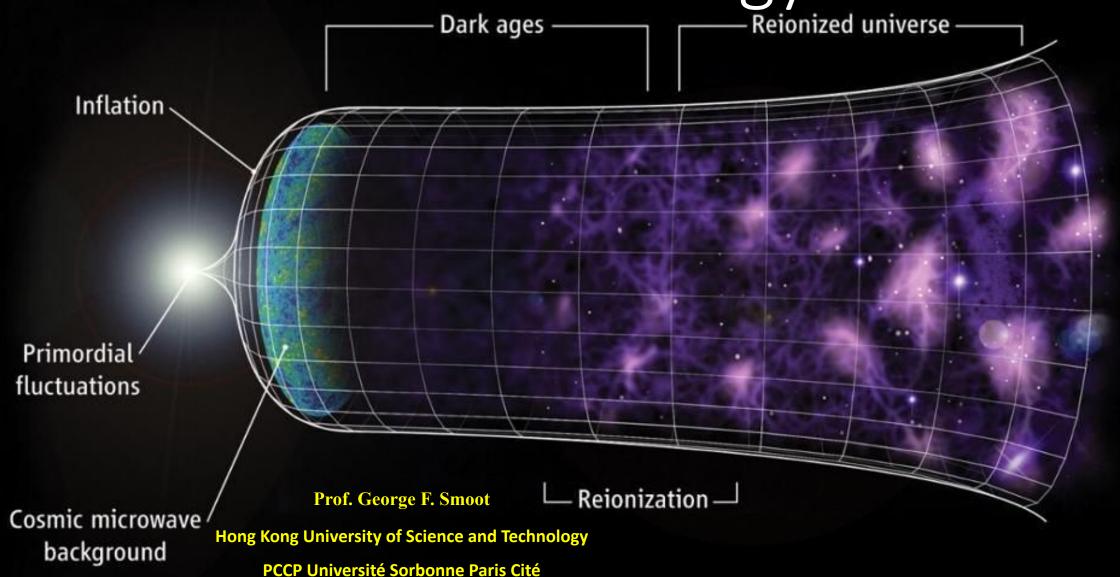
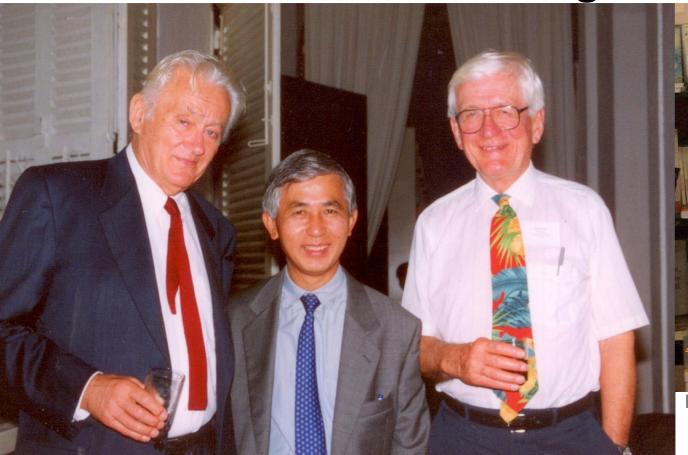
The Future of Cosmology & HEP?



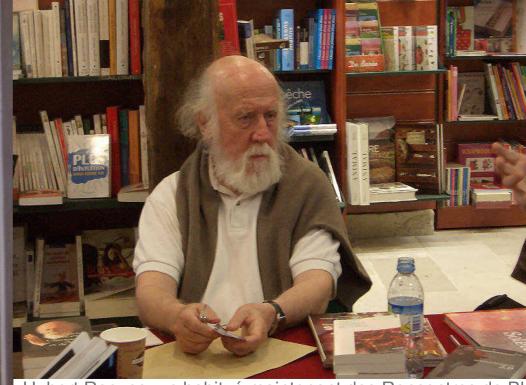
University of California at Berkeley

ECL Nazarbyaev University

Nearly 30 years of Visits to Recontres Blois, e.g. 1992



Jean Trân Thanh Vân with Georges Charpak (left), 1992 physics Nobel prize, and Norman F. Ramsey (right), 1989 physics Nobel prize



Hubert Reeves, un habitué maintenant des Rencontres de Blois





Honor of the Big Bang

In the Loire Valley theoretical physicist Jean Tran Thanh Van stages a conference to ponder the birth of the universe

By J. MADELEINE NASH

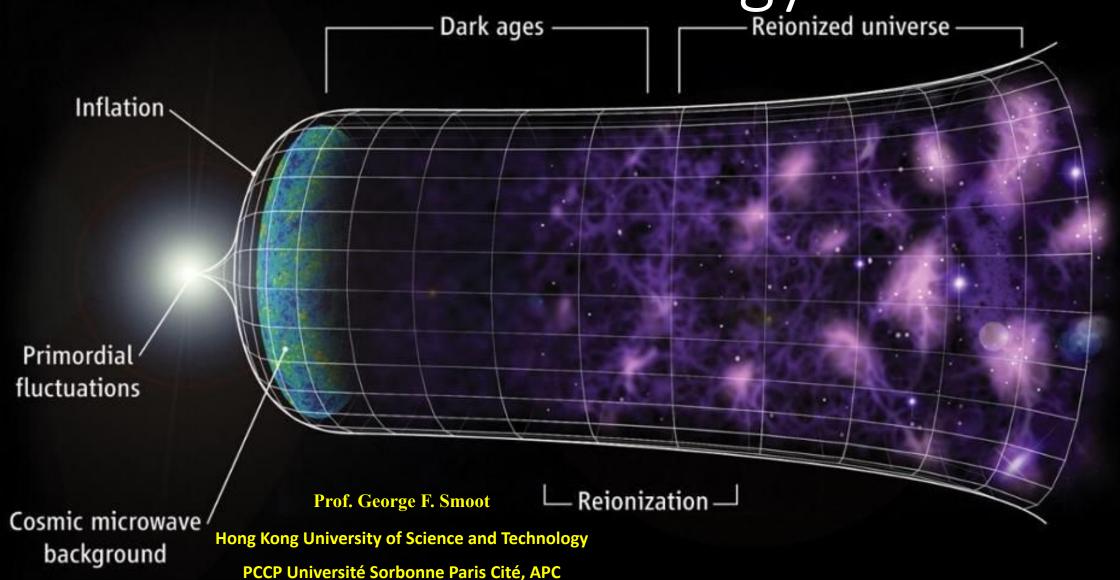
The sharp staccato of clapping hands echoes through the halls of the Châ-

chosen cosmology, a field in which physi-cists, astronomers and astrophysicists are teau de Blois. "Come, come, come, "Jean joined in the heady attempt to fashion a new creation theory. Before 1965, Tran Thanh Van urges a swirling crowd of new creation theory. Before 1965, Tran recientific the predict

particles" collectively known as wiMps. Or, adds Spiro mischievously, "it may be made of MACHOS—massive astrophysical compact halo objects."

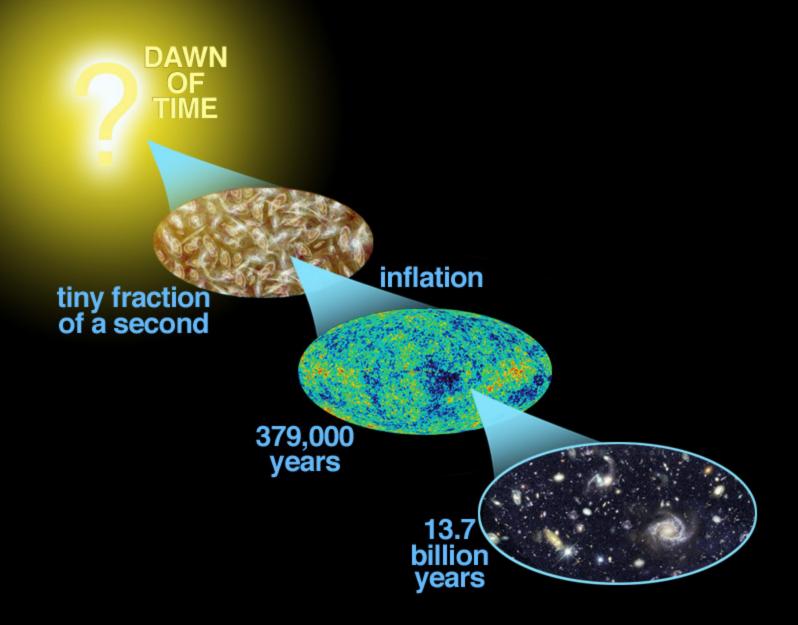
wimps? Machos? "Several centuries ago," jokes Princeton University theorist theorists and experimentalists from many countries and disciplines. This time he has lar sorts of discussions. You know, how many wimpy angels can you fit on the head

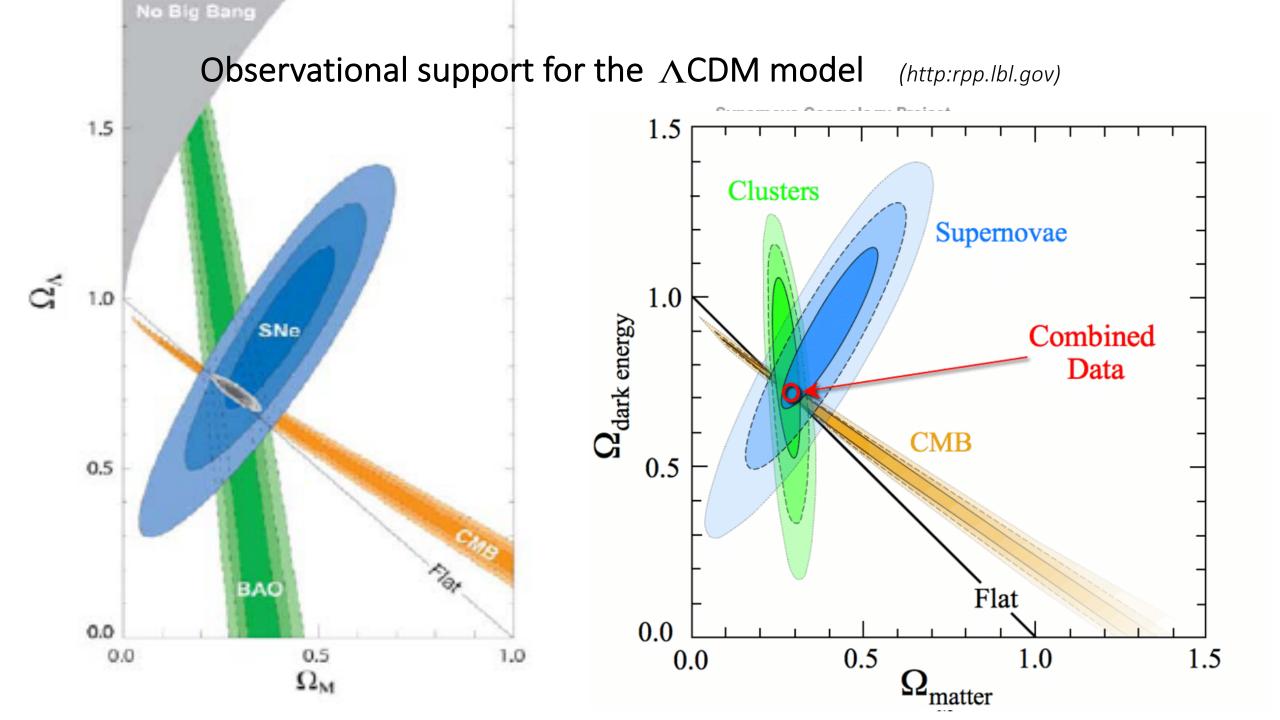
The Future of Cosmology & HEP?

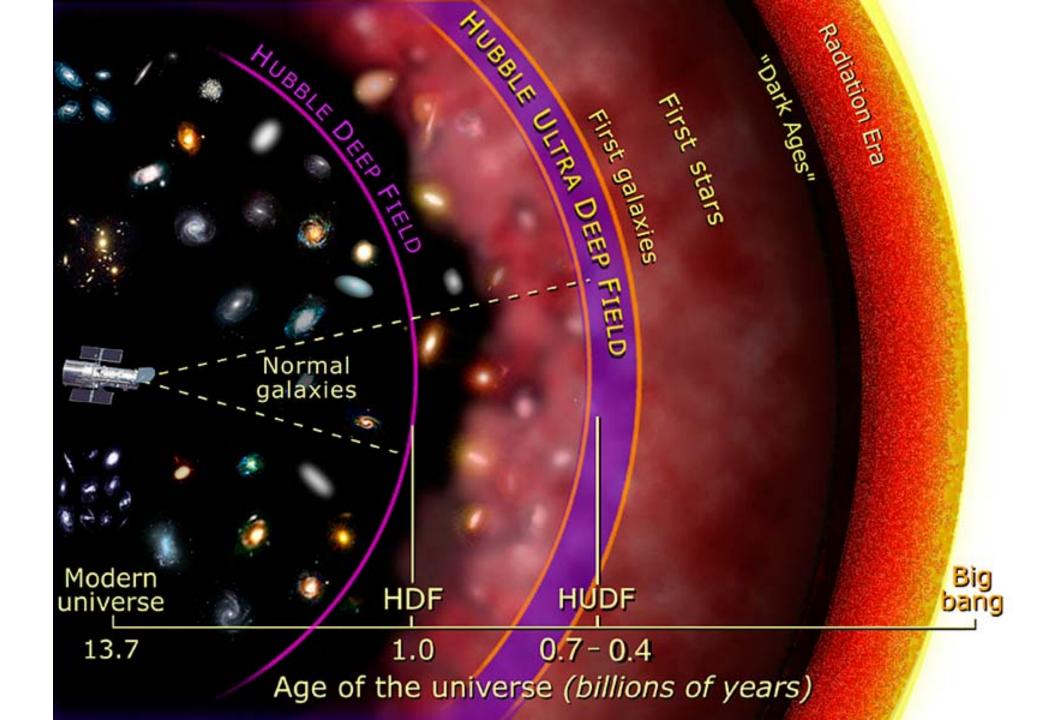


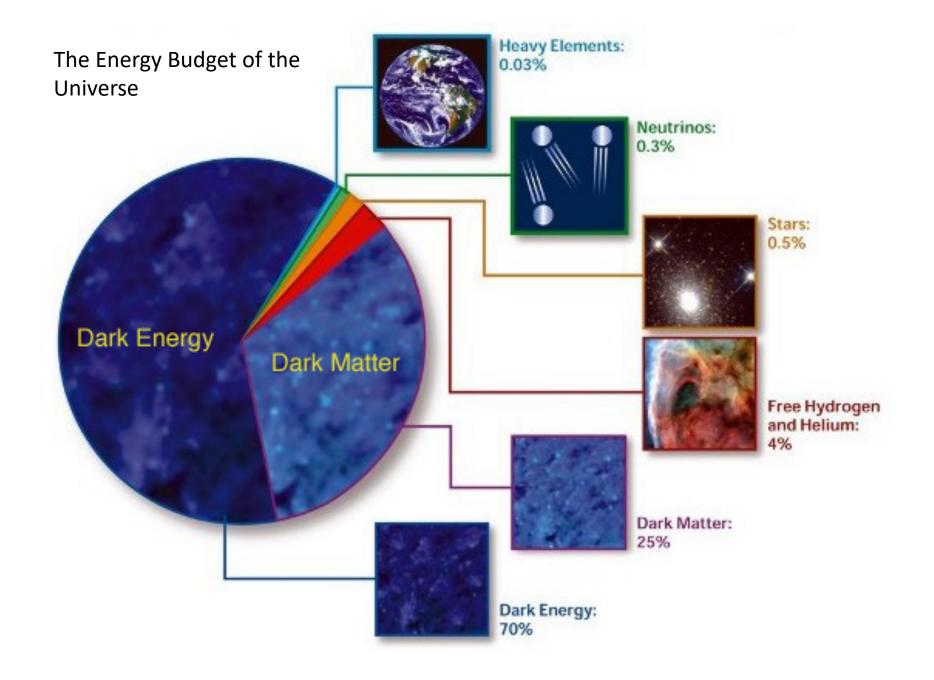
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Standard cosmological model

Cosmological principle (Isotropy and homogeneity at large scales)

Friedmann-Lemaitre-Robertson-Walker metric

$$d\tau^{2} = dt^{2} - a^{2}(t) \left[\frac{dr^{2}}{1 - kr^{2}} + r^{2}d\theta^{2} + r^{2}\sin^{2}\theta d\varphi^{2} \right]$$

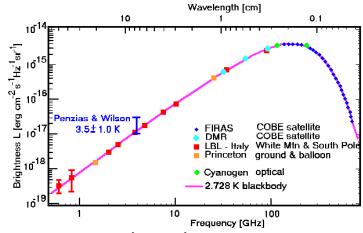
Energy-momentum (perfect fluid)

$$T_{\mu\nu} = diag(\rho, p, p, p)$$

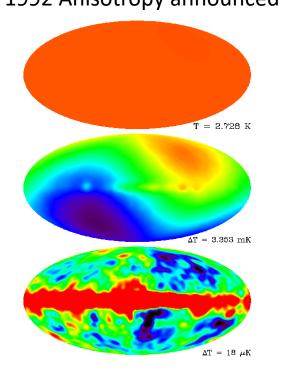
Dynamics (Friedmann equations from GR)

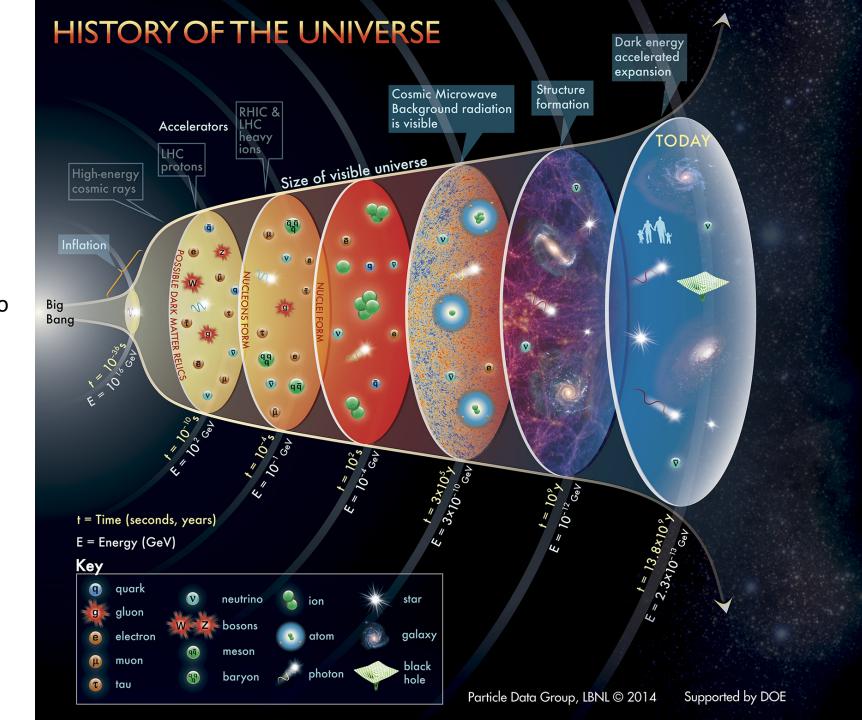
$$H^{2} \equiv \left[\frac{\dot{a}}{a^{2}}\right] = \frac{8\pi G\rho}{3} - \frac{k}{a^{2}} + \frac{\Lambda}{3} \qquad \qquad \frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p) + \frac{\Lambda}{3}$$

CMB discovered 55 years ago Penzias & Wilson 1964

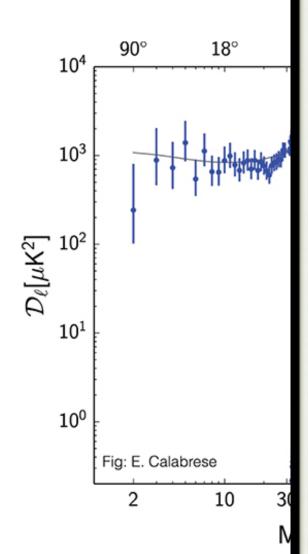


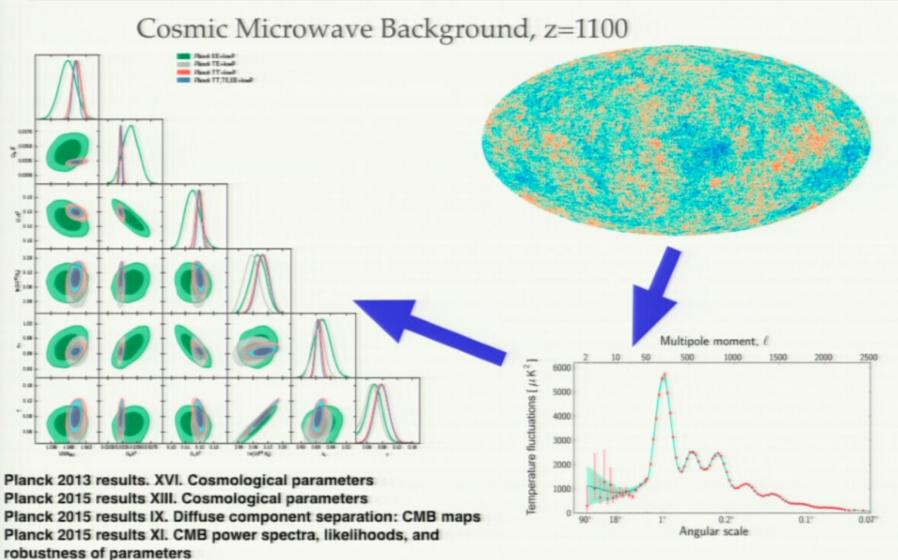
COBE launch 1989 30 years ago 1992 Anisotropy announced



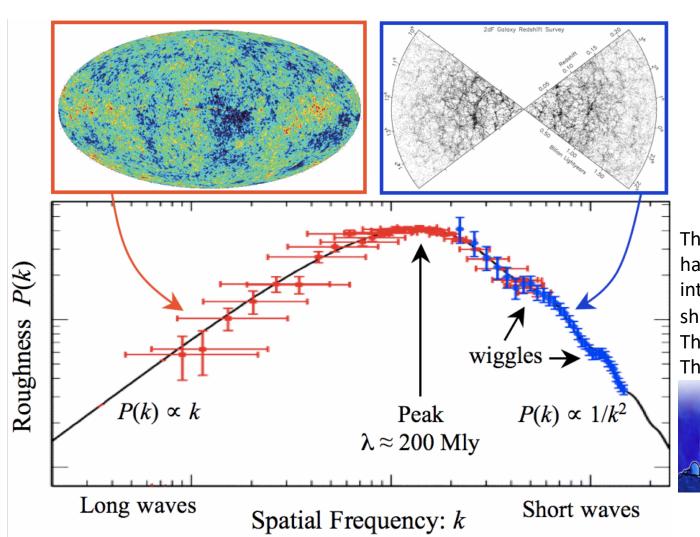


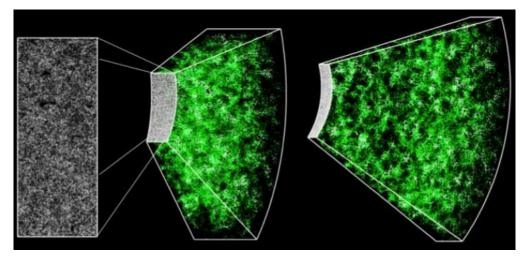
Intensity mapping: CMB





Observations show very good fit to cosmology model



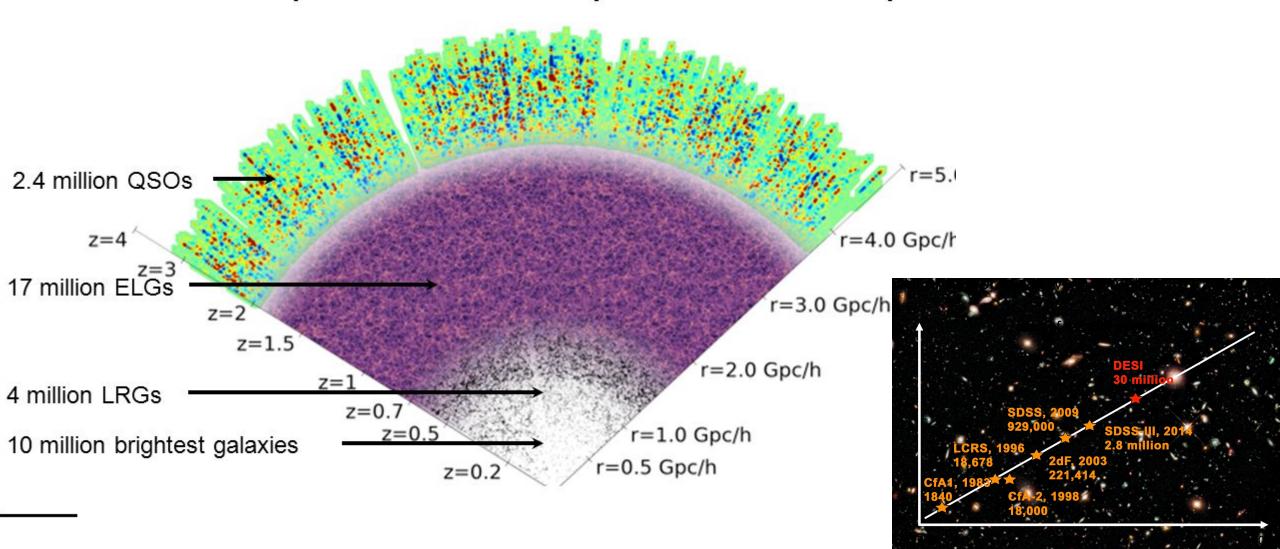


The Sloan Digital Sky Survey - Baryon Oscillation Spectroscopic Survey has transformed a two-dimensional image of the sky (left panel) into a three-dimensional map spanning distances of billions of light ye shown here from two perspectives (middle and right panels). This map includes 120,000 galaxies over 10% of the survey area. The brighter regions correspond to the regions of the Universe



The largest spectroscopic survey for dark energy

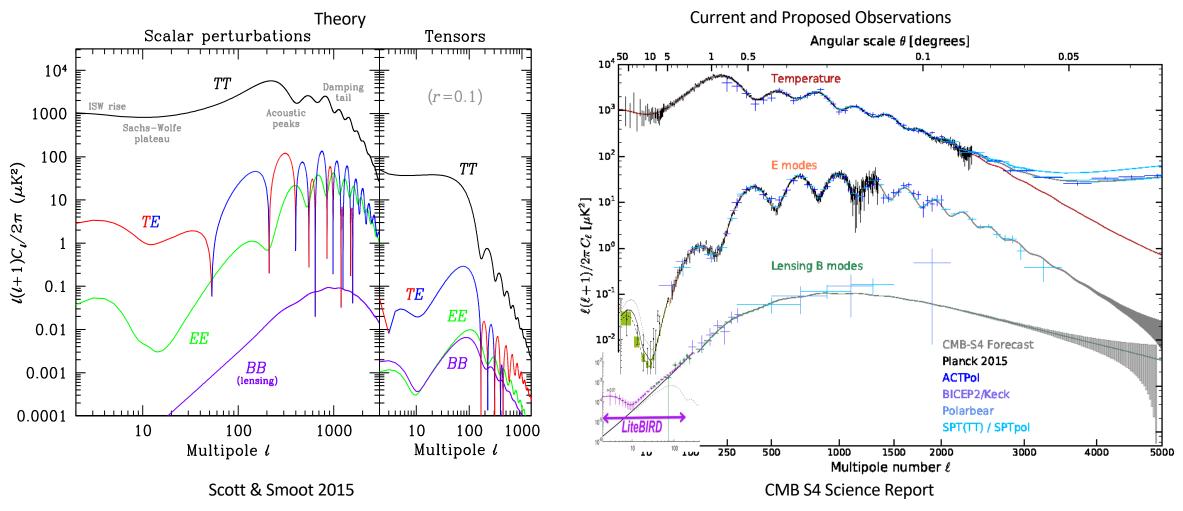
SDSS ~2h-3Gpc³ → BOSS ~6h-3Gpc³ → DESI 50h-3Gpc³



Notre point de vue sans cesse sur Our Ever-sharpening View of the l'univers embryonnaire ... Embryo Universe ... COBE 992 **WMAP** 2003 **PLANCK** 2015 In Progress Simons Array $\sim 0.3^{\circ}$ PolarBear, ACT = Atacama Cosmology Telescope Ali CPT LiteBIRD (JAXA) $\sim 0.02^{\circ}$ Tibet ~2027 **BICEP/KECK** ~0.50 CMB S4 SPT LiteBIRD South Pole Telescope

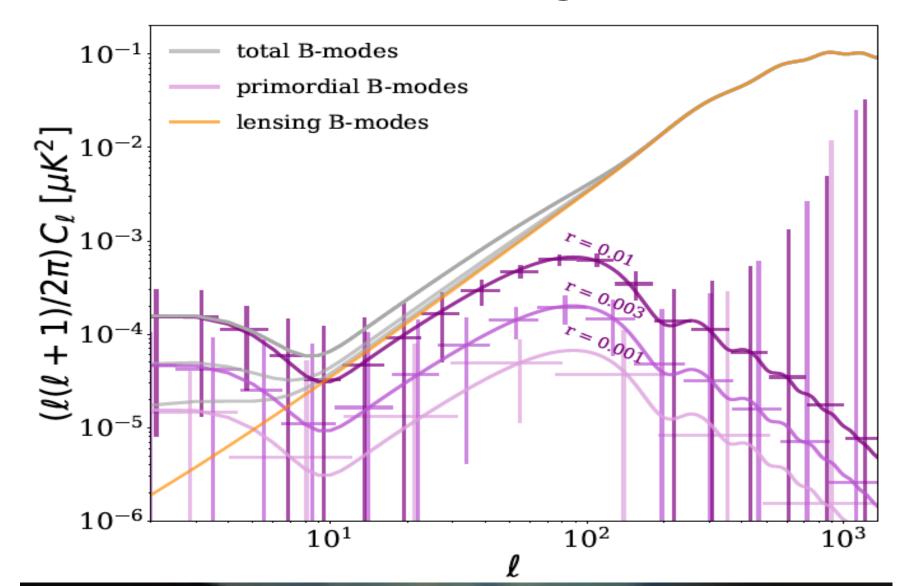
CMB Angular Power Spectra

Scalar Effects Observed Very Well; Tensor Modes not seen

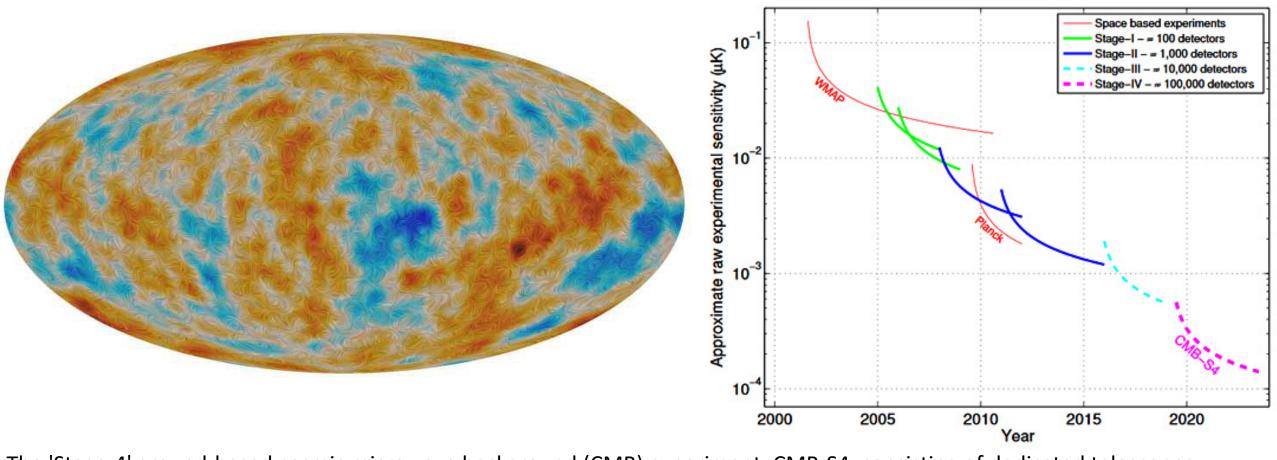


In roughly 5 years CMB observations become clean up except at large angular scales (LiteBIRD assuming r = 0.01 in 2027+) Test of Inflation to be primodial B modes/Tensors

LiteBIRD with best delensing

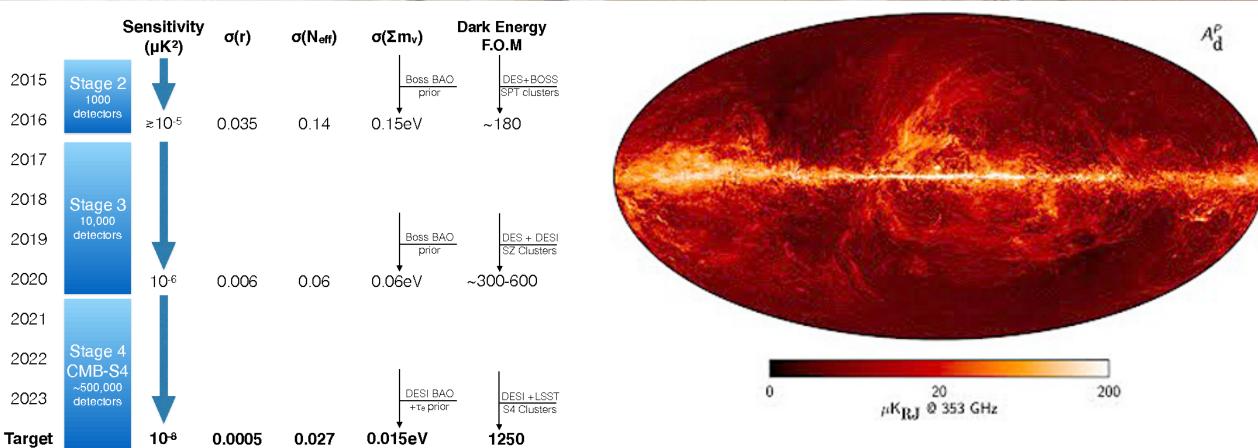


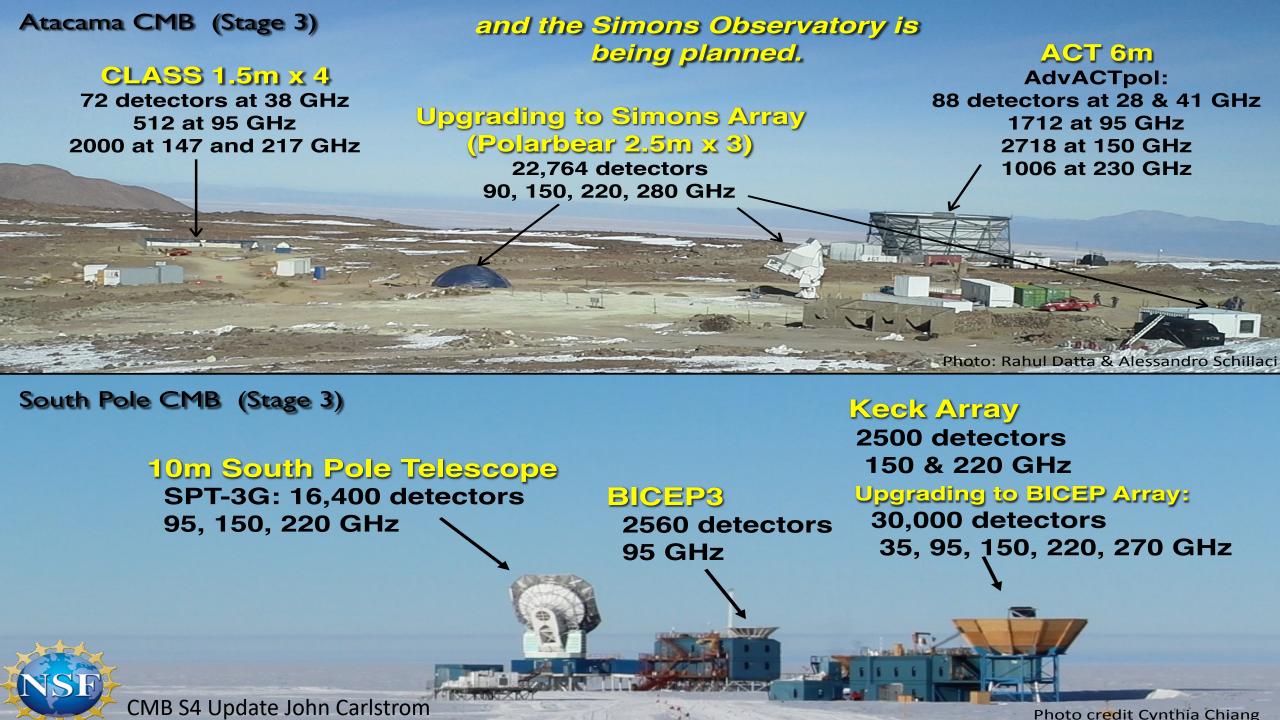
Cosmic Microwave Background (CMB) Radiation

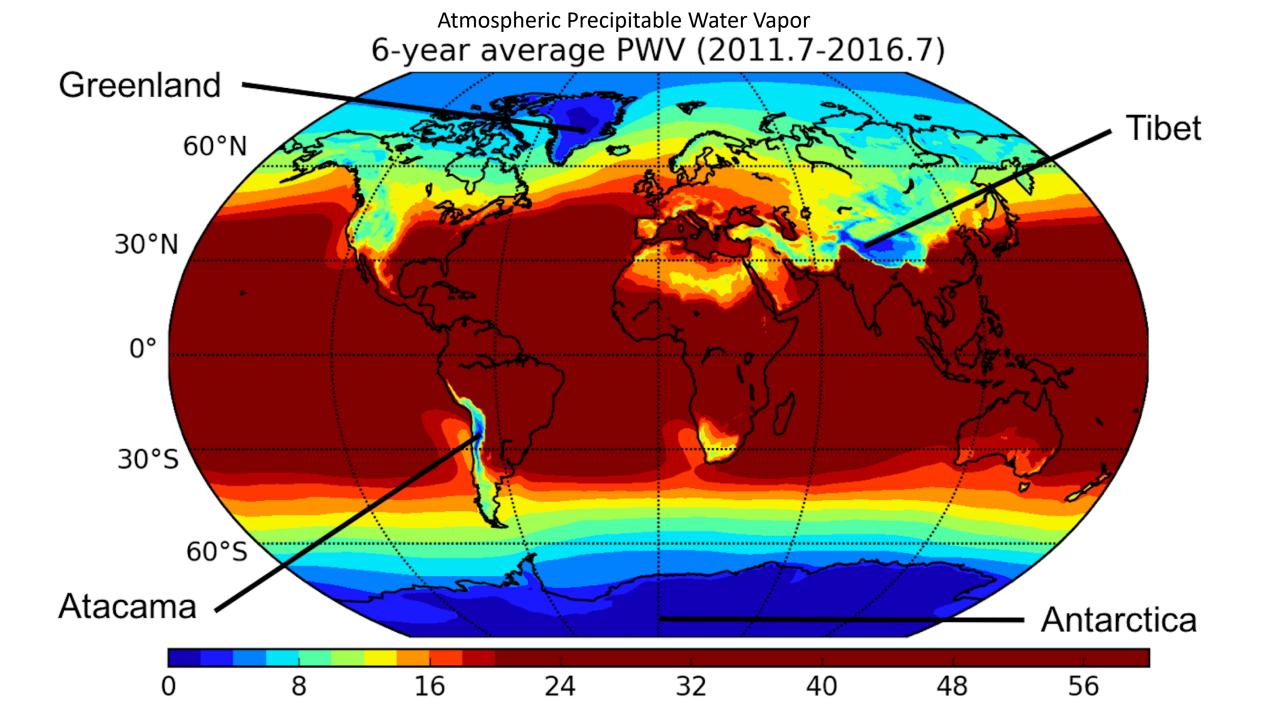


The 'Stage-4' ground-based cosmic microwave background (CMB) experiment, CMB-S4, consisting of dedicated telescopes equipped with highly sensitive superconducting cameras operating at the South Pole, the high Chilean Atacama plateau, and possibly northern hemisphere sites, will provide a dramatic leap forward in our understanding of the fundamental nature of space and time and the evolution of the Universe. CMB-S4 will be designed to cross critical thresholds in testing inflation, determining the number and masses of the neutrinos, constraining possible new light relic particles, providing precise constraints on the nature of dark energy, and testing general relativity on large scales.



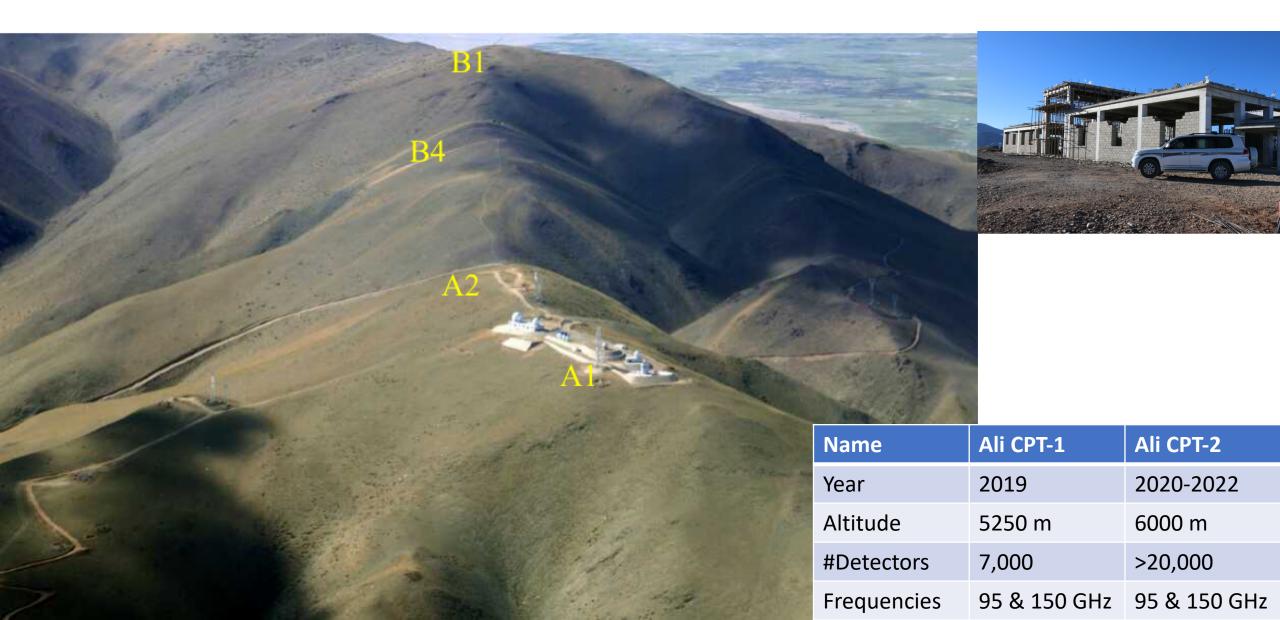




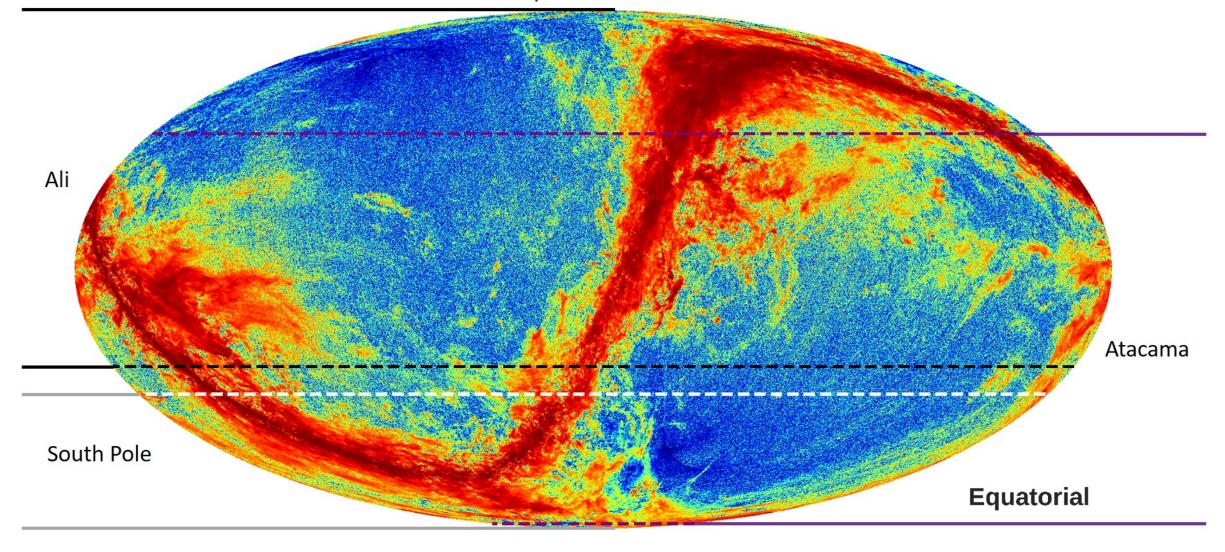




Ali CMB Program Observing Site

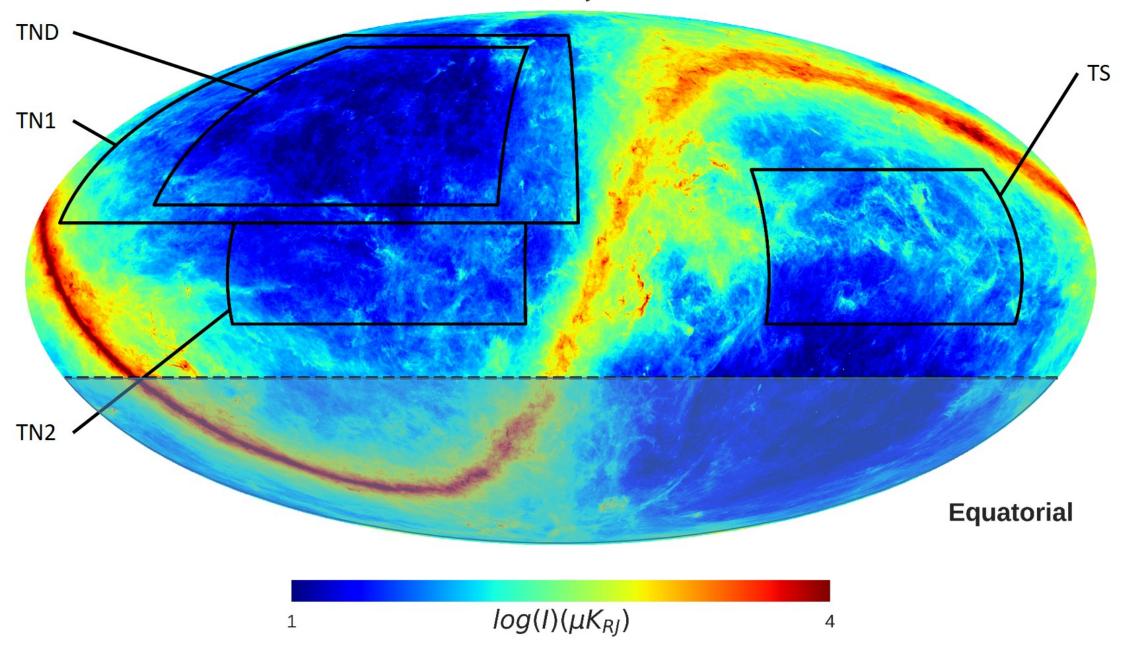




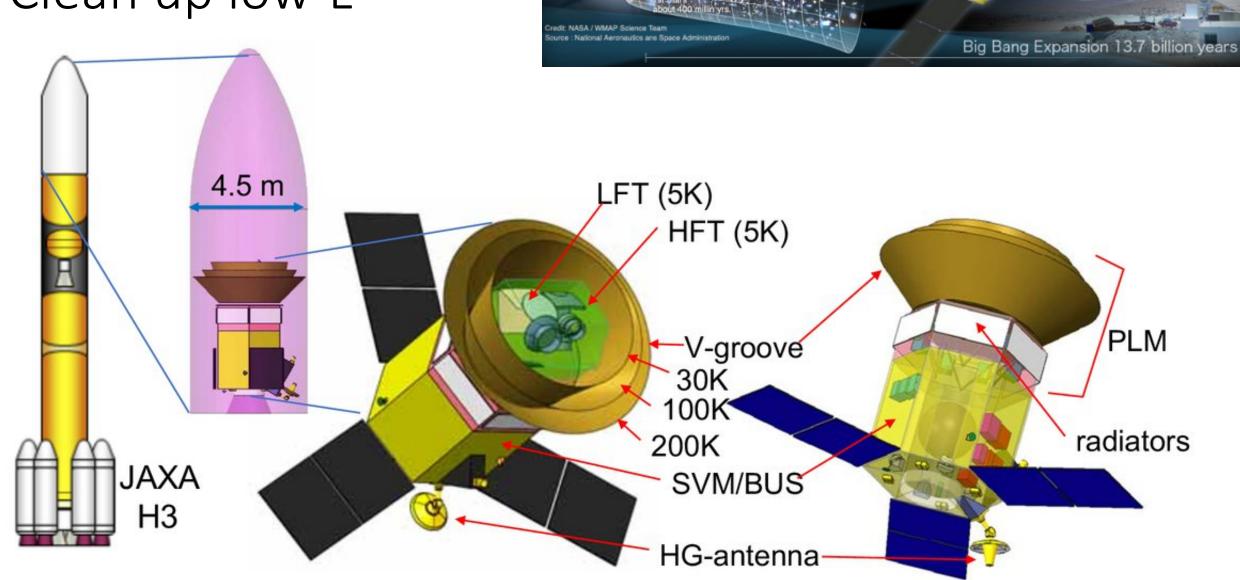




Planck dust intensity at 545.0GHz



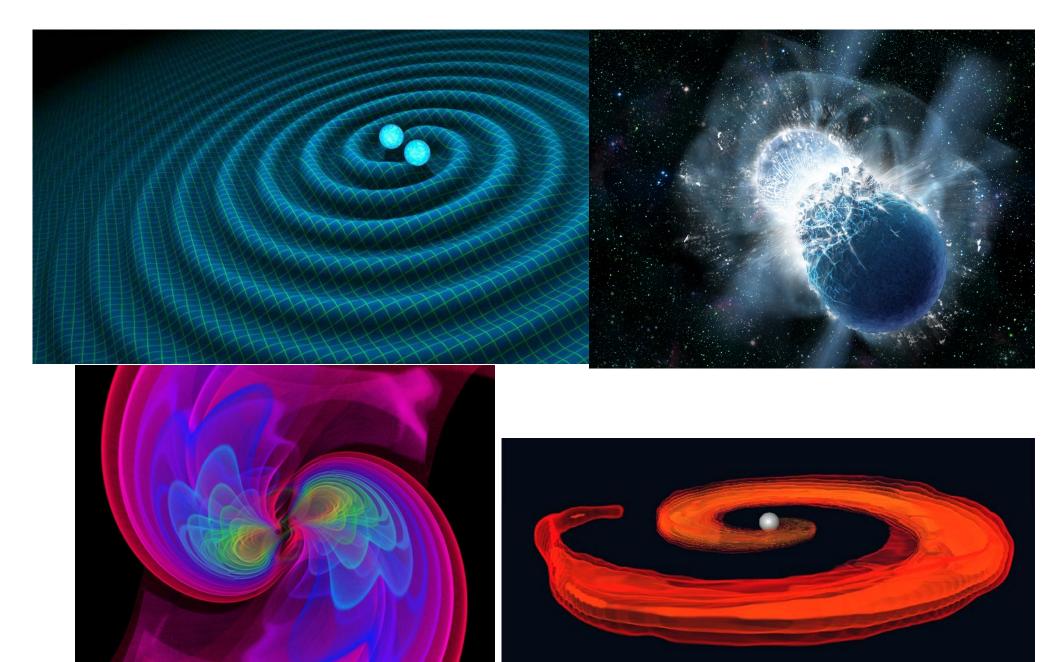
LiteBIRD – JAXA. 2027 Clean up low-L



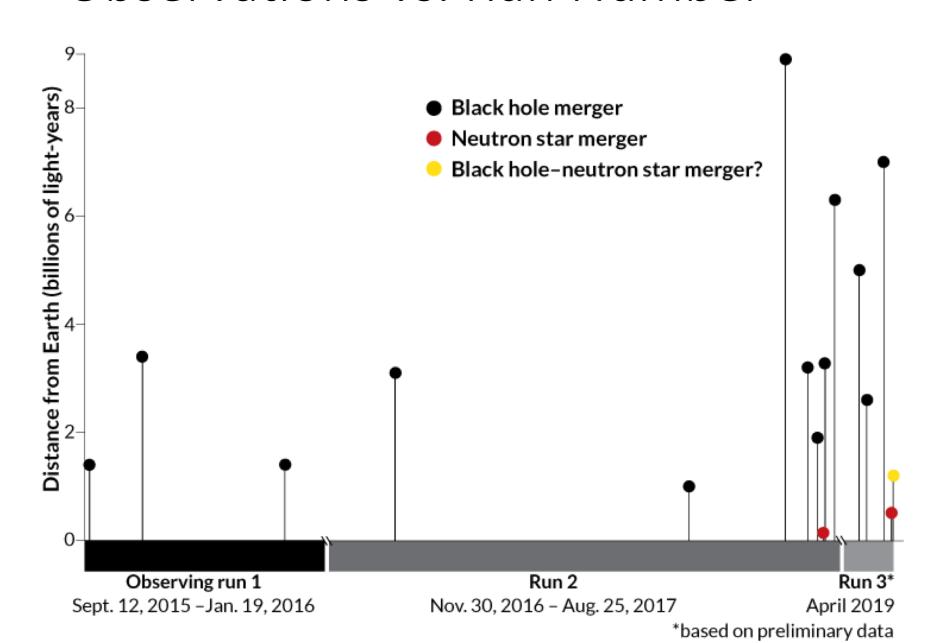
Mode Polarization

LiteBIRD

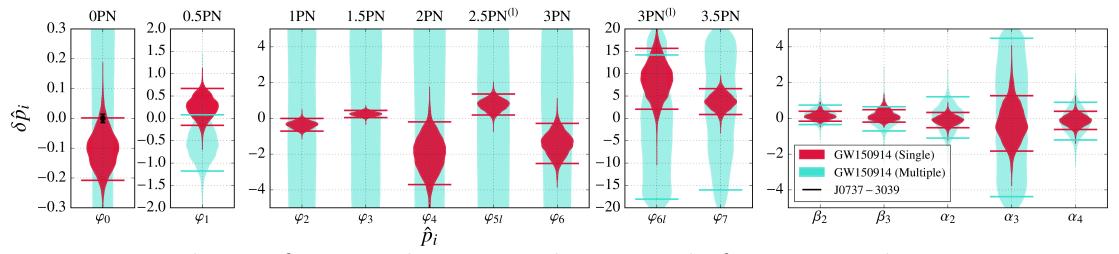
Gravitational Waves: New Kid on the Block



Observations vs. Run Number



Tests of General Relativity with GW



Generic metric theories of gravity predict up to six polarization modes for metric perturbations: two tensor (helicity ±2), two vector (helicity ±1), and two scalar (helicity 0) modes [76, 77]. GWs in GR, however, have only the two tensor modes regardless of the source properties; any detection of a non-tensor mode would be unambiguous indication of physics beyond GR. Simple two tensor polarization best fit so far.

Speed of gravtitational waves is consistent with the speed of light: $\delta v/c < 1.7 \text{ sec} / 130 \text{ mLy} = 4 \times 10^{-10}$



The Cosmological Principle

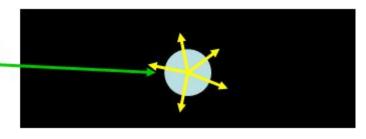
Considering the largest scales in the universe, we make the following fundamental assumptions:

1) Homogeneity: On the largest scales, the local universe has the same physical properties throughout the

Every region has the same physical properties (mass density, expansion rate, visible vs. dark matter, etc.)

2) Isotropy: On the largest scales, the local universe looks the same in any direction that one observes.

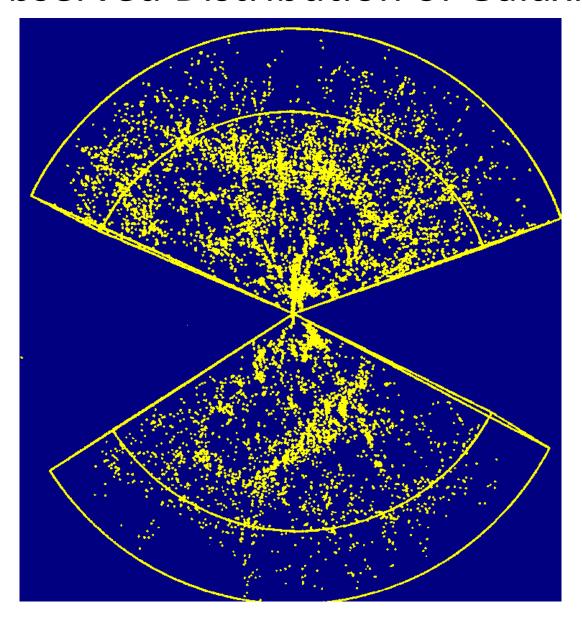
> You should see the same largescale structure in any direction.



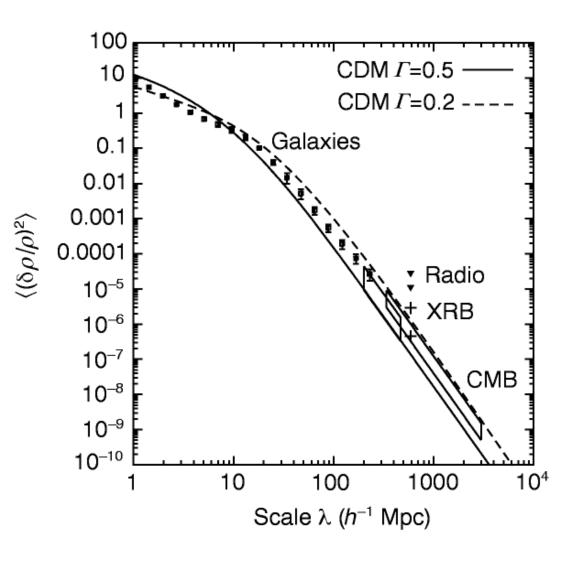
Implies Robertson-Walker Metric

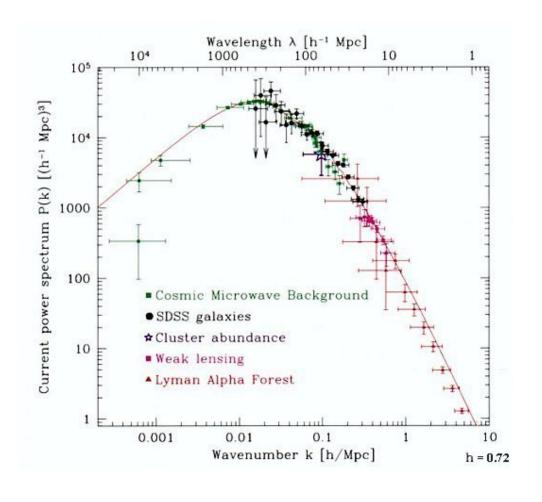
3) Universality: The laws of physics are the same everywhere in the universe.

Observed Distribution of Galaxies

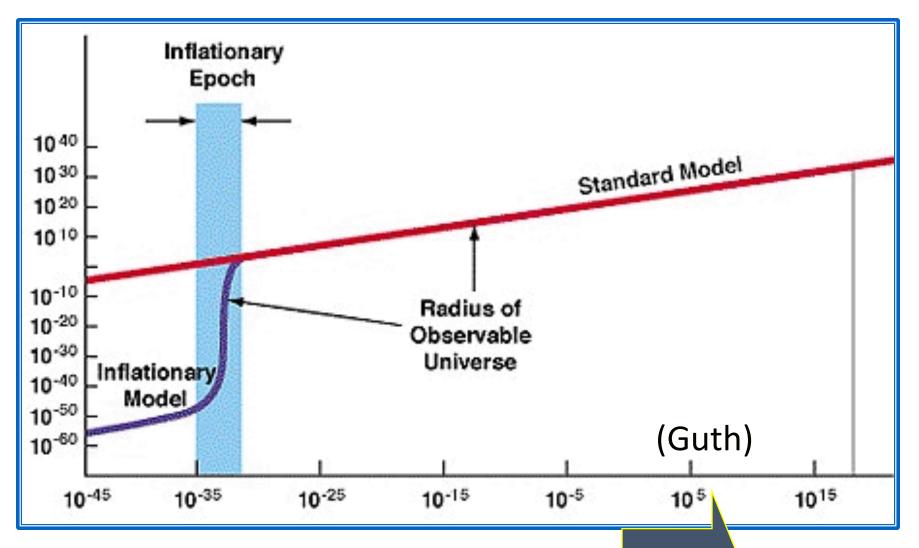


Cosmological Principle Re-Interpreted Currently





The Inflationary Universe



Time in seconds

Broken Symmetries

Cosmological Principles

- The Copernican Principle:
 We do not occupy a special place in the Universe
- The Cosmological Principle:
 The Universe is homogeneous and istotropic (no special place or direction)
- The Perfect Cosmological Principle:
 The Universe is homogeneous in space and time, and is isotropic in space.
- The Weak Anthropic Principle: Life can exist only in the Universe as it is.
- The Strong Anthropic Principle:

 The Universe is such as it is, because its purpose is to create life.

Observations

- Well Earth is pretty special for light years. There is no Planet B.
- The Universe is homogeneous only on very large scales and more in the past than now. Seems isotropic on large scales.
- Clearly wrong on the homogeneity of time on even largest observed scales.
- No clear observations supporting this concept.
- Not necessarily well defined statement.

HISTORY OF THE UNIVERSE Dark energy accelerated expansion Structure Cosmic Microwave formation Background radiation RHIC & is visible Accelerators LHC heavy TODA ions LHC Size of visible universe protons High-energy cosmic rays S *MA (E) Inflation V v 9 SSIBLE DARK MATTER RELICS Big g Bang (V) (e) e V 99 9 100/10/1 F = 10 4 E 3x1010 CeV t = 3×105 V 13.8×10°y t = Time (seconds, years) E = 2.3×10-13 GeV E = Energy (GeV) Key quark neutrino star ion gluon bosons galaxy electron atom meson muon black photon baryon hole tau Particle Data Group, LBNL © 2014 Supported by DOE

What comes after the Standard Model?

- The standard model (SM) of elementary particles involves particle symmetry and the mechanism of its breaking. There areno contradictions with experiments, but calls for extensions to solutions of its internal problems and in view of its evident incompleteness.
- The paradigm of the modern cosmology is based on inflationary models with baryosynthesis and dark matter/energy each involving physics beyond the standard model (BSM) of elementary particles. However, studies of the BSM physical basis of the modern cosmology inevitably reveals additional particle model dependent cosmological consequences that go beyond the modern Standard cosmological model. The mutual relationship of the BSM particle physics basis of the key future directions.

BSM physics and its cosmological reflections

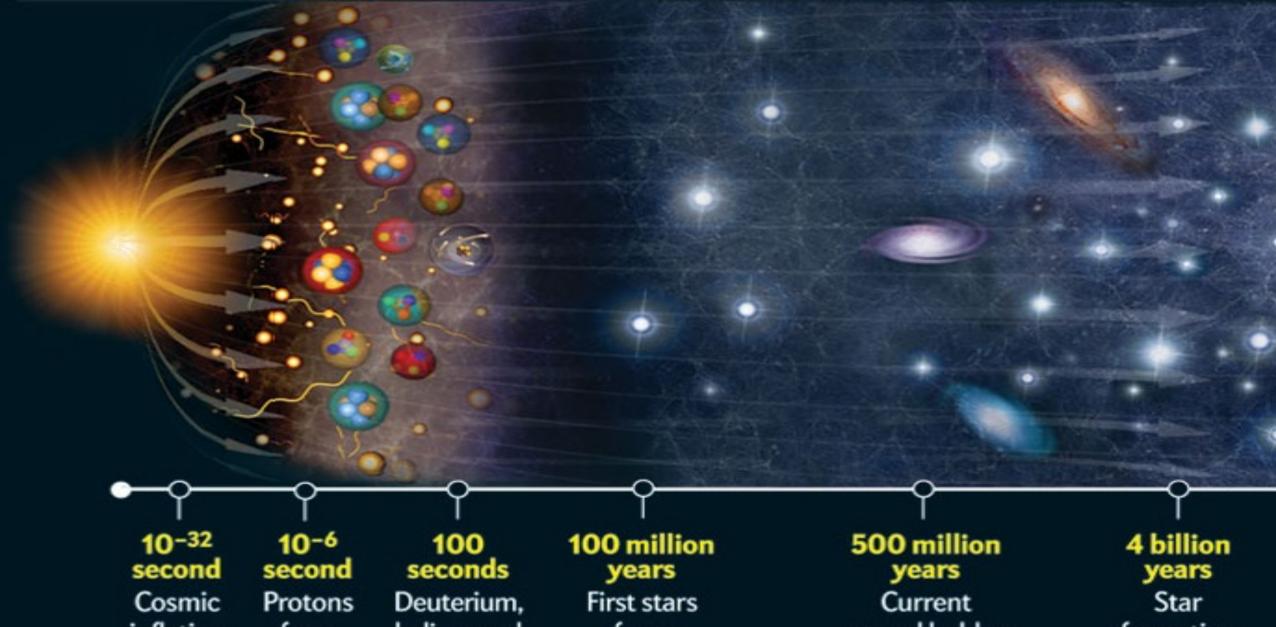
- Arguments for extension of the Standard model of particle physics
- BSM physics of the mass of neutrino
- Supersymmetry
- Composite Higgs boson
- Axion and pseudo Nambu-Goldstone models
- BSM physics of the quark-lepton families
- Mirror and Shadow Worlds
- Unification of fundamental forces
- Unification of all the four fundamental forces, including gravity

Emergent Symmetry

- A key difference between spontaneously broken symmetries and emergent symmetries is that emergent symmetries are never exact (as opposed to broken symmetries prior to breaking).
- For example, could lepton number and baryon number be only approximately conserved especially at lower energies?
- Is it possible that in the low energy ground state one has symmetries arise including Lorentz symmetry? Seems unlikely and thus it would be a fundamental property of the more basic `Lagrangian'. However there are a couple of Condensed matter examples where it does occur.
- Dynamical generation of a gauge symmetry

Theorem – S. Weinberg 1979

- The quantum field theory generated by the most general Lagrangian with some assumed symmetries will produce the most general S matrix, incorporating quantum mechanics, Lorentz invariance, unitarity, cluster decomposition, and those symmetries with no further physical content.
- it isn't any good just to present the formalism and say that it agrees with experiment
- it is very likely that any quantum theory that at sufficiently low energy and large distances looks Lorentz invariant and satisfies the cluster decomposition principle will also at sufficiently low energy look like a quantum field theory



inflation ends

form

helium and lithium are synthesized form

record holder for earliest known galaxy formation peaks

The Einstein-Hilbert action

$$\delta \int \mathcal{L} \mathrm{d}^4 x = 0$$
 Variational Principle

EH gravitational Lagrangian

$$\mathcal{L} = \sqrt{-g}R$$
, therefore $S_{\mathrm{EH}} = \int \sqrt{-g}R\,\mathrm{d}^4x$
Key is variation of the metric

Obtain the full field equations by adding the matter Lagrangian and

vary the action

$$S=rac{1}{16\pi G}S_{
m EH}+S_{
m M}$$
 We now define the energy-momentum tensor as

$$\frac{1}{\sqrt{-g}}\frac{\delta S}{\delta g^{ab}} = \frac{1}{16\pi G} \left(R_{ab} - \frac{1}{2}g_{ab}R \right) + \frac{1}{\sqrt{-g}}\frac{\delta S_{M}}{\delta g^{ab}} = 0$$

$$T_{ab} = -2\frac{1}{\sqrt{-g}} \frac{\delta S_{\rm M}}{\delta g^{ab}}.$$

This allows us to recover the complete Einstein's equation,

$$R_{ab} - \frac{1}{2}Rg_{ab} = 8\pi GT_{ab}.$$

$$\mathcal{L} = -\frac{1}{4}B_{\mu\nu}B^{\mu\nu} - \frac{1}{8}tr(\mathbf{W}_{\mu\nu}\mathbf{W}^{\mu\nu}) - \frac{1}{2}tr(\mathbf{G}_{\mu\nu}\mathbf{G}^{\mu\nu}) \qquad (U(1), SU(2) \text{ and } SU(3) \text{ gauge terms})$$

$$+(\bar{\nu}_L, \bar{\epsilon}_L) \delta^{\mu}iD_{\mu} \begin{pmatrix} \nu_L \\ e_L \end{pmatrix} + \bar{\epsilon}_R \sigma^{\mu}iD_{\mu}e_R + \bar{\nu}_R \sigma^{\mu}iD_{\mu}\nu_R + (\text{h.c.}) \qquad (\text{lepton dynamical term})$$

$$-\frac{\sqrt{2}}{v} \left[(\bar{\nu}_L, \bar{\epsilon}_L) \phi M^c e_R + \bar{\epsilon}_R \bar{M}^c \bar{\phi} \begin{pmatrix} \nu_L \\ e_L \end{pmatrix} \right] \qquad (\text{electron, muon, tauon mass term})$$

$$-\frac{\sqrt{2}}{v} \left[(-\bar{\epsilon}_L, \bar{\nu}_L) \phi^* M^{\nu} \nu_R + \bar{\nu}_R \bar{M}^{\nu} \phi^T \begin{pmatrix} -e_L \\ \nu_L \end{pmatrix} \right] \qquad (\text{neutrino mass term})$$

$$+(\bar{u}_L, \bar{d}_L) \delta^{\mu}iD_{\mu} \begin{pmatrix} u_L \\ d_L \end{pmatrix} + \bar{u}_R \sigma^{\mu}iD_{\mu}u_R + \bar{d}_R \sigma^{\mu}iD_{\mu}d_R + (\text{h.c.}) \qquad (\text{quark dynamical term})$$

$$-\frac{\sqrt{2}}{v} \left[(\bar{u}_L, \bar{d}_L) \phi M^d d_R + \bar{d}_R \bar{M}^d \bar{\phi} \begin{pmatrix} u_L \\ d_L \end{pmatrix} \right] \qquad (\text{down, strange, bottom mass term})$$

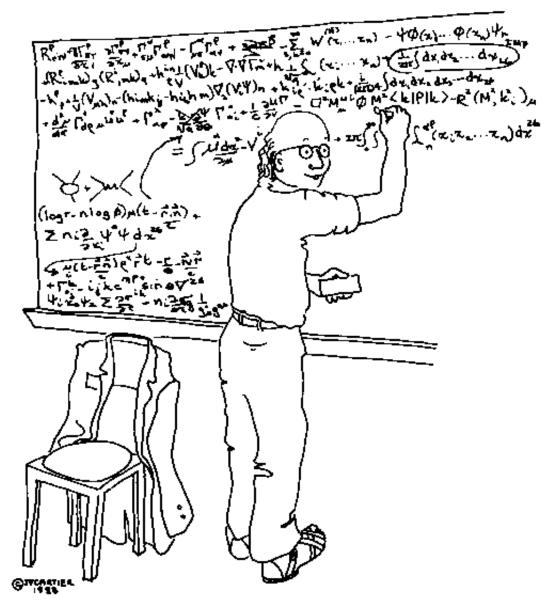
$$-\frac{\sqrt{2}}{v} \left[(-\bar{d}_L, \bar{u}_L) \phi^* M^u u_R + \bar{u}_R \bar{M}^u \phi^T \begin{pmatrix} -d_L \\ u_L \end{pmatrix} \right] \qquad (\text{up, charmed, top mass term})$$

$$+ \overline{(D_\mu \phi)} D^\mu \phi - m_h^2 [\bar{\phi} \phi - v^2/2]^2 / 2v^2. \qquad (\text{Higgs dynamical and mass term}) \qquad (1)$$

$$\frac{1}{8} g^2 \frac{g^2 v_\mu^2}{v_\mu^2} \frac{g^2 v$$

Variation of matter Lagrangian components To provide the extrememum of the action Including space time





"At this point we notice that this equation is beautifully simplified if we assume that space-time has 92 dimensions."

In a modern understanding of particle physics, global symmetries are approximate and gauge symmetries may be emergent. - Ed Witten 2017

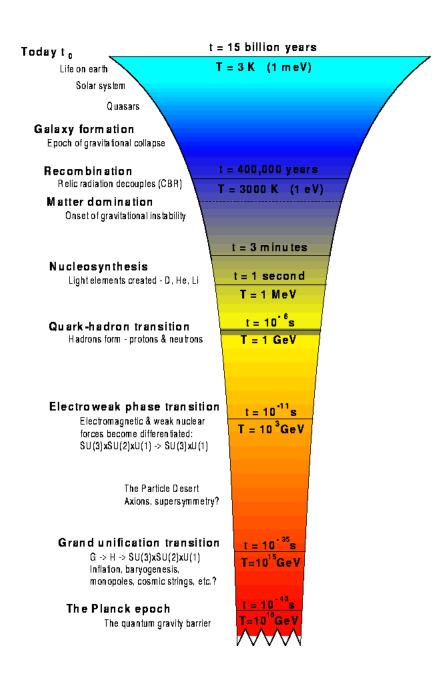
If we want to generate gravity as an emergent phenomenon, we must go farther and generate spacetime itself as an emergent phenomenon.

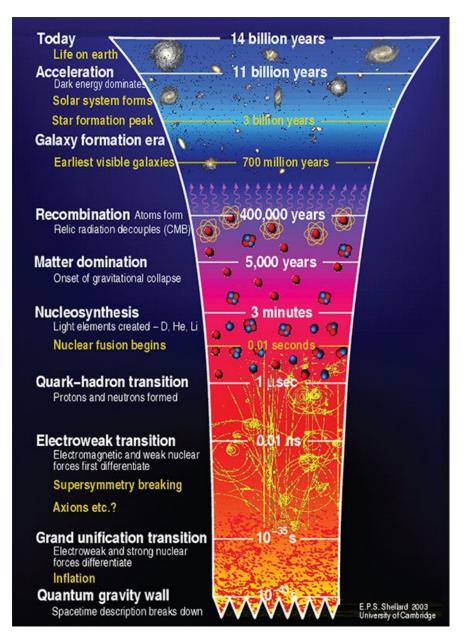
Conclusion

- Interesting observations coming in Cosmology
- Possible there will be no great surprises only improved constraints and confirmations of the Standard Model of Cosmology
- Cosmology motivates a lot of HEP program

- Possible BSM could have big surprises
- Also possible emergent spacetime and symmetries







Examples of more epochs
Or major transitions

One can imagine this Going on into the future To energy scale 10⁻²² eV

However now we see Issues at 10¹⁹ GeV (10²⁸ eV And at 10⁻²² eV but Perhaps lower much later.

We can make it so that
Every time is within
A decade of an important
Universe epoch.
Seems arbitrary but is
Predictive