - Majorana gaugino masses
 - □ A terms

superpotencial

$$\begin{split} \mu_d \, \hat{R}_d \, \hat{H}_d \, + \mu_u \, \hat{R}_u \, \hat{H}_u \\ - Y_d \, \hat{d} \, \hat{q} \, \hat{H}_d \, - Y_e \, \hat{e} \, \hat{l} \, \hat{H}_d \, + Y_u \, \hat{u} \, \hat{q} \, \hat{H}_u \\ \Lambda_d \, \hat{R}_d \, \hat{T} \, \hat{H}_d \, + \Lambda_u \, \hat{R}_u \, \hat{T} \, \hat{H}_u \, + \lambda_d \, \hat{S} \, \hat{R}_d \, \hat{H}_d \, + \lambda_u \, \hat{S} \, \hat{R}_u \, \hat{H}_u \end{split}$$

- soft-SUSY breaking terms
 - \Box B_{μ} -term

- □ soft scalar masses
- Dirac gaugino masses
- no A-terms

One way to fix it: <u>Dirac masses</u> Minimal R-Symmetric Supersymmet

Minimal R-Symmetric Supersymmetric Standardmodel (MRSSM) Kribs et.al. arXiv:0712.2039

			<i>SU</i> (3) _C	$SU(2)_L$	$U(1)_Y$	$U(1)_{R}$
	Singlet	Ŝ	1	1	0	0
Additional fields:	Triplet	Ť	1	3	0	0
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	R-Higgses	Â _u	1	2	-1/2	2
		Â _d	1	2	1/2	2

Kribs, Popitz, Weiter (2008)

MSSM vs. MRSSM

- superpotencial $\mu \hat{H}_{u} \hat{H}_{d} \qquad \bigcirc$ $-Y_{d} \hat{d} \hat{q} \hat{H}_{d} - Y_{e} \hat{e} \hat{l} \hat{H}_{d} + Y_{u} \hat{u} \hat{q} \hat{H}_{u} \bigotimes$
 - soft-SUSY breaking terms
 - $\Box \quad B_{\mu}$ term
 - **D** soft scalar masses
 - Majorana gaugino masses

0

□ A - terms

	superpotencial							
	$ \qquad \qquad$							
	soft-SUSY breaking terms							
	\square B_{μ} -term							
	soft scalar masses							
	 Dirac gaugino masses 							
	no A-terms							
	One way to fix it: <u>Dirac masses</u> <u>Minimal R-Symmetric Supersymmetric Standardmodel (MRSSM)</u> <u>Kribs et.al. arXiv:0712.2039</u>							
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Particle content summary: MSSM vs. MRSSM

different number of physical states completely new states **R-Higgs** Higgs charginos **CP-odd** charged charged neutral sgluon **CP-even MSSM** 2 2 0 0 0 1 1 MRSSM 4 3 3 2 + 22 2 1

	neutralino	gluino
MSSM	4	1
MRSSM	4	1

Majorana fermions

Dirac fermions

Exemplary mass spectrum



5

production

Previous and future low energy experiments

- As the LHC still sees nothing, we look into low energy experiments:
 - prospects for g-2 measurement

x prospect for $\mu \rightarrow e\gamma$

current: 4.2×10-13 (MEG)

 $a_{\mu}^{\exp} - a_{\mu}^{\rm SM} = (28.1 \pm 6.3^{\exp} \pm 3.6^{\rm th}) \times 10^{-10} \qquad a_{\mu}^{\exp} - a_{\mu}^{\rm SM} = (??? \pm 1.6^{\exp} \pm 3.4^{\rm th}) \times 10^{-10}$

future: $\approx 4 \times 10^{-14}$

prospect for $\mu \rightarrow e$ conversion

current: 7×10-13 (SINDRUM-II)

future: ≤10-16

Relation between $(g-2)_{\mu}$ and LFV observables



each observable requires a dedicated experiment

$(g-2)_{\mu}$ in the MSSM



and similarly for $\mu \rightarrow e\gamma$ and $\mu \rightarrow e$ - as long as tan β is not very small all considered observables are dominated by the dipole contributions and therefore strongly correlated $CR(\mu \rightarrow e) \propto \alpha \cdot BR(\mu \rightarrow e\gamma)$

$$\operatorname{CR}(\mu \to e) \le 3 \cdot 10^{-15}$$



$(g-2)_{\mu}$ in the MRSSM



there is one class of enhanced diagram though





 $\propto m_{\mu}^{2} \tan \beta \mu M_{1}$ \tilde{B} \tilde{H}_{2} \tilde{H}_{1}^{0}

 μ_R $ilde{\mu}_R$ μ_L

$(g-2)_{\mu}$ in the MRSSM

It is possible to obtain large contribution to g-2



The price to pay are light EW-inos, in tension with experiment

Photonic penguin dominance

For $|\lambda_d| \ge 1$ the dipoles dominate: g-2 scales linearly with λ_d , while $\mu \rightarrow e\gamma$ and $\mu \rightarrow e$ quadratically



- For $|\lambda_d| \ge 1$ the ratio of $\mu \to e\gamma$ over $\mu \to e$ is of the order 100, as in the MSSM where $CR(\mu \to e) \propto \alpha \cdot BR(\mu \to e\gamma)$
- Near $|\lambda_d| \approx 0$ the ratio is of order 1 or less

 $\frac{br(\mu \to e\gamma)}{CR(\mu \to e, Au)}$ a_{μ} **vs.**



- In the region dominated by the dipoles the $br(\mu \to e\gamma) \sim \sin^2 2\theta \cdot a_{\mu}^2$
- In the MRSSM this is a region of $|\lambda_d| \ge 1$, in the MSSM $\tan \beta \ge 5$

6 distinct regions of parameter space

- At least 2 light masses needed
- 6 distinct parameter regions

BL: líght M_B^D , $m_{\tilde{l}}$ BR: líght M_B^D , $m_{\tilde{e}}$ WL: líght M_W^D , $m_{\tilde{l}}$

BHL: líght M_B^D , μ_d , $m_{\tilde{l}}$ BHR: líght M_B^D , μ_d , $m_{\tilde{e}}$ WHL: líght M_W^D , μ_d , $m_{\tilde{l}}$

Only red ones exhibit λ_d or Λ_d enhancement

Numerical analysis of λ_d and Λ_d enhancement



Summary plot



Conclusions:

- Two distinct cases: $|\lambda_d| \approx 0$, $|\lambda_d| > 0$
- For large $|\lambda_d|$ observables might get dominated by photon "penguins" and strongly correlated
- Generating sufficient contribution to g-2 through large λ_d overshots LFV observables (unless one fine-tunes the mixing angle)
- Similar things happen for $\Lambda_{\rm d}$
- For $|\lambda_d| \approx 0$ the g-2 and $\mu \rightarrow e\gamma$ are still correlated but the $\mu \rightarrow e$ conversion rate can be dominated by so-called charge radius, Z-penguin and box contributions
- It is therefore possible to find a parameter points not excluded by current experimental results, within reach of the next $\mu \rightarrow e$ conversion (but not $\mu \rightarrow e\gamma$) experiment

Backup

EW sector of the MRSSM (status)

- The SM-like Higgs boson mass in the MRSSM has been calculated including full 1-loop and leading 2-loop corrections^{1,2}
- Impact of EWPO was analyzed¹
- MRSSM can predicts correct dark matter relic density while being in agreement with dark matter direct detection bounds³
 - Its EW signatures were checked against available 7 and 8 TeV data³

1. P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, JHEP 1412 (2014) 124

 P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, Adv. High Energy Phys. 2015 (2015) 760729

3. P. Dießner, J. Kalinowski, W. Kotlarski and D. Stöckinger, JHEP **1603** (2016) 007



2 component dark matter

- consider scenarios where the lightest particle with R=1 is neutralino or sneutrino with mass m_{LSP1}
- if $m_{R_1^0} < 2 m_{\text{LSP1}}$, lightest neutral R-Higgs is also stable
- two SUSY dark matter candidates with relic densities Ω_1 and Ω_2
 - requirements
 - $\Box \quad \Omega_{total} h^2 \equiv (\Omega_1 + \Omega_2) h^2 \simeq 0.11$
 - \square substantial fraction Ω_2/Ω_{total}
- (for now) best points are not collinear friendly:

$$m_{\tilde{\chi}^0_1} = 367 \text{ GeV}$$

 $m_{R^0_1} = 571 \text{ GeV}$



Sgluon pair production at 13 TeV LHC

- Analysis of the sgluon pair production with subsequent decay into $t\bar{t}$ pairs. Recasting ATLAS search in the same-sign lepton channel using 3.2/ fb of integrated luminosity
- Signal simulated at NLO using MadGraph5_aMC@NLO + FeynRules + NLOCT and matched to parton shower in the MC@NLO scheme
 - Detector response parametrized using Delphes3
- Analysis validated on background processes $t\bar{t}l^+l^-, t\bar{t}l^\pm\nu$
- Mass of pair produced real spluons decaying with $BR(O \rightarrow t\bar{t}) = 1$ excluded up to 950 GeV



Leading order analysis



LO cross-sections for sparticle production at the LHC at $\sqrt{s} = 13$ TeV

NLO improvements



reduction of theoretical uncertainty

shift of cross-sections

Comparison with the MSSM



Two possible definitions of K-factors:

- unsummed over L- and R-squarks
- ***** summed

Differential distributions



$\mu \rightarrow e\gamma$ in the MRSSM

first analysis performed by Fok and Kribs [Phys. Rev. D 82, 035010 (2010)]

