## Future Colliders and the Cosmic Frontier Part A: EWPT


http://www.physics.umass.edu/acfi/


My pronouns: he/him/his

http://tdli.sjtu.edu.cn/web/yjxy/5130001.htm

[^0]
## Feliz Fiesta del Orgullo



Cincuenta anos Stonewall


## Questions for Future Colliders

- What is the "value added"?
- What are the synergies/complementarities involving the pp, ee, and ep colliders ?
- Are there well-defined targets in mass reach and precision that would definitively address key open questions?


## The Origin of Matter

Cosmic Energy Budget


What can the LHC \& future colliders teach us about open questions in cosmology?

## Themes for This Talk

- The future collider program provides an opportunity to perform a comprehensive probe of the thermal history of EW symmetry breaking in BSM scenarios
- Many interesting aspects of dark matter/dark sector physics can be studied with future colliders $\rightarrow$ a comprehensive picture remains to be developed


## Disclaimer \& Credits

- Disclaimer: I am attempting to combine two topics into one longer talk $\rightarrow$ my apologies to anyone for omission of work that should be mentioned in a full one hour talk on either topic
- Credits: Many thanks to input l've received from week 1 speakers, Tim Tait who unfortunately had to cancel his visit, and my many collaborators


## Outline - Part A: EW Phase Transition

I. Context \& Questions
II. Models \& Phenomenology

- MSSM
- Simplified Higgs Portal
III. Theoretical Robustness
IV. Outlook


## Outline - Part B: DM

I. Context<br>II. MSSM<br>III. Simplified Models<br>IV. EW Multiplets<br>V. QCD-Like DM<br>VI. Mediators<br>VII. Outlook

## A-I. Context \& Questions

## Electroweak Phase Transition

- Higgs discovery $\rightarrow$ What was the thermal history of EWSB ?


## Thermal History of Symmetry Breaking



QCD Phase Diagram $\rightarrow$ EW Theory Analog?

## EWSB: The Scalar Potential



What was the thermal history of EWSB?

## EWSB: The Scalar Potential



What was the thermal history of EWSB?

## EW Phase Transition: St'd Model



## EW Phase Transition: St'd Model



Increasing $m_{h}$

| Lattice | Authors | $M_{\mathrm{h}}^{C}(\mathrm{GeV})$ |
| :--- | :---: | :---: |
| 4D Isotropic | [76] | $80 \pm 7$ |
| 4D Anisotropic | [74] | $72.4 \pm 1.7$ |
| 3D Isotropic | [72] | $72.3 \pm 0.7$ |
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SM EW: Cross over transition

## EW Phase Transition: St'd Model



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SM EW: Cross over transition

## Patterns of Symmetry Breaking


S. Weinberg, PRD 9 (1974) 3357

## Patterns of Symmetry Breaking



Extrema can evolve differently as $T$ evolves $\rightarrow$ rich possibilities for symmetry breaking

## Patterns of Symmetry Breaking



Extrema can evolve differently as $T$ evolves $\rightarrow$ rich possibilities for symmetry breaking

## Electroweak Phase Transition

- Higgs discovery $\rightarrow$ What was the thermal history of EWSB ?
- Baryogenesis $\rightarrow$ Was the matter-antimatter asymmetry generated in conjunction with EWSB (EW baryogenesis) ?


## Baryogenesis Scenarios



## Baryogenesis Scenarios



Era of EWSB: $t_{\text {univ }} \sim 10 \mathrm{ps}$

## EW Phase Transition: Baryogenesis



Increasing $m_{h}$


EWSB

Baryogenesis Gravity Waves Scalar DM LHC Searches

## EW Phase Transition: Baryogenesis



Increasing $m_{h}$


## EW Phase Transition: Baryogenesis



Increasing $m_{h}$


Baryogenesis Gravity Waves Scalar DM LHC Searches


## EW Phase Transition: Baryogenesis



Increasing $m_{h}$


Bubble nucleation


## EW Phase Transition: Baryogenesis



Increasing $m_{h}$


Quench

$Y_{B}$ : diffuses EWSB into interiors


## Electroweak Phase Transition

- Higgs discovery $\rightarrow$ What was the thermal history of EWSB ?
- Baryogenesis $\rightarrow$ Was the matter-antimatter asymmetry generated in conjunction with EWSB (EW baryogenesis) ?
- Gravitational waves $\rightarrow$ If a signal observed in LISA, could a cosmological phase transition be responsible ?


## Gravitational Radiation



1. Bubbles nucleate and grow
2. Expand in a plasma - create reaction fronts
3. Bubbles + fronts collide - violent process
4. Sound waves left behind in plasma
5. Turbulence; damping
[^1]
## Gravitational Radiation



Thanks: D. Weir

## A-II. Models \& Phenomenology

## Models \& Phenomenology

## What BSM Scenarios?

$S M+$ scalar singlet
$S M+$ Scalar Doublet
$(2 H D M)$

SM + Scalar Triplet

MSSM

NMSSM..

Espinosa, Quiros 93, Benson 93, Choi, Volkas 93, Vergara 96, Branco, Delepine, Emmanuel Costa, Gonzalez 98, Ham, Jeong, Oh 04, Ahriche 07, Espinosa, Quiros 07, Profumo, Ramsey-Musolf, Shaughnessy 07, Noble, Perelstein 07, Espinosa, Konstandin, No, Quiros 08, Barger, Langacker, McCaskey, Ramsey-Musolf, Shaughnessy 09, Ashoorioon, Konstandin 09, Das, Fox, Kumar, Weiner 09, Espinosa, Konstandin, Riva 11, Chung, Long 11, Barger, Chung, Long, Wang 12, Huang, Shu, Zhang 12, Fairbairn, Hogan 13, Katz, Perelstein 14, Profumo, Ramsey-Musolf, Wainwright, Winslow 14, Jiang, Bian, Huang, Shu 15, Kozaczuk 15 Cline, Kainulainen, Tucker-Smith 17, Kurup, Perelstein 17, Chen, Kozaczuk, Lewis 17, Gould, Kozaczuk, Niemi, Ramsey-Musolf, Tenkanen, Weir 19...

Turok, Zadrozny 92, Davies, Froggatt, Jenkins, Moorhouse 94, Cline, Lemieux 97, Huber 06 Froome, Huber, Seniuch 06, Cline, Kainulainen, Trott 11, Dorsch, Huber, No 13, Dorsch, Huber, Mimasu, No 14, Basler, Krause, Muhlleitner, Wittbrodt, Wlotzka 16, Dorsch, Huber, Mimasu, No 17, Bernon, Bian, Jiang 17, Andersen, Gorda, Helset, Niemi, Tenkanen, Tranberg, Vuorinen, Weir 18...

Patel, Ramsey-Musolf 12, Niemi, Patel, Ramsey-Musolf, Tenkanen, Weir 18 ...
Carena, Quiros, Wagner 96, Delepine, Gerard, Gonzalez Felipe, Weyers 96, Cline, Kainulainen 96, Laine, Rummukainen 98, Carena, Nardini, Quiros, Wagner 09, Cohen, Morrissey, Pierce 12, Curtin, Jaiswal, Meade 12, Carena, Nardini, Quiros, Wagner 13, Katz, Perelstein, Ramsey-Musolf, Winslow 14...

Pietroni 93, Davies, Froggatt, Moorhouse 95, Huber, Schmidt 01, Ham, Oh, Kim, Yoo, Son 04 Menon, Morrissey, Wagner 04, Funakubo, Tao, Yokoda 05, Huber, Konstandin, Prokopec, Schmidt 07, Chung, Long 10, Kozaczuk, Profumo, Stephenson Haskins, Wainwright 15...

## EWPT: Theory \& Phenomenology

- What models can lead to a (strong) first order electroweak phase transition (EW baryogenesis \& gravitational waves) ?
- Can they also yield contributions to $\Omega_{D M}$ ?
- How can they be tested experimentally?
- How reliably can we compute phase transition properties \& make the connection with phenomenology?


## First Order EWPT from BSM Physics

- Thermal loops involving new bosons
- T=0 loops (CW Potential)
- Change tree-level vacuum structure


## Why $T_{E W}$ Sets a Scale for Colliders

- Thermal loops involving new bosons
- T=0 loops (CW Potential)
- Change tree-level vacuum structure


## EWPT "Poster Child": MSSM Light Stop Scenario



Thermal loops

## EW Phase Transition: SUSY



Increasing $m_{h}$
$\longleftarrow$ New scalars

Light RH stops also affect Higgs properties

Curtin, Jaiswal, Meade 1203.2932
$M S S M+\delta \lambda_{4}\left(H_{u}^{\dagger} H_{u}\right)^{2}$


Katz, Perelstein, R-M,
Winslow 1509.02934

## Strong $1^{\text {st }}$ Order EWPT



Definitive probe of the possibilities $\rightarrow$ LHC + next generation colliders

## Higgs Portal: Simple Scalar Extensions

| Extension | DOF | EWPT | DM |
| :--- | :---: | :---: | :---: |
| Real singlet: $\mathrm{Z}_{2}$ | $\mathbf{1}$ | $\checkmark$ | $\checkmark$ |
| Real singlet: $\mathrm{Z}_{2}$ | $\mathbf{1}$ | $\nearrow$ | $\nearrow$ |
| Complex Singlet | $\mathbf{2}$ | $\nearrow$ | $\nearrow$ |
| EW Multiplets | $3+$ | $\nearrow$ | $\nearrow$ |

May be low-energy remnants of UV complete theory \& illustrative of generic features

## Higgs Portal: Simple Scalar Extensions



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## Higgs Portal: Simple Scalar Extensions

| This talk | Extension | DOF | EWPT | DM |
| :---: | :---: | :---: | :---: | :---: |
|  | Real singlet: $\chi_{8}$ | 1 | $\checkmark$ | * |
|  | Real singlet: $Z_{2}$ | 1 | $\checkmark$ | $\checkmark$ |
|  | Complex Singlet | 2 | $\checkmark$ | $\checkmark$ |
| This talk | EW Multiplets | 3+ | $\checkmark$ | $\checkmark$ |

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## Higgs Portal: Simple Scalar Extensions

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## Simplest Extension

## Standard Model + real singlet scalar

```
singlet EW Phase Transition
Driven
(lots of) Motivation
\(\Rightarrow\) Neutral Naturalness
\(\Rightarrow\) Higgs Portal (Dark Sectors)
\(\Rightarrow\) Non-minimal SUSY (e.g. NMSSM)
\(\Rightarrow\) Warped Extra Dim (dilaton...)
...
```


## Simplest Extension

Standard Model + real singlet scalar

$$
V_{\mathrm{HS}}=\frac{a_{1}}{2}\left(H^{\dagger} H\right) S+\frac{a_{2}}{2}\left(H^{\dagger} H\right) S^{2}
$$

- Strong first order EWPT
- Two mixed singlet-doublet states


## EW Phase Transition: New Scalars



Increasing $m_{h}$

Real Singlet: $\quad \phi \rightarrow S$
Simplest Extension:
two states $h_{1} \& h_{2}$


## EW Phase Transition: Singlet Scalars



Increasing $\boldsymbol{m}_{h}$

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SM EW: Cross over transition

## EW Phase Transition: Singlet Scalars



Increasing $m_{h}$
$\longleftarrow$ New scalars

Real Singlet: $\quad \phi \rightarrow S$
Simplest Extension: two states $h_{1} \& h_{2}-h, S$ mixtures


Profumo, MJRM, Shaugnessy ‘07

## EW Phase Transition: Singlet Scalars



Collider probes

- Resonant di-Higgs production
- Precision Higgs measurements
- Non-resonant di-Higgs \& exotic Higgs decays



## EW Phase Transition: New Scalars



Increasing $m_{h}$
Resonant di-Higgs production


No \& RM, arXiv:1310.6035 : LHC Discovery w/ $100 \mathrm{fb}^{-1}$


## EW Phase Transition: Singlet Scalars

SFOEWPT Benchmarks: Resonant di-Higgs \& precision Higgs studies


Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442;
Li et al, 1906.05289

## EW Phase Transition: New Scalars



Modified Higgs Self-Coupling



## EW Phase Transition: Singlet Scalars



Modified Higgs Self-Coupling


Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble \& Perelstein 0711.3018


## EW Phase Transition: Singlet Scalars



Modified Higgs Self-Coupling


Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble \& Perelstein 0711.3018


53
Thanks: M. Cepeda

## EW Phase Transition: Singlet Scalars



Modified Higgs Self-Coupling


Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble \& Perelstein 0711.3018



Thanks: M. Cepeda

## EW Phase Transition: New Scalars



Thanks: J. M. No, M. Cepeda

## EW Phase Transition: Singlet Scalars



Singlet-like pair production (off shell)


Chen, Kozaczuk, Lewis 2017


## Higgs Portal: Simple Scalar Extensions

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## The Simplest Extension

## DM Scenario



## EW Phase Transition: Two-Step



Profumo, R-M, Shaugnessy 2007
Epsinosa, Konstandin, Riva 2011
Curtain, Meade, Yu: arXiv: 1409.0005
Jiang, Bian, Huang, Shu 1502.07574

## EW Phase Transition: Singlet Scalars



Curtain, Meade, Yu: arXiv: 1409.0005
$Z_{2}$ symmetric real singlet extension

- Loop-induced 1-step transition
- 2-step transition for $\mu_{S}{ }^{2}<0$

* Singlet two step: see also Profumo, R-M, Shaugnessy 2007,


## EW Phase Transition: Singlet Scalars



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VBF @ 100 TeV pp:

$$
p p \rightarrow h j j, h \rightarrow \text { invis }
$$

* Singlet two step: see also Profumo, R-M, Shaugnessy 2007,
 Epsinosa, Konstandin, Riva 2011


## EW Phase Transition: DM Direct Detection



Curtain, Meade, Yu: arXiv: 1409.0005
$Z_{2}$ symmetric real singlet extension

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- 2-step transition for $\mu_{S}{ }^{2}<0$

Scalar singlet DM: direct detection


## Higgs Portal: Simple Scalar Extensions

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May be low-energy remnants of UV complete theory \& illustrative of generic features

## Real Triplet

$$
\Sigma^{0}, \Sigma^{+}, \Sigma^{-} \quad \sim(1,3,0)
$$

$$
V_{H \Sigma}=\frac{a_{1}}{2} H^{\dagger} \Sigma H+\frac{a_{2}}{2} H^{\dagger} H \operatorname{Tr} \Sigma^{2}
$$

EWPT: $a_{1,2} \neq 0 \quad \&<\Sigma^{0}>\neq 0$
DM \& EWPT: $\left.a_{1}=0 \quad \&<\Sigma^{0}\right\rangle=0$

## Real Triplet

$$
\sum^{0}, \Sigma^{+}, \Sigma^{-} \sim(1,3,0) \quad \begin{aligned}
& \text { Fileviez-Perez, Patel, Wang, R-M: PRD 79: } \\
& 055024 \text { (2009); 0811.3957 [hep-ph] }
\end{aligned}
$$

$$
V_{H \Sigma}=\quad+\frac{a_{2}}{2} H^{\dagger} H \operatorname{Tr} \Sigma^{2}
$$

$$
\begin{aligned}
& \text { EWPT: } a_{1,2} \neq 0 \quad \&<\Sigma^{0}>\neq 0 \\
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\end{aligned}
$$

DM Stability

## EW Multiplets: EWPT


$\longleftarrow$ New scalars

## EW Multiplets: EWPT


$\longleftarrow$ New scalars

- Thermal loops
- Tree-level barrier


## EW Multiplets: One-Step EWPT




Increasing $m_{h}$

$\longleftarrow$ New scalars

- One-step: Sym phase $\rightarrow$ Higgs phase


## EW Multiplets: Two-Step EWPT




Increasing $m_{h}$

$\longleftarrow$ New scalars

- One-step: Sym phase $\rightarrow$ Higgs phase
- Two-step: successive EW broken phases


## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$


Potential B
 phases

## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$
$\longleftarrow$ New scalars

- Step 1: thermal loops
- Step 2: tree-level barrier



## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$


## EW Multiplets: Two-Step EWPT




## EW Multiplets: 2HDM



Increasing $m_{h}$



Difference between Synmetric - Broken phase in CW piece guaranteed for large BSM mass splitting!

## A-III. Theoretical Robustness

## Theory Meets Phenomenology

A. Non-perturbative

- Most reliable determination of character of EWPT \& dependence on parameters
- Broad survey of scenarios \& parameter space not viable
B. Perturbative
- Most feasible approach to survey broad ranges of models, analyze parameter space, \& predict experimental signatures
- Quantitative reliability needs to be verified


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B. Perturbative mark pert
- NB: Recible approach to survey broad ranges of models, analyze parameter space, \& predict experimental signatures
- Quantitative reliability needs to be verified


## EWPT \& Perturbation Theory

## Expansion parameter



SM lattice studies: $g_{\text {eff }} \sim 0.8$ in vicinity of EWPT for $m_{H} \sim 70 \mathrm{GeV}$

## EW Multiplets: One-Step EWPT?



Increasing $m_{h}$
$\longleftarrow$ New scalars

- One-step: thermal loops



## Benchmarking PT: Recent Progress

Meeting ground: 3-D high-T effective theory


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Lattice simulations exist

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## Benchmarking PT: Recent Progress

## Meeting ground: 3-D high-T effective theory



- Assume BSM fields are "heavy" or "supeheavy": integrate out
- Effective "SM-like" theory parameters are functions of BSM parameters
- Use existing lattice computations for SM-like effective theory \& matching onto full theory to determine FOEWPT-viable parameter space regions


## Benchmarking PT: Recent Progress

## Meeting ground: 3-D high-T effective theory



Lattice simulations exist (e.g., Kajantie et al '95)

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## Benchmarking PT: Recent Progress

## Meeting ground: 3-D high-T effective theory



## Real Triplet \& EWPT



## Real Triplet \& EWPT



## Real Triplet \& EWPT



## Real Triplet \& EWPT



## Real Triplet Example: Lessons

- Initial non-perturbative studies using 3d EFT reveals regions of FOEWPT \& crossover transition not evident in PT
- Next generation circular e+e- and pp colliders likely necessary to access these region: a first order transition $\rightarrow$ Observable shift in $h \rightarrow \gamma \gamma$ rate
- Next generation colliders will have needed sensitivity


## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$


Patel, R-M: arXiv 1212.5652 ; Blinov et al: 1504.05195

## Real Triplet \& EWPT



## Scalar Singlets \& EWPT: Collider Reach

SFOEWPT Benchmarks: Resonant di-Higgs \& precision Higgs studies


Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442

## Real Singlet \& EWPT: Lattice "Repurpose"



## Heavy Real Singlet \& EWPT: Probes



## Heavy Real Singlet: EWPT \& GW

Non-dynamical heavy BSM scalars


- One-step
- Non-perturbative


## Heavy Real Singlet: EWPT \& GW



- One-step
- Non-perturbative


## A-IV. EWPT Outlook

## Questions for Future Colliders

- What is the "value added"?
- What are the synergies/complementarities involving the pp, ee, and ep colliders ?
- Are there well-defined targets in mass reach and precision that would definitively address key open questions?


## EWPT

- Value added

Extend reach significantly beyond HL-LHC

- Synergy/complementarity

Look for correspondence between new states (hh mode) and modified Higgs couplings (ee \& hh modes)

- Well-defined target in mass and/or precision


Singlets: $100 \mathrm{TeV}+30 a b^{-1}$
EW Multiplets: < 10\% on h $\gamma \gamma$


[^0]:    Future Colliders Workshop
    IFT Madrid, June-July 2019

[^1]:    Thanks: D. Weir

