

Study of small scale position-sensitive scintillator detector for gamma-ray spectroscopy

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11 July 2019

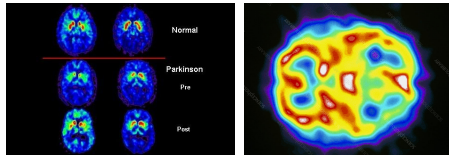
- 1 Introduction & motivation
- 2 Experimental setup
- 3 Results

1 Introduction & motivation

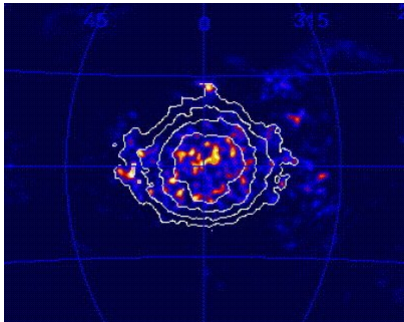
2 Experimental setup

3 Results

- Particle physics
- Nuclear physics
- **Medical physics**



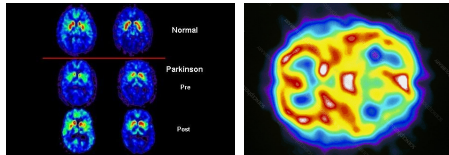
PET brain scan & SPECT brain during migraine
Credit: Dept. of Nuclear Medicine, Charing Cross Hospital / Science Photo Library



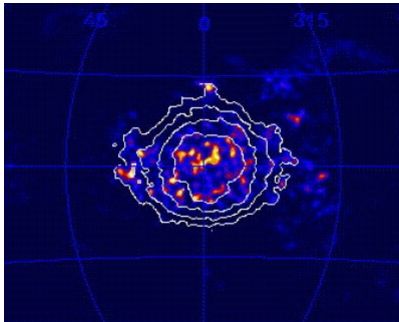
SPI image of 511 keV emission of centre of our Galaxy
Credit: J. Knödlseder (CESR) & SPI team

- **Astronomy**

- Particle physics
- Nuclear physics
- **Medical physics**



PET brain scan & SPECT brain during migraine
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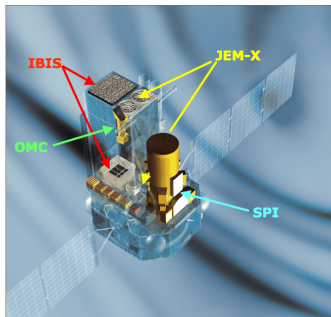
- **Astronomy**
- Project NDeGRA "Novel Detectors for Gamma-Ray Astronomy"

SPI Spectrometer on Integral

- Energy range: 8 keV - 8 MeV
- Energy resolution:
2.2 keV FWHM @ 1.33 MeV
- 19 Ge detectors
(Operating T of 85 K)

IBIS Imager on board the Integral Satellite

- Energy range: 15 keV - 10 MeV
- Energy resolution:
8% @ 100 keV & 10% @ 1 MeV



Integral
Credit: ESA

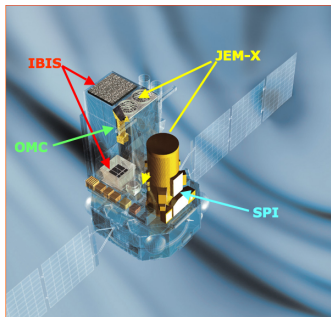
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IBIS Imager on board the Integral Satellite

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8% @ 100 keV & 10% @ 1 MeV

- Long development time
- High risk for mission execution
- Expensive



Integral
Credit: ESA

- SPI + IBIS weight:
over 2 tones
- Complex systems
- Difficult to improve

Different approach

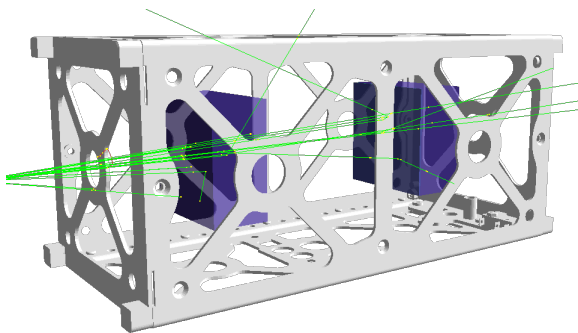
- Small and light
- Easy to build
- Easy to improve
- Cheap (to launch)
- Repeatedly reproduced

Different approach

- Small and light
 - Easy to build
 - Easy to improve
 - Cheap (to launch)
 - Repeatedly reproduced
- Imaging & spectroscopy
 - Energy range: 50 keV - 2 MeV

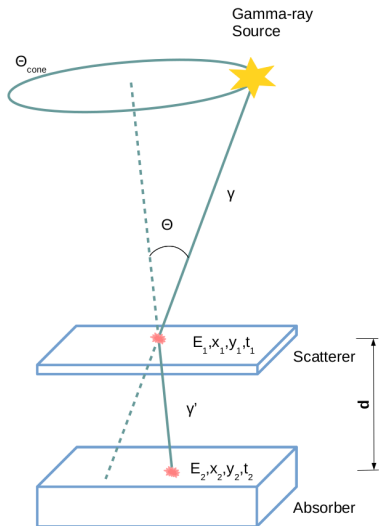
Different approach

- Small and light
- Easy to build
- Easy to improve
- Cheap (to launch)
- Repeatedly reproduced
- Imaging & spectroscopy
- Energy range: 50 keV - 2 MeV
- Nanosatellite:
up to 3 CubeSat units
(each $10 \times 10 \times 10 \text{ cm}^3$)



How we can do that?

Compton telescope



$$\cos\Theta = 1 - m_e c^2 \left(\frac{1}{E_2} - \frac{1}{E_1 + E_2} \right)$$

$$d = (x_1, y_1) - (x_2, y_2)$$

$$\text{Source}(x, y) = (d, \Theta)$$

- Fast γ - γ coincidences

1 Introduction & motivation

2 Experimental setup

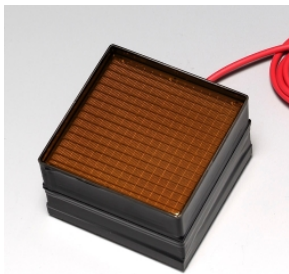
3 Results

- Scintillator: **CeBr₃**
 - Good Energy resolution
< 5% 662 keV
 - Fast scintillator
< 1 ns time resolution
(as LaBr₃Ce)
 - No internal radiation
 - High light output



Credit: SCIONIX

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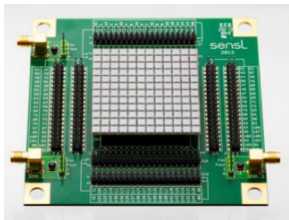
Credit: HAMAMATSU



Credit: SCIONIX

- Light sensor: **MAPMT**
 - High-voltage operation
 - High Gain apx. 10^6

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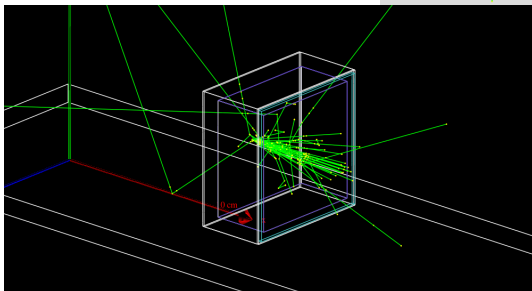
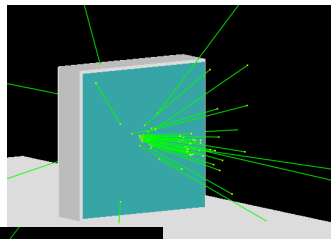
Credit: SENSL



Credit: SCIONIX

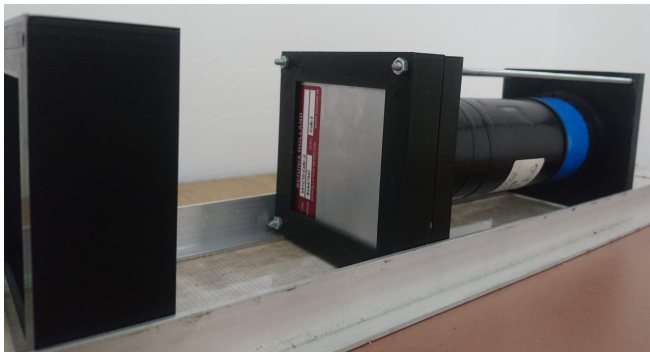
- Light sensor: **SiPM**
 - Low-voltage operation ✓
 - High Gain apx. 10^6

- Geometry - G4GDMLParser class
(generated GDML file directly from CAD file)
- Beam - G4GeneralParticleSource (GPS)
class (macro file)



Experimental setup for CeBr₃ energy resolution estimations

- CeBr₃ - thickness 25 & 10 mm
 - 51 x 51 mm²
- PMT - XP20D0B, Photonis
 - 51 mm, round tube
- HV - N1470, CAEN (operating HV -1000 V)
- ¹³⁷Cs, ²²Na & ²⁴¹Am

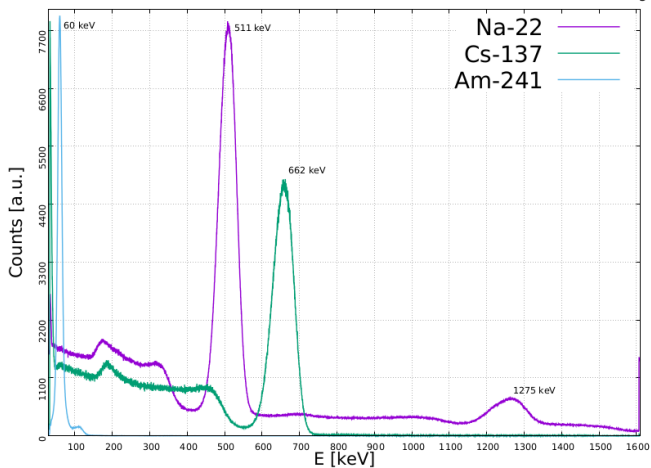


1 Introduction & motivation

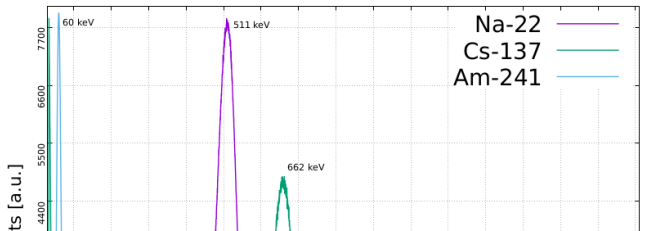
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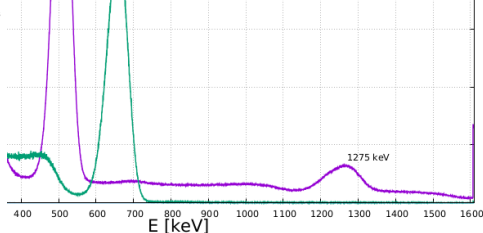
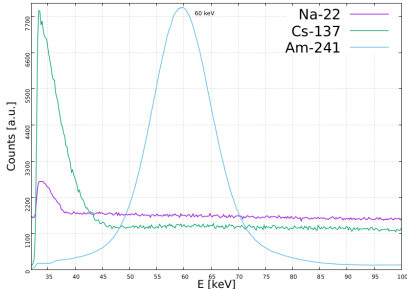
Energy spectrum of γ -rays from ²⁴¹Am, ¹³⁷Cs & ²²Na sources for 25mm CeBr₃



Energy spectrum of γ -rays from ^{241}Am , ^{137}Cs & ^{22}Na sources for 25mm CeBr₃

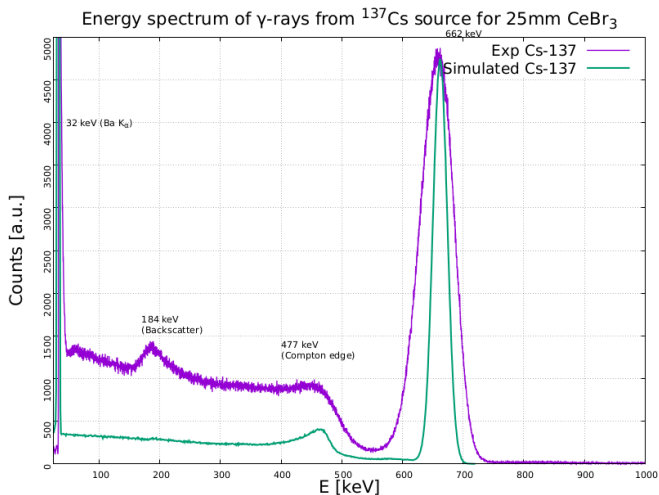


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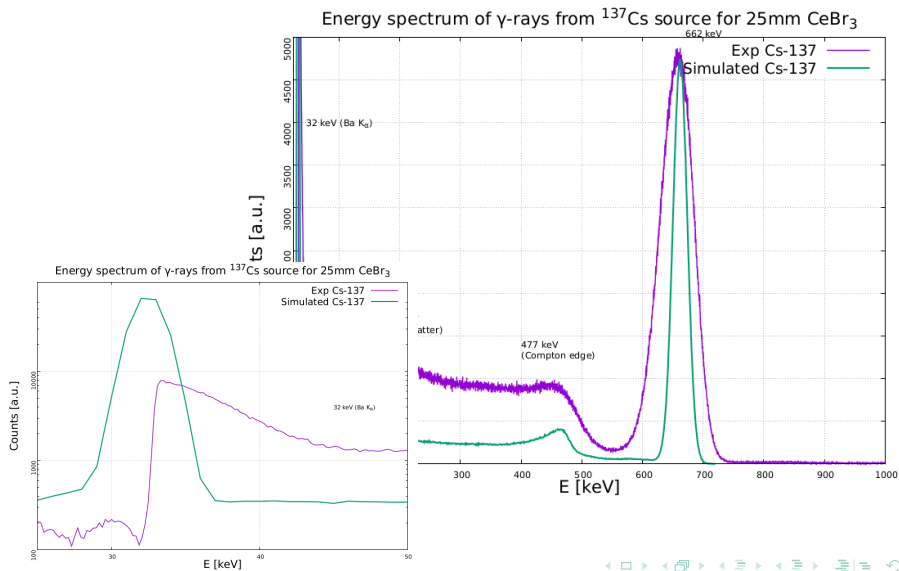


- CeBr₃ - thickness 25 mm
 - 18% @ 60 keV ²⁴¹Am
 - 8.7% @ 511 keV ²²Na
 - 9.0% @ 662 keV ¹³⁷Cs
(vs. 4%)
 - 8.1% @ 1274 keV ²²Na
 - CeBr₃ - thickness 10 mm
 - 30% @ 60 keV ²⁴¹Am
 - 9.8% @ 511 keV ²²Na
 - 9.4% @ 662 keV ¹³⁷Cs
(vs. 4%)
 - 7.2% @ 1274 keV ²²Na
-
- NOTE: PMT is smaller than the size of CeBr₃

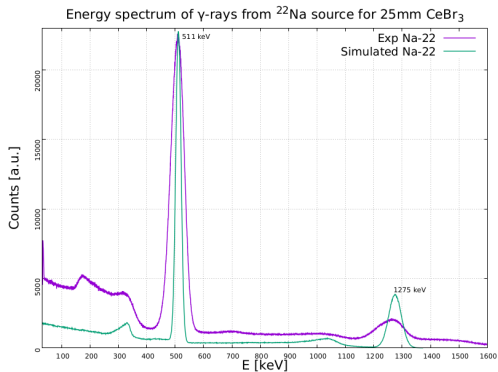
Comparison between the experimental and simulated spectrums for 25 mm thicker CeBr₃ with ¹³⁷Cs



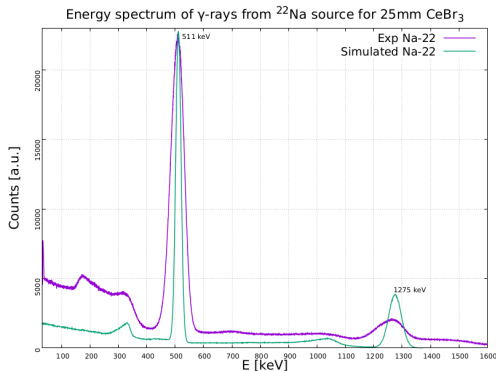
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Comparison between the experimental and simulated spectrums for 25 mm thicker CeBr₃ with ²²Na



Comparison between the experimental and simulated spectrums for 25 mm thicker CeBr₃ with ²²Na



● Experimental data

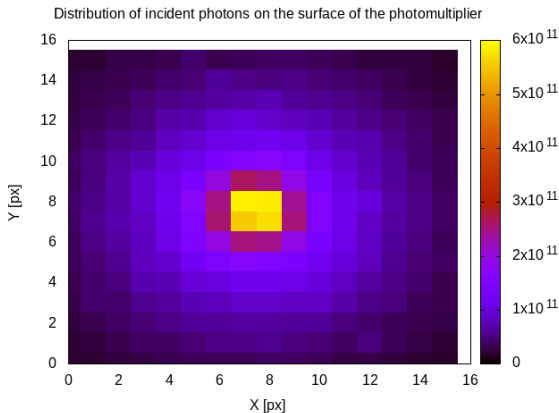
- 18% @ 60 keV ²⁴¹Am
- 8.7% @ 511 keV ²²Na
- 9.0% @ 662 keV ¹³⁷Cs
- 8.1% @ 1274 keV ²²Na

● Simulated data

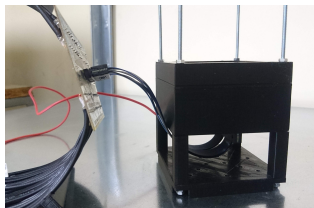
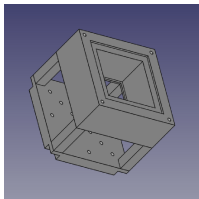
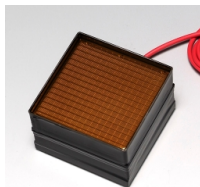
- 5.1% @ 60 keV ²⁴¹Am
- 4.3% @ 511 keV ²²Na
- 4.4% @ 662 keV ¹³⁷Cs
- 3.9% @ 1274 keV ²²Na

Preliminary results from simulations for the position resolution of 25 mm CeBr₃

- One pixel = anode size
 $2.8 \times 2.8 \text{ mm}^2$
- Position resolution:
 $< 5.3 \text{ mm } (2\sigma)$



- MAPMT - H9500, Hamamatsu
 - 16 x 16 anodes
 - anode size $2.8 \times 2.8 \text{ mm}^2$
- HV N1470, CAEN



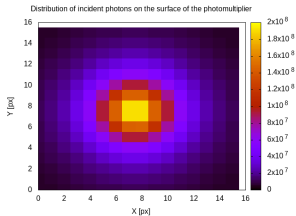
- Anode uniformity:
one photoelectron calibration

- **Thicker than 10 mm scintillator** for scatterer or even a different concept?
- **Promising** preliminary results for the position resolution.
- **Good** results from simulations - input information for the specification of the readout electronics and a better understanding of the experimental setup.
- **Outlook**
 - Investigating the potential issues CeBr_3 bad energy resolution.
 - Construct and build a new system for testing the individual channel response of the MAPMT.
 - Preparing the SiPM for operation and test.

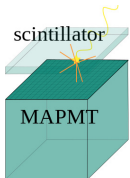


Thanks to my colleagues:
S. Ivanov, G. Georgiev, V. Kozhuharov and S. Lalkovski

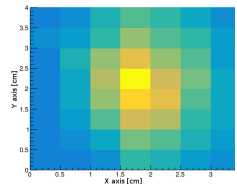
Time for questions?



- Geant4 simulations



- Monolithic plastic scintillator
- MAPMT
- $^{90}\text{Sr}/^{90}\text{Y}$
- Position resolution < 8 mm



Springer, in press, Chapter DOI: 10.1007/978-3-030-22204-8_40