

Recent highlights from VERITAS

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on behalf of the VERITAS Collaboration
UCSC-SCIPP





31st Rencontres de Blois **June 2019**

VERITAS in few lines

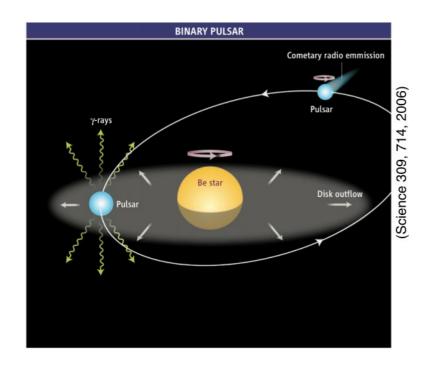


- 4 x 12m Cherenkov telescopes in Arizona (+31° 40′ 30.21″, −110° 57′ 7.77″)
- **First light:** April 2007
- Energy range: from ~ 85 GeV to >30 TeV
- Sensitivity: $\sim 6 \times 10^{-13} {\rm erg~cm^{-2}s^{-1}}$ at 1 TeV in 50 h (~ 10% of the Crab flux in 25 min)
- Ang. Resolution: < 0.1 deg at 1 TeV
- Field of view: 3.5 deg

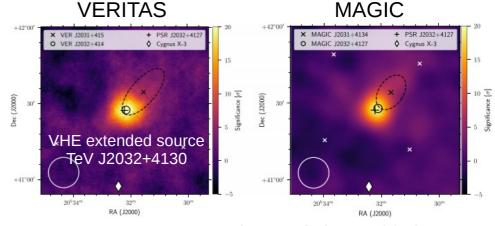


Galactic sky – *TeV emission from binary systems*

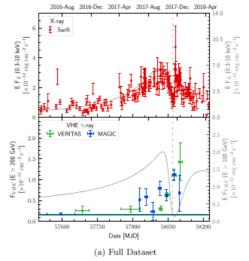
Periastron Observations of TeV Gamma-Ray Emission from a Binary System with a 50-year Period pulsar/Be star binary system PSR J2032+4127 / MT91 213

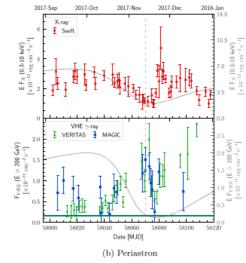


- TeV emission likely coming from the puslar wind nebulae
- Complex process driving the VHE, models need to be revised



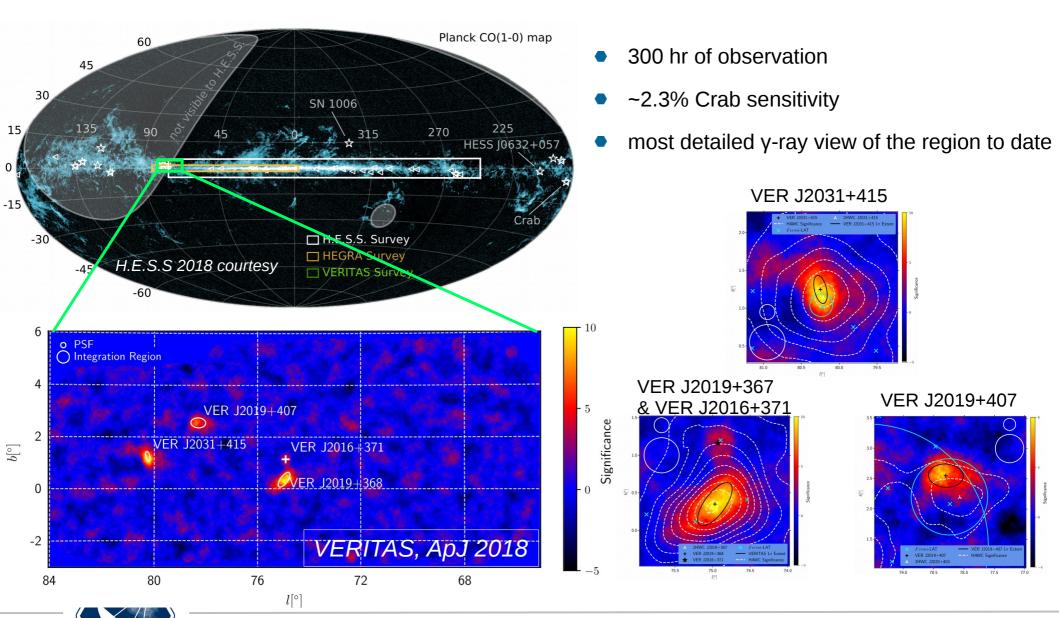
VERITAS & MAGIC, ApJ 2018





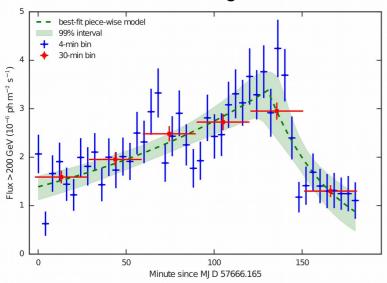


Galactic sky – A Very High Energy γ-Ray Survey towards the Cygnus Region of the Galaxy

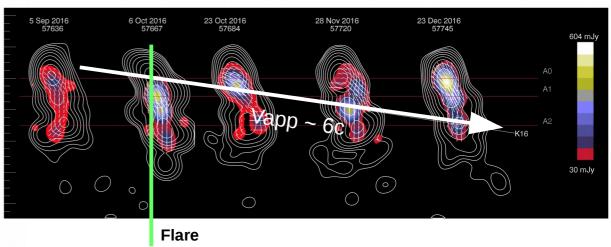


AGN - Fast flare of BL Lacertae in Oct 2016

VERITAS flare lightcurve

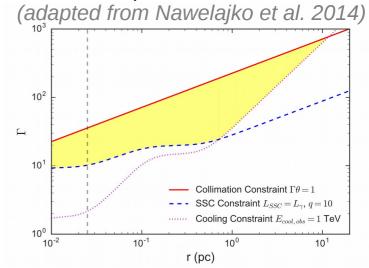


Fast knot ejection potentially linked to the flare



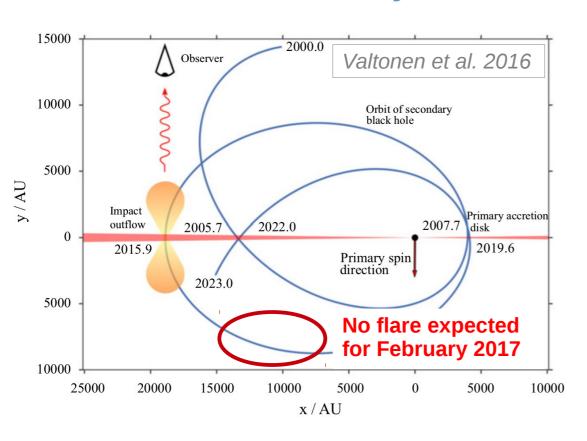
- Peak flux ~ 180% Crab, fast decay
- Rise: 140^{+25}_{-11} min, decay: 36^{+8}_{-7} min
- Intrinsic flow suggested to be faster than the measured speed in radio
- Origin of the flare: Shock and/or magnetic reconnection?

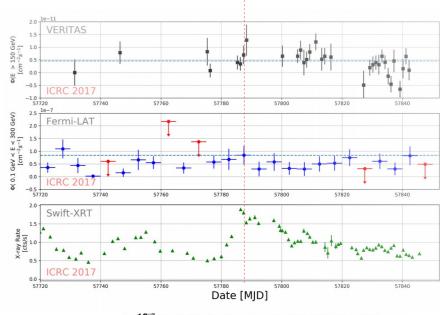
Parameter space of the VHE location





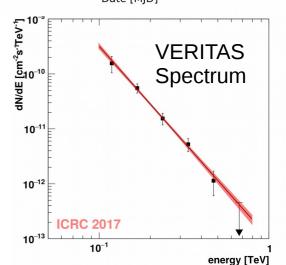
AGN – Intriguing activity of the SMBH binary OJ 287 in February 2017





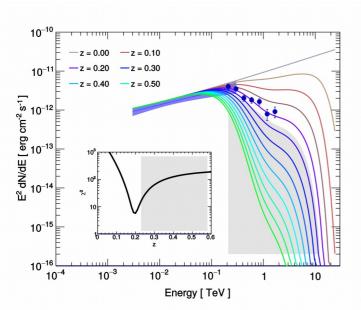
- Likely a flare from the primary black hole
- Multi-wavelength study and interpretation in process...



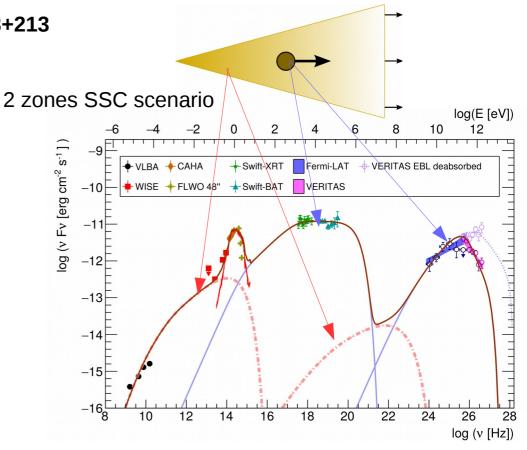


AGN – Others

Ultra high frequency BL Lac (UHBL) HESS J1943+213



Improved the redshift constraint: z < 0.23



But Also:

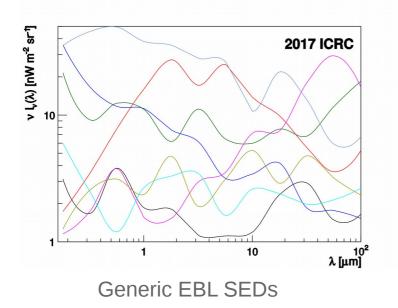
- Mrk 501 showing extreme behaviour (MAGIC & VERITAS, A&A 2018)
- Bright flares of the radiogalaxy NGC 1275 in 2016 & 2017 (See Lucy Fortson talk!)
- Detection of the radiogalaxy 3C 264 (See Lucy Fortson talk!)

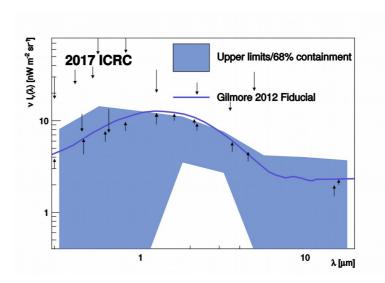


Probing the foreground fields

Extragalactic background light (EBL)

- "model-independent" method, based on M. Lorentz et al. (2015)
- An ensemble of possible EBL SEDs are generated (2nd order splines), based on a grid of points in EBL density versus EBL wavelength space
- 480,000 EBL SEDs were considered

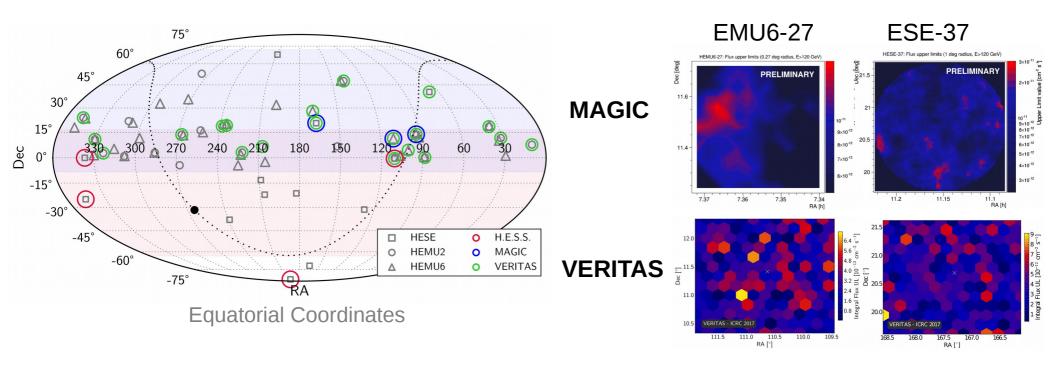




68% confidence band on the EBL shape, based on the observed source spectra of 8 blazars



Neutrino counterpart search



- The apparent isotropy of the astrophysical events seems to favor a dominant extragalactic component.
- Set strong constraints on the gamma-ray flux of potential neutrino sources detected by IceCube
- Example of synergies between multiple IACT collaborations (VERITAS, MAGIC, HESS, FACT)

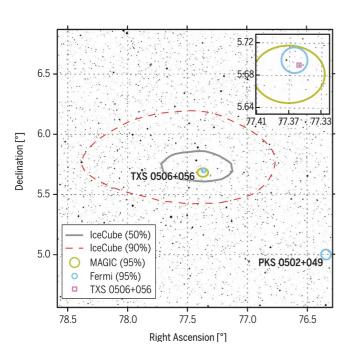


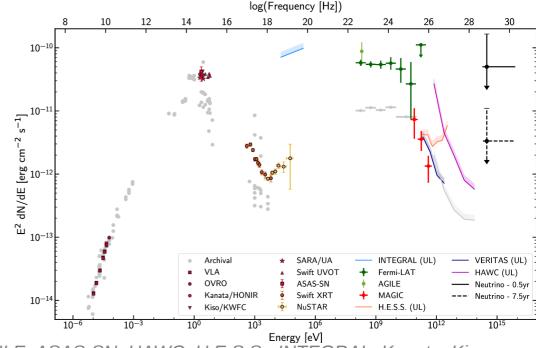
Multi messengers – The first neutrino blazar?

September 22 2017: IceCube alert

September 28 2017: Fermi-LAT report detection of a gamma-ray flare from the blazar TXS 0506+056, consistent with the location of neutrino event IC 170922A (Atel #10791)

October 04 2017: MAGIC report a detection in VHE (Atel #10817)



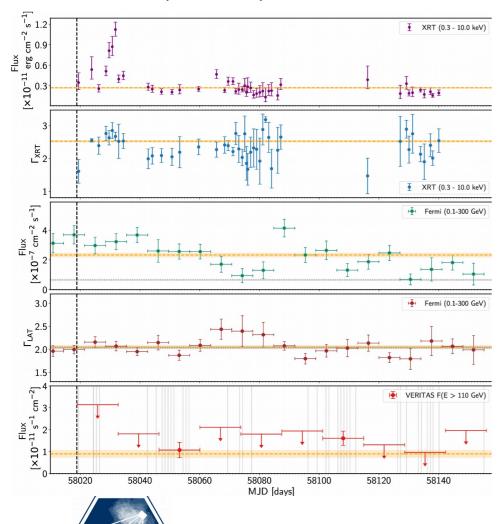


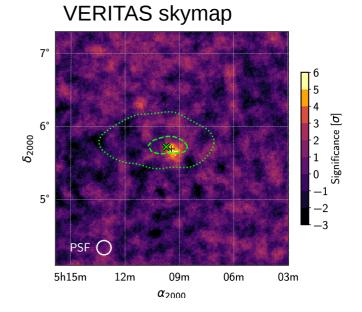
The IceCube Collaboration, Fermi-LAT, MAGIC, AGILE, ASAS-SN, HAWC, H.E.S.S., INTEGRAL, Kanata, Kiso, Kapteyn, Liverpool Telescope, Subaru, Swift/NuSTAR, VERITAS, and VLA/17B-403 teams (Science 2018)

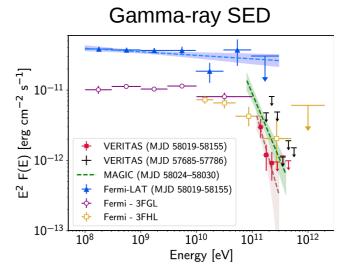


Multi messengers – *VERITAS observations TXS 0506+056 /IceCube-170922A*

5.8 sigma detection by VERITAS from 35h observing time Weak flux: $F(>110 \text{ GeV}) \sim 1.6\% \text{ Crab}$







Multi messengers – GRBs and Gravitational waves follow-up

GRBs entering in the VHE area

Congrats to our VHE colleagues!

GRB 190114C (MAGIC) GRB 180720B (H.E.S.S.)

VERITAS is also intensively monitoring GRBs e.g.

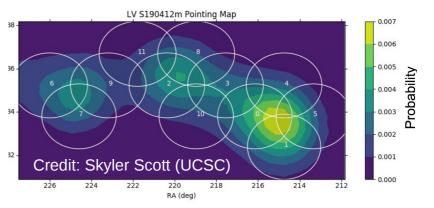
"A Strong Limit on the Very-high-energy Emission from GRB 150323A" (VERITAS, ApJ 2018)

- observation 270 s after the onset of BAT, and only
 135 s after the main BAT emission peak
- VHE non-detection constrains the wind density parameter

Gravitational wave events soon?

- 3 successful follow-ups since the Ligo/Virgo Observing run 3 (O3, from Apr 1rst 2019)
- Roughly ~1-2 GW alerts per week
- No VHE counterpart detected yet

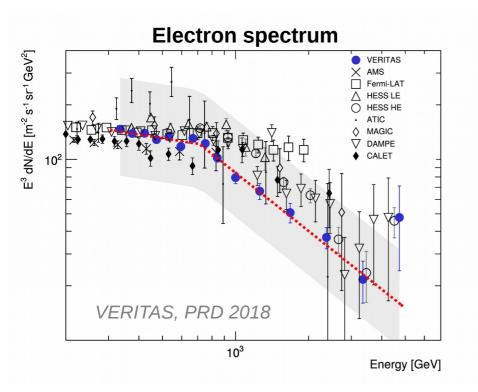
Example of follow-up for the event LV S190412



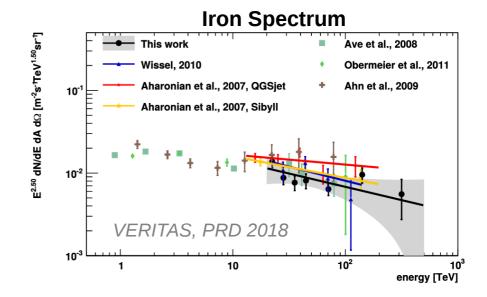
5 min per pointing → sensitivity ~50% Crab



Cosmic rays measurements at TeV energies



- ~ 300h of observation
- Spectrum from 300 GeV to 5 TeV
- Broken Power law with a break ~710 GeV



- Spectrum from 20 TeV to 500 TeV
- Power law shape

$$\frac{\mathrm{d}F}{\mathrm{d}E} = f_0 \cdot \left(\frac{E}{E_0}\right)^{-\gamma}$$

$$\gamma = 2.82 \pm 0.30 (\text{stat.})^{+0.24}_{-0.27} (\text{syst.})$$

$$f_0 = (4.82 \pm 0.98(\text{stat.})^{+2.12}_{-2.70}(\text{syst.})) \cdot 10^{-7} \,\text{m}^{-2} \,\text{s}^{-1} \,\text{TeV}^{-1} \,\text{sr}^{-1}$$

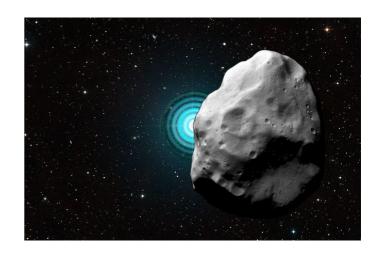
$$E_0 = 50 \,\text{TeV}$$



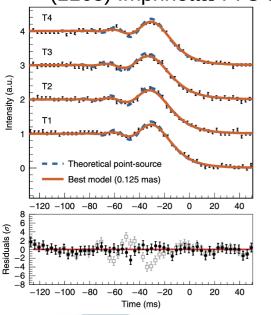
Extreme optical performances of VERITAS

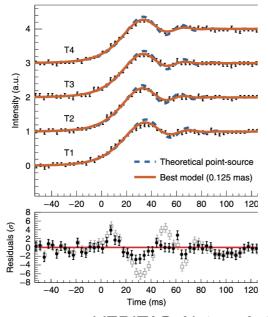
Direct measurement of the occulted stars' angular diameter at the ≤0.1 mas scale

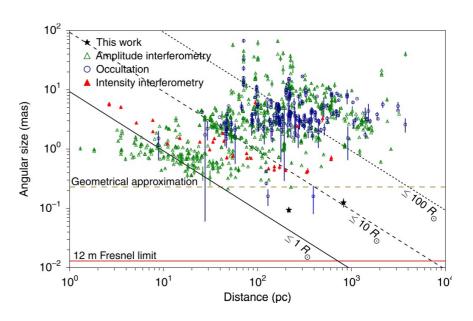
A resolution never achieved before with optical measurement









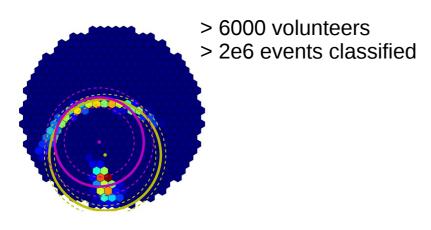


VERITAS, Nature Astronomy 2019

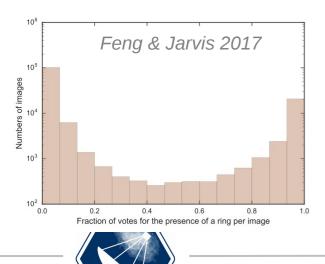


Public outreach – A citizen-science approach to muon events in IACT data

Muon Hunter 1

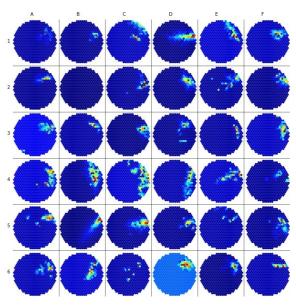


Citizen classification used as an input for CNN training





Muon Hunter 2.0: Return of the Ring



Check the results of unsupervised machine learning classification

Become a muon hunter here:

https://www.zooniverse.org/projects/dwright04/muon-hunters-2-dot-0

VERITAS is moving forward

CTA protoptype pSCT inaugurated at the VERITAS site in Jan. 2019



