Nuclear physics of dark matter direct detection



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MH, Klos, Menéndez, Schwenk PLB 746 (2015) 410, PRL 119 (2017) 181803, PRD 94 (2016) 063505, 99 (2019) 055031
 Fieguth, MH, Klos, Menéndez, Schwenk, Weinheimer PRD 97 (2018) 103532
 XENON collaboration + MH, Klos, Menéndez, Schwenk PRL 122 (2019) 071301



Direct detection of dark matter

Rate for WIMP-nucleus scattering: factorization

$$\frac{\mathrm{d}R}{\mathrm{d}E_{\mathrm{r}}} = \underbrace{\frac{\sigma_{\chi N}^{\mathrm{SI}}}{m_{\chi}\mu_{N}^{2}}}_{\mathrm{particle + hadronic physics}} \times \underbrace{\left|\mathcal{F}_{+}^{\textit{M}}(q^{2})\right|^{2}}_{\mathrm{nuclear physics}} \times \underbrace{\rho_{0} \int_{v_{\mathrm{min}}}^{v_{\mathrm{esc}}} \frac{f(\mathbf{v},t)}{v} \mathrm{d}^{3}v}_{\mathrm{astrophysics}}$$

- This talk:
 - Nuclear responses: $\mathcal{F}_{+}^{M}(q^2), \ldots$
 - Hadronic matrix elements: relating $\sigma_{\gamma N}^{SI}$ to BSM operators, ...
 - Uncertainties: known and unknown unknowns

Direct detection of dark matter

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- This talk:
 - Nuclear responses: $\mathcal{F}_{+}^{M}(q^{2}), \ldots$
 - Hadronic matrix elements: relating $\sigma_{\gamma N}^{SI}$ to BSM operators, ...
 - Uncertainties: known and unknown unknowns
- Ultimate aim: controlled approximations at all stages
 - Effective field theories (EFTs) to connect scales
 - EFTs + many-body methods for nuclear structure
 - EFTs/lattice QCD/dispersion relations for hadronic matrix elements
 - → Trace all of this back to QCD as far as possible "ab initio"



Direct detection of dark matter: scales

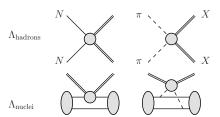
DESM Scale Λ_{BSM} : \mathcal{L}_{BSM}

- **Effective Operators**: $\mathcal{L}_{\text{SM}} + \sum\limits_{i,k} \frac{1}{\Lambda_{\text{BSM}}^{i}} \mathcal{O}_{i,k}$
- Integrate out EW physics

- **Nuclear scale**: $\langle \mathcal{N} | H_l | \mathcal{N} \rangle$
 - $\hookrightarrow \text{nuclear wave function } \text{chiral EFT, NREFT}$

 $\Lambda_{\rm EW}$

Direct detection of dark matter: scales



- Hadronic scale: nucleons and pions
 offective interaction Hamiltonian A
 - \hookrightarrow effective interaction Hamiltonian H_I Chiral EFT
- **Nuclear scale**: $\langle \mathcal{N} | \mathcal{H}_l | \mathcal{N} \rangle$
- → nuclear wave function Chiral EFT, NREFT
- Typical WIMP-nucleon momentum transfer

$$|\mathbf{q}_{\mathrm{max}}| = 2\mu_{\mathcal{N}\chi}|\mathbf{v}_{\mathrm{rel}}| \sim 200\,\mathrm{MeV} \qquad |\mathbf{v}_{\mathrm{rel}}| \sim 10^{-3} \qquad \mu_{\mathcal{N}\chi} \sim 100\,\mathrm{GeV}$$

- Chiral EFT: pions, nucleons, and WIMPs as degrees of freedom
 Prézeau et al. 2003, Cirigliano et al. 2012, Menéndez et al. 2012, Klos et al. 2013, MH et al. 2015, Bishara et al. 2017 ...
- NREFT: all degrees of freedom integrated out but nucleons and WIMPs
 Fan et al. 2010, Fitzpatrick et al. 2012, Anand et al. 2013...

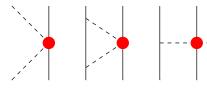


Chiral EFT: a modern approach to nuclear forces

- Traditionally: meson-exchange potentials
- Chiral effective field theory
 - Based on chiral symmetry of QCD
 - Expansion in M_{π}/Λ_{χ} , $\Lambda_{\chi}\sim 600\,\mathrm{MeV}$
 - Low-energy constants
 - Hierarchy of multi-nucleon forces
 - Consistency of NN and 3N
 - \hookrightarrow modern theory of nuclear forces
- Long-range part related to pion-nucleon scattering

	2N force	3N force	4N force		
LO	X 	_	—		
NLO	XHHM	_	_		
N²LO	취석	HH HX X	—		
N³LO	X4444	国 洲 IX			

Figure taken from 1011.1343

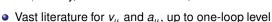


Chiral EFT: currents

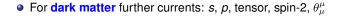


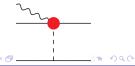
- Coupling to external sources $\mathcal{L}(v_{\mu}, a_{\mu}, s, p)$
- Same LECs appear in axial current

 $\hookrightarrow \beta$ decay, neutrino interactions, dark matter

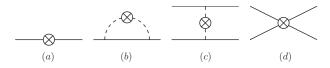


- With unitary transformations: Kölling et al. 2009, 2011, Krebs et al. 2016, 2019
- Without unitary transformations: Park et al. 2003, Pastore et al. 2008,
 Baroni et al. 2015



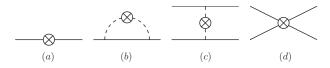


Chiral EFT for dark matter: diagrams



- One-body (1b) and two-body (2b) diagrams:
 - Leading 1b responses (a): standard spin-independent (SI) and spin-dependent (SD) interactions
 - Subleading 1b responses (a): non-relativistic expansion of nucleon form factors
 - Radius corrections (b): q^2 expansion of nucleon form factors
 - Two-body currents (c), (d): pion exchange, contact operators

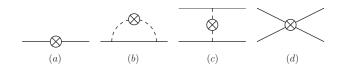
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- Nuclear responses:
 - SI: coherent sum over all A nucleons ("M- response")
 - SD: proportional to $\langle S \rangle$ (combination of Σ' , Σ'')
 - Spin-orbit: coherent sum over partially filled shells (Φ")
 - Two-body currents: coherent over all A, but suppressed in chiral counting



Chiral EFT vs. NREFT

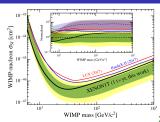


NREFT

- Integrate out the pion, expand in q and v
- Large number of operators \mathcal{O}_i already at 1b level, $> 2 \times 10$ at $\mathcal{O}(q^2, vq, v^2)$
- 2b corrections would require many more operators in addition
- Chiral EFT
 - Hierarchy predicted from chiral symmetry
 - Including 2b corrections much more efficient
 - \hookrightarrow fewer parameters



Case 1: spin-independent scattering



XENON1T 2018

SI scattering: only keep isoscalar, fully coherent response

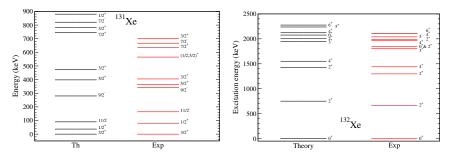
$$\frac{\mathrm{d}\sigma_{\chi N}}{\mathrm{d}q^2} = \frac{\sigma_{\chi N}^{\mathrm{SI}}}{4\mu_N^2 v^2} \left| \mathcal{F}_+^{M}(q^2) \right|^2 \qquad \mu_N = \frac{m_N m_\chi}{m_N + m_\chi}$$

Traditionally, "Helm form factor" Helm 1956, Lewin, Smith 1996

$$\mathcal{F}_{+}^{M}(q^{2})|_{\text{Helm}} = A \frac{3j_{1}(qr_{n})}{qr_{n}} e^{-\frac{1}{2}q^{2}s^{2}}$$

- → purely phenomenological
- Modern many-body methods: significant step forward, but not fully "ab initio" yet

Spectra and shell-model calculation



- Shell-model diagonalization for Xe isotopes with ¹⁰⁰Sn core
- Currently phenomenological shell-model interaction
 - \hookrightarrow chiral-EFT-based interactions in the future
- Progress in ab-initio methods: In-Medium Similarity Renormalization Group . . .
- Already possible for "light" nuclei

Further cross checks: charge radii and neutron skin

	19 _F	28 _{Si}	⁴⁰ Ar	⁷⁴ Ge	132 _{Xe}
$\sqrt{\langle r^2 \rangle_{\rm ch}}$ [fm] (th)	2.83	3.19	3.43	4.08	4.77
(exp)	2.898(2)	3.122(2)	3.427(3)	4.0742(12)	4.7808(49)
$\sqrt{\langle r_n^2 \rangle} - \sqrt{\langle r_p^2 \rangle}$ [fm]	0.02	0	0.11	0.17	0.28
shell-model interaction	USDB	USDB	SDPF.SM	RG	GCN

Excellent agreement for charge radii

$$\langle r_{\rm ch}^2 \rangle = \langle r_{p}^2 \rangle + \langle r_{E,p}^2 \rangle + \frac{N}{Z} \langle r_{E,n}^2 \rangle + \langle r_{\rm rel}^2 \rangle + \langle r_{\rm spin-orbit}^2 \rangle$$

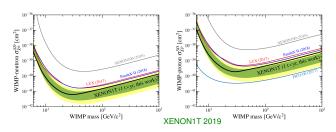
- Point-neutron radii more uncertain, can be tested by measuring weak radii
- Related to structure factors by

$$\langle r_{p}^{2} \rangle = -\frac{3}{Z} \frac{d}{dq^{2}} \left(\mathcal{F}_{+}^{M}(q^{2}) + \mathcal{F}_{-}^{M}(q^{2}) \right) \big|_{q^{2}=0} \qquad \langle r_{n}^{2} \rangle = -\frac{3}{N} \frac{d}{dq^{2}} \left(\mathcal{F}_{+}^{M}(q^{2}) - \mathcal{F}_{-}^{M}(q^{2}) \right) \big|_{q^{2}=0}$$

 Further cross checks: electric quadrupole and magnetic dipole moments, transition matrix elements



Case 2: spin-dependent scattering



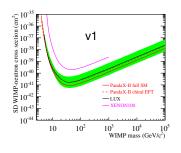
SD scattering: typically proton- or neutron-only

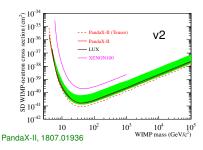
$$\frac{\mathsf{d}_{\chi N}}{\mathsf{d}q^2} = \frac{\sigma_{\chi N}^{\mathsf{SD}}}{3\mu_N^2 v^2} \frac{\pi}{2J+1} S_N(q^2)$$

- Xe sensitive to proton spin due to two-body currents
 Klos, Menéndez, Gazit, Schwenk 2013
- Power of EFT: parameters (low-energy constants) fixed from related processes



A note on spin-dependent scattering



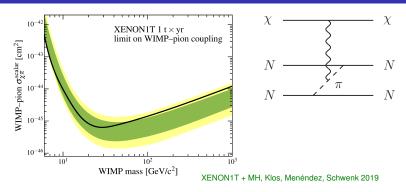


- Convention for SD scattering goes back at least to Engel, Pittel, Vogel 1992
 - \hookrightarrow axial-vector-axial-vector current $\bar{\chi}\gamma^{\mu}\gamma_5\chi\,\bar{q}\gamma_{\mu}\gamma_5q$ (motivated by SUSY)
- In QCD

$$\begin{split} \langle \textit{N}(\textit{p}')|\bar{q}\gamma_{\mu}\gamma_{5}\tau^{3}\textit{q}|\textit{N}(\textit{p})\rangle &= \langle \textit{N}(\textit{p}')|\gamma^{\mu}\gamma_{5}\textit{G}_{\textit{A}}(\textit{q}^{2})\tau^{3} + \gamma_{5}\frac{\textit{q}^{\mu}}{2m_{N}}\textit{G}_{\textit{P}}(\textit{q}^{2})\tau^{3}|\textit{N}(\textit{p})\rangle \\ \textit{G}_{\textit{A}}(0) &= \textit{g}_{\textit{A}} \qquad \textit{G}_{\textit{A}}(\textit{q}^{2}) - \frac{\textit{q}^{2}}{4m_{N}^{2}}\textit{G}_{\textit{P}}(\textit{q}^{2}) = \mathcal{O}(\textit{M}_{\pi}^{2}) \end{split}$$

- Induced pseudoscalar $G_P(q^2)$ neglected in v1, "improving" the LUX limits
 - → need to use consistent conventions for meaningful comparison!

Case 3: WIMP-pion scattering



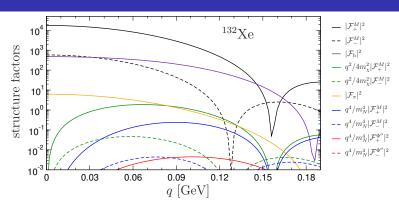
One way to illustrate 2b effects: WIMP-pion scattering

$$\frac{\mathrm{d}\sigma_{\chi\mathcal{N}}}{\mathrm{d}q^2} = \frac{\sigma_{\chi\pi}^{\mathrm{scalar}}}{\mu_{\pi}^2 v^2} \left| \mathcal{F}_{\pi}(q^2) \right|^2 \qquad \mu_{\pi} = \frac{m_{\chi} M_{\pi}}{m_{\chi} + M_{\pi}}$$

• Expression in terms of cross section depends on underlying operator, here for a scalar $\bar{\chi}\chi\bar{q}q$



Full set of coherent contributions



$$\begin{split} \frac{\mathrm{d}\sigma_{\chi\mathcal{N}}}{\mathrm{d}q^2} &= \frac{1}{4\pi v^2} \bigg| \sum_{I=\pm} \left(c_I^M - \frac{q^2}{m_N^2} \dot{c}_I^M \right) \mathcal{F}_I^M(q^2) + c_\pi \mathcal{F}_\pi(q^2) + c_b \mathcal{F}_b(q^2) + \frac{q^2}{2m_N^2} \sum_{I=\pm} c_I^{\Phi''} \mathcal{F}_I^{\Phi''}(q^2) \bigg|^2 \\ &+ \frac{1}{4\pi v^2} \sum_{i=5,8,11} \bigg| \sum_{I=\pm} \xi_i(q,v_T^\perp) c_I^{M,i} \mathcal{F}_I^M(q^2) \bigg|^2 \\ &+ \frac{1}{v^2(2J+1)} \Big(|a_+|^2 S_{00}(q^2) + \operatorname{Re}\left(a_+ a_+^*\right) S_{01}(q^2) + |a_-|^2 S_{11}(q^2) \Big) \end{split}$$

Role of the coefficients

- Coefficients c, a: combinations of Wilson coefficients (BSM operators) and hadronic matrix elements
- In above examples expressed in terms of single-particle cross sections

$$\sigma_{\chi N}^{\rm SI} = \frac{\mu_N^2}{\pi} \big| c_+^M \big|^2 \qquad \sigma_{\chi N}^{\rm SD} = \frac{3\mu_N^2}{\pi} \big| a_+ \big|^2 \qquad \sigma_{\chi \pi}^{\rm scalar} = \frac{\mu_\pi^2}{4\pi} \big| c_\pi \big|^2$$

- But: to constrain a particular BSM operator/model: need to disentangle
 → matching relations
- Example: Dirac spin-1/2 WIMP, scalar operator $\mathcal{L}^{\text{BSM}} = C_q^{\text{SS}} \bar{\chi} \chi m_q \bar{q} q$

$$c_{\pm}^{M} = \frac{1}{2} \left(f_{p} \pm f_{n} \right) + \cdots \qquad f_{N} = m_{N} \sum_{q=u,d,s} C_{q}^{SS} f_{q}^{N} + \cdots \qquad \langle N | m_{q} \bar{q} q | N \rangle = m_{N} f_{q}^{N}$$

 \hookrightarrow scalar couplings of the nucleon f_q^N



ChiralEFT4DM: a PYTHON package for dark matter

ChiralEFT4DM: all results available as PYTHON package at

- Includes:
 - (Quasi-) Coherent structure factors for F, Si, Ar, Ge, Xe
 - Nucleon matrix elements and matching relations for spin-1/2 and spin-0 WIMP
 - 1b and 2b responses up to third chiral order
 - $S, P, V, A, T, \theta^{\mu}_{\mu}$, and spin-2 effective operators that can lead to a coherent response
 - Convolution with Standard Halo Model

Uncertainties

EFT truncations:

- Order-by-order convergence
- Bayesian interpretation Furnstahl, Phillips, Wesolowski 2015 . . .

Low-energy constants:

- Single-nucleon couplings mostly well determined from simpler processes
- Two-nucleon parameters more uncertain

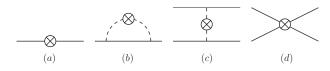
Many-body methods:

- Not fully systematic yet, but progress towards ab-initio calculations
- Compare to observables: spectra, radii, EM transitions

• Hadronic matrix elements:

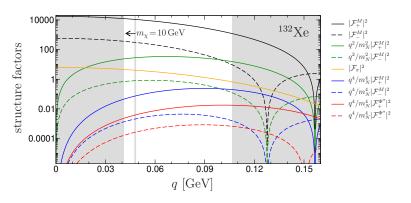
- Vector and axial-vector couplings well known
- Tension between lattice QCD and phenomenology for scalar couplings

Conclusions



- Chiral EFT for WIMP-nucleus scattering
 - Connects nuclear and hadronic scales
 - Systematic approach to nuclear responses thanks to EFT
- Uncertainties
 - EFT truncation errors: currents, few-nucleon interaction
 - Low-energy constants: input from few-nucleon observables
 - Many-body method: towards ab-initio for "heavy" nuclei
 - Hadronic matrix elements: for BSM matching

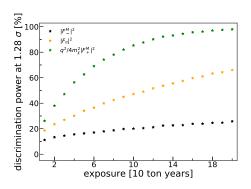
Discriminating different response functions



- White region accessible to XENON-type experiment
- Can one tell these curves apart in a realistic experimental setting?
- Consider XENON1T-like, XENONnT-like, DARWIN-like settings



Discriminating different response functions



- DARWIN-like setting, $m_{\chi} = 100 \,\text{GeV}$
- q-dependent responses more easily distinguishable
- If interaction not much weaker than current limits, DARWIN could discriminate most responses from standard SI structure factor

Higgs Portal dark matter

Higgs Portal: WIMP interacts with SM via the Higgs

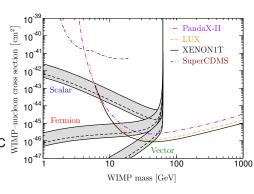
Scalar: H[†] H S²

• Vector: $H^{\dagger}HV_{\mu}V^{\mu}$

• Fermion: H[†]H ff

• If $m_h > 2m_\chi$, should happen at the LHC

 \hookrightarrow limits on invisible Higgs decays



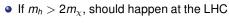
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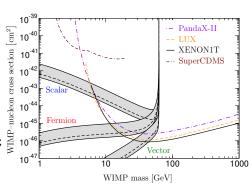
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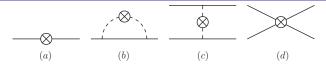
Translation requires input for Higgs-nucleon coupling

$$f_{N} = \sum_{q=u,d,s,c,b,t} f_{q}^{N} = \frac{2}{9} + \frac{7}{9} \sum_{q=u,d,s} f_{q}^{N} + \mathcal{O}(\alpha_{s}) \qquad m_{N} f_{q}^{N} = \langle N | m_{q} \bar{q} q | N \rangle$$

Issues: input for f_N = 0.260...0.629 outdated, 2b currents missing



Higgs-nucleon coupling



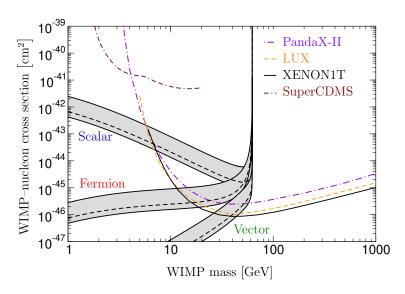
One-body contribution

$$f_N^{1b} = 0.307(9)_{ud}(15)_s(5)_{pert} = 0.307(18)$$

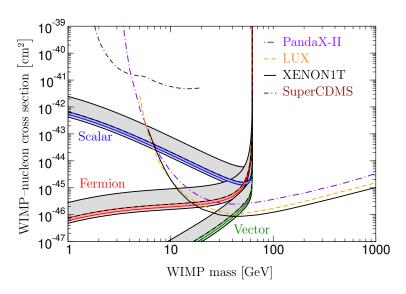
- Limits on WIMP-nucleon cross section subsume 2b effects
 - → have to be included for meaningful comparison
- Two-body contribution
 - Need s and θ^{μ}_{μ} currents
 - Treatment of θ^μ_μ tricky: several ill-defined terms combine to $\langle \Psi | T + V_{NN} | \Psi \rangle = E_{\rm b}$
 - A cancellation makes the final result anomalously small

$$f_N^{\text{2b}} = [-3.2(0.2)_A(2.1)_{\text{ChEFT}} + 5.0(0.4)_A] \times 10^{-3} = 1.8(2.1) \times 10^{-3}$$

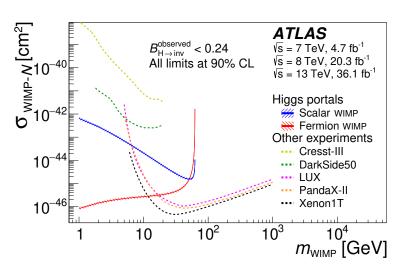
Improved limits for Higgs Portal dark matter



Improved limits for Higgs Portal dark matter

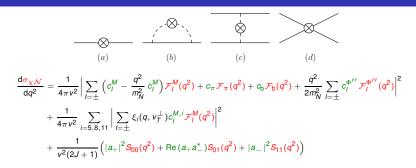


Improved limits for Higgs Portal dark matter



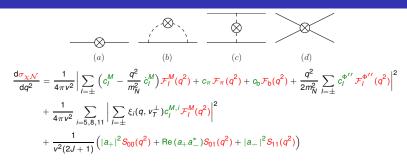
ATLAS 2019

Cross section and nuclear structure factors



- Decomposition into nuclear structure factors \mathcal{F} , S_{ii} and coefficients c, a
- Three classes of contributions:
 - (Sub-) Leading 1b responses (a): $c_l^M \mathcal{F}_l^M(q^2)$, $c_l^{\phi''} \mathcal{F}_l^{\phi''}(q^2)$, $|a_{\pm}|^2 S_{ii}(q^2)$
 - Radius corrections (b): $\dot{c}_{i}^{M}\mathcal{F}_{i}^{M}(q^{2})$
 - Two-body currents (c), (d): $c_{\pi}\mathcal{F}_{\pi}(a^2)$, $c_{h}\mathcal{F}_{h}(a^2)$
- (a)+(b) essentially **nucleon form factors**, but (c)+(d) genuinely new effects

Cross section and nuclear structure factors



- Nuclear structure interpretation:
 - $\mathcal{F}_{l}^{M}(q^{2})$: $\mathbb{1} \Rightarrow$ charge distribution (coherent), $\mathcal{F}_{\pm}^{M}(0) = Z \pm N$
 - $\mathcal{F}_{l}^{\Phi''}(q^{2})$: $i\mathbf{S}_{N} \cdot (\mathbf{q} \times \mathbf{v}^{\perp}) \Rightarrow$ spin-orbit interaction (quasi-coherent)
 - $S_{ij}(q^2)$: $\mathbf{S}_{\chi} \cdot \mathbf{S}_N$, $\mathbf{S}_{\chi} \cdot \mathbf{q} \mathbf{S}_N \cdot \mathbf{q} \Rightarrow$ spin average (not coherent), $S_{00}(0) \pm S_{01}(0) + S_{11}(0) = \frac{(2J+1)(J+1)}{4\pi J} |\langle \mathbf{S}_{p/n} \rangle|^2$
- Coefficients: convolution of Wilson coefficients and nucleon matrix elements

$$c_{\pm}^{M} = \frac{\zeta}{2} (f_p \pm f_n) + \cdots \qquad f_N = \frac{m_N}{\Lambda^3} \sum_{q=u,d,s} C_q^{SS} f_q^N + \cdots \qquad \langle N | m_q \bar{q} q | N \rangle = m_N f_q^N$$

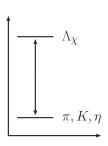
Chiral Perturbation Theory

Effective theory of QCD based on chiral symmetry

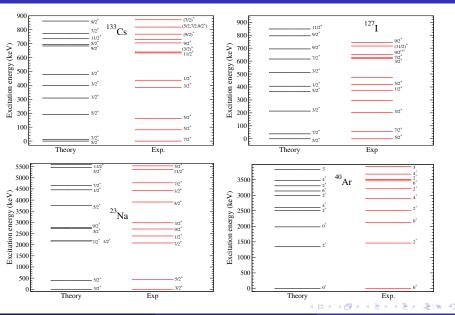
$$\mathcal{L}_{\text{QCD}} = \bar{q}_L i \not\!\!D q_L + \bar{q}_R i \not\!\!D q_R - \bar{q}_L \mathcal{M} q_R - \bar{q}_R \mathcal{M} q_L - \frac{1}{4} \textit{G}^a_{\mu\nu} \textit{G}^{\mu\nu}_a$$

- Expansion in momenta p/Λ_{χ} and quark masses $m_q \sim p^2$ \hookrightarrow scale separation
- Breakdown scale: $\Lambda_{\chi} = M_{\rho} \dots 4\pi F_{\pi} \sim 1 \text{ GeV}$
- Two variants
 - SU(2): u- and d-quark dynamical, m_s fixed at physical value \hookrightarrow expansion in M_π/Λ_χ , usually nice convergence
 - SU(3): u-, d-, and s-quark dynamical

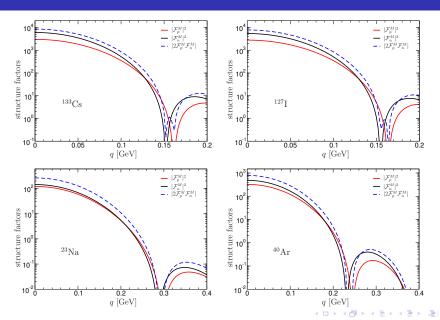
 ⇔ expansion in M_K/Λ_χ, sometimes tricky



Spectra



Structure factors



Chiral counting

ullet Effective WIMP Lagrangian for spin-1/2 SM singlet χ Goodman et al. 2010

$$\begin{split} \mathcal{L}_{\chi} &= \frac{1}{\Lambda^{3}} \sum_{q} \left[C_{q}^{SS} \bar{\chi} \chi \, m_{q} \bar{q} q + C_{q}^{PS} \bar{\chi} i \gamma_{5} \chi \, m_{q} \bar{q} q + C_{q}^{SP} \bar{\chi} \chi \, m_{q} \bar{q} i \gamma_{5} q + C_{q}^{PP} \bar{\chi} i \gamma_{5} \chi \, m_{q} \bar{q} i \gamma_{5} q \right] \\ &+ \frac{1}{\Lambda^{2}} \sum_{q} \left[C_{q}^{VV} \bar{\chi} \gamma^{\mu} \chi \, \bar{q} \gamma_{\mu} q + C_{q}^{AV} \bar{\chi} \gamma^{\mu} \gamma_{5} \chi \, \bar{q} \gamma_{\mu} q + C_{q}^{VA} \bar{\chi} \gamma^{\mu} \chi \, \bar{q} \gamma_{\mu} \gamma_{5} q + C_{q}^{AA} \bar{\chi} \gamma^{\mu} \gamma_{5} \chi \, \bar{q} \gamma_{\mu} \gamma_{5} q \right] \\ &+ \frac{1}{\Lambda^{3}} \left[C_{g}^{S} \bar{\chi} \chi \, \alpha_{S} G_{\mu\nu}^{a} G_{a}^{\mu\nu} \right] \end{split}$$

Chiral power counting

$$\partial = \mathcal{O}(p), \qquad m_q = \mathcal{O}(p^2) = \mathcal{O}(M_\pi^2), \qquad a_\mu, v_\mu = \mathcal{O}(p), \qquad \frac{\partial}{m_N} = \mathcal{O}(p^2)$$

- \hookrightarrow construction of effective Lagrangian for nucleon and pion fields
- \hookrightarrow organize in terms of **chiral order** ν , $\mathcal{M} = \mathcal{O}(p^{\nu})$

Chiral counting: summary

	Nucleon		V		Α		Nucleon	S	P
WIMP		t	x	t	x	WIMP			
	1b	0	1 + 2	2	0 + 2		1b	2	1
V	2b	4	2 + 2	2	4 + 2	S	2b	3	5
	2b NLO	_	_	5	3 + 2		2b NLO	_	4
	1b	0 + 2	1	2 + 2	0		1b	2 + 2	1 + 2
Α	2b	4 + 2	2	2 + 2	4	P	2b	3 + 2	5+2
	2b NLO	_	_	5 + 2	3		2b NLO	_	4 + 2

- ullet +2 from NR expansion of WIMP spinors, terms can be dropped if $m_\chi \gg m_N$
- Red: all terms up to $\nu = 3$
- 1b: one-body (single-nucleon), 2b: two-body, 2b NLO: two-body at (nominally) next-to-leading order

Example: chiral counting in scalar channel

Leading pion-nucleon Lagrangian

$$\mathcal{L}_{\pi N}^{(1)} = \bar{\Psi} \bigg[i \gamma_{\mu} \big(\partial^{\mu} - i \textbf{\textit{v}}^{\mu} \big) - \textit{m}_{N} + \frac{g_{A}}{2} \gamma_{\mu} \gamma_{5} \Big(2 \frac{\textbf{\textit{a}}^{\mu}}{F_{\pi}} - \frac{\partial^{\mu} \pi}{F_{\pi}} \Big) + \cdots \bigg] \Psi$$

→ no scalar source!

	Nucleon	s
WIMP		
	1b	2
S	2b	3

Example: chiral counting in scalar channel

Leading pion–nucleon Lagrangian

$$\mathcal{L}_{\pi N}^{(1)} = \bar{\Psi} \bigg[i \gamma_{\mu} \big(\partial^{\mu} - i \textbf{\textit{v}}^{\mu} \big) - \textit{m}_{N} + \frac{g_{A}}{2} \gamma_{\mu} \gamma_{5} \Big(2 \frac{\textbf{\textit{a}}^{\mu}}{F_{\pi}} \Big) + \cdots \bigg] \Psi$$

- → no scalar source!
- Scalar coupling

$$f_N = \frac{m_N}{\Lambda^3} \sum_{q=u,d,s} C_q^{SS} f_q^N + \cdots \qquad \langle N | m_q \bar{q} q | N \rangle = f_q^N m_N$$

Nucleon S
WIMP

1b 2
S 2b 3

- \hookrightarrow for q = u, d related to **pion–nucleon** σ -term $\sigma_{\pi N}$
- Chiral expansion

$$\sigma_{\pi N} = -4c_1 M_{\pi}^2 - \frac{9g_A^2 M_{\pi}^3}{64\pi F_{\pi}^2} + \mathcal{O}(M_{\pi}^4) \qquad \dot{\sigma} = \frac{5g_A^2 M_{\pi}}{256\pi F_{\pi}^2} + \mathcal{O}(M_{\pi}^2)$$

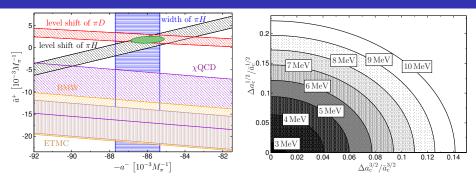
- \hookrightarrow slow convergence due to strong $\pi\pi$ rescattering
- \hookrightarrow use phenomenology for the full scalar form factor!



Extracting $\sigma_{\pi N}$ from πN scattering: low-energy theorem

- No scalar probe, but still relation to experiment! How?
 - → low-energy theorem
- Topic for another talk MH, Ruiz de Elvira, Kubis, Meißner PRL 115 (2015) 092301, PLB 760 (2016) 74, JPG 45 (2018) 024001
 - Goes back to Cheng, Dashen; Brown, Pardee, Peccei 1971
 - Relates $\sigma_{\pi N}$ to πN scattering amplitude, but at **unphysical kinematics**
 - No chiral logs at one-loop order! Bernard, Kaiser, Meißner 1996
 - Protected by SU(2)
 - \hookrightarrow expected correction: $\sigma_{\pi N} M_{\pi}^2/m_N^2 \sim 1 \text{ MeV}$

$\pi N \sigma$ -term: a lingering tension with lattice QCD

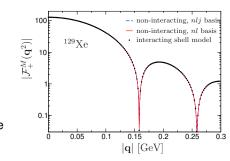


- Independent experimental constraints from pionic atoms and low-energy cross sections agree at the level of $\sigma_{\pi N} = 58(5)\,\text{MeV}$
- This needs to be resolved: rare opportunity to benchmark lattice BSM matrix elements from experiment

Scalar two-body currents: oscillator model

$$\begin{split} \mathcal{F}_{\pi}(\mathbf{q}^2) &= \frac{M_{\pi}}{2} \left(\frac{g_A}{2F_{\pi}} \right)^2 \sum_{n_1 l_1 n_2 l_2} \sum_{\tau_1 \tau_2} \int \frac{\mathrm{d}^3 \rho_1 \mathrm{d}^3 \rho_2 \mathrm{d}^3 \rho_1' \mathrm{d}^3 \rho_2'}{(2\pi)^6} R_{n_1 l_1}(|\mathbf{p}_1'|) R_{n_2 l_2}(|\mathbf{p}_2'|) R_{n_1 l_1}(|\mathbf{p}_1|) R_{n_2 l_2}(|\mathbf{p}_2|) \\ &\times \frac{(2l_1 + 1)(2l_2 + 1)}{16\pi^2} P_{l_1}(\hat{\mathbf{p}}_1' \cdot \hat{\mathbf{p}}_1) P_{l_2}(\hat{\mathbf{p}}_2' \cdot \hat{\mathbf{p}}_2) (2\pi)^3 \delta^{(3)}(\mathbf{p}_1 + \mathbf{p}_2 - \mathbf{p}_1' - \mathbf{p}_2' - \mathbf{q}) \\ &\times (3 - \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2) \frac{\mathbf{q}_1^{\mathrm{ex}} \cdot \mathbf{q}_2^{\mathrm{ex}}}{((\mathbf{q}_1^{\mathrm{ex}})^2 + M_{\pi}^2)((\mathbf{q}_2^{\mathrm{ex}})^2 + M_{\pi}^2)} \end{split}$$

- Two-body current defines genuinely new structure factor
- Checked the oscillator model for 1b case \hookrightarrow reproduces perfectly the L=0 multipole



Scalar two-body currents: numerical estimates

Early claims: could be as large as 60% Prézeau, Kurylov, Kamionkowski, Vogel 2003

$$\frac{\mathcal{A}_{\pi\pi}}{\mathcal{A}_{NN}} \simeq (0.21 \pm 0.08) r \frac{\mathcal{N}_{\pi\pi}}{A} \qquad r = \frac{S_u m_u + S_d m_d}{(S_u m_u + S_d m_d) \frac{\epsilon_s^0}{2} + \sum_{q=s,c,b,t} S_q m_q \epsilon_s^q}$$

• They find $1 < \frac{N_{\pi\pi}}{A} < 2$ and consider r as large as 1.5. This brings you to 60%.

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- They find $1 < \frac{N_{\pi\pi}}{A} < 2$ and consider r as large as 1.5. This brings you to 60%.
- Decomposition is actually scale dependent, they quote

$$\epsilon_s^0 = \langle N | \bar{u}u + \bar{d}d | N \rangle \simeq 16 \pm 8$$

ullet Without such cancellations ($m{C}_s^{SS} = m{C}_g^{\prime S} = 0$) MH et al. 2016

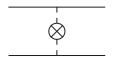
$$2\frac{2f_{\pi}}{f_{p}+f_{n}}\frac{\mathcal{F}_{\pi}(0)}{A}=-9\%$$

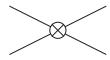
- Claimed enhancement not at all related to power counting, but to cancellations in BSM parameter space
- To actually check power counting: scalar current in light nuclei + lattice

 σ -terms work in progress, NPLQCD 2013, 2018, Körber et al. 2017



Contact terms





- Scalar source suppressed for (N[†]N)²
 - → long-range contribution dominant (in Weinberg counting)
- Typical size (5–10)%
 - → reflected by results for structure factors
 - \hookrightarrow more important in case of cancellations
- Contact terms do appear for other sources, e.g. θ^μ_μ
 - \hookrightarrow related to nuclear binding energy E_b
- Same structure factor in spin-2 two-body currents

Coherence effects

Six distinct nuclear responses

Fitzpatrick et al. 2012, Anand et al. 2013

•
$$M \leftrightarrow \mathcal{O}_1 \leftrightarrow SI$$

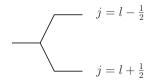
•
$$\Sigma', \Sigma'' \leftrightarrow \mathcal{O}_4, \mathcal{O}_6 \leftrightarrow SD$$

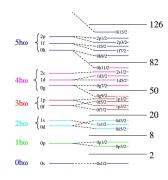
- $\Phi'' \leftrightarrow \mathcal{O}_3 \leftrightarrow$ quasi-coherent, spin-orbit operator
- Δ , $\tilde{\Phi}'$: not coherent

Quasi-coherence of Φ"

- Spin-orbit splitting
- Coherence until mid-shell
- About 20 coherent nucleons in Xe
- Interference $M-\Phi'' \leftrightarrow \mathcal{O}_1-\mathcal{O}_3$
- Coherent 2b currents:

 - Vector $\propto N Z$





Spin-2 and coupling to the energy-momentum tensor

- Effective Lagrangian truncated at dim-7, but if WIMP heavy $m_\chi/\Lambda = \mathcal{O}(1)$
 - → heavy-WIMP EFT Hill, Solon 2012, 2014

$$\mathcal{L} = \frac{1}{\Lambda^4} \bigg\{ \sum_q \frac{C_q^{(2)}}{\bar{\chi}} \bar{\chi} \gamma_\mu i \partial_\nu \chi \frac{1}{2} \bar{q} \Big(\gamma^{\{\mu} i \mathcal{D}_-^{\nu\}} - \frac{m_q}{2} g^{\mu\nu} \Big) q + \frac{C_g^{(2)}}{g} \bar{\chi} \gamma_\mu i \partial_\nu \chi \Big(\frac{g_{\mu\nu}}{4} G_{\lambda\sigma}^a G_a^{\lambda\sigma} - G_a^{\mu\lambda} G_{a\lambda}^{\nu} \Big) \bigg\}$$

- → similar two-body current as in scalar case, pion pdfs, EMC effect
- Coupling of trace anomaly θ^{μ}_{μ} to $\pi\pi$

$$\theta^{\mu}_{\mu} = \sum_{q} m_{q} \bar{q} q + \frac{\beta_{\text{QCD}}}{2g_{s}} G^{a}_{\mu\nu} G^{\mu\nu}_{a} \quad \Leftrightarrow \quad \langle \pi(p') | \theta_{\mu\nu} | \pi(p) \rangle = \rho_{\mu} \rho'_{\nu} + \rho'_{\mu} \rho_{\nu} + g_{\mu\nu} \left(M_{\pi}^{2} - p \cdot p' \right)$$

 \hookrightarrow probes gluon Wilson coefficient C_g^S

