

# Test of tri-direct CP symmetry models by neutrino oscillations Jian Tang

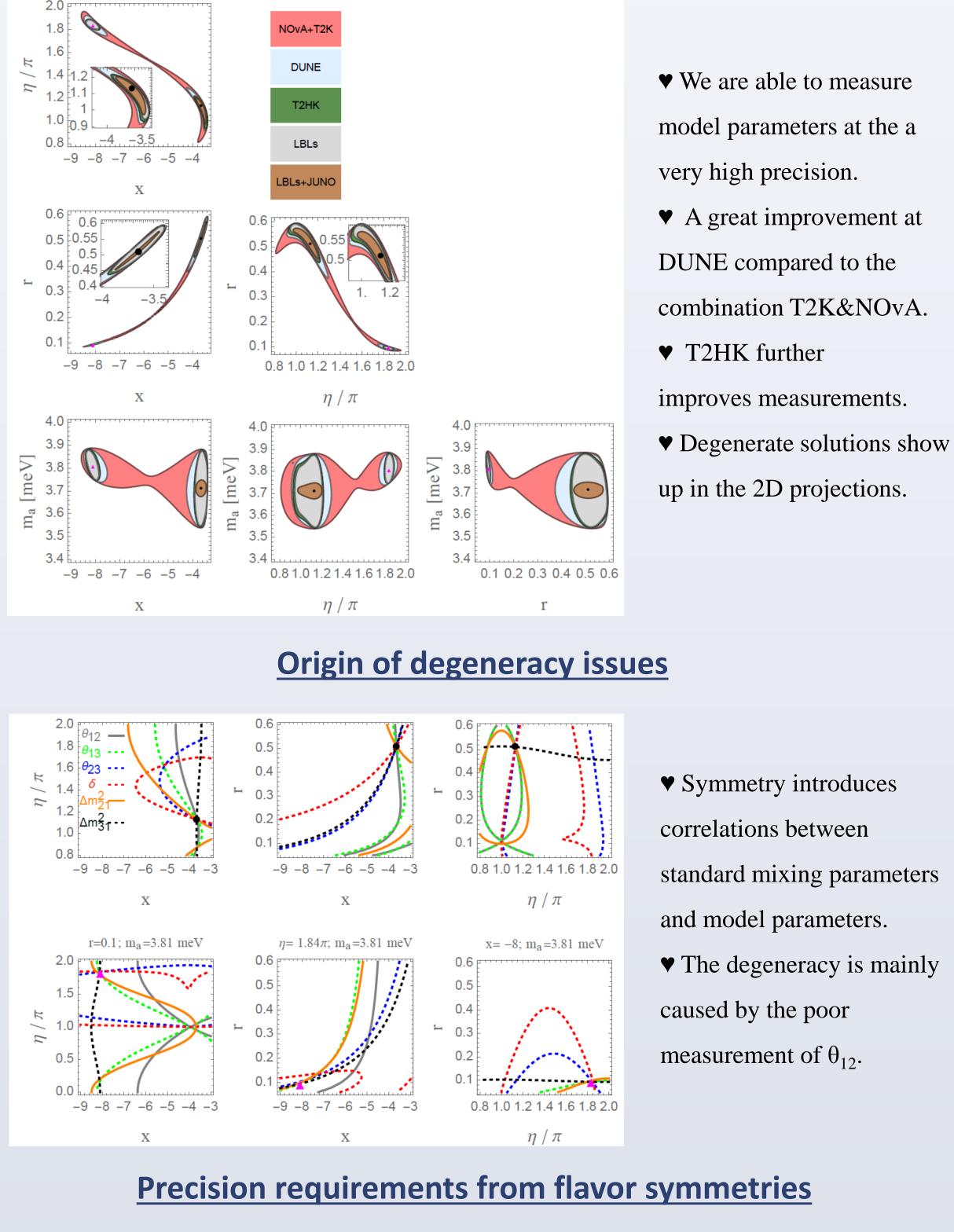


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#### **Motivations**

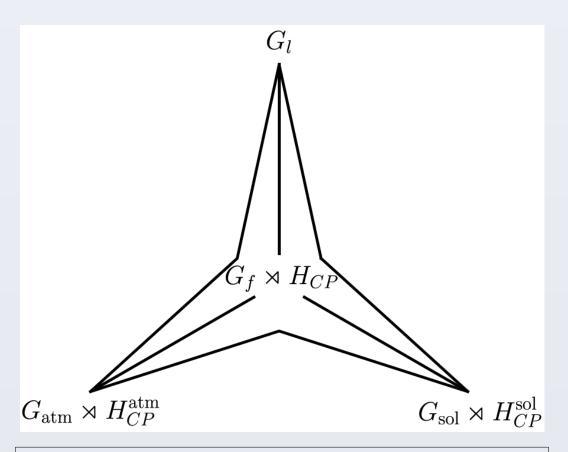
- 1) It is important to extend the Standard Model to naturally generate tiny neutrino masses. Flavor symmetry helps to reduce the degrees of freedom.
- 2) How powerful is it to reach precision measurement of standard neutrino mixing parameters to test flavor-symmetry models?
- How can we implement the over-constrained mixing parameters predicted by the flavor-3) symmetry models?
- 4) Are there any new features like degeneracy issues and how to break degeneracies?
- Is it possible to check the sum rules predicted by models? 5)
- At which level are we able to exclude a class of flavor-symmetry models? 6)





#### **Constrain model parameters**

#### **Review of tri-direct CP symmetry models**



Sketch of the tri-direct CP approach for two right-handed neutrino models

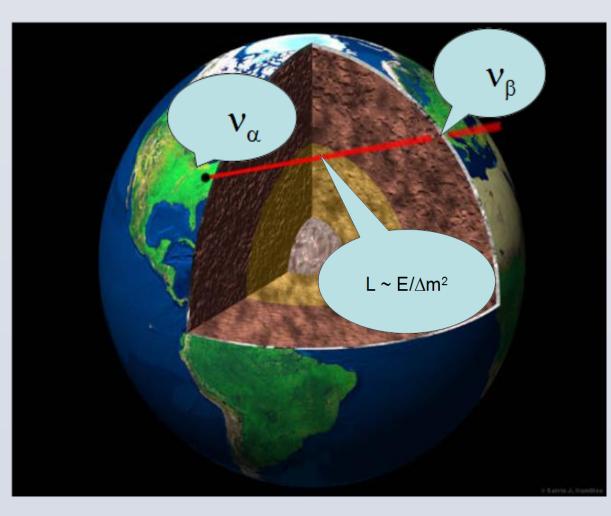
- The flavor group S4 and CP is broken to the subgroups. The residual symmetries are associated with the atmospheric and solar flavor sectors.
- The charged lepton mass matrix is diagonal.
- Structure of the neutrino and charged lepton mass matrices arise from the vacuum alignment of flavon fields which are fixed by the residual symmetry.
- Only four parameters  $m_a$ ,  $m_s$ ,  $\eta$  and x are involved to describe both neutrino masses and lepton mixing parameters.

| $m_D =$ | $\begin{pmatrix} y_a \ \omega y_a \ \omega^2 y_a \end{pmatrix}$ | $\begin{pmatrix} y_s \\ xy_s \\ xy \end{pmatrix}$ | $m_{\nu} = m_a$ | $\begin{pmatrix} 1\\ \omega\\ \omega^2 \end{pmatrix}$ | $\omega \ \omega^2 \ 1$ | $\begin{pmatrix} \omega^2 \\ 1 \\ \omega \end{pmatrix}$ | $+ e^{i\eta}m_s$ | $\begin{pmatrix} 1 \\ x \\ x \end{pmatrix}$ | $\begin{array}{c} x \\ x^2 \\ x^2 \end{array}$ | $\begin{pmatrix} x \\ x^2 \\ x^2 \end{pmatrix}$ |
|---------|---|---|-----------------|---|-------------------------|---|------------------|---|--|---|
|         | $\sqrt{\omega^{-}y_{a}}$  | $xy_s$  |                 | $\langle \omega^2 \rangle$                            | T                       | $\omega$ )  |                  | $\backslash x$                              | <i>x</i> -                                     | $x^2$   |

Best-fit values based on the latest global fit NuFit4.0 with priors taken into account.

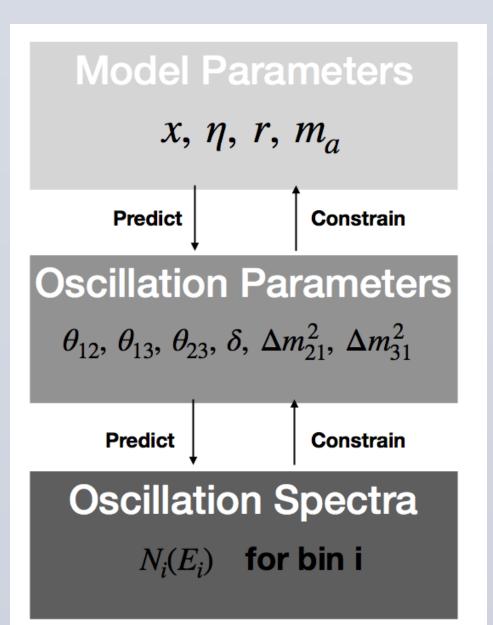
| $\Delta \chi^2$ | x = x | $\eta/\pi$ | r     | $m_a/~{ m meV}$ | $\theta_{12}/^{\circ}$ | $	heta_{13}/^{\circ}$ | $	heta_{23}/^{\circ}$ | $\delta/^{\circ}$ | $\Delta m^2_{21}/10^{-5} {\rm eV^2}$ | $\Delta m^2_{31}/10^{-3} {\rm eV}^2$ |
|-----------------|-------|------------|-------|-----------------|------------------------|-----------------------|-----------------------|-------------------|--------------------------------------|--------------------------------------|
| 4.98            | -3.65 | 1.13       | 0.511 | 3.71            | 35.25                  | 8.63                  | 46.98                 | 278.96            | 7.39                                 | 2.525                                |

## **Working principle of neutrino oscillations**



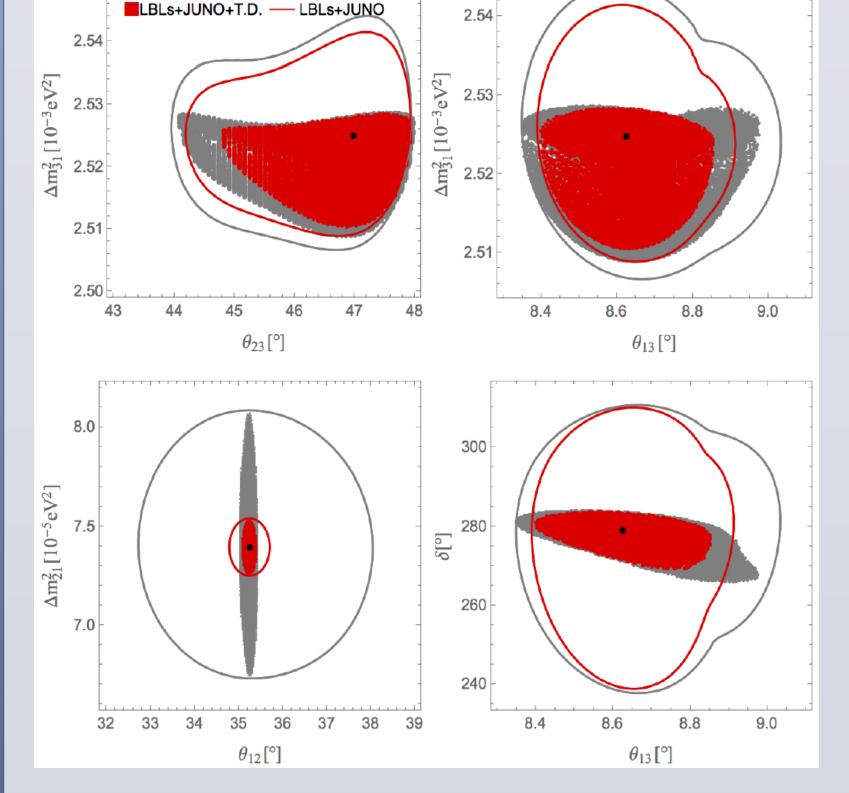
- high-energy protons hit the target station to produce charged mesons which can decay to generate neutrinos.
- 2) Near detector: flux measurements, cancellation of the systematic uncertainties...
- 3) Far detector: detection of oscillated neutrinos, reconstruction of oscillation probabilities, conduct physics analysis.
- 4) Running experiments: T2K, NOvA
- 5) Next generations: T2HK, DUNE, JUNO

### **Strategies to test the flavor-symmetry models**



$$\overrightarrow{\mathcal{M}} = \{x, \eta, m_a, r\}$$
$$\chi^2(\overrightarrow{\mathcal{M}}) = \sum_{i=1}^N \frac{\left[\mu_i(\overrightarrow{\mathcal{O}}(\overrightarrow{\mathcal{M}})) - n_i\right]^2}{\sigma_i^2}$$

$$\overrightarrow{\mathcal{O}} = \{\theta_{12}, \theta_{13}, \theta_{23}, \delta_{\mathrm{CP}}, \Delta m_{21}^2, \Delta m_{31}^2\}$$
$$\chi^2(\overrightarrow{\mathcal{O}}) = \sum_{i=1}^N \frac{\left[\mu_i(\overrightarrow{\mathcal{O}}) - n_i\right]^2}{\sigma_i^2}$$



LBLs+T.D.

- LBLs

incorporate the model symmetry.

♥ Shaded regions

♥ JUNO is good at precision measurement of  $\theta_{12}$  and might help to break degeneracies.

♥ Gray and red regions highlights the contribution by JUNO.

♥ Shape of contours in the projected parameter space can give us hints of the underlying theory.

## References

- 1) I. Esteban, M. C. Gonzalez-Garcia, A. Hernandez-Cabezudo, M. Maltoni, and T. Schwetz, JHEP 01, 106 (2019), 1811.05487.
- 2) G.-J. Ding, S. F. King, and C.-C. Li, JHEP 12, 003 (2018), 1807.07538.
- 3) G.-J. Ding, Y.-F. Li, J. Tang, and T.-C. Wang (2019), 1905.12939, to appear in PRD.
- 4) J. Tang and T.-C. Wang (2019), arXiv: 1907.01371,

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