

**The Modern Physics of Compact Stars and
Relativistic Gravity**

**Vacuum fluxes from a plate in de
Sitter spacetime with compact
dimensions**

Yerevan, Armenia,
September 17-21, 2019

A. A. Saharian, D. H. Simonyan, T. A. Petrosyan

Yerevan State University,

Gourgen Sahakyan Chair of Theoretical Physics

Outline

- Geometry of the problem and the Hadamard function
- VEV of the field squared
- VEV of the energy momentum tensor: energy flux
- Conclusions

Geometry of the problem

- We consider combined effects of **topology** and the **presence of a boundary** on the VEVs of the field squared and energy-momentum tensor for charged scalar field on the background of dS spacetime with an arbitrary number of toroidally compactified spatial dimensions

- **dS spacetime** is described by the line element:

$$ds^2 = dt^2 - e^{2t/\alpha} \sum_{l=1}^D (dz^l)^2,$$

Parameter α is related to the positive **cosmological constant** Λ by the formula $\alpha^2 = D(D-1)/(2\Lambda)$

- With **conformal time** $\tau = -\alpha e^{-t/\alpha}$ $-\infty < \tau \leq 0$ the line element takes conformally flat form $ds^2 = (\alpha/\tau)^2 \eta_{\mu\nu} dx^\mu dx^\nu$

- **Spatial Topology** $R^p \times (S^1)^q,$

Uncompact dimensions $\mathbf{x}_p = (x^1, \dots, x^p)$ $-\infty < x^l < \infty, l = 1, \dots, p$

Compact dimensions $\mathbf{x}_q = (x^{p+1}, \dots, x^D)$ $L_l: 0 \leq x^l \leq L_l, l = p+1, \dots, D.$

$q = D - p$

↑
Length of the compact dimension

- Physical (comoving) length of compact dimension $\alpha L_l / \eta, \eta = |\tau|$

Geometry of the problem

- In addition, we assume **planar boundary** at $x^p = 0$ on which the field obeys the **Robin** boundary condition

$$(1 + \beta D_p)\varphi = 0, \quad x^p = 0$$

Special cases: **Dirichlet**
and **Neumann** conditions

- We investigate boundary-induced effects

Complex scalar field

■ We consider a **complex scalar field** $\varphi(x)$ with a curvature coupling parameter ξ , in the presence of a classical abelian gauge field A_μ . The corresponding field equation has the form

$$(D_\mu D^\mu + m^2 + \xi R) \varphi(x) = 0.$$

■ In the most important special cases of **minimally** and **conformally** coupled scalars one has $\xi = 0$ and $\xi = (D-1)/(4D)$ respectively

- $D_\mu = \nabla_\mu + ieA_\mu$ gauge-covariant derivative
- e coupling between the scalar and gauge fields
- $R = D(D+1)/\alpha^2$ scalar curvature

■ One should specify the **periodicity conditions** along compact dimensions:

$$\varphi(t, \mathbf{x}_p, \mathbf{x}_q + L_l \mathbf{e}_l) = e^{i\alpha_l} \varphi(t, \mathbf{x}_p, \mathbf{x}_q).$$

$\alpha_l = \text{const}$, \mathbf{e}_l is the unit vector along the dimension $l = p+1, \dots, D$

Gauge transformation

- We consider the simplest configuration of the gauge field with $A_\mu = \text{const}$
The gauge field can be excluded from the field equation by the gauge transformation

$$A'_\mu = A_\mu + \partial_\mu \omega, \quad \varphi'(x) = e^{-ie\omega} \varphi(x), \quad \omega = -A_\mu x^\mu$$

- In the new gauge one has $A'_\mu = 0$ and $D'_\mu = \nabla_\mu$

- Quasi-periodicity condition after gauge transformation

$$\varphi'(t, \mathbf{x}_p, \mathbf{x}_q + L_l \mathbf{e}_l) = e^{i\tilde{\alpha}_l} \varphi'(t, \mathbf{x}_p, \mathbf{x}_q)$$

- The phases in the periodicity conditions and the value of the gauge field are related to each other through a gauge transformation.

$$\tilde{\alpha}_l = \alpha_l + eA_l L_l$$

Effect of gauge field

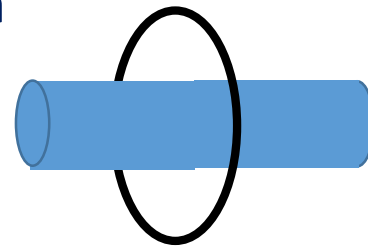
Although the field strength vanishes, a constant gauge field shifts the phases in the periodicity conditions along compact dimensions

This leads to the **Aharonov-Bohm-like effects** on the expectation values

The shift due to the gauge field may be written in the form,

$$eA_l L_l = -eA_l L_l = -2\pi\Phi_l / \Phi_0, \text{ where } \Phi_0 = 2\pi / e$$

is the **flux quantum** and Φ_l is the flux enclosed by the circle corresponding to the l -th compact dimension



Complete set of solutions and Hadamard function

Mode-sum for the Hadamard function

Mode functions are specified by the set of quantum numbers

$$G^{(1)}(x, x') = \sum_{\sigma} \sum_{s=\pm} \varphi_{\sigma}^{(s)}(x) \varphi_{\sigma}^{(s)*}(x')$$

$$\sigma = (\mathbf{k}_p, \mathbf{n}_q) \quad \mathbf{n}_q = (n_{p+1}, \dots, n_D)$$

Complete set of solutions to the field equation

We assume that the scalar field is prepared in the **Bunch-Davies vacuum**

$$\varphi_{\sigma}^{(+)}(x) = C_{\sigma}^{(+)} \eta^{D/2} H_{\nu}^{(1)}(K\eta) \cos(k_p x^p + \alpha(k_p)) e^{i\mathbf{k}_{p-1} \mathbf{x}_{p-1} + i\mathbf{k}_q \mathbf{x}_q}$$

$$\nu = [D^2/4 - D(D+1)\xi - m^2\alpha^2]^{1/2} \quad \varphi_{\sigma}^{(-)}(x) = \varphi_{\sigma}^{(+)*}(x)$$

Momentum along uncompactified dimensions $H_{\nu}^{(1)}(z)$ Hankel function

$$-\infty < k_l < +\infty, \quad l = 1, \dots, p, \quad 0 \leq k_p < +\infty$$

Along compactified dimensions

$$k_l = (2\pi n_l + \tilde{\alpha}_l) / L_l, \quad n_l = 0, \pm 1, \pm 2, \dots, \quad l = p+1, \dots, D$$

Complete set of solutions and Hadamard function

- The **Hadamard function** is split into two parts

$$G^{(1)}(x, x') = G_0^{(1)}(x, x') + G_b^{(1)}(x, x')$$

without boundary

boundary-induced part

$$G_b^{(1)}(x, x') = \frac{(\eta\eta')^{D/2}}{2^p \pi^p V_q \alpha^{D-1}} \int d\mathbf{k}_{p-1} e^{i\mathbf{k}_{p-1} \Delta \mathbf{x}_{p-1}} \sum_{\mathbf{n}_q} e^{i\mathbf{k}_q \Delta \mathbf{x}_q} \int_k^\infty du e^{-u(x^p + x^{p'})} \frac{\beta u + 1}{\beta u - 1} \\ \times \{K_\nu(\eta y)[I_{-\nu}(\eta' y) + I_\nu(\eta' y)] + [I_{-\nu}(\eta y) + I_\nu(\eta y)]K_\nu(\eta' y)\}_{y=\sqrt{u^2 - k^2}}$$

VEV of the field squared

- Given the Hadamard function, the vacuum expectation value of the field squared is evaluated as the coincidence limit

$$\langle \varphi \varphi^\dagger \rangle = \frac{1}{2} \lim_{x' \rightarrow x} G^{(1)}(x, x')$$

- $\langle \varphi \varphi^\dagger \rangle = \langle \varphi \varphi^\dagger \rangle_0 + \langle \varphi \varphi^\dagger \rangle_b$

without boundary

boundary-induced part

$$\langle \varphi \varphi^\dagger \rangle_b = \frac{2^{1-p} \pi^{-(p+1)/2} \eta^D}{\Gamma((p-1)/2) V_q \alpha^{D-1}} \sum_{\mathbf{n}_q} \int_0^\infty dx x^p f(\sqrt{x^2 + k_{\mathbf{n}_q}^2}, x^p) h_\nu(\eta x)$$

$$h_\nu(u) = \int_0^1 dz z (1 - z^2)^{(p-3)/2} g_\nu(uz)$$

$$g_\nu(x) = K_\nu(x) [I_{-\nu}(x) + I_\nu(x)],$$

$$k_{\mathbf{n}_q}^2 = \sum_{l=p+1}^D (2\pi n_l + \tilde{\alpha}_l)^2 / L_l^2$$

$$f(u, x^p) = \frac{e^{-2ux^p}}{u} \frac{\beta u + 1}{\beta u - 1}$$

VEV of the field squared

- For a **conformally coupled massless** scalar field

$$\langle \varphi \varphi^+ \rangle_b = \frac{(\eta/\alpha)^{D-1}}{(4\pi)^{p/2} \Gamma(p/2) V_q} \sum_{\mathbf{n}_q} \int_{k_{\mathbf{n}_q}}^{\infty} dx (x^2 - k_{\mathbf{n}_q}^2)^{p/2-1} e^{-2xx^p} \frac{\beta x + 1}{\beta x - 1}$$

Conformal factor

Conformally related to the result in Minkowski bulk

- For **Dirichlet** and **Neumann** boundary conditions

$$\langle \varphi \varphi^+ \rangle_b = \mp \frac{2\eta^D (2x^p)^{2-p}}{(2\pi)^{p/2+1} V_q \alpha^{D-1}} \int_0^{\infty} dy y g_\nu(\eta y) \sum_{\mathbf{n}_q} f_{p/2-1}(2x^p \sqrt{y^2 + k_{\mathbf{n}_q}^2})$$

VEV of the energy-momentum tensor

- Vacuum expectation value of the energy-momentum tensor

$$\langle T_{ik} \rangle_{p,q} = \frac{1}{2} \lim_{x' \rightarrow x} \partial_i \partial'_k G_{p,q}(x, x') + \frac{1}{2} \left[\left(\xi - \frac{1}{4} \right) g_{ik} \nabla_l \nabla^l - \xi \nabla_i \nabla_k - D \xi g_{ik} / \alpha^2 \right] G_{p,q}(x, x)$$

- $\langle T_{ik} \rangle = \langle T_{ik} \rangle_0 + \langle T_{ik} \rangle_b$

without boundary

boundary-induced part

$$\langle T_i^k \rangle_b = \frac{A_p}{V_q \alpha^{D+1}} \sum_{\mathbf{n}_q} \int_0^\infty dx x^{p-D} f(\sqrt{x^2 + k_{\mathbf{n}_q}^2}, x^p) \times \left[\hat{F}_i^k(\eta) h_\nu^{(0)}(\eta x) + \eta^2 x^2 F_i^k h_\nu^{(1)}(\eta x) \right],$$

- The only nonzero off-diagonal component $\langle T_0^p \rangle$

- Boundary-induced part obeys the **covariant continuity equation**

$$\nabla_k \langle T_i^k \rangle_b = 0$$

VEV of the energy-momentum tensor

- Off-diagonal component $\langle T_0^p \rangle_b$ appears (energy flux)

$$\langle T_0^p \rangle_b = \frac{\pi^{-(p+1)/2} (\eta/\alpha)^{D+1}}{2^{p-2} \Gamma((p-1)/2) V_q} \sum_{\mathbf{n}_q} \int_0^\infty dy y F_\nu(\eta y) \\ \times \int_0^\infty dw w^{p-2} \frac{\beta u + 1}{\beta u - 1} e^{-2x^p u} \Big|_{u=\sqrt{y^2+w^2+k_{\mathbf{n}_q}^2}},$$

where $F_\nu(x) = \left[\left(\xi - \frac{1}{4} \right) x \partial_x + (D+1)\xi - \frac{D}{4} \right] g_\nu(x)$

- In special cases of Dirichlet and Neumann boundary conditions

$$\langle T_0^p \rangle = \mp \frac{4(\eta/\alpha)^{D+1}}{(2\pi)^{p/2+1} V_q (2x^p)^{p-1}} \sum_{\mathbf{n}_q} \int_0^\infty dy y f_{p/2}(2x^p \sqrt{y^2 + k_q^2}) F_\nu(\eta y)$$

- For Dirichlet and Neumann boundary conditions the energy flux has opposite signs

Asymptotic behavior

- For $x^p / \eta \ll 1$ (x^p / η is the proper distance from the boundary in units of the curvature radius $1/\alpha$)

the influence of the **gravitational field is weak**

- For $x^p / \eta \square 1$

- $\nu^2 > 0$, in addition $x^p \gg L_l$

$$\langle \varphi \varphi^+ \rangle_b \approx \mp \frac{2^{2\nu-p} \Gamma(\nu) \Gamma(p/2 - \nu) \eta^{D-2\nu}}{(2\pi)^{p/2+1} V_q \alpha^{D-1} (x^p)^{3p/2-1-2\nu}} \longrightarrow \text{For } \tilde{\alpha}_l = 0, l = p+1, \dots, D$$

Power-law decay

$$\langle \varphi \varphi^+ \rangle_b \approx \mp \frac{2^{2\nu-p} \Gamma(\nu) \eta^{D-2\nu} k_{(0)}^{(p-1)/2-\nu} e^{-2x^p k_{(0)}}}{(2\pi)^{(p+1)/2} V_q \alpha^{D-1} (x^p)^{p-\nu-1/2}} \longrightarrow \text{If, at least one } \tilde{\alpha}_l \neq 0$$

Exponential decay

- $\nu^2 < 0$, in addition $x^p \gg L_l$

$$\langle \varphi \varphi^+ \rangle_b \approx \mp \frac{\pi^{-p/2-1} A_p \eta^D}{2^{p+1} V_q \alpha^{D-1} (x^p)^p} \cos [2|\nu| \ln (2x^p / \eta) + \phi_1] \uparrow$$

For $\tilde{\alpha}_l = 0, l = p+1, \dots, D$

If, at least one $\tilde{\alpha}_l \neq 0$

$$\langle \varphi \varphi^+ \rangle_b \approx \mp \frac{\pi^{-p/2} \eta^D k_{(0)}^{(p-1)/2} e^{-2x^p k_{(0)}} \cos [|\nu| \ln (4x^p / \eta^2 k_0) + \phi_2]}{2^{p+1} V_q \alpha^{D-1} (x^p)^{(p+1)/2} \sqrt{|\nu| \sinh(\pi|\nu|)}} \nearrow$$

Asymptotic behavior

- At **early** stages of the expansion ($t \rightarrow -\infty$)

$$\langle T_0^p \rangle \approx \mp \frac{2D (\xi - \xi_D) (\eta/\alpha)^D}{(2\pi)^{(p+1)/2} \alpha V_q (2x^p)^p} \sum_{\mathbf{n}_q} f_{(p+1)/2}(2x^p k_{\mathbf{n}_q})$$

- At **late** stages of the expansion ($t \rightarrow +\infty$)

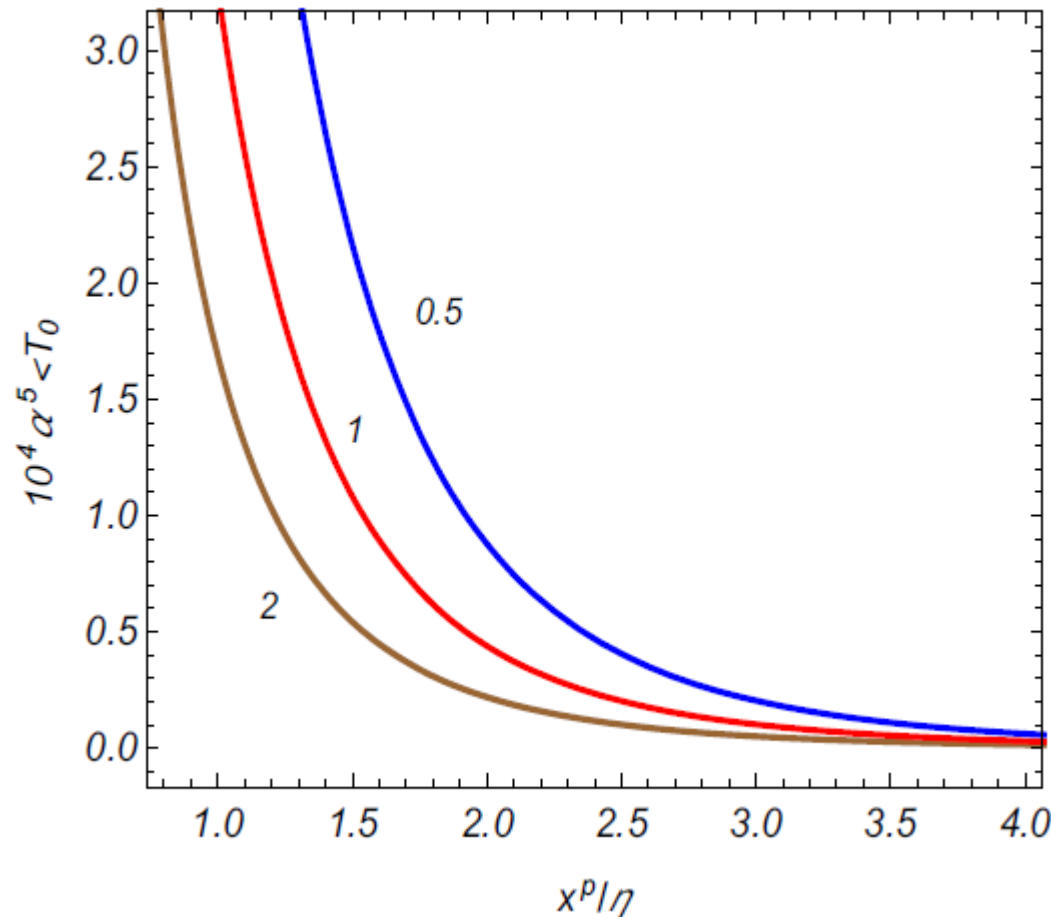
- $\nu^2 > 0$ Decay is **monotonically**

$$\langle T_0^p \rangle \approx \mp 2^{p/2+\nu+2} \Gamma(\nu) (\eta/\alpha)^{D+1-2\nu} \frac{2\nu (1/4 - \xi) + (D+1)\xi - D/4}{(4\pi)^{p/2+1} \alpha^{2\nu} V_q (2x^p)^{p+1-2\nu}} \sum_{\mathbf{n}_q} g_{p/2+1-\nu}(2x^p k_{\mathbf{n}_q})$$

- $\nu^2 < 0$ Decay is **oscillatory**

$$\langle T_0^p \rangle \approx - \frac{(\eta/\alpha)^{D+1}}{2^{p-1} \pi^{(p+1)/2} V_q} \sum_{\mathbf{n}_q} A(x^p, k_{\mathbf{n}_q}) \sin[2|\nu| \ln(\eta/2) - \phi(x^p, k_{\mathbf{n}_q})]$$

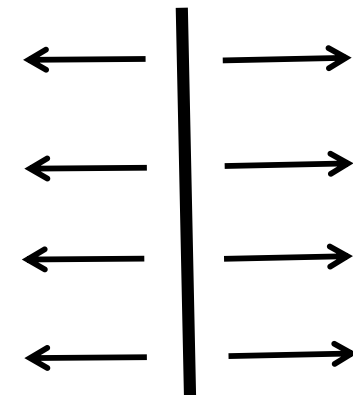
Energy flux



$$D = 4$$

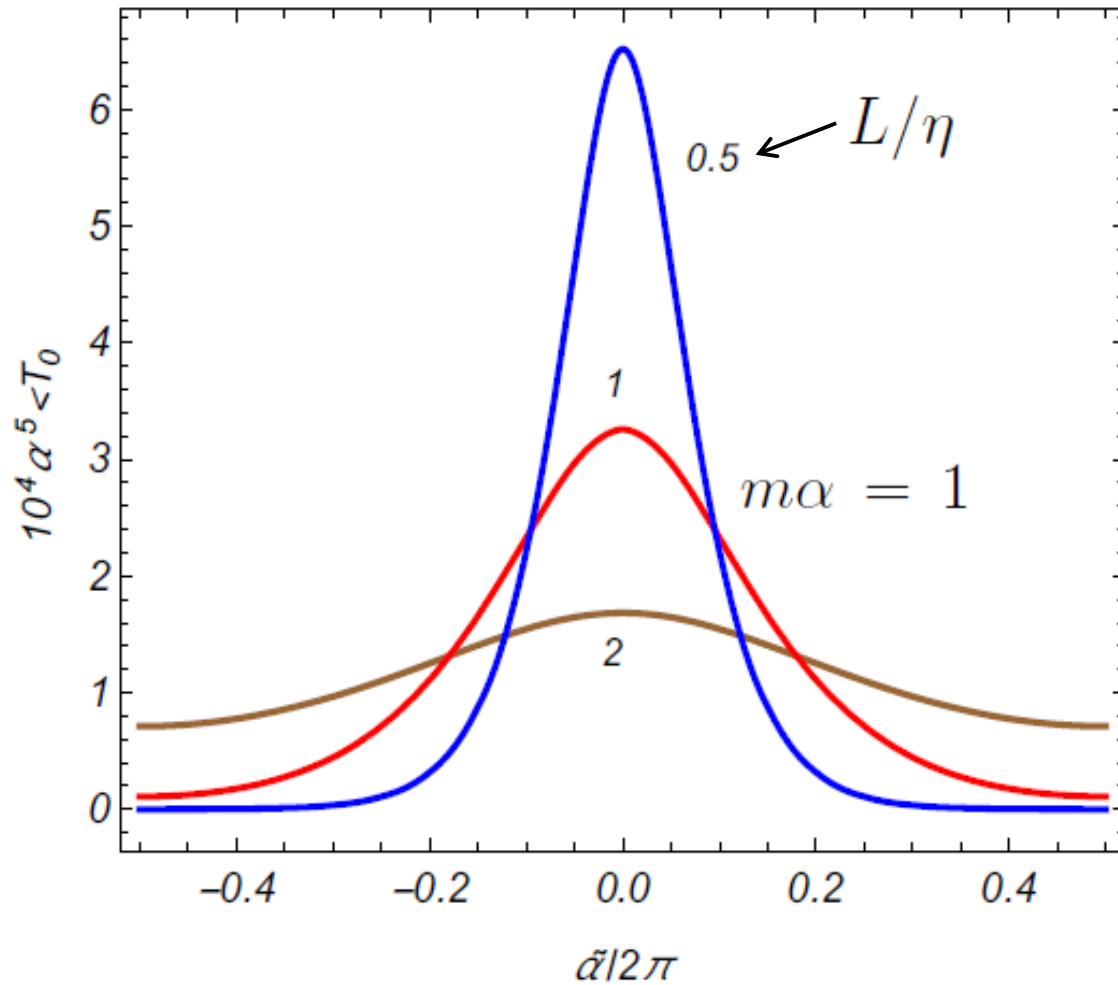
Model with a single compact dimension
Dirichlet boundary condition

$$m\alpha = 1, \tilde{\alpha} = \pi/4$$



Energy flux is directed from the boundary (to the boundary for **Neumann** b.c.)

Energy flux



$$D = 4$$

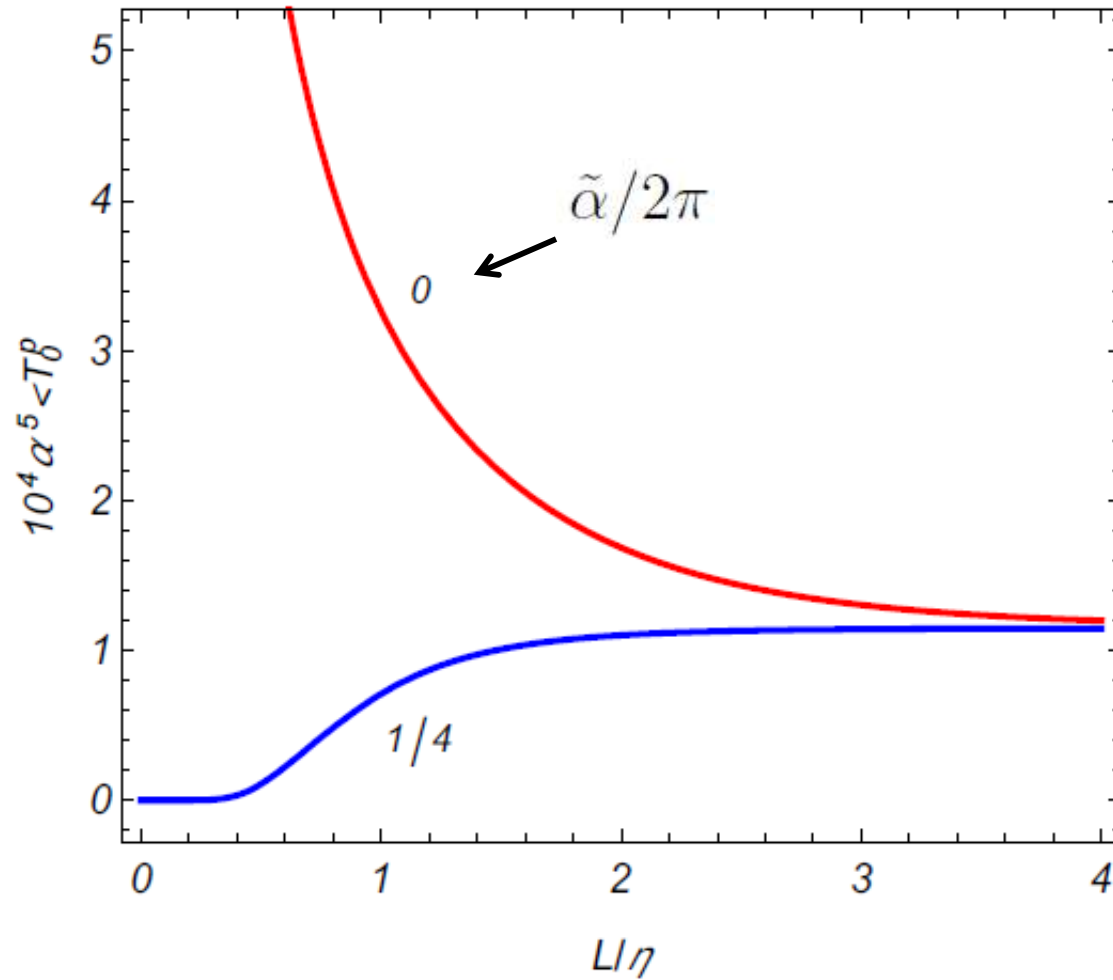
Model with a single compact dimension
Dirichlet boundary condition

$$x^p / \eta = 1 \quad m\alpha = 1$$

Energy flux

$$D = 4$$

Model with a single compact dimension



Conclusion

The presence of a planar boundary give rise to the **energy flux** in the vacuum state along the direction normal to the plate

Vacuum expectation values are decomposed into the **boundary-free** and **boundary-induced parts**

The flux is an **even periodic** function of magnetic fluxes enclosed by compact dimensions with the period equal to flux quantum

Near the boundary the influence of the gravitational field on boundary-induced quantum effects is **weak**

THANK YOU