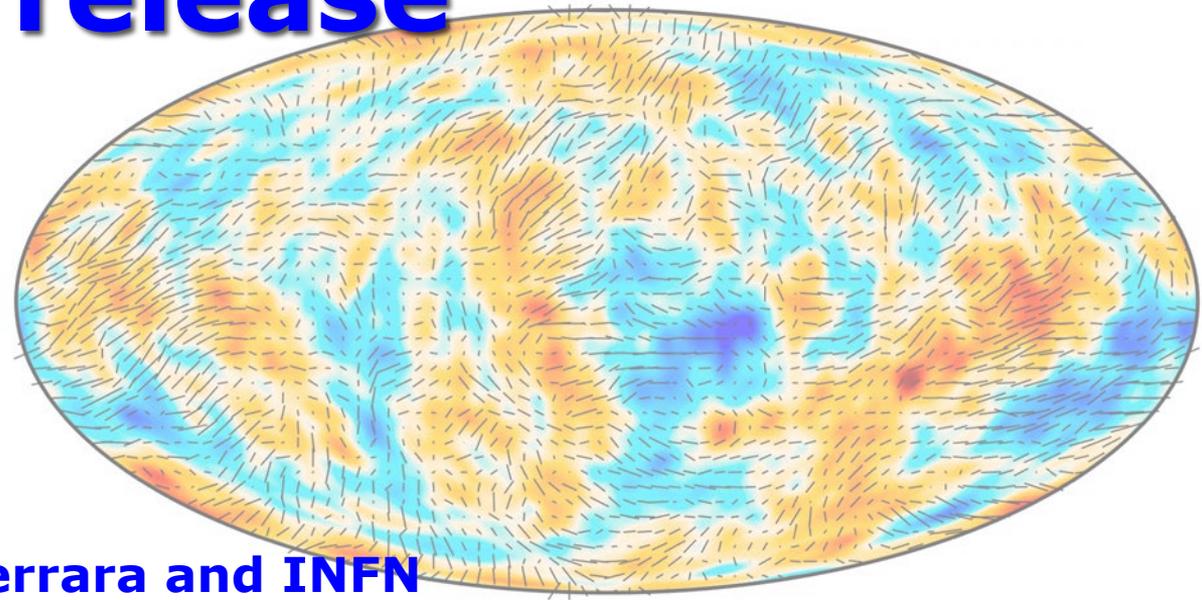




planck

Cosmology after the Planck legacy release



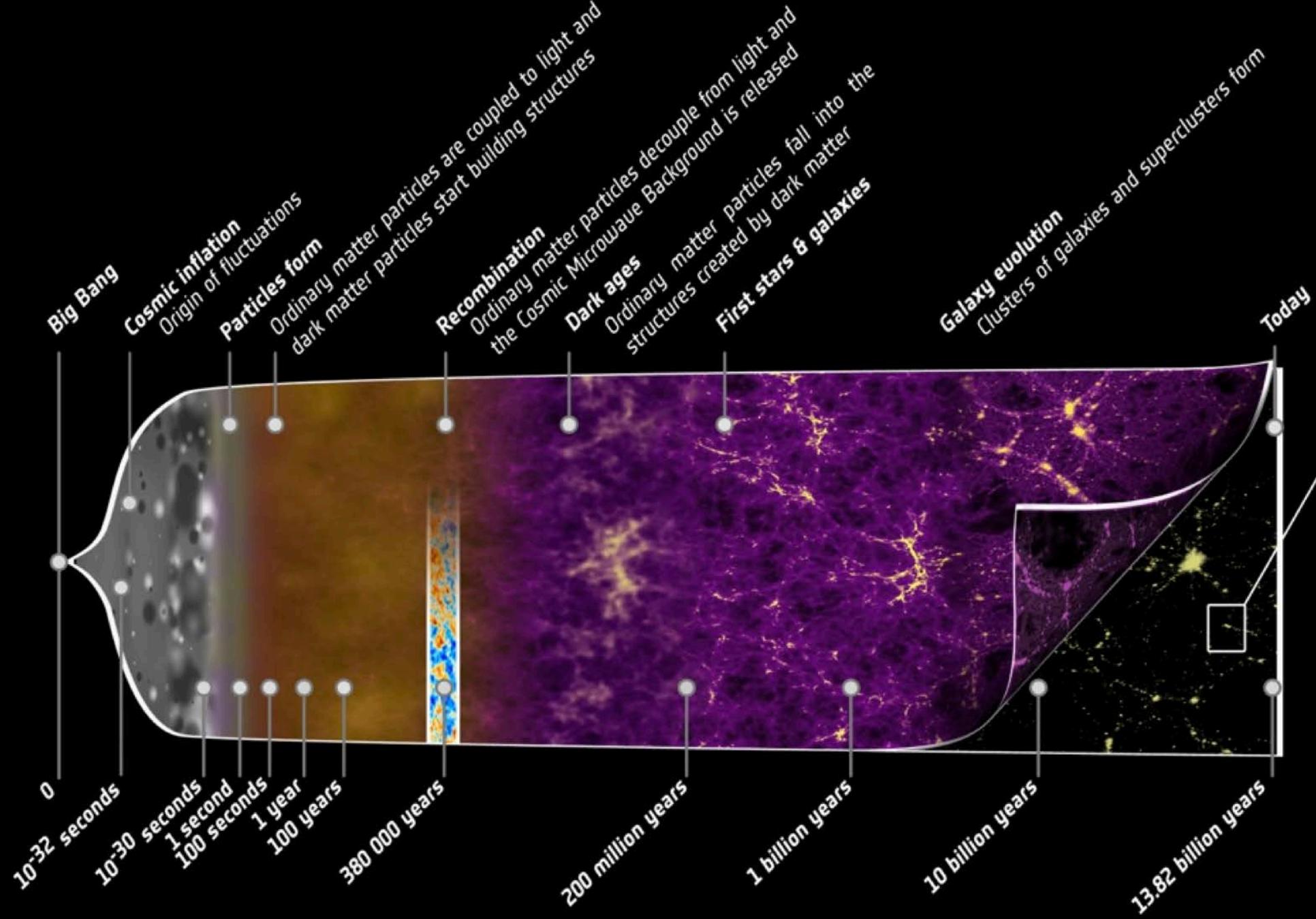
Paolo Natoli

Università di Ferrara and INFN

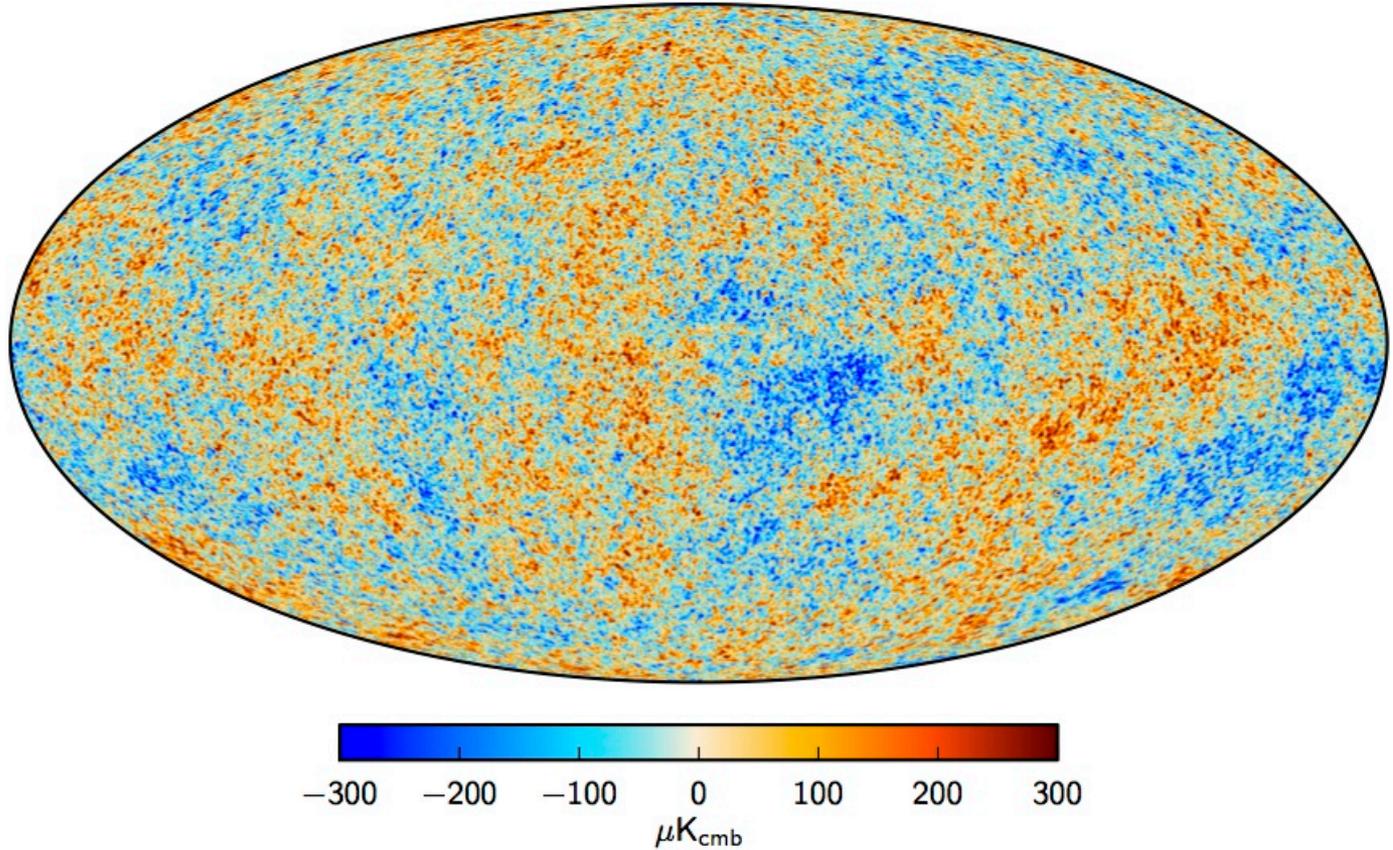
*XVIII International Workshop on Neutrino Telescopes
Venezia, 21 March 2019*



UNIVERSITÀ
DEGLI STUDI
DI FERRARA
- EX LABORE FRUCTUS -



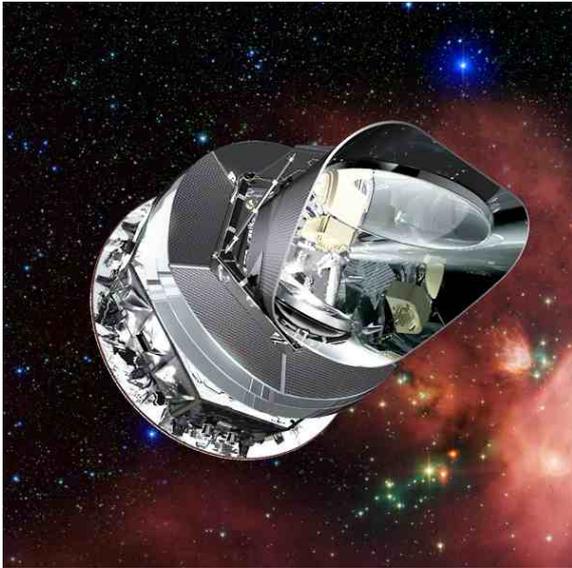
THE COSMIC MICROWAVE BACKGROUND



The CMB is a blackbody radiation with $T=2.7$ K extremely uniform across the whole sky; it is the relic radiation emitted at the time the nuclei and electrons recombined to form neutral hydrogen, when the Universe was $\sim 400,000$ years old (the so-called last scattering surface, LSS).

Its tiny ($\sim 10^{-5}$) temperature and polarization anisotropies encode a wealth of cosmological information.

THE PLANCK SATELLITE



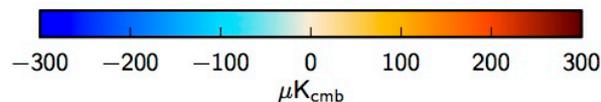
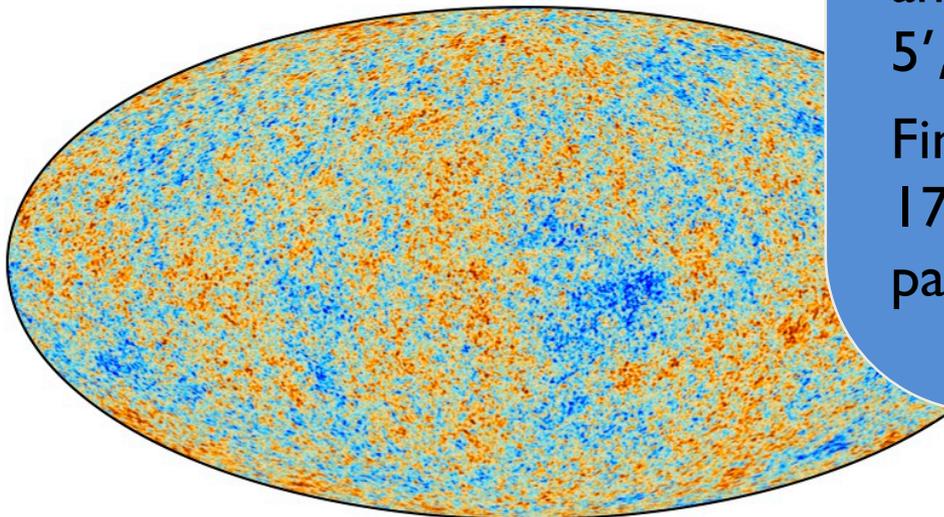
Planck is the 3rd generation ESA satellite devoted to CMB

Ultimate characterization of the temperature anisotropies

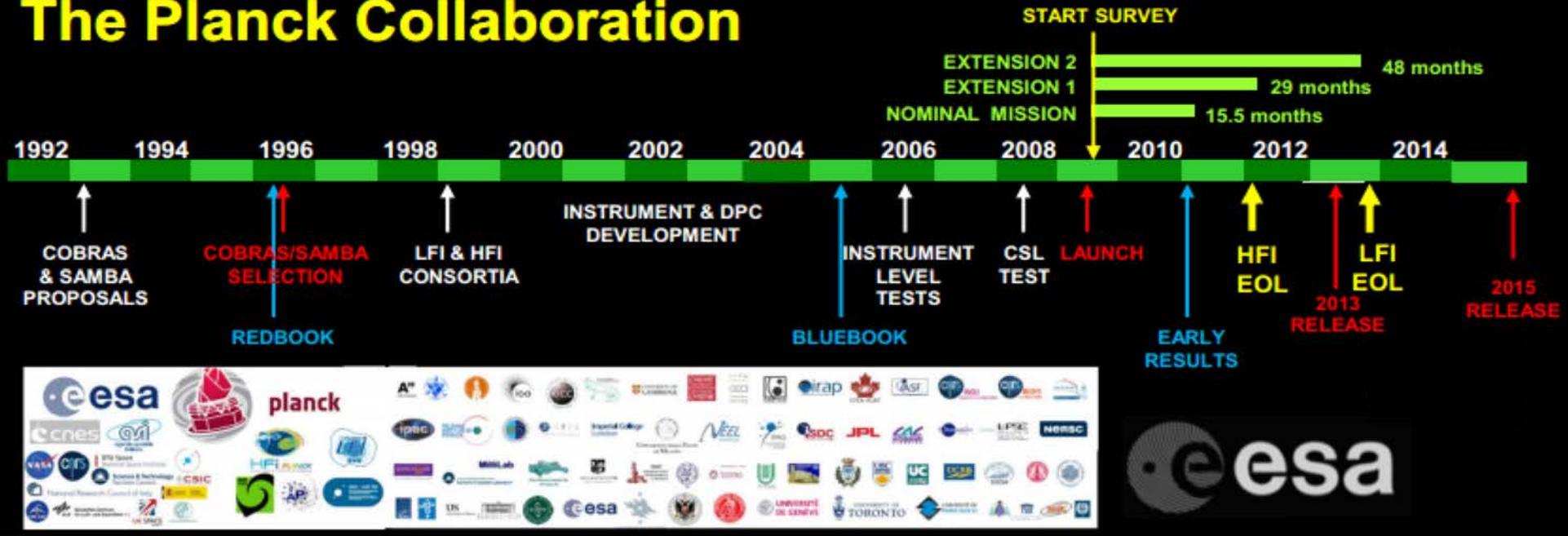
74 detectors (radiometers and bolometers) in 9 frequency bands from 30 to 857 GHz

angular resolution between 30' and 5', $\Delta T/T \sim 2 \times 10^{-6}$

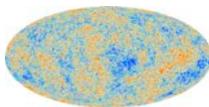
Final (legacy) release took place on 17 July 2018, for data and (most) papers.



The Planck Collaboration



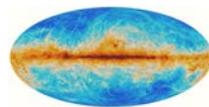
May 2009: Launched from Kourou



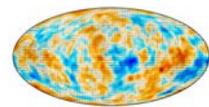
Mar 2013: Data Release and Cosmology Results
 Nominal Mission Temperature data 32 papers



Oct 2013: Planck 'Shut Down' 52 papers / intermediate results

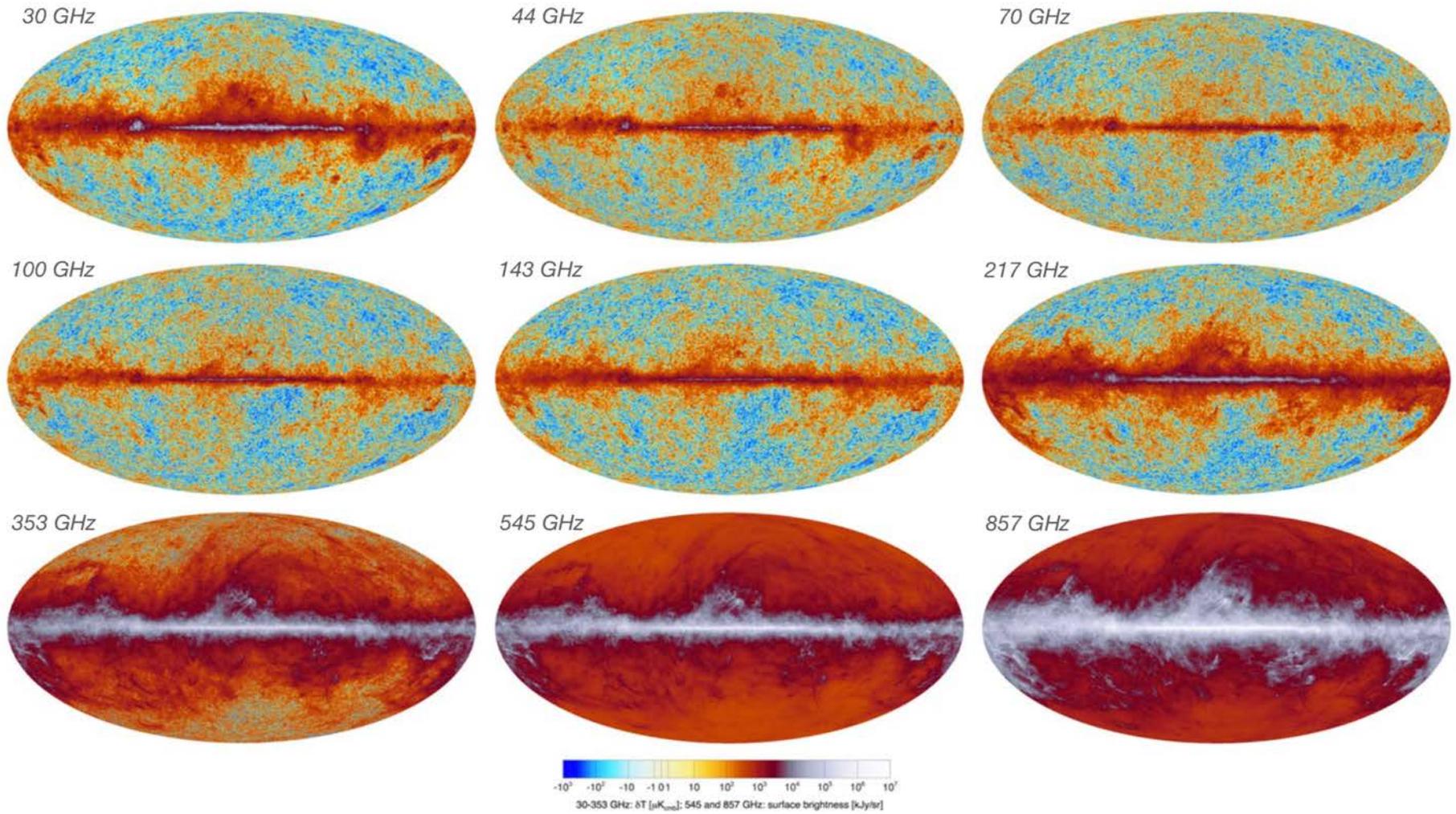


Feb 2015: Data Release and Cosmology Results
 Full Mission Temperature and (preliminary) Polarization data 28 papers



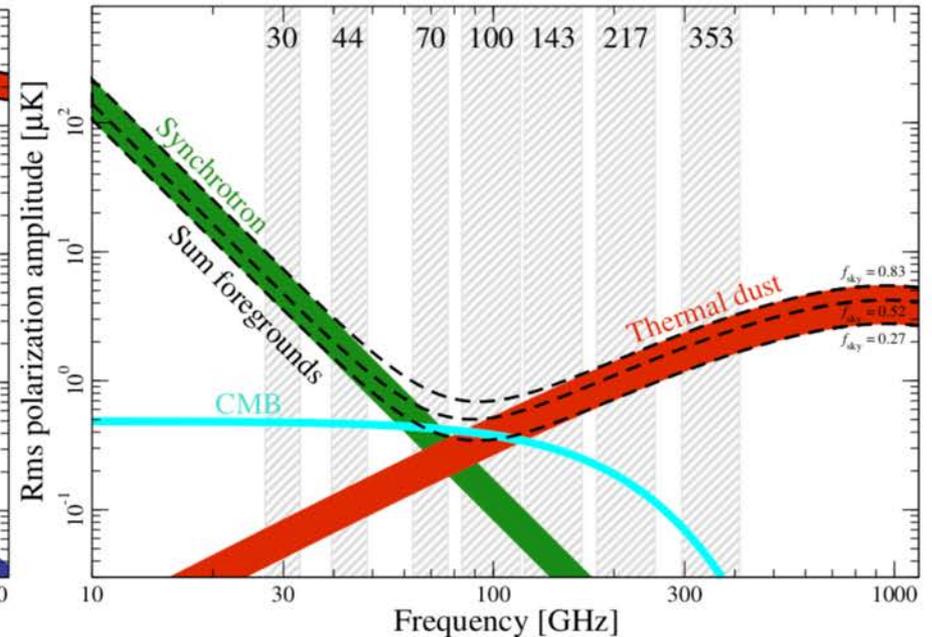
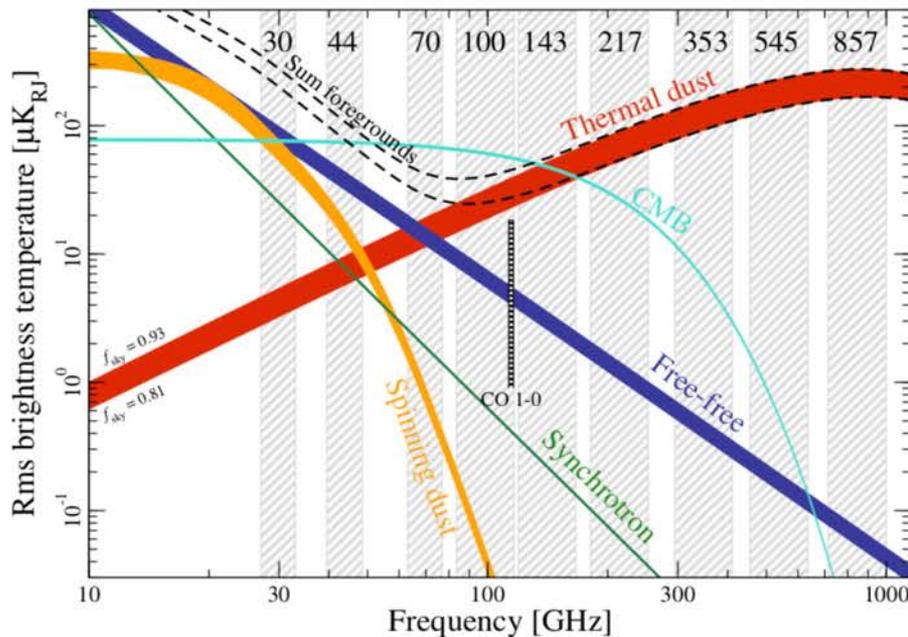
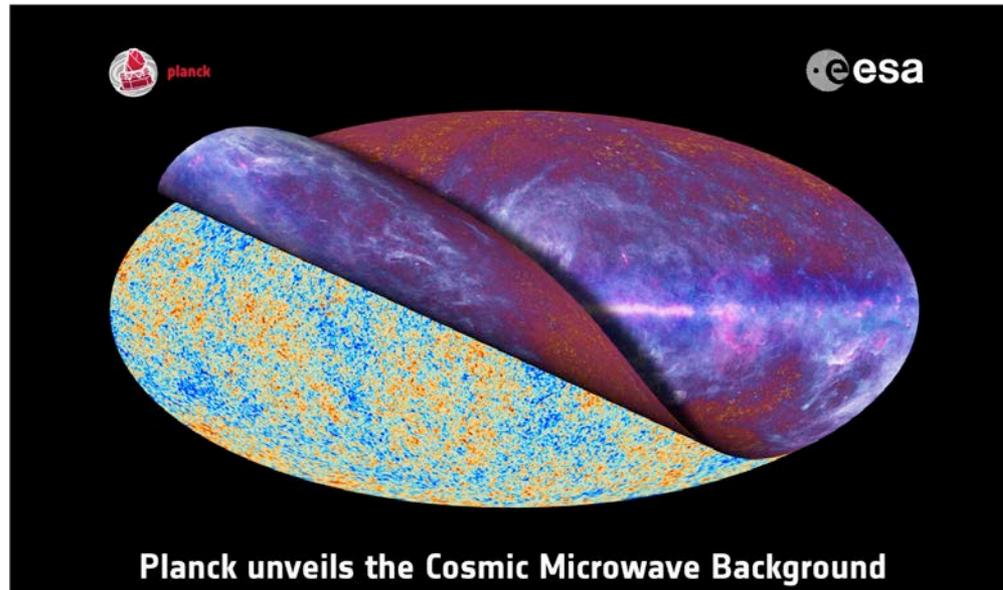
Jul 2018: Legacy Data & Paper Release 9 papers (+3 to appear soon)

THE TEMPERATURE SKY AS SEEN BY PLANCK 2018

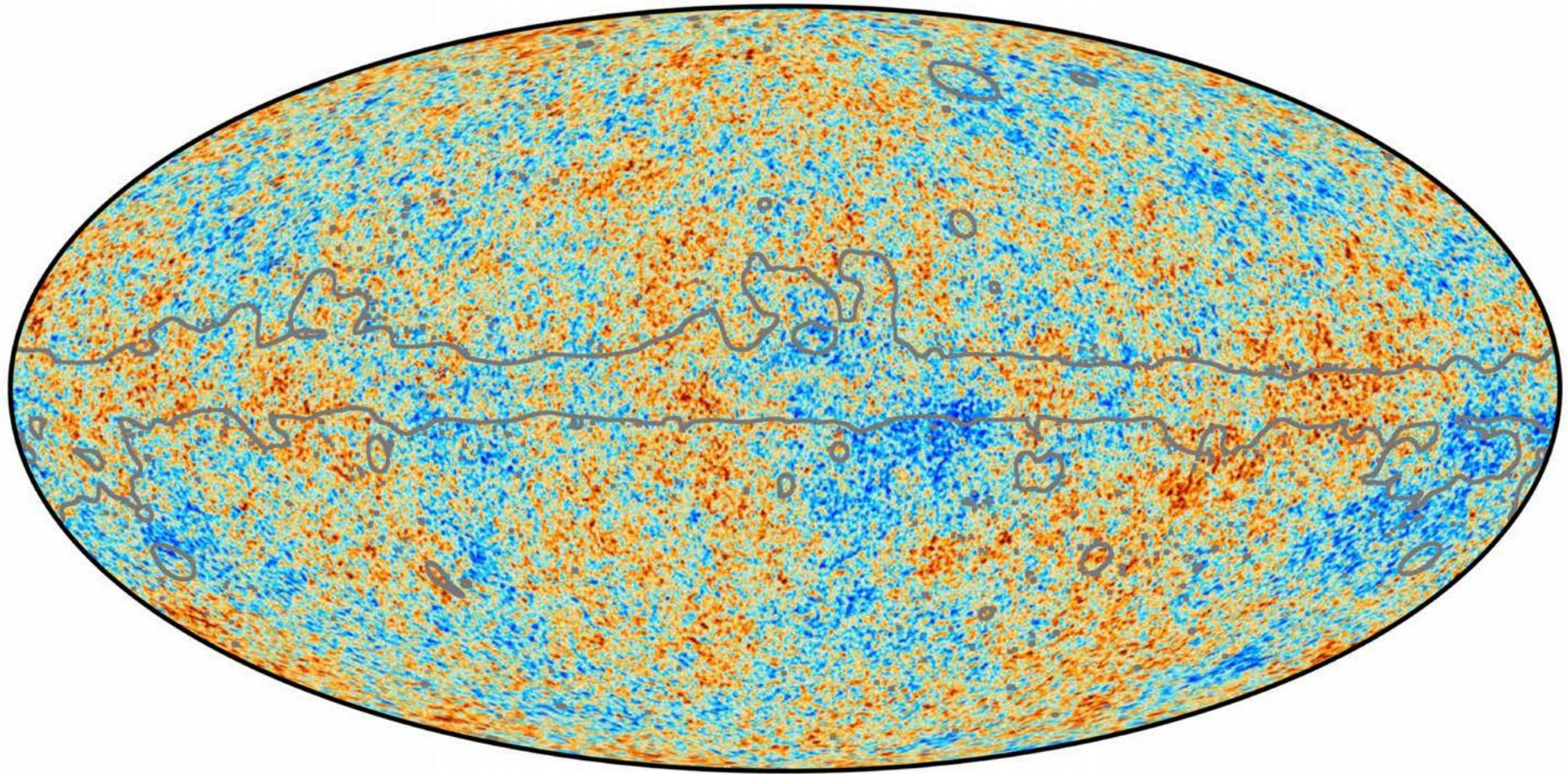


UNVEILING THE CMB SKY

The *ultimate* measurement of the CMB temperature anisotropy field

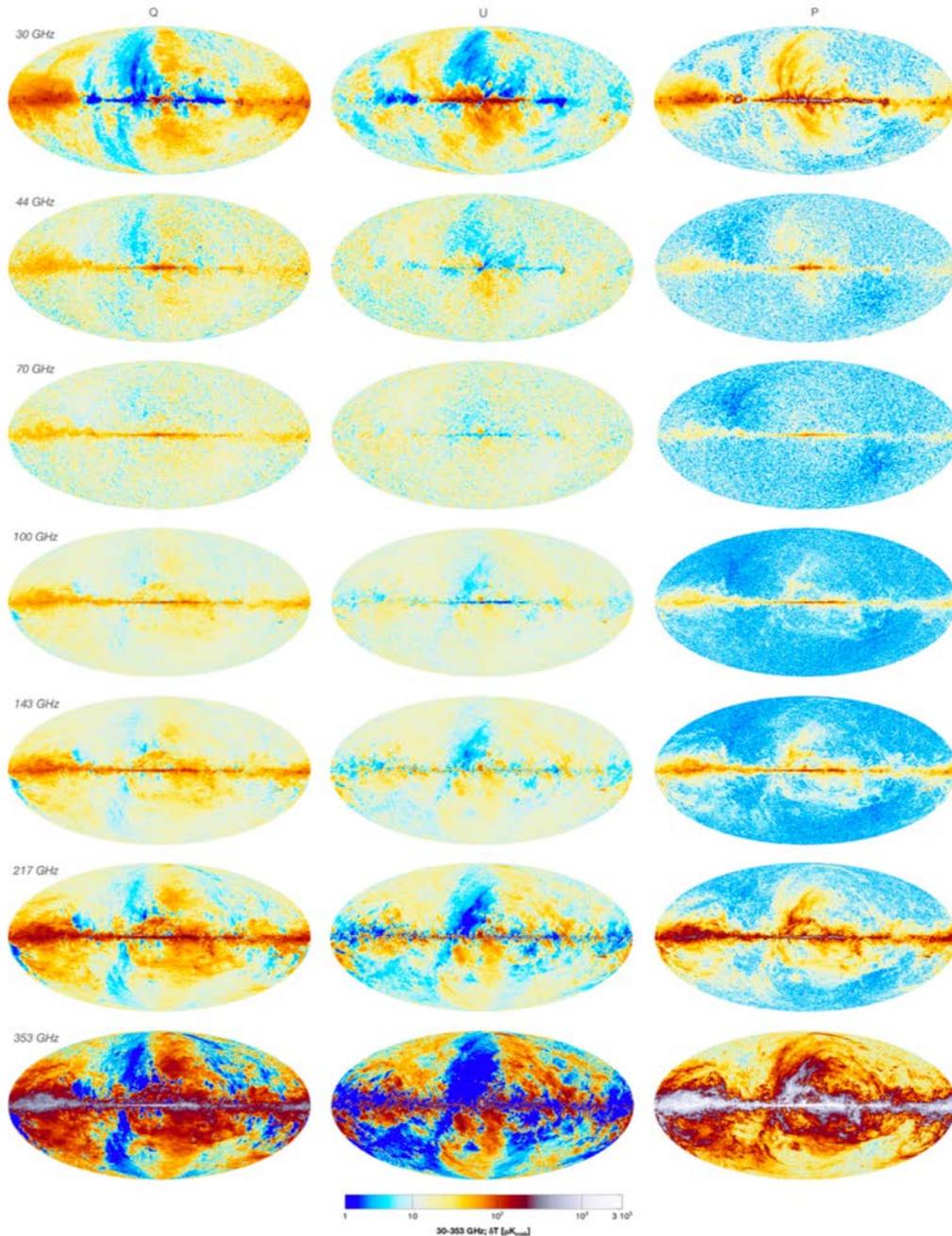


PLANCK: TEMPERATURE ANISOTROPIES



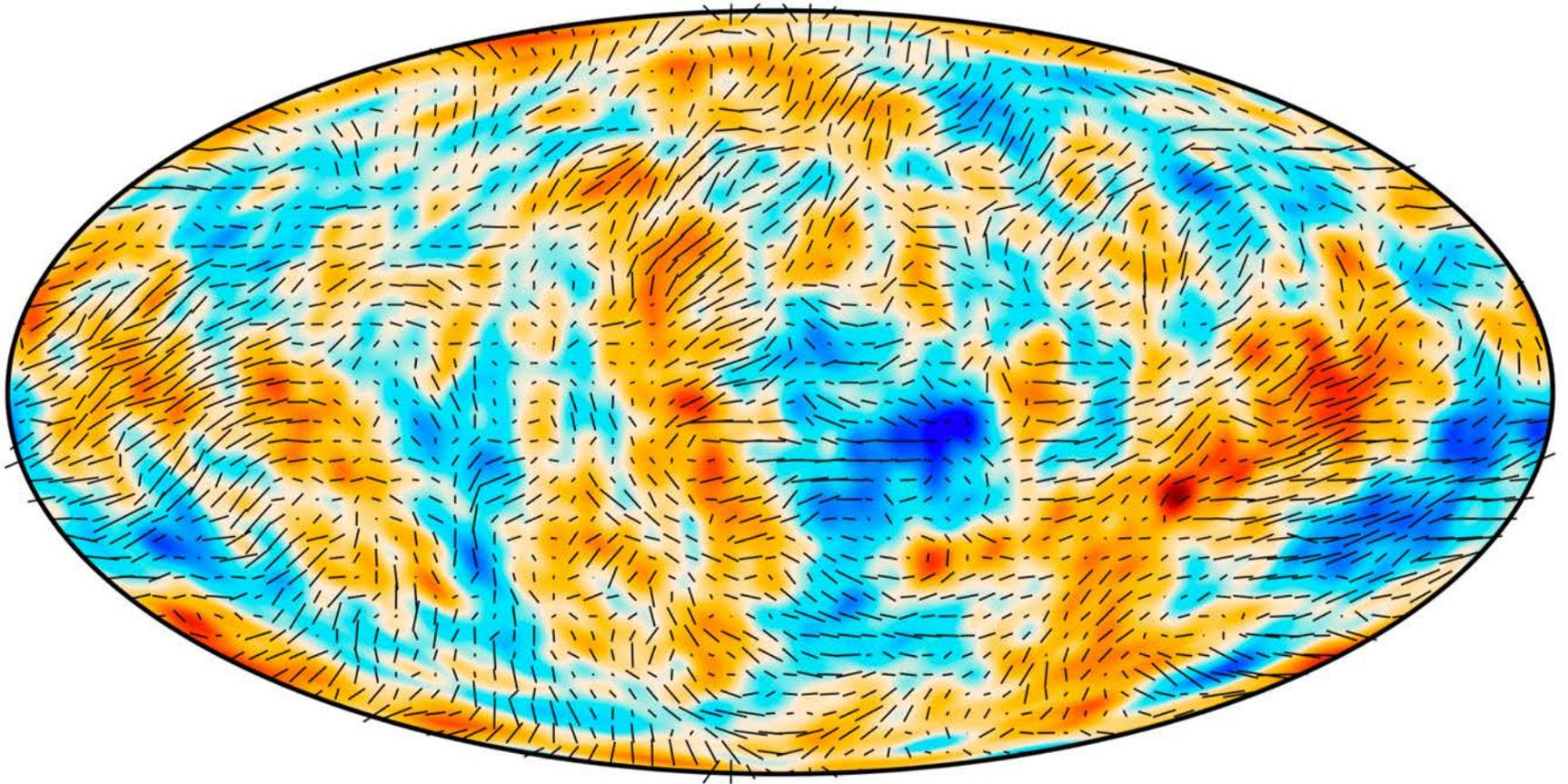
-300  300 μK

THE POLARIZATION SKY AS SEEN BY PLANCK 2018



Significant reduction of large scale polarization systematics in 2018

PLANCK: POLARIZATION ANISOTROPIES



1 0.41 μK

-160



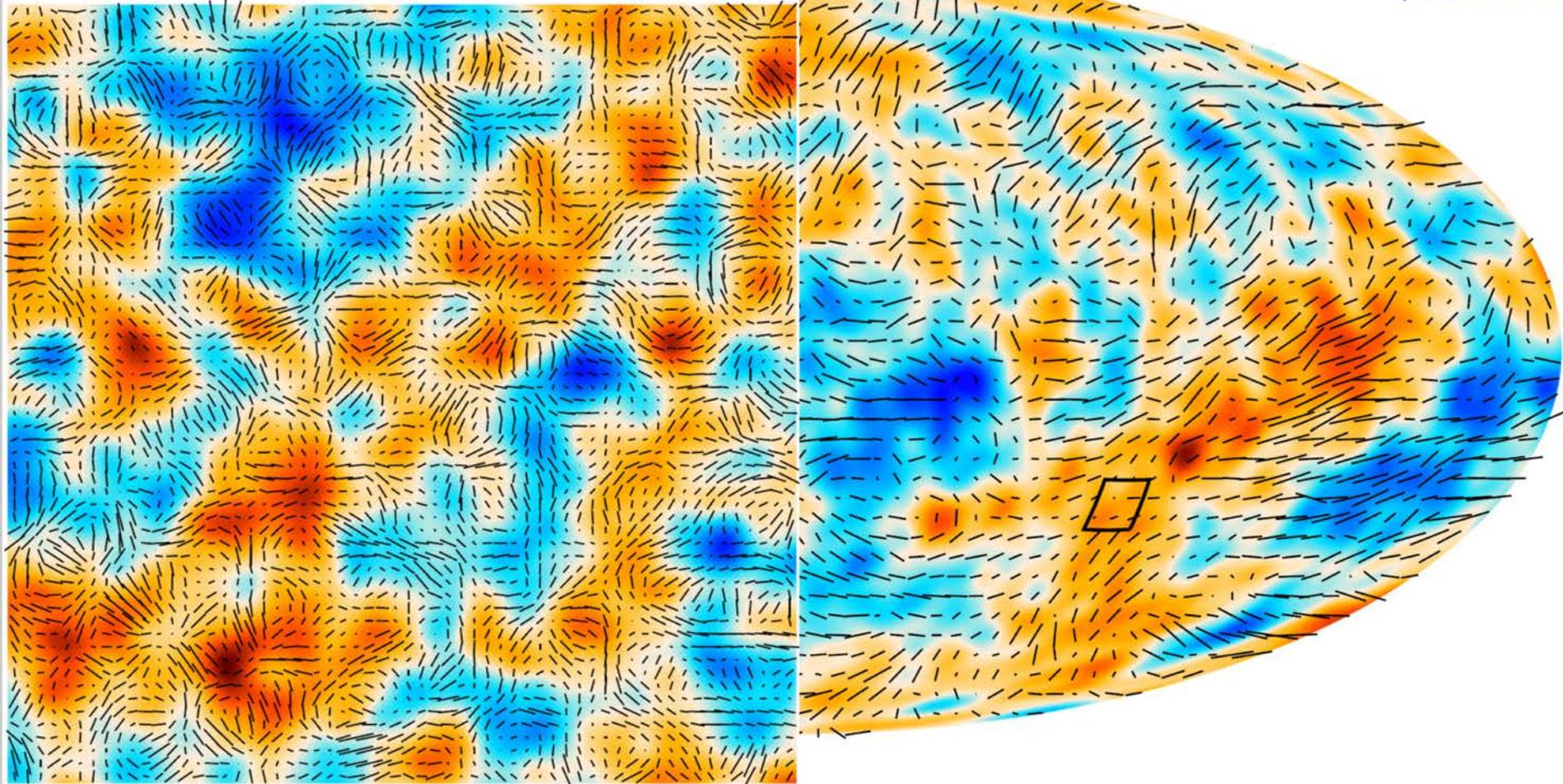
160 μK

Temperature smoothed to 5 degrees

PLANCK: POLARIZATION ANISOTROPIES

$10^\circ \times 10^\circ$, smoothed at $20'$

(Planck 2018 I)



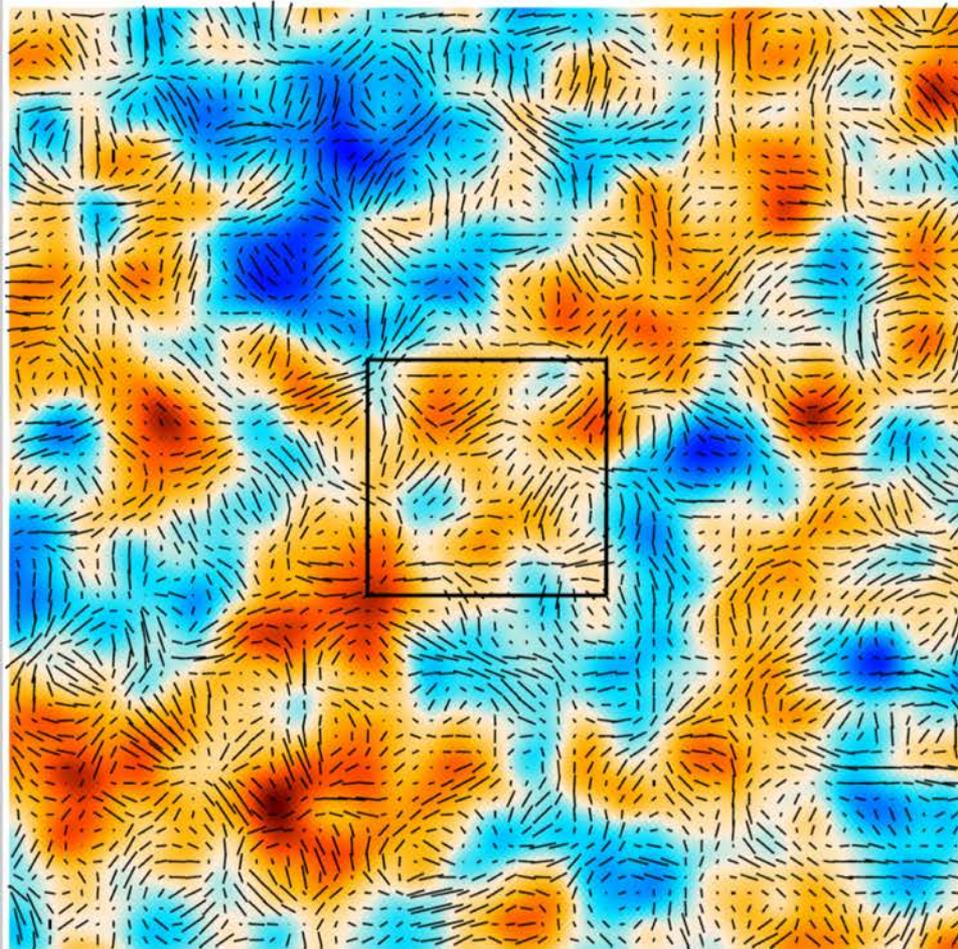
-201 μK 309 μK

13.7 μK

(276.4, -29.8) Galactic

PLANCK: POLARIZATION ANISOTROPIES

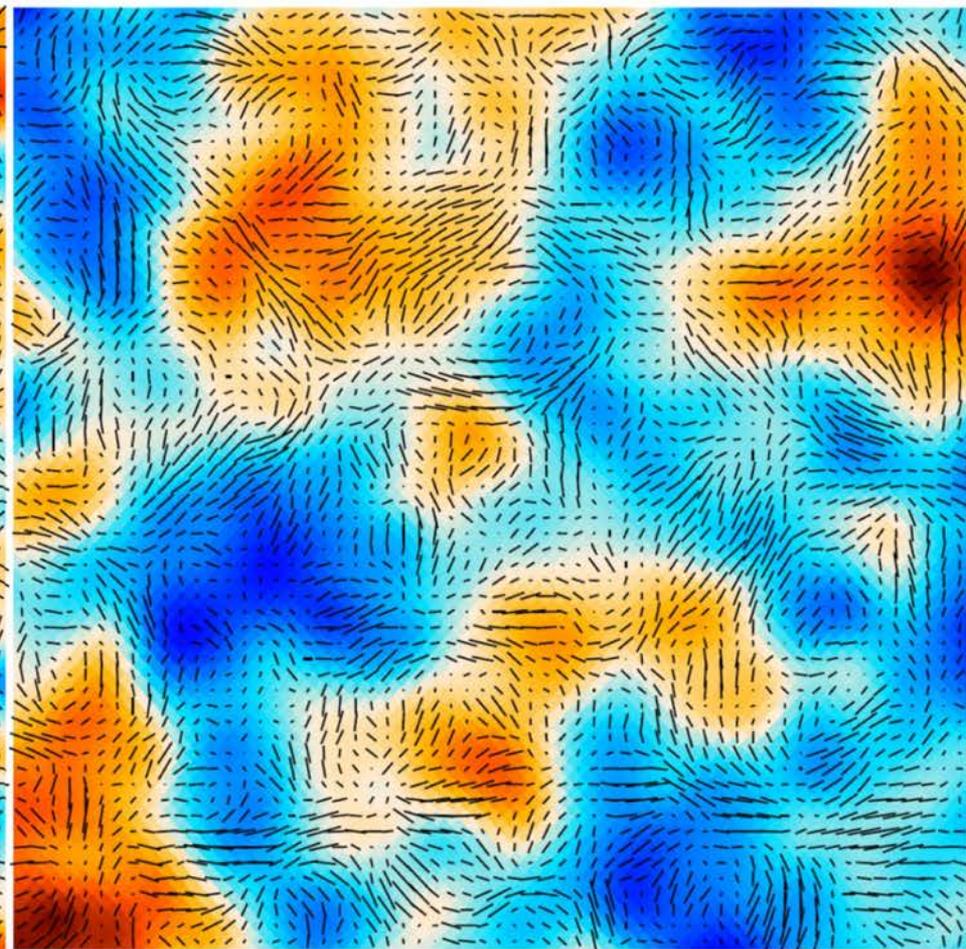
$10^\circ \times 10^\circ$, smoothed at $20'$



-201 309 μK
(276.4, -29.8) Galactic

13.7 μK

$2.5^\circ \times 2.5^\circ$, smoothed at $7'$

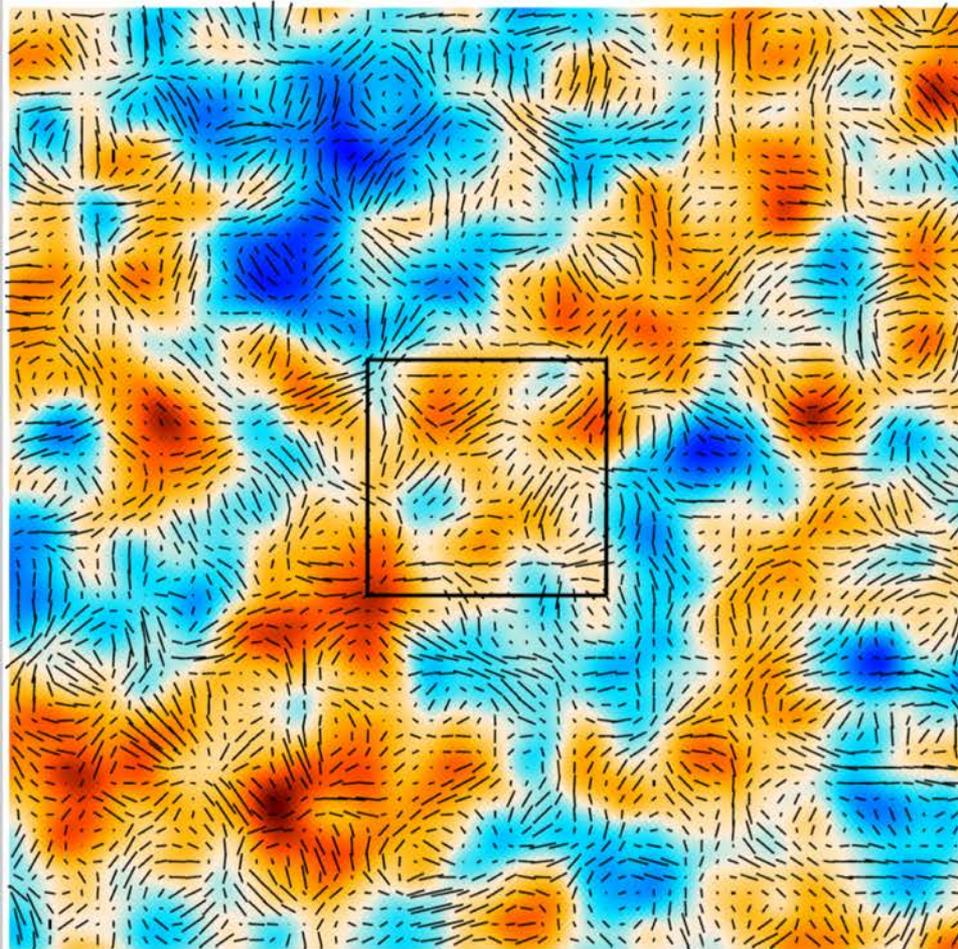


-67 311 μK
(276.4, -29.8) Galactic

36.1 μK

PLANCK: POLARIZATION ANISOTROPIES

$10^\circ \times 10^\circ$, smoothed at $20'$

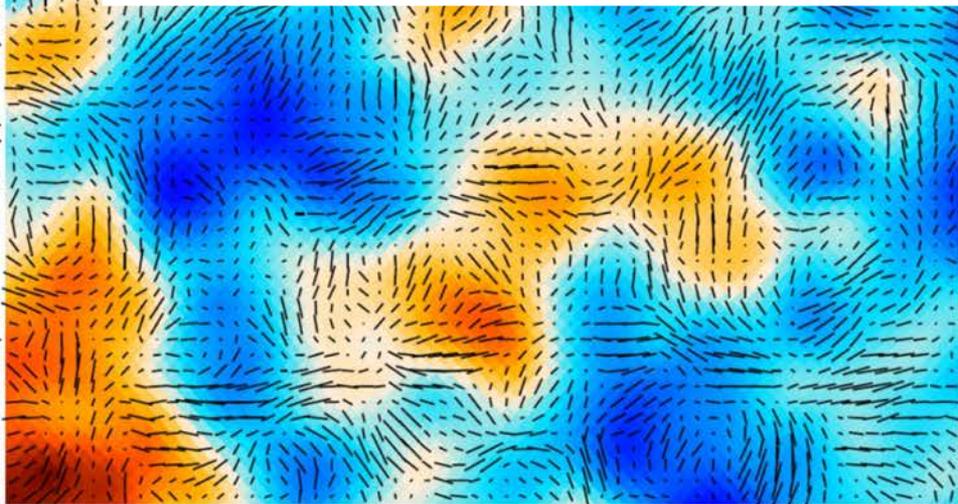
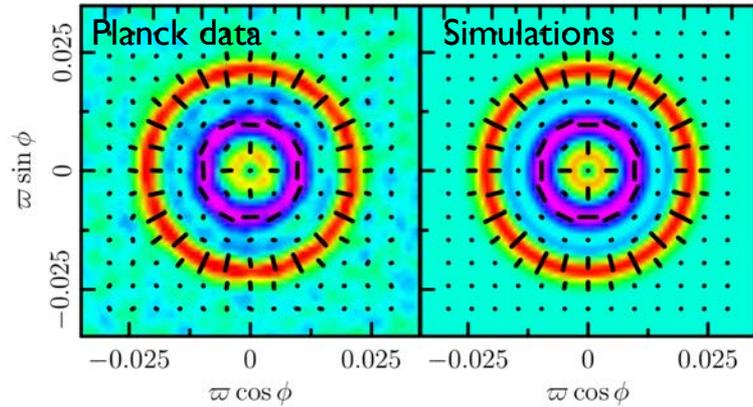
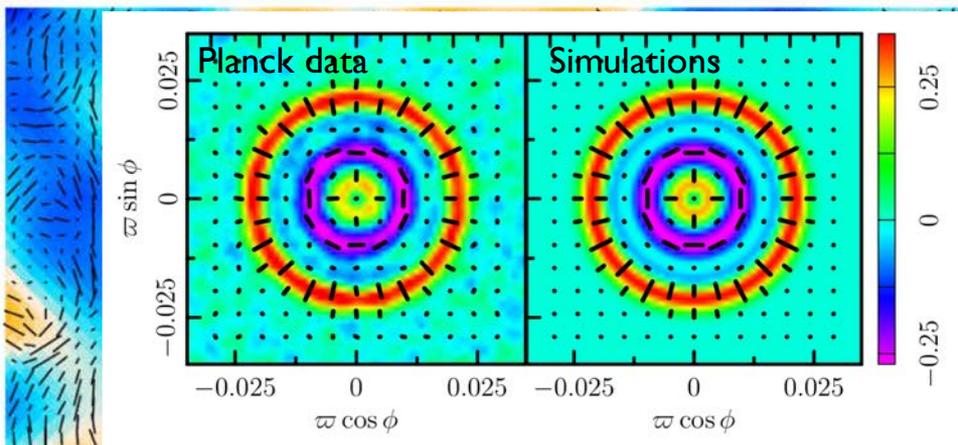


-201  309 μK

13.7 μK

(276.4, -29.8) Galactic

$2.5^\circ \times 2.5^\circ$, smoothed at $7'$

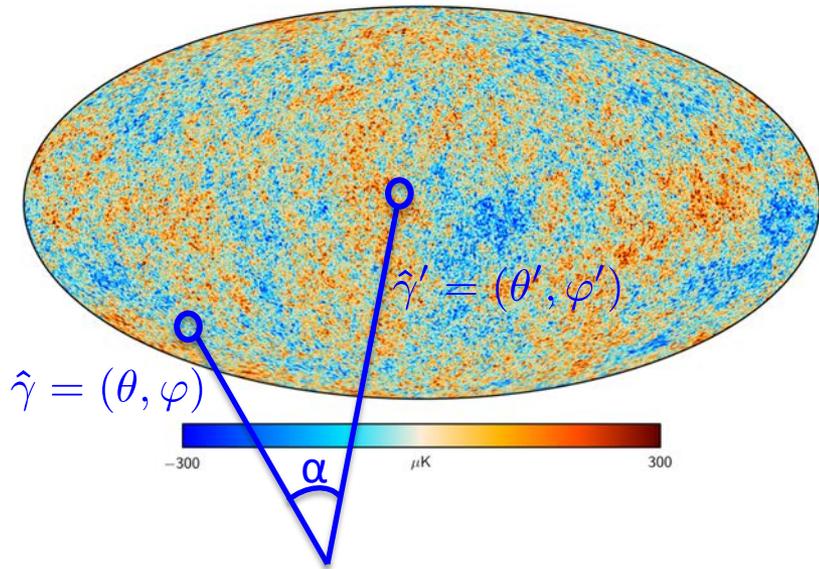


-67  311 μK

36.1 μK

(276.4, -29.8) Galactic

STATISTICAL DESCRIPTION



CORRELATION FUNCTIONS

$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \right\rangle \quad \leftarrow \text{from Inflation}$$

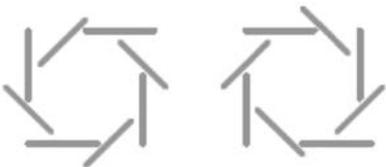
$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \frac{\Delta T}{T}(\vec{\gamma}'') \right\rangle$$

$$\left\langle \frac{\Delta T}{T}(\vec{\gamma}) \frac{\Delta T}{T}(\vec{\gamma}') \frac{\Delta T}{T}(\vec{\gamma}'') \frac{\Delta T}{T}(\vec{\gamma}''') \right\rangle$$

...



E modes



B modes

POLARIZATION

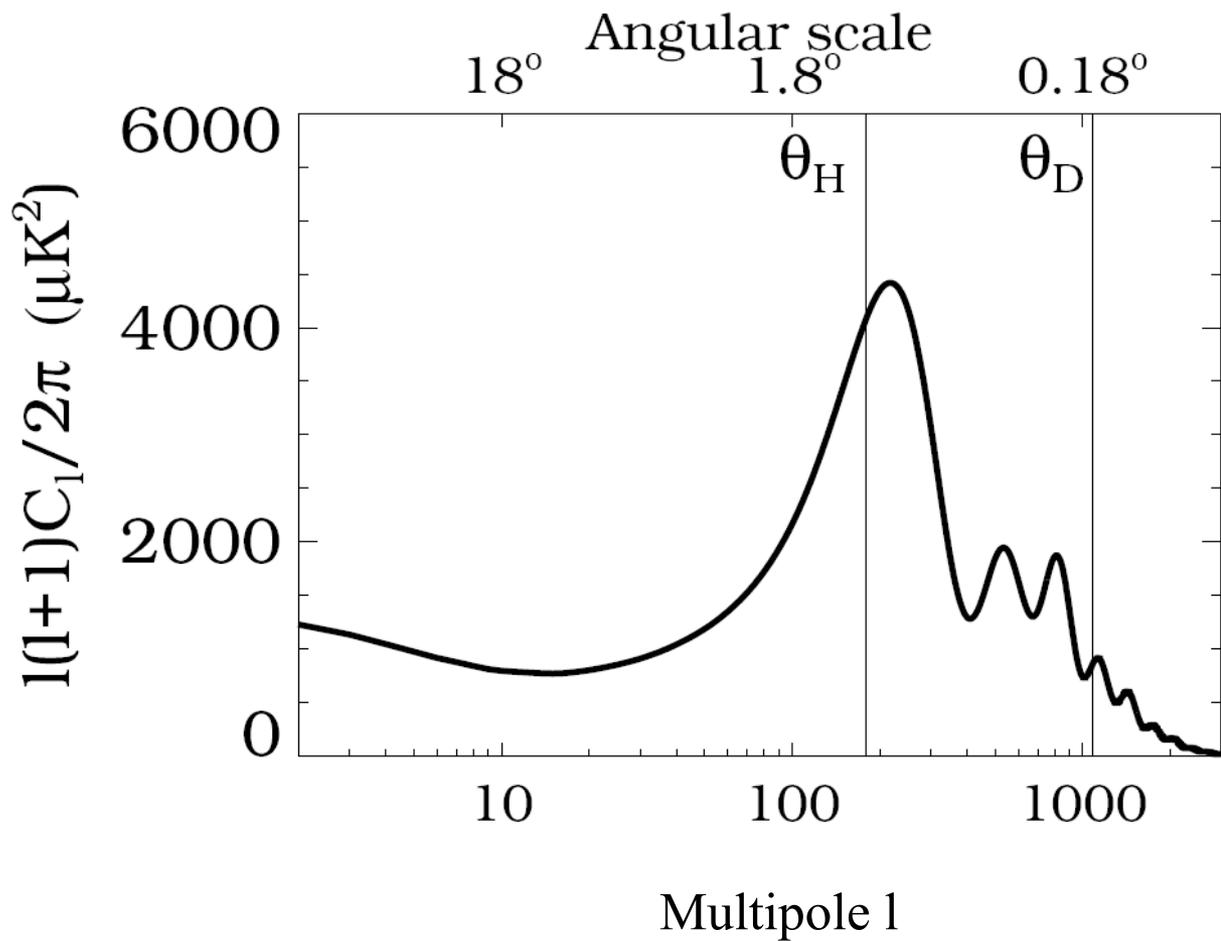
$$\mathbf{P}(\hat{\gamma}) = \nabla \mathbf{E} + \nabla \times \mathbf{B}$$

E-modes: even under parity

B-modes: odd under parity

Density perturbations \rightarrow E-modes

Gravitational Waves \rightarrow E- and B-modes

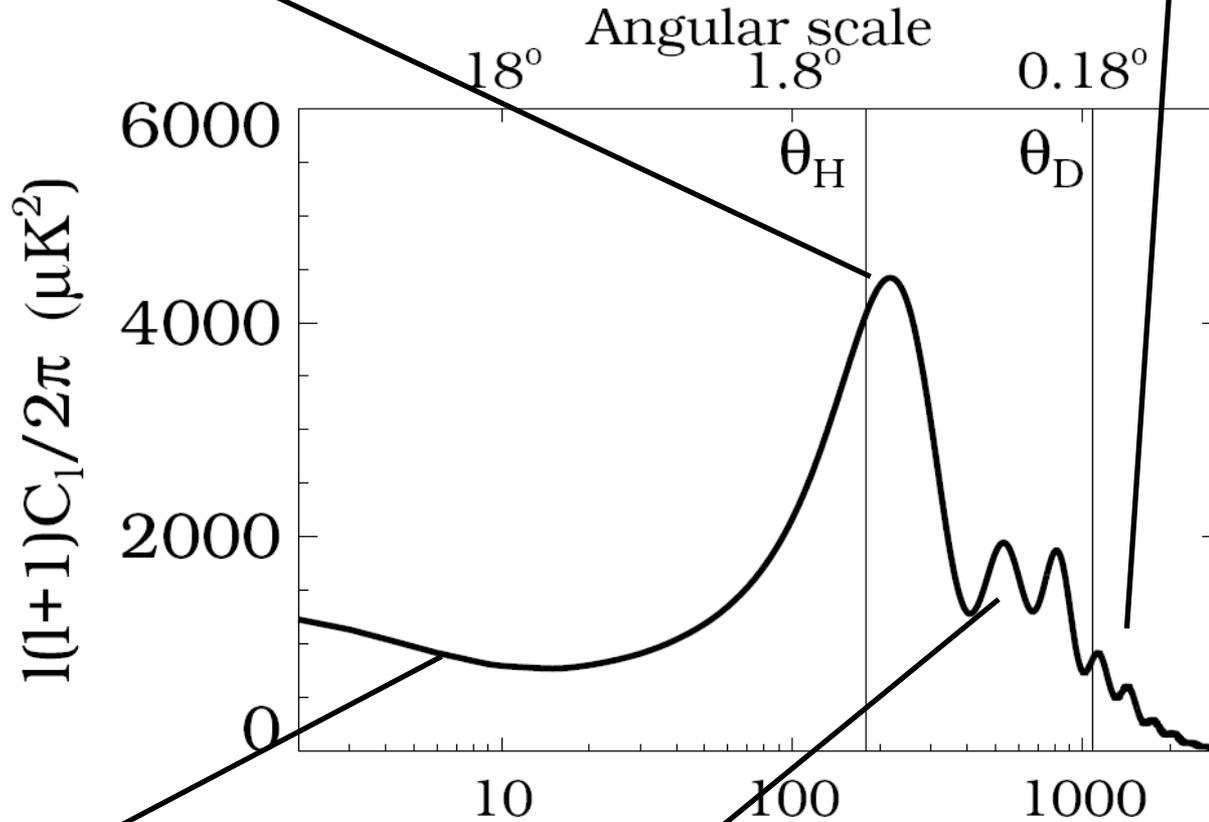


spatial curvature
relative abundance of matter and radiation
distance to the last scattering surface

H_0, Ω_m, Ω_k

Photon diffusion length at recombination
Slope of the primordial spectrum

$N_{\text{eff}}, \Omega_b, Y_p, n_s$



+ Overall power
 $A_s e^{-2\tau}$

+ low-ell
polarization
(not shown)
Reionization
history

τ

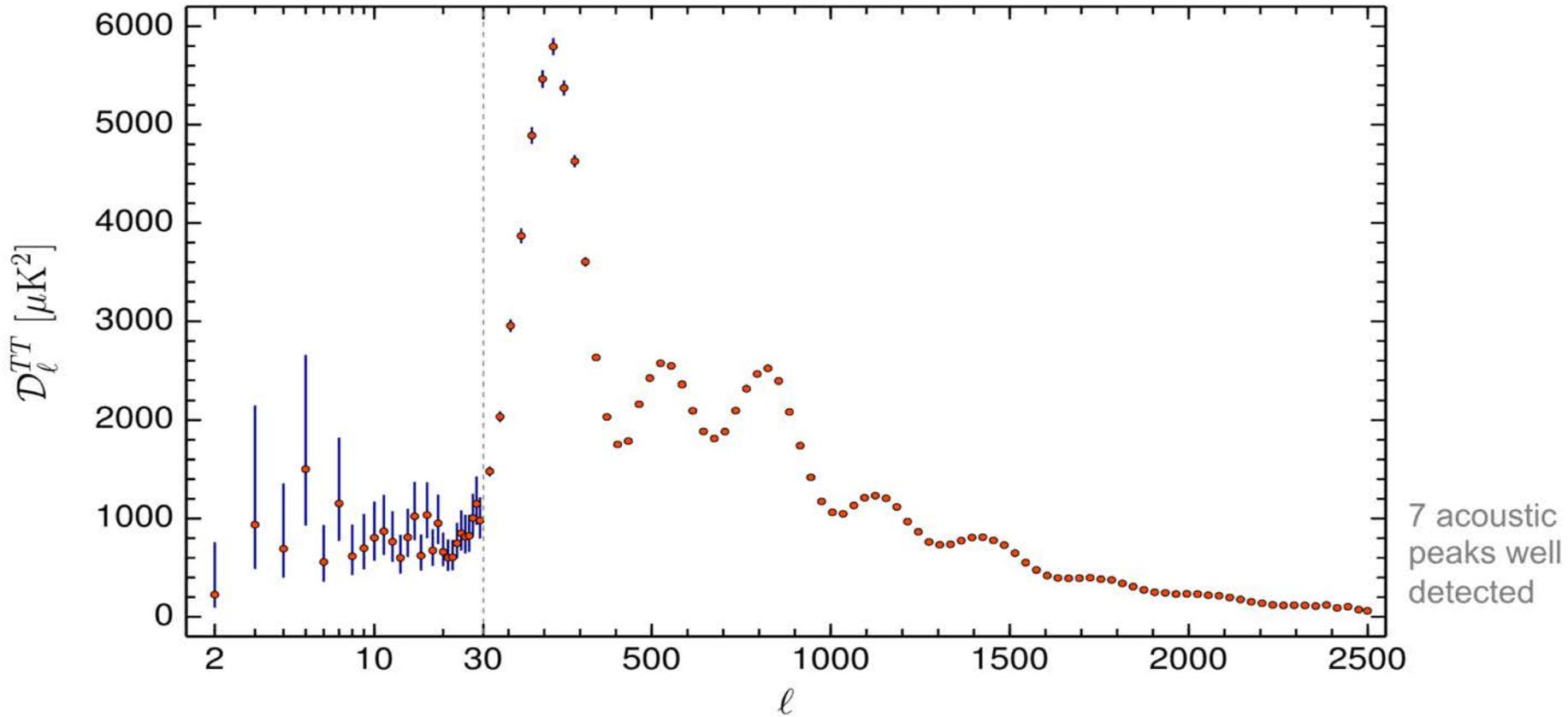
Primordial power spectrum
late time expansion

A_s, Ω_Λ

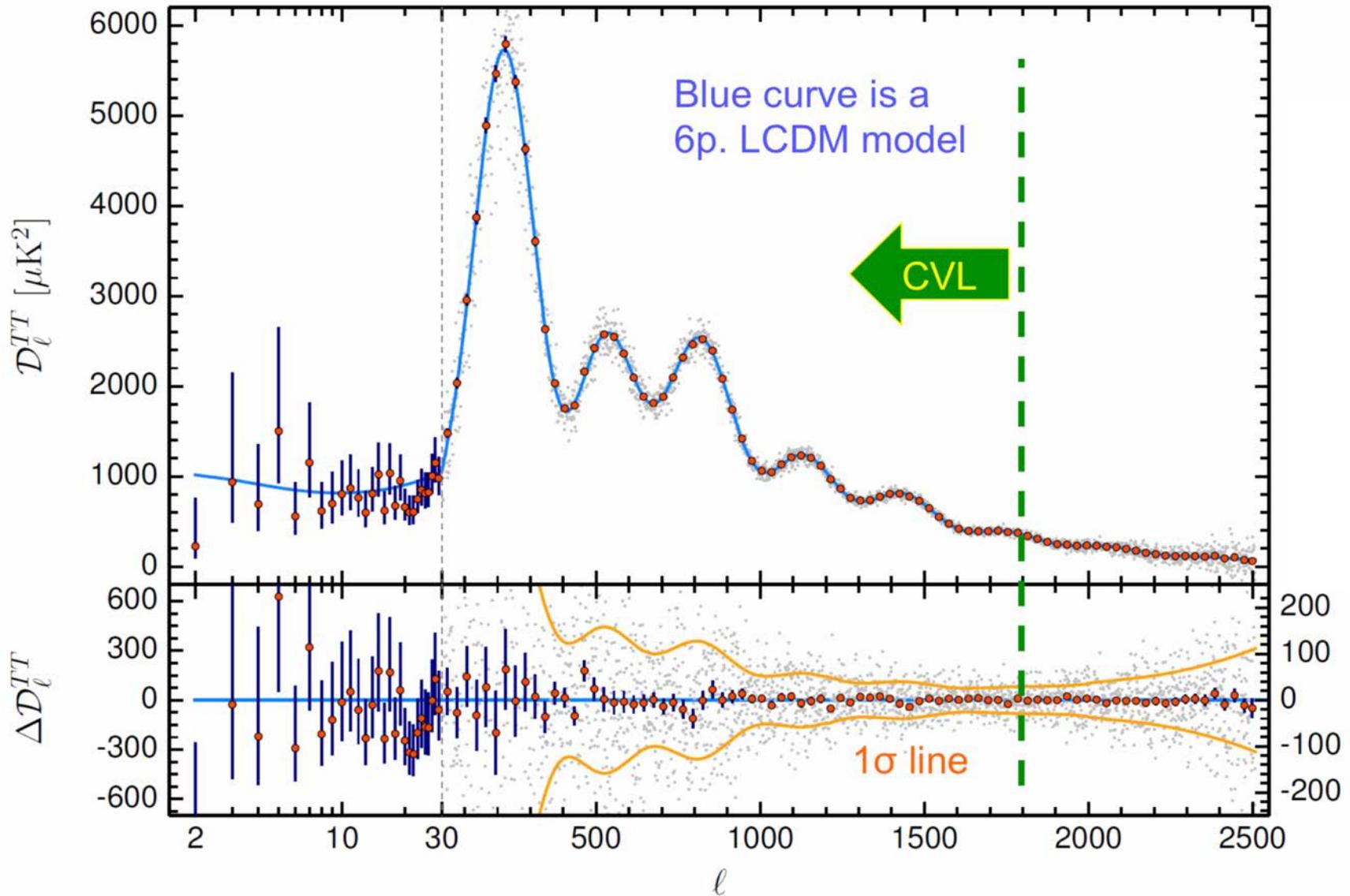
Baryon abundance

Ω_b

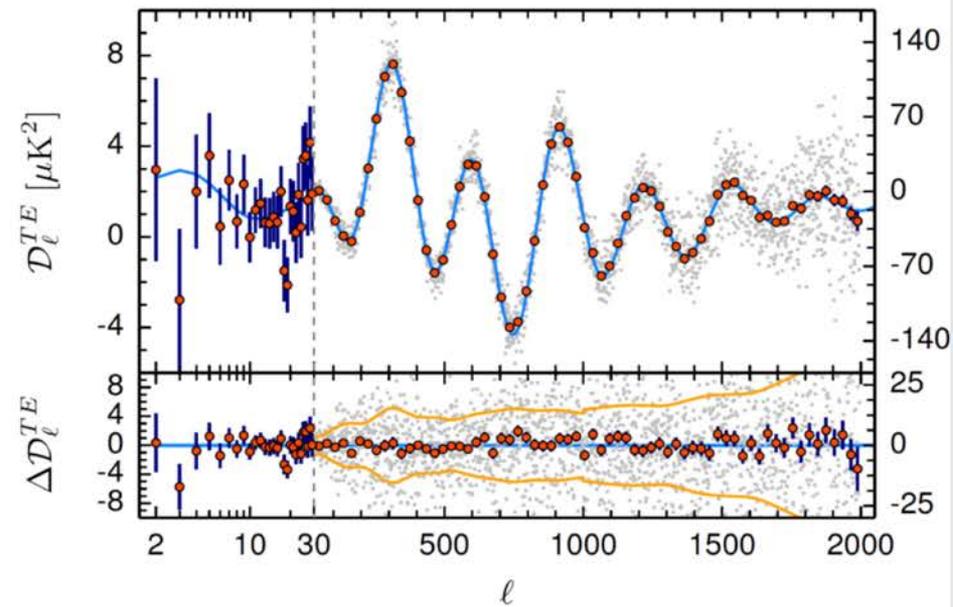
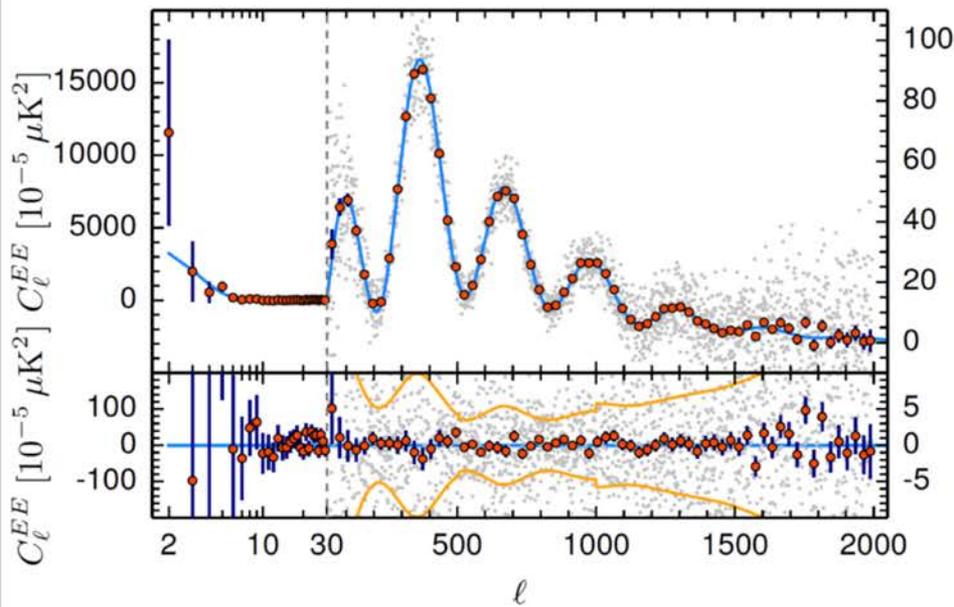
Planck 2018 TT power spectrum



Planck 2018 TT power spectrum

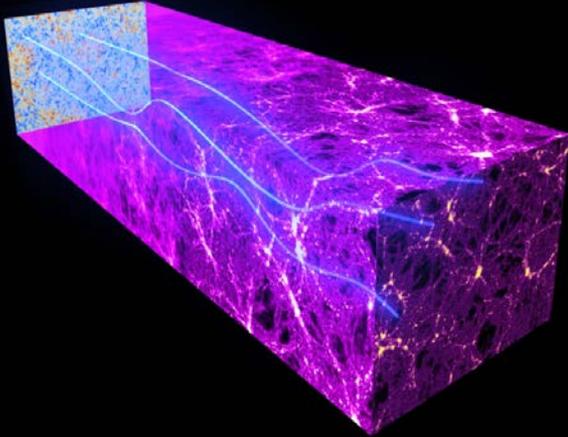


Planck 2018 TE, EE power spectra



Blue line is not a fit, but a prediction given the TT spectrum!

CMB is sensitive to the late-time density field, too....

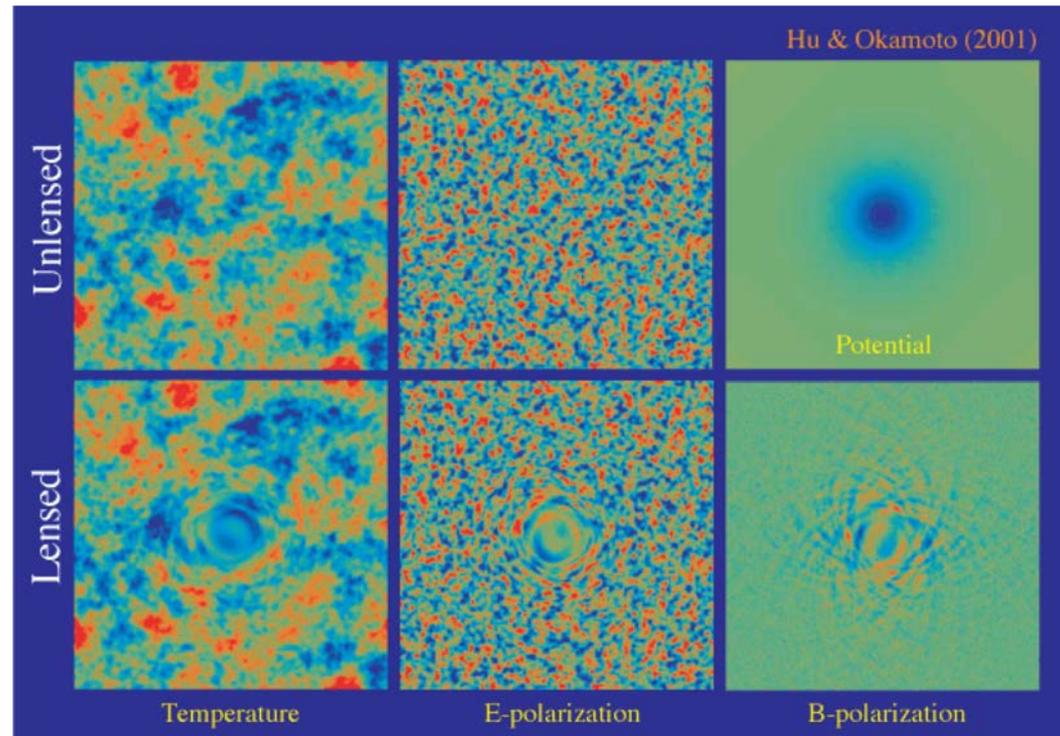


Line-of-sight integral of the gravitational potentials

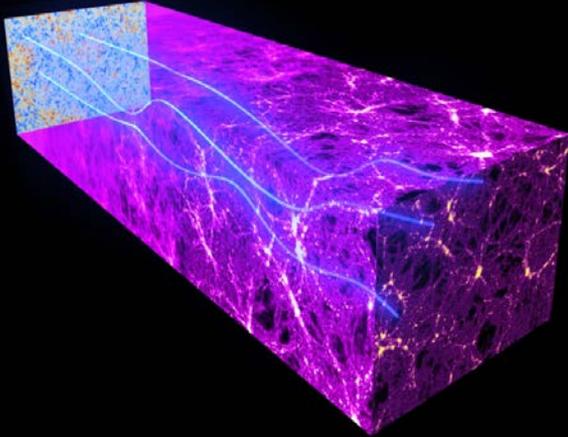
$$\phi(\hat{n}) = - \int_0^{\chi_*} d\chi \frac{\chi_* - \chi}{\chi_* \chi} (\Phi + \Psi)$$

Deflection field

$$\vec{d} = \vec{\nabla} \phi$$



CMB is sensitive to the late-time density field, too....



Deflection field

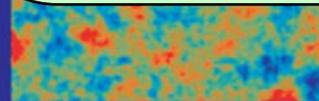
$$\vec{d} = \vec{\nabla} \phi$$

Line-of-sight integral of the gravitational potentials

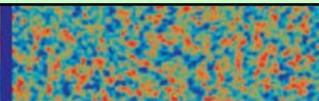
$$\phi(\hat{n}) = - \int_0^{\chi_*} d\chi \frac{\chi_* - \chi}{\chi_* \chi} (\Phi + \Psi)$$

Measures deflection of light due to intervening structures
(average deflection angle is ~ 2.5 arcmin)

Gives integrated information about the matter distribution between us and the last scattering surface



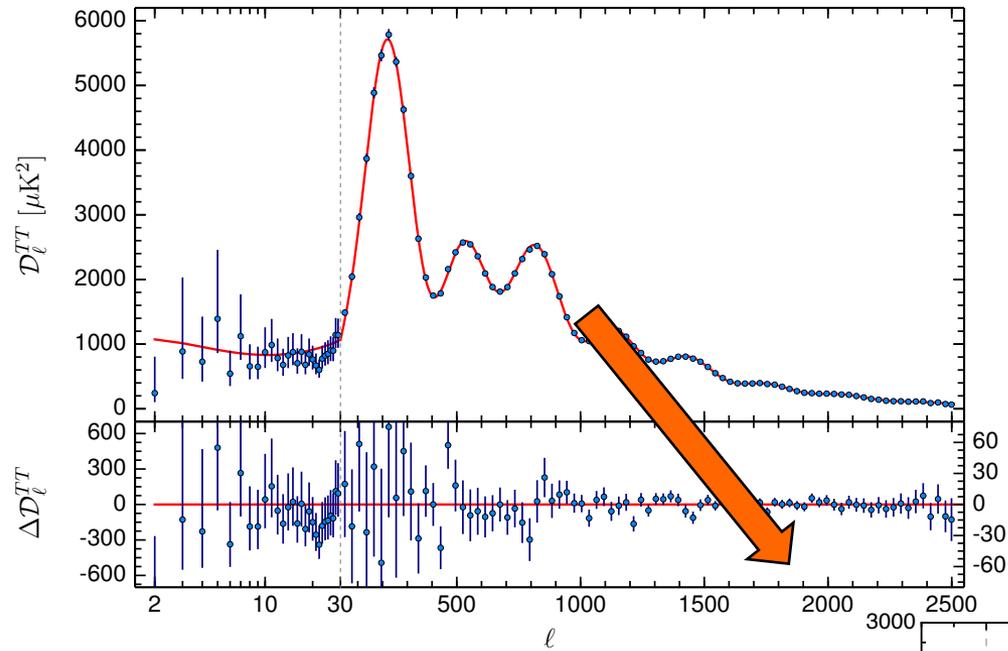
Temperature



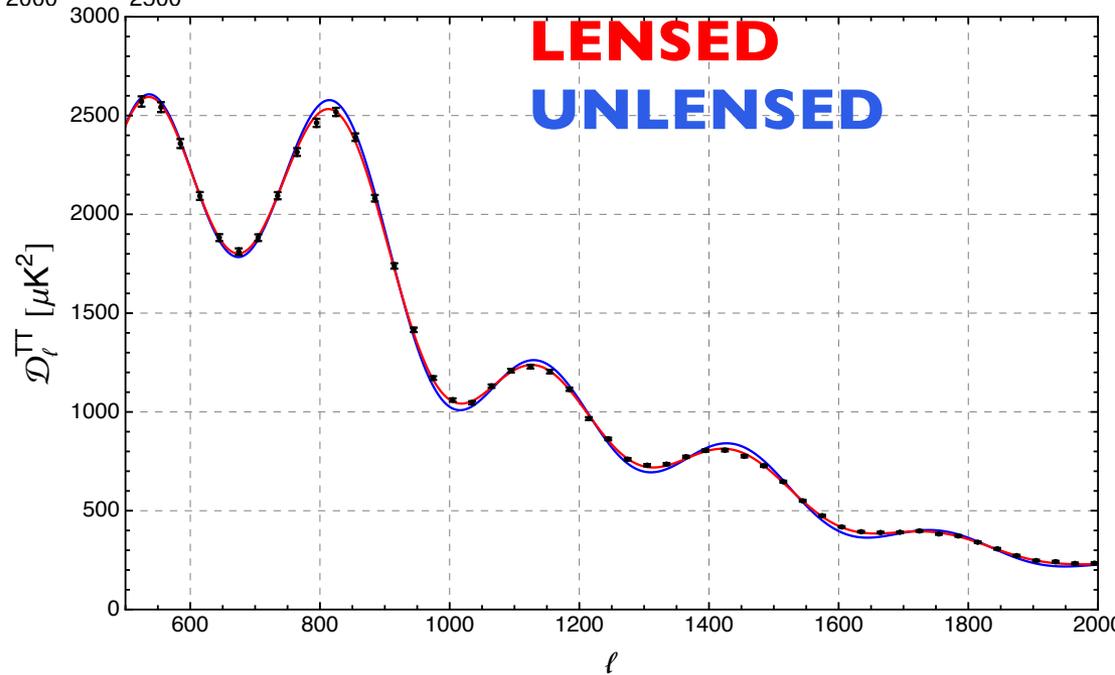
E-polarization



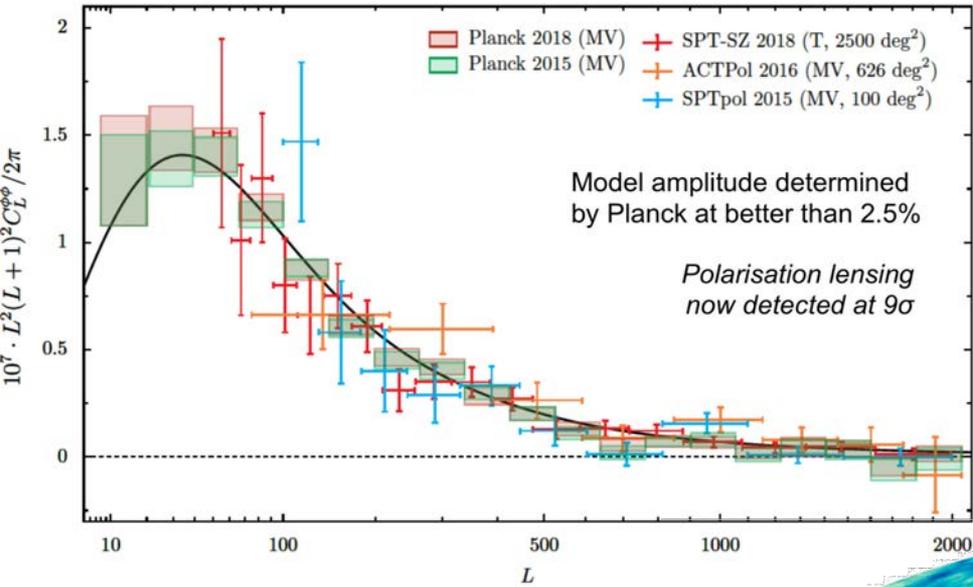
B-polarization



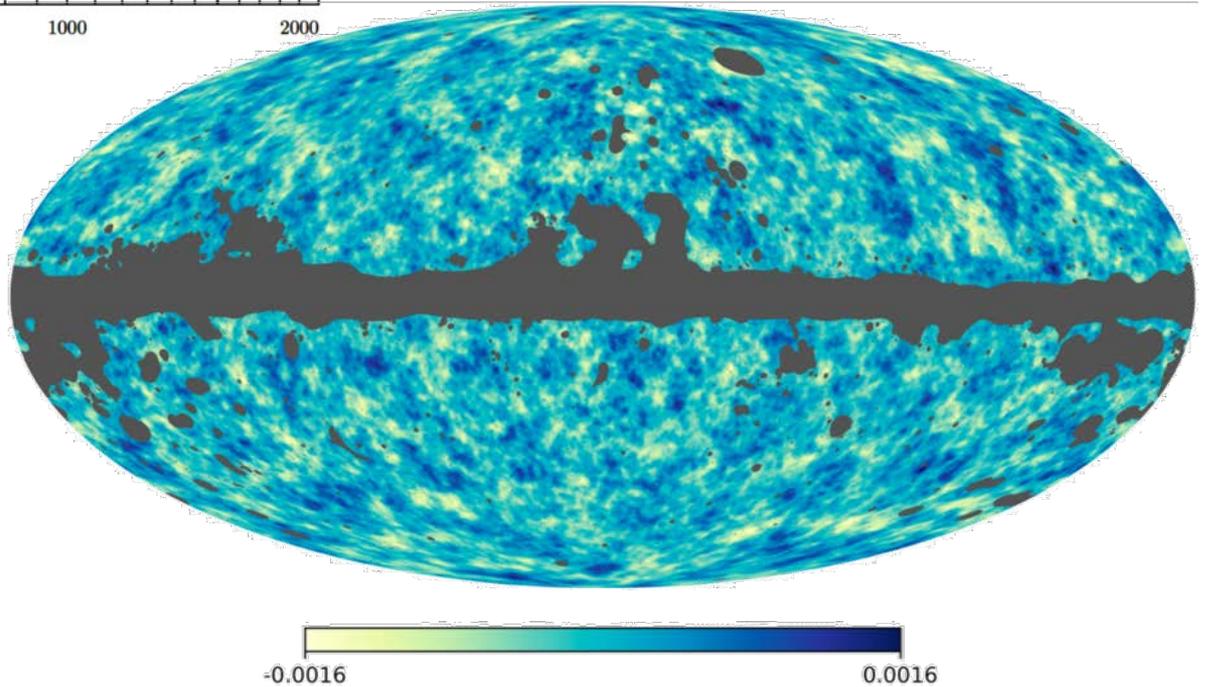
Lensing smooths the peaks
of the CMB power
spectrum...
... and introduces non-
gaussianities in the map
(nonzero 4-point c.f.)



LENSING



Lensing potential estimated from the four-point correlation function



Λ CDM 6 parameter fit

(Planck temperature, polarization and lensing)

		Mean	Stdev	Rel. err.
primary	$\Omega_b h^2$ Baryon density	0.02237	0.00015	0.007
	$\Omega_c h^2$ Dark matter density	0.1200	0.0012	0.01
	100θ CMB acoustic scale	1.04092	0.00031	0.0003
	τ Optical depth to last scattering surface	0.0544	0.0073	0.13
	$\ln(A_s 10^{10})$ Primordial amplitude of perturbation	3.044	0.014	0.007
	n_s Primordial Scalar spectral index	0.9649	0.0042	0.004
derived	H_0 Hubble parameter today	67.36	0.54	0.008
	Ω_m Total matter density	0.3153	0.0073	0.023
	σ_8 Matter perturbation amplitude	0.8111	0.0060	0.007

Λ CDM 6 parameter fit

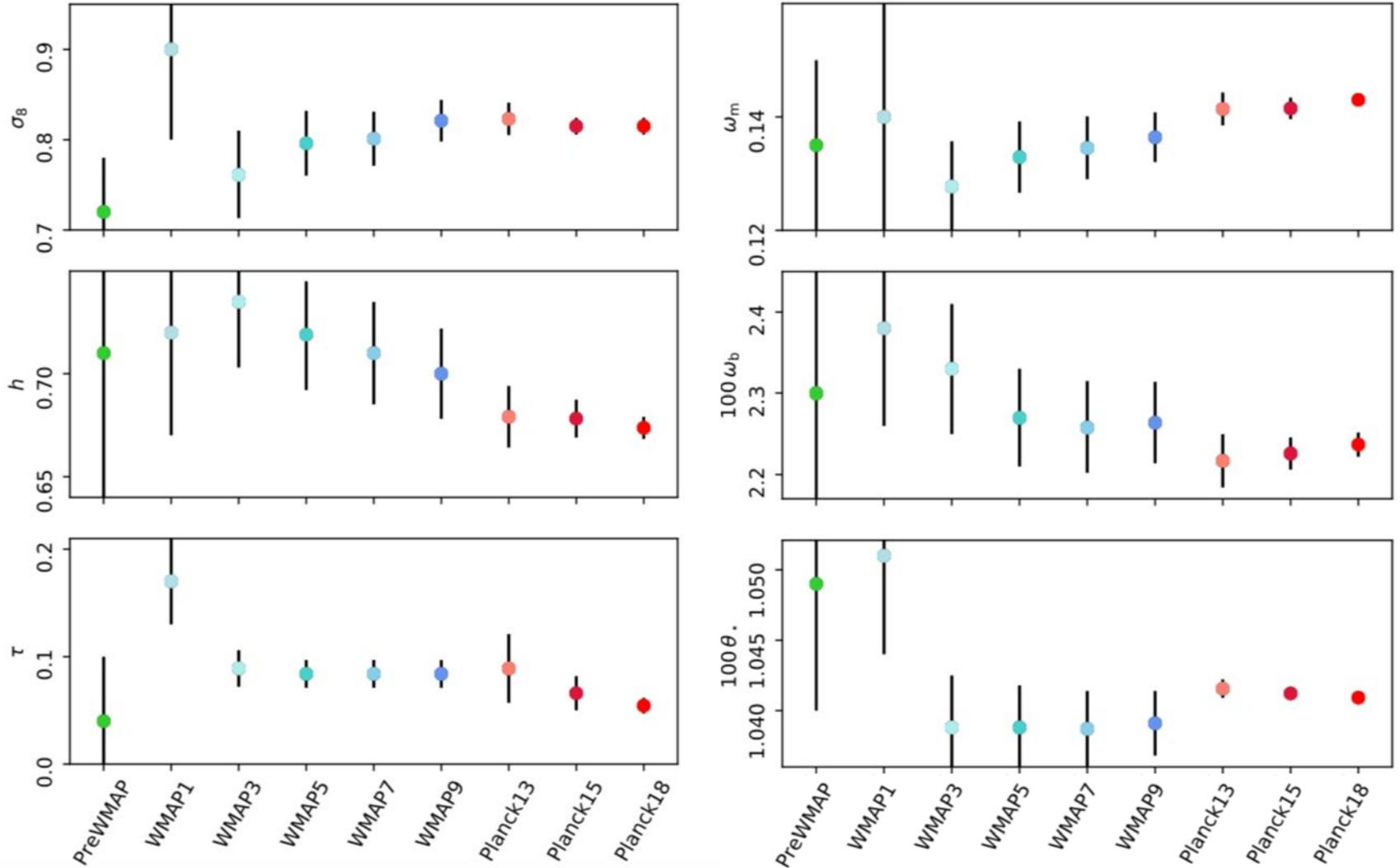
(Planck temperature, polarization and lensing)

Highlights:

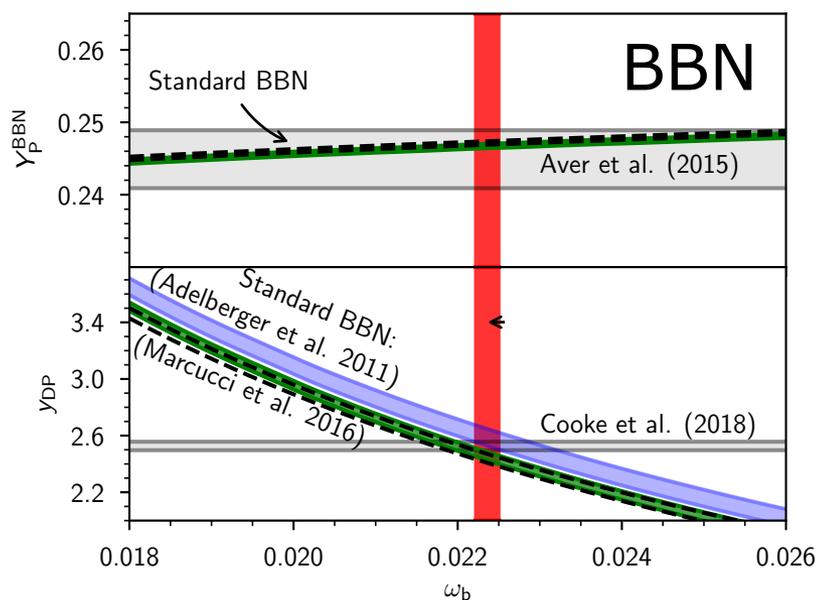
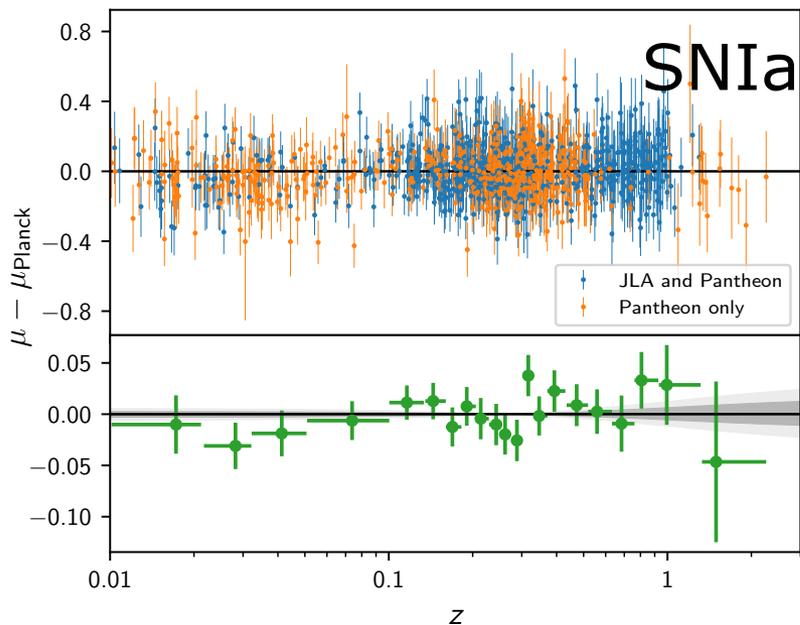
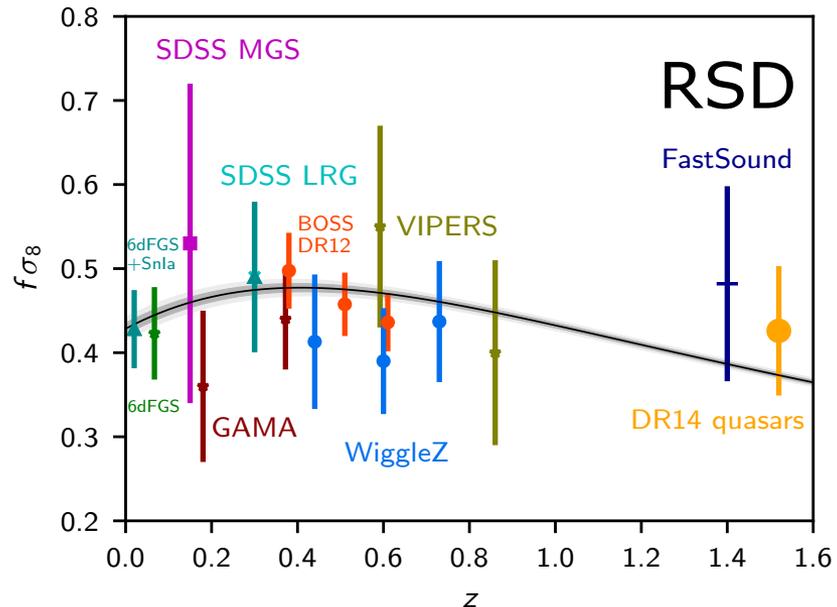
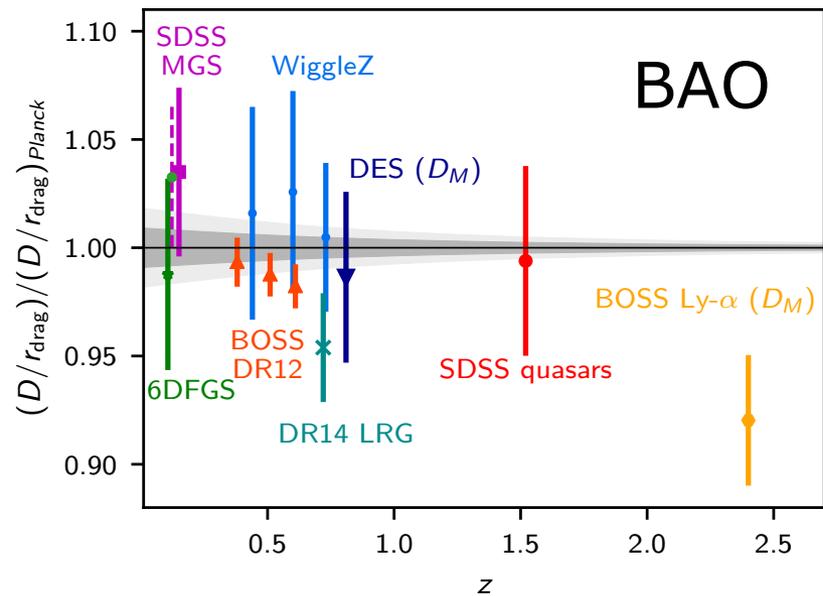
1. Best determination of H_0 to date (indirect, in strong tension with direct measurements)
2. Scalar spectral index is now 8 σ away from 1 (a signature of inflation). Even in extended
3. Optical depth τ greatly improved after taming of large-angle polarization systematics. Still, at 13% relative error, by far the worst parameter determined from CMB

	Mean	Stdev	Rel. err.
$\Omega_b h^2$ Baryon density	0.02237	0.00015	0.007
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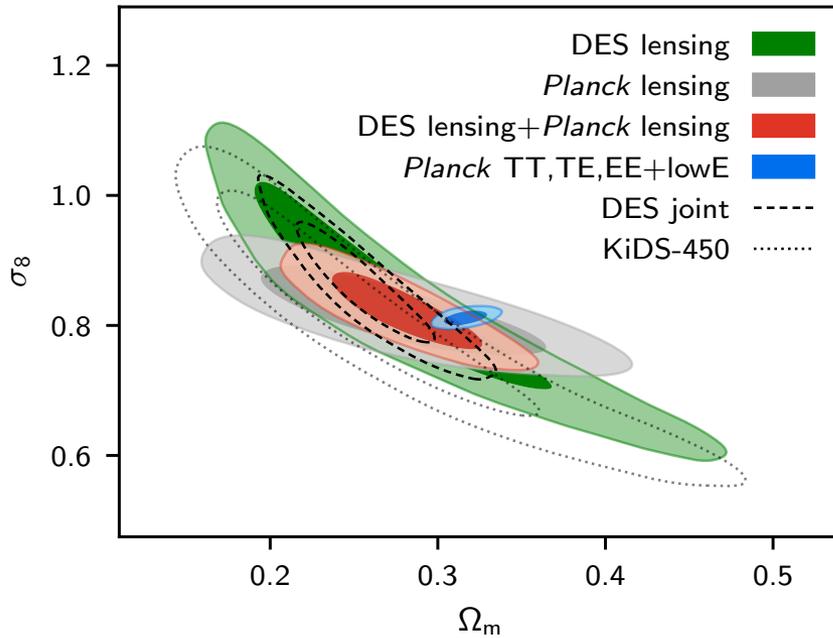
Improvement in parameter accuracy



Consistency with other datasets



Tensions with other datasets...



Mild tension with DES year I results

$$S_8 \equiv \sigma_8 (\Omega_m / 0.3)^{0.5}$$

DES Joint
 $S_8 \equiv 0.792 \pm 0.024$
 $\Omega_m = 0.257^{+0.023}_{-0.031}$

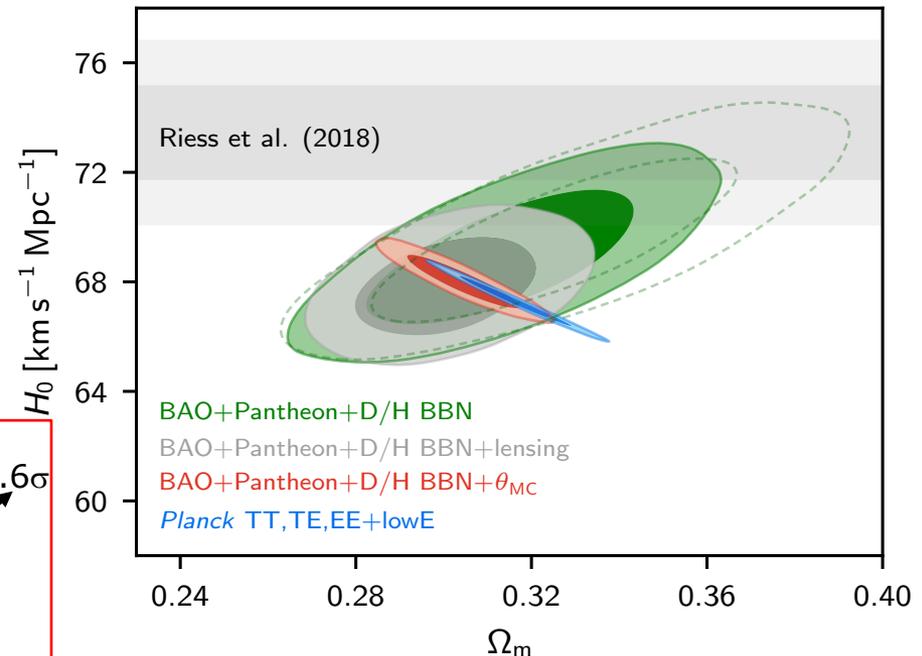
Planck TTTEE+lowE+lensing
 $S_8 = 0.832 \pm 0.013$
 $\Omega_m = 0.315 \pm 0.007$

Strong tension with H_0 distance ladder measurements.

$H_0 = 67.36 \pm 0.54$ km/s/Mpc Planck Λ CDM
 $H_0 = 73.5 \pm 1.6$ km/s/Mpc SH0ES (Riess+ 18)

Inverse distance ladder:
 $H_0 = 67.9 \pm 1.3$ km/s/Mpc BAO+D/H+CMB lensing

3.6 σ

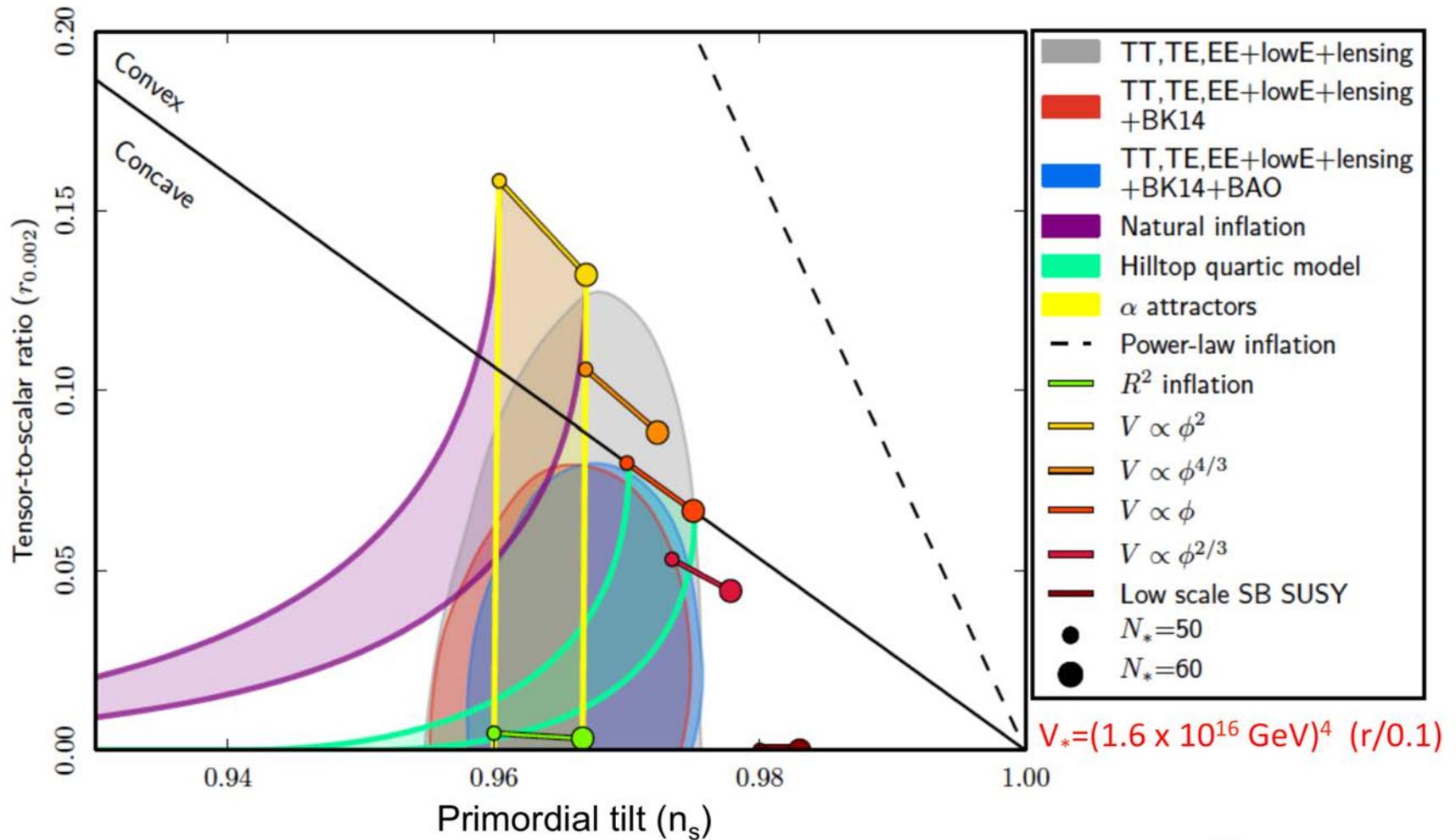


Now > 4 σ , next talk by Martina Gerbino

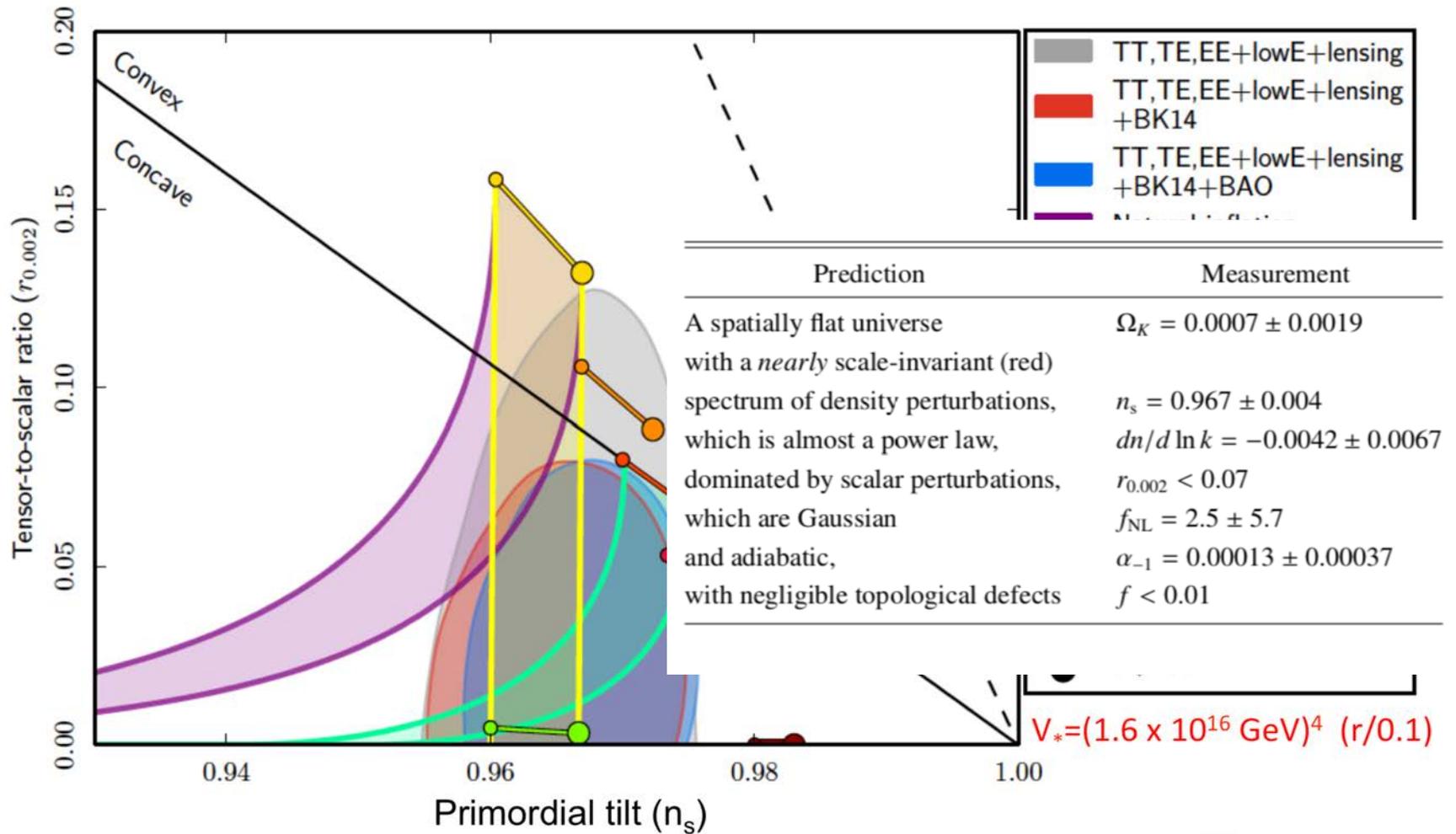
Λ CDM 6 parameter fit + extensions (where surprises might hide)

- Tensor modes, i.e. primordial gravitational waves, $r = A_T/A_s$
- Running spectral index $dn_s/d\ln k$
- Primordial non Gaussianity f_{NL}
- Non adiabatic (isocurvature) primordial perturbations
- Dark energy equation of state, w
- Spatial curvature $\Omega_k = 1 - \Omega_m - \Omega_\Lambda$
- Neutrino masses Σm_ν
- Number of relativistic species N_{eff}
- ...

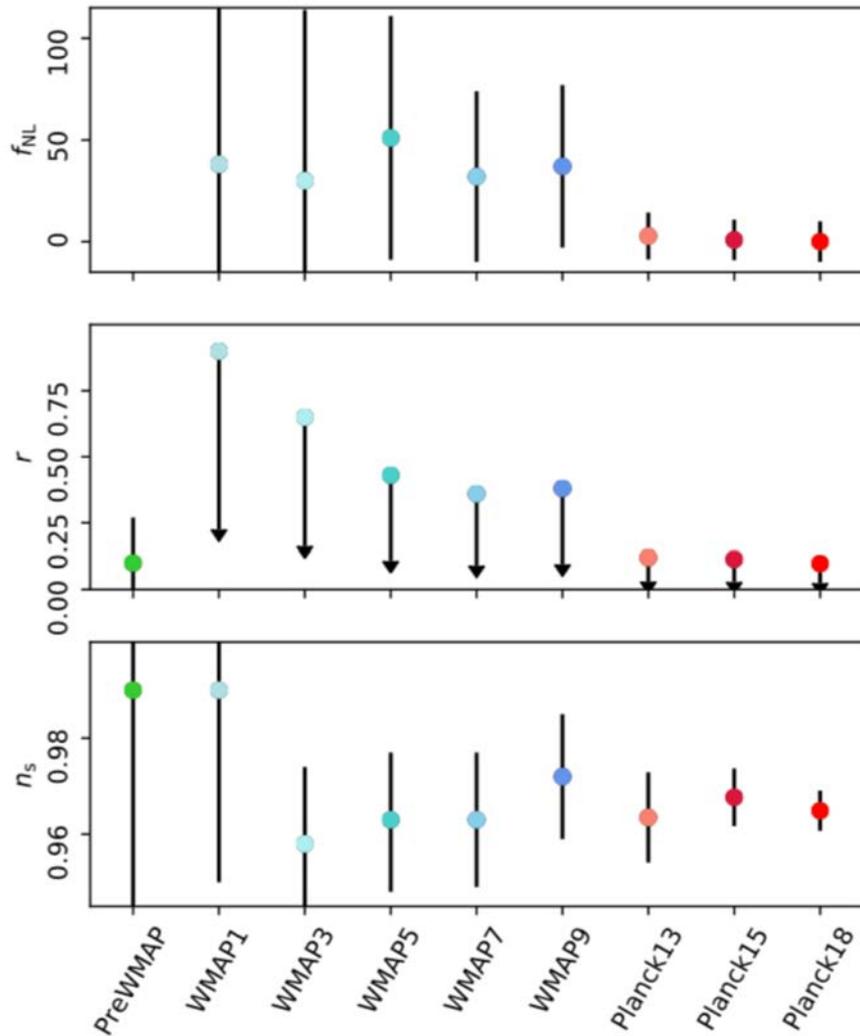
Constraints for tensor perturbations



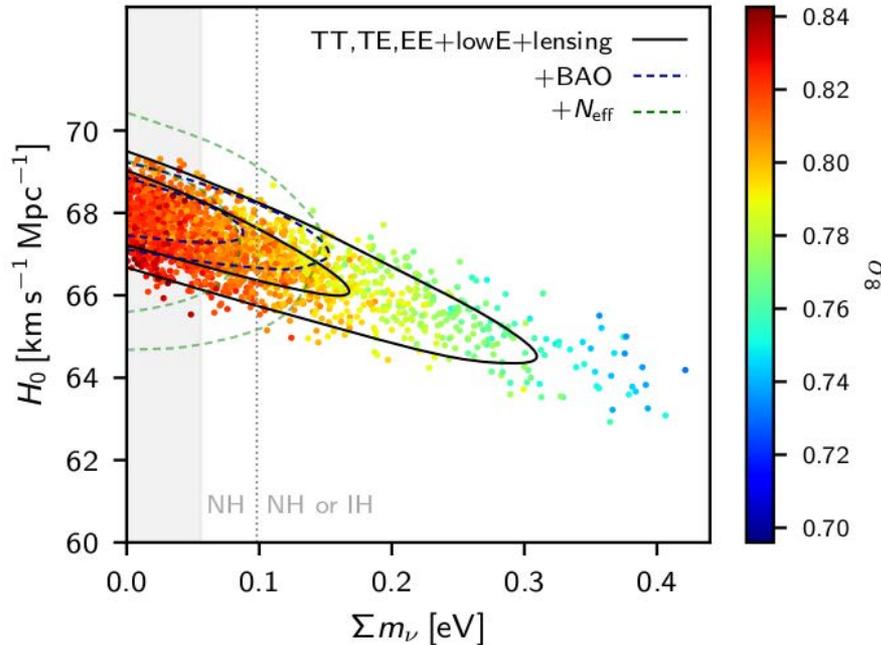
Constraints for tensor perturbations



Improvement in inflationary parameters



Neutrino legacy of Planck: Σm_ν

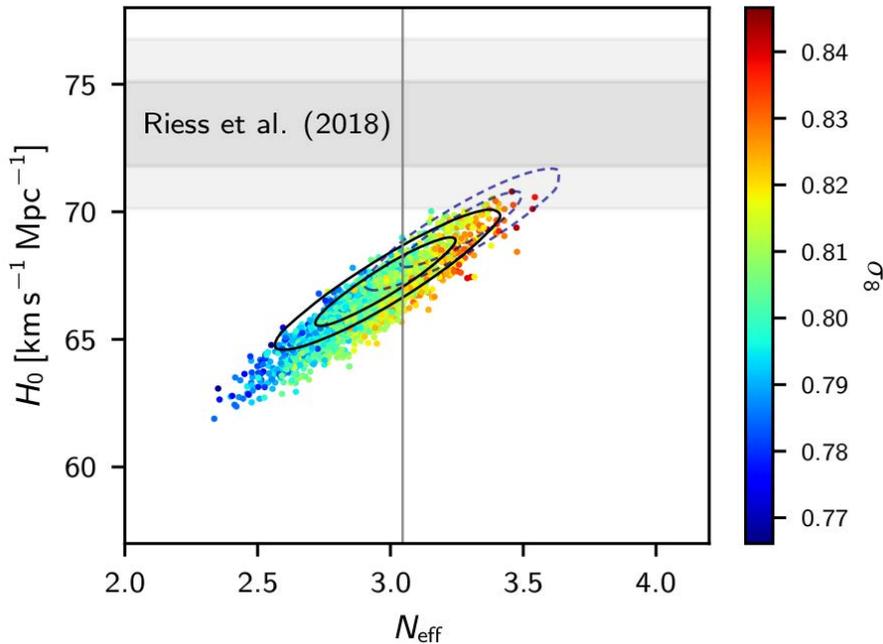


- Tightest constraint from a single experiment
- First constraint exploiting the information encoded in the CMB weak lensing
- One order of magnitude better than present kinematic constraints, already at the same level than future expectations for KATRIN
- The combined limits from Planck and large scale structure probes are starting to corner the inverted hierarchy scenario

$m_\nu < 0.44 \text{ eV}$ (95%CL, TT + lowE + lensing)

$m_\nu < 0.13 \text{ eV}$ (95% CL, TT+lowE+lensing+BAO)

Neutrino legacy of Planck: N_{eff}



$$N_{eff} = 3.00^{+0.57}_{-0.53} \quad (95\% \text{ CL, TT+lowE})$$

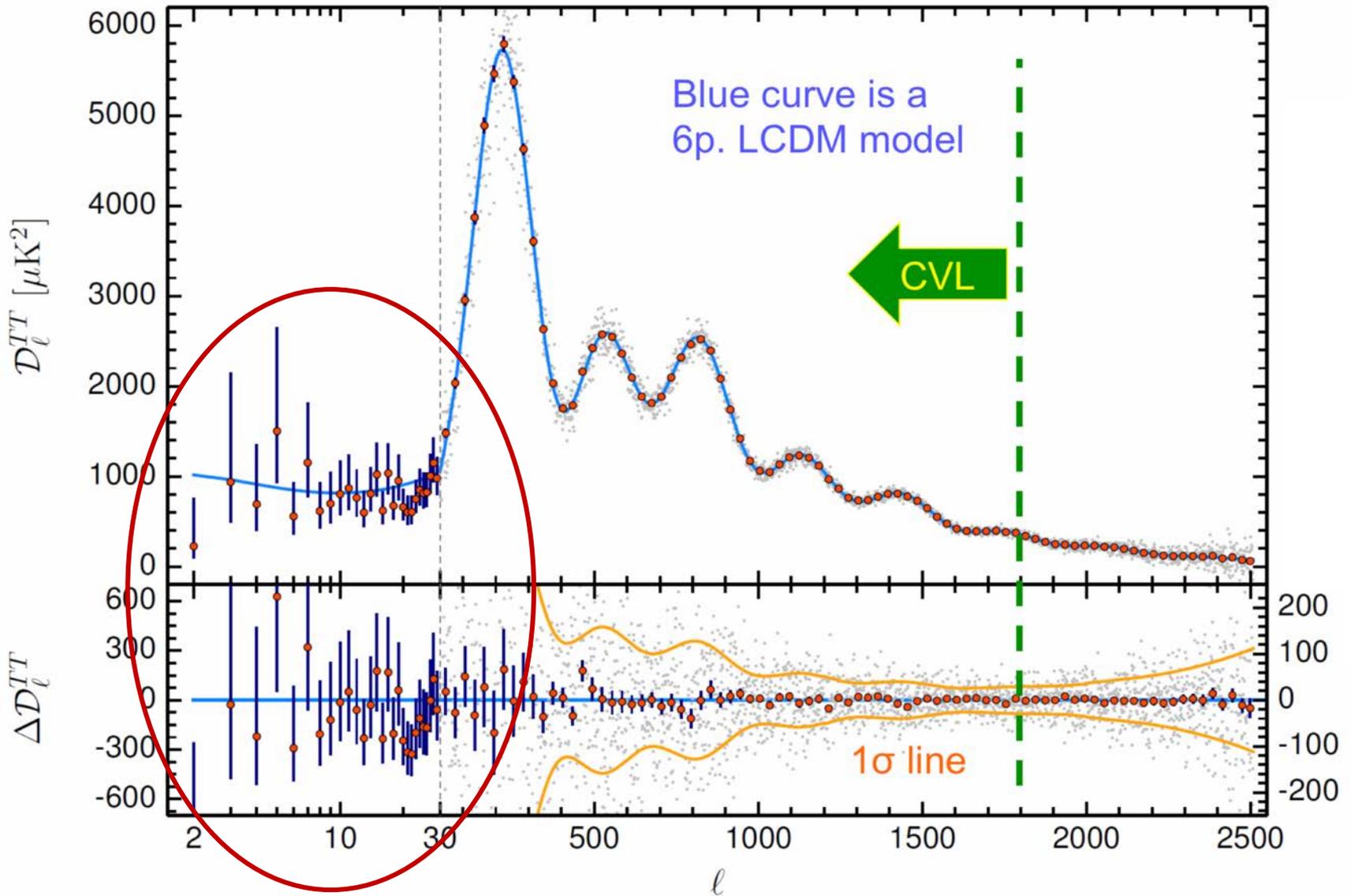
$$N_{eff} = 3.11^{+0.44}_{-0.43} \quad (95\% \text{ CL, TT+lowE+lensing+BAO})$$

- Effective number of relativistic species is consistent with the standard expectation $N_{eff} = 3.046$
- Data are consistent with these relativistic species behaving as free-streaming neutrinos – a strong indication that they are indeed the SM neutrinos!
- A fourth thermalized species ($N_{eff}=4$) is excluded at 3.5 to 6 σ , depending on the dataset
- A light sterile neutrino species is allowed if not thermalized. Still, the sterile neutrino interpretation of the short-baseline anomalies is excluded by Planck

Anomalies in the CMB field

- At large angles, the CMB field is known to exhibit anomalies:
 - Lack of power
 - Hemispherical asymmetry
 - Even-odd asymmetry
 - And others...
- For temperature, Planck has reached cosmic variance. For polarization, there is much room for improvement.

Planck 2018 TT power spectrum



Scale-invariance of the large-scale perturbations is a prediction of single-field, slow-roll inflation.

Transition from a pre-inflationary “fast-roll” phase to slow-roll would suppress power in the primordial spectrum.

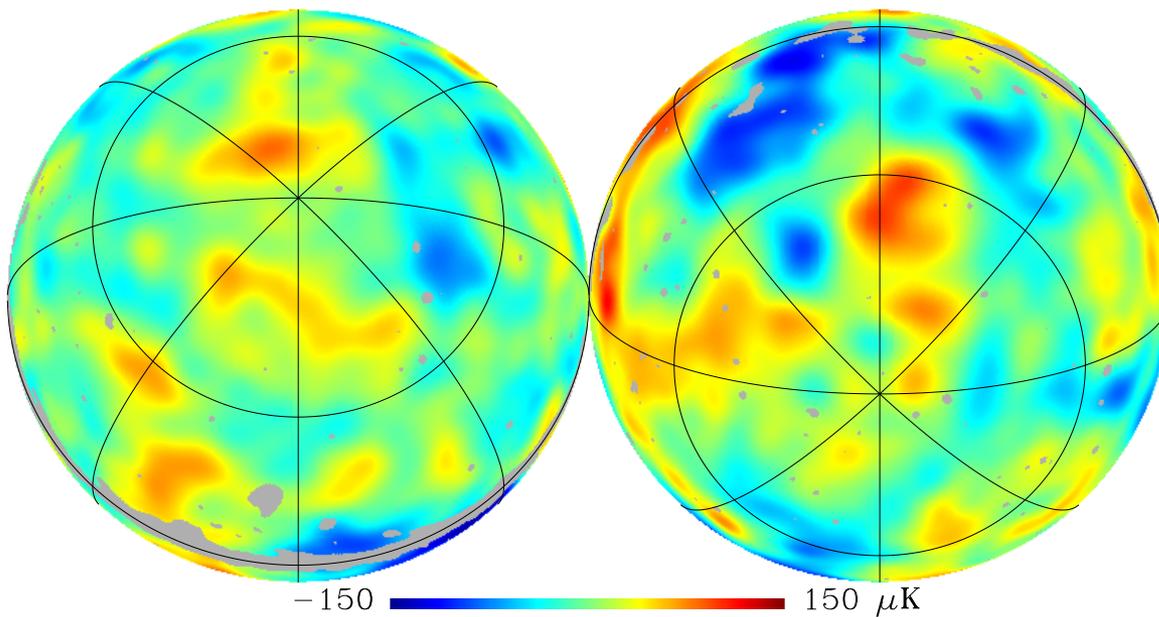
Are we seeing relics of a decelerating inflaton?

See e.g. Contaldi, Peloso, Kofman, Linde (2003); Destri, de Vega, Sanchez (2010); Dudas, Kitazawa, Patil, Sagnotti (2012); Kitazawa, Sagnotti (2014)

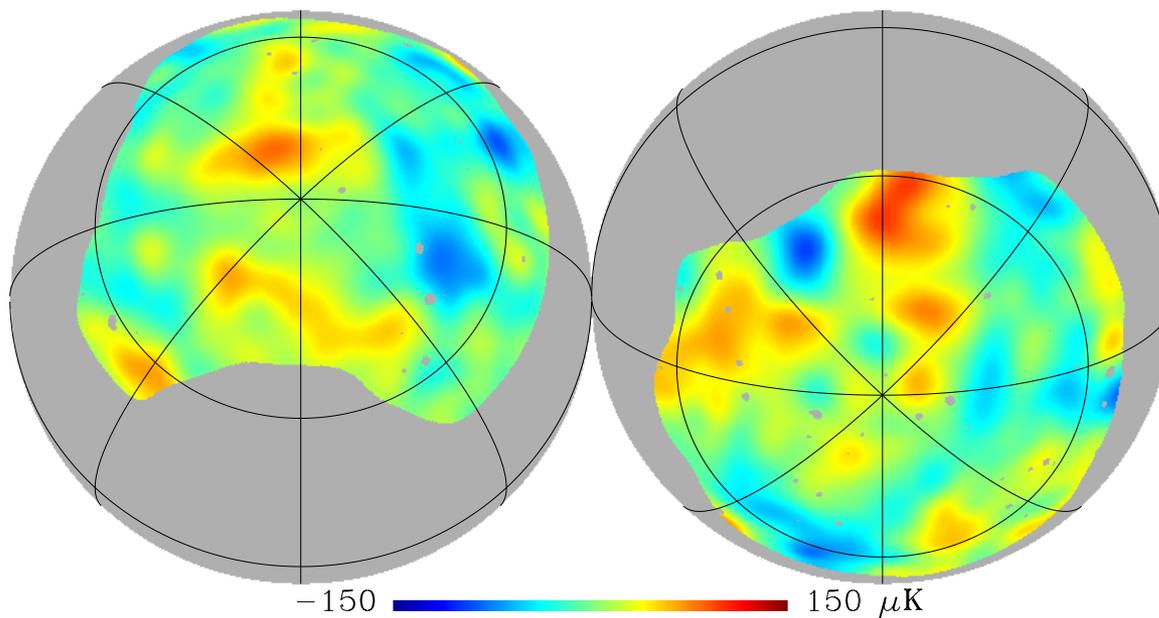
$$P(k) \sim \frac{k^3}{[k^2 + \Delta^2]^{2 - \frac{n_s}{2}}}$$

↙
~ scale that enters the horizon
at the onset of slow roll

standard mask

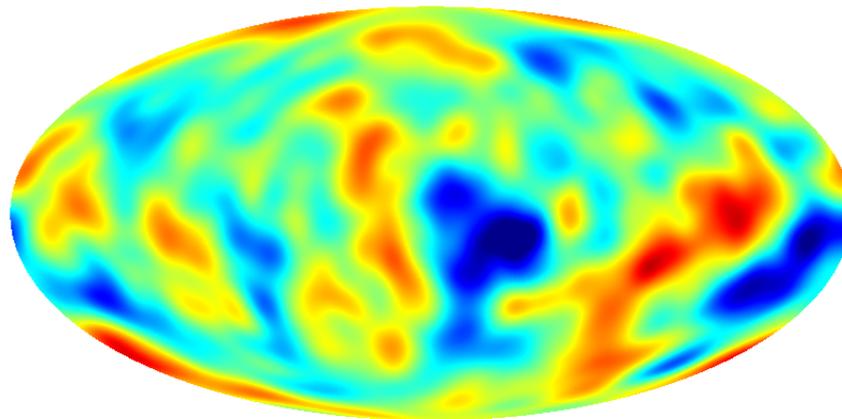


extended mask



A. Gruppuso, N.
Kitazawa, N.
Mandolesi, PN, A.
Sagnotti 2017

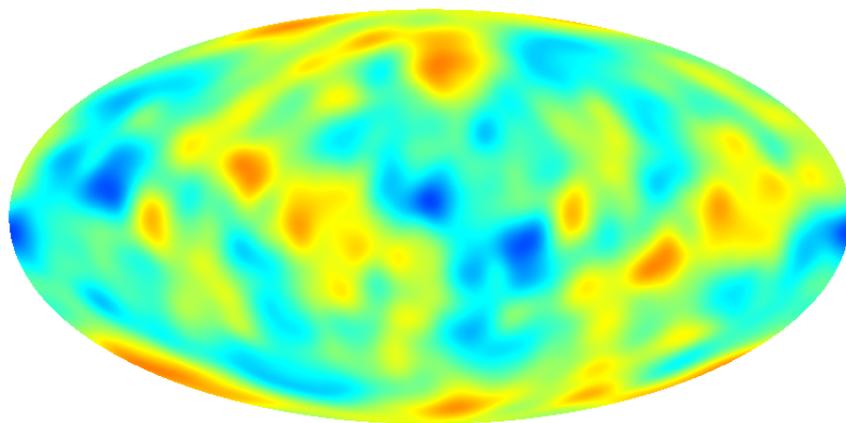
even+odd



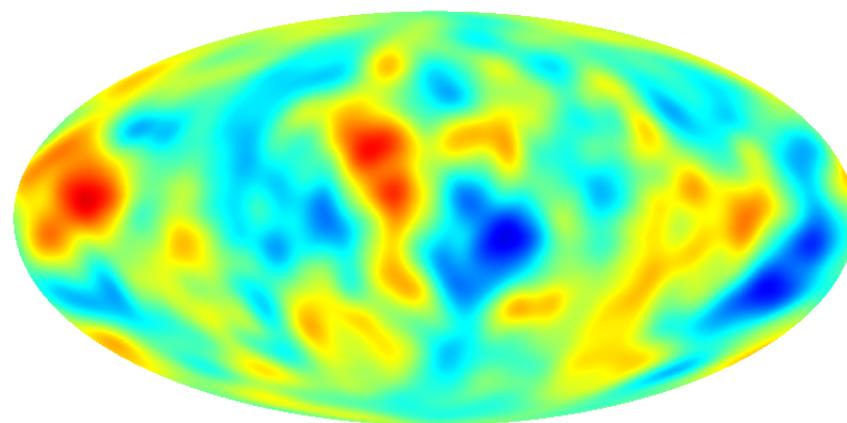
Planck 2015 data

-100.0 100.0 μK

even

-100.0 100.0 μK

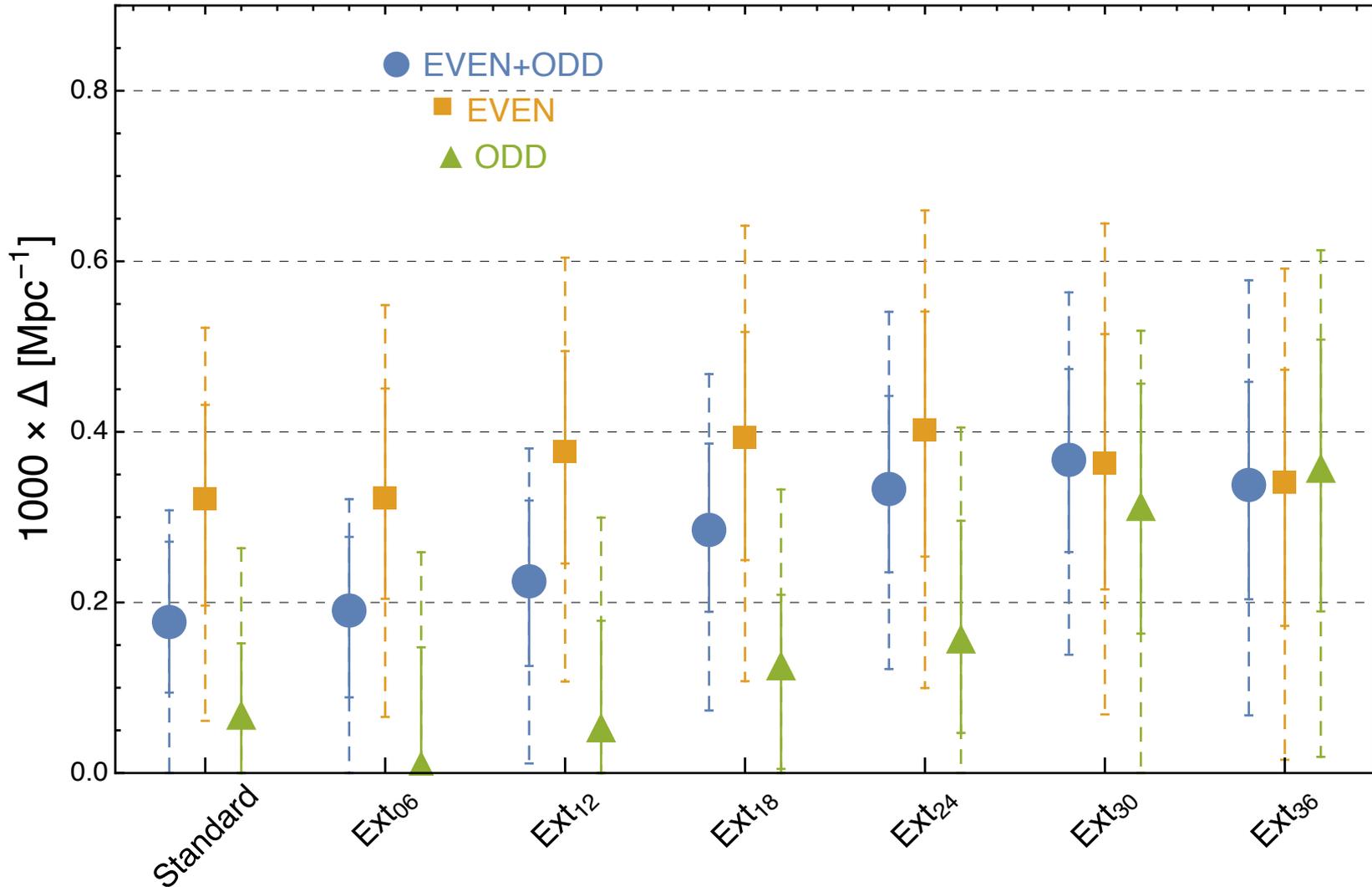
odd

-100.0 100.0 μK

$$\left(\frac{\delta T}{T}(\hat{n}) \right)_{\pm} \equiv \frac{1}{2} \left[\frac{\delta T}{T}(\hat{n}) \pm \frac{\delta T}{T}(-\hat{n}) \right]$$

A. Gruppuso, N. Kitazawa, M. Lattanzi, N. Mandolesi, PN, A. Sagnotti 2017

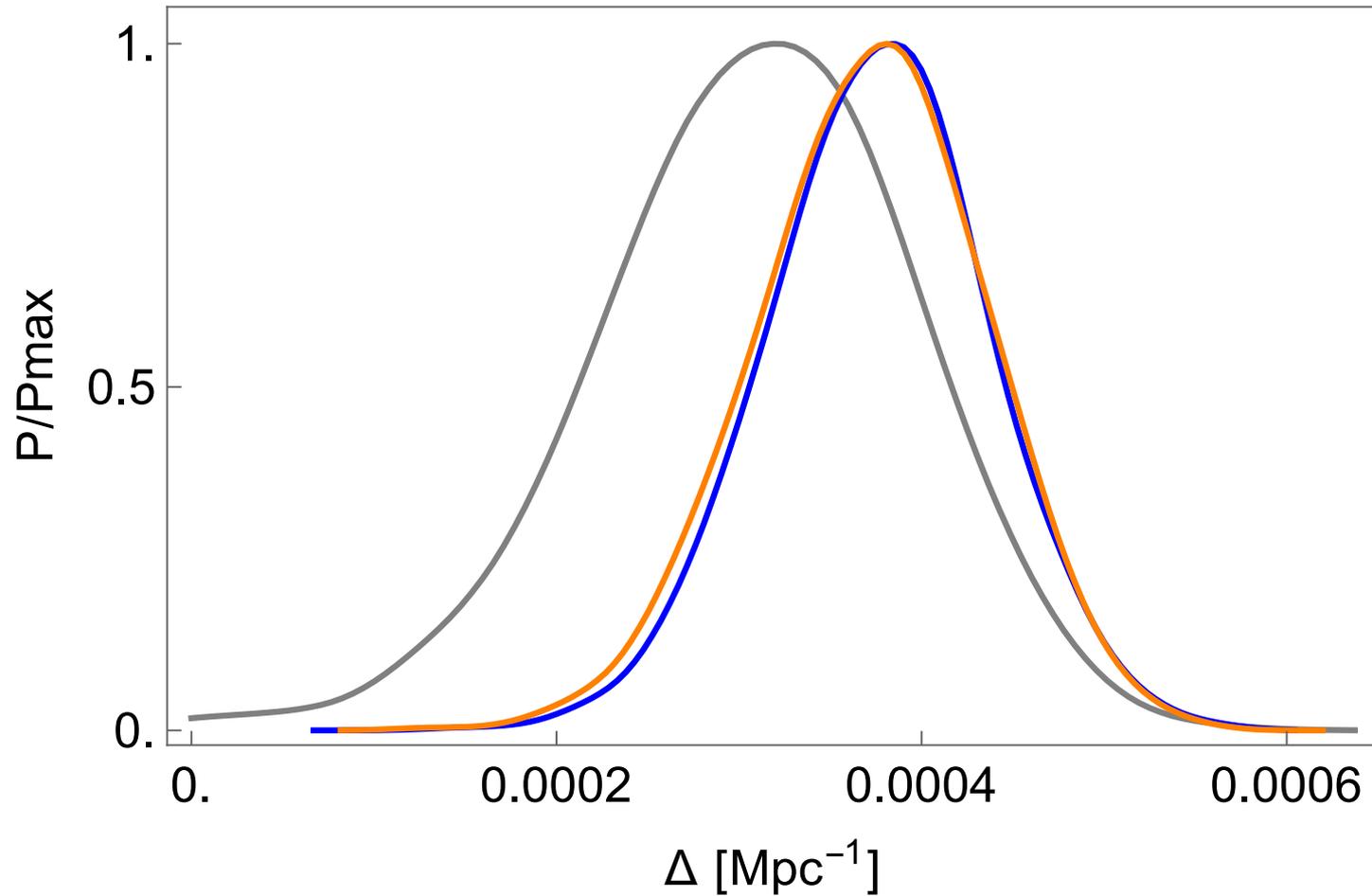
Constraints on Δ from Planck 2015



A. Gruppuso, N. Kitazawa, M. Lattanzi, N. Mandolesi,
PN, A. Sagnotti 2017

- The even multipoles are consistently lower than the Λ CDM expectation, independently on the galactic masking
- The odd multipoles are consistent with the Λ CDM expectation for the smaller masks (more sky). In larger masks (less sky), they are consistent with the even multipoles (and then have low power)
- The power at large scales is concentrated around the galactic plane, in the odd multipoles

Forecasted constraints on Δ from future experiments



- Grey:** Planck-like noise, standard masking
- Orange:** CVL large-scale pol., ext30 mask
- Blue:** CVL large-scale pol., full sky

PRESENT AND FORTHCOMING CMB PROBES

Ground



POLARBEAR



ACTPol

Atacama,
Chile

In addition,
ABS, CLASS, POLARBEAR-2,
Simons Array, Adv-ACTPol, ...



BICEP1 BICEP2 DASI QUAD KECK
SPTPol

South
Pole

In addition, BICEP3, POLAR, QUBIC, ...

Data SIO, NOAA, U.S. Navy, NGA, GEBCO
© 2011 Inev/Geosistemas SRL

Balloon

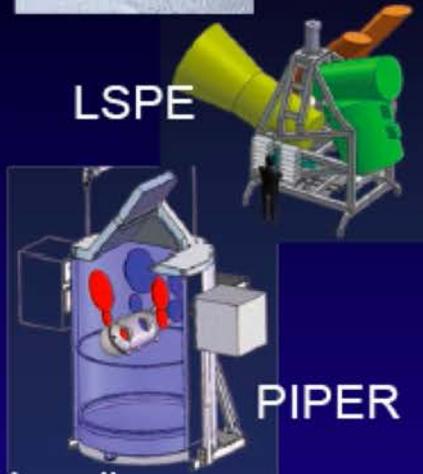


EBEX



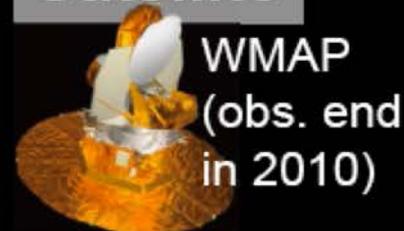
SPIDER

LSPE



PIPER

Satellite



WMAP
(obs. end
in 2010)



Planck



LiteBIRD



PIXIE



CoRE+

In addition, QUIJOTE in Canary island, AMiBA in Hawaii

PRESENT AND FORTHCOMING CMB PROBES

Ground



POLARBEAR



ACTPol

Atacama, Chile

In addition, ABS, CLASS, POLARBEAR-2, Simons Array, Adv-ACTPol, ...

and many more!



South Pole

In addition, BICEP3, POLAR, QUBIC, ...

Data SIO, NOAA, U.S. Navy, NGA, GEBCO
© 2011 Inev/Geosistemas SRL

Balloon



EBEX



SPIDER



LSPE

SWIPE

PIPER

Satellite



WMAP
(obs. end in 2010)



Planck



LiteBIRD



PIXIE



CORE+

**Retired
(legacy remains)**

**Not funded!
Not funded!
(India interested)**

In addition, QUIJOTE in Canary island, AMiBA in Hawaii

Conclusions

- Planck has delivered its final (legacy) release
- It has provided the ultimate (cosmic variance limited) measurement of CMB anisotropy
- ... But just opened the door of CMB polarization (which was never designed to measure, by the way)
- It has fulfilled its promise of measuring the fundamental cosmological parameters to percent accuracy
- And brought remarkable constraints on particle physics parameters as well, excluding a fourth fully thermalized neutrino and constraining the total neutrino masses in the 100 meV range.
- Has measured well one relevant inflationary parameter, the primordial spectral index, allowing constraints on the inflationary paradigm
- Yet has uncovered several tensions with astrophysical measurements, which may or may not hint at new physics.
- Intrinsic anomalies do exist in the large-angle CMB field, which may also be a tracer of something new.
- If these tension/anomalies are really hinting at new physics, its signature in the CMB is scant. Accurate measurements are needed to pin down the issue.
- Primordial gravitational waves remain unseen.
- To exploit the wealth of information that still is in the CMB, we need to cope with the extraordinary complexity of the sky. This can be credibly done only with a future space mission.

The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



planck



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy)

with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.



DTU Space
National Space Institute



National Research Council of Italy



X COSPAR 2018, July 2018

