



# Space charge effects on Landau damping from octupoles

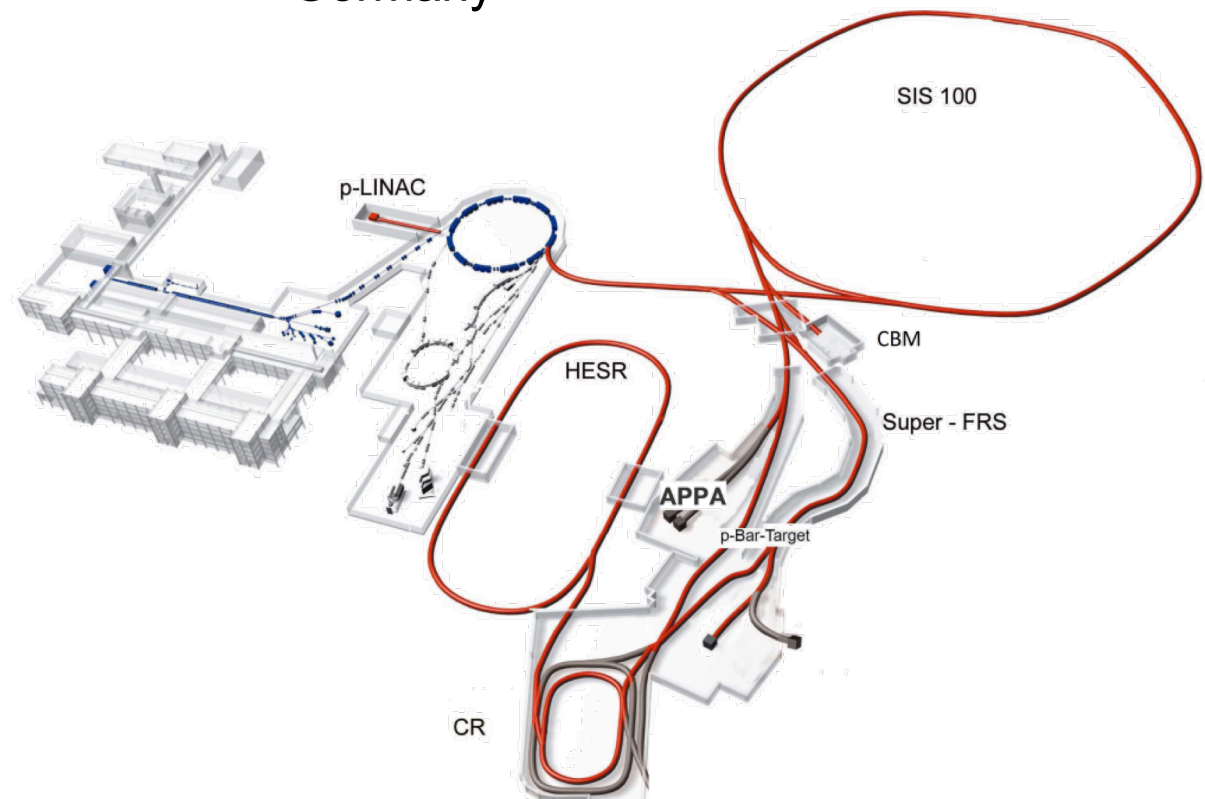
Vladimir Kornilov  
GSI Darmstadt, Germany

# FAIR project at GSI

## SIS100

- C:  $5 \times \text{SIS18} = 1\,083.6 \text{ m}$
- Ions H-U.
- $\text{U}^{28+}$   $0.2 \rightarrow 1.5 \text{ GeV/u}$
- $\text{p}^+$   $4 \rightarrow 29 \text{ GeV}$
- High-intensity, low-loss operation
- Coherent instabilities (head-tail, see the poster) is an important issue
- Space charge  $\Delta Q_{\text{SC}}$  up to 0.3
- Under construction, commissioning 2025.

The SIS100 synchrotron is the central accelerator of the FAIR Project at GSI Helmholtzzentrum in Darmstadt, Germany



## Octupole magnets

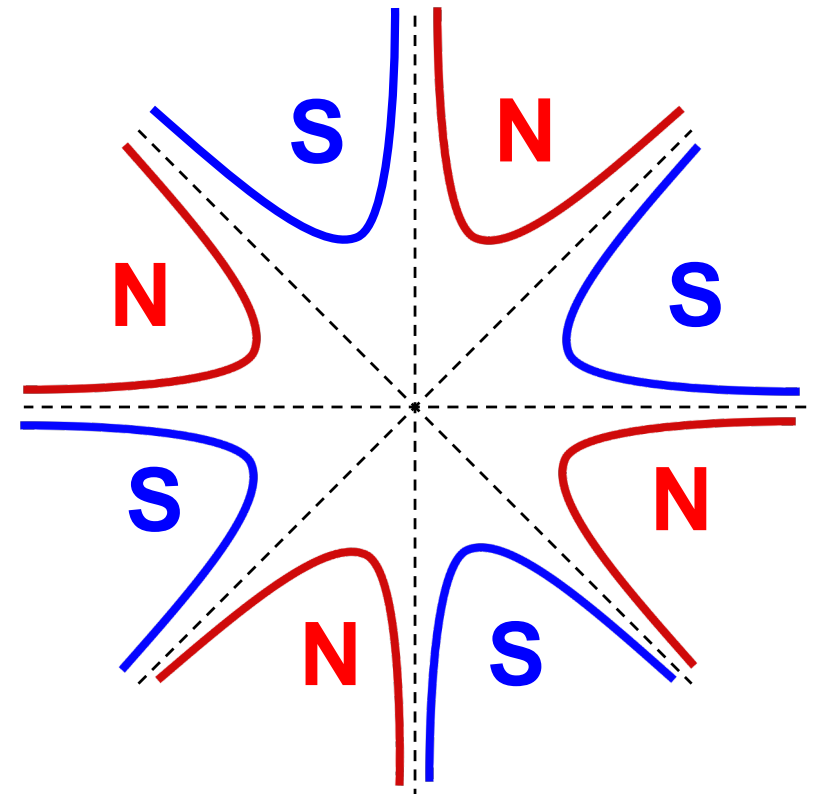
$$B_x = O_3(3x^2y - y^3)$$

$$B_y = O_3(x^3 - 3xy^2)$$

$$\Delta Q_x = \left( \int \frac{K_3 \beta_x^2}{16\pi} ds \right) J_x - \left( \int \frac{K_3 \beta_x \beta_y}{8\pi} ds \right) J_y$$

$$\Delta Q_y = \left( \int \frac{K_3 \beta_y^2}{16\pi} ds \right) J_y - \left( \int \frac{K_3 \beta_x \beta_y}{8\pi} ds \right) J_x$$

- Amplitude-dependent betatron tune shifts
- Tune spread provides Landau damping
- Used since 70s in many machines
- The cornerstone of the mitigation scheme for SIS100 (beam  $\varnothing \approx 30\text{mm}$ )
- Can reduce the Dynamic Aperture



Schematic yoke profile of an octupole magnet

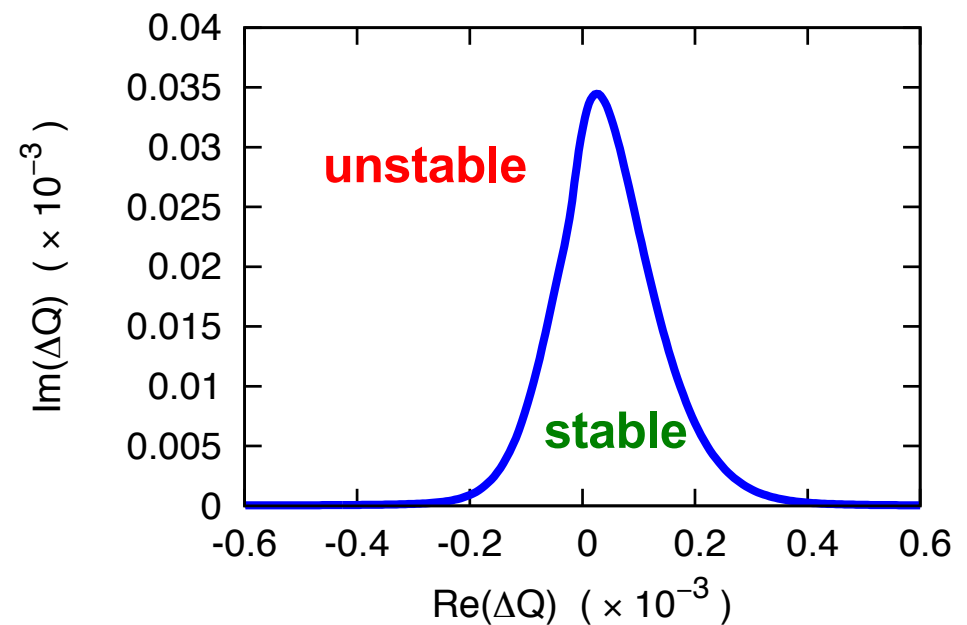
## Analytic calculations

For Landau damping due to octupoles only, the dispersion relation has been commonly used

$$\Delta Q_{\text{coh}} \int \frac{1}{\Delta Q_{\text{oct}} - \Omega/\omega_0} J_x \frac{\partial \psi_{\perp}}{\partial J_x} dJ_x dJ_y = 1$$

Discussed here extensively

But, with space charge,  
it is more difficult



# Landau damping with space charge

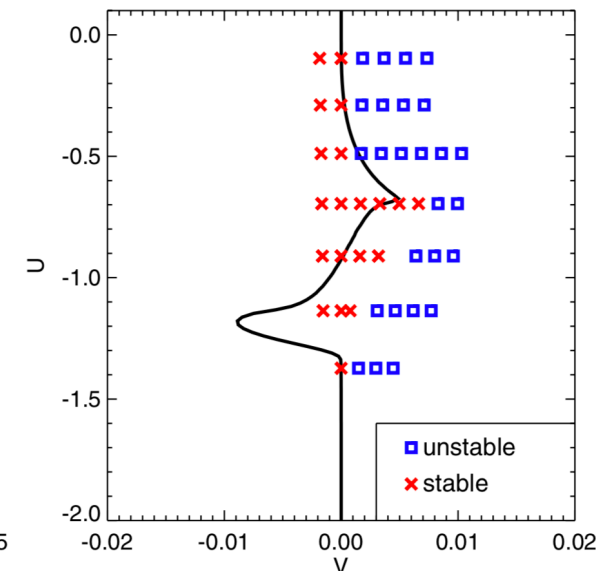
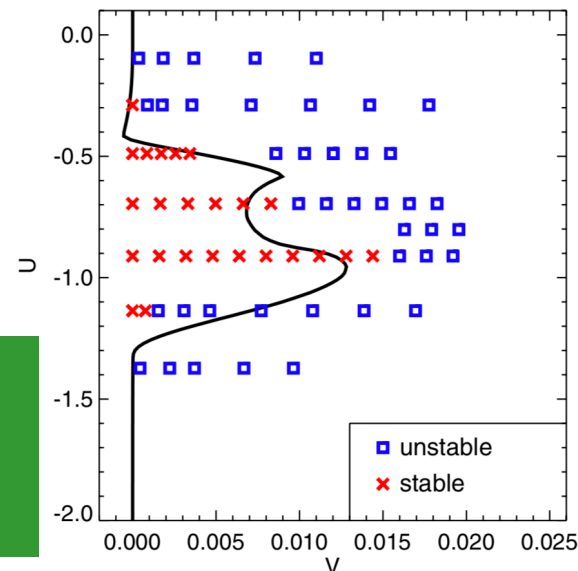
The dispersion relation (D.Möhl, H.Schönauer, 1974)

$$\int \frac{\Delta Q_{\text{coh}} - \Delta Q_{\text{sc}}}{\Delta Q_{\text{ex}} + \Delta Q_{\text{sc}} - \Omega/\omega_0} J_x \frac{\partial f}{\partial J_x} dJ_x dJ_y = 1$$

which is very good  
for monotonic  $(\Delta Q_{\text{ex}} + \Delta Q_{\text{sc}})$ ,  
but can predict  
an antidamping  
for a complicated  
 $(\Delta Q_{\text{ex}} + \Delta Q_{\text{sc}})(J_x, J_y)$

Note the negative shift  
 $U \sim \text{Im}(\Delta Q)$  due to (linear)  
effect of space charge!

V.Kornilov, O.Boine-Frankenheim, I.Hofmann,  
PRSTAB 11, 014201 (2008)



## Landau damping with space charge

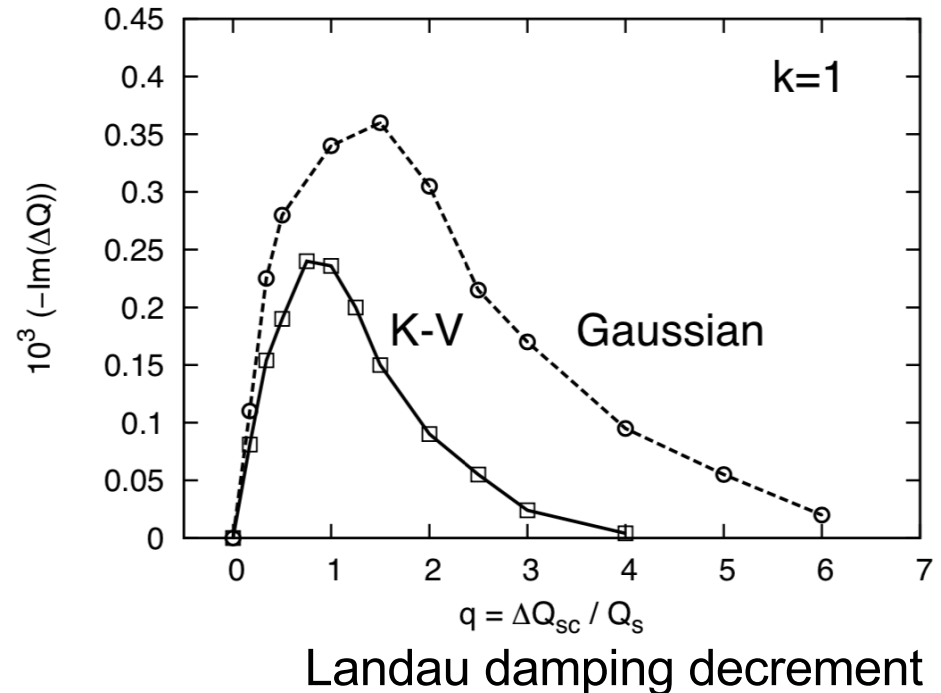
The dispersion relation (D.Möhl, H.Schönauer, 1974)

$$\int \frac{\Delta Q_{\text{coh}} - \Delta Q_{\text{sc}}}{\Delta Q_{\text{ex}} + \Delta Q_{\text{sc}} - \Omega/\omega_0} J_x \frac{\partial f}{\partial J_x} dJ_x dJ_y = 1$$

Predicts (correctly!) for coasting beams, that there is NO Landau damping due to space charge only (even with the overlap coh. mode  $\leftrightarrow$  inc. spectrum). But, there is damping due to space charge only in bunches.

Cannot be directly extended to bunches

V.Kornilov, O.Boine-F, arXiv:1709.01425 (2017)  
V.Kornilov, O.Boine-F, PRSTAB 13, 114201 (2010)



# Particle-in Cell Tracking

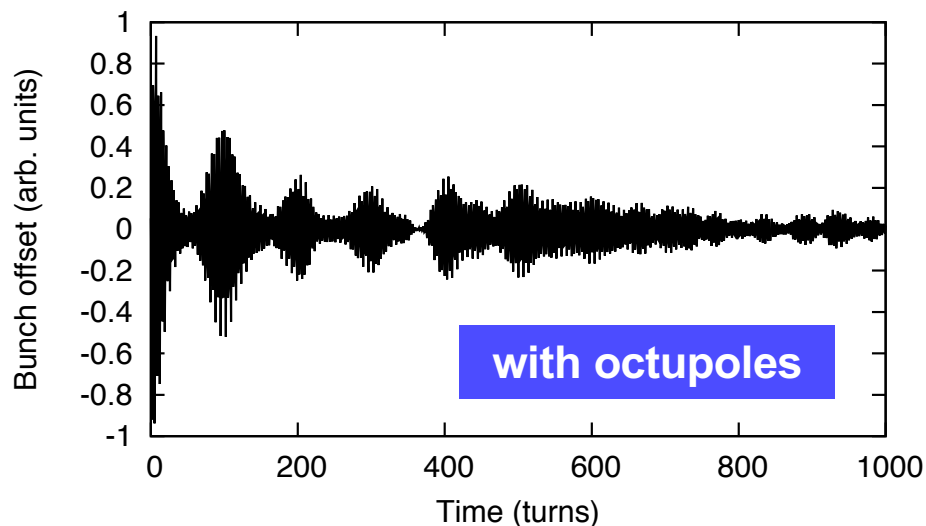
For the mitigation in SIS100, we need the quantitative predictions for Landau damping due to combinations of octupoles and space charge

## The PIC code PATRIC (development GSI Darmstadt)

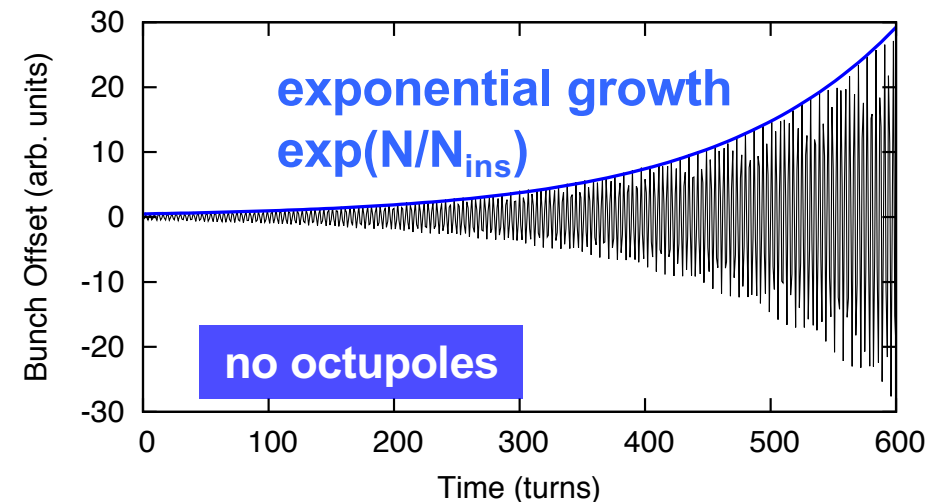
- 2.5D sliced bunches
- Self-consistent space-charge
- Impedances, Wakes, Image Charges
- Tune shifts, spectra, instabilities, decoherence (bunched beams, coasting beams, with space charge, without space charge) verified with analytical theories
  - V. Kornilov and O. Boine-Frankenheim, Proc. of ICAP2009, San Francisco (2009)
  - O.Boine-Frankenheim, V.Kornilov, Proc. of ICAP2006 (2006)
  - V.Kornilov, HB2016, July 3-8, Malmö, Sweden (2016)
- Verified vs. HEADTAIL (2005)
- Landau damping simulations, head-tail modes with space-charge
  - V.Kornilov, O.Boine-Frankenheim, PRSTAB 13, 114201 (2010)
  - V.Kornilov, O.Boine-Frankenheim, PRSTAB 15, 114201 (2012)
  - V.Kornilov, O.Boine-Frankenheim, arXiv:1709.01425 (2017)
- This work: the constant focusing, linear rf bucket, 3D Gaussian  $3.5\sigma$

## Particle tracking simulations

- Start with a tiny perturbation
- Apply a wake field (resistive-wall  $W(z) = w_0/\sqrt{z}$ )
- Apply octupoles



Stabile due to octupoles

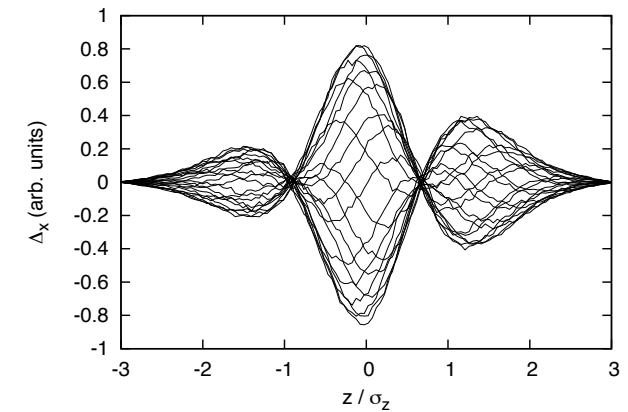
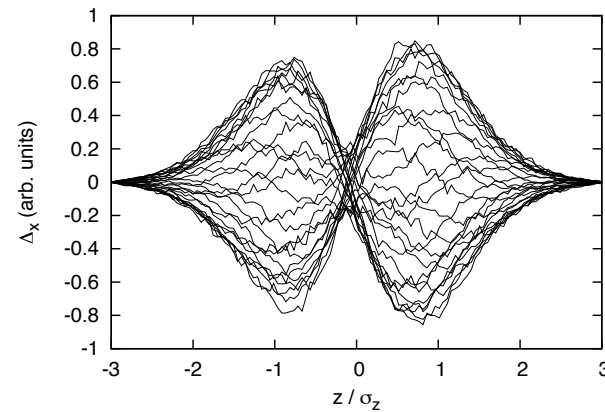
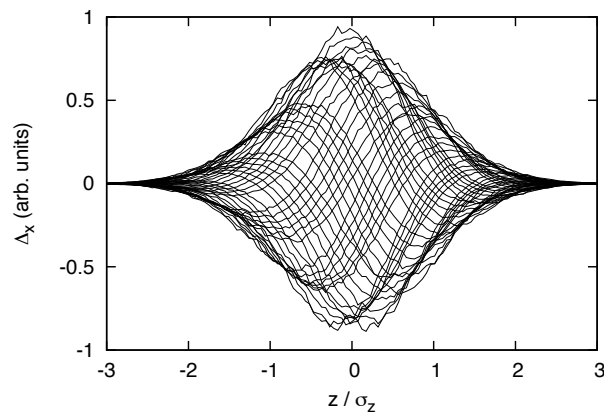


Unstable



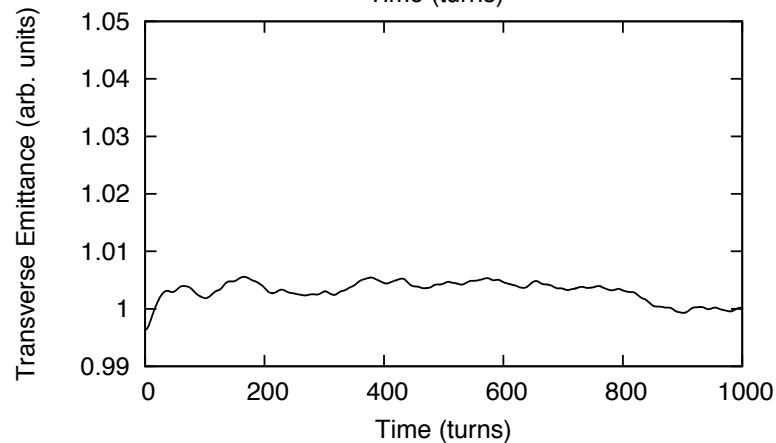
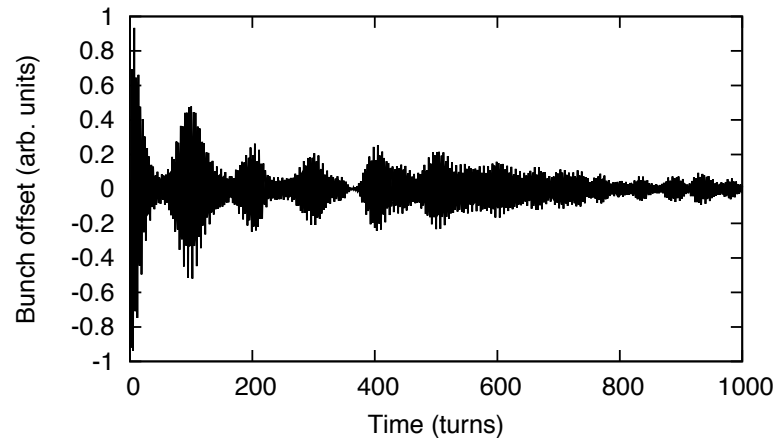
# Particle tracking simulations

## Simulation scans for $k=0$ , $k=1$ and $k=2$ modes

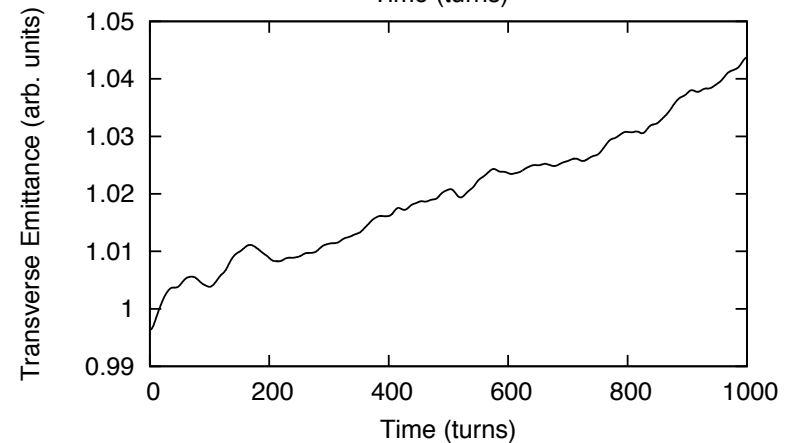
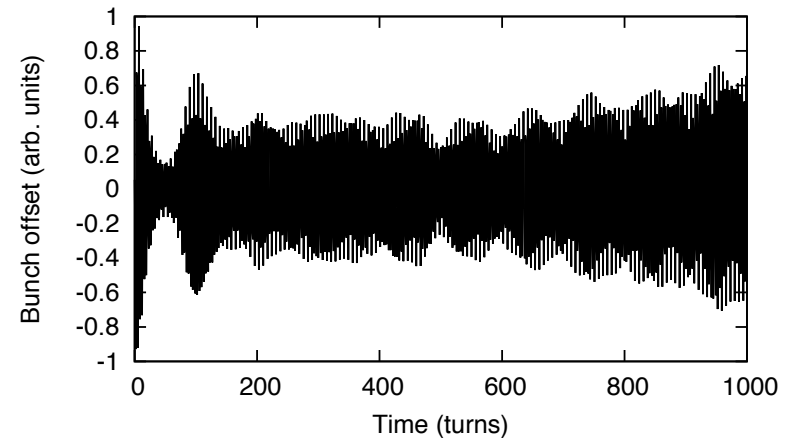


- Varying the chromaticity, the same RW wake
- Intra-Bunch oscillation
- Even the  $k=0$  mode is not a rigid dipole mode

# Particle tracking simulations



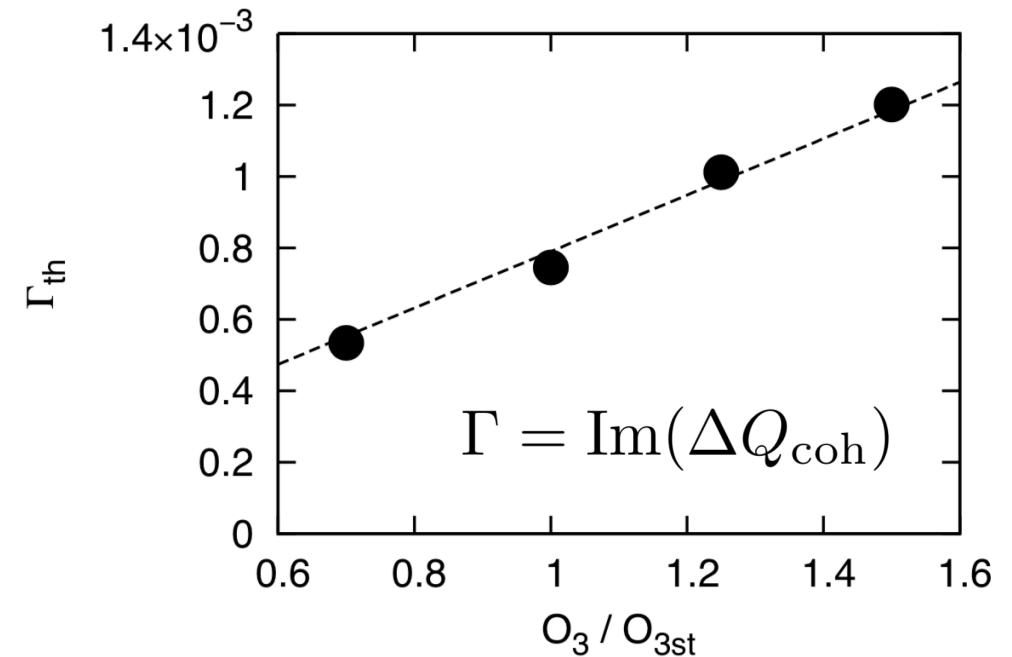
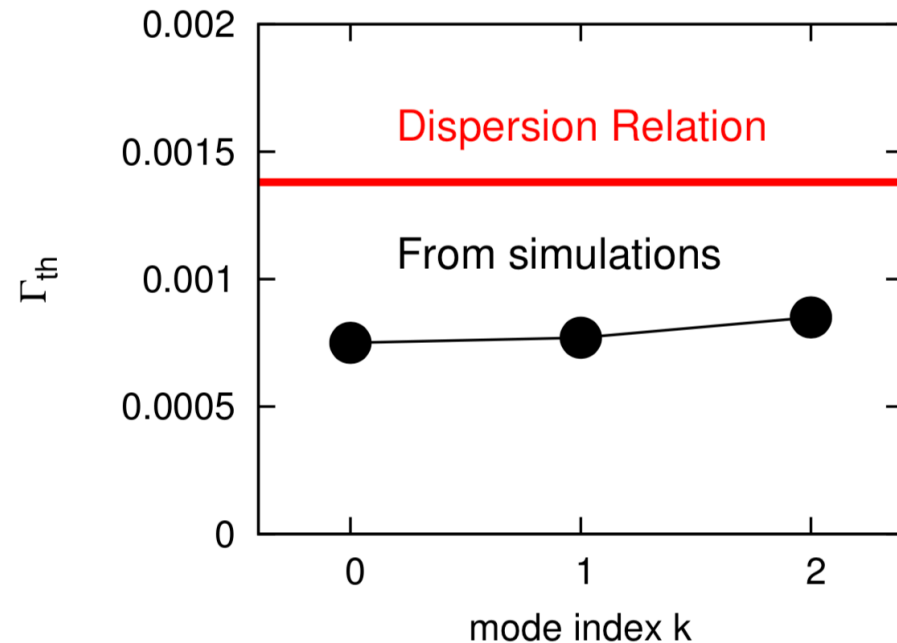
Stabile due to octupoles



Above the threshold

# Damping due to octupoles

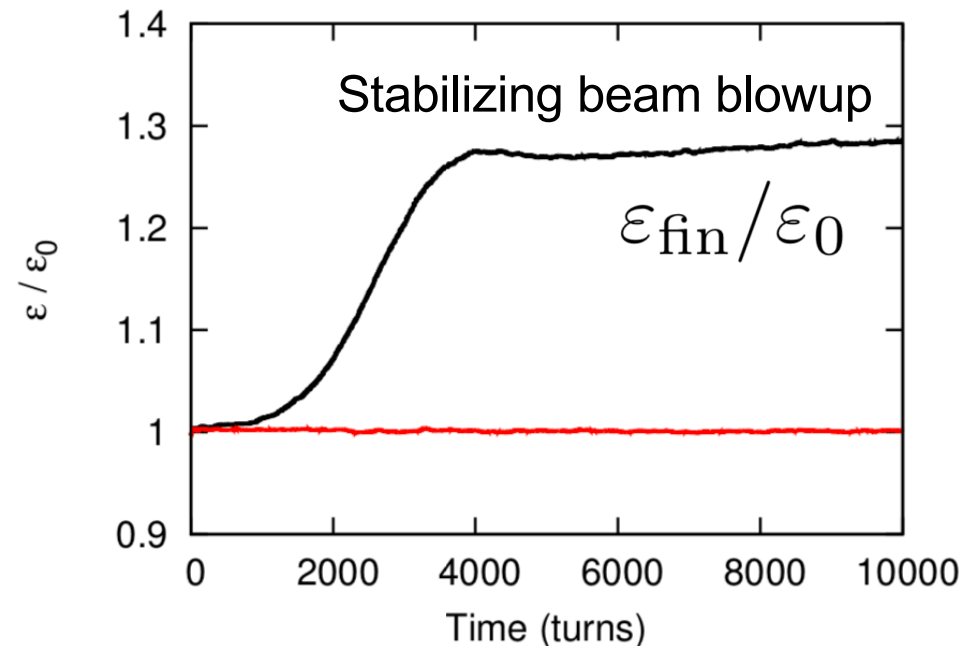
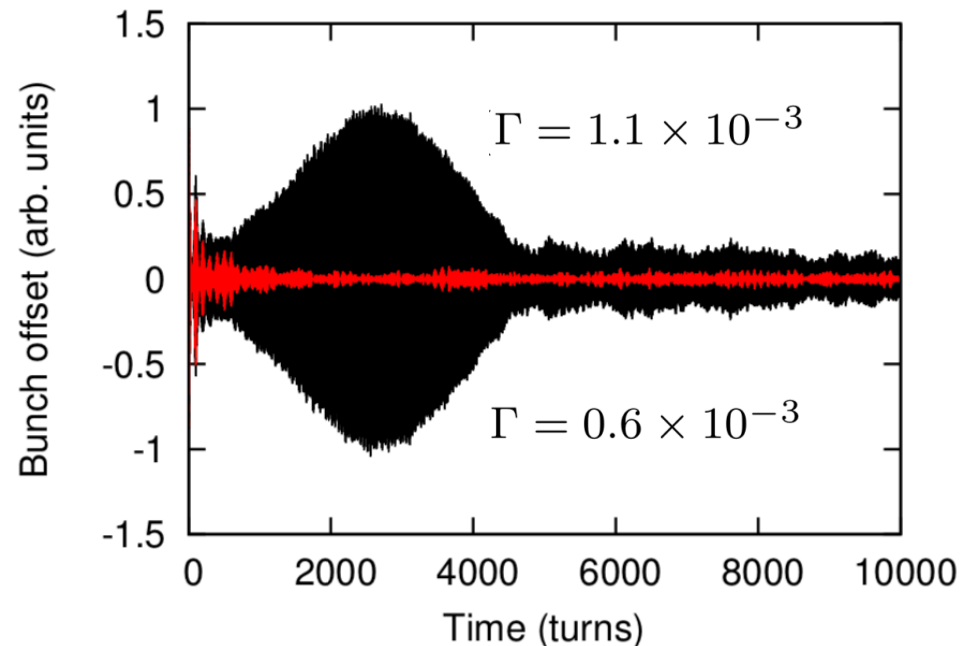
## Simulation scans



- $\Gamma_{th}$  is the stability threshold for a fixed octupole scheme
- Octupoles provide similar Landau damping to  $k=0$ ,  $k=1$ ,  $k=2$ .
- Clear linear dependency of the damping on the octupole power

## Damping due to octupoles

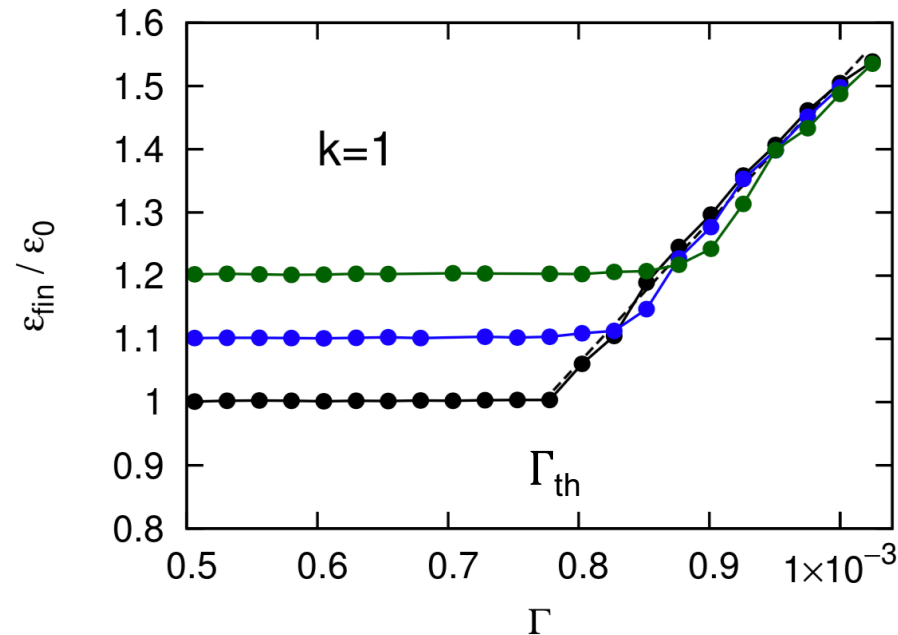
### Octupoles for modes slightly above the stability threshold



Advantage of the octupoles:  
The tune spread is proportional to the transverse emittance.  
Oscillations cause a beam blowup and stabilize the beam.

## Damping due to octupoles

### Octupoles for modes slightly above the stability threshold



Stabilizing  
beam blowup

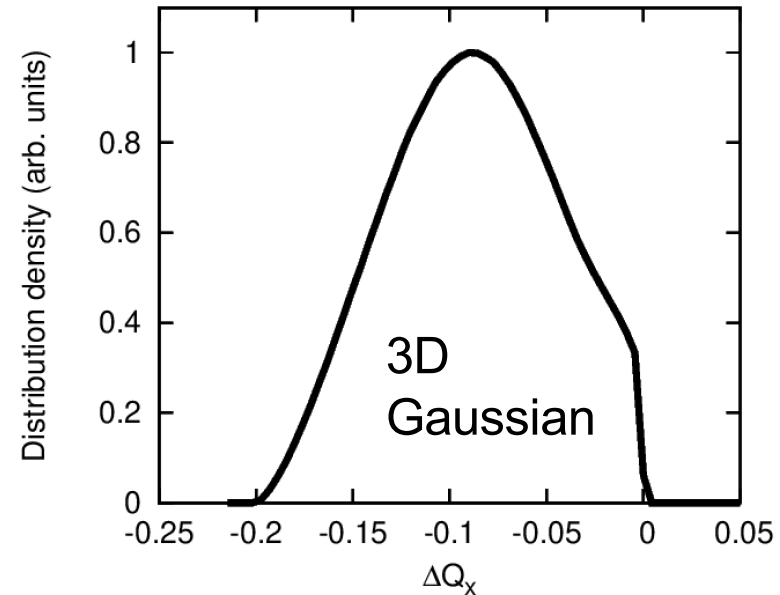
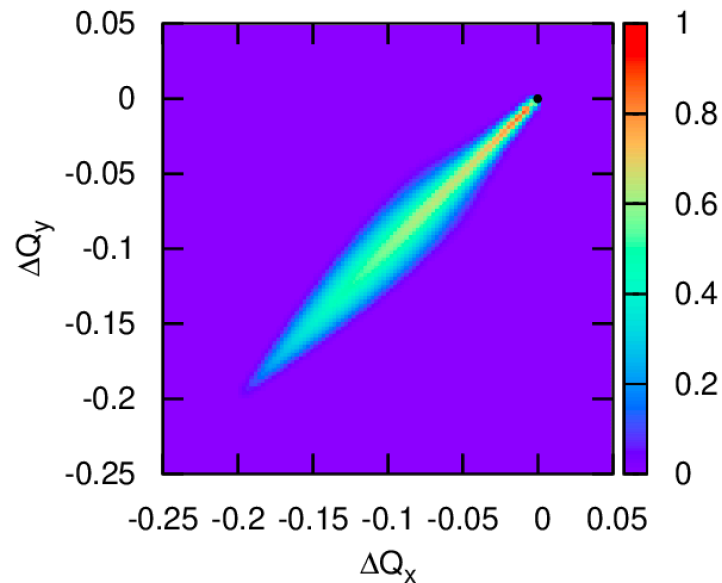
$$\Gamma = \text{Im}(\Delta Q_{\text{coh}})$$

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The tune spread is proportional to the transverse emittance.  
Oscillations cause a beam blowup and stabilize the beam.

# Tune shifts due to space charge

## Tune footprints: space charge



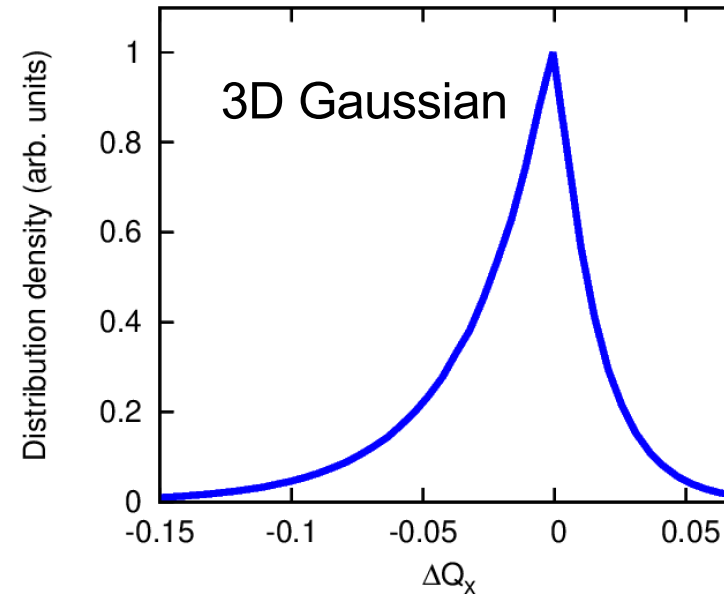
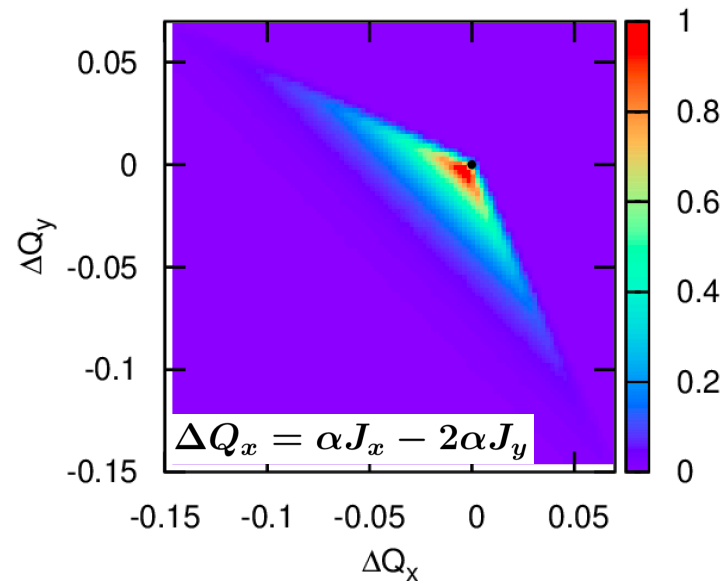
Space charge  $q = \frac{\Delta Q_{sc}}{Q_s}$

Octupoles  $q_4 = \frac{\Delta Q_\sigma}{Q_s}$

Average  $\langle \Delta Q_x \rangle = -0.0878$   
 RMS  $\delta(\Delta Q_x) = 0.0429$   
 $q=10$   
 for  $Q_s=0.01$

# Tune shifts due to octupoles

## Tune footprints: octupoles



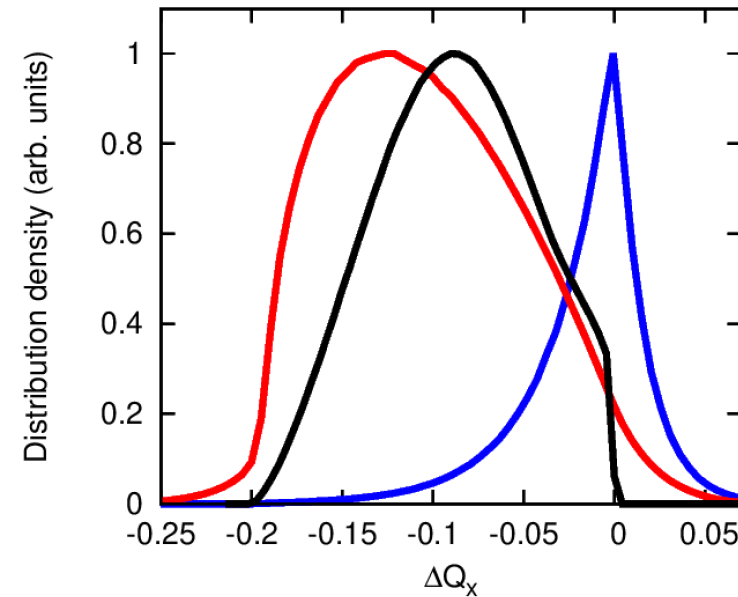
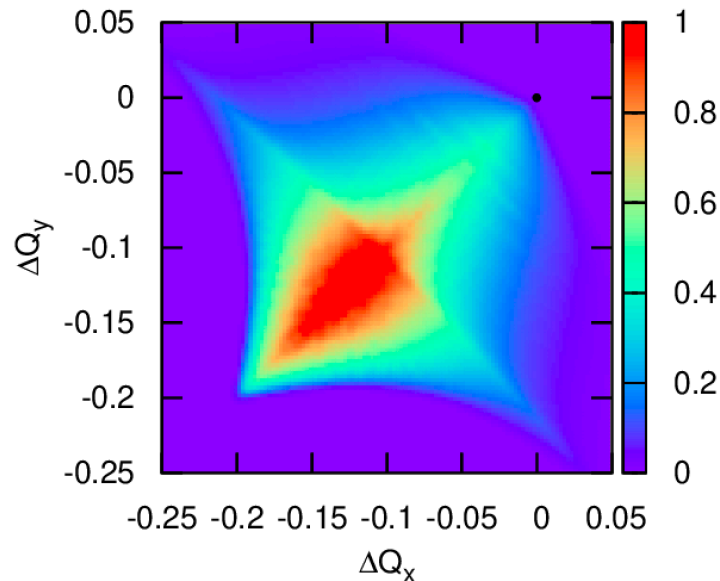
Space charge  $q = \frac{\Delta Q_{sc}}{Q_s}$

Octupoles  $q_4 = \frac{\Delta Q_\sigma}{Q_s}$

Average  $\langle \Delta Q_x \rangle = -0.0157$   
 RMS  $\delta(\Delta Q_x) = 0.034$   
 $\Delta Q_\sigma = 0.008$   
 $q_4 = 0.8$   
 $\Delta Q_\sigma$  for a particle with  $a_x = \sigma_x, a_y = 0$

# Tune shifts due to octupoles

## Tune footprints: octupoles + space charge



Octupoles  
Space charge  
Combined

Space charge  $q = \frac{\Delta Q_{sc}}{Q_s}$

Octupoles  $q_4 = \frac{\Delta Q_\sigma}{Q_s}$

Average  $\langle \Delta Q_x \rangle = -0.104$   
 RMS  $\delta(\Delta Q_x) = 0.0534$   
 $q_4 = 0.8$   
 $q = 10$



## Landau damping for the k=0 mode

### Results of the simulation scans

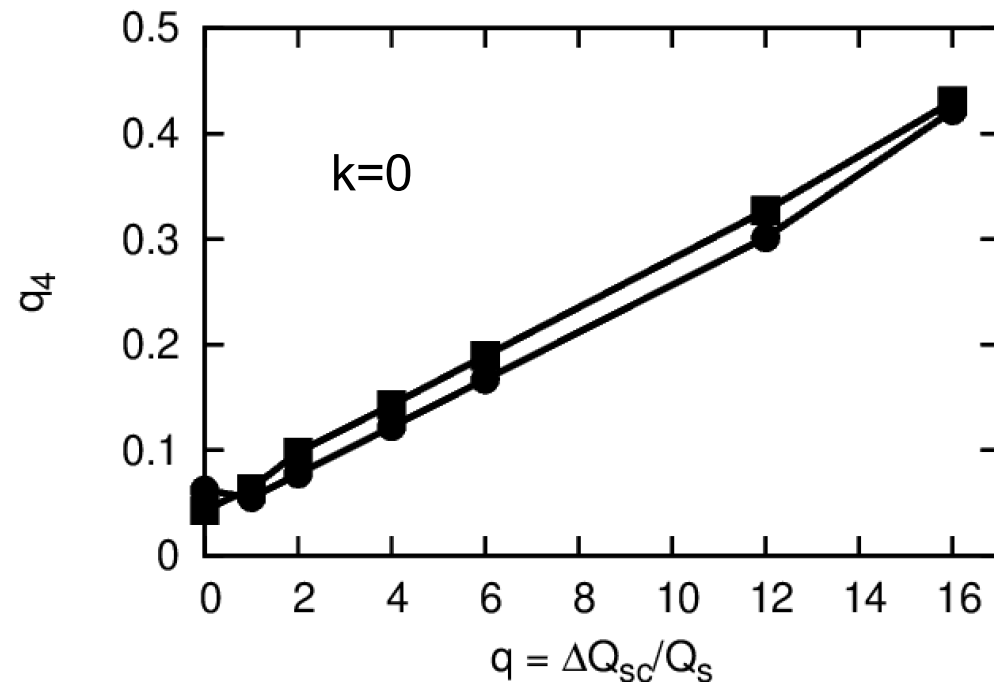
$$q_4 = \frac{\Delta Q_\sigma}{Q_s}$$

Minimal octupole power for the stability as a function of the space-charge strength

Octupole polarity:

Circles: positive  $q_4$

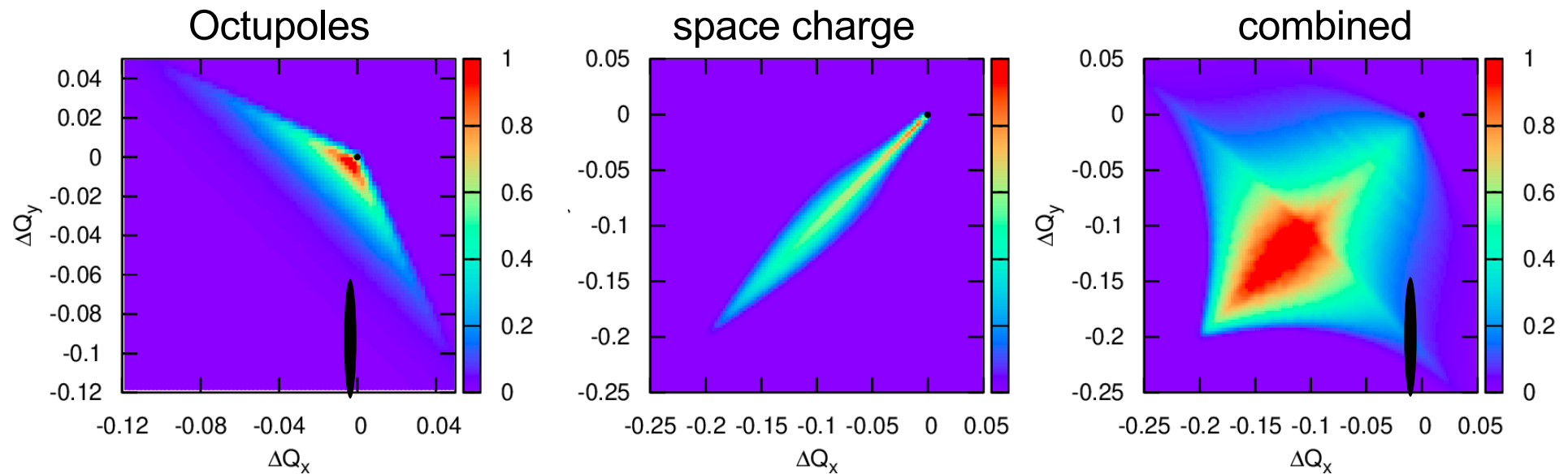
Squares: negative  $q_4$



- For stability, higher octupole power at stronger space charge is needed
- Loss of Landau damping due to space charge
- As predicted, no Landau damping due to space charge for k=0

# Tune footprint for octupole + space charge

## Tune footprints: octupoles + space charge



$$q_4=0.8$$

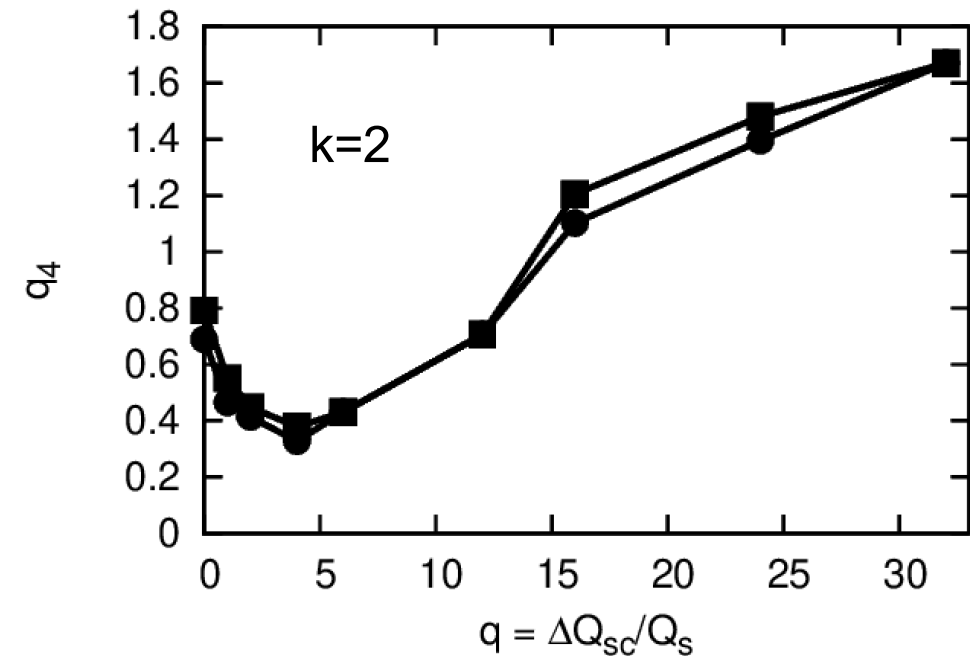
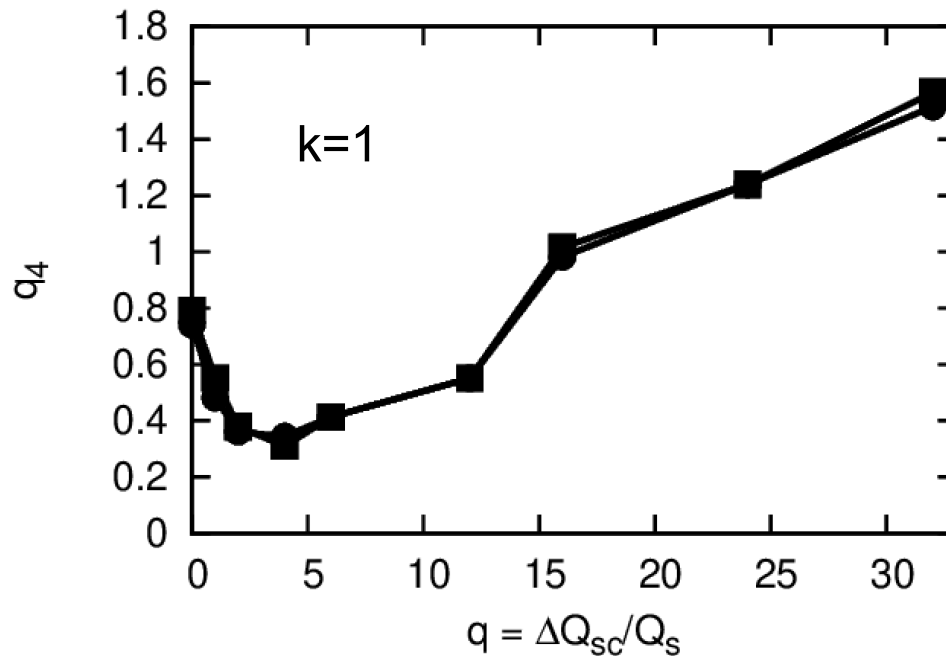
$$q=10$$

Black: schematic for the frequency of the coherent mode

Similar to the case of a coasting beam:  
loss of Landau damping due to space charge

# Landau damping for $k=1$ , $k=2$ modes

## Results of the simulation scans



$$q_4 = \frac{\Delta Q_\sigma}{Q_s}$$

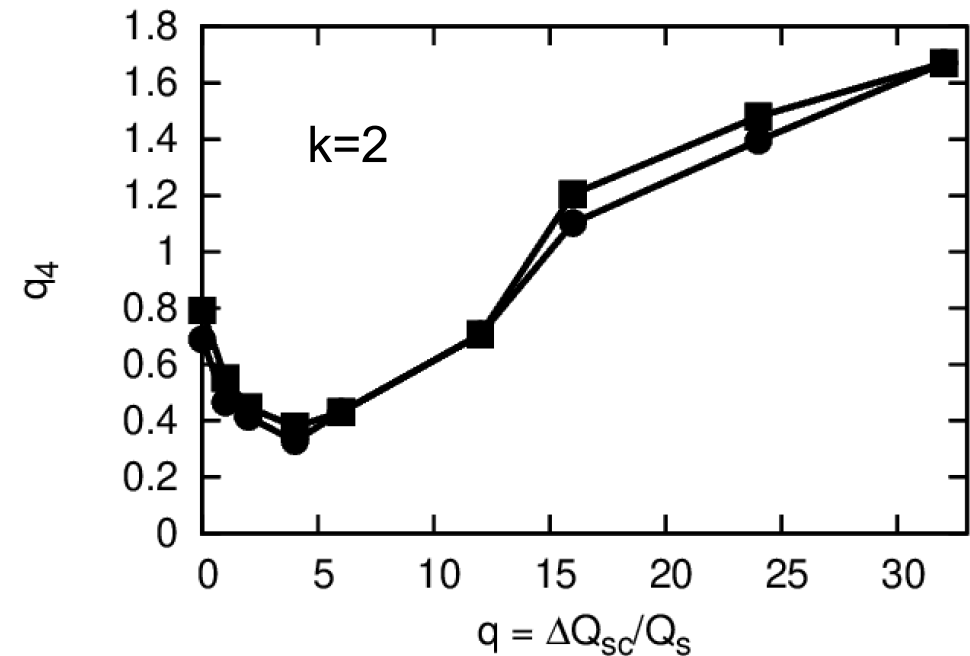
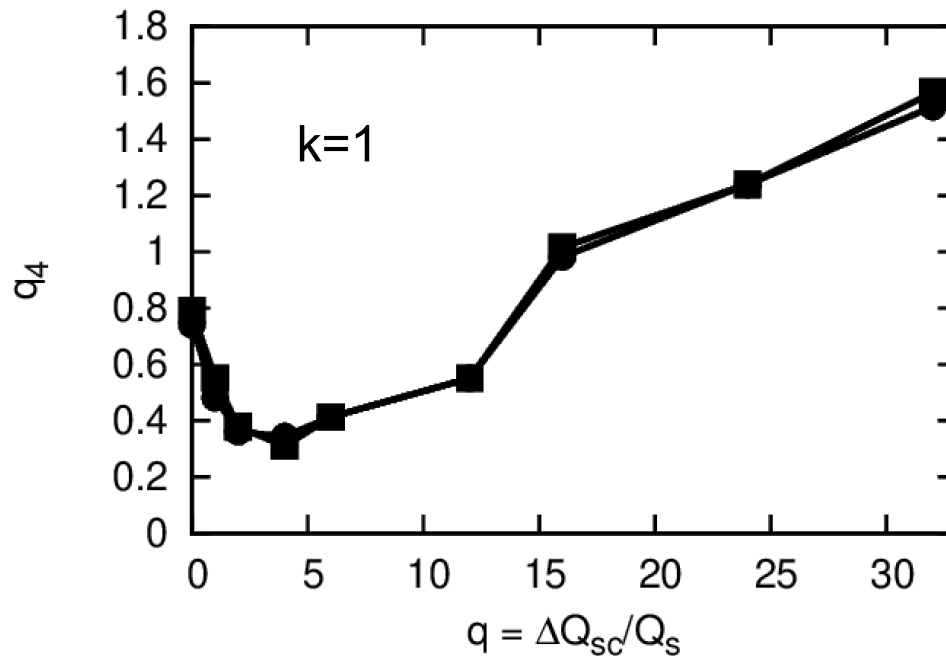
Minimal octupole power for the stability as a function of the space-charge strength

Octupole polarity: Circles: positive  $q_4$

Squares: negative  $q_4$

# Landau damping for $k=1$ , $k=2$ modes

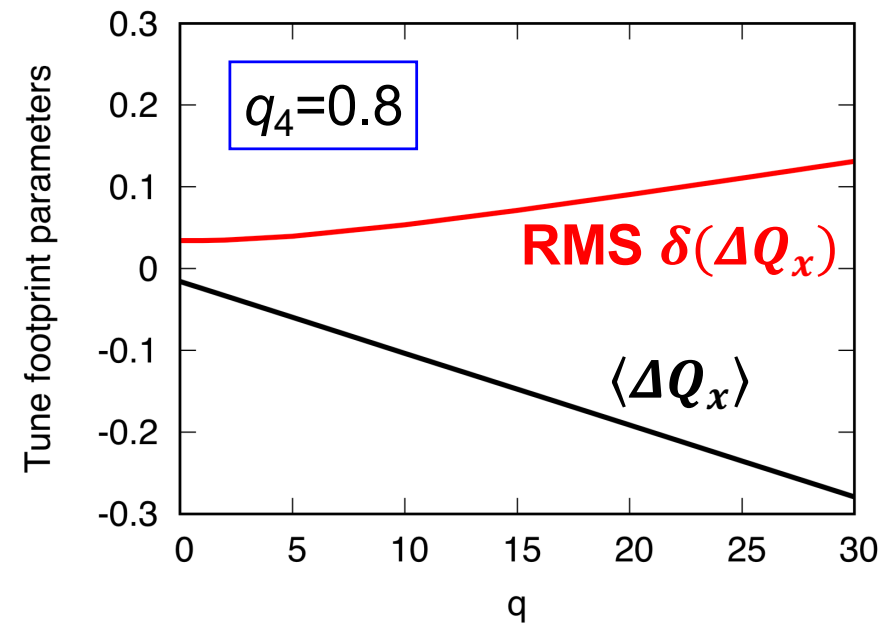
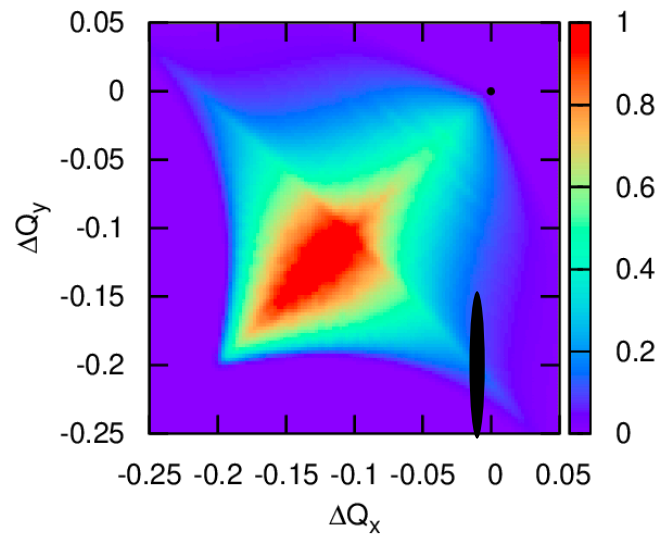
## Results of the simulation scans



- For stability, higher octupole power at stronger space charge is needed
- Loss of Landau damping due to space charge
- Additional Landau damping due to space charge at medium  $q$

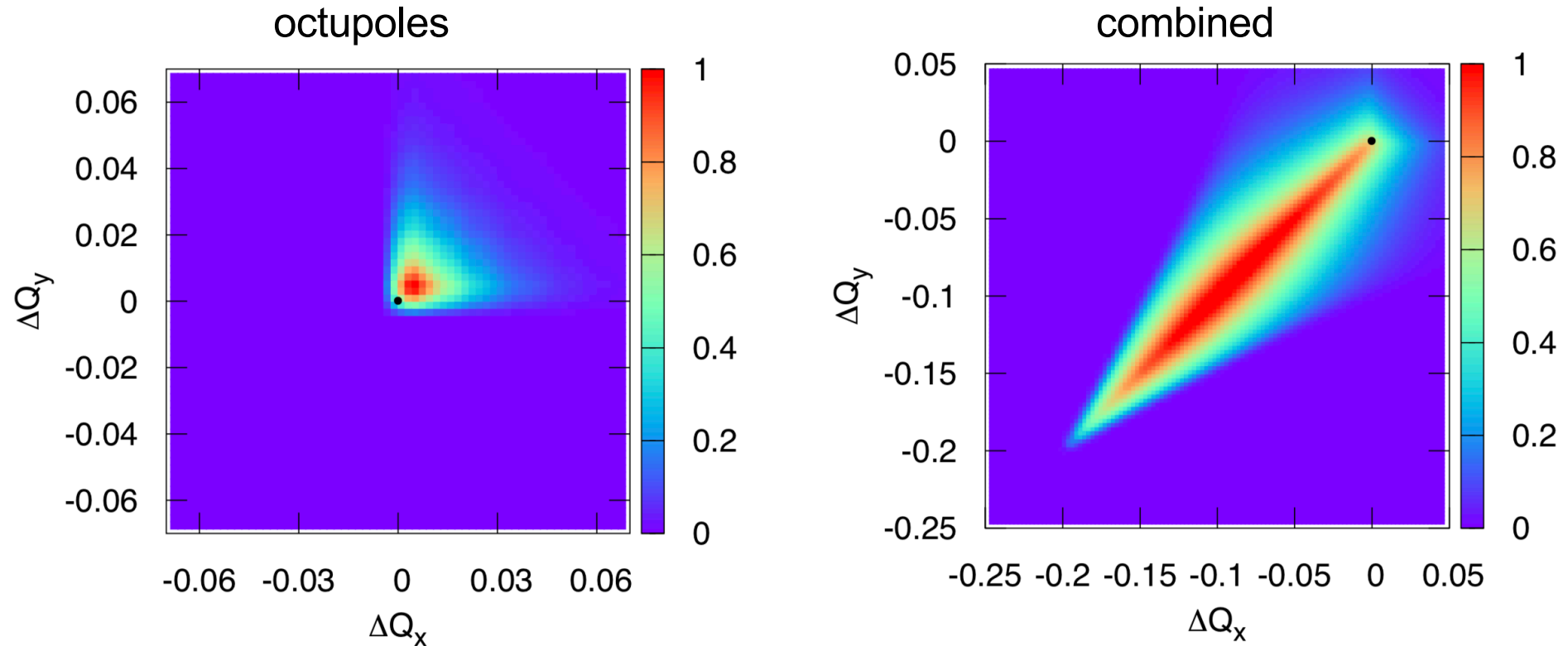
# Landau damping for $k=1$ , $k=2$ modes

Tune footprints: octupoles + space charge



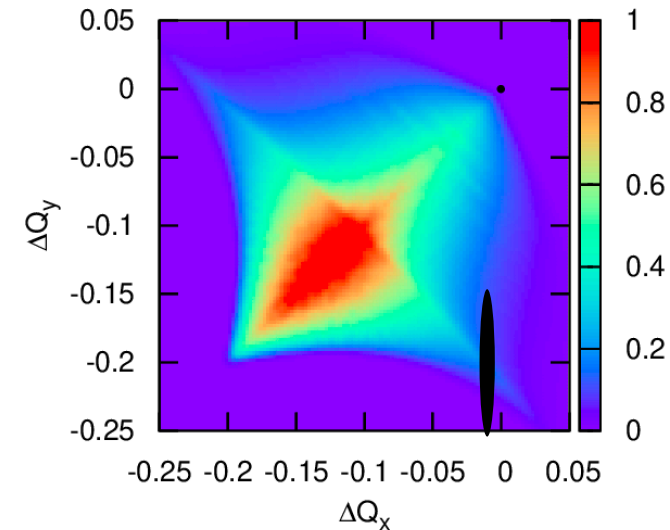
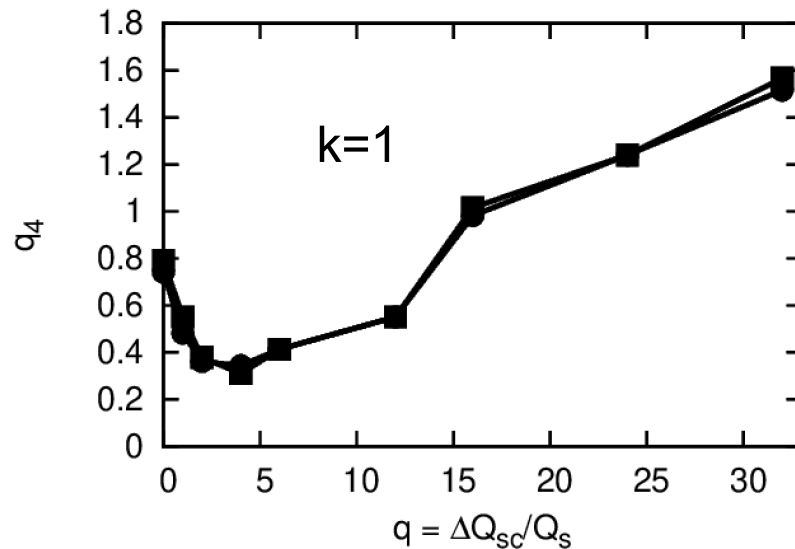
Loss of Landau damping due to space charge

# Octupole optimization



If groups of octupoles are available (SIS100: 2), optimization of the power scheme is possible

## Conclusions



- Octupoles scheme is a cornerstone of the mitigation at SIS100
- Octupoles: the stabilizing beam blowup
- Important role of space charge for the mitigation using octupoles
- Loss of Landau damping due to space charge (still enough for SIS100)
- Additional Landau damping due to space charge at medium  $q$  for  $k>0$
- The flexible octupole scheme provides optimized footprints for stability