

31st Rencontres de Blois – June 5, 2019

**An Updated Measurement of the
Hubble Constant from the H_0 Lenses in
COSMOGRAIL's Wellspring (H0LiCOW)**

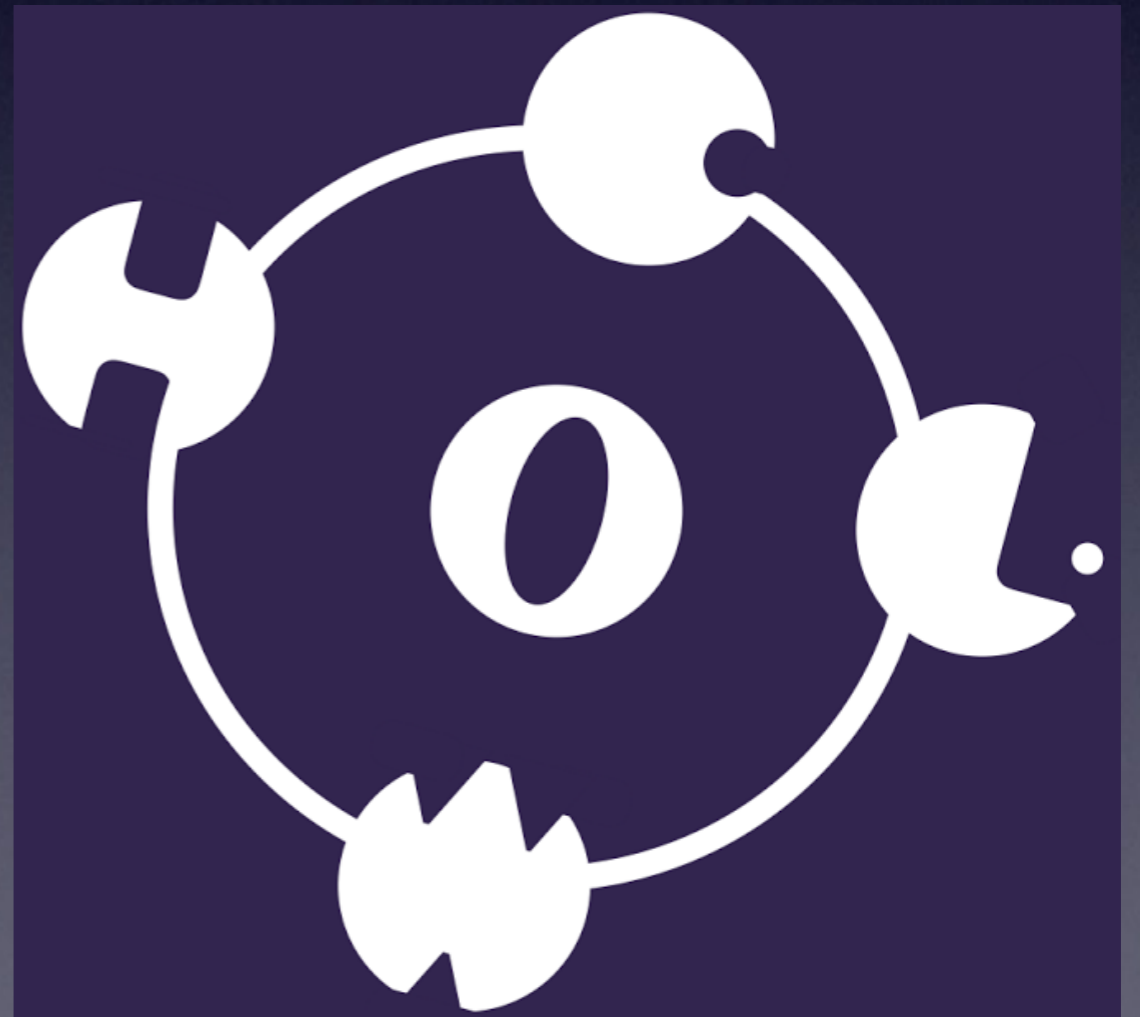
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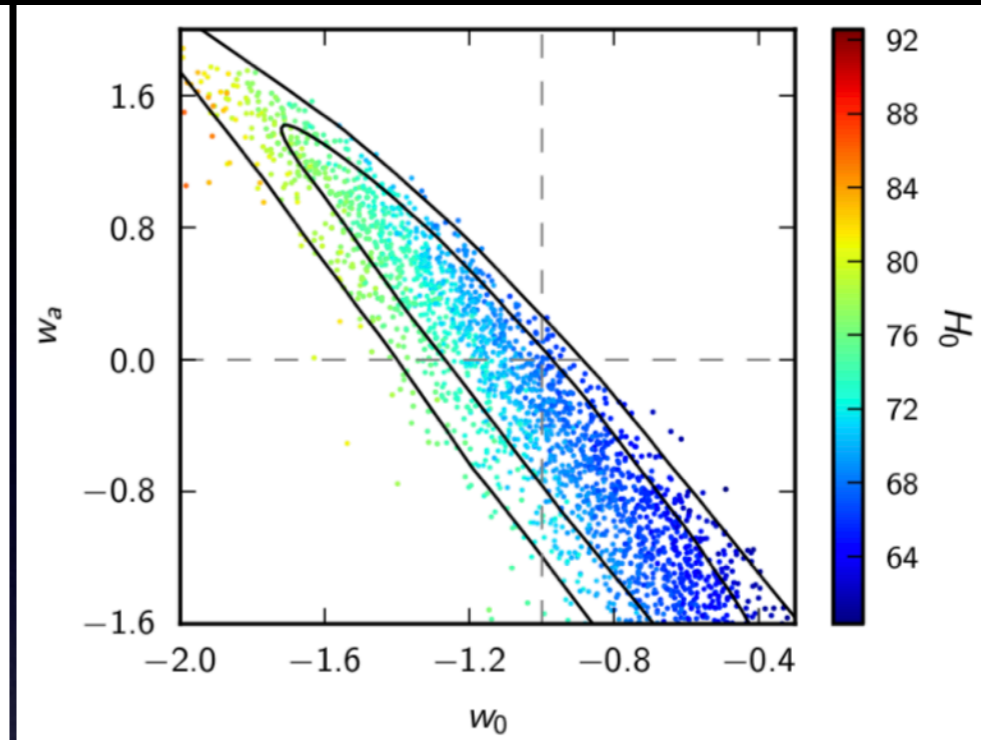
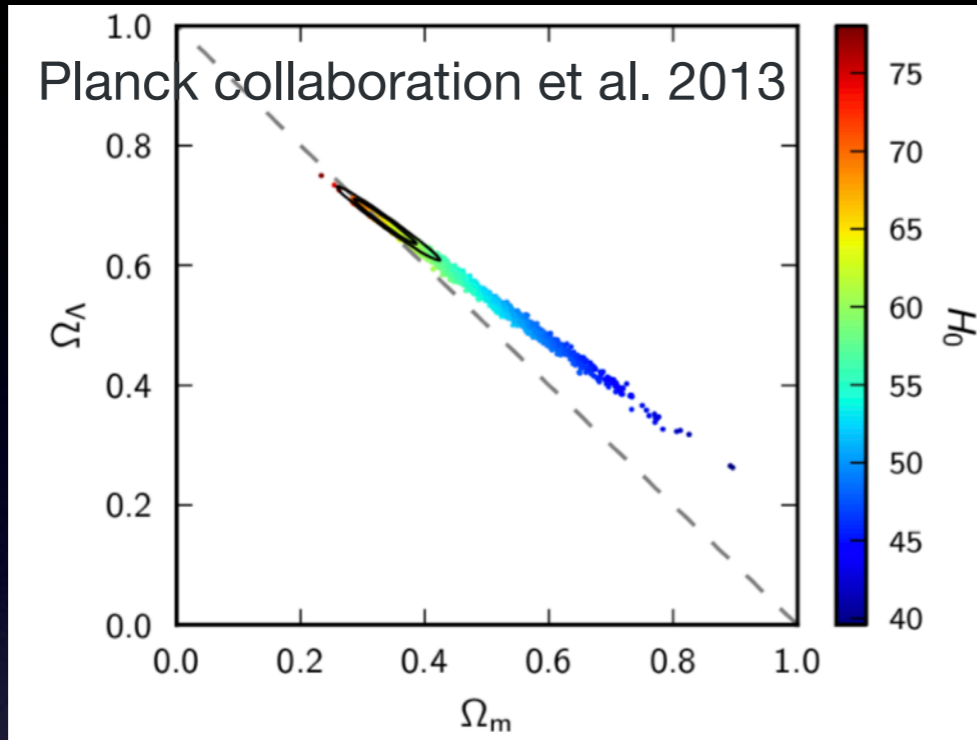
Outline

- Motivation: the Hubble Constant
- Constraining cosmological parameters with gravitational lens time delays
- H0LiCOW and our newest results
- Towards the future



Motivation: The Hubble Constant H_0

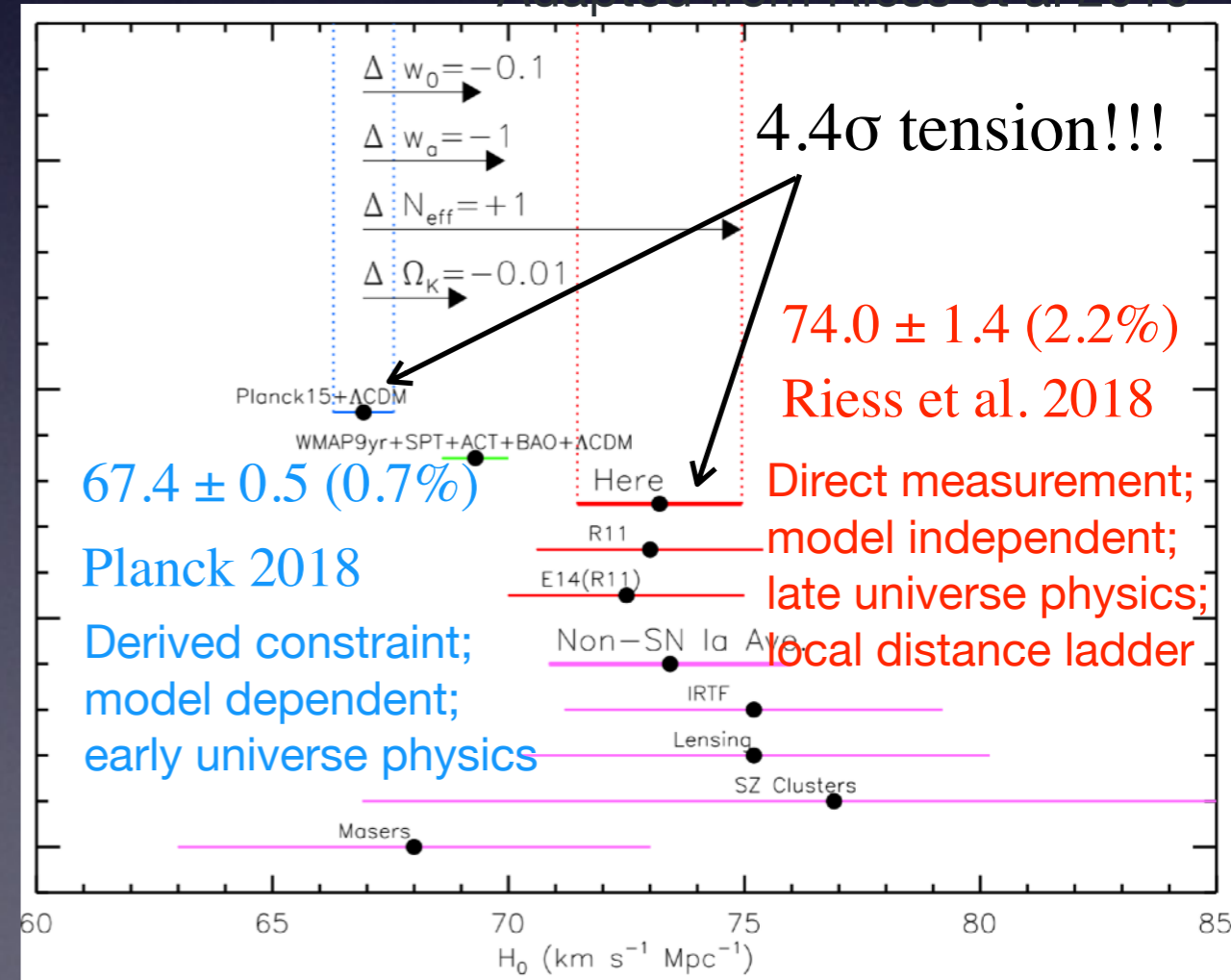
- H_0 is tightly constrained by Planck, but only under in the 6-parameter flat Λ CDM



Currently, there is tension between H_0 measurements from different probes...

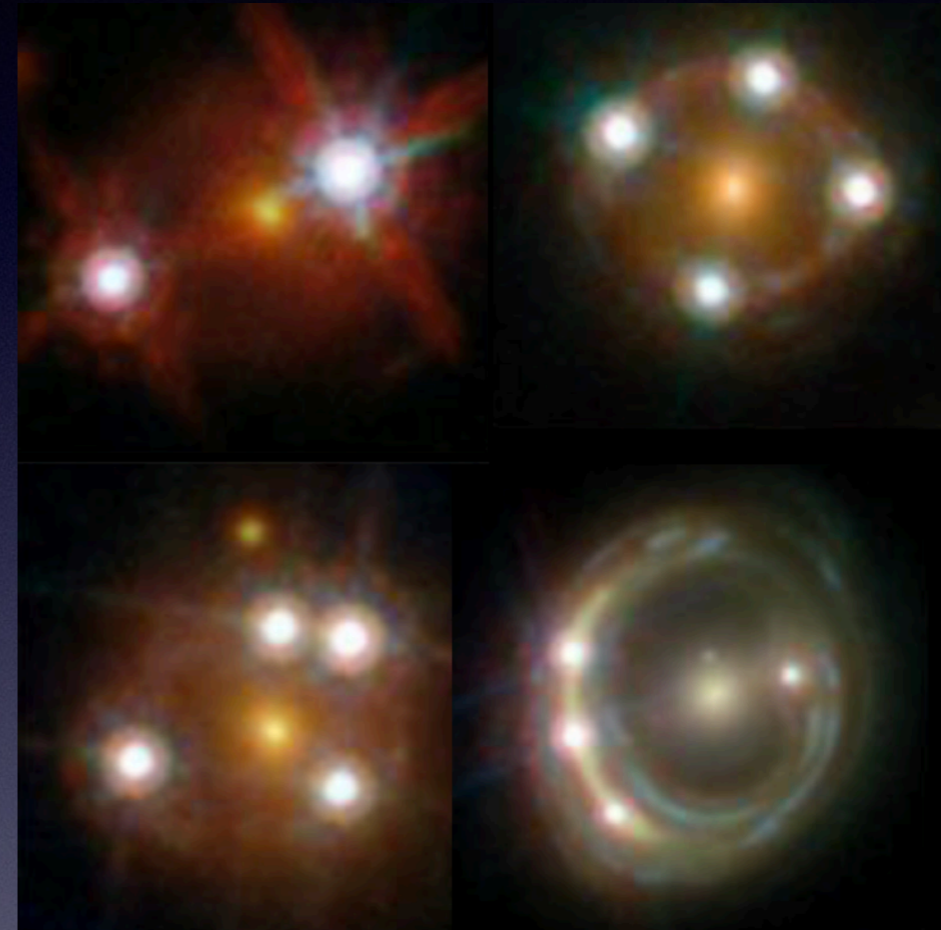
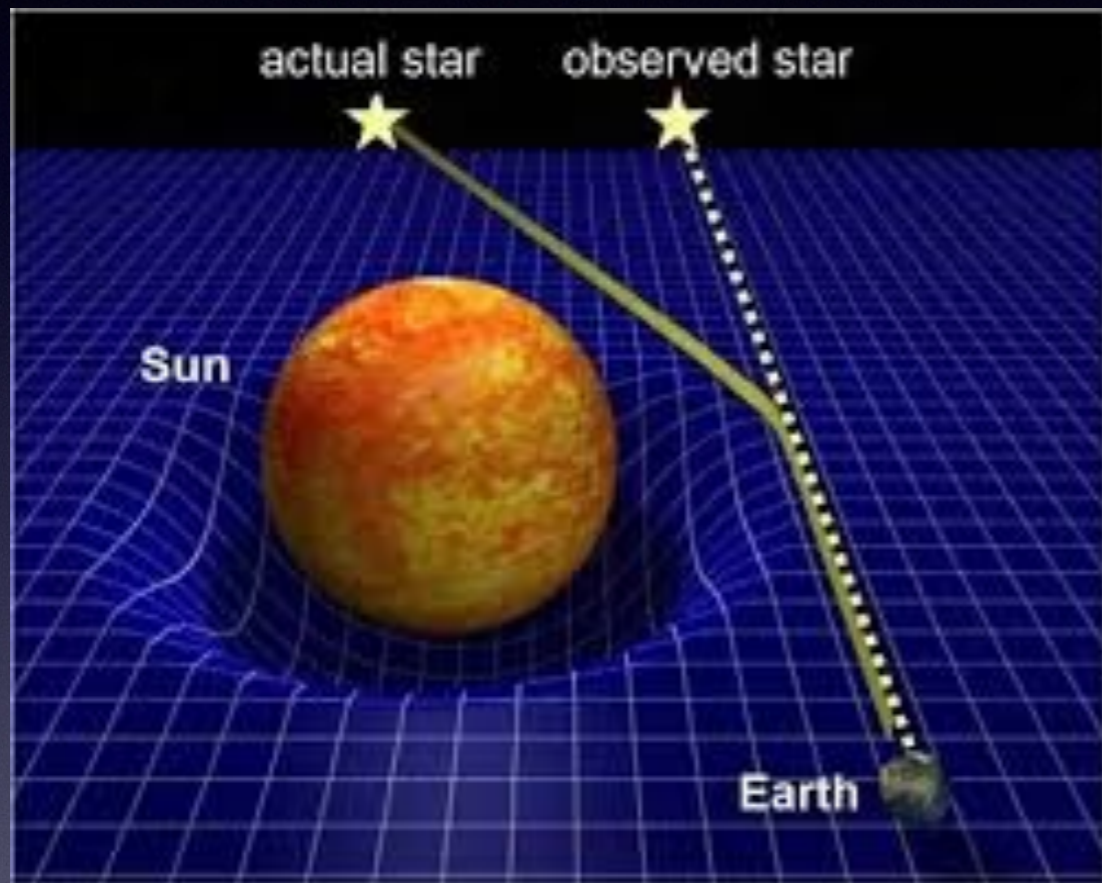
- 1) systematic measurement errors?
- 2) New physics?

An independent method of comparable precision is needed to test for hidden systematics and check for consistency



Constraining cosmology with GL time delays: I

- General Relativity: mass curves spacetime -> deflection of light rays

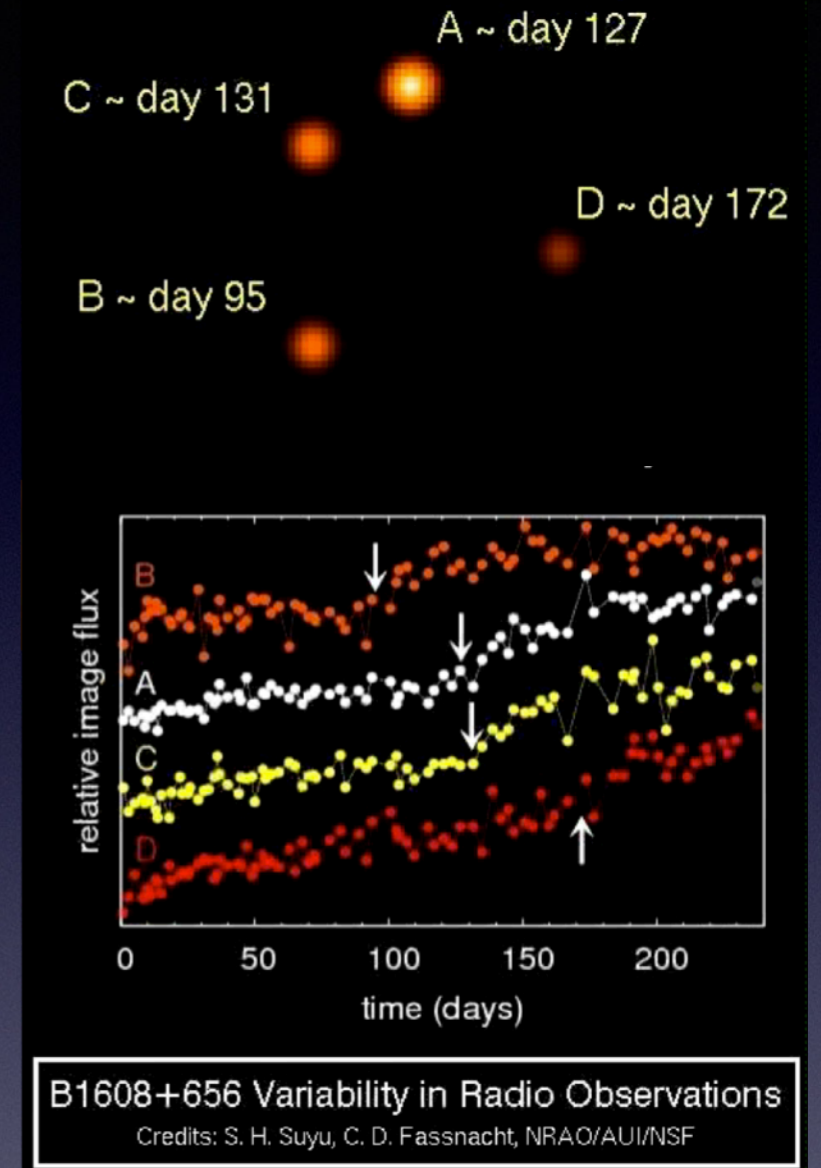
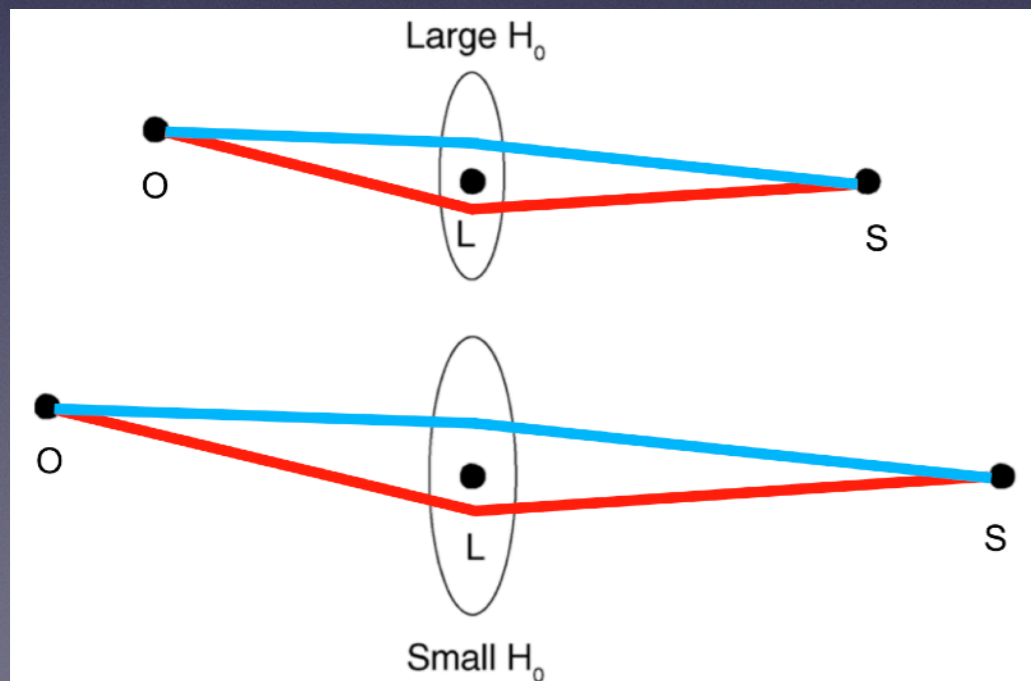


Hubble Telescope images of GLs

- With the right alignment, massive galaxies acting as GLs produce multiple images of background sources

Constraining cosmology with GL time delays: II

- Quasars are powered by accretion into SMBH
- Quasars “flicker” in time

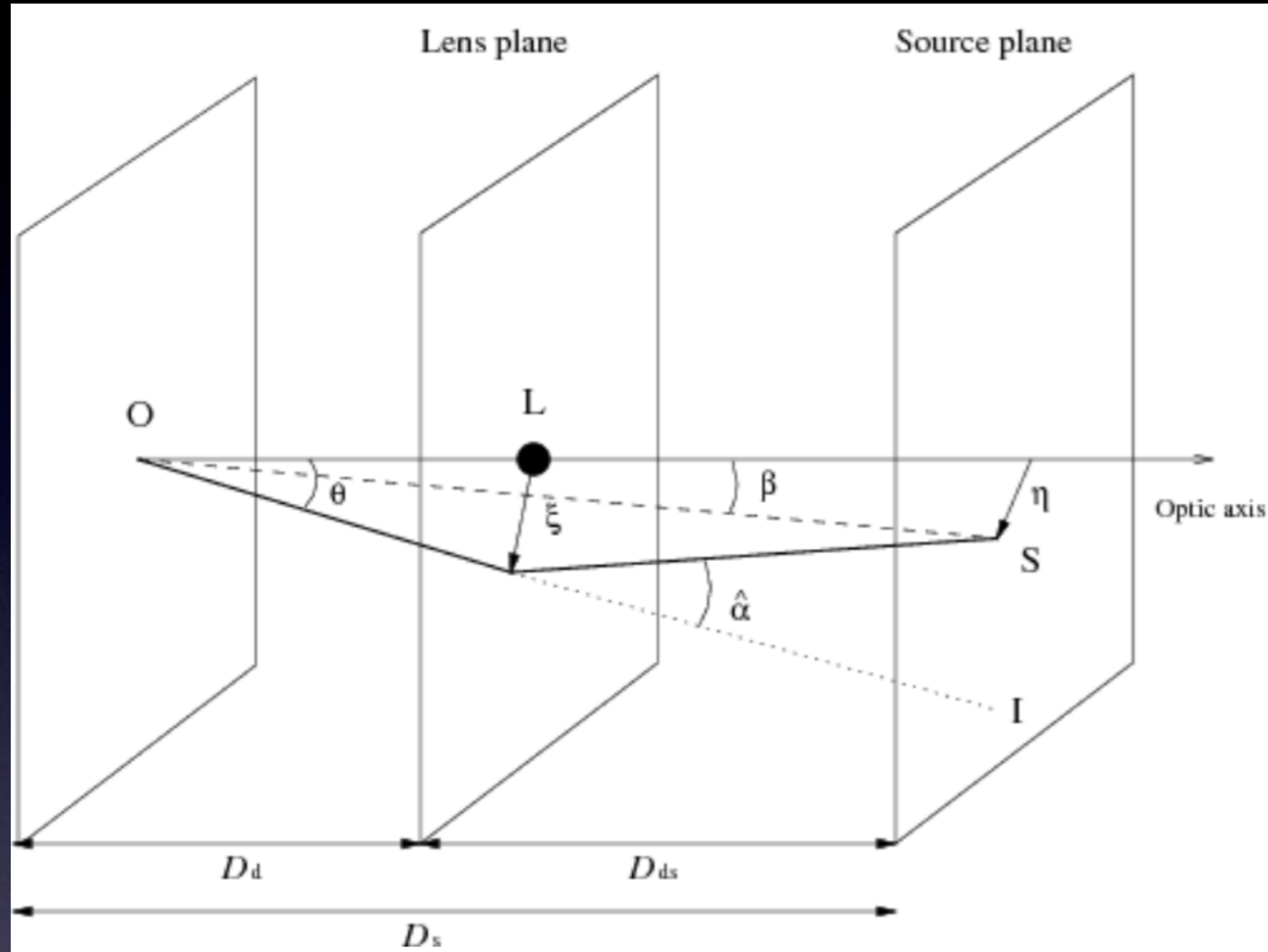


$$v = H_0 d$$

$$\frac{\text{length} - \text{length}}{c} = \Delta t$$

$$\Delta t \propto \frac{1}{c} \frac{1}{H_0}$$

Constraining cosmology with GL time delays: III



Angular diameter distances:

$$D(z) = \frac{c}{H_0} \frac{1}{1+z} \int_0^z \frac{dz'}{\sqrt{\Omega_m(1+z')^3 + \Omega_\Lambda}}$$

$$t(\theta, \beta) = \frac{1}{c} \frac{D_d D_s}{D_{ds}} (1+z_d) \phi \left[\frac{(\theta - \beta)^2}{2} - \psi(\theta) \right]$$

$$D_{\Delta t} \propto \frac{1}{H_0} \begin{array}{l} \text{time delay} \\ \text{distance} \end{array}$$

We measure relative time delays between pairs of images:

$$\Delta t = \frac{1}{c} D_{\Delta t} \Delta \phi$$

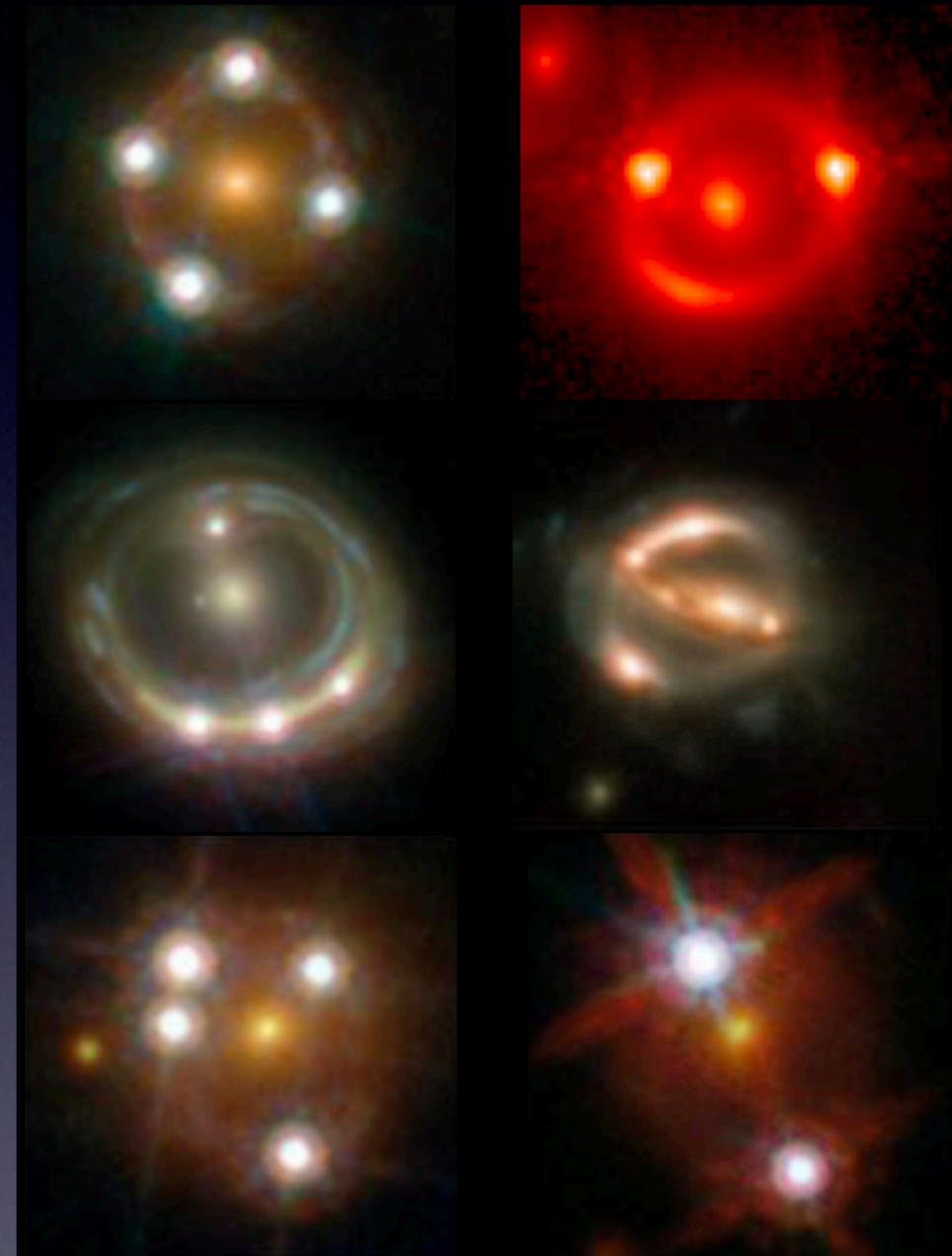
↑ monitoring
 ← lens modeling

$D_{\Delta t}$ encloses the cosmological information, and is primarily sensitive to H_0

- based on angular diameter distances → immune to dust, no calibration issues
- an absolute distance; provides a distance-redshift relation
- Based on well-understood physics (GR)

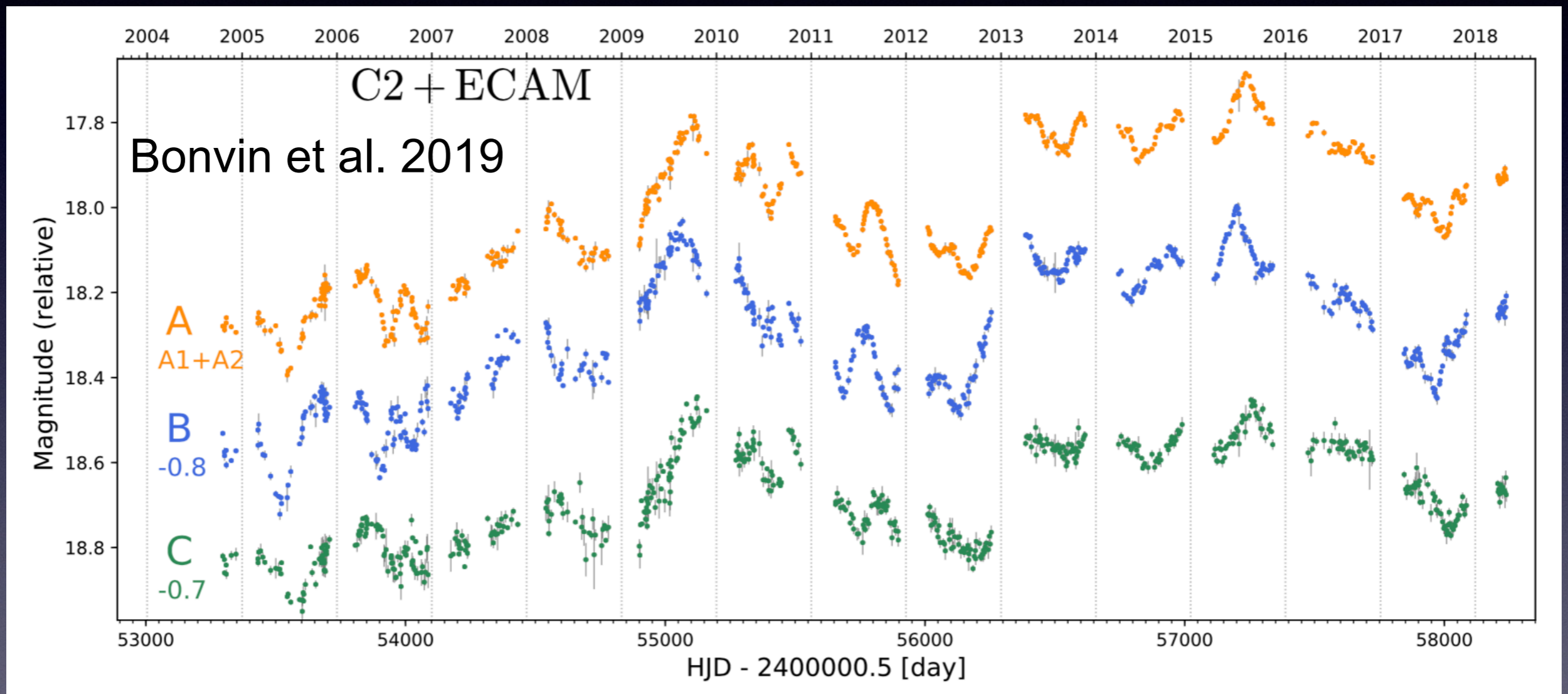
HOLiCOW: our mission

- We focus on GL systems for which we can gather an exhaustive set of data, and account in detail for all known systematics
- a single lens with well-measured time delays can be used to measure time delay distances to 6-7% uncertainty (random and systematic; Suyu et al. 2010, 2013)
- Thus, when doing joint inference we insure both precision and accuracy



HOLiCOW: what do we need?

- Time delays measured of ~ 10 years from COSmological MOnitoring of GRAVitational Lenses (COSMOGRAIL)



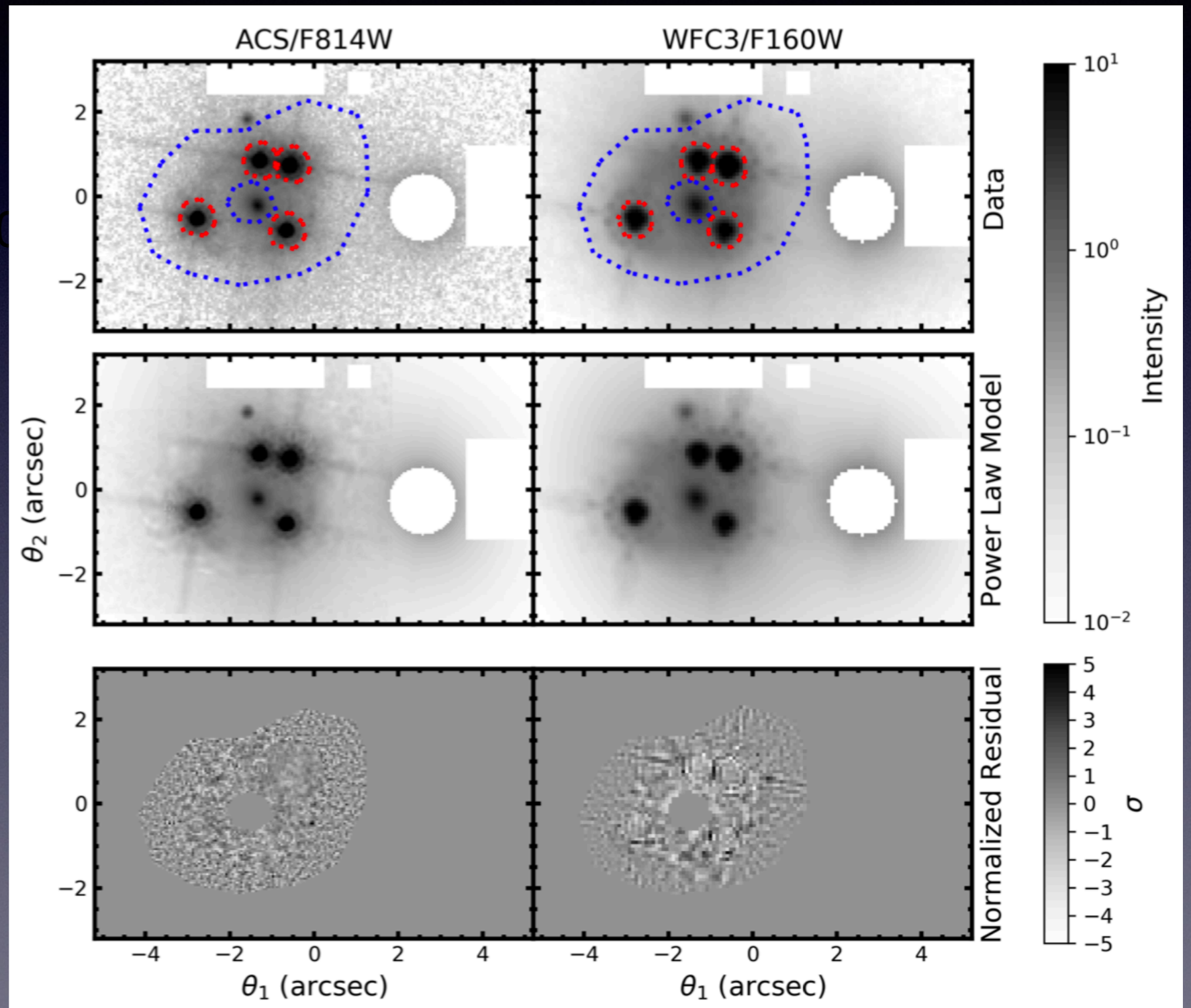
$$\Delta t = \frac{1}{c} D_{\Delta t} \Delta \phi$$

HOLiCOW: what do we need?

- High-resolution Hubble Telescope images of the lenses, using the constraints from 1000s of pixels in the arcs to constrain the mass potential

Bonvin et al. 2019

Rusu et al. 2019



$$\Delta t = \frac{1}{c} D_{\Delta t} \Delta \phi$$

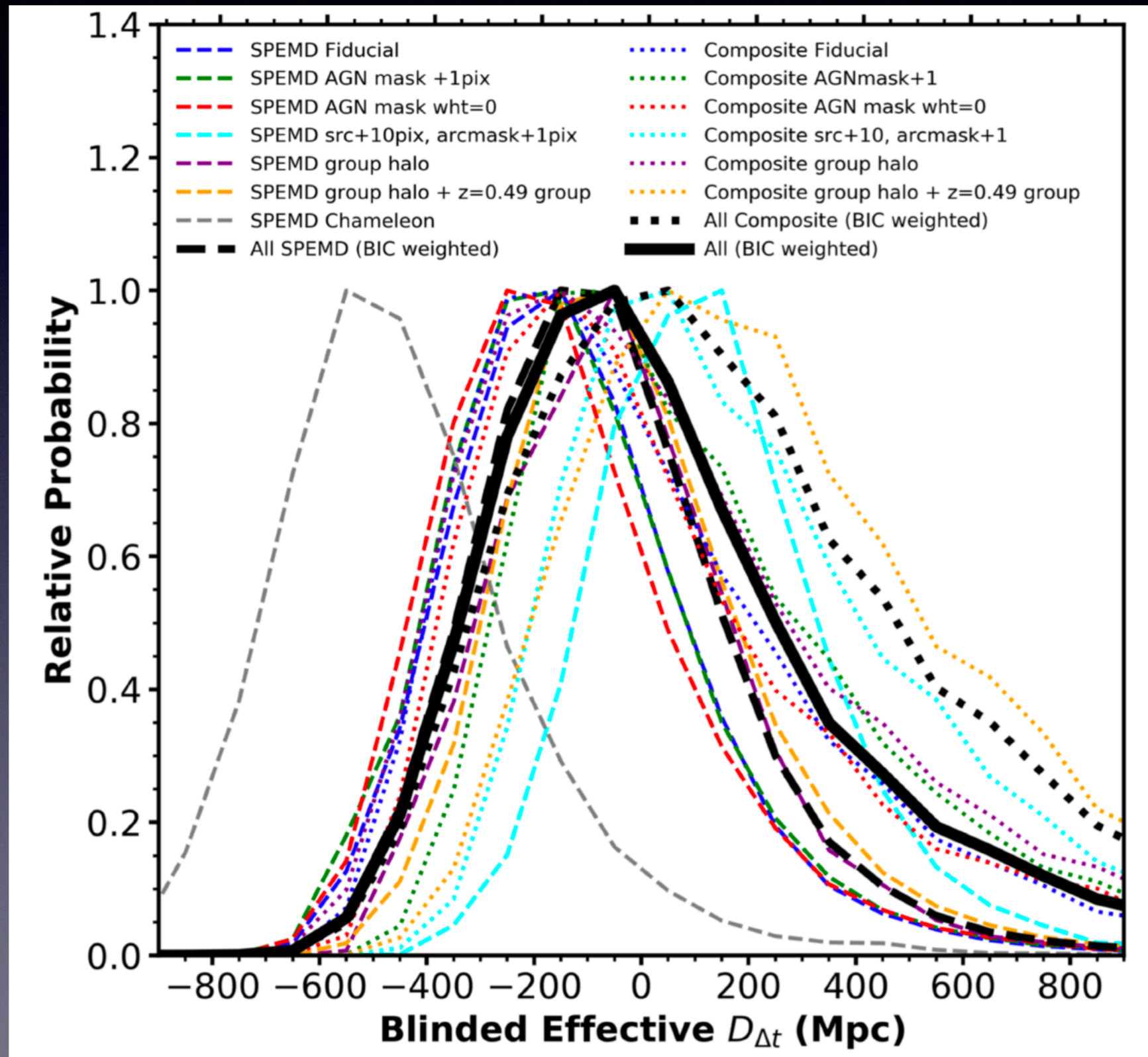
HOLiCOW: results from the latest GL

- The probability distribution for the time delay distance:

- test for systematics

- blinded determination with centroids offset to origin, following SNe Ia approach in Conley et al. 2006

Rusu et al. 2019

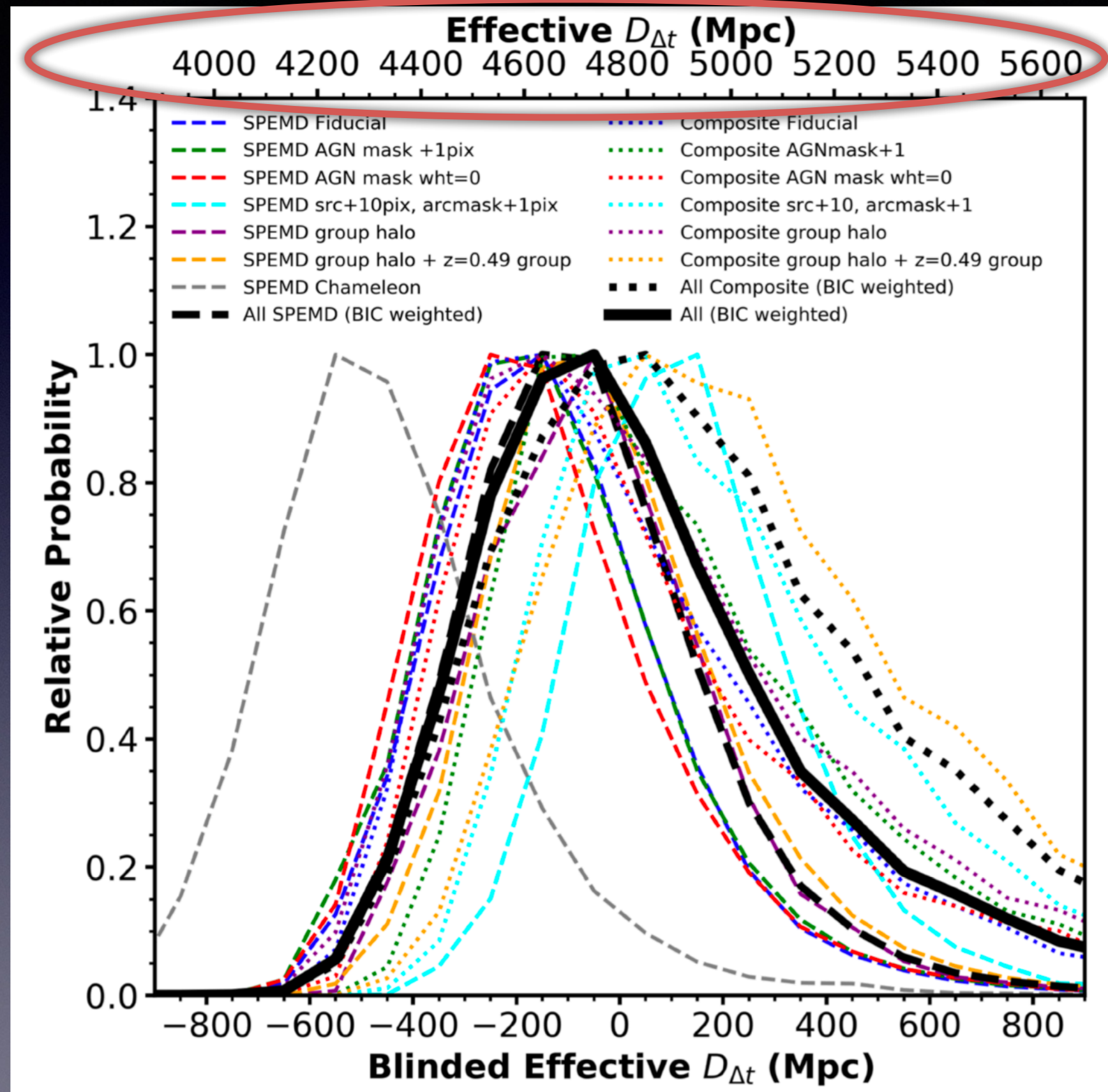


HOLiCOW: results from the latest GL

unblinding on 7 May 2019:

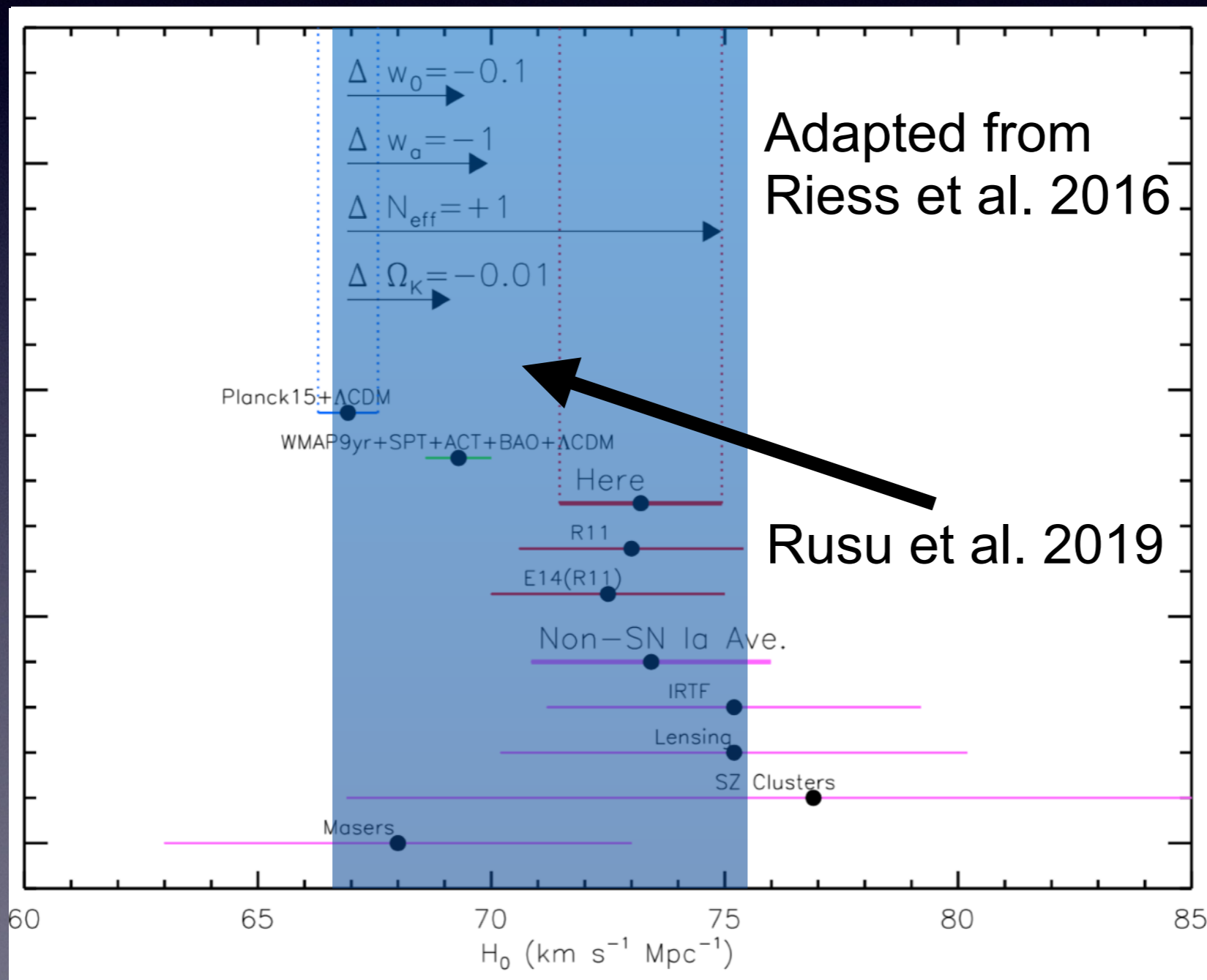
~ 6.6 % uncertainty

$$D_{\Delta t} \propto \frac{1}{H_0}$$



HOLiCOW: results from the latest GL

- A single lens produces results in agreement with the other probes, but poor precision
- At present, we have independent H0 measurements from 5 lenses
- Compute Bayes factor F to test if the 5 lenses are consistent with same cosmology



$$F = \frac{P(d_1, \dots, d_5 | H^{\text{global}})}{\prod_{I=1, \dots, 5} P(d_i | H^{\text{ind}, i})} > 1$$

HOLiCOW: results from 5 lenses

UNPUBLISHED

HOLiCOW: results from 5 lenses

UNPUBLISHED

HOLiCOW: results from 5 lenses

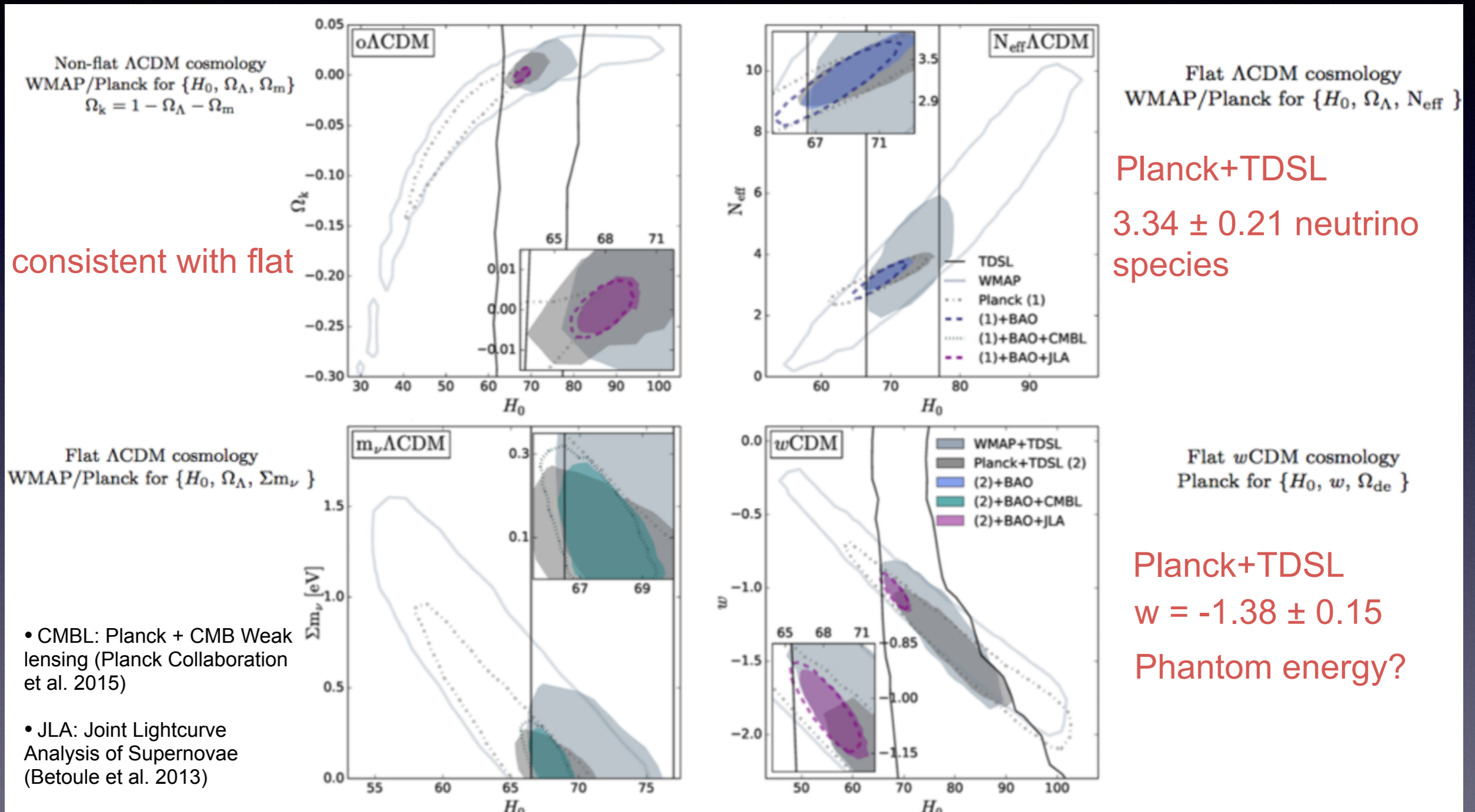
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HOLiCOW: results from 5 lenses

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- Tighter constraints are obtained by combining with other probes (complementarity -> break parameter degeneracies)

- Combining 3 H0LiCOW lenses with other probes (Bonvin et al. 2017):



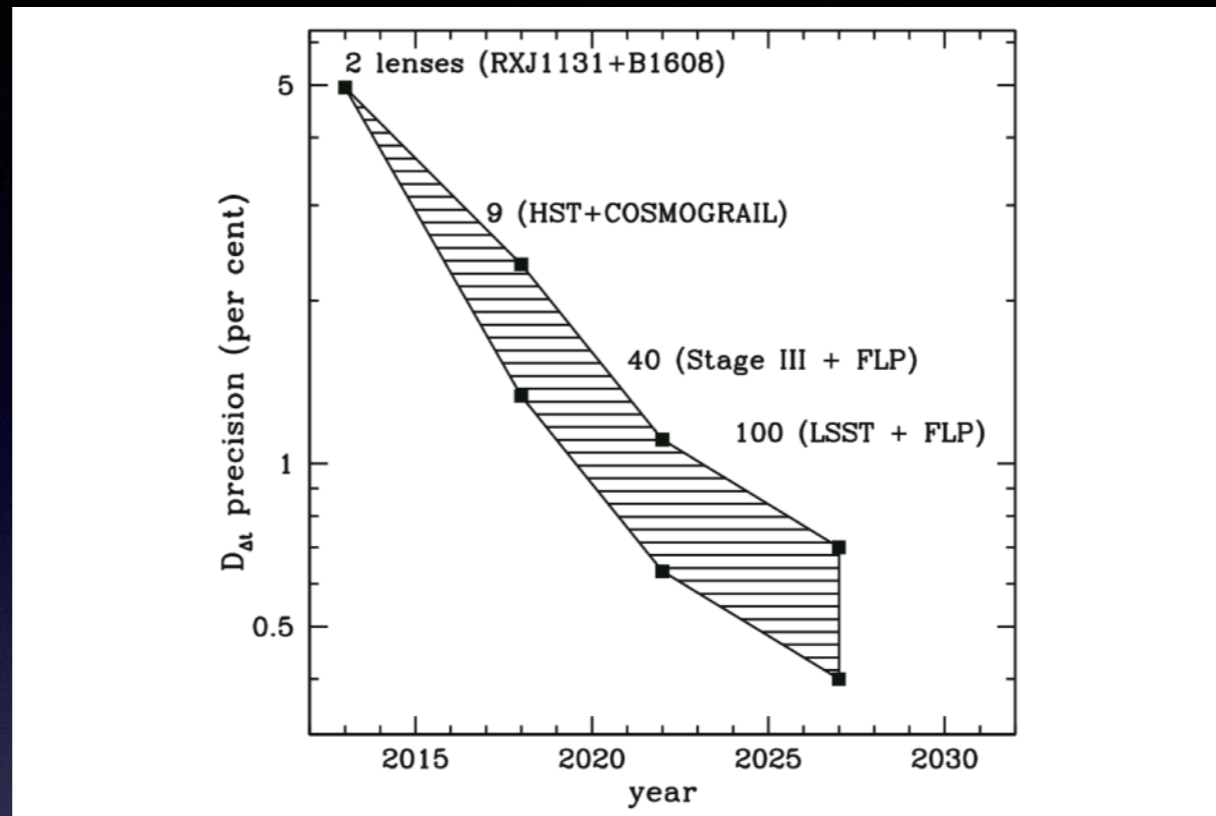
- Wong et al. 2019 will show the combination with 6 lenses

HOLiCOW: the future

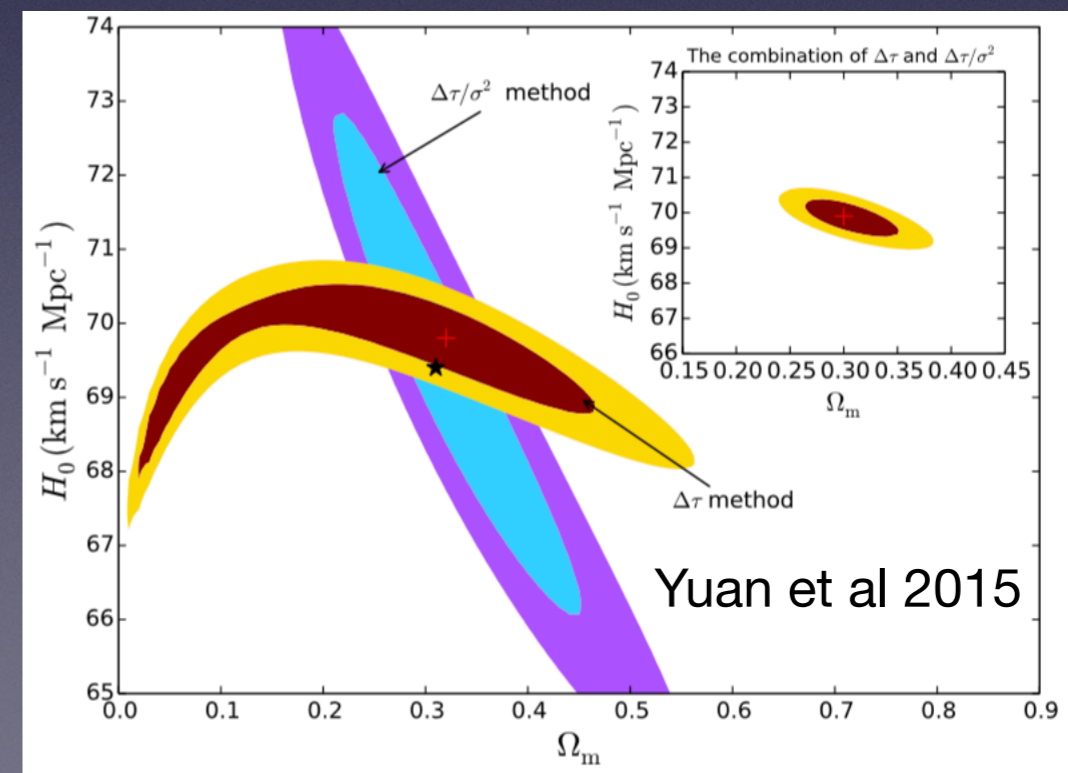
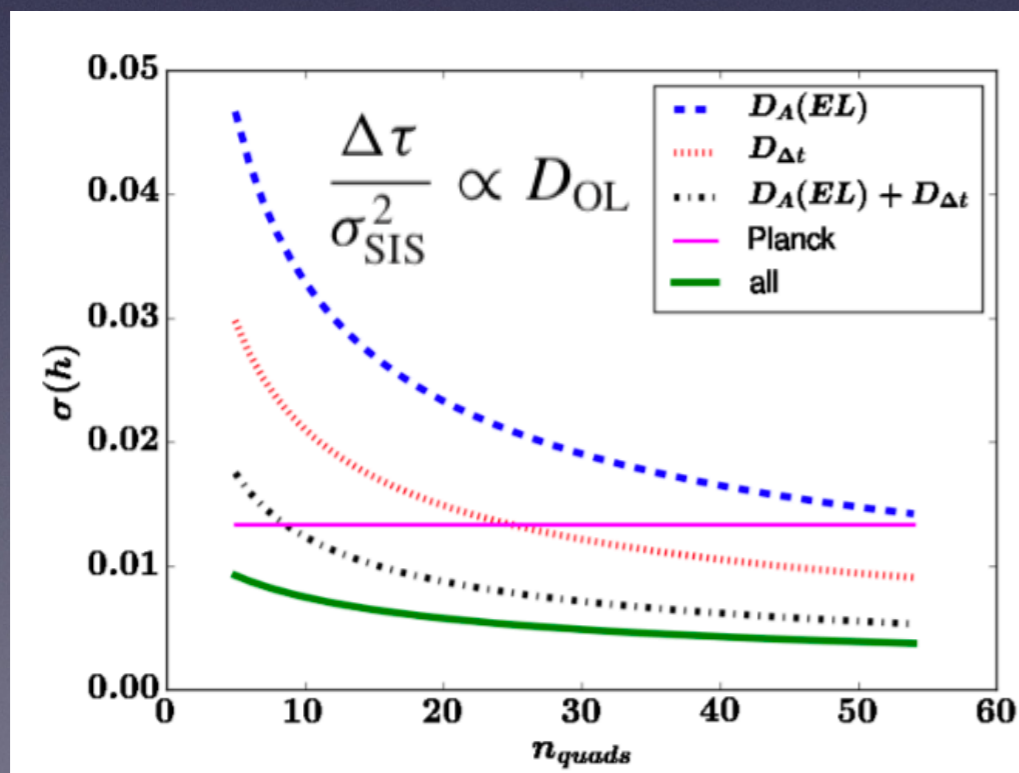
- Things are accelerating:
 - Suyu et al. 2010
 - Suyu et al. 2013
 - Wong et al. 2017
 - Birrer et al. 2019
 - Rusu et al 2019
 - Chen et al. 2019, in prep.
- Milestone paper: Wong et al. 2019, in prep. Will combine 6 lenses
- We are making progress at improving our technique:
 - Theoretical progress on breaking the mass-sheet degeneracy
 - We will acquire spatially resolved kinematics to constrain anisotropy
 - We are improving imaging modeling using adaptive optics
- LSST will significantly expand the known sample of lensed quasars

HOLiCOW: the future

- Treu & Marshall 2016 extrapolate 1% precision from HOLiCOW with ~40 lenses



- Jee et al. 2016: We can achieve 1% on H0 in Λ CDM with just ~15 lenses



Summary

- H0 measurements from different lenses give mutually consistent results; 5 lenses constrain H0 to X.X% in flat Λ CDM, in good agreement with SH0ES and inconsistent with Planck at X.X σ
- Milestone paper Wong et al. 2019 in prep. will combine the 6-lens H0LiCOW measurement with different cosmological probes
- H0LiCOW is expected to reach 1% precision on H0 from $> \sim 15$ lenses in a few years