

# Consequences of CP transformed mixed $\nu_\mu$ - $\nu_\tau$ antisymmetry

Probir Roy

Centre for Astroparticle Physics and Space Science  
Bose Institute, Kolkata

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**Collaborators** R. Sinha and A. Ghosal of SINP

# PLAN OF THE TALK

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# 1. INTRODUCTION

Statement of  $CP^{\theta\mu\tau A}$  :  $\mathcal{G}^\theta M_\nu \mathcal{G}^\theta = -M_\nu^*$  Sinha, Roy, Ghosal (2018)  
under  $\nu_{L\ell} \rightarrow i\mathcal{G}_{lm}^\theta \gamma^0 \nu_{Lm}^C$

with generator  $\mathcal{G}^\theta = i \begin{pmatrix} -1 & 0 & 0 \\ 0 & -\cos \theta & \sin \theta \\ 0 & \sin \theta & \cos \theta \end{pmatrix}.$

Mixing angle  $\theta \in [0, \pi/2]$  with  $\theta = \pi/2 \leftrightarrow$  exact CP transformed  $\nu_\mu \nu_\tau$  antisymmetry. Most general form of  $M_\nu$  is

$$M_\nu^{CP\theta A} = \begin{pmatrix} ix & a_1 + ia_2 & a_1 t_{\frac{\theta}{2}}^{-1} - ia_2 t_{\frac{\theta}{2}} \\ a_1 + ia_2 & y_1 + iy_2 & y_1 c_\theta s_\theta^{-1} + ic \\ a_1 t_{\frac{\theta}{2}}^{-1} - ia_2 t_{\frac{\theta}{2}} & y_1 c_\theta s_\theta^{-1} + ic & -y_1 + i(y_2 + 2cc_\theta s_\theta^{-1}) \end{pmatrix}$$

with 7 real parameters  $x, a_{1,2}, c, y_{1,2}, \theta$ . Here  $c_\theta \equiv \cos \theta, s_\theta \equiv \sin \theta, t_{\frac{\theta}{2}} \equiv \tan \frac{\theta}{2}.$

## 2. NEUTRINO MIXING ANGLES AND PHASES

Lam's observation:

Lam (2007)

$$\mathcal{G}^\theta U^* = U \tilde{d}, \quad \tilde{d} = \text{diag}(\tilde{d}_1, \tilde{d}_2, \tilde{d}_3), \quad d_{1,2,3} = \pm 1.$$

$$U = \text{diag}(e^{i\phi_1}, e^{i\phi_2}, e^{i\phi_3}) U_{\text{PMNS}},$$

$$U_{\text{PMNS}} = \begin{pmatrix} c_{12}c_{13} & e^{i\frac{\alpha}{2}}s_{12}c_{13} & s_{13}e^{-i(\delta - \frac{\beta}{2})} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & e^{i\frac{\alpha}{2}}(c_{12}c_{23} - s_{12}s_{13}s_{23}e^{i\delta}) & c_{13}s_{23}e^{i\frac{\beta}{2}} \\ s_{12}s_{23} - c_{12}s_{13}c_{23}e^{i\delta} & e^{i\frac{\alpha}{2}}(-c_{12}s_{23} - s_{12}s_{13}c_{23}e^{i\delta}) & c_{13}c_{23}e^{i\frac{\beta}{2}} \end{pmatrix}.$$

Algebraic matching leads to

$$e^{i\alpha} = \tilde{d}_1 \tilde{d}_2, \quad e^{2i(\delta - \frac{\beta}{2})} = \tilde{d}_1 \tilde{d}_3$$

$$\Rightarrow \alpha = 0 \text{ or } \pi, \text{ and } \beta = 2\delta \text{ or } 2\delta - \pi.$$

Moreover,  $\cot 2\theta_{23} = \cot \theta \cos(\phi_2 - \phi_3)$ ,  $\sin \delta = \pm \sin \theta / \sin 2\theta_{23}$ ,  
i.e.  $\theta \rightarrow \pi/2 \Rightarrow \theta_{23} \rightarrow \frac{\pi}{4}$ . In general,  $\theta_{23} \neq \pi/4$  and  $\delta \neq 0$  or  $\pi$ .

### 3. NUMERICAL ANALYSIS

Input mixing angles and mass-squared differences from latest global analysis.

Esteban et al (2017)

Neutrino mass sum  $m_1 + m_2 + m_3 < 0.17$  eV from Planck data.

Aghanim et al (2016)

Table: Input  $3\sigma$  ranges used in the analysis

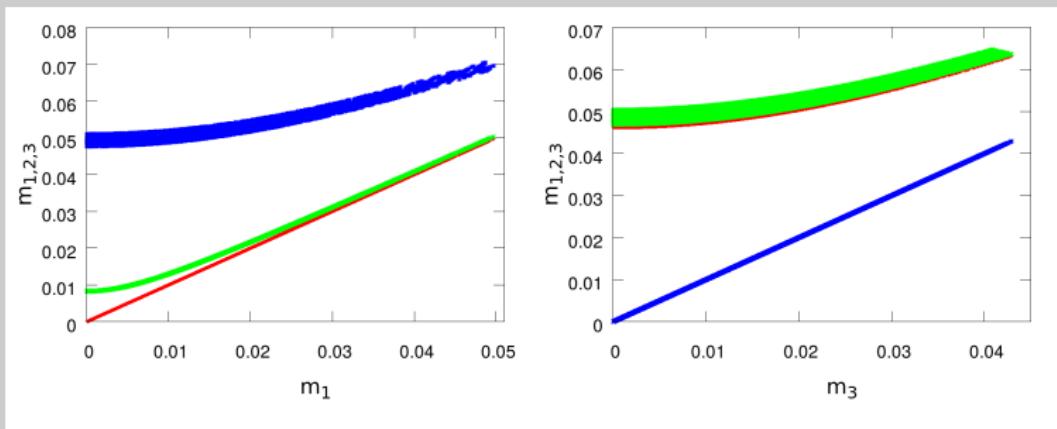
Values	$\theta_{12}$ degrees	$\theta_{23}$ degrees	$\theta_{13}$ degrees	$\Delta m_{21}^2$ $10^{-5}$ eV $^2$	$ \Delta m_{31}^2 $ $10^{-3}$ (eV $^2$ )
NO	31.42 to 36.05	40.3 to 51.5	8.09 to 8.98	6.80 to 8.02	2.399 to 2.593
IO	31.43 to 36.06	41.3 to 51.7	8.14 to 9.01	6.80 to 8.02	2.399 to 2.593

Table: Output values of the parameters of  $M_\nu$

Values	$10^3 x$	$10^3 a_1$	$10^3 a_2$	$10^3 y_1$	$10^3 y_1$	$10^3 c$	$\theta(^{\circ})$
NO	-22 to 22	-45 to 45	-32 to 32	-35 to 35	-45 to 45	-35 to 35	12 to 164
IO	-25 to 25	-45 to 45	-4 to 4	-25 to 25	-35 to 35	-25 to 25	2 to 156

Table: Predictions on the light neutrino masses.

Normal Ordering ( $m_3 > m_2$ )			Inverted Ordering ( $m_3 < m_1$ )		
$10^3 m_1$ (eV)	$10^{-3} m_2$ (eV)	$10^3 m_3$ (eV)	$10^3 m_1$ (eV)	$10^3 m_2$ (eV)	$10^3 m_3$ (eV)
$8.4 \times 10^{-2} - 49$	$9 - 51$	$50 - 71$	$48 - 64$	$49 - 66$	$4.4 \times 10^{-2} - 42$



Neutrino masses for normal (left) and inverted (right) ordering against the lightest mass eigenvalue. The red, green and blue bands refer to  $m_1$ ,  $m_2$  and  $m_3$  respectively.

## 4. NEUTRINOLESS DOUBLE BETA DECAY



Half-life  $T_{1/2}^{0\nu} = G_{0\nu} |\mathcal{M}|^2 |M_\nu^{ee}|^2 m_e^{-2}$ ,

$G_{0\nu}$  = two-body phase space factor,  $\mathcal{M}$  = nuclear matrix element,

$$M_\nu^{ee} = c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{i\alpha} + s_{13}^2 m_3 e^{i(\beta - 2\delta)}$$

Four cases in our model.

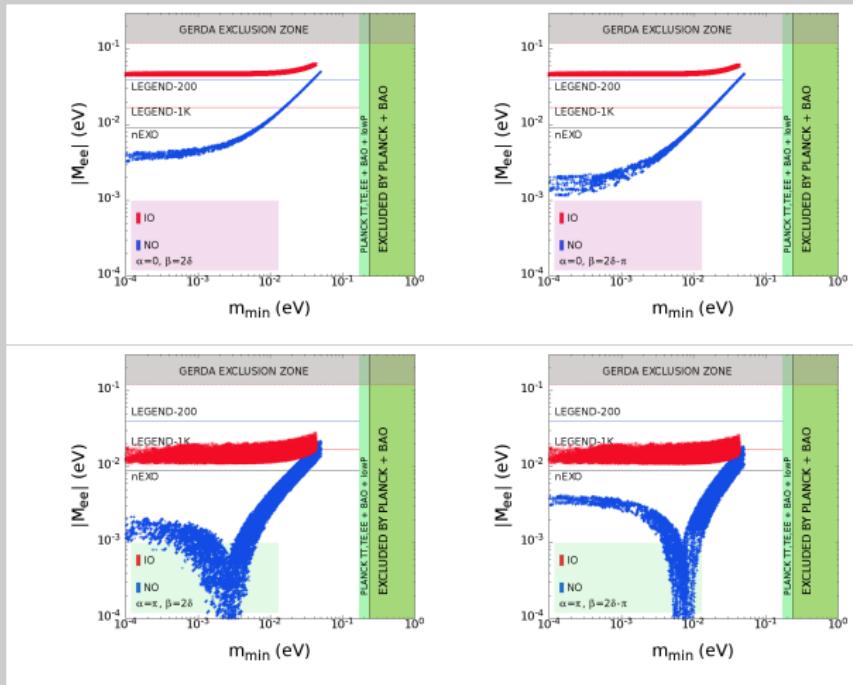
(i)  $|M_\nu^{ee}| = c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 + s_{13}^2 m_3$  for  $\alpha = 0, \beta = 2\delta$ ,

(ii)  $|M_\nu^{ee}| = c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 - s_{13}^2 m_3$  for  $\alpha = 0, \beta = 2\delta - \pi$ ,

(iii)  $|M_\nu^{ee}| = c_{12}^2 c_{13}^2 m_1 - s_{12}^2 c_{13}^2 m_2 + s_{13}^2 m_3$  for  $\alpha = \pi, \beta = 2\delta$  and

(iv)  $|M_\nu^{ee}| = c_{12}^2 c_{13}^2 m_1 - s_{12}^2 c_{13}^2 m_2 - s_{13}^2 m_3$  for  $\alpha = \pi, \beta = 2\delta - \pi$ .

# Plots of $|M_{\nu}^{ee}|$ versus the minimum neutrino mass $m_{min}$



The four plots correspond to four possible choices of  $\alpha$  and  $\beta$ .

Predicted signal below the reach of GERDA phase II but reachable by LEGEND-200, LEGEND-1K and nEXO. Failure of nEXO to see any signal would rule out our model for IO.

# 5. CP ASYMMETRY IN NEUTRINO OSCILLATIONS

Experimental CP asymmetry

$$A_{\mu e} \equiv \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} = \frac{2\sqrt{P_{31}}\sqrt{P_{21}} \sin \Delta_{32} \sin \delta}{P_{31} + P_{21} + 2\sqrt{P_{31}}\sqrt{P_{21}} \cos \Delta_{32} \cos \delta}$$

with

$$\sqrt{P_{31}} \simeq s_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} - aL)}{\Delta_{31} - aL} \Delta_{31},$$

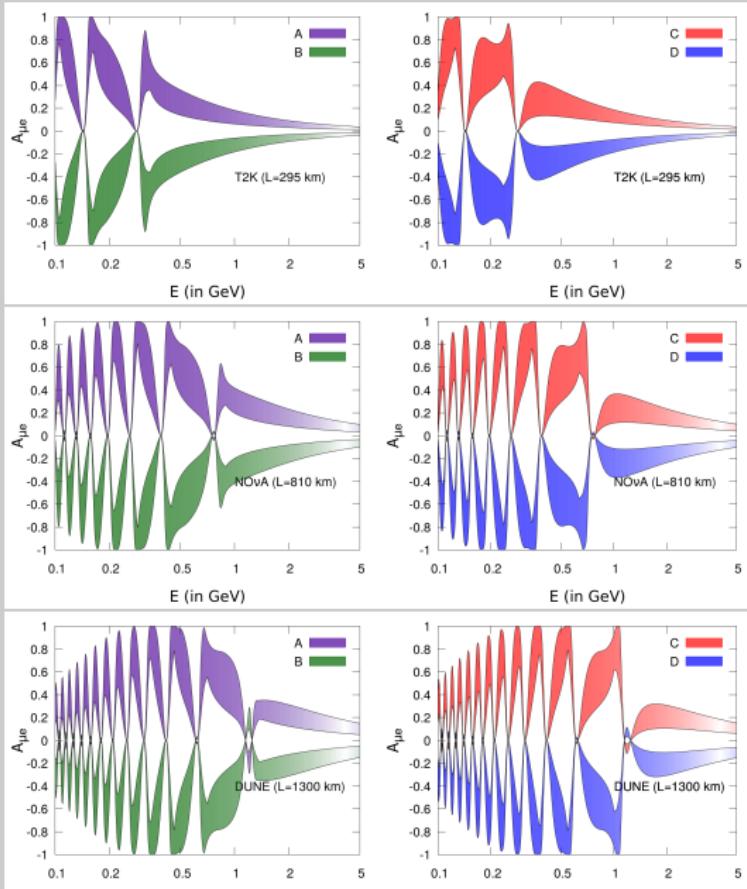
$$\sqrt{P_{21}} \simeq c_{23} \sin 2\theta_{12} \frac{\sin(aL)}{aL} \Delta_{21},$$

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{4E}, \quad a = \frac{G_F N_e}{\sqrt{2}} \simeq 3500 \text{ km}^{-1}, \quad N_e = \text{electron density in the medium}$$

$\sin \delta$  and  $\cos \delta$  can have four different combinations.

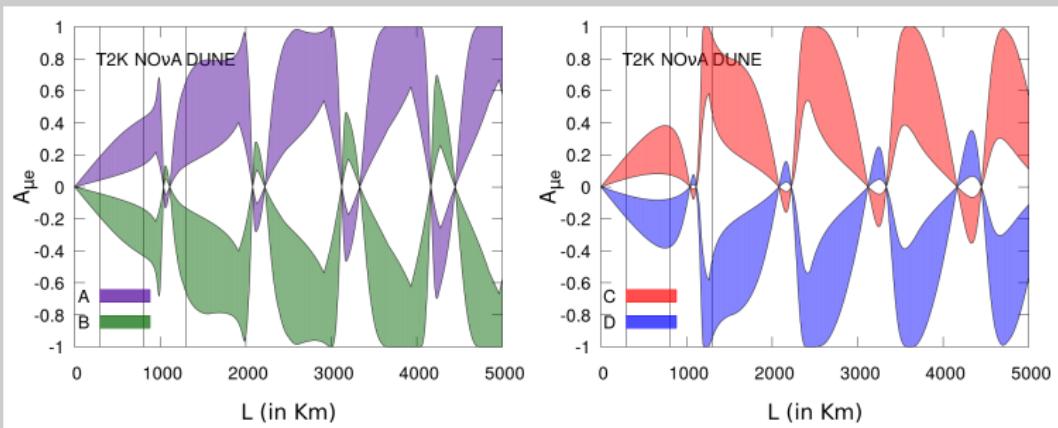
Table: Four possibilities for  $A_{\mu e}$

Possibilities	$\sin \delta$	$\cos \delta$
Case A	$+\sin \theta (\sin 2\theta_{23})^{-1}$	$+(\sin 2\theta_{23})^{-1} \sqrt{\cos^2 \theta \sin^2 2\theta_{23} - \sin^2 \theta \cos^2 2\theta_{23}}$
Case B	$-\sin \theta (\sin 2\theta_{23})^{-1}$	$+(\sin 2\theta_{23})^{-1} \sqrt{\cos^2 \theta \sin^2 2\theta_{23} - \sin^2 \theta \cos^2 2\theta_{23}}$
Case C	$+\sin \theta (\sin 2\theta_{23})^{-1}$	$-(\sin 2\theta_{23})^{-1} \sqrt{\cos^2 \theta \sin^2 2\theta_{23} - \sin^2 \theta \cos^2 2\theta_{23}}$
Case D	$-\sin \theta (\sin 2\theta_{23})^{-1}$	$-(\sin 2\theta_{23})^{-1} \sqrt{\cos^2 \theta \sin^2 2\theta_{23} - \sin^2 \theta \cos^2 2\theta_{23}}$



Plots of  $A_{\mu e}$  against beam energy  $E$  for different baselines lengths of T2K, NO $\nu$ A and DUNE respectively.

The numerical distinction between NO and IO is insignificant for the  $3\sigma$  range of  $\theta_{23}$ .



CP asymmetry parameter  $A_{\mu e}$  vs. baseline length  $L$  for cases A,B,C,D.

1. For a fixed beam energy of  $E = 1\text{GeV}$ .
2. Plots are practically indistinguishable for NO and IO.
3. The bands are due to  $3\sigma$  range  $\theta_{23}$  while the other parameters are kept at their best fit values.
4. Experimental data should distinguish among the four cases and yield information on  $\delta$ .

## 6. FLAVOR FLUX RATIOS AT NEUTRINO TELESCOPES

Source: Cosmic  $p p$  collisions (TeV-PeV)

$$\rightarrow \pi^+ \pi^- \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu \rightarrow e^+ e^- 2\nu_\mu 2\bar{\nu}_\mu \nu_e \bar{\nu}_e$$

$$\Rightarrow \{\phi_{\nu_e}^S, \phi_{\bar{\nu}_e}^S, \phi_{\nu_\mu}^S, \phi_{\bar{\nu}_\mu}^S, \phi_{\nu_\tau}^S, \phi_{\bar{\nu}_\tau}^S\} = \phi_0 \left\{ \frac{1}{6}, \frac{1}{6}, \frac{1}{3}, \frac{1}{3}, 0, 0 \right\}.$$

Source: Cosmic  $p\gamma$  collisions (GeV- $10^2$ GeV)

$$\rightarrow \pi^+ \rightarrow \mu^+ \nu_\mu \rightarrow e^+ \nu_e + \bar{\nu}_\mu.$$

$$\Rightarrow \{\phi_{\nu_e}^S, \phi_{\bar{\nu}_e}^S, \phi_{\nu_\mu}^S, \phi_{\bar{\nu}_\mu}^S, \phi_{\nu_\tau}^S, \phi_{\bar{\nu}_\tau}^S\} = \phi_0 \left\{ \frac{1}{3}, 0, \frac{1}{3}, \frac{1}{3}, 0, 0 \right\}.$$

With  $\phi_\ell^S \equiv \phi_{\nu_\ell}^S + \phi_{\bar{\nu}_\ell}^S$ ,

$$\{\phi_e^S, \phi_\mu^S, \phi_\tau^S\} = \phi_0 \left\{ \frac{1}{3}, \frac{2}{3}, 0 \right\}$$

for both sources,  $\phi_0$  = overall normalization.

Flux at source  $S \rightarrow$  flux at telescope  $T$  changed by neutrino oscillations averaged over many periods.

Effectively,  $P(\nu_m \rightarrow \nu_\ell) = P(\bar{\nu}_m \rightarrow \bar{\nu}_\ell) \simeq \sum_i |U_{ei}|^2 |U_{mi}|^2$  and

$$\phi_\ell^T = \sum_i \sum_m \phi_m^S |U_{\ell i}|^2 |U_{mi}|^2 = \frac{\phi_0}{3} \sum_i |U_{\ell i}|^2 (|U_{ei}|^2 + 2|U_{\mu i}|^2).$$

It follows from the unitarity of  $U$  that

$\phi_\ell^T = \frac{\phi_0}{3} [1 + \sum_i |U_{\ell i}|^2 (|U_{\mu i}|^2 - |U_{\tau i}|^2)]$  which vanishes for exact  $\mu\tau$  symmetry or antisymmetry, but is nonzero in general.

Neglect  $\mathcal{O}(\sin^2 \theta_{13}) \approx 0.01$  terms and define flavor flux ratios

$$R_e \equiv \phi_e (\phi_\mu + \phi_\tau)^{-1}, R_\mu \equiv \phi_\mu (\phi_e + \phi_\tau)^{-1}, R_\tau \equiv \phi_\tau (\phi_\mu + \phi_e)^{-1}.$$

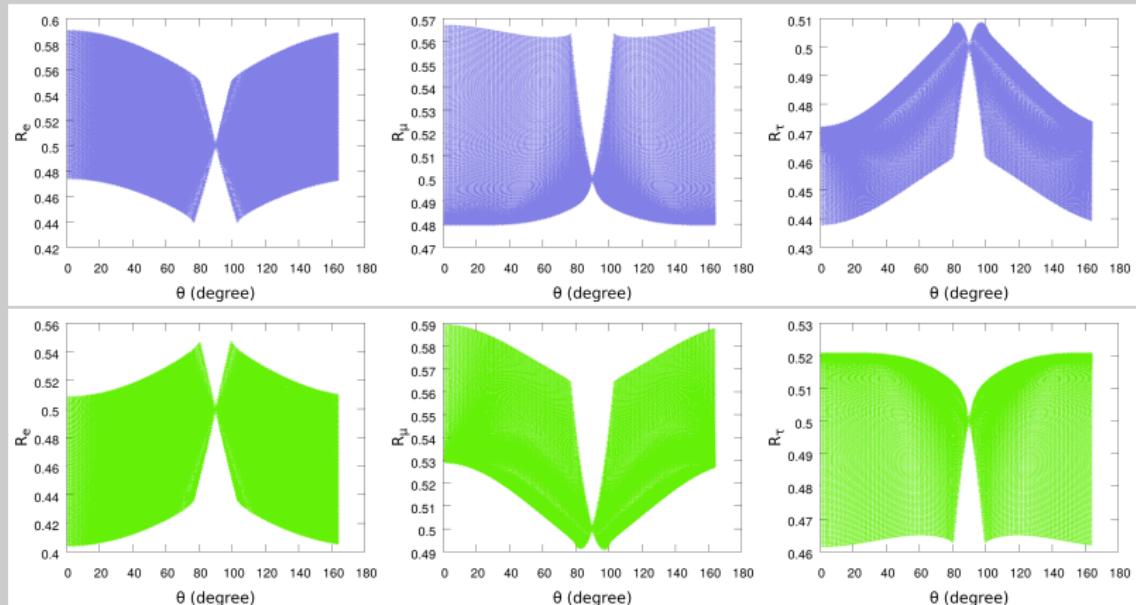
Now,

$$R_e \approx \frac{1 + \frac{1}{2} \sin^2 2\theta_{12} \cos 2\theta_{23} + \frac{1}{2} \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta}{2 - \frac{1}{2} \sin^2 2\theta_{12} \cos 2\theta_{23} - \frac{1}{2} \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta},$$

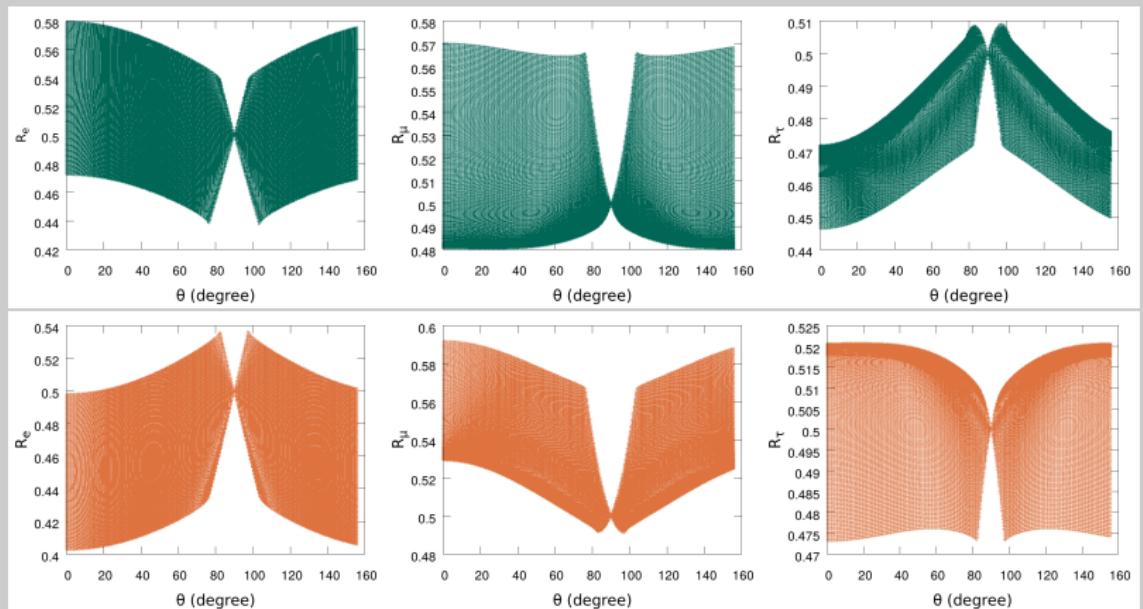
$$R_\mu \approx \frac{1 + \{c_{23}^2 (1 - \frac{1}{2} \sin^2 2\theta_{12}) - s_{23}^2\} \cos 2\theta_{23} - \frac{1}{4} \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta (4c_{23}^2 - 1)}{2 - \cos^2 2\theta_{23} + \frac{1}{2} \sin^2 2\theta_{12} \cos 2\theta_{23} c_{23}^2 + \frac{1}{4} (3 - 4s_{23}^2) \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta},$$

$$R_\tau \approx \frac{1 + \{s_{23}^2 (1 - \frac{1}{2} \sin^2 2\theta_{12}) + c_{23}^2\} \cos 2\theta_{23} - \frac{1}{4} \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta (4s_{23}^2 - 1)}{2 + \cos^2 2\theta_{23} + \frac{1}{2} \sin^2 2\theta_{12} \cos 2\theta_{23} c_{23}^2 + \frac{1}{4} (3 - 4c_{23}^2) \sin 4\theta_{12} \sin 2\theta_{23} s_{13} \cos \delta}.$$

Dependence on  $\cos \delta$  makes  $R_\ell$  double-valued except at  $\theta = \pi/4$   
 $(\cos \delta = 0$  when  $R_e = R_\mu = R_\tau = \frac{1}{2}$ ).



Flux ratios  $R_{e,\mu,\tau}$  vs.  $\theta$  for NO; range of  $\theta$ :  $12^\circ - 164^\circ$ .  
The upper (lower) panel corresponds to  $\cos \delta \geq 0 (\leq 0)$ .



Flux ratios  $R_{e,\mu,\tau}$  vs.  $\theta$  for IO; range of  $\theta$ :  $2^\circ - 156^\circ$   
The upper (lower) panel corresponds to  $\cos \delta \geq 0 (\leq 0)$ .

Continuous bands because of  $3\sigma$  variation in input parameters.  
Drastic change in  $R_e$  from  $1/2$  (as  $\theta$  moves away from  $\pi/2$ ) can be used to pinpoint  $\theta$ .

## 7. CONCLUSIONS

- CP transformed mixed  $\nu_\mu$ - $\nu_\tau$  antisymmetry in  $M_\nu$  is proposed.
- With input neutrino neutrino mixing angles and mass-squared differences ( $3\sigma$ ), ranges of values of neutrino masses for NO and IO given.
- Specific prediction on the  $\beta\beta0\nu$  process to be tested crucially by nEXO.
- Neutrino CP asymmetry  $A_{\mu e}$ , when measured, should be able to provide information on the CP phase  $\delta$ .
- Specific predictions on neutrino-antineutrino flavor flux ratios to be measured in neutrino telescopes.