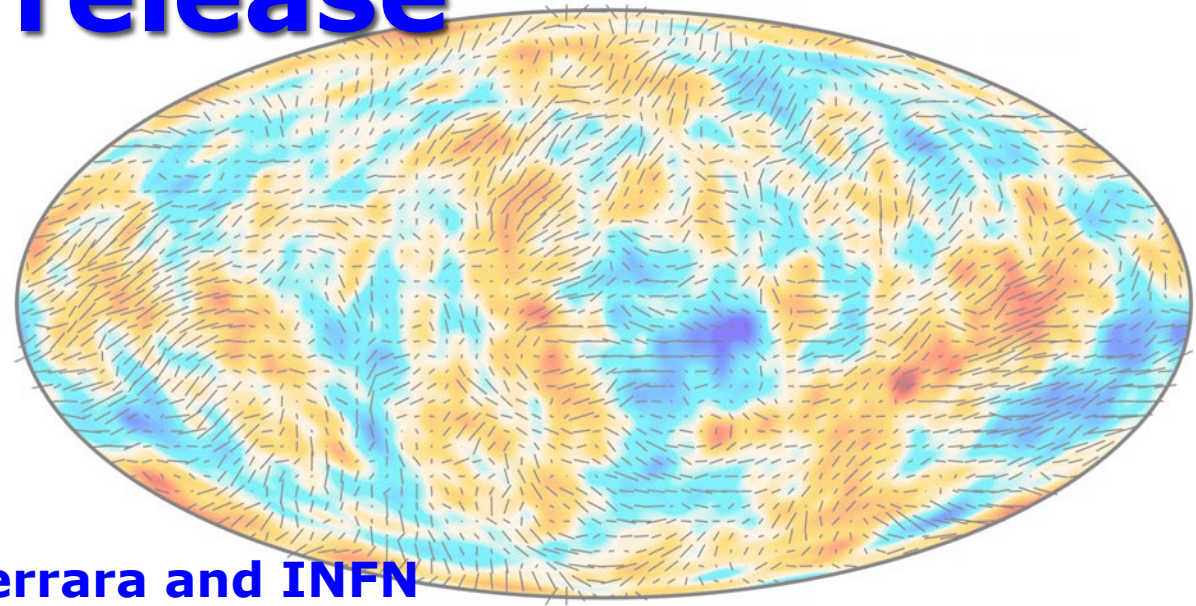




planck

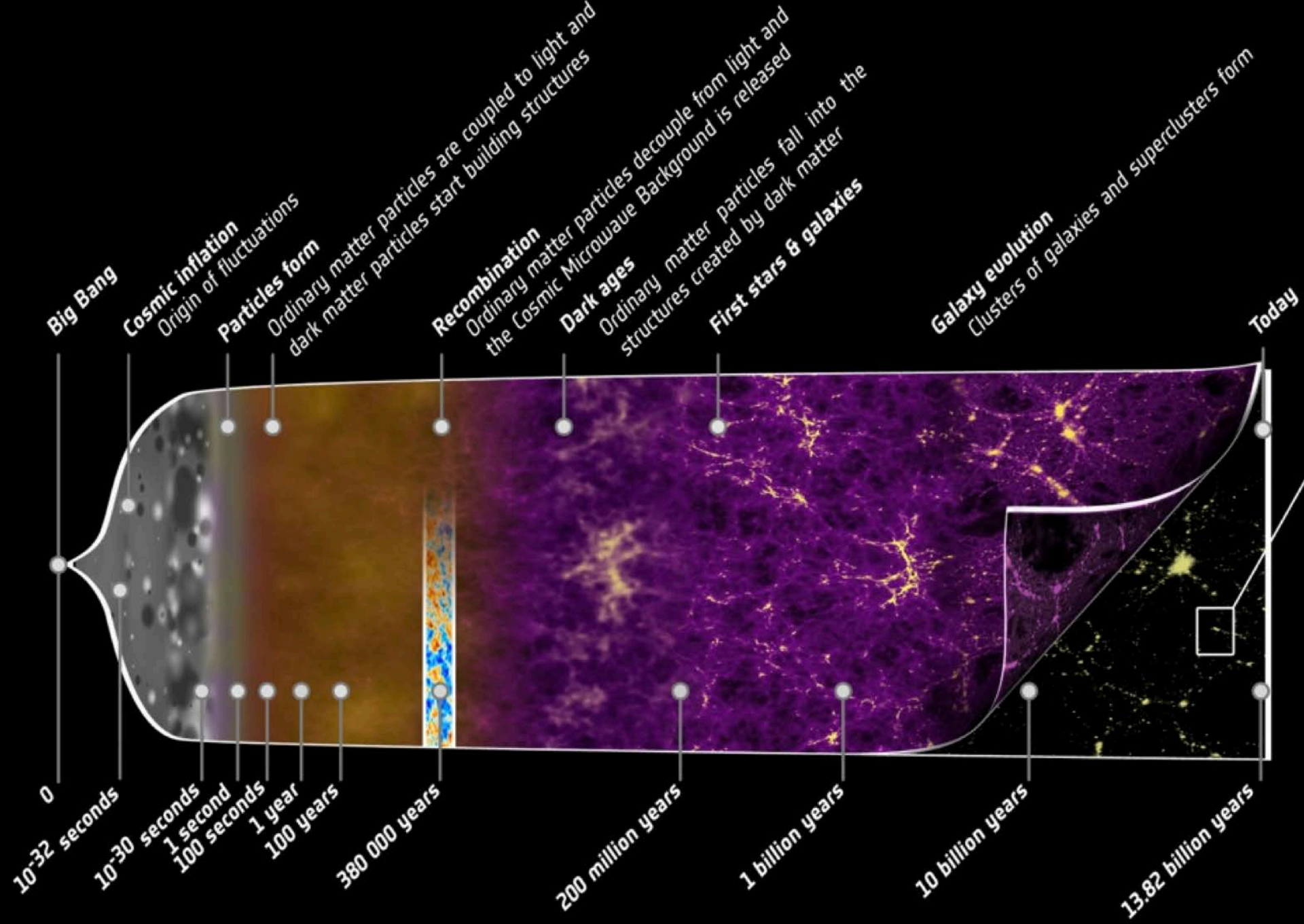
# Cosmology after the Planck legacy release



**Paolo Natoli**

**Università di Ferrara and INFN**

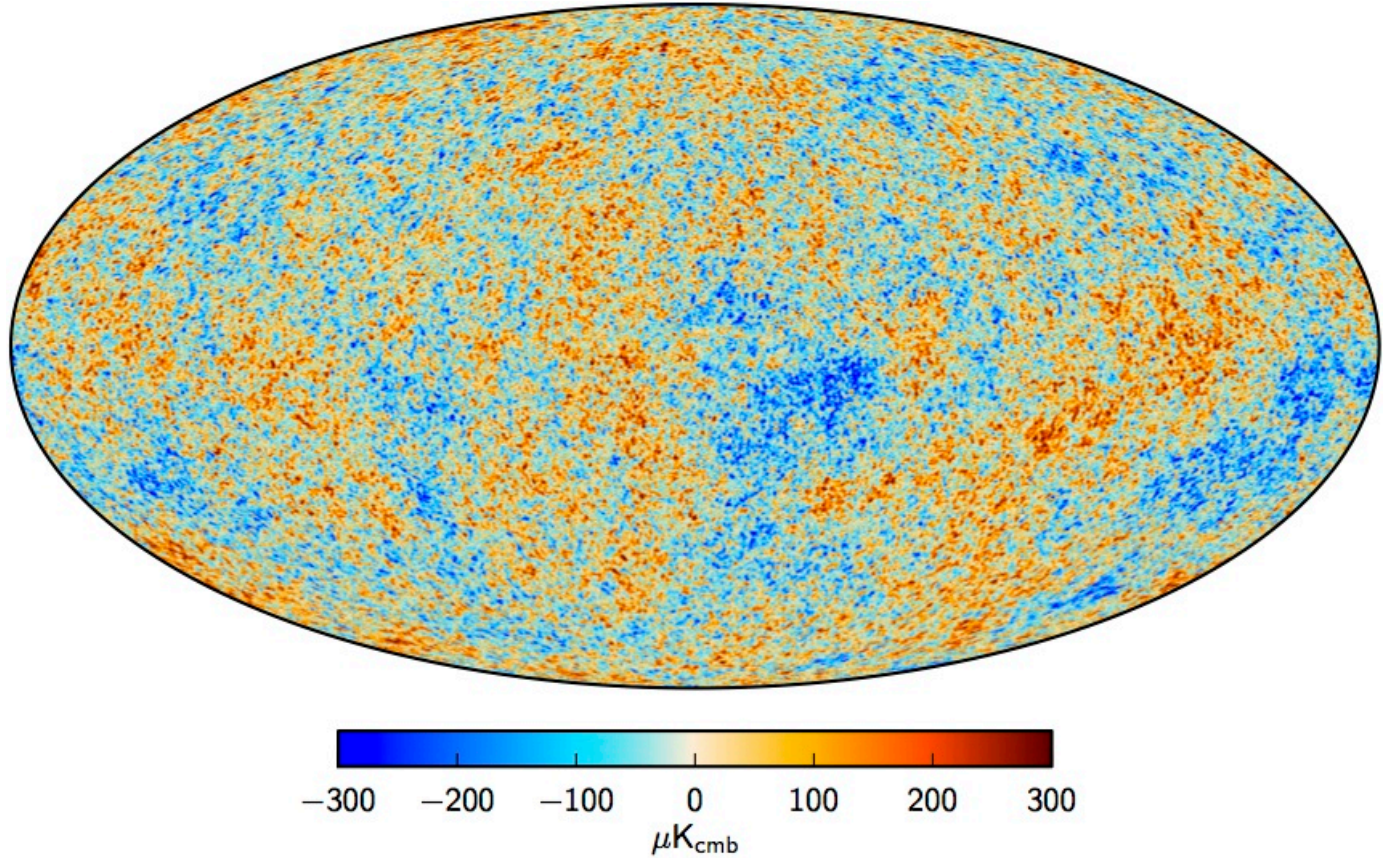
*XVIII International Workshop on Neutrino Telescopes  
Venezia, 21 March 2019*





# THE COSMIC MICROWAVE BACKGROUND

---



The CMB is a blackbody radiation with  $T=2.7$  K extremely uniform across the whole sky; it is the relic radiation emitted at the time the nuclei and electrons recombined to form neutral hydrogen, when the Universe was  $\sim 400,000$  years old (the so-called last scattering surface, LSS).

Its tiny ( $\sim 10^{-5}$ ) temperature and polarization anisotropies encode a wealth of cosmological information.

# THE PLANCK SATELLITE



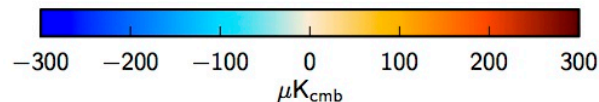
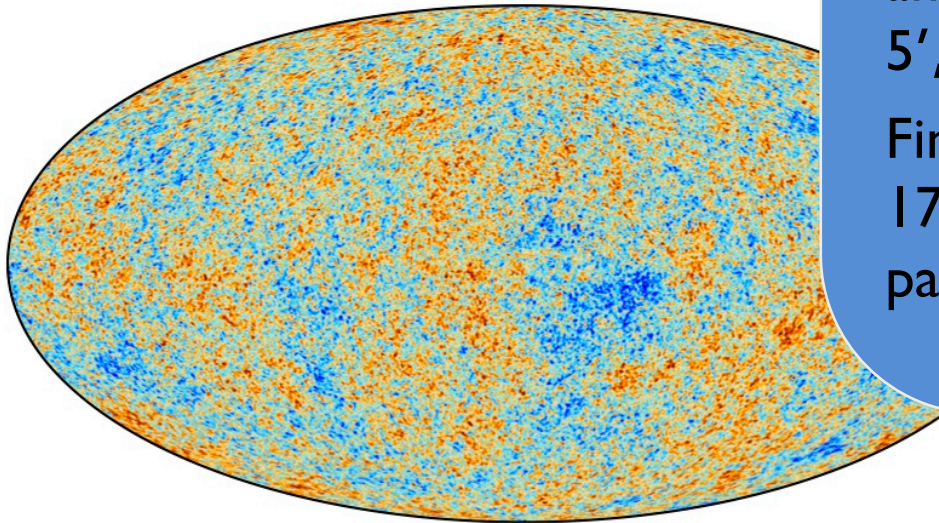
Planck is the 3<sup>rd</sup> generation ESA satellite devoted to CMB

Ultimate characterization of the temperature anisotropies

74 detectors (radiometers and bolometers) in 9 frequency bands from 30 to 857 GHz

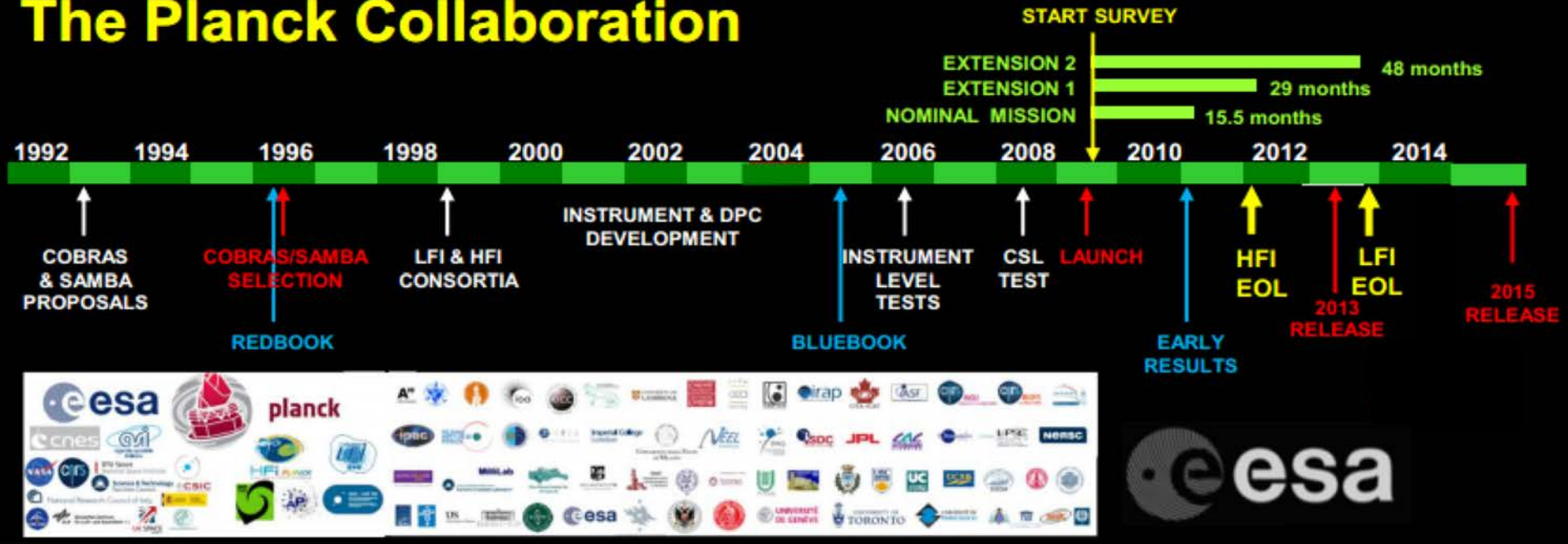
angular resolution between 30' and 5',  $\Delta T/T \sim 2 \times 10^{-6}$

Final (legacy) release took place on 17 July 2018, for data and (most) papers.

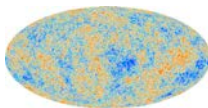




# The Planck Collaboration



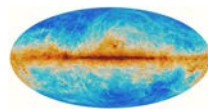
**May 2009:** Launched from Kourou



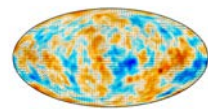
**Mar 2013:** Data Release and Cosmology Results  
 Nominal Mission Temperature data 32 papers



**Oct 2013:** Planck 'Shut Down' 52 papers / intermediate results

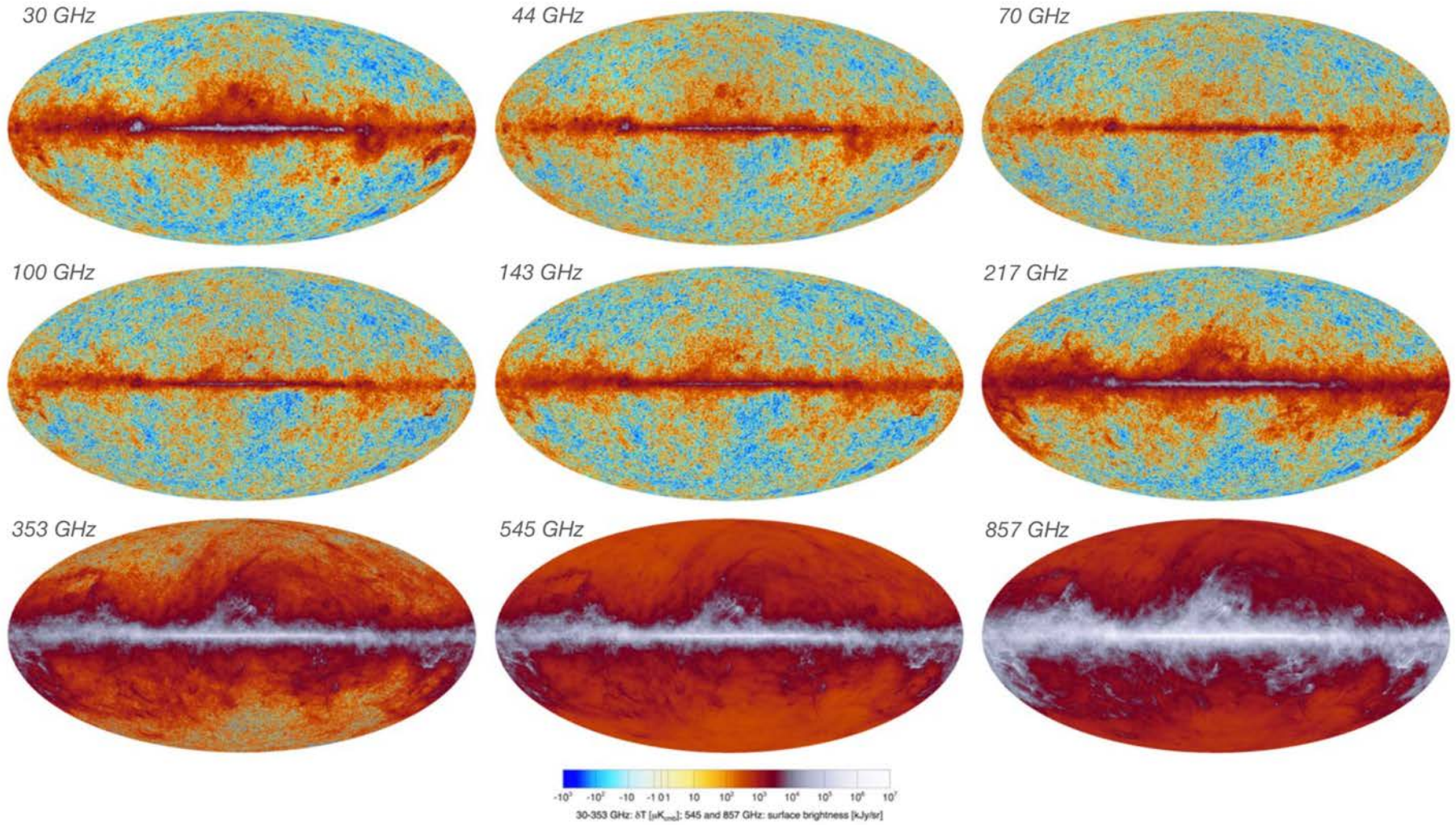


**Feb 2015:** Data Release and Cosmology Results  
 Full Mission Temperature and (preliminary) Polarization data 28 papers



**Jul 2018:** Legacy Data & Paper Release 9 papers (+3 to appear soon)

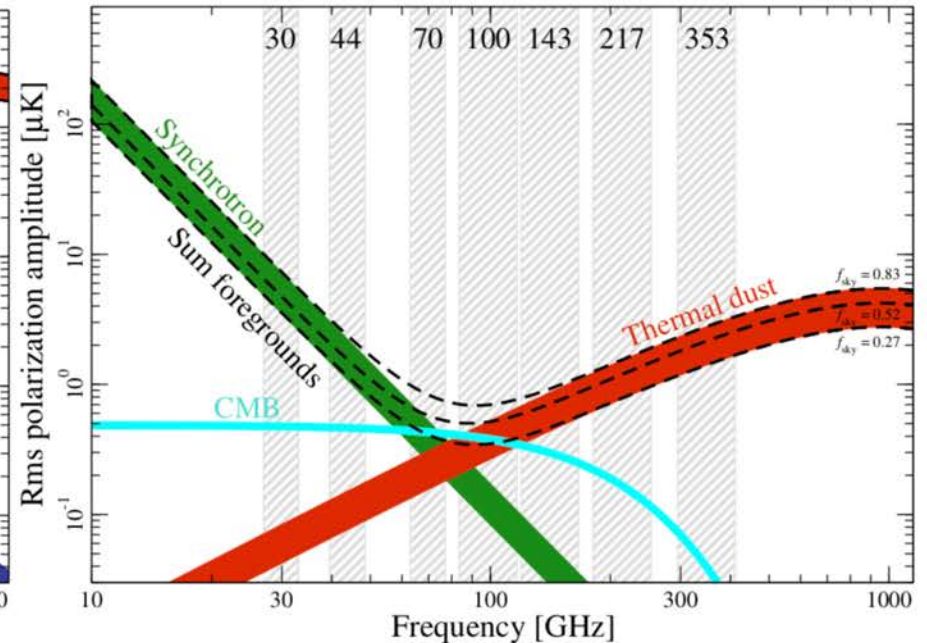
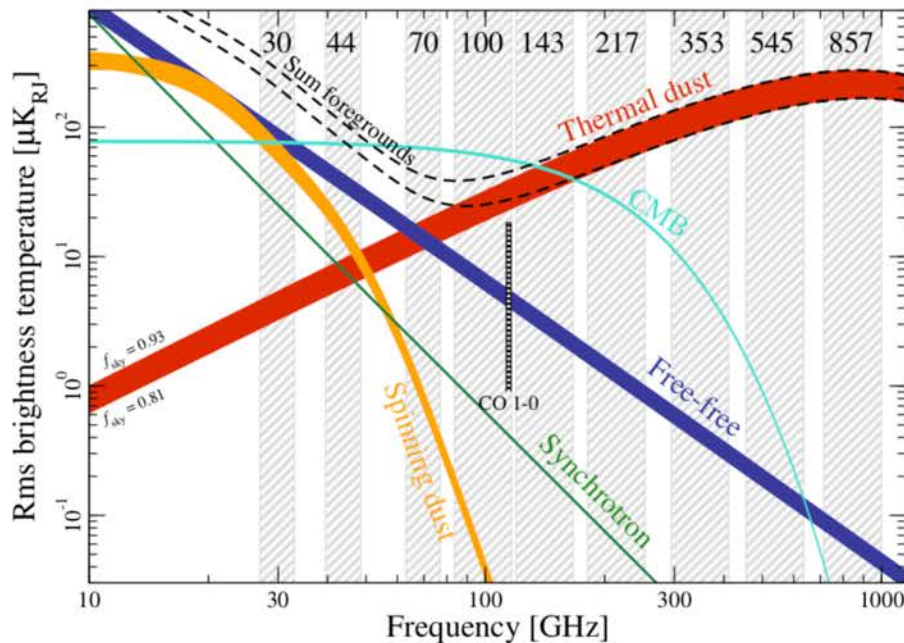
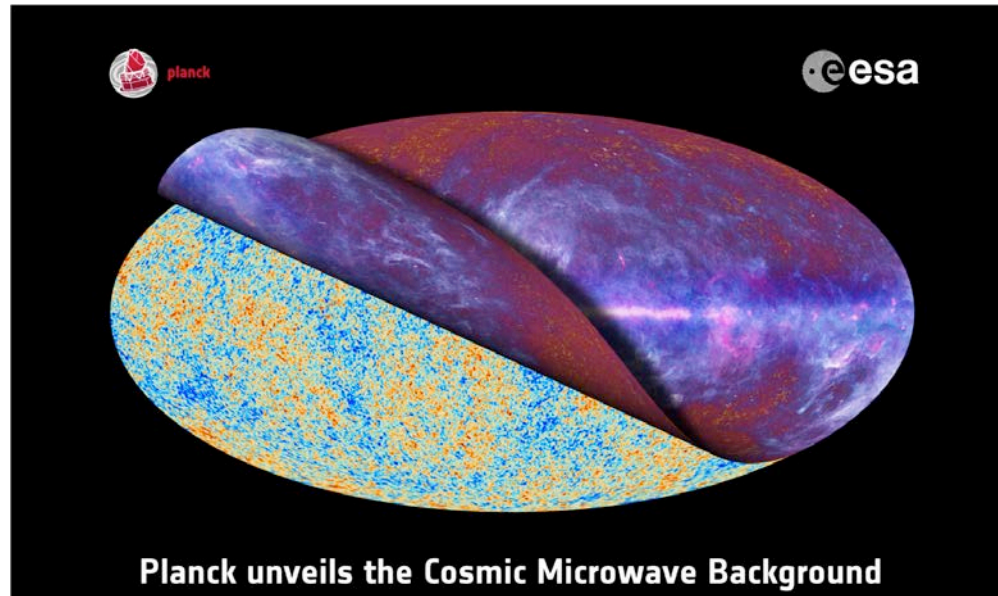
# THE TEMPERATURE SKY AS SEEN BY PLANCK 2018





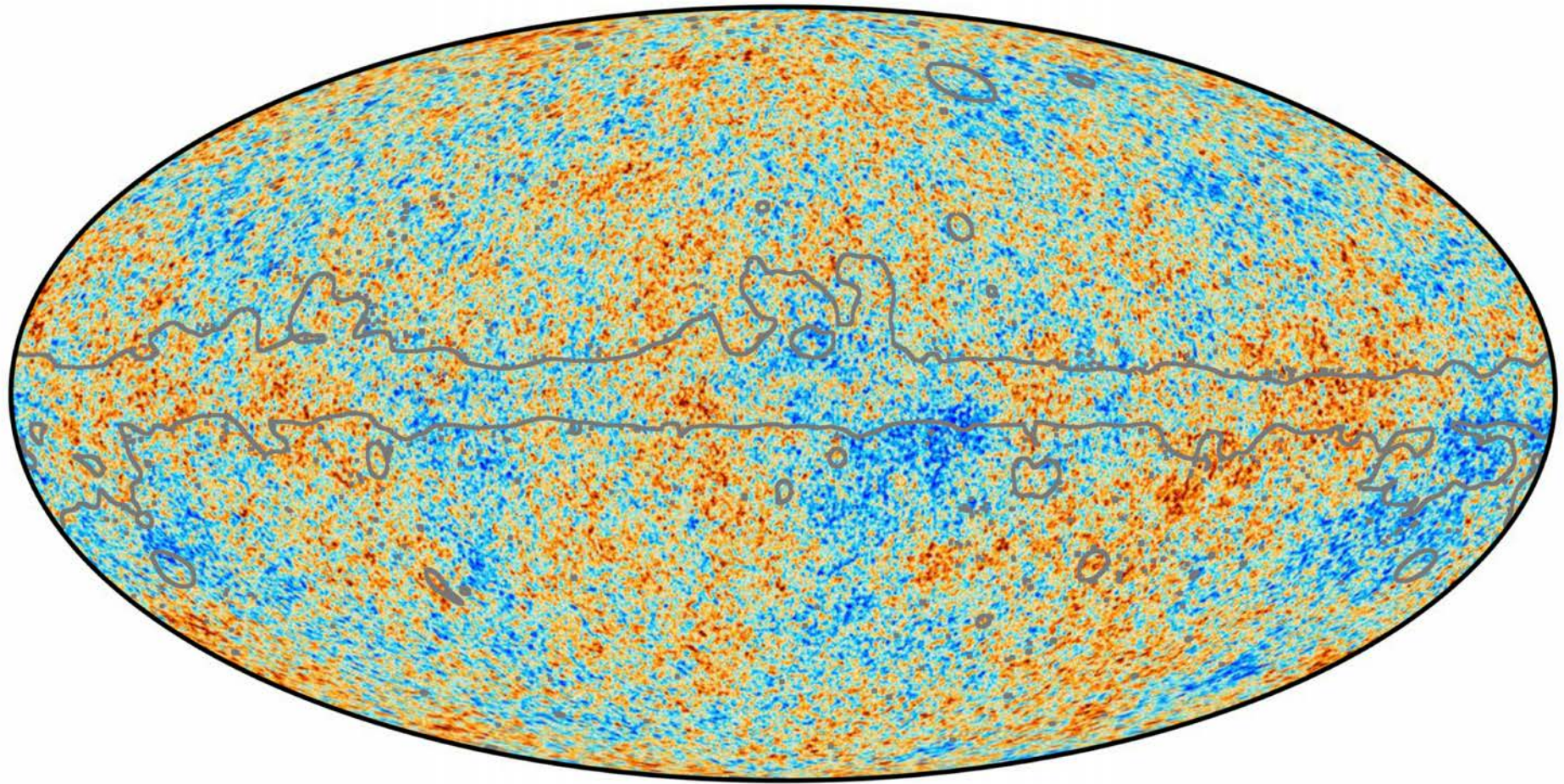
# UNVEILING THE CMB SKY

The *ultimate* measurement of the CMB temperature anisotropy field





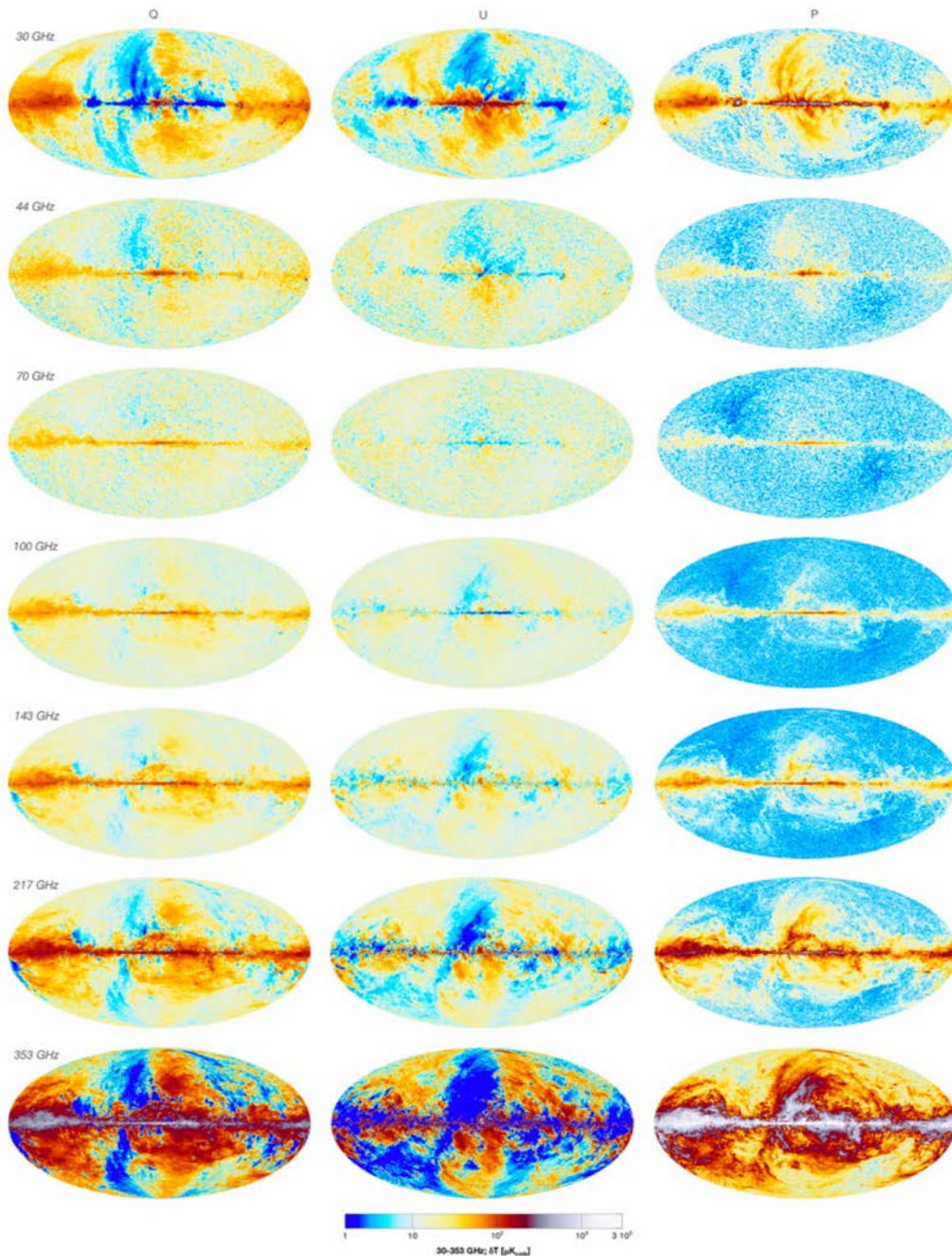
# PLANCK: TEMPERATURE ANISOTROPIES



-300  300  $\mu\text{K}$

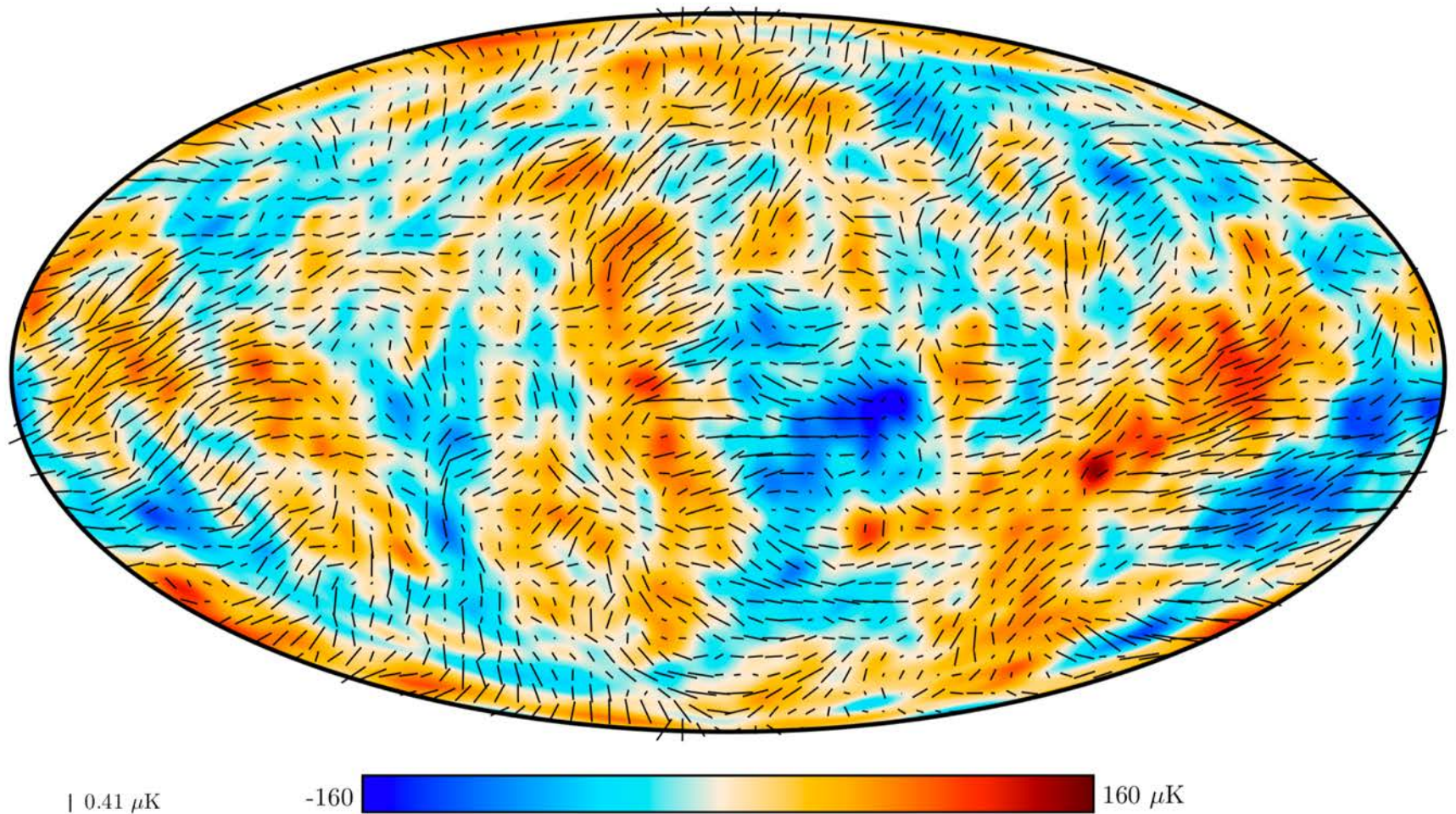


# THE POLARIZATION SKY AS SEEN BY PLANCK 2018



Significant reduction of large scale polarization systematics in 2018

# PLANCK: POLARIZATION ANISOTROPIES



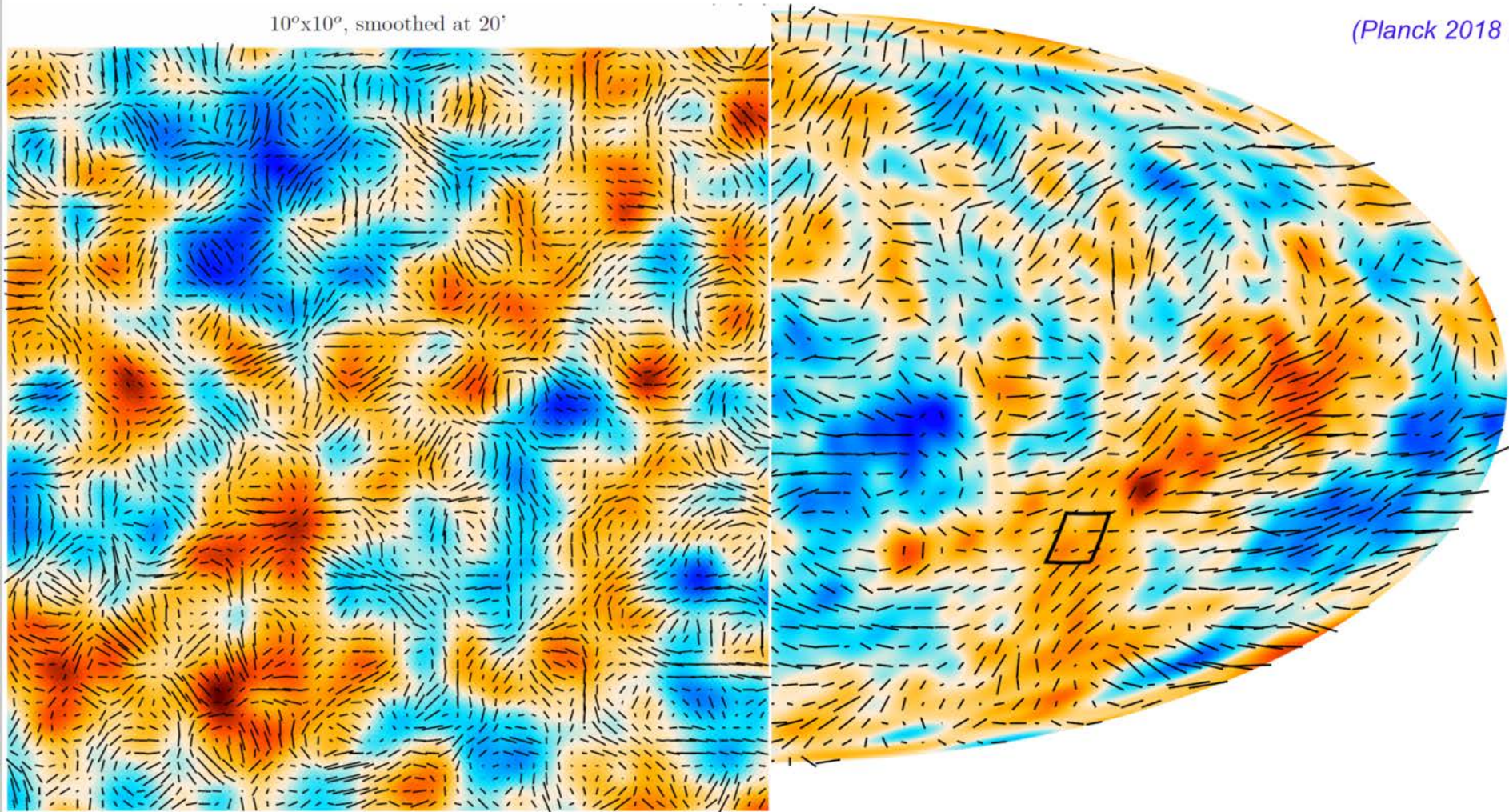
Temperature smoothed to 5 degrees



# PLANCK: POLARIZATION ANISOTROPIES

$10^\circ \times 10^\circ$ , smoothed at  $20'$

(Planck 2018 I)



$13.7 \mu\text{K}$

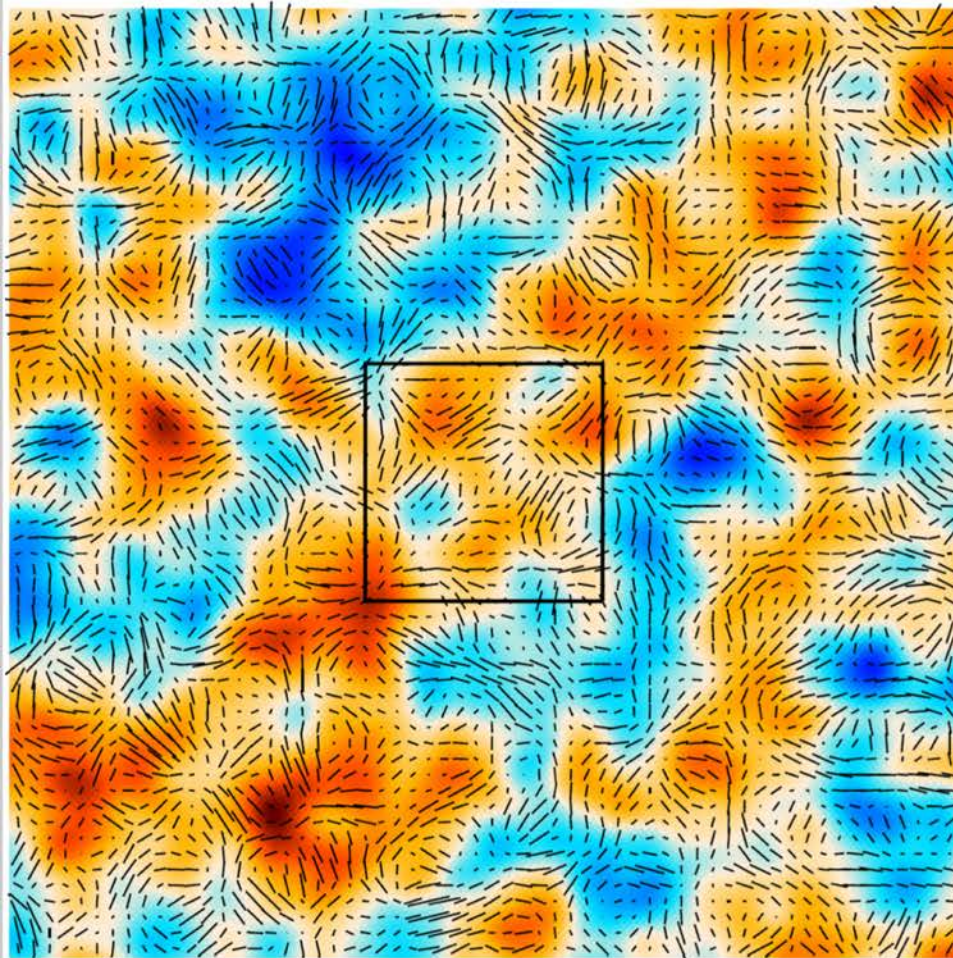
-201  309  $\mu\text{K}$

(276.4, -29.8) Galactic



# PLANCK: POLARIZATION ANISOTROPIES

$10^\circ \times 10^\circ$ , smoothed at  $20'$

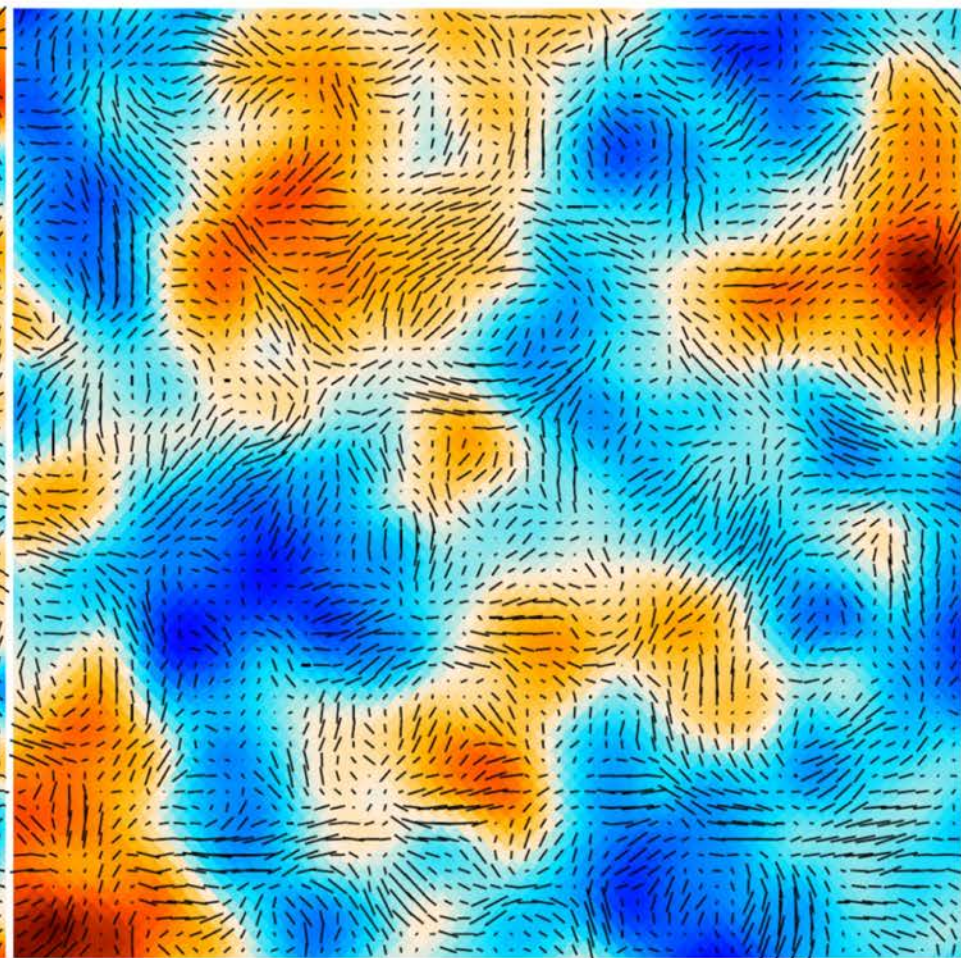


-201  309  $\mu\text{K}$

13.7  $\mu\text{K}$

(276.4, -29.8) Galactic

$2.5^\circ \times 2.5^\circ$ , smoothed at  $7'$



-67  311  $\mu\text{K}$

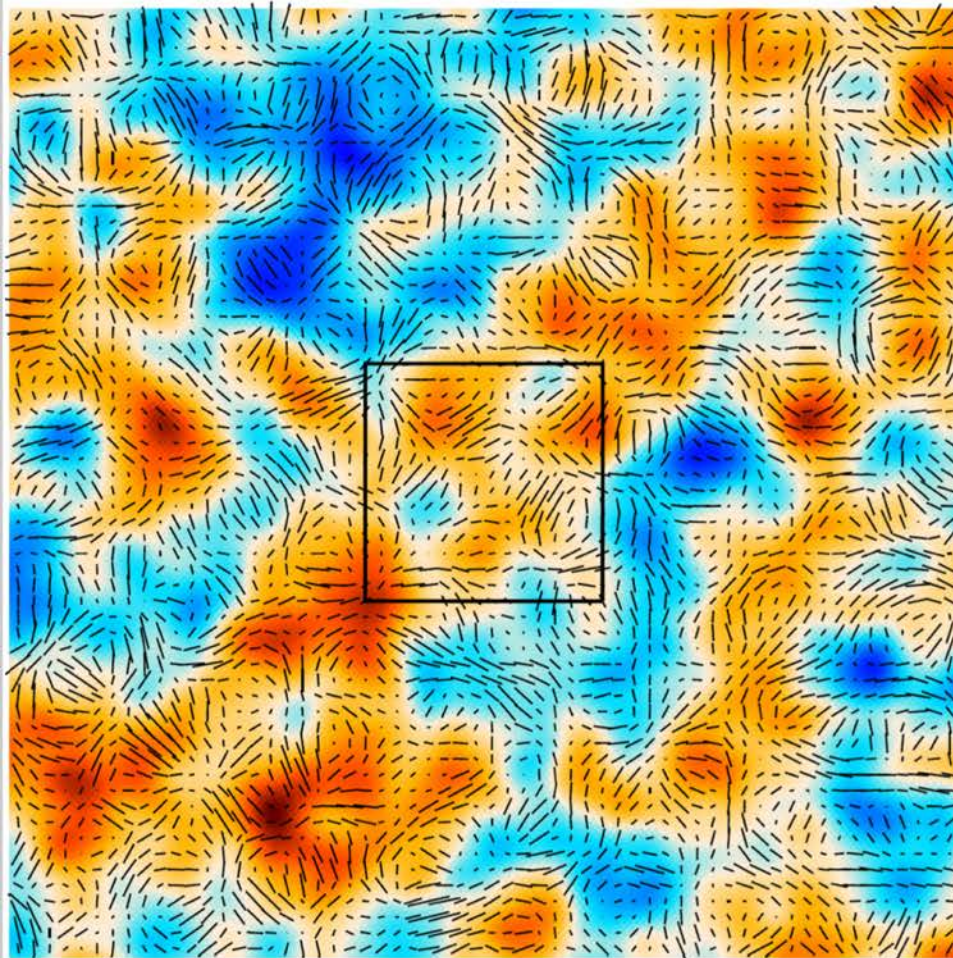
36.1  $\mu\text{K}$

(276.4, -29.8) Galactic



# PLANCK: POLARIZATION ANISOTROPIES

$10^\circ \times 10^\circ$ , smoothed at  $20'$

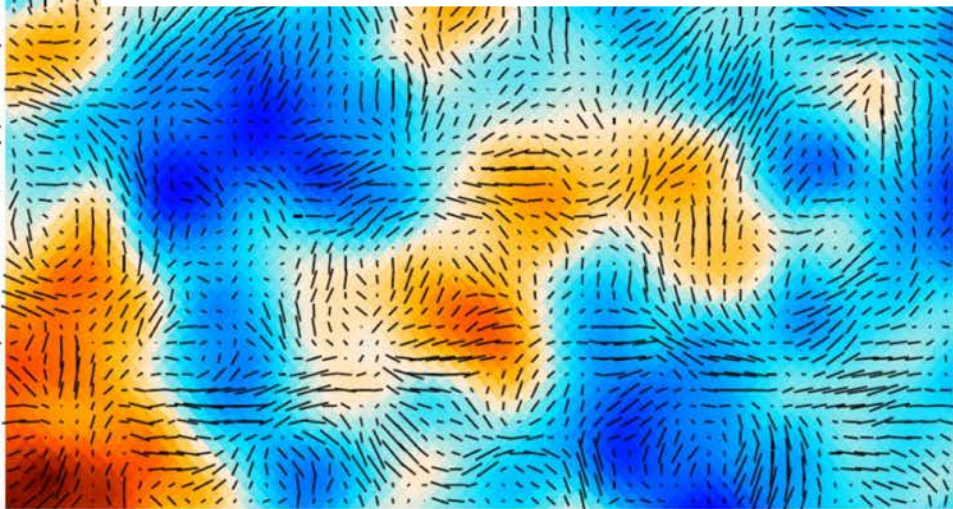
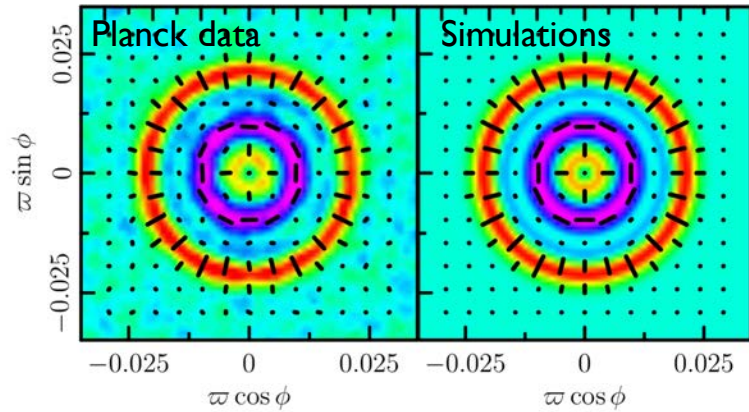
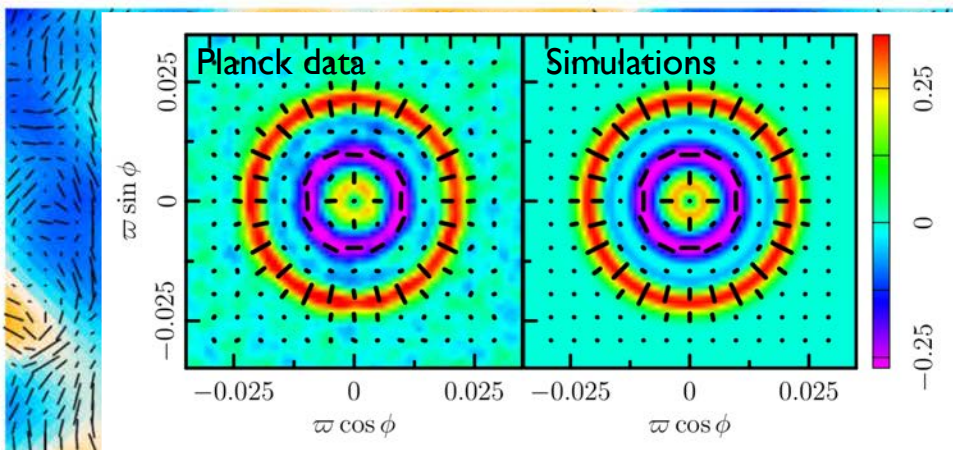


-201  309  $\mu\text{K}$

13.7  $\mu\text{K}$

(276.4, -29.8) Galactic

$2.5^\circ \times 2.5^\circ$ , smoothed at  $7'$



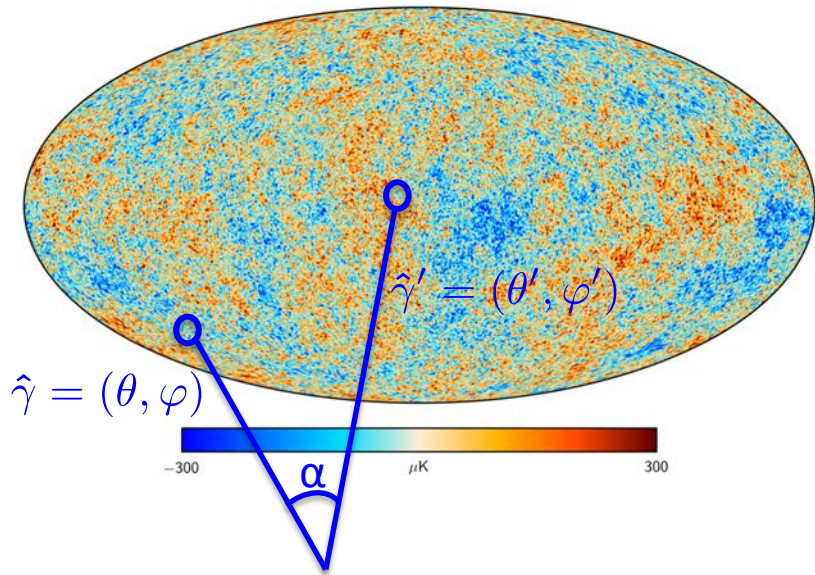
-67  311  $\mu\text{K}$

36.1  $\mu\text{K}$

(276.4, -29.8) Galactic



# STATISTICAL DESCRIPTION



## CORRELATION FUNCTIONS

$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \right\rangle \quad \leftarrow \text{from Inflation}$$

$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \frac{\Delta T}{T}(\vec{\gamma}'') \right\rangle$$

$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \frac{\Delta T}{T}(\vec{\gamma}'') \frac{\Delta T}{T}(\vec{\gamma}''') \right\rangle$$

...



E modes



B modes

## POLARIZATION

$$\mathbf{P}(\hat{\gamma}) = \nabla \mathbf{E} + \nabla \times \mathbf{B}$$

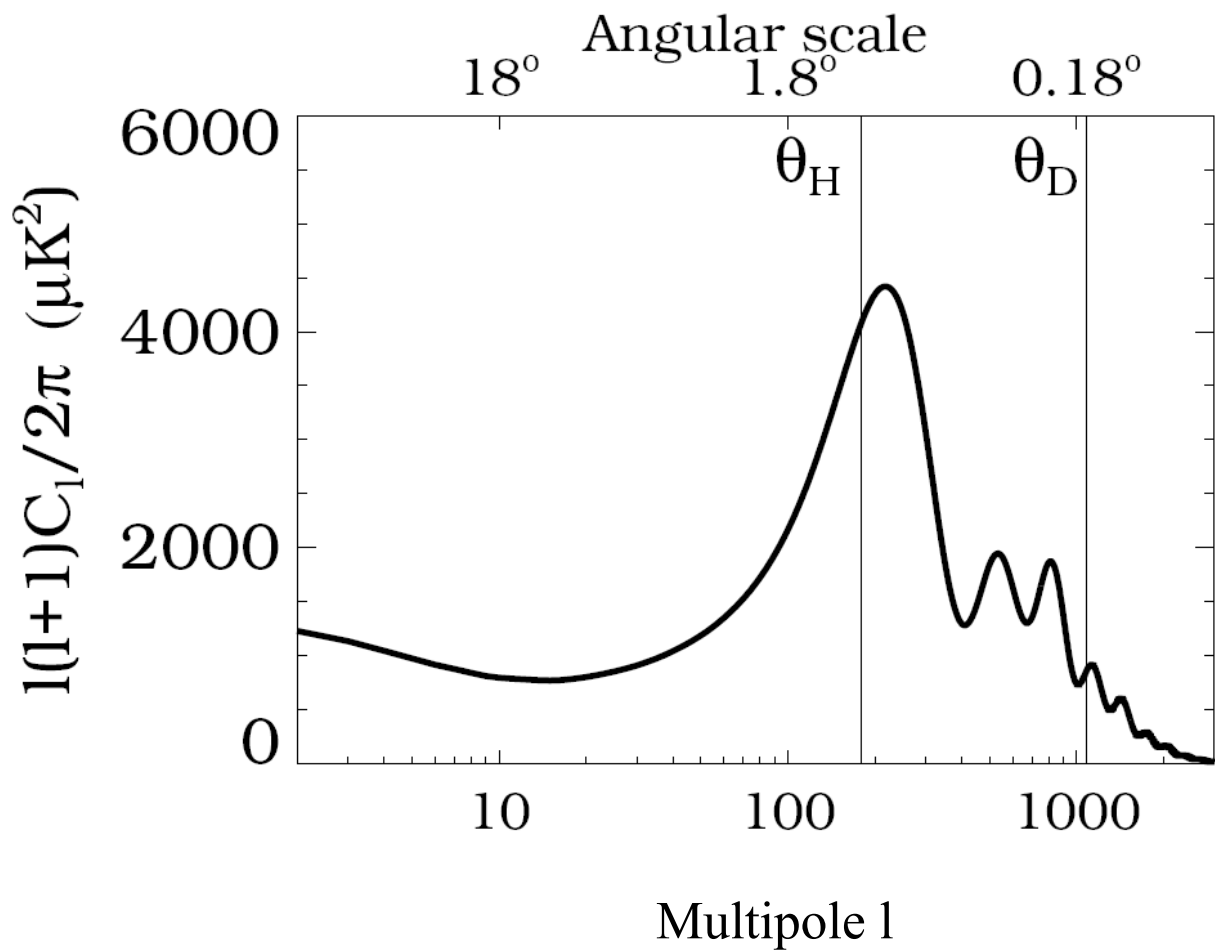
**E-modes:** even under parity

**B-modes:** odd under parity

Density perturbations -> E-modes

Gravitational Waves -> E- and B-modes



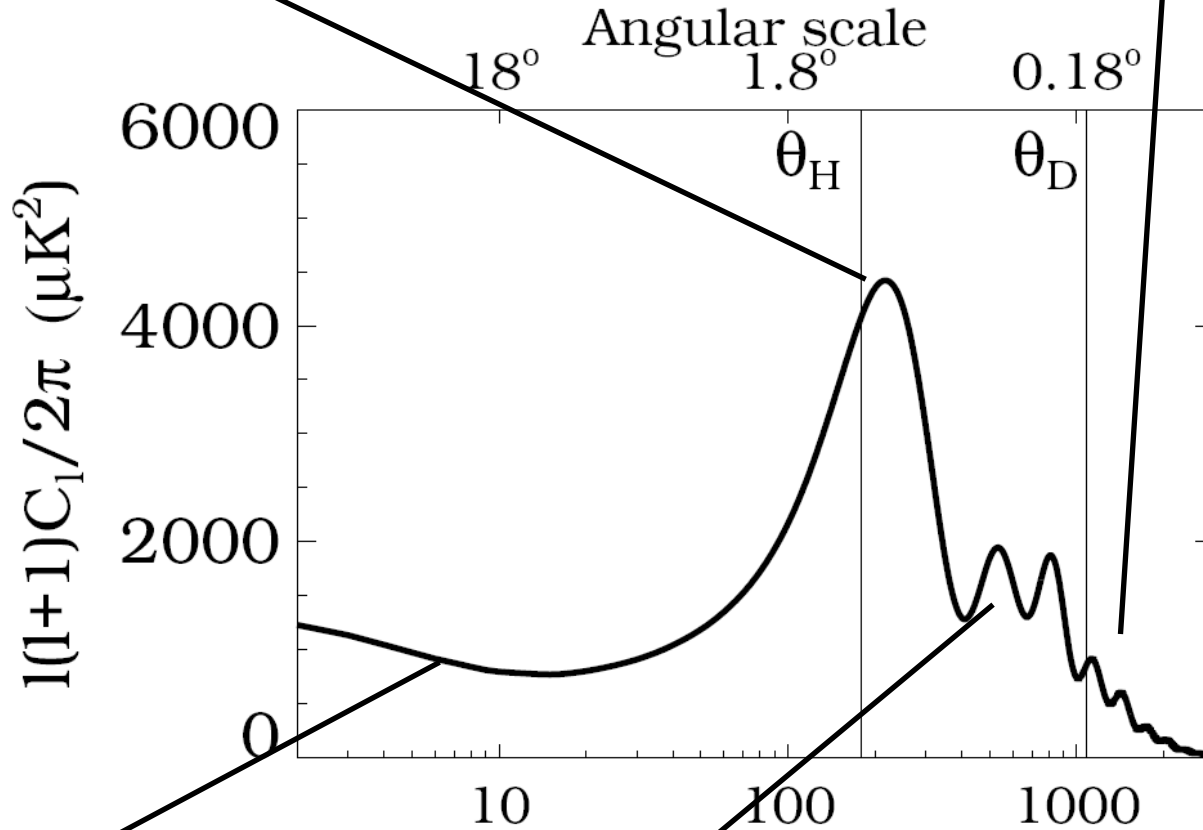


spatial curvature  
relative abundance of matter and radiation  
distance to the last scattering surface

$H_0, \Omega_m, \Omega_k$

Photon diffusion length at recombination  
Slope of the primordial spectrum

$N_{\text{eff}}, \Omega_b, Y_p, n_s$



+ Overall power  
 $A_s e^{-2\tau}$

+ low-ell  
polarization  
(not shown)  
Reionization  
history

$\tau$

Primordial power spectrum  
late time expansion

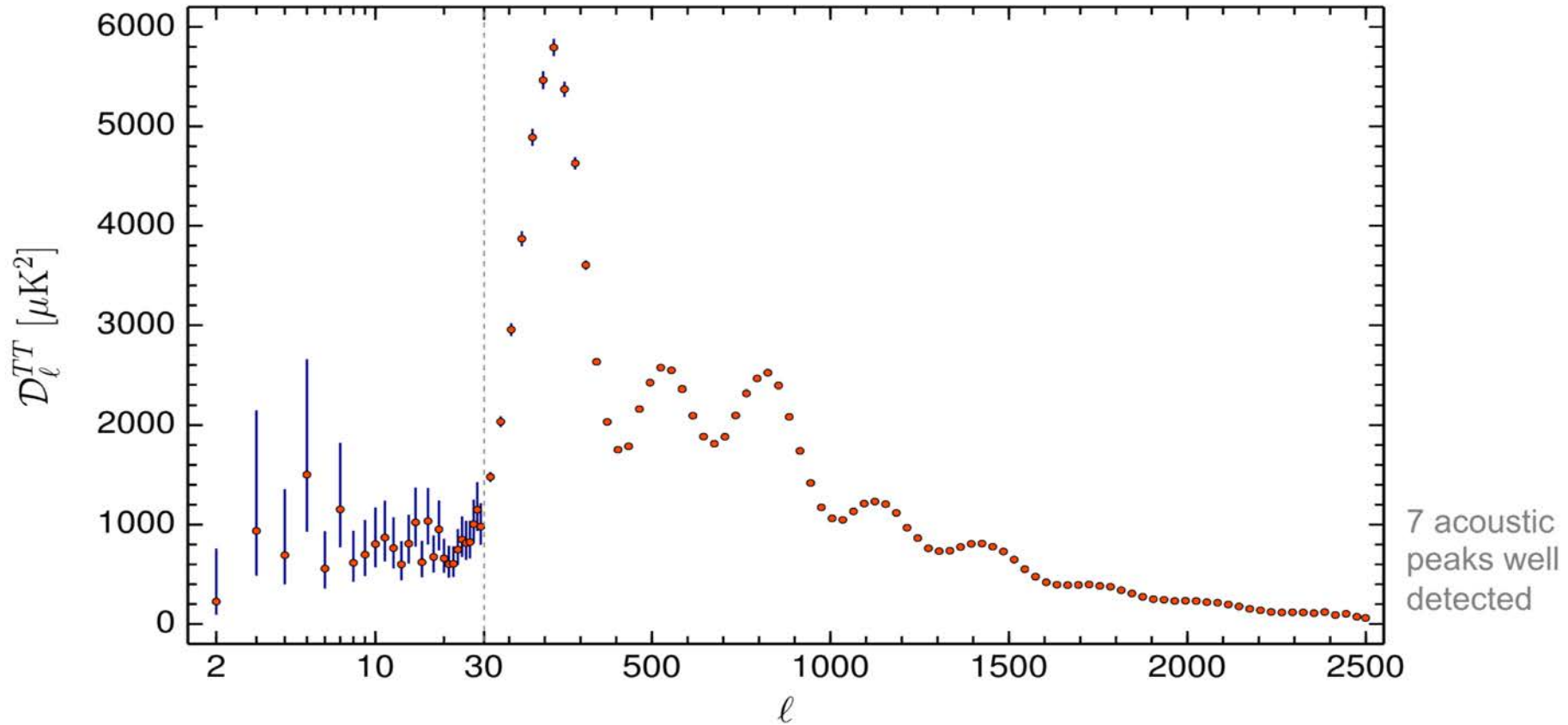
$A_s, \Omega_\Lambda$

Baryon abundance

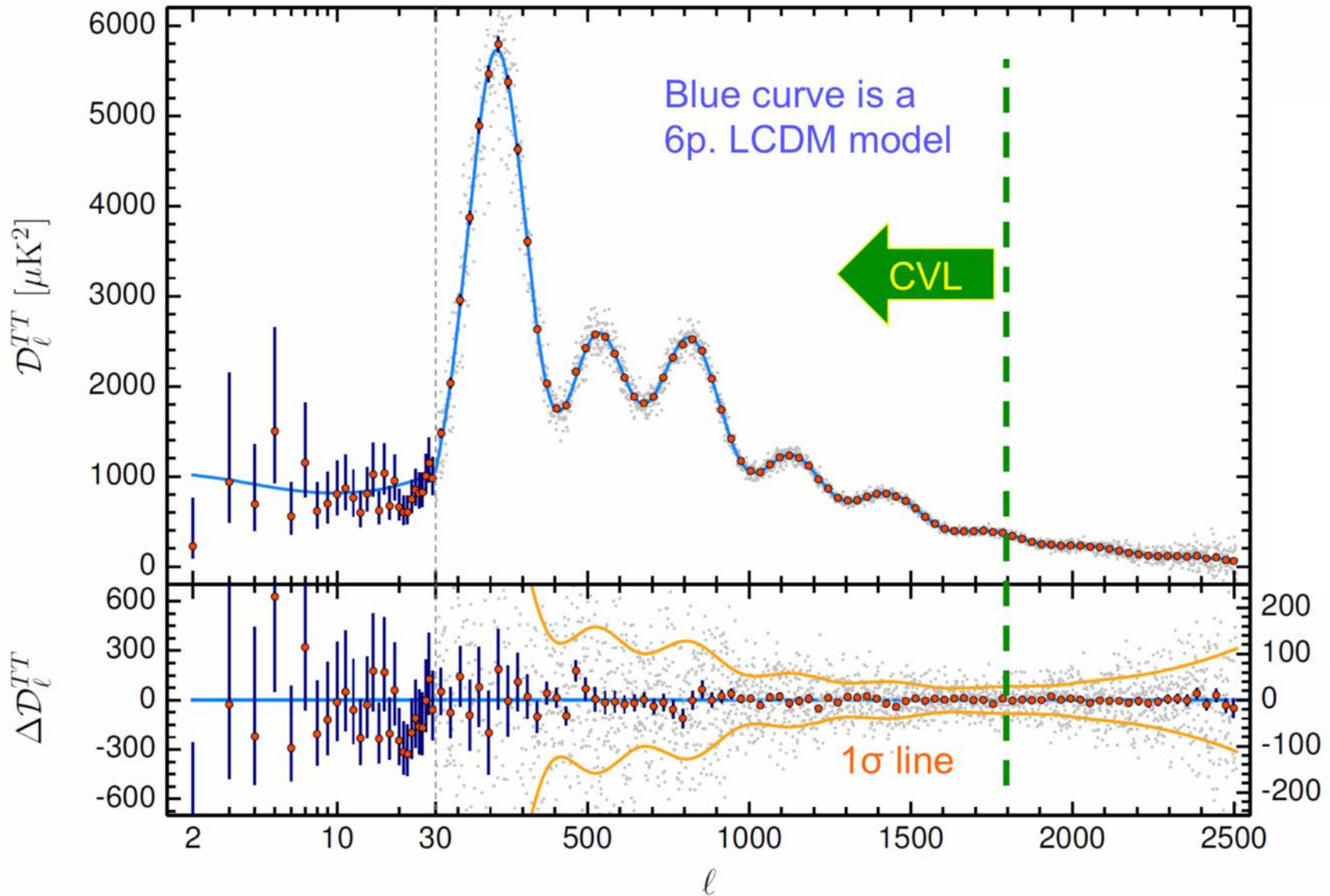
$\Omega_b$



# Planck 2018 TT power spectrum

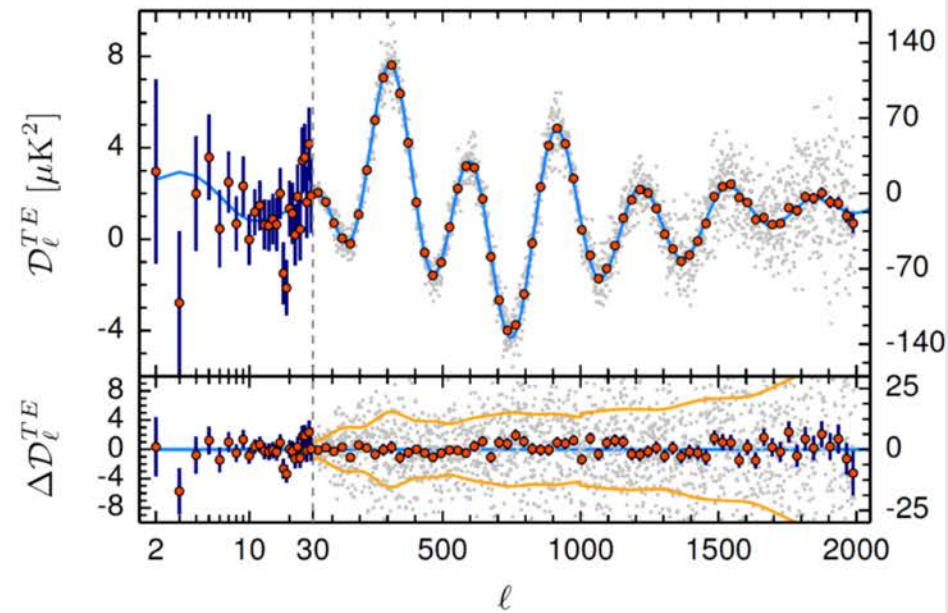
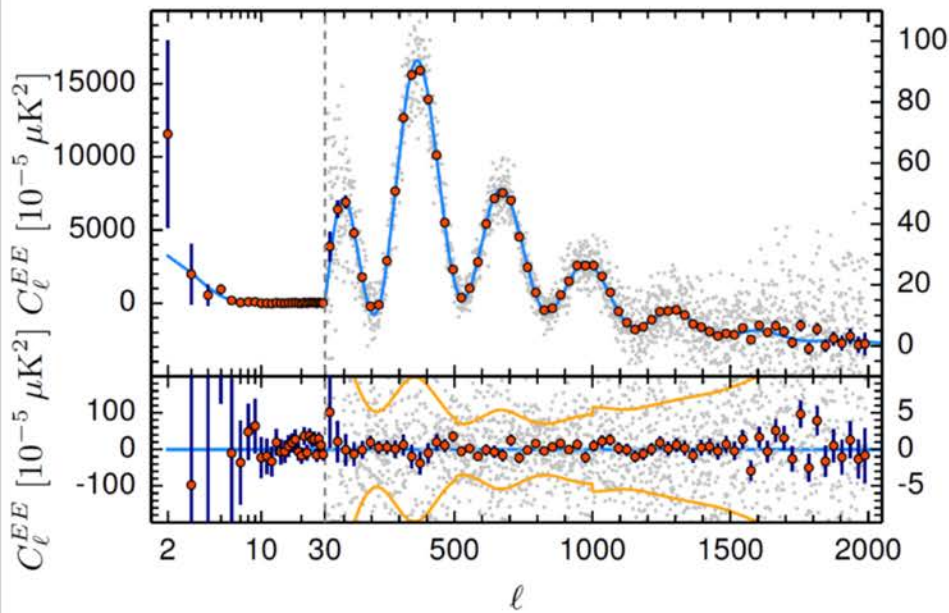


# Planck 2018 TT power spectrum



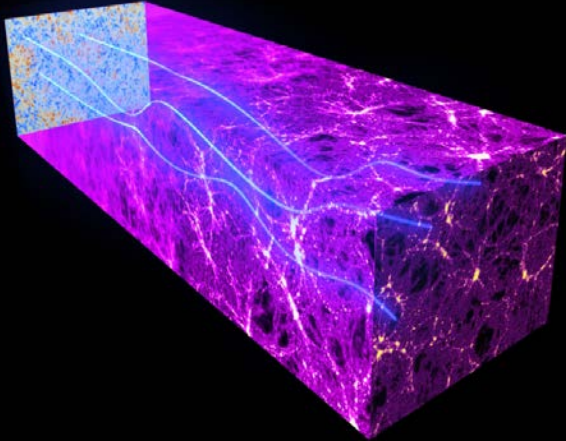


# Planck 2018 TE, EE power spectra



Blue line is not a fit, but a prediction given the TT spectrum!

CMB is sensitive to the late-time density field, too....

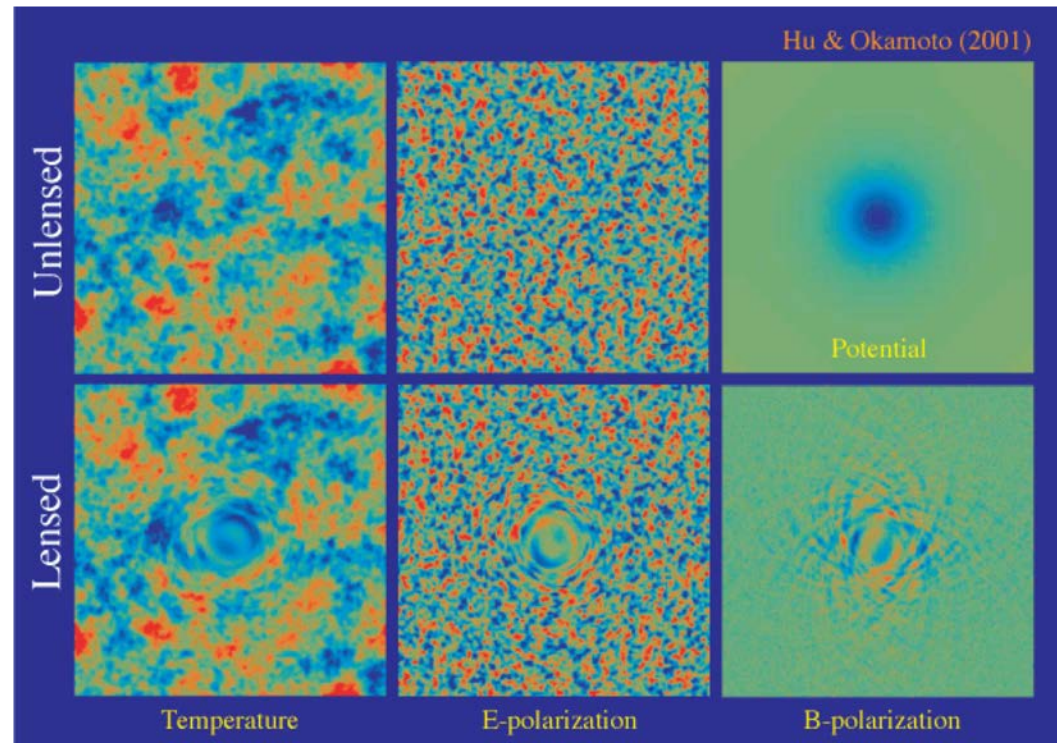


Line-of-sight integral of the gravitational potentials

$$\phi(\hat{n}) = - \int_0^{\chi_*} d\chi \frac{\chi_* - \chi}{\chi_* \chi} (\Phi + \Psi)$$

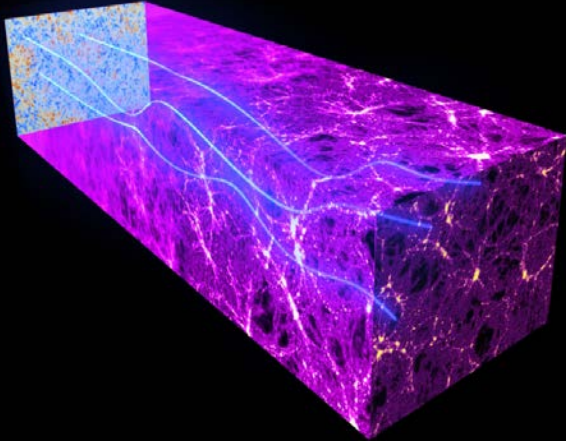
Deflection field

$$\vec{d} = \vec{\nabla} \phi$$





CMB is sensitive to the late-time density field, too....



Deflection field

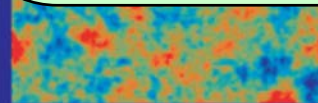
$$\vec{d} = \vec{\nabla} \phi$$

Line-of-sight integral of the gravitational potentials

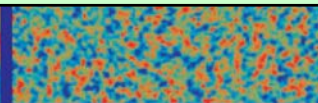
$$\phi(\hat{n}) = - \int_0^{\chi_*} d\chi \frac{\chi_* - \chi}{\chi_* \chi} (\Phi + \Psi)$$

Measures deflection of light due to intervening structures  
(average deflection angle is  $\sim 2.5$  arcmin)

Gives integrated information about the matter distribution between us and the last scattering surface



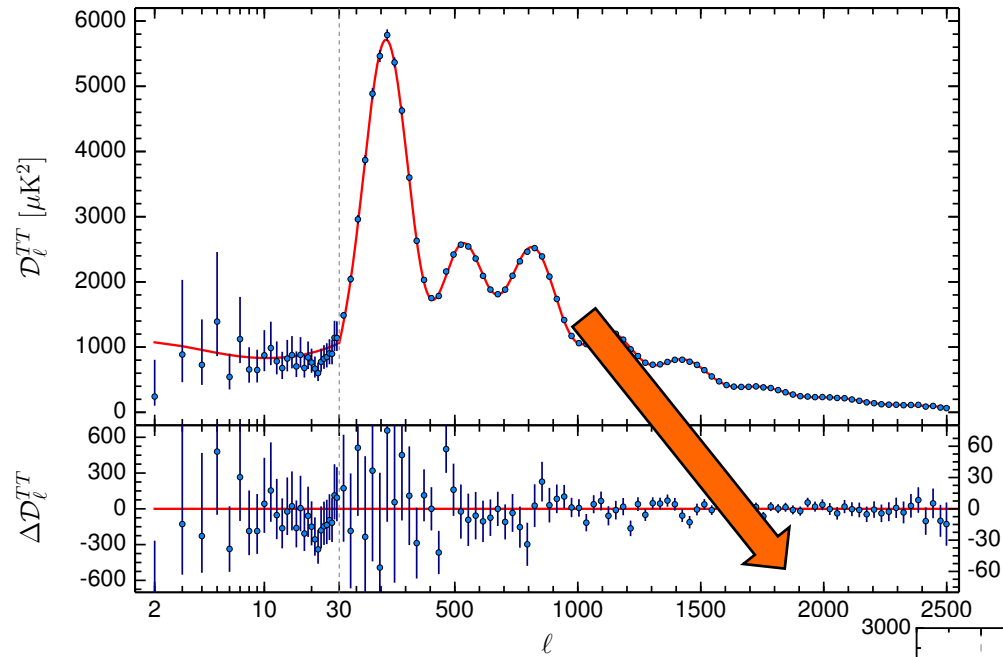
Temperature



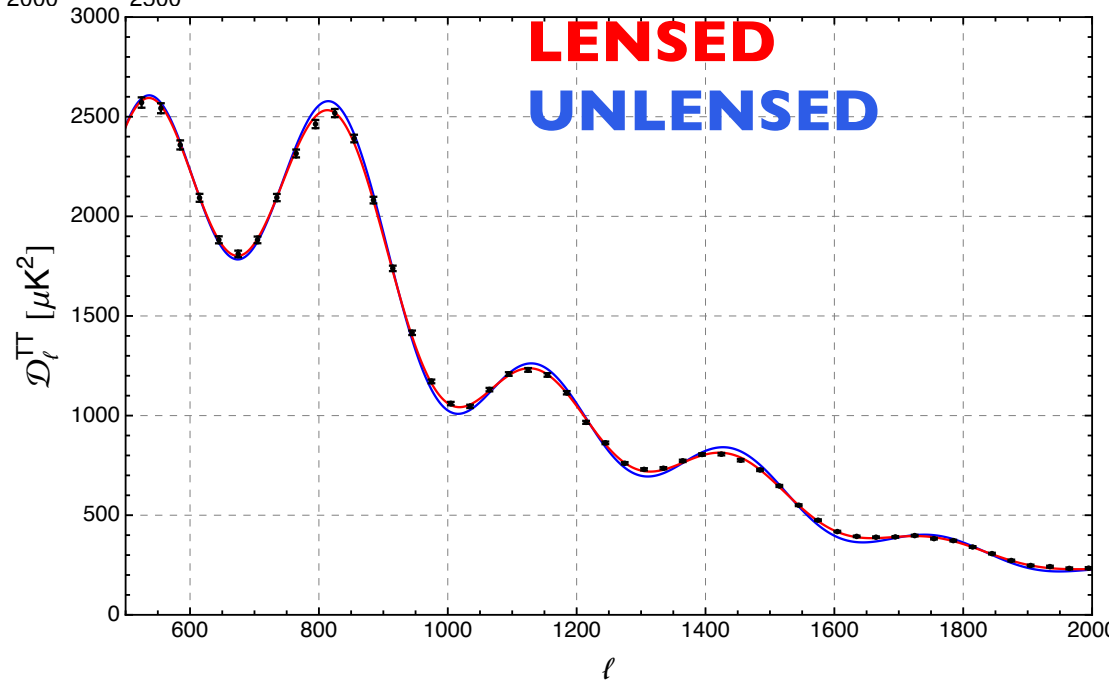
E-polarization



B-polarization

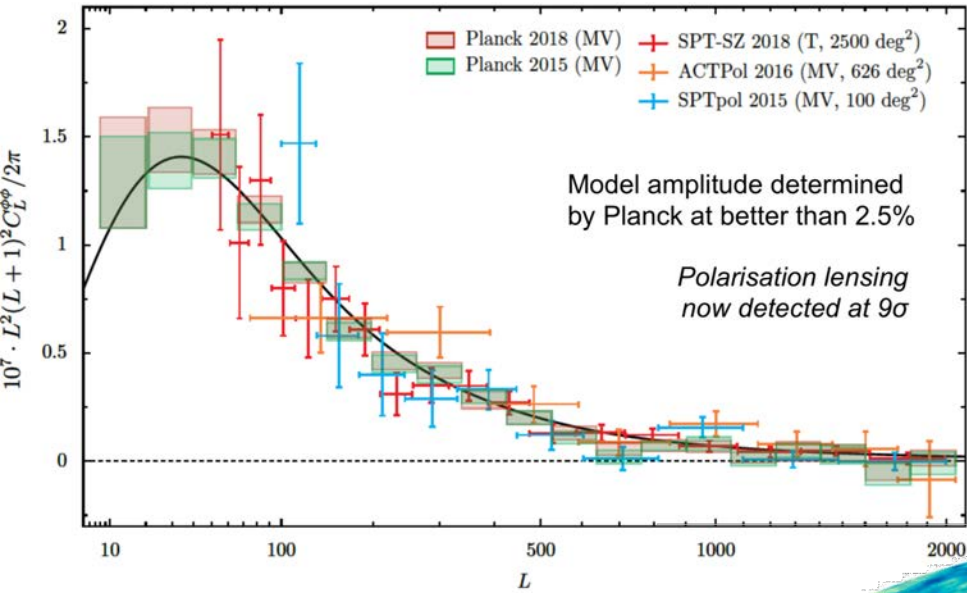


Lensing smooths the peaks  
of the CMB power  
spectrum...  
... and introduces non-  
gaussianities in the map  
(nonzero 4-point c.f.)

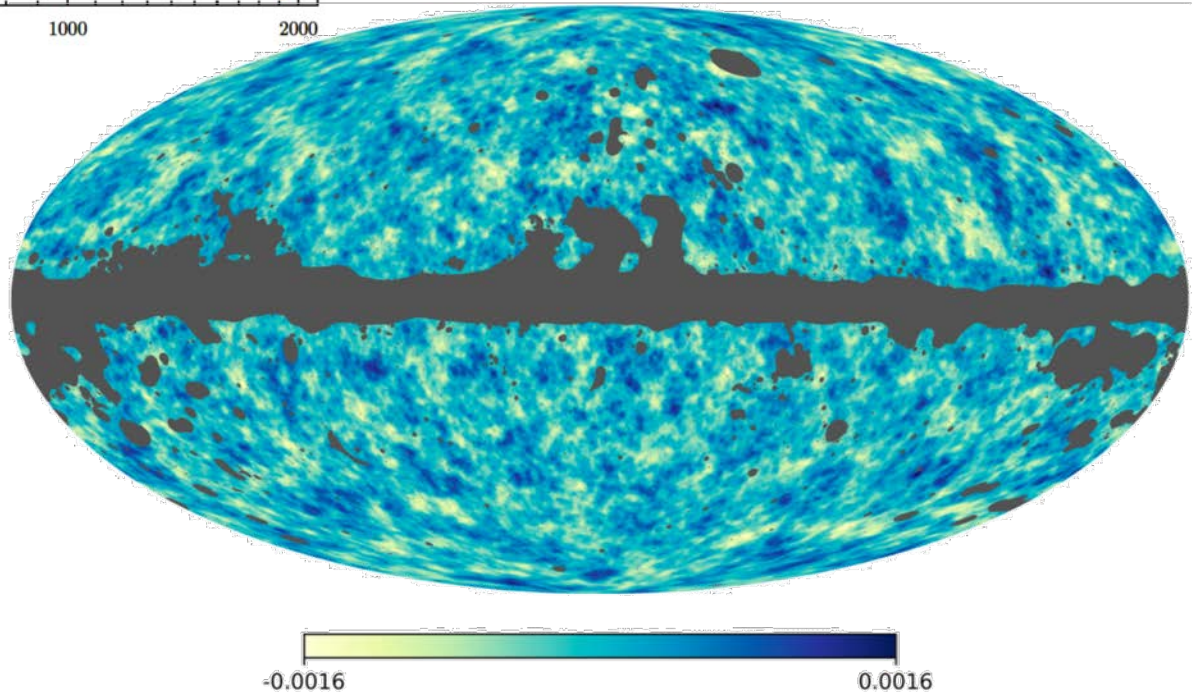




# LENSING



Lensing potential estimated from the four-point correlation function



# $\Lambda$ CDM 6 parameter fit

(Planck temperature, polarization and lensing)

		Mean	Stdev	Rel. err.
primary	$\Omega_b h^2$ Baryon density	0.02237	0.00015	0.007
	$\Omega_c h^2$ Dark matter density	0.1200	0.0012	0.01
	$100\theta$ CMB acoustic scale	1.04092	0.00031	0.0003
	$\tau$ Optical depth to last scattering surface	0.0544	0.0073	0.13
	$\ln(A_s 10^{10})$ Primordial amplitude of perturbation	3.044	0.014	0.007
	$n_s$ Primordial Scalar spectral index	0.9649	0.0042	0.004
derived	$H_0$ Hubble parameter today	67.36	0.54	0.008
	$\Omega_m$ Total matter density	0.3153	0.0073	0.023
	$\sigma_8$ Matter perturbation amplitude	0.8111	0.0060	0.007



# $\Lambda$ CDM 6 parameter fit

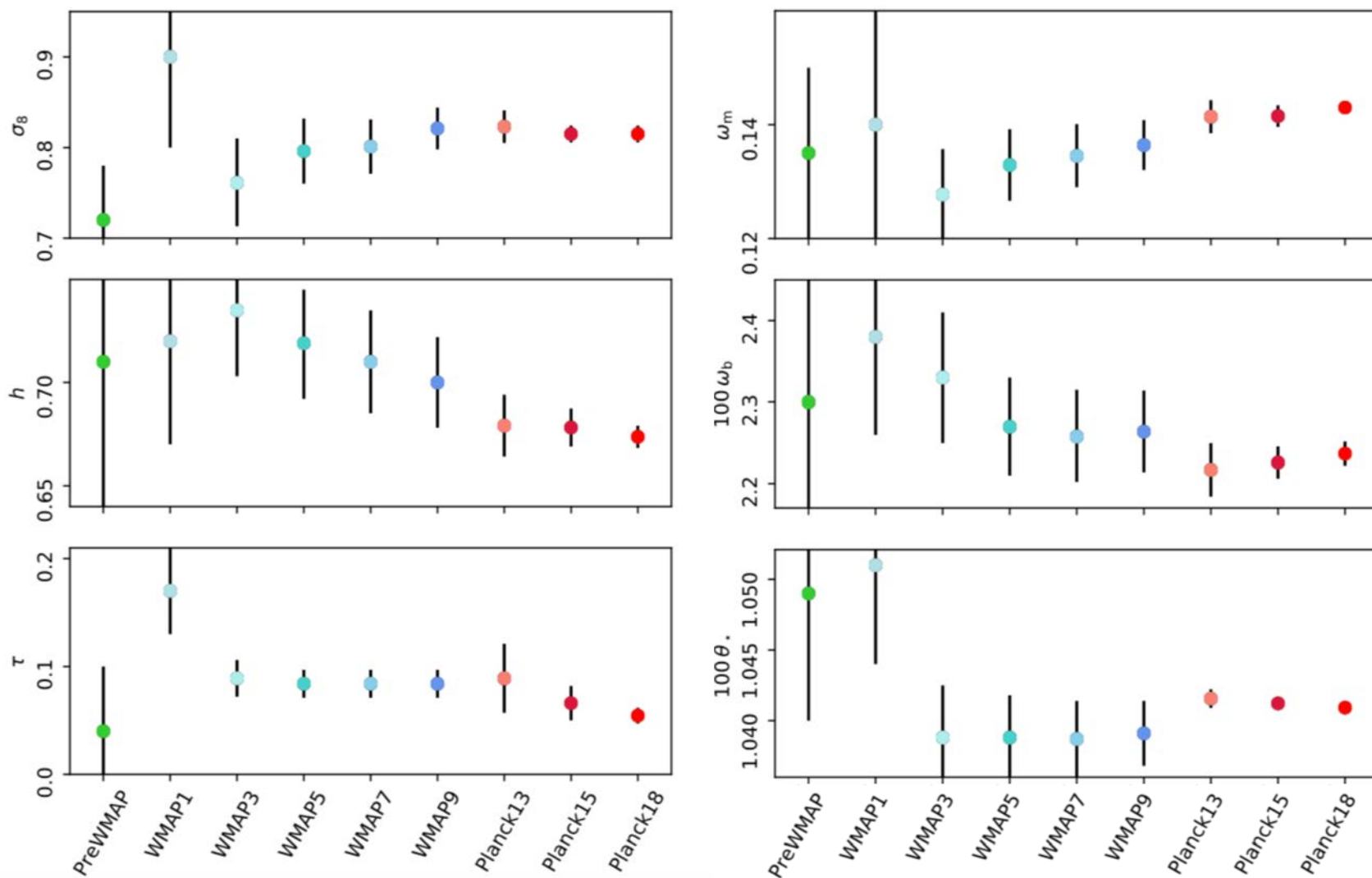
## (Planck temperature, polarization and lensing)

### Highlights:

1. Best determination of  $H_0$  to date (indirect, in strong tension with direct measurements)
2. Scalar spectral index is now 8  $\sigma$  away from 1 (a signature of inflation). Even in extended
3. Optical depth  $\tau$  greatly improved after taming of large-angle polarization systematics. Still, at 13% relative error, by far the worst parameter determined from CMB

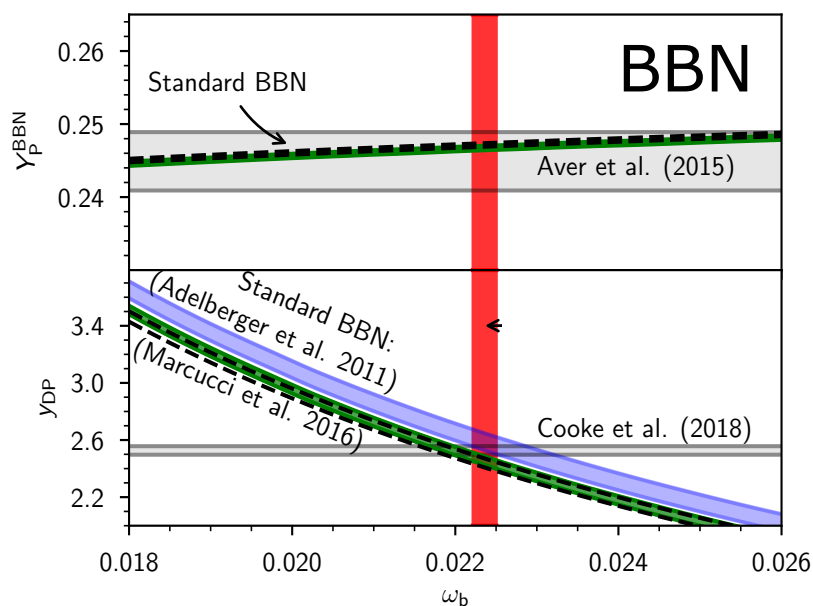
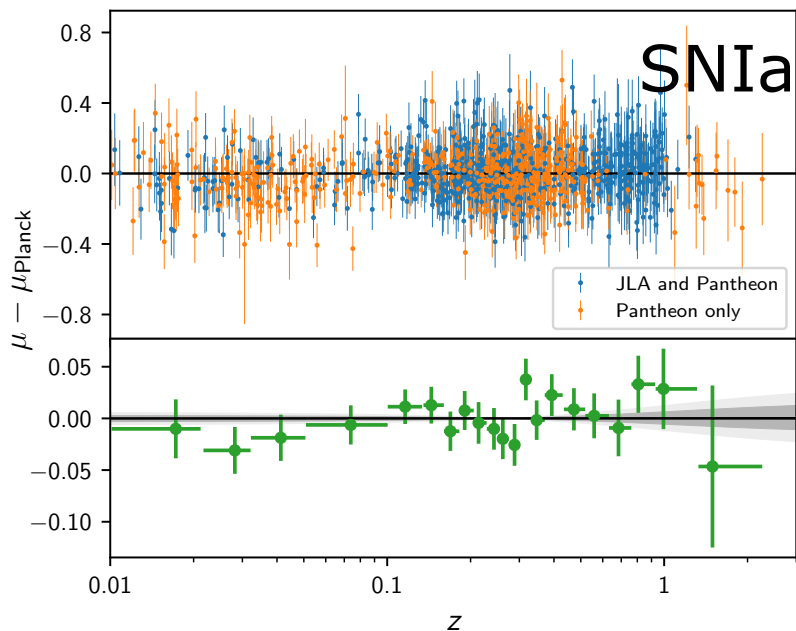
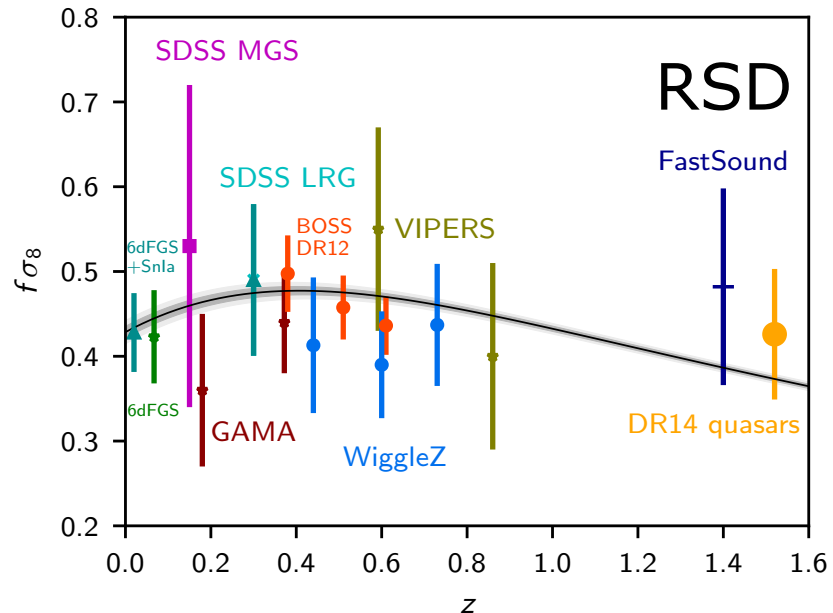
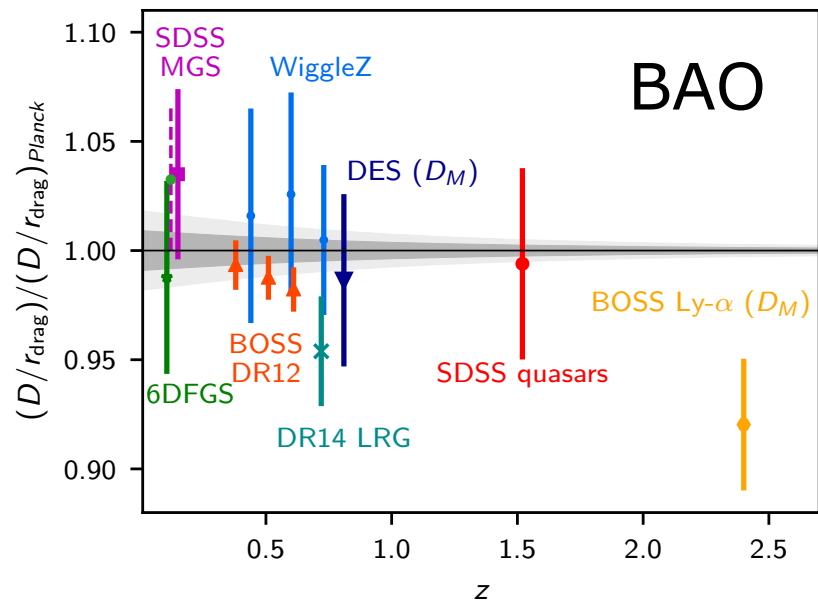
	Mean	Stdev	Rel. err.
$\Omega_b h^2$ Baryon density	0.02237	0.00015	0.007
$\Omega_c h^2$ Dark matter density	0.1200	0.0012	0.01
$100\theta$ CMB acoustic scale	1.04092	0.00031	0.0003
$\tau$ Optical depth to reionization	0.0544	0.0073	0.13
$\ln(A_s 10^{10})$ Primordial amplitude of perturbation	3.044	0.014	0.007
$n_s$ Primordial Scalar spectral index	0.9649	0.0042	0.004
$H_0$ Hubble parameter today	67.36	0.54	0.008
$\Omega_m$ Total matter density	0.3153	0.0073	0.023
$\sigma_8$ Matter perturbation amplitude	0.8111	0.0060	0.007

# Improvement in parameter accuracy

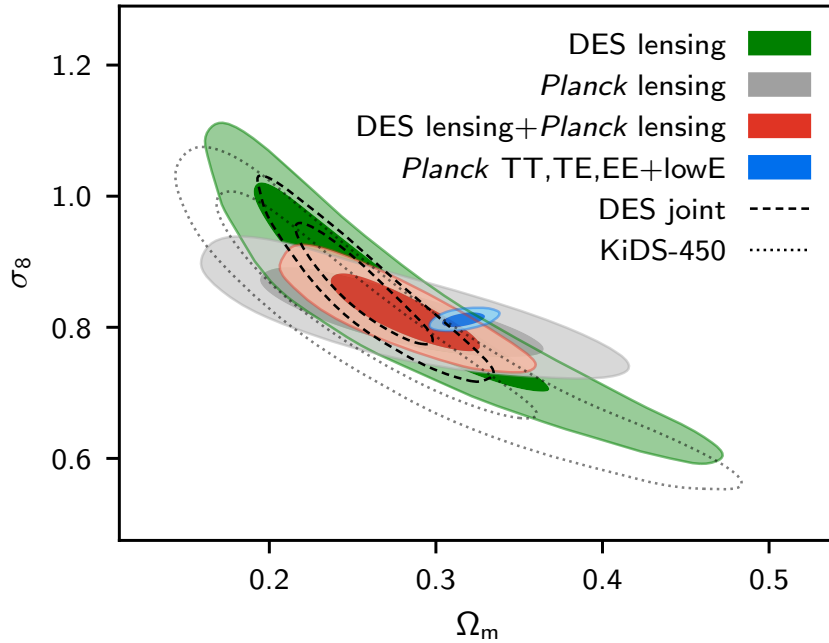




# Consistency with other datasets



# Tensions with other datasets...



## Mild tension with DES year I results

$$S_8 \equiv \sigma_8 (\Omega_m / 0.3)^{0.5}$$

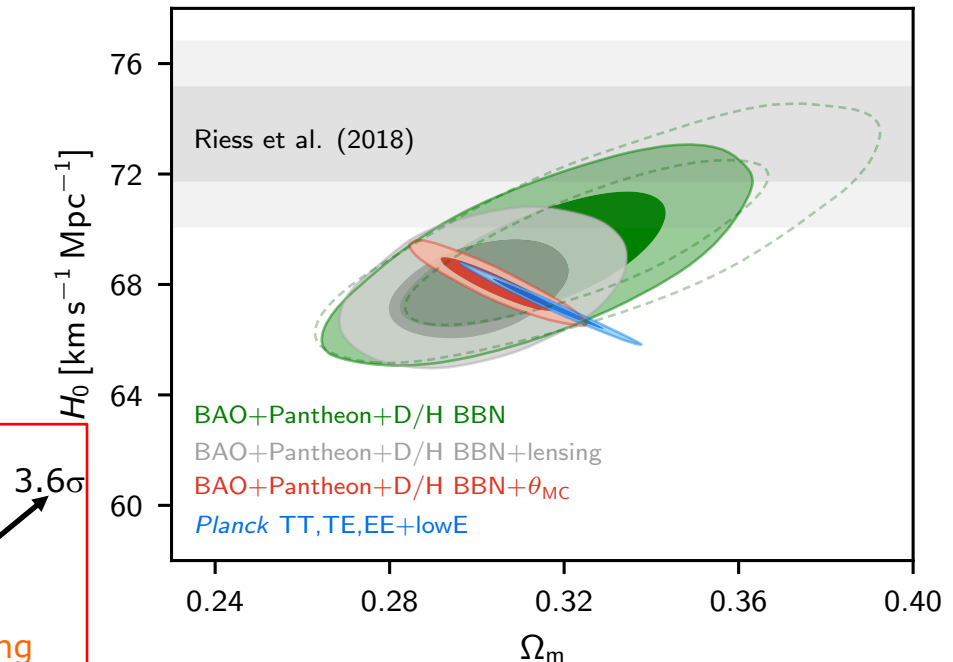
**DES Joint**  
 $S_8 = 0.792 \pm 0.024$   
 $\Omega_m = 0.257^{+0.023}_{-0.031}$

**Planck TTTEE+lowE+lensing**  
 $S_8 = 0.832 \pm 0.013$   
 $\Omega_m = 0.315 \pm 0.007$

Strong tension with  $H_0$  distance ladder measurements.

$H_0 = 67.36 \pm 0.54$  km/s/Mpc Planck  $\Lambda$ CDM  
 $H_0 = 73.5 \pm 1.6$  km/s/Mpc SH0ES (Riess+ 18)

Inverse distance ladder:  
 $H_0 = 67.9 \pm 1.3$  km/s/Mpc BAO+D/H+CMB lensing



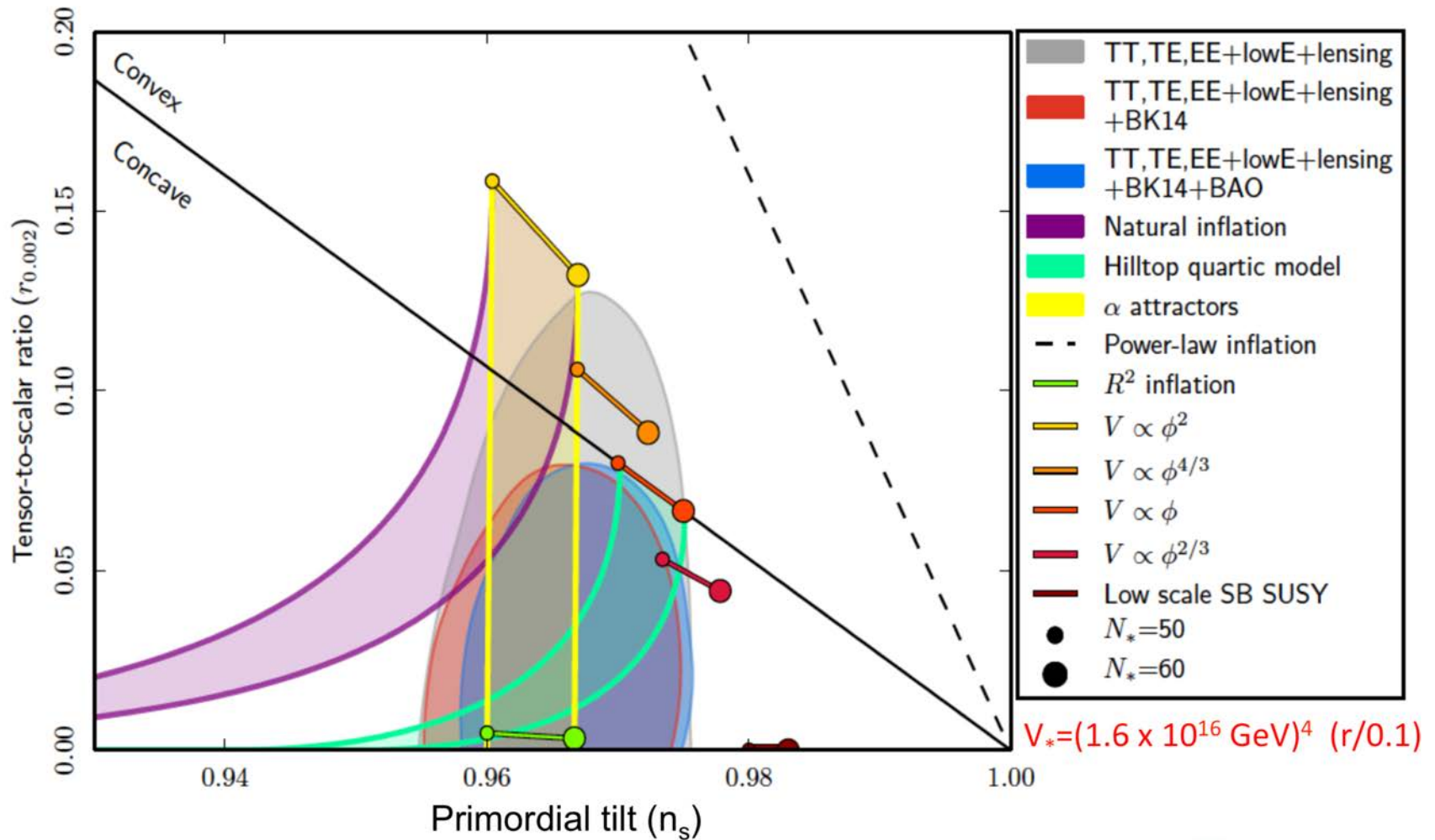
Now > 4 $\sigma$ , next talk by Martina Gerbino



# $\Lambda$ CDM 6 parameter fit + extensions (where surprises might hide)

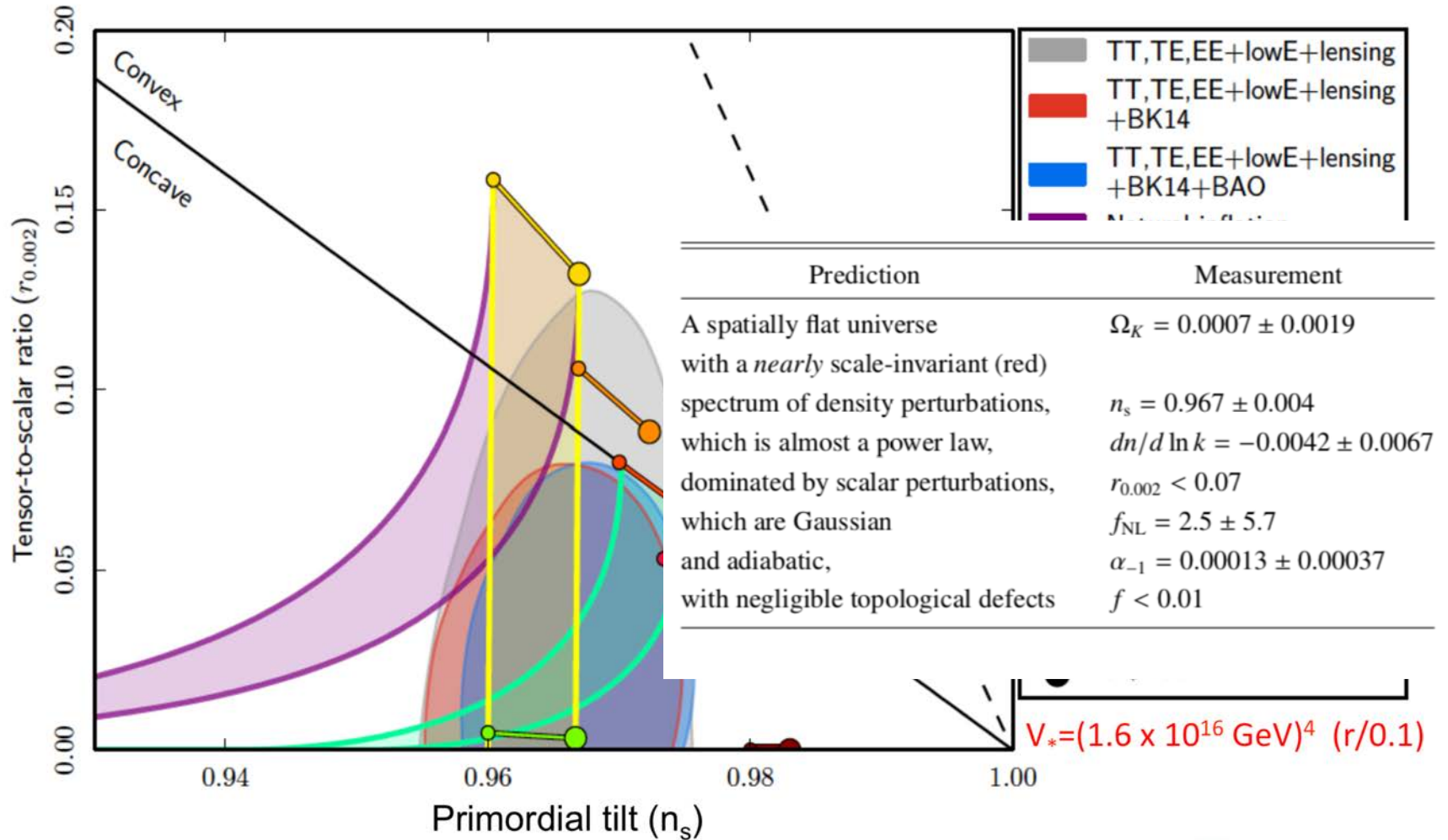
- Tensor modes, i.e. primordial gravitational waves,  $r = A_T/A_s$
- Running spectral index  $dn_s/d\ln k$
- Primordial non Gaussianity  $f_{NL}$
- Non adiabatic (isocurvature) primordial perturbations
- Dark energy equation of state,  $w$
- Spatial curvature  $\Omega_k = 1 - \Omega_m - \Omega_\Lambda$
- Neutrino masses  $\Sigma m_\nu$
- Number of relativistic species  $N_{eff}$
- ...

# Constraints for tensor perturbations

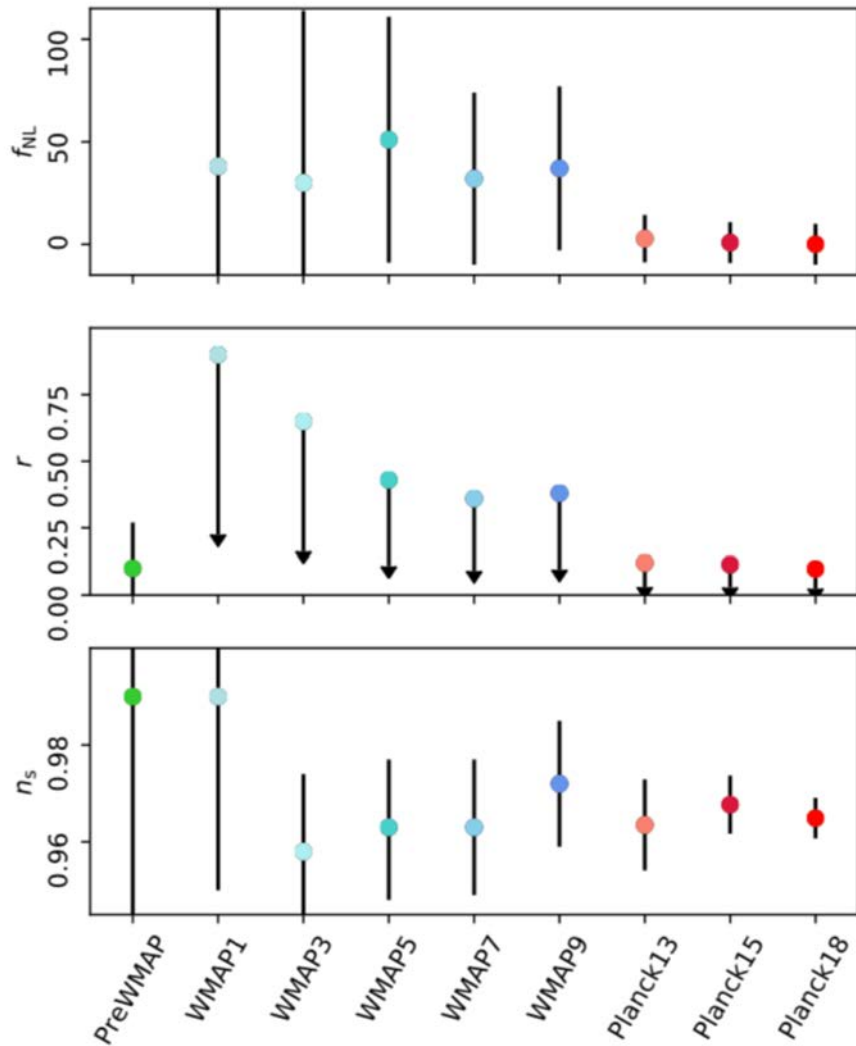




# Constraints for tensor perturbations

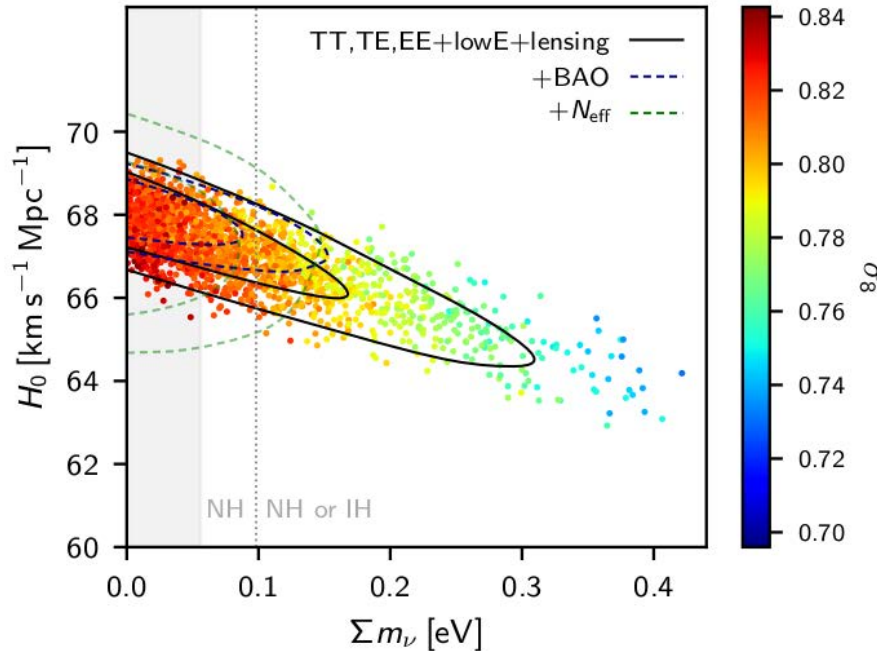


# Improvement in inflationary parameters





# Neutrino legacy of Planck: $\Sigma m_\nu$

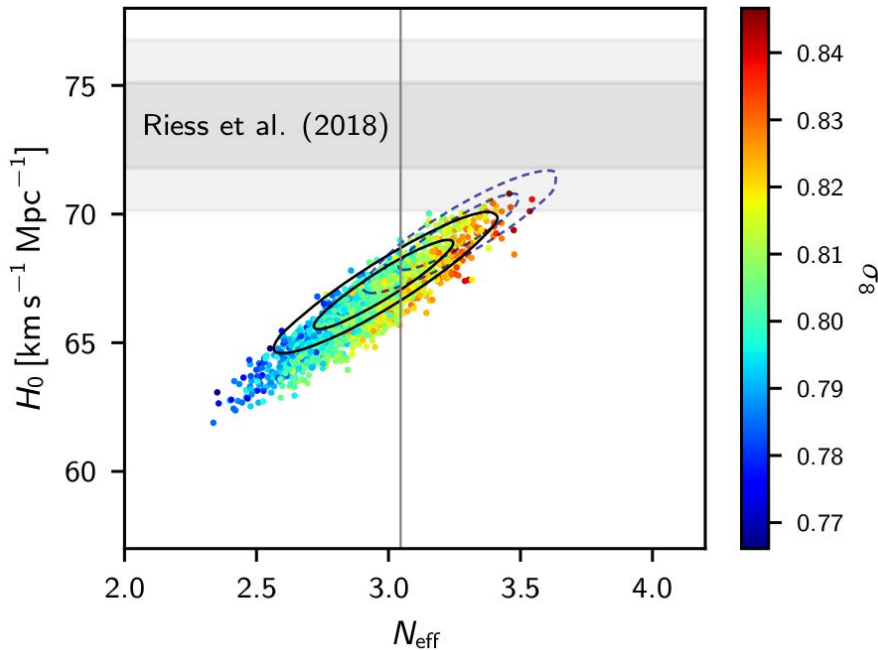


- Tightest constraint from a single experiment
- First constraint exploiting the information encoded in the CMB weak lensing
- One order of magnitude better than present kinematic constraints, already at the same level than future expectations for KATRIN
- The combined limits from Planck and large scale structure probes are starting to corner the inverted hierarchy scenario

$m_\nu < 0.44 \text{ eV}$  (95%CL, TT + lowE + lensing)

$m_\nu < 0.13 \text{ eV}$  (95% CL, TT+lowE+lensing+BAO)

# Neutrino legacy of Planck: $N_{\text{eff}}$



$$N_{\text{eff}} = 3.00^{+0.57}_{-0.53} \quad (95\% \text{ CL, TT+lowE})$$

$$N_{\text{eff}} = 3.11^{+0.44}_{-0.43} \quad (95\% \text{ CL, TT+lowE+lensing+BAO})$$

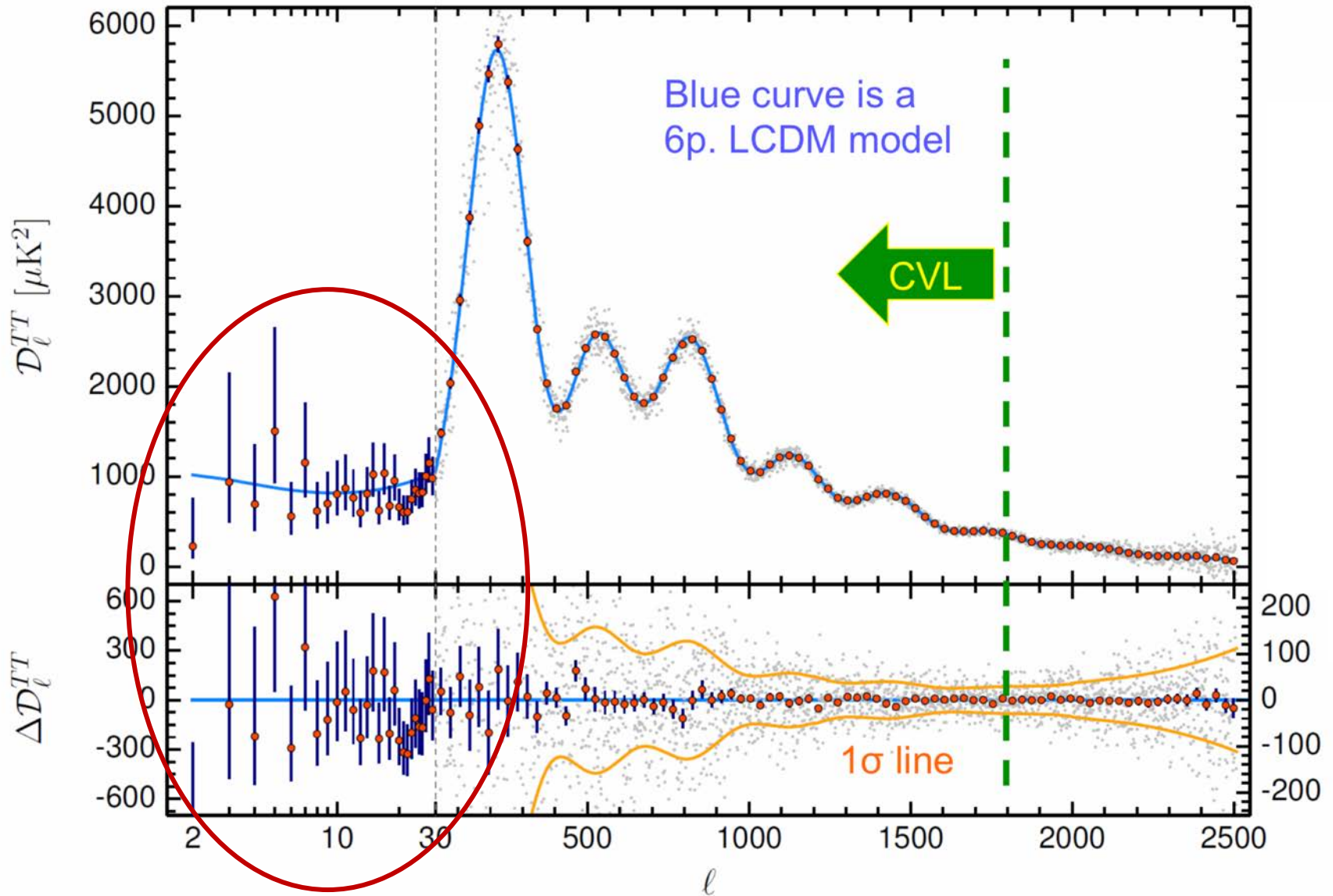
- Effective number of relativistic species is consistent with the standard expectation  $N_{\text{eff}} = 3.046$
- Data are consistent with these relativistic species behaving as free-streaming neutrinos – a strong indication that they are indeed the SM neutrinos!
- A fourth thermalized species ( $N_{\text{eff}}=4$ ) is excluded at 3.5 to 6  $\sigma$ , depending on the dataset
- A light sterile neutrino species is allowed if not thermalized. Still, the sterile neutrino interpretation of the short-baseline anomalies is excluded by Planck



# Anomalies in the CMB field

- At large angles, the CMB field is known to exhibit anomalies:
  - Lack of power
  - Hemispherical asymmetry
  - Even-odd asymmetry
  - And others...
- For temperature, Planck has reached cosmic variance. For polarization, there is much room for improvement.

# Planck 2018 TT power spectrum



Scale-invariance of the large-scale perturbations is a prediction of single-field, slow-roll inflation.

Transition from a pre-inflationary “fast-roll” phase to slow-roll would suppress power in the primordial spectrum.

Are we seeing relics of a decelerating inflaton?

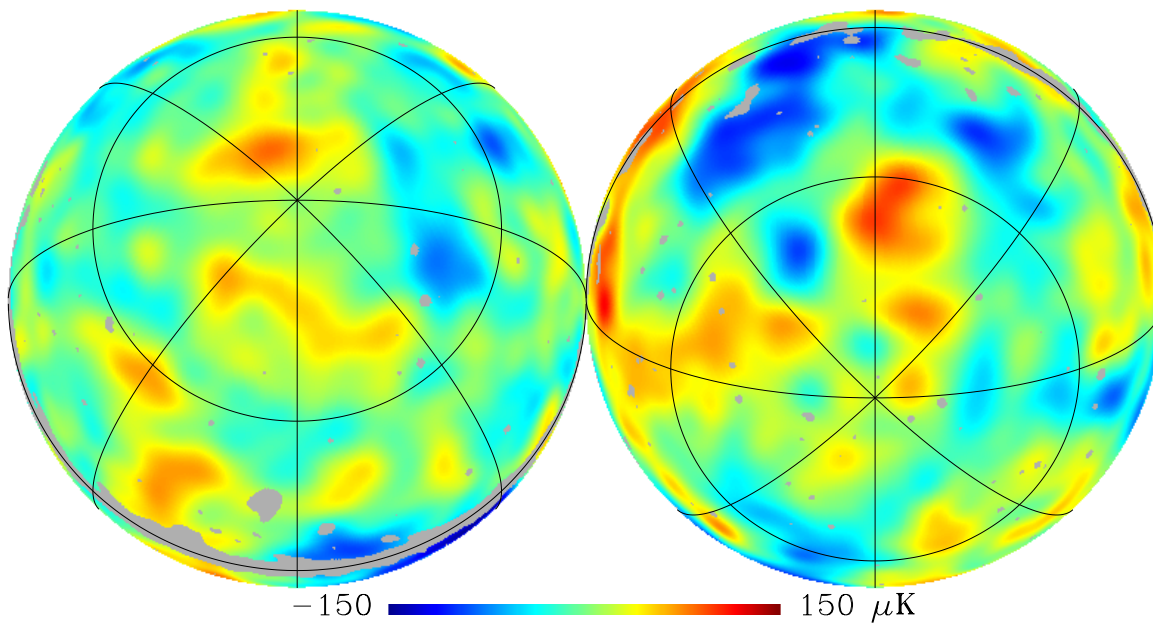
See e.g. Contaldi, Peloso, Kofman, Linde (2003); Destri, de Vega, Sanchez (2010); Dudas, Kitazawa, Patil, Sagnotti (2012); Kitazawa, Sagnotti (2014)

$$P(k) \sim \frac{k^3}{[k^2 + \Delta^2]^{2 - \frac{n_s}{2}}}$$

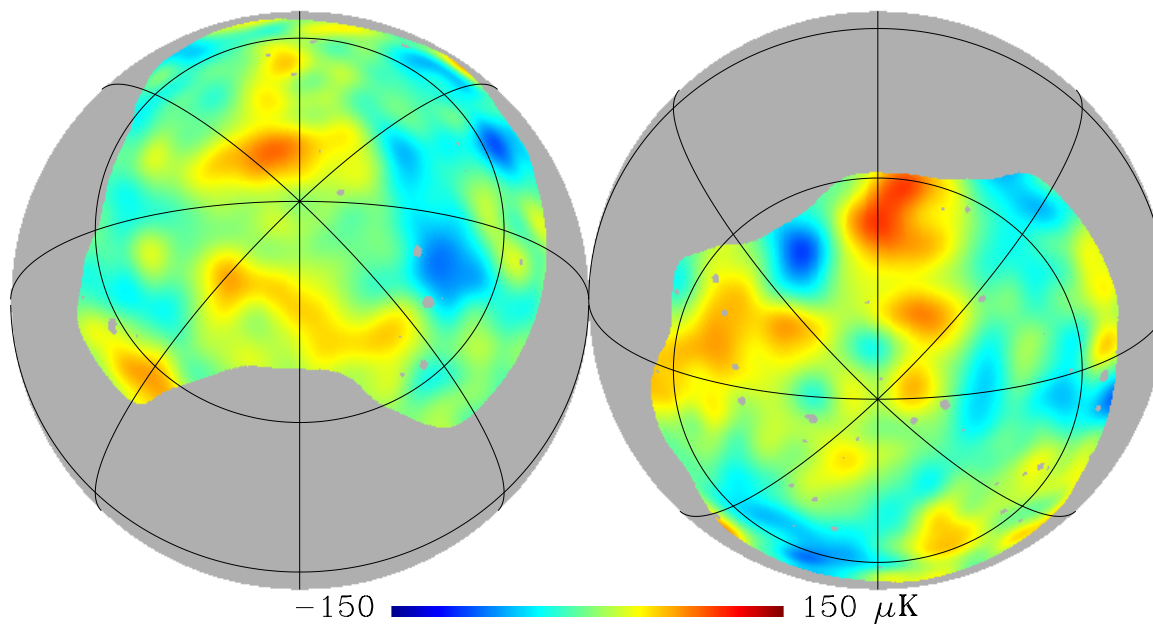
↙  
~ scale that enters the horizon  
at the onset of slow roll



standard mask

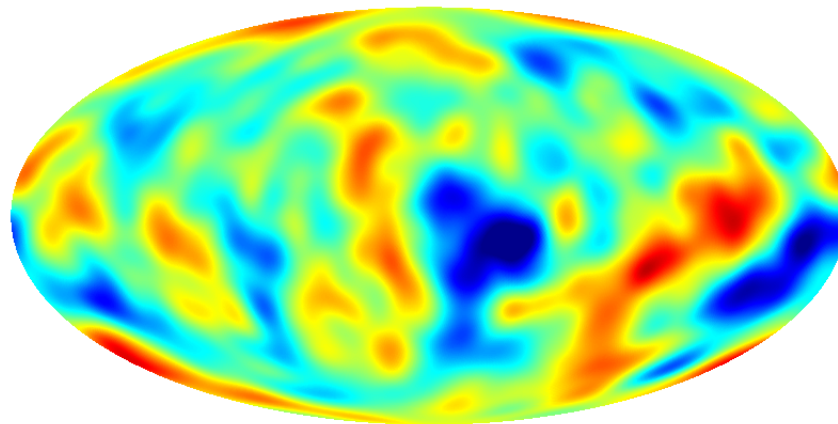


extended mask



A. Gruppuso, N.  
Kitazawa, N.  
Mandolesi, PN, A.  
Sagnotti 2017

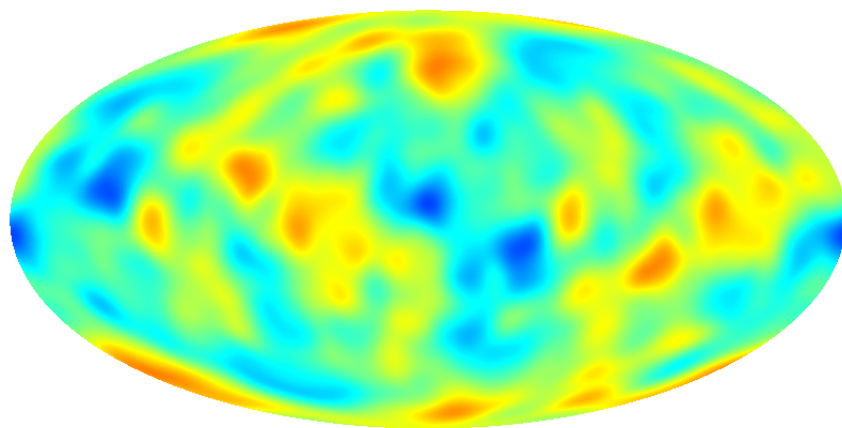
even+odd



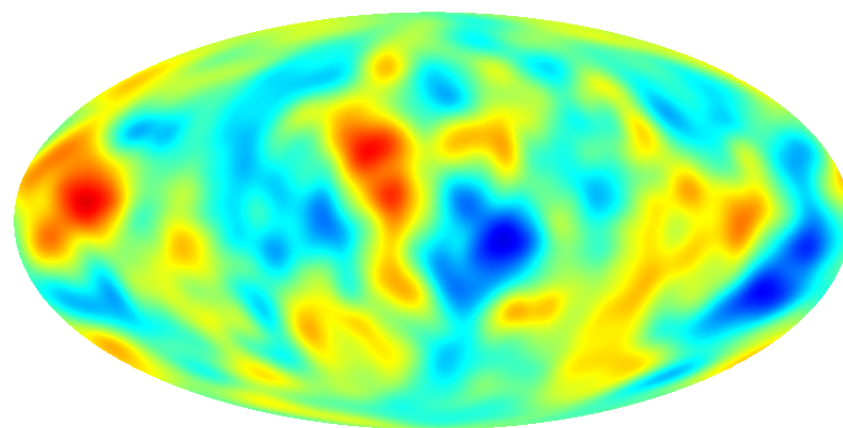
Planck 2015 data

-100.0 100.0  $\mu\text{K}$ 

even

-100.0 100.0  $\mu\text{K}$ 

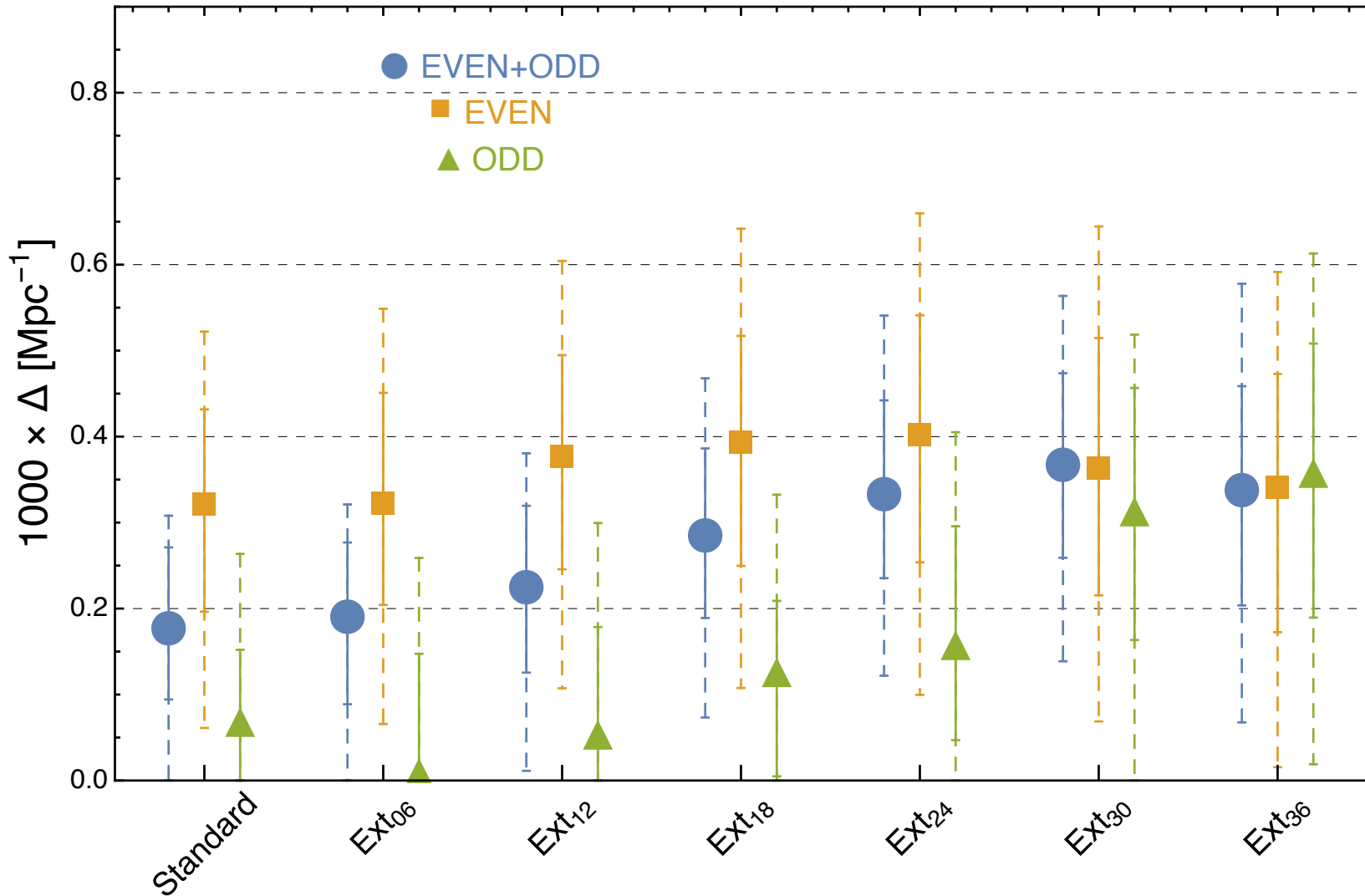
odd

-100.0 100.0  $\mu\text{K}$ 

$$\left( \frac{\delta T}{T}(\hat{n}) \right)_{\pm} \equiv \frac{1}{2} \left[ \frac{\delta T}{T}(\hat{n}) \pm \frac{\delta T}{T}(-\hat{n}) \right]$$

A. Gruppuso, N. Kitazawa, M. Lattanzi, N. Mandolesi, PN, A. Sagnotti 2017

# Constraints on $\Delta$ from Planck 2015

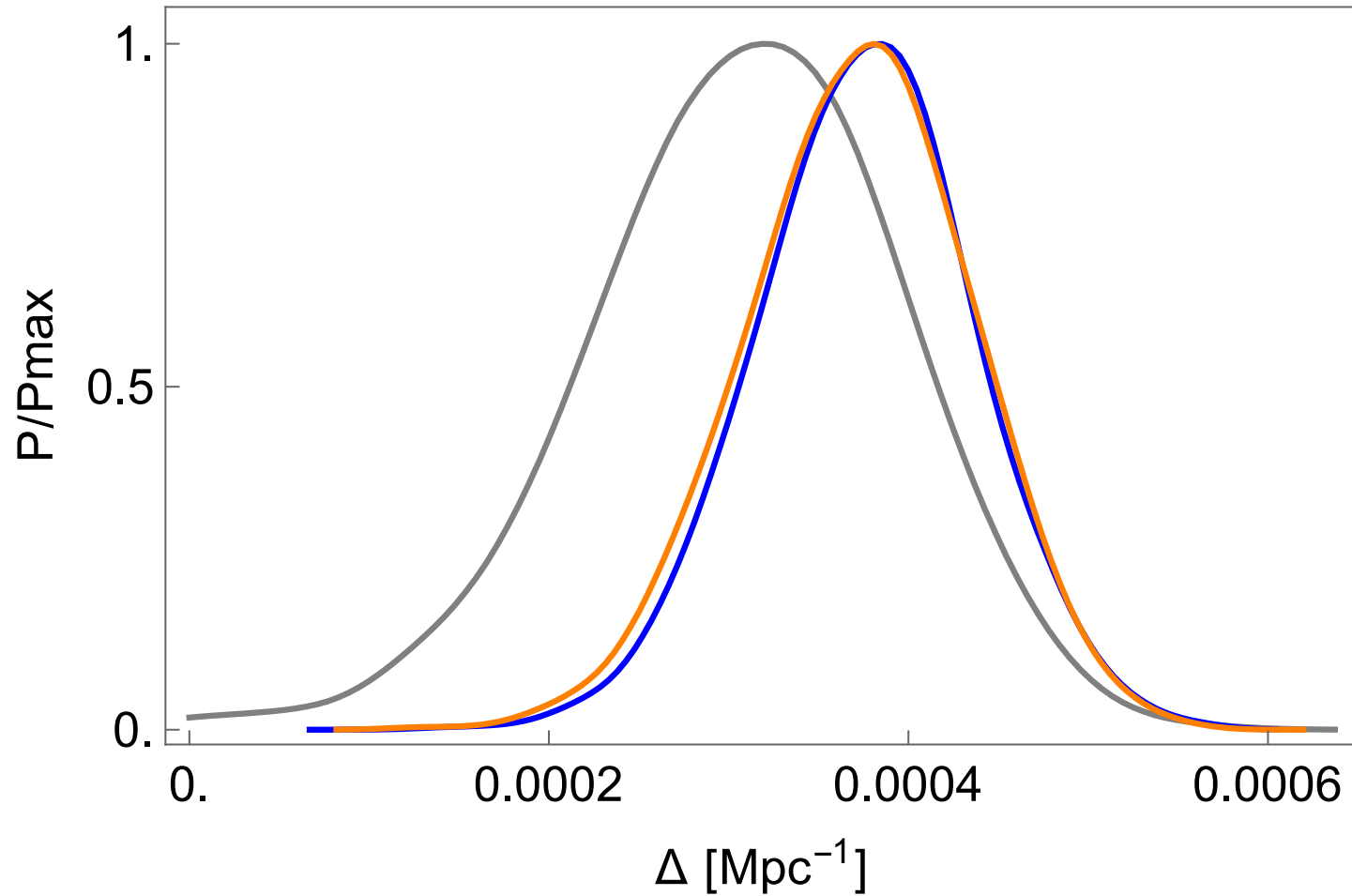


A. Gruppuso, N. Kitazawa, M. Lattanzi, N. Mandolesi,  
PN, A. Sagnotti 2017



- The even multipoles are consistently lower than the  $\Lambda$ CDM expectation, independently on the galactic masking
- The odd multipoles are consistent with the  $\Lambda$ CDM expectation for the smaller masks (more sky). In larger masks (less sky), they are consistent with the even multipoles (and then have low power)
- The power at large scales is concentrated around the galactic plane, in the odd multipoles

# Forecasted constraints on $\Delta$ from future experiments



- Grey:** Planck-like noise, standard masking
- Orange:** CVL large-scale pol., ext30 mask
- Blue:** CVL large-scale pol., full sky

# PRESENT AND FORTHCOMING CMB PROBES

## Ground



POLARBEAR



ACTPol

Atacama,  
Chile

In addition,  
ABS, CLASS, POLARBEAR-2,  
Simons Array, Adv-ACTPol, ...



BICEP1 BICEP2 DASI QUAD KECK  
SPTPol

South  
Pole

In addition, BICEP3, POLAR, QUBIC, ...

Data SIO, NOAA, U.S. Navy, NGA, GEBCO  
© 2011 Inuv/Geosistemas SRL

## Balloon

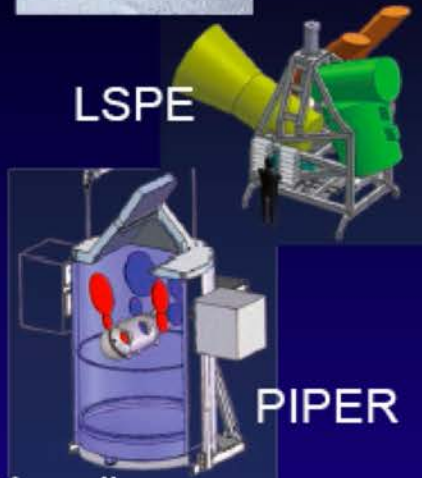


EBEX



SPIDER

LSPE



PIPER

## Satellite



WMAP  
(obs. end  
in 2010)



Planck



LiteBIRD



PIXIE



CoRE+

In addition, QUIJOTE in Canary island, AMiBA in Hawaii



# PRESENT AND FORTHCOMING CMB PROBES

## Ground



POLARBEAR



ACTPol

Atacama, Chile

In addition, ABS, CLASS, POLARBEAR-2, Simons Array, Adv-ACTPol, ...

**and many more!**

BICEP1 BICEP2 QUAD KECK



In addition, BICEP3, POLAR, QUBIC, ...

Data SIO, NOAA, U.S. Navy, NGA, GEBCO © 2011 Inev/Geosistemas SRL

## Balloon



EBEX



SPIDER



LSPE SWIPE

**LSPE r STRIP**

PIPER

## Satellite



WMAP (obs. end in 2010)



Planck

**Retired (legacy remains)**



LiteBIRD



PIXIE

**Not funded! Not funded! (India interested)**



CORE+

In addition, QUIJOTE in Canary island, AMiBA in Hawaii

# Conclusions

- Planck has delivered its final (legacy) release
- It has provided the ultimate (cosmic variance limited) measurement of CMB anisotropy
- ... But just opened the door of CMB polarization (which was never designed to measure, by the way)
- It has fulfilled its promise of measuring the fundamental cosmological parameters to percent accuracy
- And brought remarkable constraints on particle physics parameters as well, excluding a fourth fully thermalized neutrino and constraining the total neutrino masses in the 100 meV range.
- Has measured well one relevant inflationary parameter, the primordial spectral index, allowing constraints on the inflationary paradigm
- Yet has uncovered several tensions with astrophysical measurements, which may or may not hint at new physics.
- Intrinsic anomalies do exist in the large-angle CMB field, which may also be a tracer of something new.
- If these tension/anomalies are really hinting at new physics, its signature in the CMB is scant. Accurate measurements are needed to pin down the issue.
- Primordial gravitational waves remain unseen.
- To exploit the wealth of information that still is in the CMB, we need to cope with the extraordinary complexity of the sky. This can be credibly done only with a future space mission.



# The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.

X COSPAR 2018, July 2018