TeV Scale Seesaw Mechanism, Singlet Scalar Dark Matter and Electroweak Vacuum Stability Based on PhysRevD.96.055020<sup>1</sup>

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- Introduction
- TeV Scale Seesaw model : Inverse Seesaw
- Singlet Scalar DM
- $\bullet~$  SM + Inverse Seesaw + Singlet Scalar and EW vacuum stability
- Numerical Analysis and Results

- The Standard Model (SM) of particle physics is a very successful theory
- Neutrino oscillation  $\implies$  neutrinos have Mass and Mixing
- The first indication towards the need for a theory beyond SM
- Another issue that the SM does not have an answer to : the existence of Dark Matter (DM)
- These issues could be addressed either by extending just the particle content or by extending the gauge group
- It is important to study the implications of the BSM models that can solve these issues

- The most natural approach towards understanding the sub-eV neutrino mass scale
- Neutrinos are Majorana particles and the lepton number must be explicitly violated at a high-energy scale
- Tree level exchange of some heavy particle present at a higher energy  $\implies$  Effective dimension-5 operator \*  $\frac{\kappa_5 LL\Phi\Phi}{M}$  at low scale



• Introduce 3 heavy right handed Majorana neutrinos ( $N_R$ ) into the Standard Model : Type-1 seesaw mechanism <sup>†</sup>

<sup>†</sup>Minkowski(1977), Mohapatra and Senjanovic (1980) 🧃 🕨

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<sup>\*</sup>Weinberg (1979)

### TeV Scale Seesaw Mechanism

- The minimal type 1 seesaw model is not testable
- Motivates us to look for testable TeV scale seesaw models
- To the type-1 seesaw picture , add 3 additional gauge-singlet neutrinos with opposite lepton number,  $S_R^i$  (i = 1, 2, 3)

$$-L_{\nu} = \bar{I}_{L} Y_{\nu} H^{c} N_{R} + \bar{N}_{R}^{c} M_{S} S_{R} + \frac{1}{2} \bar{S}_{R}^{c} M_{\mu} S_{R} + \text{h.c.}$$

• Once the Higgs field H acquire a vev (v),  $M_D = Y_
u v/\sqrt{2}$  ,

$$-L_{mass} = \frac{1}{2} \begin{pmatrix} \bar{\nu}_L & \bar{N}_R^c & \bar{S}_R^c \end{pmatrix} \begin{pmatrix} 0 & M_D & 0 \\ M_D^T & 0 & M_S \\ 0 & M_S^T & M_\mu \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \\ S_R \end{pmatrix} + \text{h.c.}$$

• The mass scales of three sub-matrices of M may naturally have a hierarchy  $M_S>>M_D>>M_\mu$ 

• 
$$\implies$$
  $M_{\nu} = M_D(M_S^T)^{-1}M_{\mu}M_S^{-1}M_D^T$ : Inverse Seesaw Mechanism

- The smallness of  $M_{\nu}$  is naturally attributed to both the smallness of  $M_{\mu}$  and the smallness of  $\frac{M_D}{M_S}$ .
- $M_{\nu} \approx O(0.1) \, eV$  can easily be achieved from  $\frac{M_D}{M_S} \approx 10^{-2}$  and  $M_{\mu} \approx O(1 \, keV)$
- Lepton number is softly broken by  $M_{\mu}$
- $M_{
  u}$  goes to 0 in the limit of  $M_{\mu}$  going to 0
- Heavy neutrino masses : a few 100 GeV to a few TeV
- Can give large unitarity violation and lepton flavour violating radiative decays
- BR  $(\mu 
  ightarrow e \gamma)~pprox~10^{-14}$

### Singlet Scalar with a $Z_2$

- $\bullet\,$  Nearly 95 percent of the Universes matter density is dark ;  $\sim\,$  26 percent DM
- Among the various models of DM that are proposed, the most minimal extension of the SM : Higgs portal models
- Here, we add a real scalar singlet, A, to the SM with a discrete  $Z_2$  symmetry and 0 vev
- The new scalar potential becomes,

$$V = V_{SM} + \frac{1}{2}m_A^2 A^2 + \frac{\kappa}{2}H^{\dagger}H A^2 + \frac{\lambda_A}{4}A^4$$
$$V_{SM} = m^2 H^{\dagger}H + \lambda (H^{\dagger}H)^2$$

 In a model with SM extended by a TeV seesaw + real singlet scalar, could there be a connection between the two seemingly disconnected sectors ? YES

- $\bullet$  Quantum loop corrections will make the mass parameter and coupling dependent on the energy scale  $\Lambda$
- The  $\lambda$  for the quartic term is running with  $\Lambda$  as :  $\Lambda \frac{d}{d\Lambda} = \beta_{\lambda}$
- At one-loop order,

$$\beta_{\lambda SM} = \frac{1}{(4\pi)^2} \left[ 24\lambda^2 - 6y_t^4 + \frac{3}{8} \left( 2g^4 + (g^2 + g'^2)^2 \right) + (-9g^2 - 3g'^2 + 12y_t^2) \lambda \right]$$

• The relative sign between bosonic and fermionic contributions would dramatically affect the UV behaviour of the theory

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### Vacuum Stability in the Standard Model

- Due to the heavy quarks contribution
- $\lambda$  runs as  $-y_t^4$  : The large top yukawa coupling will pull  $\lambda$  down to negative values at higher energies
- Then the potential might develop a new minimum at a higher energy scale ⇒ The EW vacuum may be unstable due to quantum tunnelling
- In SM,  $\lambda$  becomes negative at an energy scale of  $10^9-10^{10}~{\rm GeV}$  depending on the values of  $\alpha_S$  and  $y_t$  used
- This is not a threat to the theory as long the decay time is greater than the age of the universe ⇒ SM vacuum is metastable
- This gives a bound on  $\lambda$  :

$$\lambda(\Lambda_B) > \lambda_{\min}(\Lambda_B) = \frac{-0.06488}{1 - 0.00986 \ln(v/\Lambda_B)}$$

#### Effective Higgs Potential and $\lambda_{eff}$

• The tree level Higgs potential in the SM is given by,

$$V(h) = \frac{m^2}{2}h^2 + \frac{\lambda}{4}h^4$$

• This will get corrections from higher order loop diagrams

$$V_1^{SM+A+\nu}(h) = V_1^{SM}(h) + V_1^A(h) + V_1^{\nu}(h)$$

 For h >> v, the effective potential could be approximated as, (vacuum instability appears at a scale >> ew vacuum)

$$V_{eff}^{SM+A+
u} = \lambda_{eff}(h,t)rac{h^4}{4}$$

$$\lambda_{eff}(h) \,=\, \lambda_{eff}^{SM}(h) \,+\, \lambda_{eff}^{A}(h) \,+\, \lambda_{eff}^{
u}(h)$$

• Modified RGE for  $\lambda$  :  $\beta_{\lambda} = \beta_{\lambda SM} + \frac{1}{(4\pi)^2} [+ 4\kappa^2 - 2\text{Tr}(Y_{\nu}^{\dagger}Y_{\nu})^2]$ 

- The vacuum stability/metastability analysis is done considering following constraints on both sectors.
- Bounds on the scalar sector:
  - Perturbative unitarity bound.
  - Dark matter relic abundance.
  - Bounds on dark matter mass and Higgs portal coupling from Higgs invisible decay width, direct (LUX-2016) and indirect detection (Fermi-LAT).
- The low energy constraints on the lepton sector:
  - Oscillation data : Mixing angles and Mass squared differences.
  - Constraints on the non-unitarity of PMNS matrix.
  - Bounds on heavy neutrino masses and light-heavy mixing from LEP experiments.
  - BR  $(\mu \rightarrow e\gamma) < 4.2 \times 10^{-13}$ .









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## Phase Diagram in the $Tr[Y_{\nu}^{\dagger}Y_{\nu}] - \kappa$ Plane



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- The lack of experimental testability of canonical type 1 seesaw mechanism motivates us to consider TeV scale seesaw mechanisms
- One of the most studied TeV scale seesaw models is the inverse seesaw model where the smallness of  $m_{\nu}$  is naturally attributed to the smallness of a LNV parameter  $M_{\mu}$
- As a result of low mass thresholds, large values of  $Y_{\nu}$  get constraints from vacuum stability considerations
- Adding the extra scalar allows us to have even larger values of  $Y_{\nu}$ , giving rise to larger values of *LFV* decay rates and non-unitarity, and also having a stable DM candidate at the same time
- Two seemingly disjoint sectors connected by vacuum stability constraints

# THANK YOU

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