

**Where we stand on B-decay Discrepancies
– Global Fits –**

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CNRS, Annecy

Minimal TH considerations

(before any fit)

EFT considerations

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well-suited to UV-complete models

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⇒ Accurate $B_s \rightarrow \mu\mu$ measurement

- present single-measurement error $\simeq 20\%$
- exp combi may soon be able to confirm or exclude deviations of C_{10} of $O(10\%)$

More TH considerations

③ *Pattern of Lepton Universality Violation in $b \rightarrow s$*

The observed new-physics hierarchy:

effects in $ee \ll effects in \mu\mu \ll (allowed) effects in \tau\tau$

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Data now allow to disprove some, and to make more precise some other of these considerations

Weak-Effective-Theory

Global fits

1-Wilson-coeff. picture

[Aebischer *et al.*, 2019]

| Coeff. | best fit | 1σ | 2σ | pull |
|---------------------------------------|----------|----------------|----------------|-------------|
| $C_9^{bs\mu\mu}$ | -0.97 | [-1.12, -0.81] | [-1.27, -0.65] | 5.9σ |
| $C_9^{\prime bs\mu\mu}$ | +0.14 | [-0.03, +0.32] | [-0.20, +0.51] | 0.8σ |
| $C_{10}^{bs\mu\mu}$ | +0.75 | [+0.62, +0.89] | [+0.48, +1.03] | 5.7σ |
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| $C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$ | -0.53 | [-0.61, -0.45] | [-0.69, -0.37] | 6.6σ |

Similar fits (w/ non-identical conclusions) performed in:

[Algueró *et al.*, 1903.09578] [Ciuchini *et al.*, 1903.09632]

[Datta *et al.*, 1903.10086] [Kowalska *et al.*, 1903.10932]

[Arbey *et al.*, 1904.08399]

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- Two scenarios stand out: C_9 alone or $C_9 = -C_{10}$ ($\mu\mu$ -channel only)
- $C_9 = -C_{10}$ now better than C_9 alone
- C_{10} alone also ok, but $B \rightarrow K^* \mu\mu$ unresolved

C_9 vs. $C_9 = -C_{10}$

What makes $C_9 = -C_{10}$ more significant than C_9 alone?

$$C_9 \text{ vs. } C_9 = -C_{10}$$

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A concurrence of effects, mostly:

- $B_s \rightarrow \mu\mu$ (new average)
- $\Lambda_b \rightarrow \Lambda \mu\mu$: A_{FB} and BR from LHCb
- $\Delta F = 2$ (mostly ϵ_K and ΔM_s)

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How can ϵ_K (that doesn't depend on C_{10}) increase the C_{10} significance?

Role of correlated quantities

Consider the following two observables

- $O_1 \equiv BR(B_s \rightarrow \mu\mu)$
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Now consider the corresponding terms in the χ^2 function

(with $D_i \equiv O_i^{exp} - O_i^{th}$)

$$\chi^2 = \frac{D_1^2(C_{10})}{\sigma_1^2} + \frac{D_2^2}{\sigma_2^2} + \rho_{12} \frac{D_1(C_{10})D_2}{\sigma_1\sigma_2}$$

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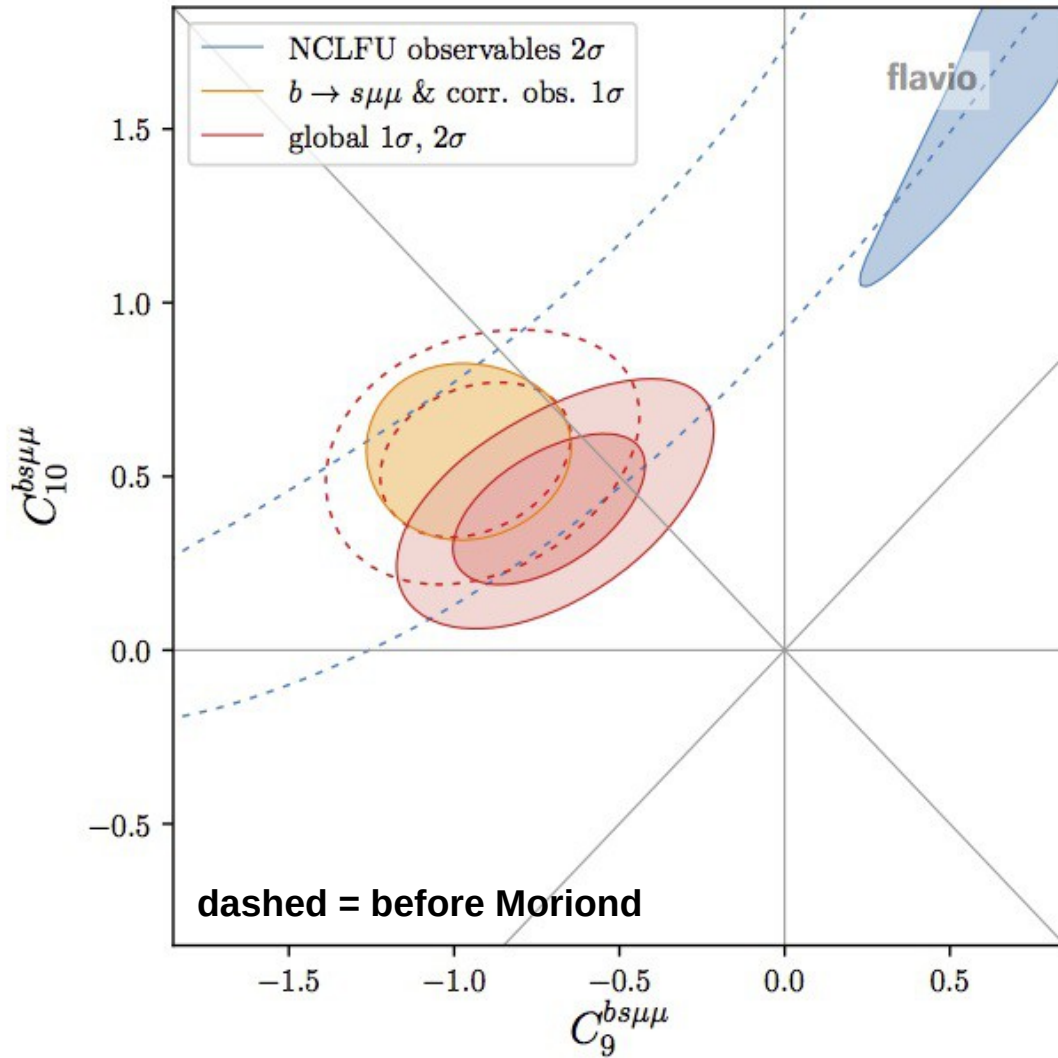
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- The 3rd term depends on C_{10} through D_1
- This term will influence the C_{10} value, depending on $\text{sign}(D_2 \rho_{12})$

C_9 vs. $C_9 = -C_{10}$

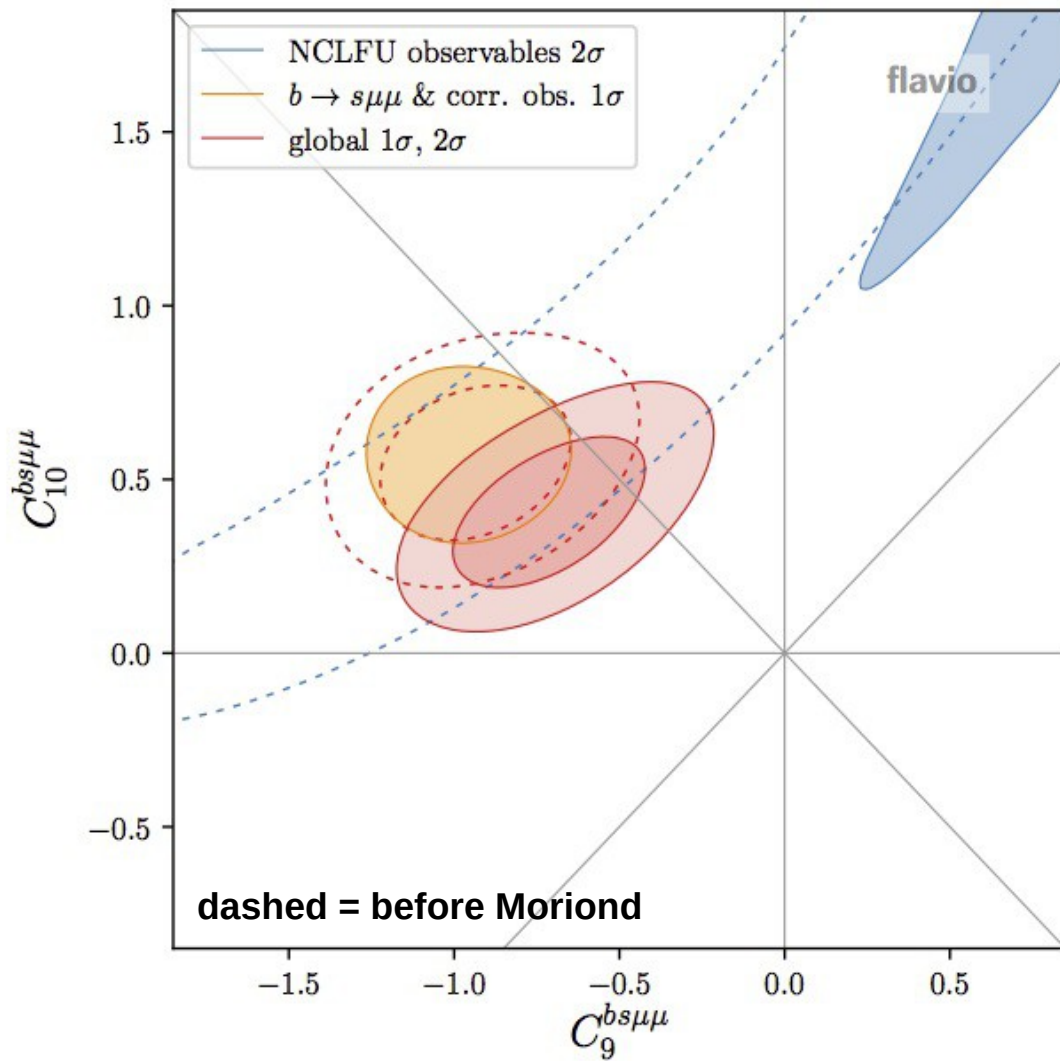


Main point

- **Before Moriond**

$R_{K^{(*)}}$ & $b \rightarrow s\mu\mu$ in nearly perfect agreement

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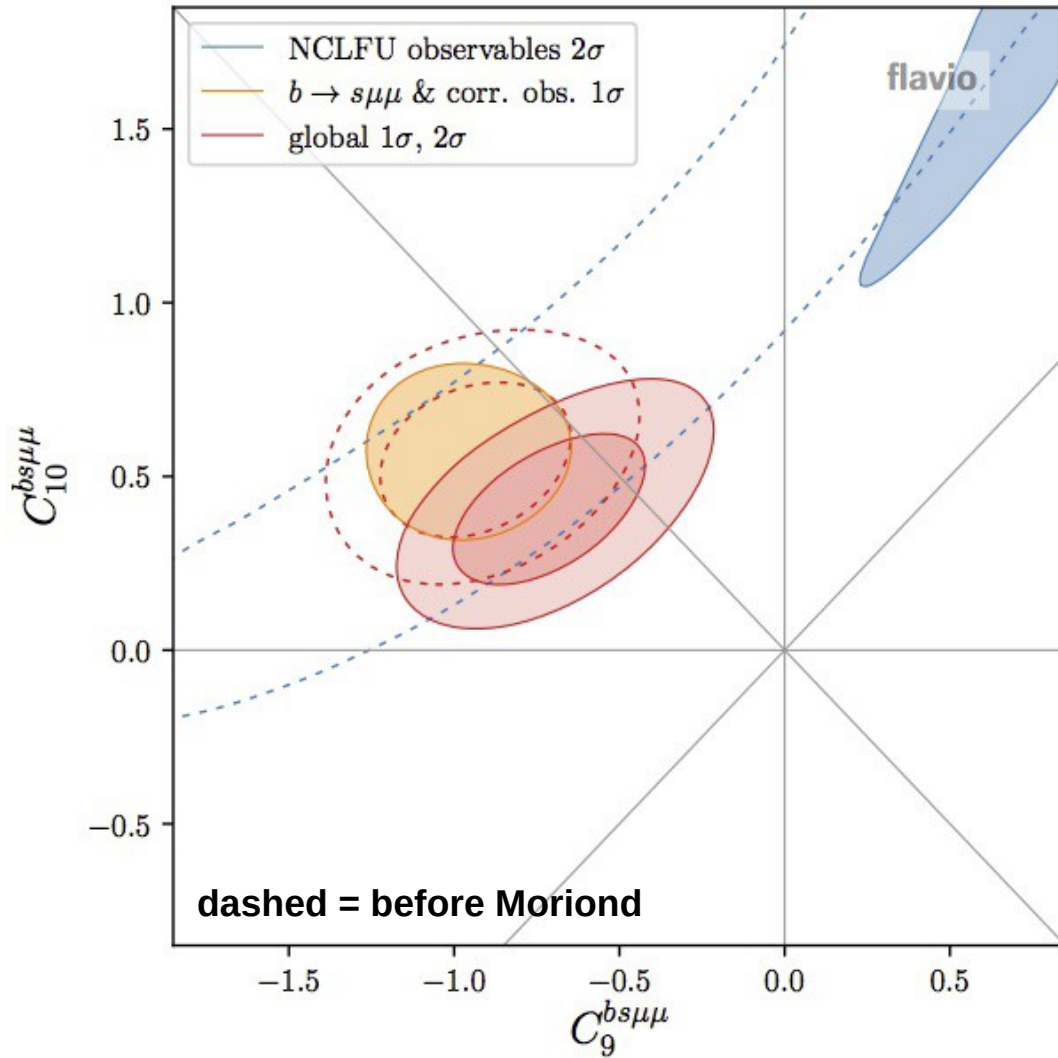
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some tension in C_9 dir.

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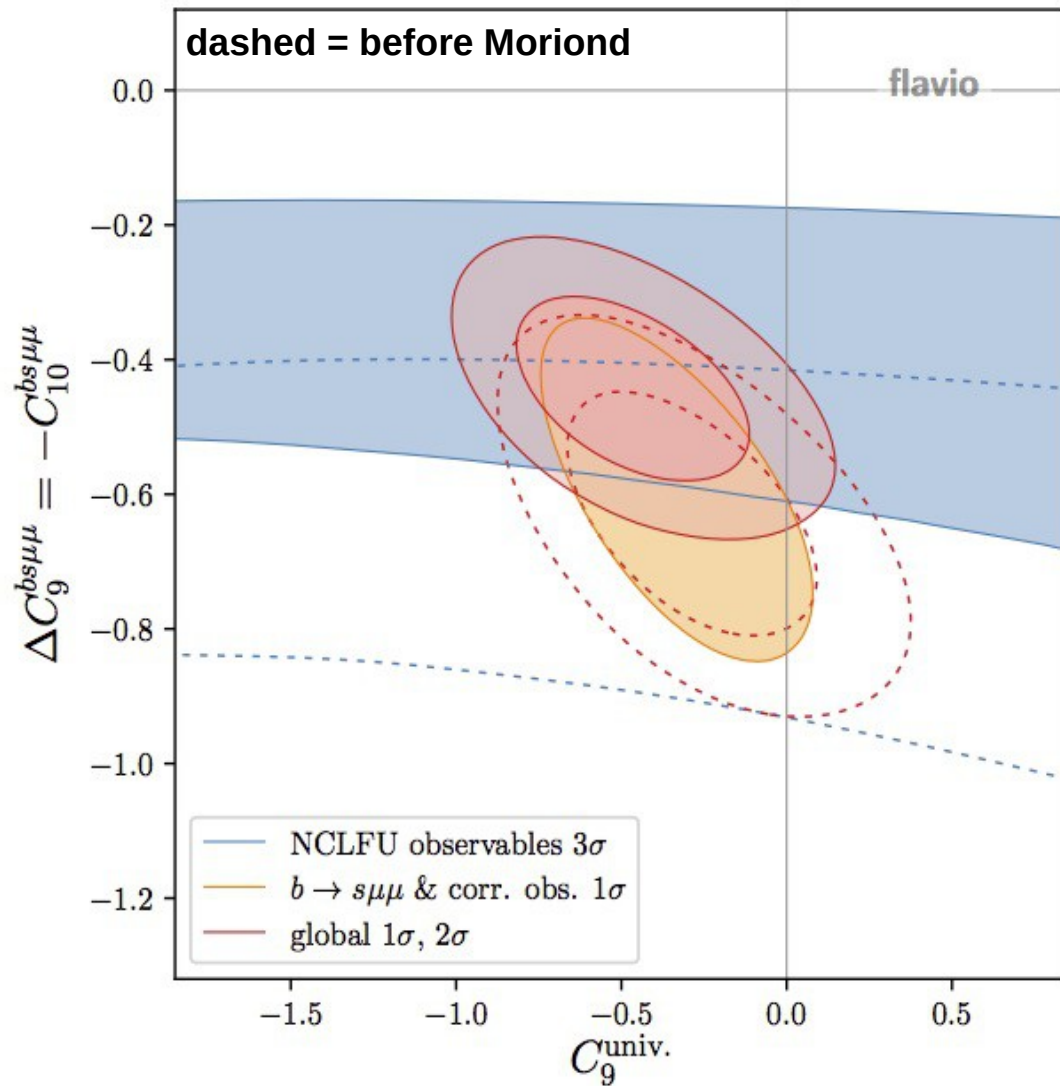
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Note

a $C_9^{univ.}$ component would shift $b \rightarrow s\mu\mu$ data but not $R_{K^{(*)}}$

Univ. vs. non-univ. Wilson coeffs.



Notes

y-axis: μ -specific shift in
 $C_9 = -C_{10}$

x-axis: additional, lepton-univ.
shift in C_9 only

After Moriond

Data tend to prefer $C_9^{univ.} \neq 0$

Both $C_9 = -C_{10}$ and $C_9^{univ.}$

well justified above the EW scale.

The SMEFT picture

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Assume BSM d.o.f. to occur at a scale $\Lambda \gg M_{EW}$



Dynamics below Λ described by ops.

- constructed with SM fields only
- and invariant under the full SM group

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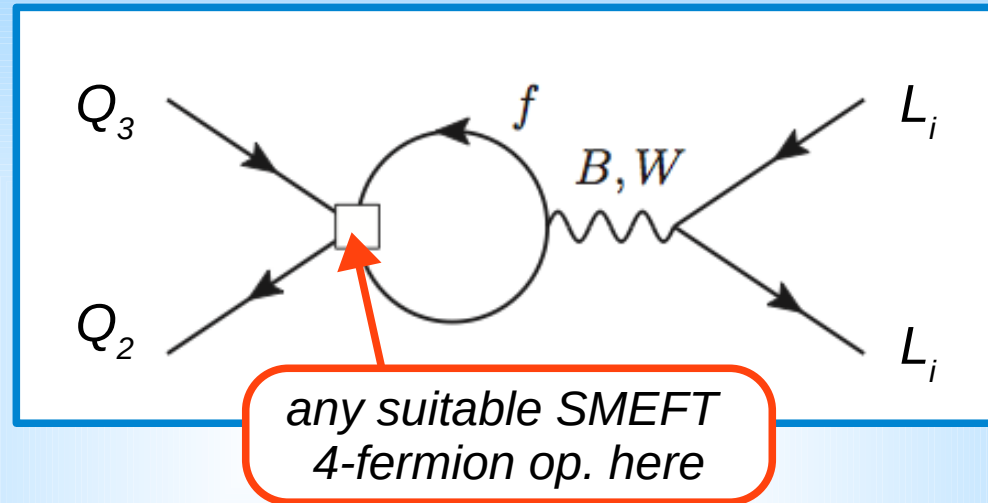
- 1 Contributions to muonic $C_9 = -C_{10}$ may come from SMEFT ops. directly matching onto $O_{9,10}$

$$[O_{LQ}^{(1)}]_{2223} = \bar{L}_2 \gamma^\lambda L_2 \cdot \bar{Q}_2 \gamma_\lambda Q_3$$

$$[O_{LQ}^{(3)}]_{2223} = \bar{L}_2 \gamma^\lambda \sigma^a L_2 \cdot \bar{Q}_2 \gamma_\lambda \sigma^a Q_3$$

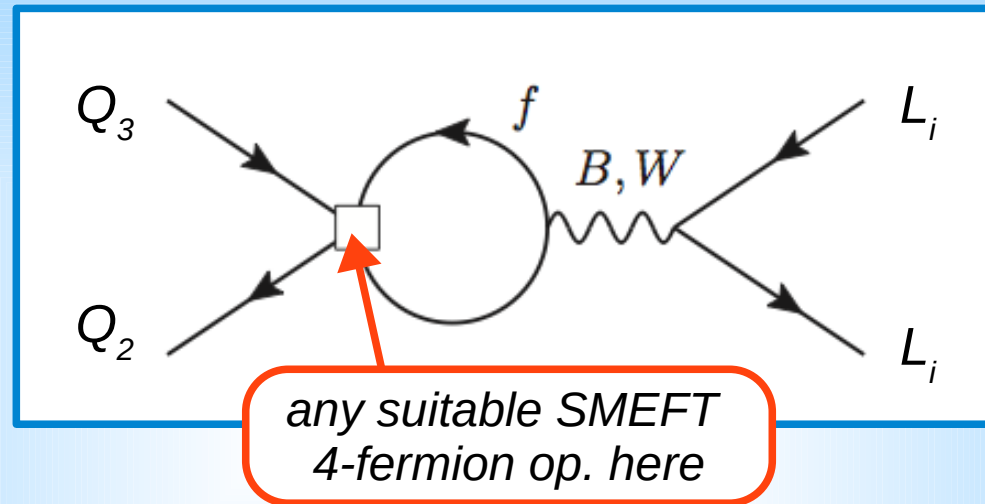
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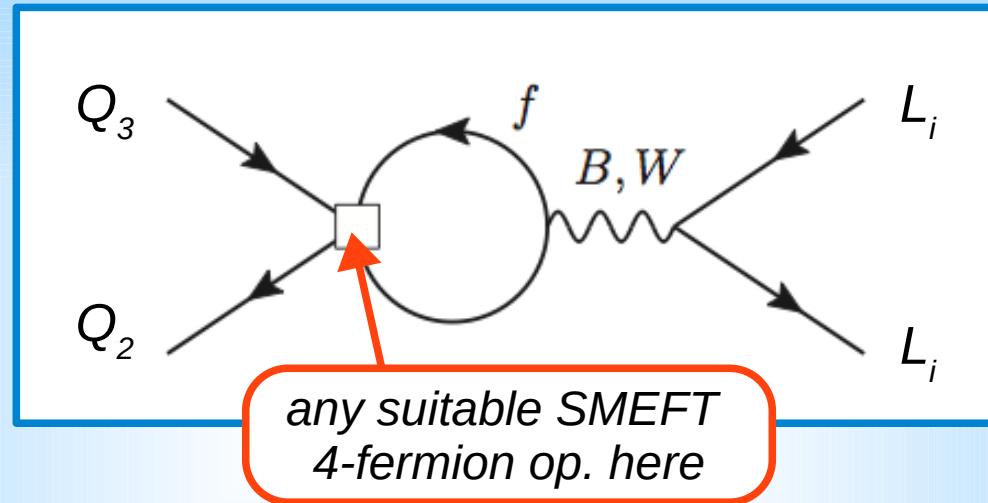
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- Case $f = \tau$ allows natural connection (right sign & size) with $[O_{LQ}^{(1)}]_{3323}$ & $[O_{LQ}^{(3)}]_{3323}$ responsible for $b \rightarrow c \tau \nu$ [Crivellin, Greub, Müller, Saturnino, 2018]

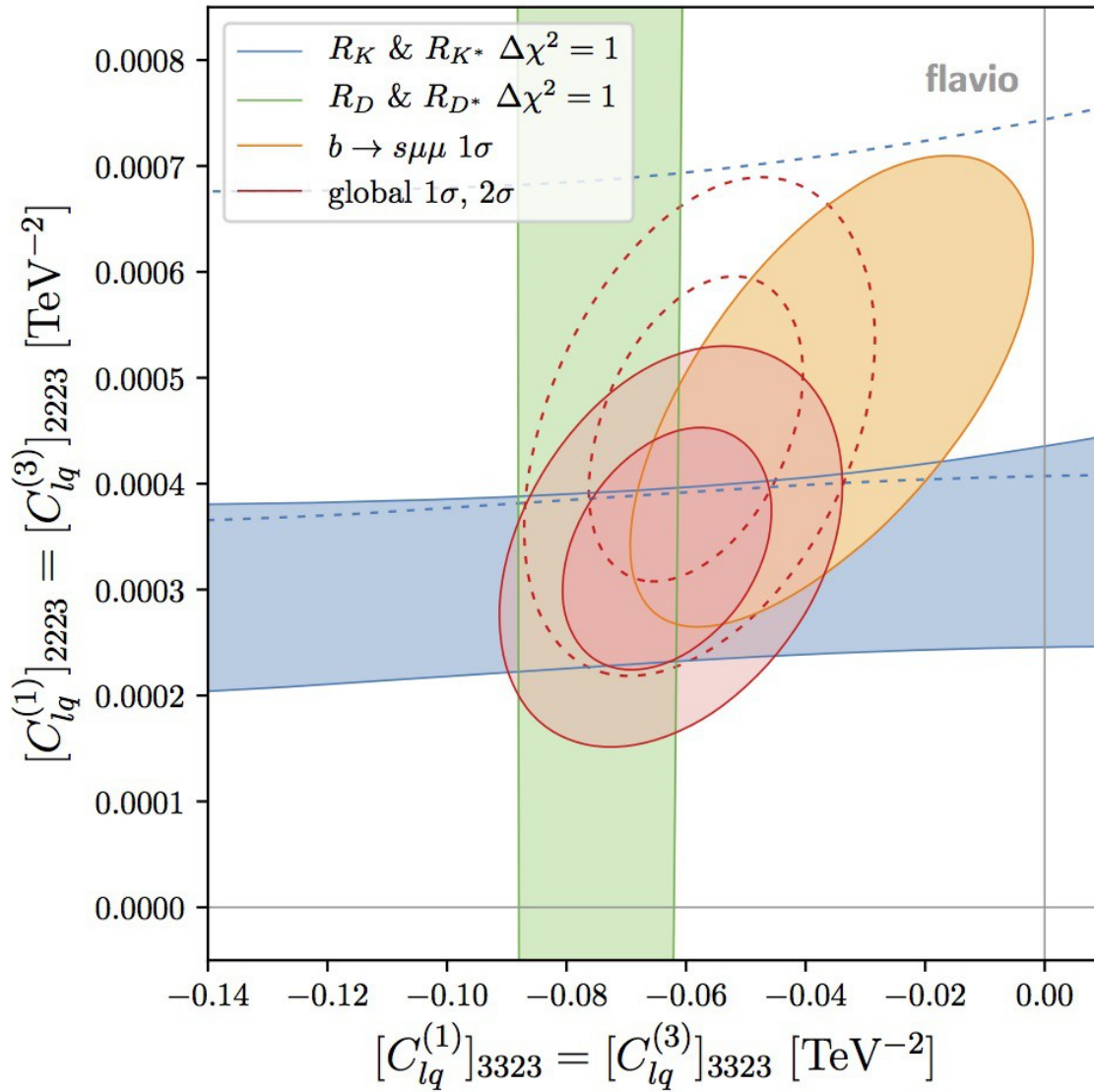
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- *Caveat: need $[C_{LQ}^{(1)}]_{3323} \simeq [C_{LQ}^{(3)}]_{3323}$ to avoid $B \rightarrow K(^*) \nu \nu$ constraint* [Buras-Girrbach-Niehoff-Straub, 2014]

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



Before Moriond (dashed)

$R_{K^{(*)}}$ (blue) and $b \rightarrow s \mu\mu$ (orange) were in perfect agreement

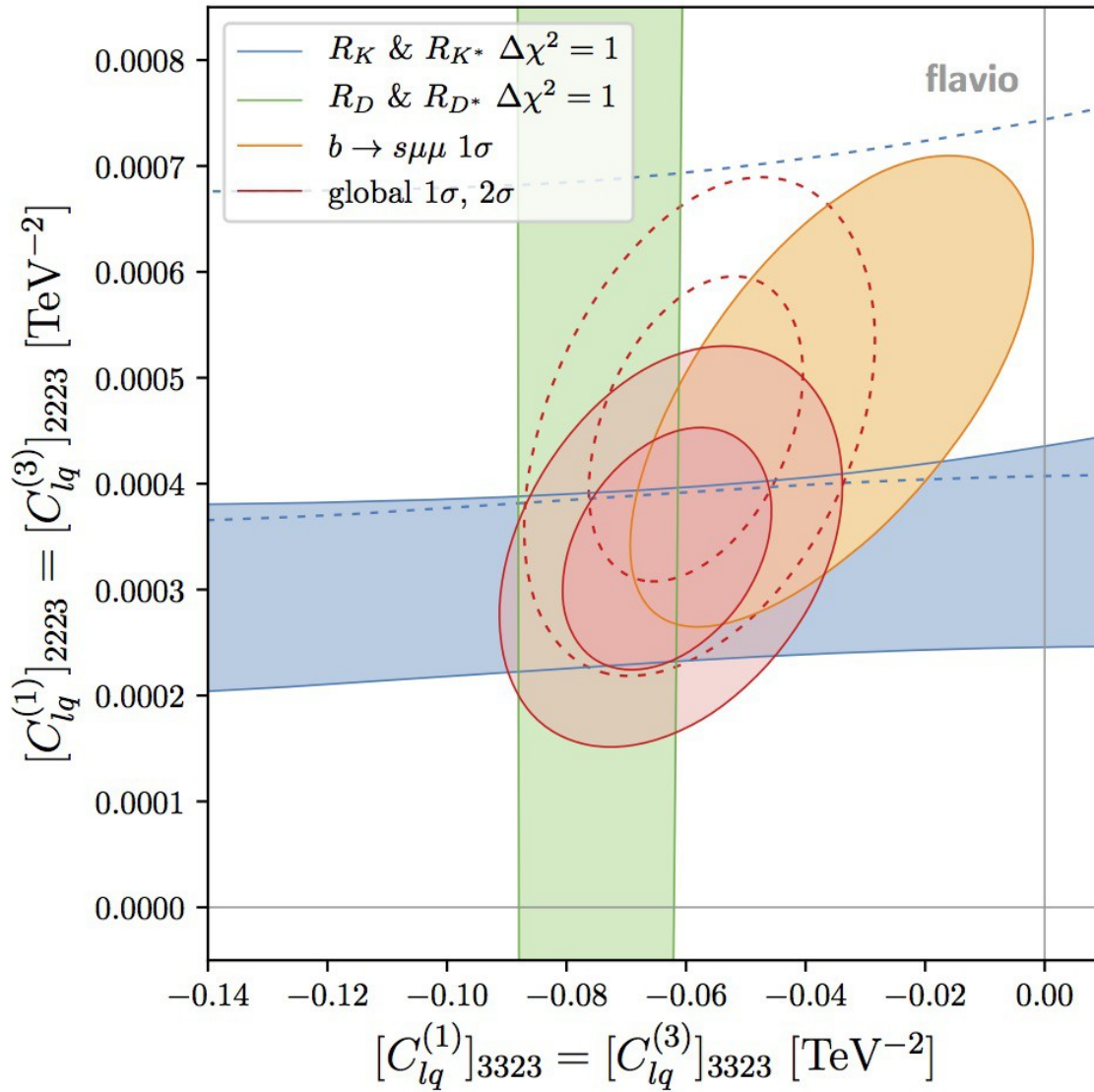


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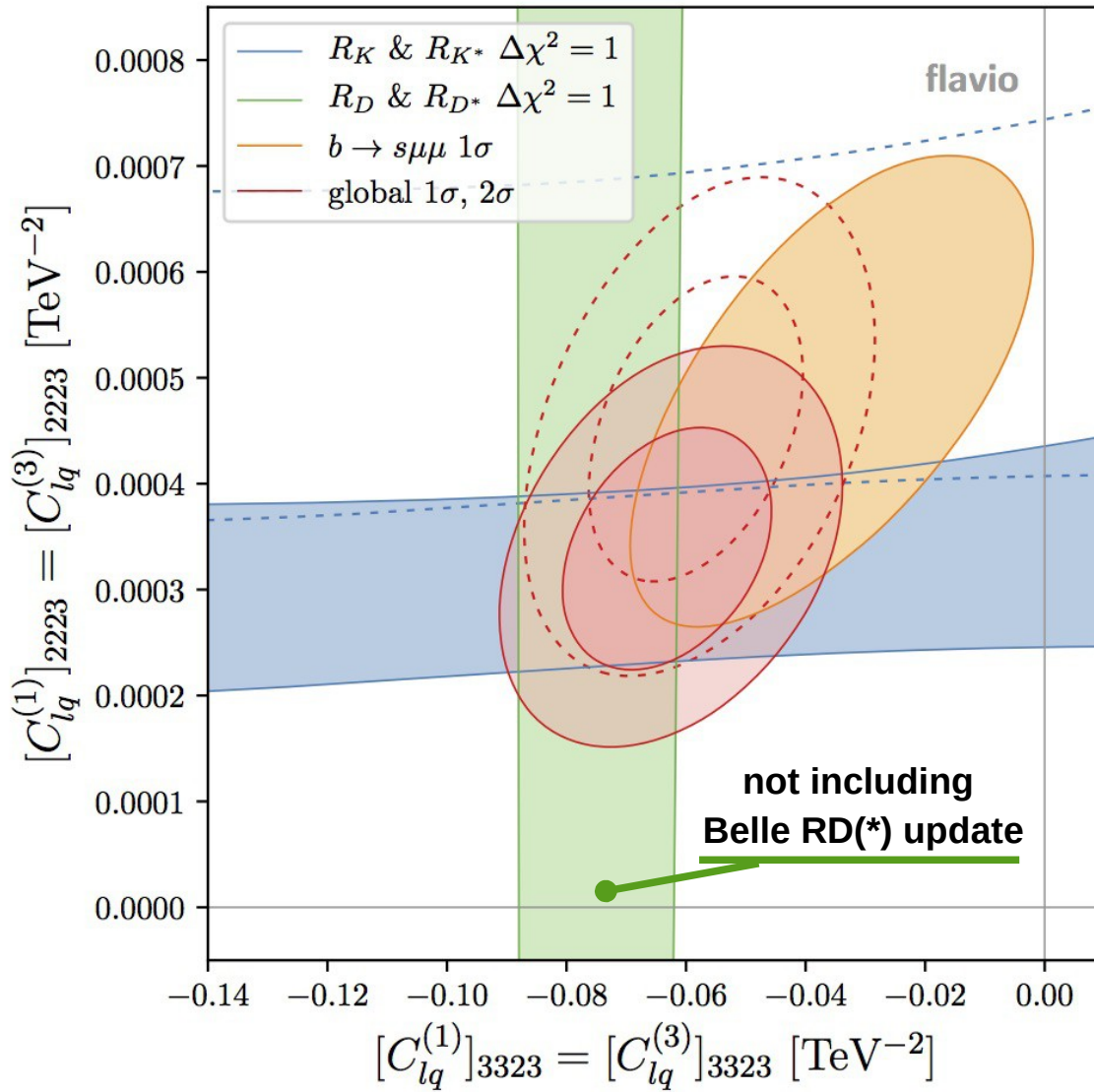


in a region with x-axis values well below 0



Quantitative agreement with the $R_{D^{(*)}}$ constraint (green)

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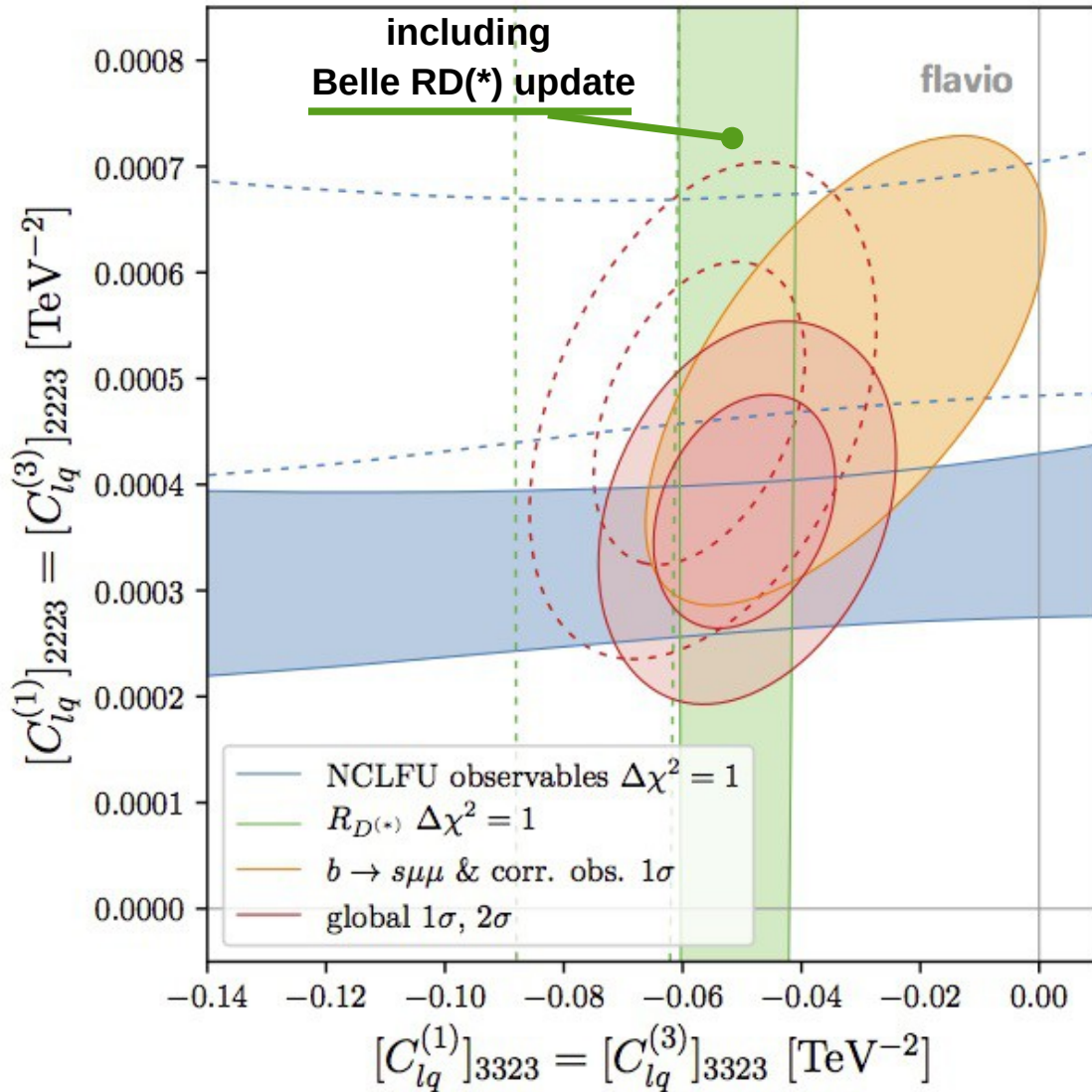


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Conclusions

Post-Moriond updates imply a nicely coherent TH picture, with

- $R_K & R_{K^*} <$ by $O(20\%)$ than SM
- $R_D & R_{D^*} >$ by $O(10\%)$ (not more) than SM
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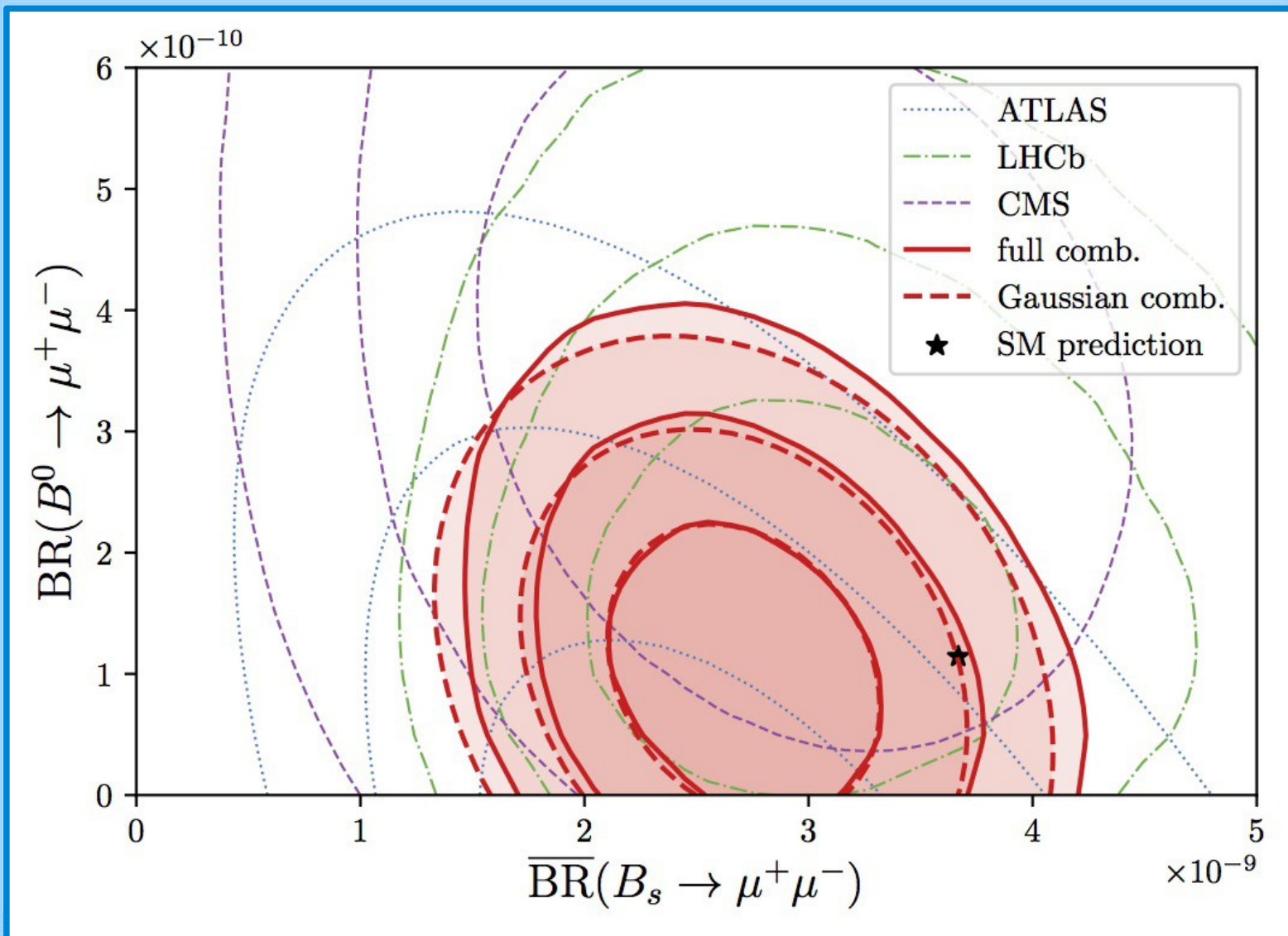
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We'll know soon (?) whether this is all just a happy coincidence

$B_s \rightarrow \mu\mu$ average

w/o CMS PAS BPH-16-004 (Aug. 2019 update)

Credit: Peter Stangl



$B_s \rightarrow \mu\mu$ average

with CMS PAS BPH-16-004 (Aug. 2019 update)

Credit: Peter Stangl

