



Status of the detector design studies for ESSvSB

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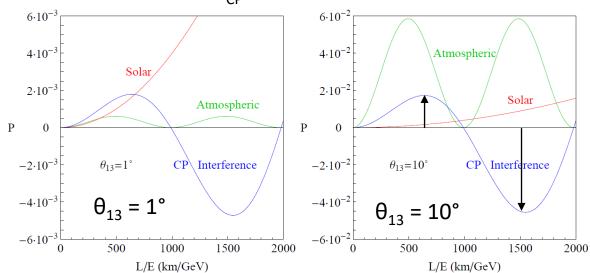
Introduction



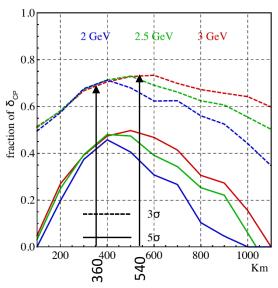
"Discovery and measurement of leptonic CP violation using an intensive neutrino Super Beam generated with the exceptionally powerful ESS linear accelerator"

15 institutes from 11 European countries, including CERN and ESS

Measure δ_{CP} at 2nd oscillation maximum



Prospective mine locations



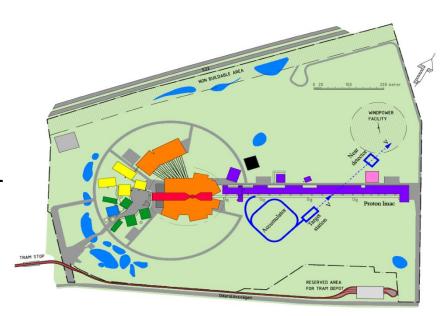
E. Baussan et al., Nucl. Phys. B 885, 127 (2014)

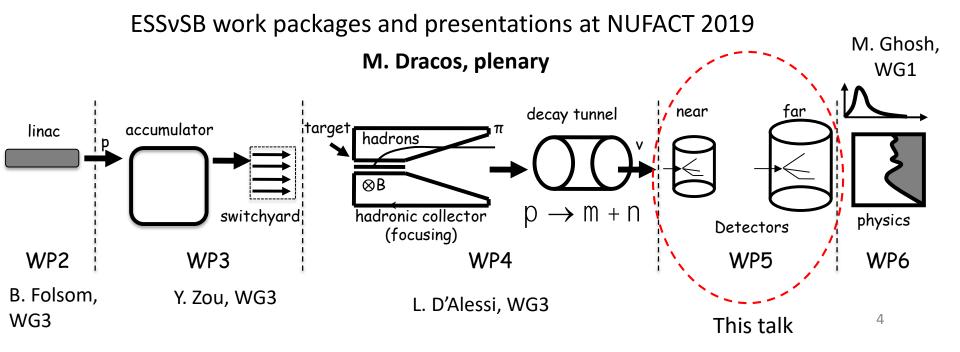
P. Coloma and E. Fernandez-Martinez, arXiv:1110.4583v2

Introduction

- 2.0-GeV protons
- 5-MW beam power
- 14-Hz repetition rate, 2.86-ms pulses
- 10¹⁵ protons per pulse, > 10²³ per year

Upgrades desired for ESSvSB and other experiments





The near detector (ND)

Characterization of the neutrino beam near the production point

- Energy and flavor measurement neutrino flux
- Neutrino interaction cross section measurement

Aim: < 5% systematic uncertainties in signal/background normalization
5% signal, 10% background uncertainties assumed in white paper
[E. Baussan et al., Nucl. Phys. B 885, 127 (2014)]

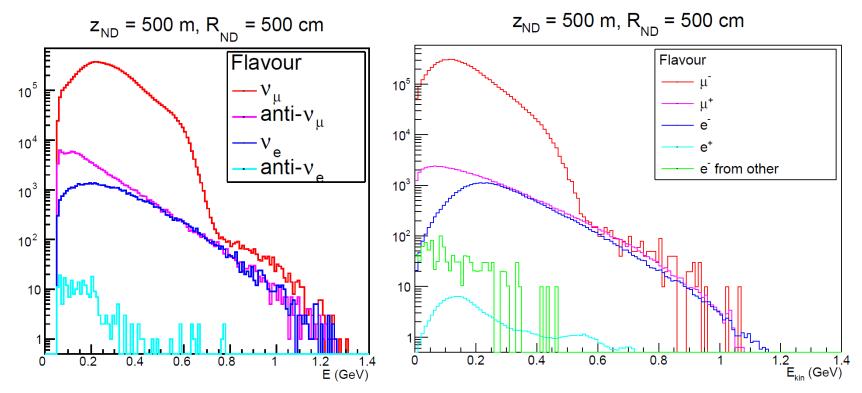
ND design criteria/requirements:

- Large mass to provide a sufficient number of neutrino interactions
- Acceptance for charged leptons (muons and electrons) in large scattering angle
- Capability to reconstruct and identify short tracks of low-energy hadrons around the interaction vertex

ND – background

Characterization of the neutrino beam near the production point

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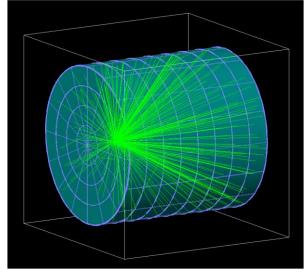


Neutrinos/antineutrinos

Charged leptons from neutrino interactions

ND designs

Water Cherenkov detector



A. Burgman, MSc thesis presentation

PMT or LAPPD readout

Fiducial volume: 100-1000 tonnes

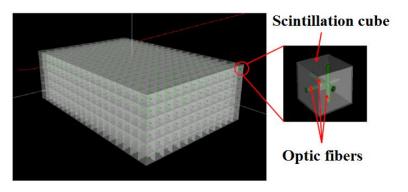
Simulations implemented in:

WCSim – Geant4 and ROOT-based water

Cherenkov detector simulation, open source

EsbRoot – based on FairRoot framework

Super Fine-grained detector (SuperFGD)



A. Blondel et al., Jour. Instrum. 13, P02006 (2018)

 $1 \times 1 \times 1$ cm³ plastic scintillator cubes with wavelength-shifting fibers inside magnetic field

Mass: 1-10 tonnes

Tracking and momentum measurement

Simulated in EsbRoot

WC-ND development

Motivation: same neutrino interaction material and detection mechanism as for the water Cherenkov far detector (FD)

→ Possible reduction or cancellation in systematic uncertainties in energy reconstruction and flavor identification between ND and FD

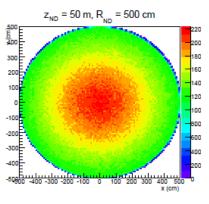
Questions to be addressed as a function of ND geometry:

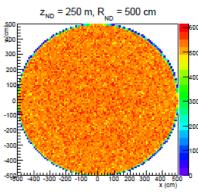
- Which reconstruction algorithm, especially at < 200 MeV energies?
- Keeping neutrino flavor misidentification rates to << 1%
- Energy/momentum/vertex reconstruction performance

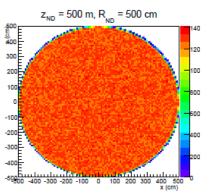
Several conclusions drawn from beam profile and charged lepton propagation in water (next slide)

Detector performance and geometry under investigation

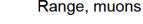
WC-ND physical requirements

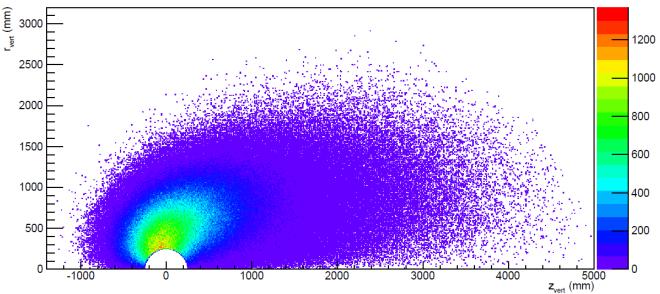






Neutrino spatial profile





- The length of the near detector: $L_{ND} \gg 3$ m.
- The radius of the near detector: $R_{ND} \gg 2$ m, $R_{ND} \le \frac{z_{ND}}{50}$.
- The distance between the neutrino beam production and the near detector: $z_{ND} \ge 200$ m.
- ► The detector must have a space resolution smaller than 10 cm.
- ► The detector must have a time resolution shorter than 100 ps.

Conclusions from MSc thesis by A. Burgman, Lund University (2015) $Z_{ND} = 500 \text{ m}$ proposed for more shielding from cosmic background

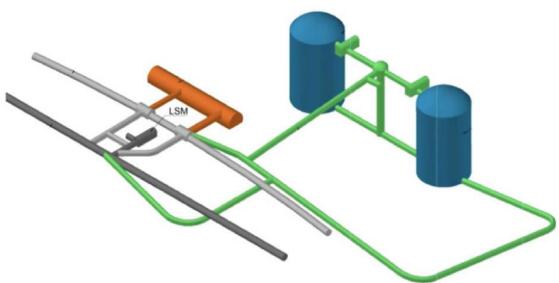
The far detector (FD)

MEMPHYS-like Water Cherenkov detector

(MEgaton Mass PHYSics)

L. Agostino et al., Phys. Rev. ST Accel. Beams 16, 061001 (2013)



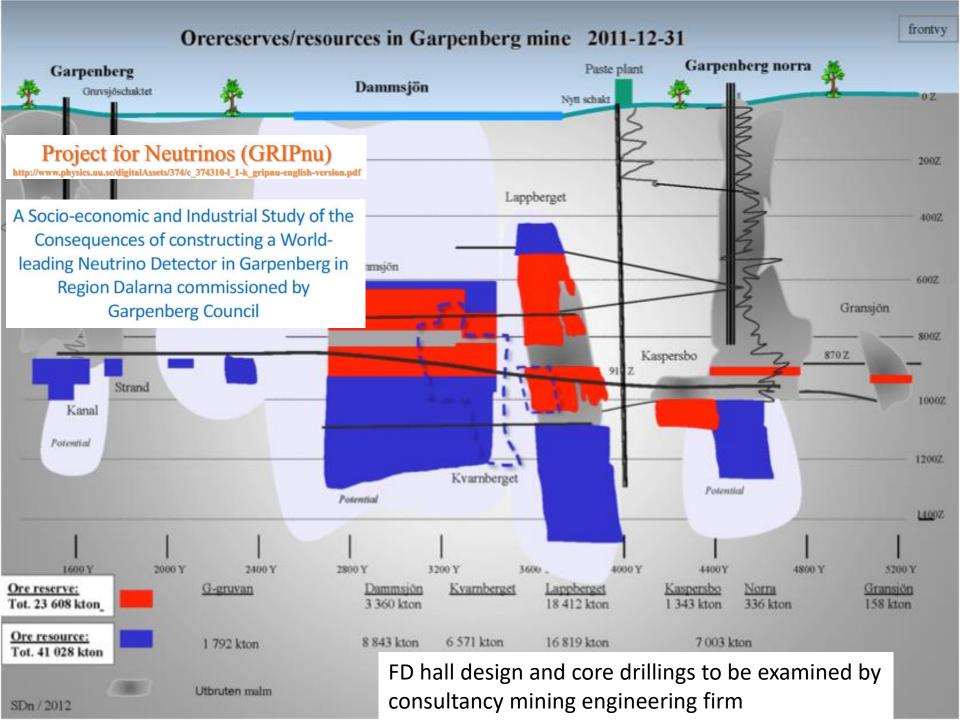


- Neutrino oscillations
- Proton decay
- Astroparticles
- Neutrinos from supernovae
- Solar/atmospheric neutrinos

500-kt fiducial volume (~20 x SuperK)

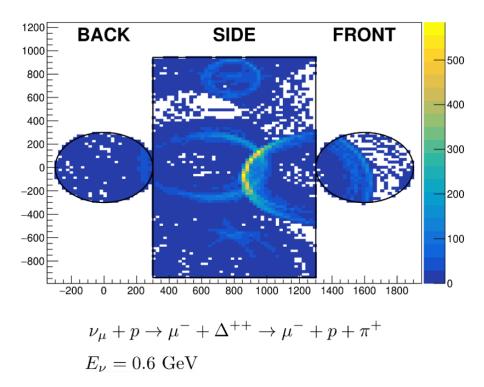
~240k 8-inch PMTs, 30% optical coverage

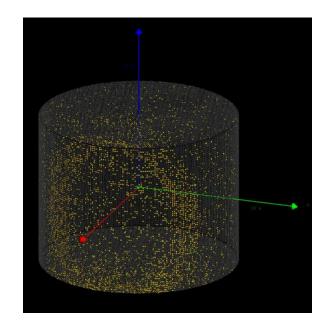
Cost: ~€700M



FD simulations

Simulated in EsbRoot and WCSim event displays

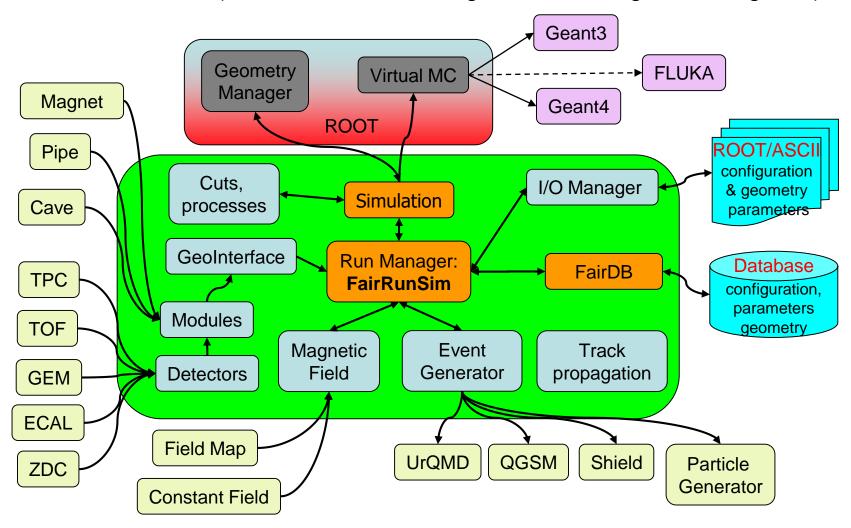




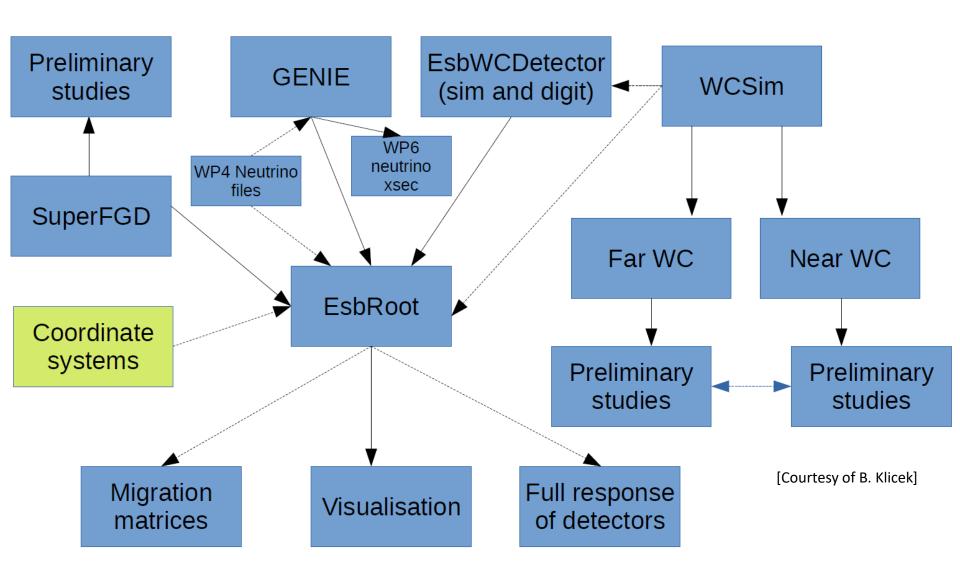
- Charged lepton simulation and hit digitization complete
- Working on consistency in physics parameters and simulation output between EsbRoot and WCSim
- Assessing different reconstruction strategies and design dimensions

Software development - simulations

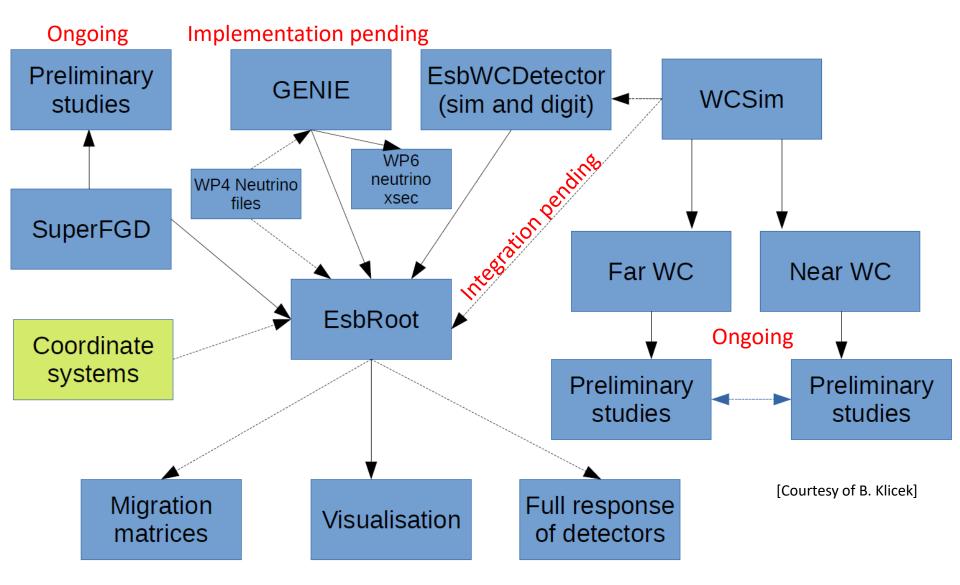
FairRoot framework (material from K. Gertsenberger, ESSvSB meeting in Strasbourg, 2018)



Detector software block diagram



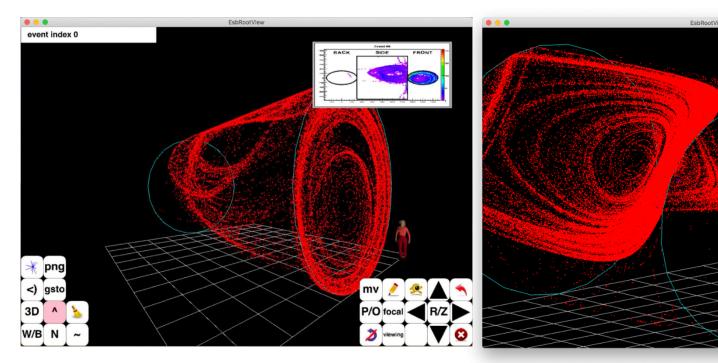
Detector software block diagram



3rd software workshop held in July 2019

Event display tool - EsbRootView

Courtesy of G. Barrand, CNRS/IN2P3/LAL



- 2D/3D event topologies
- Rotatable view
- Web browser/app development
- Timing information?

CCQE interaction inside SuperFGD, Chernkov radiation in water detector

Important cross checking tool for event reconstruction and PID

Summary and outlook

Initial designs and performance studies

- Water Cherenkov ND/FD
- SuperFGD tracker

Software frameworks

- EsbRoot, WCSim implemented
- Reconstruction algorithm under development

Tasks:

- Integrate GENIE Monte-Carlo neutrino interaction generator (arXiv:1510.05494) into EsbRoot
- Develop reconstruction algorithm for both ND and FD
- Finalize detector location, dimensions, components, and data analysis strategies



Design studies to be completed by the end of 2021 with a conceptual design report

Acknowledgements

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ESSvSB WP5 collaborators:

G. Barrand, E. Baussan, M. Bogomilov, J. Cederkäll, P. Christiansen, M. Dracos, T. Ekelöf, G. Fanourakis, G. Gokbulut, B. Klicek, M. Oglakci, J. Park, G. Petkov, G. Stavropoulos, T. Tolba, A. K. Topaksu, R. Tsenov,

and O. Zormpa

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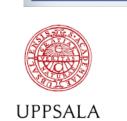












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