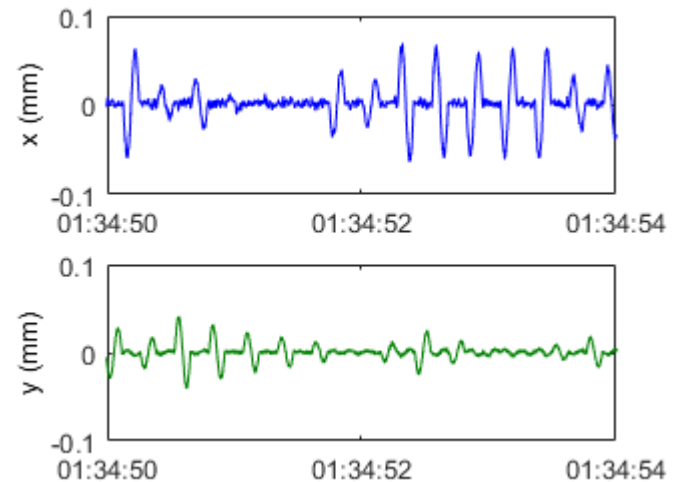
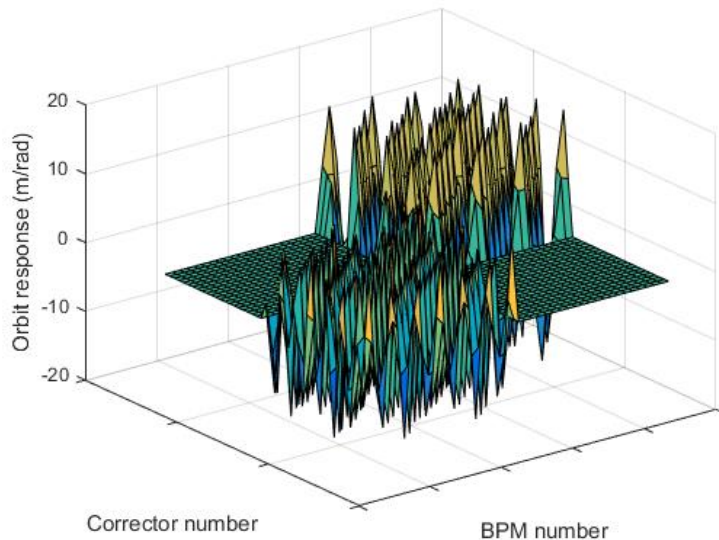


# Optics Measurements using Fast Orbit Feedback Data



*Ian Martin*

*With thanks: G. Rehm, M. Furseman, V. Smaluk, Z. Martí, A. Franchi*

Workshop on Beam Tests and Commissioning of Low Emittance Storage Rings

Karlsruhe Institute of Technology

20<sup>th</sup> February 2019

# Talk Outline

## *Introduction:*

why use fast orbit data for LOCO?

## *Streamlined LOCO procedure:*

fast orbit response matrix measurements  
choice of excitation frequency  
Python implementation  
status at Diamond

## *Examples from other facilities:*

NSLS-II: algorithm comparisons; multi-corrector excitation  
ALBA: off-energy measurements; non-linear lattice  
ESRF: ID coupling compensation

## *Summary*

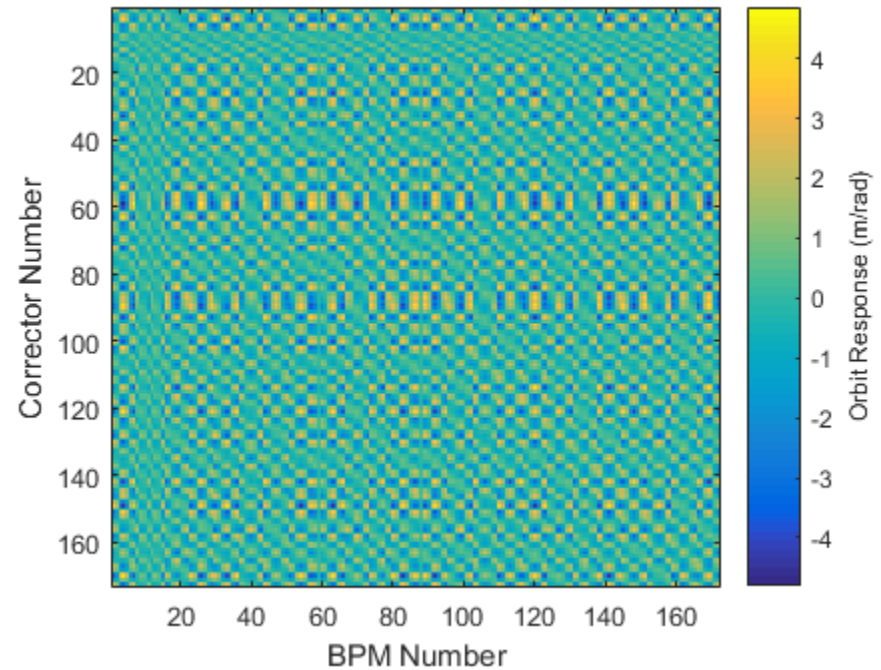
# Introduction

Excellent control of linear optics is mandatory for latest generation of light sources (source size, tune-stability, resonance control, coupling, vertical dispersion, lifetime, inj. efficiency, ...).

Many techniques exist, however, LOCO[1]-style algorithms based on closed orbit response matrix (ORM) measurements are typically applied.

## The problem:

- ORM large (Diamond-II: 252 BPMs x 252 CM x 2 planes); can take ~1h per correction cycle
- Significant drift can occur during measurement (particularly after fresh injection)
- Orbit stability can affect accuracy
- Invasive, so cannot be measured during user beam



# Introduction

## Standard LOCO Method

- 1) Measure ORM (+Dispersion, +BPM noise)
  - Correctors stepped up/down via matlab script over EPICS (MML)
  - Wait for magnet to reach set-point + fresh SA-BPM data (~1s)
  - Cycle through each HCM then VCM
- 2) Convert measured data to LOCO input file
- 3) Run LOCO analysis
- 4) Apply results

## Standard parameter fit for Diamond-I:

- Individual quadrupole gradients (248)
- Individual skew-quadrupole gradients (98)
- BPM and CM gains / rolls

Typically takes ~20 minutes to acquire measured data, plus ~15-20 minutes to complete processing and to apply corrections.

# Orbit Response Matrix Measurement

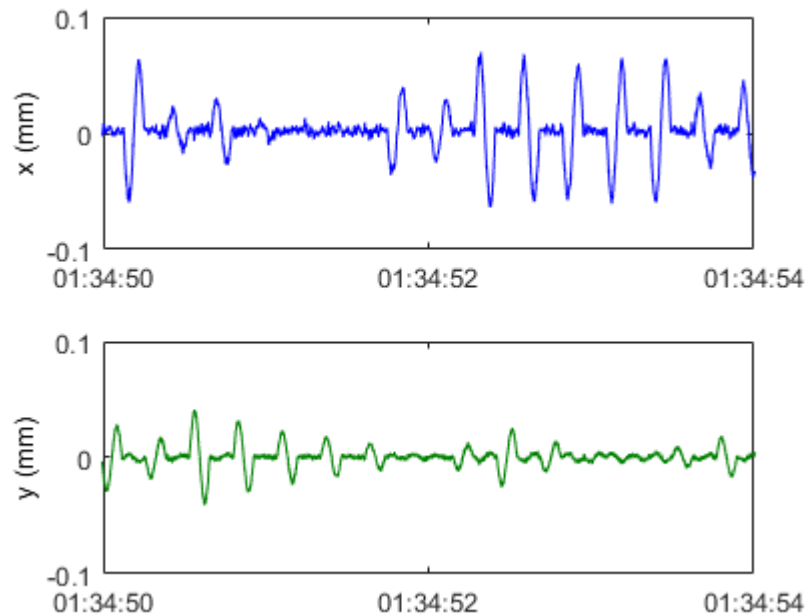
## Implementation

- Each feedback node can produce sine-wave excitation on each corrector with programmable amplitude, frequency, duration and synchronised start time
- Configured using Python script
- Orbit data extracted from 1 kHz FA data stream
- Amplitude extracted from measured orbit at BPM  $m$  and corrector  $n$ :

$$A_{m,n} = \langle 2 \times z_m(t) \times \sin(2\pi f t) \rangle$$

$$R_{m,n} = A_{m,n} / \theta_n$$

- Number of cycles, excitation frequency, choice of correctors, delay between corrector all configurable
- Particular choice depends on context (e.g. low alpha mode, fast coupling correction, ...)



# Introduction

## Now

All stages automated via python interface, with Matlab stages launched in batch mode [2].

- 1) Measure ORM using fast orbit feedback network [3]
  - Programmed sine-wave excitation applied to each corrector in turn
  - Orbit data taken from fast acquisition (FA) BPM data and post-processed
  - Horizontal then vertical alternated
- 2) Convert measured data to LOCO input file
- 3) Run LOCO analysis (uses Matlab Parallel Computing Toolbox)
- 4) User queried if they would like to apply the results

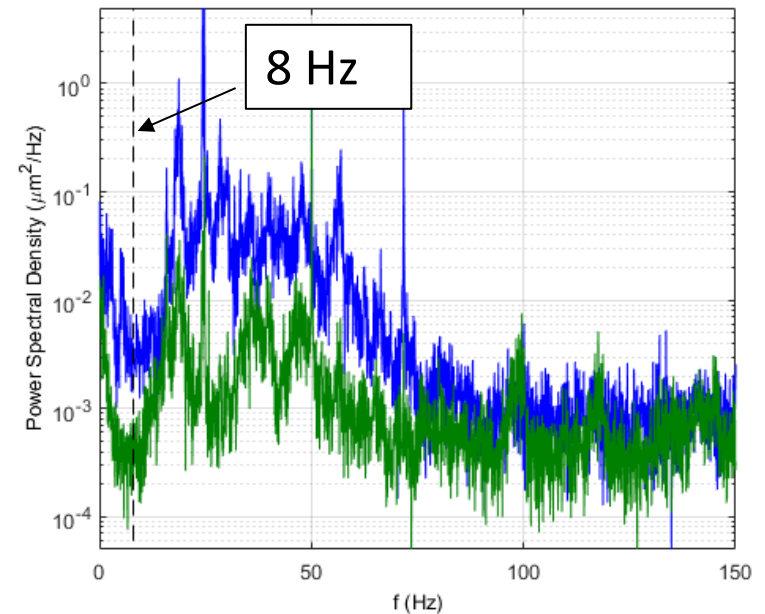
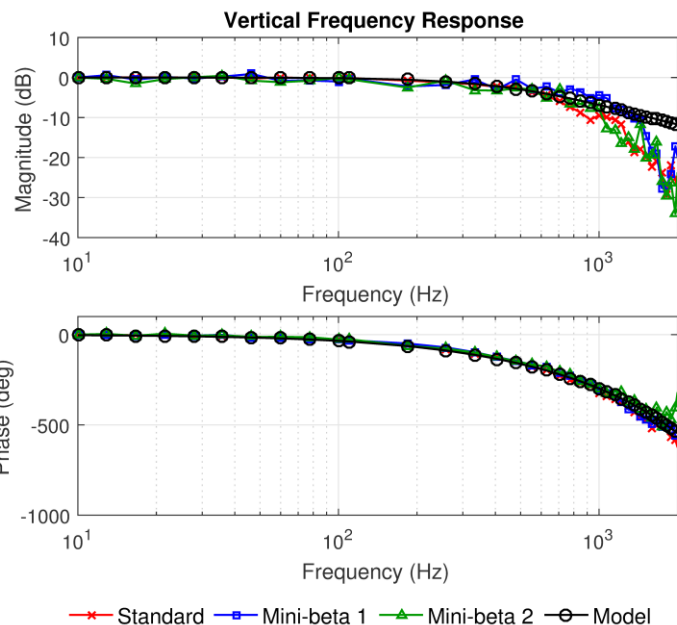
Keep the same parameter fits as previously

ORM measurement takes 52 seconds. Complete correction cycle ~5 minutes.

# Choice of Excitation Frequency

Choice of excitation frequency is a trade off between reducing measurement time and improving accuracy

Noise spectrum at Diamond suggests 8 Hz is optimal



Freq. response of correctors affects measurement

Depends on whole system (magnet, PS, vacuum chamber, ...)

Currently using 8 Hz for embedded correctors on sextupoles, 1 Hz for discrete correctors in mini-beta straights



# Python Interface

**PLOCO**

**Orbit**

FOFB Off ☒ SOFB Off ☒ RFFB Off ☒  
Orbit OK ☒ Std. X 1.0031 Std. Y 0.8335

**Tunes**

TFB Off ☒ Actual X nan Actual Y nan  
Tunes OK ☒ Target X 0.1890 Target Y 0.2770

**Current**

Top Up Off ☒ MTGen 101.01  
Current OK ☒ DCCT -0.00

**Vertical Emittance**

VEFB Off ☒ Target 8.00 Actual 7.99

Custom Configuration ...

Default

Cancel Run

**PLOCO - Custom Configuration**

File Help

Dataroot /dls/ops-physics/diamonddata

Norm Freq (Hz) 8.0

Slow Freq (Hz) 1.0

Cycles 1.0

Excite Gap (0.1 ms) 100

Num Std 100

Num Iters 5

Use Wigglers ☐ I12 ☐ I15

Fit Param 27

SVD Threshold 0.0002

Weights 0.001 0.0005 0.001

Correctors Add Remove

	MB1	MB2	1	2	3	4	5	C1	C2	6	7	8	10
1			1	2	3	4	5			6	7		
2			8			9	10	11	12		13	14	15
3			16	17	18	19	20			21	22		
4			23	24	25	26	27			28	29		
5			30	31	32	33	34			35	36		
6			37	38	39	40	41			42	43		
7			44	45	46	47	48			49	50		
8			51	52	53	54	55			56	57		
9	58	59	60	61	62	63	64			65	66		
10			67	68	69	70	71			72	73		
11			74	75	76	77				78	79		
12			80	81	82	83	84			85	86		
13	87	88	89	90	91	92	93			94	95		
14			96	97	98	99	100			101	102		
15			103	104	105	106	107			108	109		
16			110	111	112	113	114			115	116		
17			117	118	119	120	121			122	123		
18			124	125	126	127	128			129	130		
19			131	132	133	134	135			136	137		
20			138	139	140	141	142			143	144		
21			145	146	147	148	149			150	151		
22			152	153	154	155	156			157	158		
23			159	160	161	162	163			164	165		
24			166	167	168	169	170			171	172		

Apply Cancel

**console.log (/dls/ops-physics/diamonddata/DIAD/LOCO/)**

File Edit View Search Tools Documents Help

Open Save Undo

console.log X

```
Done
  BPM data saved to /dls/ops-physics/diamonddata/DIAD/
  BPMDData_19-01-07_18-24-49.mat
  The total measurement time was 0.23 minutes.

Measuring Dispersion

=====
Generating amplitude files.
=====

Done.

=====
Measuring Fast Response Matrix.
=====

Connecting to PVs
Building excitations
Getting current fatimestamp
Starting excitation
Now 512 ticks behind timestamp
Excitation will complete in 57.12 seconds
data.shape (57727, 174, 2)
Calculating response matrix from data
rms.shape (2, 2, 172, 173)
Removing bad bps from response matrix
rms.shape (2, 2, 172, 171)
Saving response matrix
[[[71h[9]=

=====
Converting matrix from raw format.
=====

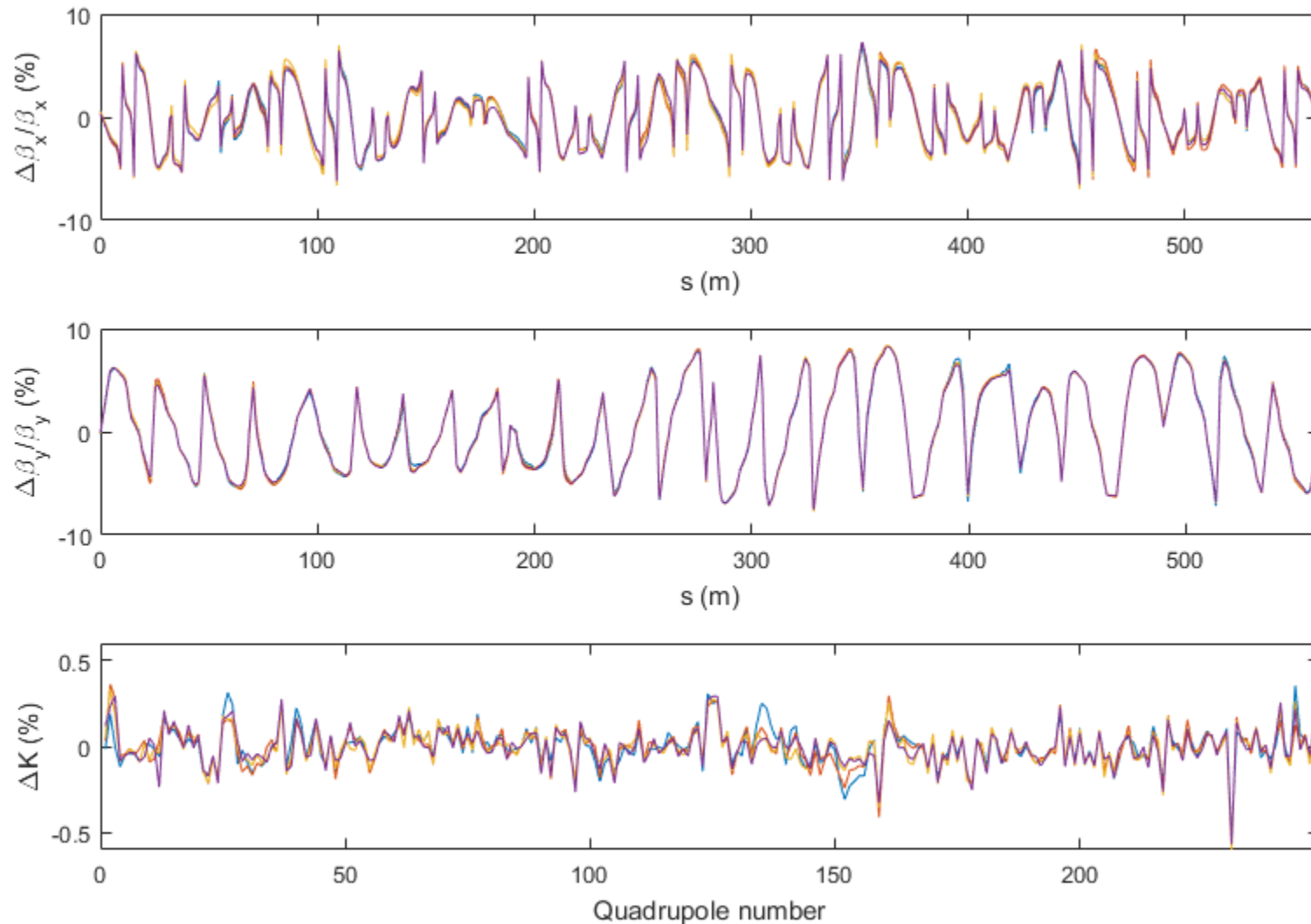
Done.

=====
Combining data into loco file
=====
```

Plain Text Tab



# Results



— Standard LOCO

— Fast ORM #1

— Fast ORM #1

— Fast ORM #1

# Status

- New LOCO application has been in routine operation since 2014
- Used by anyone (no specialist knowledge). Now applied for many situations:
  - Normal/skew quadrupoles for ID compensation
  - Re-correct the coupling (after beam trip)
  - Correct machine for MD studies (injection studies, resonant spin, pinhole calibration, ID studies, tune scans, ....)
- Different configurations in use for different operating modes:
  - User optics: single cycle, 100 samples for BPM sigma
  - Low alpha: 5 cycles, 200 samples for BPM sigma
- Parallelisation in Matlab allows increased number of iterations during LOCO fit to improve the convergence
- Tests with higher excitation frequencies show comparable results for optics; main impact is reduced fitted gain for correctors (attenuation / phase delay)
- Can further reduce acquisition and fit times by using fewer correctors; main impact is again on gain / roll values for correctors
- Using multiple excitations in parallel at different frequencies has been demonstrated. Can measure complete ORM in <10 s, however, it is the post-processing of the measured data that is limiting factor at present.

# NSLS-II: Algorithms Cross-Check

LOCO (DC and AC) results compared with various turn-by turn (TBT) schemes [4]:

4 TBT based algorithms:

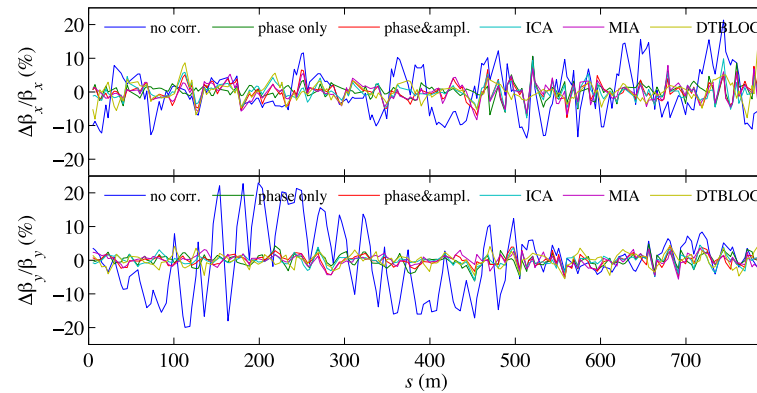
- weighted correction of betatron phase and amplitude [5]
- independent component analysis [6]
- model-independent analysis [7]
- driving-terms-based linear optics characterization [8]

2 orbit-based algorithms:

- standard (DC) LOCO [1]
- AC LOCO [2, 9]

⇒ LOCO-based algorithms measurement precision higher

⇒ TBT-based algorithms less time-consuming



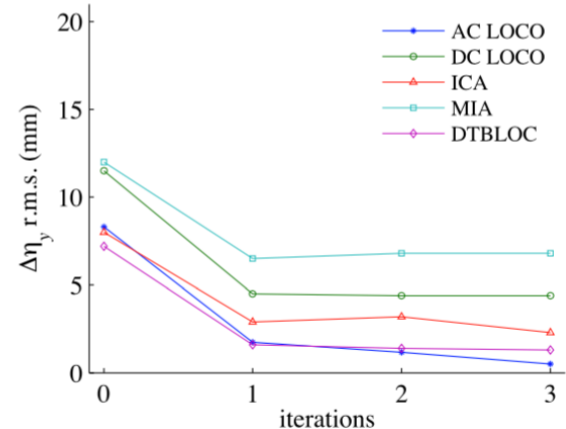
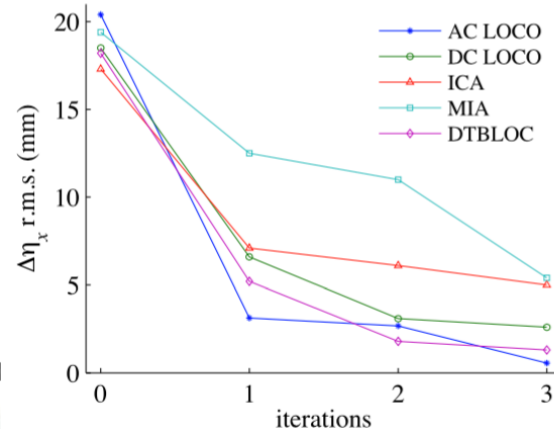
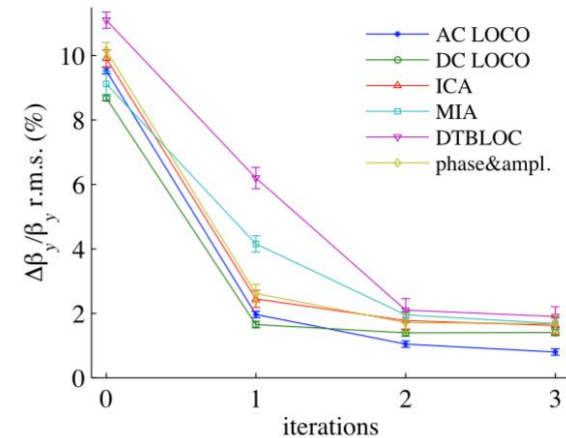
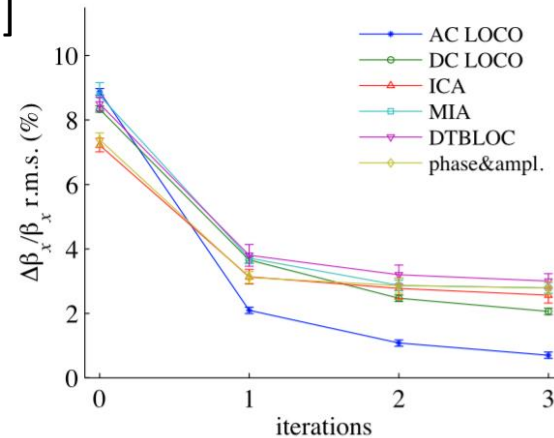
“Bad” lattice (random errors added to NSLS-II quadrupoles):

$$\Delta\beta_x/\beta_x = 8\% \text{ r.m.s.}$$

$$\Delta\beta_y/\beta_y = 10\% \text{ r.m.s.}$$

$$\Delta\eta_x = 18 \text{ mm r.m.s.}$$

$$\Delta\eta_y = 8 \text{ mm r.m.s.}$$



Ian Martin, I

# NSLS-II: Multi-Frequency Excitation

- Fast corrector signal bandwidth is small (0.2 Hz for a 5 sec measurement), it provides an opportunity of simultaneous excitation of beam oscillations via multiple fast correctors with different frequencies separated by an interval of  $\Delta f$ .
- Can potentially reduce the measurement time to be comparable with TBT-based methods
- Frequency range for the multiple excitations depends upon on the frequency-dependent signal-to-noise ratio of the system.

- Beam oscillation measured by BPMs is a finite-time sine wave, Fourier transform of which is proportional to:

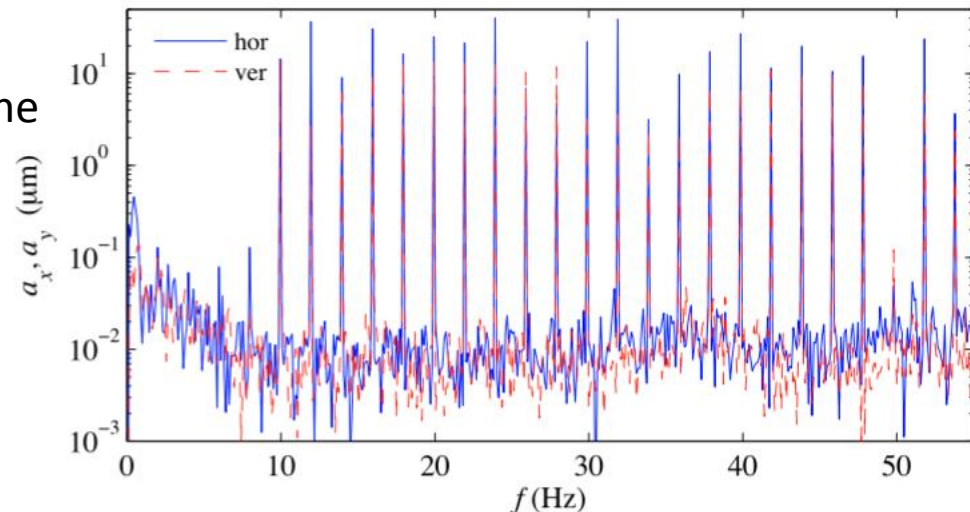
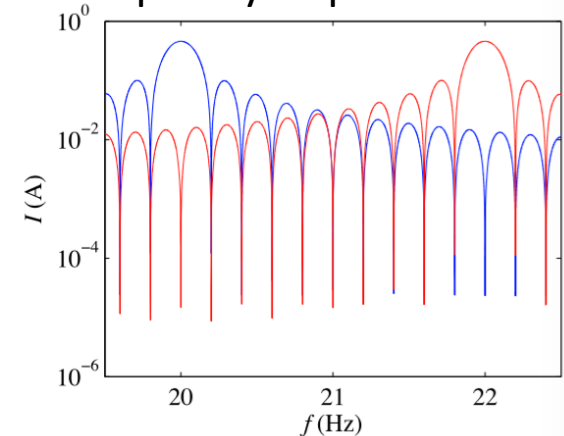
$$\text{sinc}(\pi T \Delta f) = \sin(\pi T \Delta f) / \pi T \Delta f$$

where  $\Delta f = f - f_0$ , and  $f_0$  is the excitation frequency.

- This function has zero values at  $\Delta f = k/T$ , where  $k$  is an integer, so we can choose any of these frequencies.

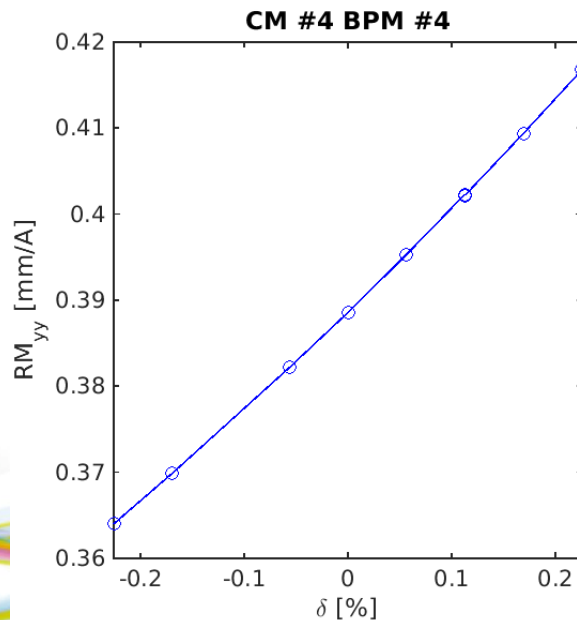
NSLS-II:  $T = 5$  s,  $k = 10 \Rightarrow \Delta f = 2$  Hz

- 23 fast correctors, AC driving signals with the frequencies of 10 Hz; 12 Hz; ...; 54 Hz;
- Slow orbit feedback on to minimize drift
- 10 measurement sets for statistical errors
- Measured r.m.s. errors  $\sim 20$  nm

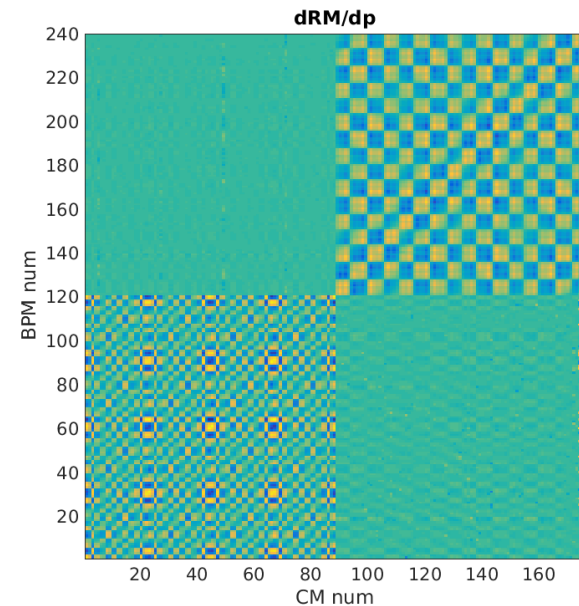


# ALBA: Off-Energy Fast ORM

- Increased accuracy allows  $\partial ORM / \partial \delta$  to be evaluated by measuring  $ORM$  at several values of RF frequency [10]
- Effective only for sextupoles at dispersive locations
- Discrepancies found to persist when fitting the sextupoles against the LOCO model



**FORM** element as a function of  $\delta p/p$  induced by a RF frequency change.



**FORM derivate** with respect to  $\delta p/p$ . Could be used to fit non linear fields.

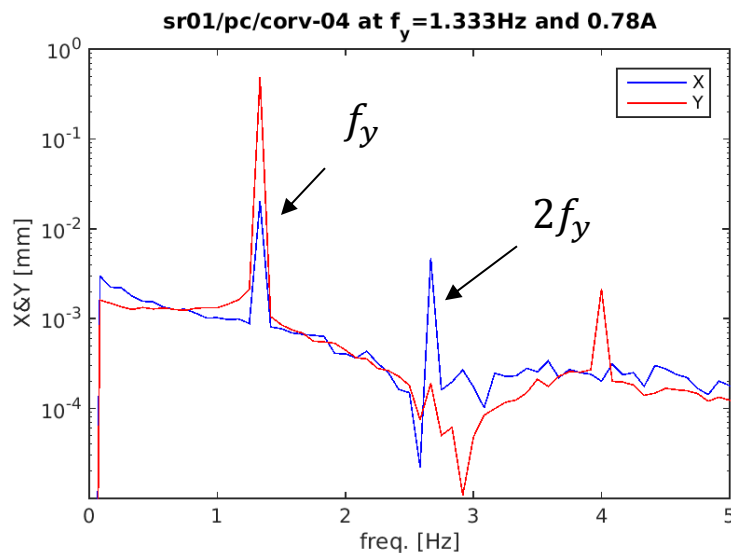
# ALBA: Non-Linear Fast ORM

- Sextupolar fields induce harmonics of the CM waveform frequency ( $ORM^2$ )
- Excite the beam at  $f_x$  and  $f_y$ :

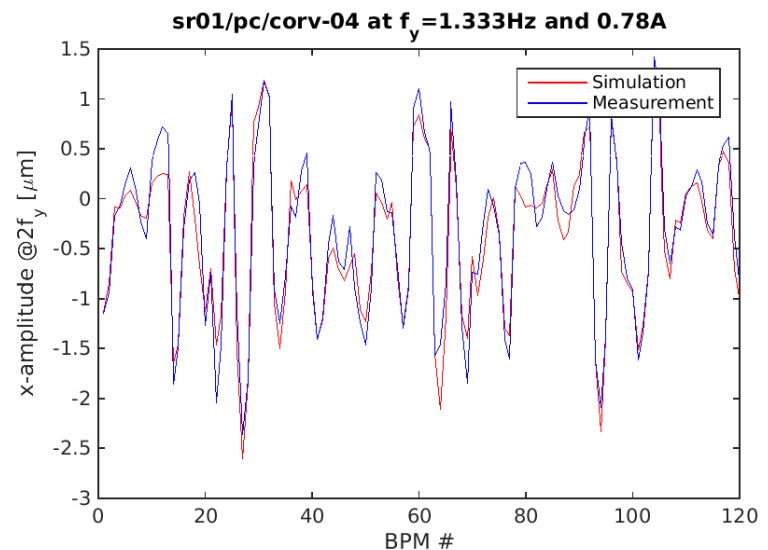
$$B_y = m(x^2 - y^2) \quad \text{- horizontal orbit contains } 2f_x \text{ and } 2f_y$$

$$B_x = 2mxy \quad \text{- vertical orbit contains } f_x - f_y \text{ and } f_x + f_y$$

- Requires large amplitude excitation to be visible
- Fit of the measured quantities against the model not yet attempted



**Discrete FFT** of the BPM signals for a particular single vertical CM.



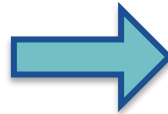
**Amplitude of the  $2f_y$**  line for the horizontal BPMs.



# ESRF: ID Compensation

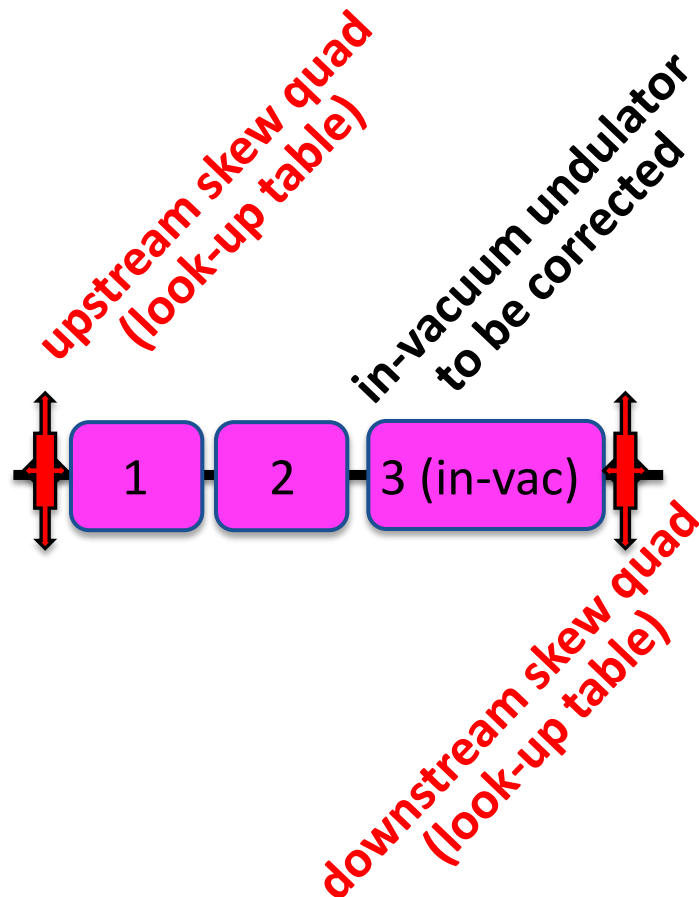
Measure and fit AC ORM to infer ID gap-dependent skew quad field, calibrate corrector coils and check look-up tables

$$\delta O_{xy} = O_{xy}^{(corr. coils)} - O_{xy}^{(corr. off)}$$

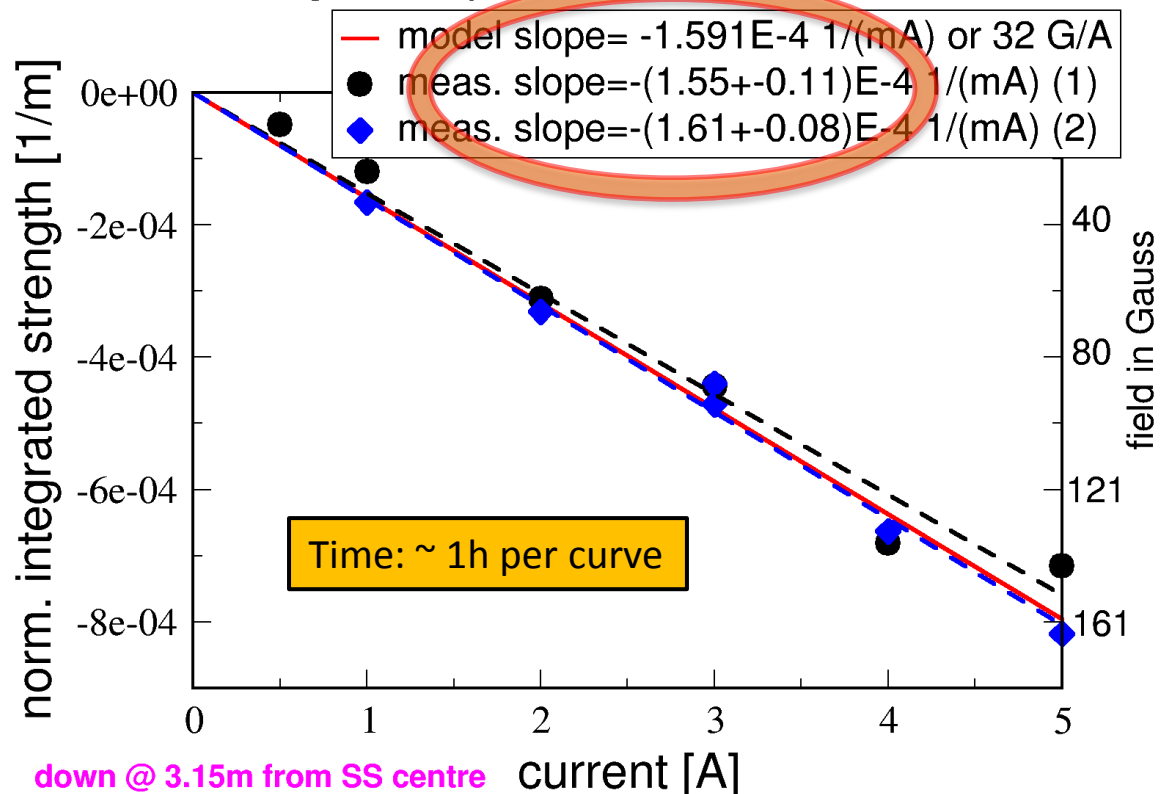


$$\delta O_{xy} = M \delta J_I$$

$\delta J_I$  skew quad field



calibrating skew quad id13\_corr DOWN via AC ORM



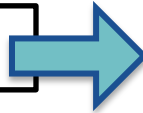
down @ 3.15m from SS centre



# ESRF: ID Compensation

Measure and fit AC ORM to infer ID gap-dependent skew quad field, calibrate corrector coils and check look-up tables

$$\delta O_{xy} = O_{xy}^{(ID \text{ gap closed})} - O_{xy}^{(ID \text{ gap open})}$$

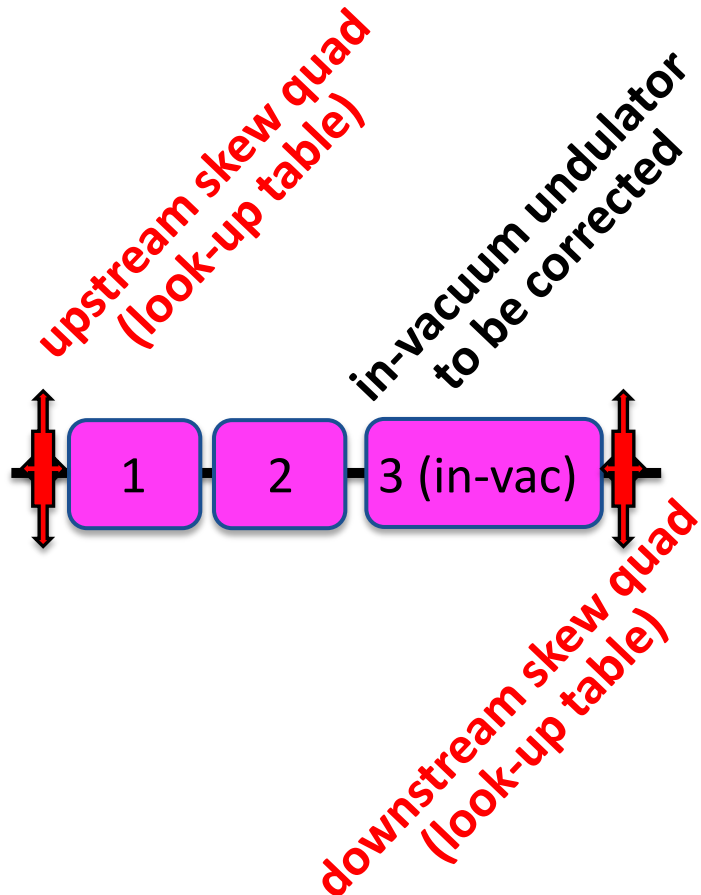


$$\delta O_{xy} = M \delta J_l$$

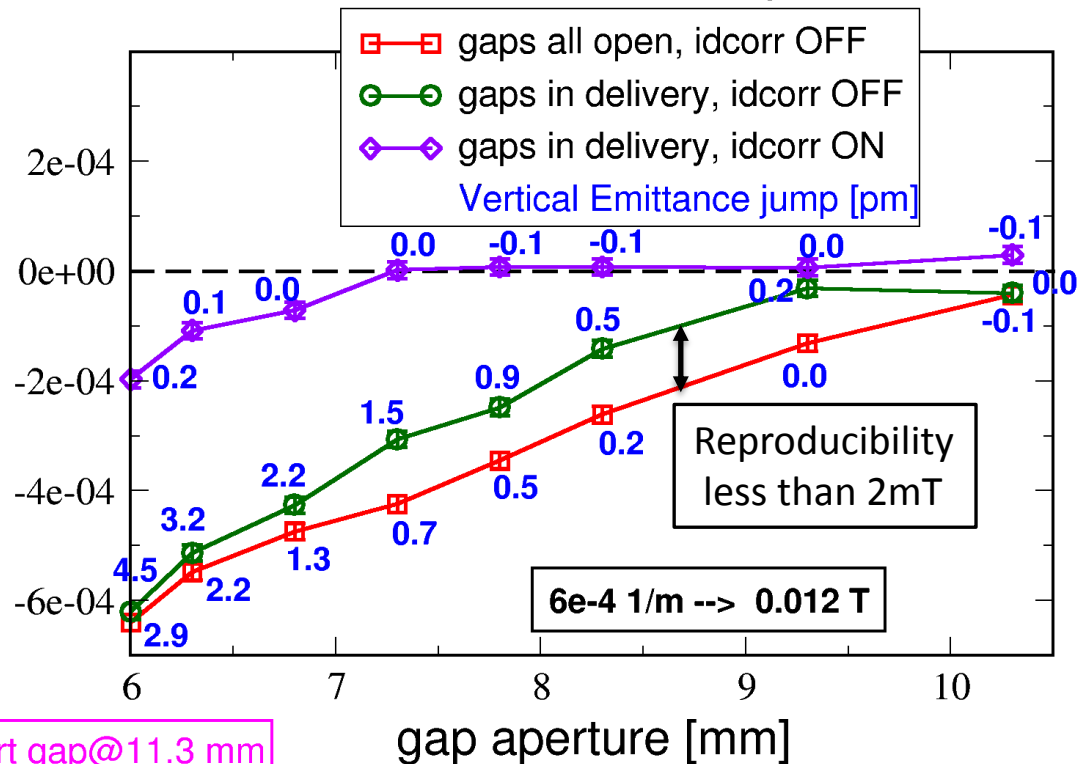
$\delta J_l$  skew quad field

Time: ~ 1h per curve

skew quad Vs ID13 IVU18C gap via AC ORM



norm. integrated strength [1/m]



# Summary

Fast ORM measurements have many benefits:

- substantial reduction in measurement / correction times
- minimises impact from machine drift and / or hysteresis effects
- improved accuracy compared to DC LOCO or TBT-based algorithms
- enables frequent optics correction and use by non-experts
- enables new types of measurement (off-energy ORM, nonlinear ORM)

Potential future developments:

- Integrate measurement with fast orbit feedback
- Transfer existing Matlab code to Python to streamline data acquisition and processing
- Investigate small amplitude / long duration excitation to enable data acquisition during user time

# References

- [1] J. Safranek, NIMA, 388, 27-36, (1997)
- [2] I.P.S. Martin et al., in Proc. IPAC'14, Dresden, Germany, TUPRI083, (2014)
- [3] G. Rehm et al. in Proc BIW'10, Santa Fe, U.S.A, MOCNB01, (2010)
- [4] V. Smaluk et al., IPAC'16, THPMR008, (2016)
- [5] G. Wang et al., BNL Tech. Note 168, (2015)
- [6] X. Huang et al., PRSTAB 8, 064001 (2005)
- [7] J. Irwin et al., PRL 82, 1684 (1999)
- [8] Y. Hidaka et al., NA-PAC'16, TUPOB52, (2016)
- [9] X. Yang et al., PRAB 20, 054001 (2017)
- [10] Z. Martí et al., IPAC'17, Copenhagen, Denmark, MOPAB102, (2017)

