

# Fermion Singlet Dark Matter in Scotogenic $B - L$ Model

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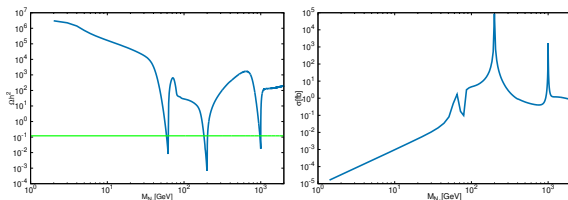
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$$M_{h_1} = 125 \text{ GeV}, M_{h_2} = 400 \text{ GeV}, M_{Z_{B-L}} = 2 \text{ TeV}, \sin \gamma = 0.1$$

- To bring more allowed parameter space for dark matter mass we need more cross-section by having extra channels.

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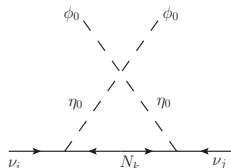
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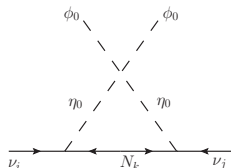
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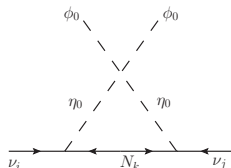
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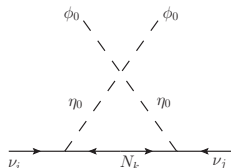
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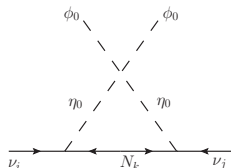
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- We can show that the high mass scale is reduced
- We can probe this in forthcoming experiments

## Fermion Singlet Dark Matter in Scotogenic $B - L$ Model

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- By combining these ideas we studied extra channel contributions on relic density
- And also studied the allowed parameter space from neutrino, direct detection and collider experiments.

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Fields	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$	$U(1)_{B-L}$	$Z_2$
$N_R$	1	1	0	-1	-
$\eta$	1	2	$\frac{1}{2}$	0	-
$\chi$	1	1	0	2	+

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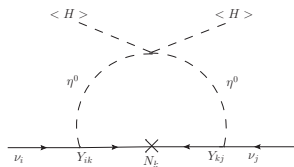
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$\chi$	1	1	0	2	+

$$\mathcal{L}_Y = \sum_{j,k=1}^3 -y_{jk} \bar{\ell}_{jL} N_{kR} \tilde{\eta} - \lambda_{jk} (\overline{N_{jR}})^c N_{kR} \chi + h.c. - V(H, \chi, \eta)$$

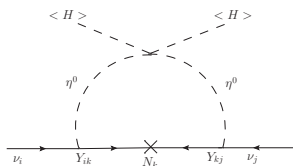
Where

$$\begin{aligned} V(H, \chi, \eta) = & -\mu_H^2 H^\dagger H + \lambda_H (H^\dagger H)^2 \\ & -\mu_\chi^2 \chi^\dagger \chi + \lambda_\chi (\chi^\dagger \chi)^2 + \mu_\eta^2 \eta^\dagger \eta + \lambda_\eta (\eta^\dagger \eta)^2 \\ & + \lambda_{H\chi} (H^\dagger H) (\chi^\dagger \chi) + \lambda_{H\eta} (H^\dagger H) (\eta^\dagger \eta) + \lambda_{\chi\eta} (\chi^\dagger \chi) (\eta^\dagger \eta) \\ & + \lambda_1 (\eta^\dagger H) (H^\dagger \eta) + \frac{\lambda_2}{2} \left[ (H^\dagger \eta)^2 + h.c. \right] \end{aligned}$$

## Radiative Neutrino Mass

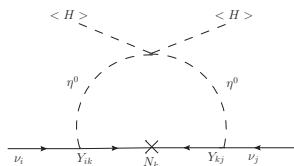


## Radiative Neutrino Mass



$$(m_\nu)_{ij} = \sum_k \frac{y_{ik} y_{kj} M_k}{32\pi^2} \left[ \frac{M_{\eta R}^2}{M_{\eta R}^2 - M_k^2} \log \left( \frac{M_{\eta R}^2}{M_k^2} \right) - \frac{M_{\eta I}^2}{M_{\eta I}^2 - M_k^2} \log \left( \frac{M_{\eta I}^2}{M_k^2} \right) \right]$$

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A typical choice of Yukawa couplings,

$$\begin{aligned} y_{e1} &= 0.00001, y_{e2} = 0.621, y_{e3} = 0.0001, \\ y_{\mu 1} &= 0.00001, y_{\mu 2} = 1.15, y_{\mu 3} = 0.485, \\ y_{\tau 1} &= 0.045, y_{\tau 2} = 0.3, y_{\tau 3} = 0.765. \\ M_{N_1} &= 99 \text{ GeV}, M_{N_2} = 10^4 \text{ GeV}, M_{N_3} = 10^3 \text{ GeV}, \\ M_{\eta R} &= 106.00000001 \text{ GeV}, M_{\eta I} = 106 \text{ GeV}. \end{aligned}$$



## Relic Density of $N_1$ in Scotogenic $B - L$ model

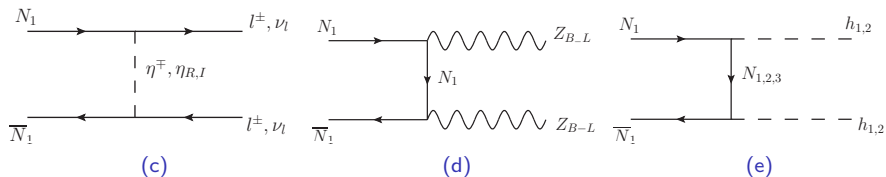
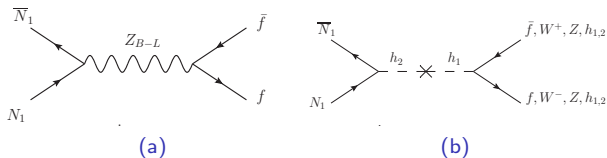


Figure: DM annihilation channels.

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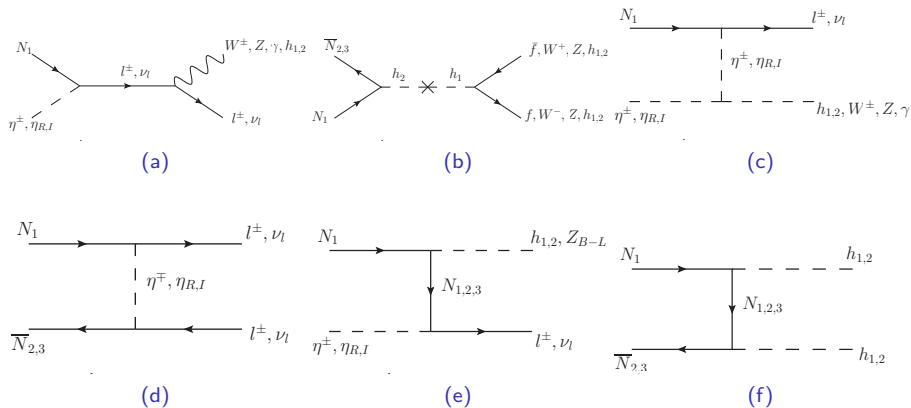
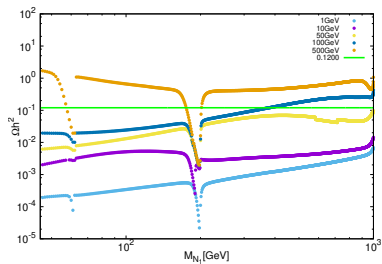


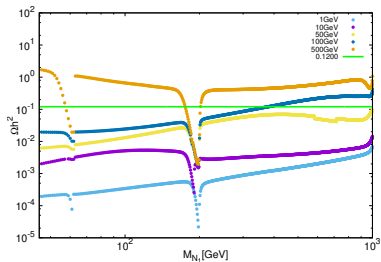
Figure: DM co-annihilation channels.

## Relic Density of $N_1$ in Scotogenic $B - L$ model



$$\begin{aligned}\delta M &= M_{\eta_R, \eta_I, \eta^\pm} - M_{N_1} \\ M_{N_2} &= 1010 \text{ GeV}, M_{N_3} = 1020 \text{ GeV}, \\ M_{h_1} &= 125 \text{ GeV}, M_{h_2} = 400 \text{ GeV}, \\ M_{Z_{B-L}} &= 2 \text{ TeV}, \sin \gamma = 0.1\end{aligned}$$

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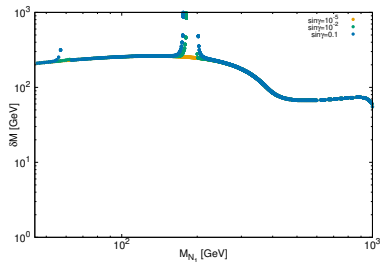


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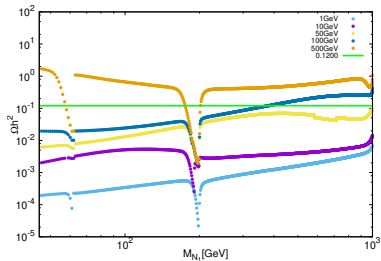
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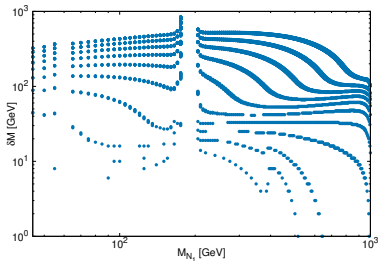
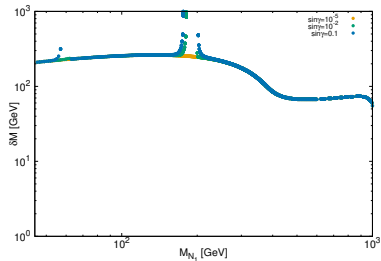


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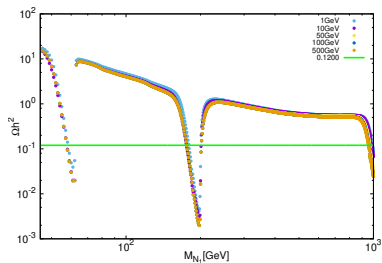
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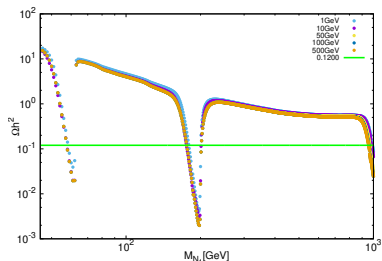


## Relic Density of $N_1$ in Scotogenic $B - L$ model



$$\begin{aligned}\delta M &= M_{N_{2,3}} - M_{N_1} = 1, 10, 50, 100, 500 \text{ GeV} \\ M_{\eta_{\pm}} &= 1030 \text{ GeV}, \\ M_{\eta_R} &= 1040 \text{ GeV}, \\ M_{\eta_I} &= 1050 \text{ GeV}, \\ M_{h_1} &= 125 \text{ GeV}, \\ M_{h_2} &= 400 \text{ GeV}, \\ M_{Z_{B-L}} &= 2 \text{ TeV}, \\ \sin \gamma &= 0.1\end{aligned}$$

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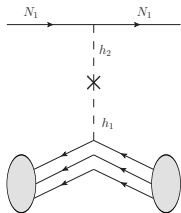
$$M_{h_2} = 400 \text{ GeV},$$

$$M_{Z_{B-L}} = 2 \text{ TeV},$$

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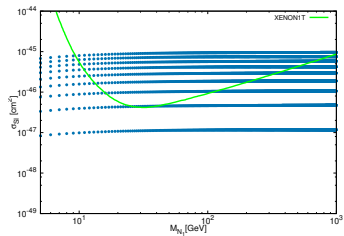
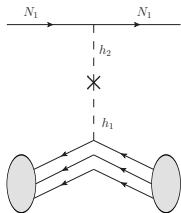
- The mass-splitting between  $\delta M = M_{\eta_R, \eta_I, \eta_{\pm}} - M_{N_1}$  is more effective than the mass-splitting between  $\delta M = M_{N_{2,3}} - M_{N_1}$ .

## Direct Detection of Dark Matter

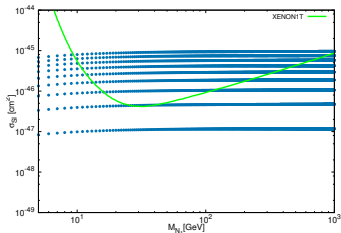
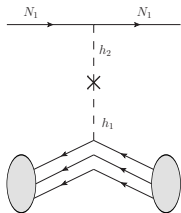




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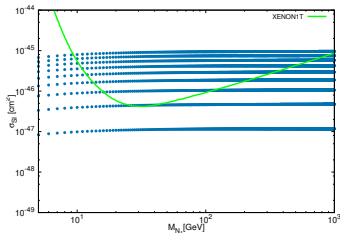
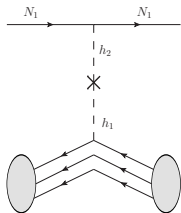


- Spin-independent elastic scattering cross-section of DM per nucleon

$$\sigma_{SI}^{h_1 h_2} = \frac{\mu_r^2}{\pi A^2} \left( \frac{\lambda_{11} \sin 2\gamma}{2\sqrt{2}} \right)^2 \left[ \frac{1}{M_{h_2^2}} - \frac{1}{M_{h_1^2}} \right]^2$$

$$\times \left[ Z \left( \frac{m_p}{v} \right) (f_{Tu}^p + f_{Td}^p + f_{Ts}^p + \frac{2}{9} f_{Tg}^p) + (A - Z) \left( \frac{m_n}{v} \right) (f_{Tu}^n + f_{Td}^n + f_{Ts}^n + \frac{2}{9} f_{Tg}^n) \right]^2$$

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$$\lambda_{11} \sin 2\gamma \geq 0.2.$$

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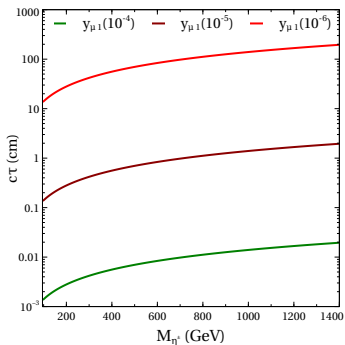
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## Collider Signature

- $\eta^\pm$  can be next to LSP.
- If the mass splitting between  $M_{\eta^\pm}$  and  $M_{N_1}$  is small.

The displaced Vertex signature:

$$\Gamma_{\eta^\pm \rightarrow N_1 \mu} = \frac{y_{\mu 1}^2 \left( m_{\eta^\pm}^2 - (m_{N_1} + m_\mu)^2 \right)}{8m_{\eta^\pm} \pi} \sqrt{1 - \left( \frac{m_{N_1} - m_\mu}{m_{\eta^\pm}} \right)^2} \sqrt{1 - \left( \frac{m_{N_1} + m_\mu}{m_{\eta^\pm}} \right)^2}$$

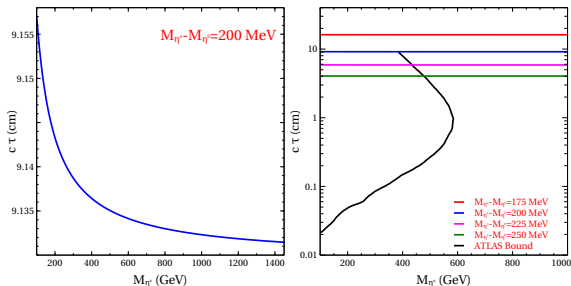


## Collider Signature

- If the mass splitting between  $M_{\eta^\pm}$  and  $M_{\eta^0}$  is of order of 100 MeV.

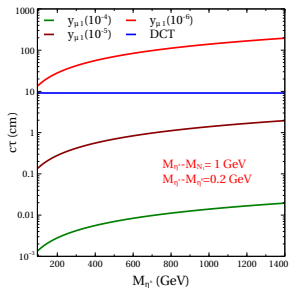
The disappearing charged track signature:

$$\Gamma_{\eta^\pm \rightarrow \eta^0 \pi} = \frac{f_\pi^2 g^4}{m_W^4} \frac{(m_{\eta^\pm}^2 - m_{\eta^0}^2)^2}{512 m_{\eta^\pm} \pi} \sqrt{1 - \left(\frac{m_{\eta^0} - m_\pi}{m_{\eta^\pm}}\right)^2} \sqrt{1 - \left(\frac{m_{\eta^0} + m_\pi}{m_{\eta^\pm}}\right)^2}$$



## Collider Signature

- The comparison between the leptonic decay mode and pionic decay mode for different values of Yukawa couplings.



## Conclusion

- By extending the SM with  $U(1)_{B-L} \times Z_2$  symmetry, we studied co-annihilation effects on relic density
- and the corresponding parameter space allowed from neutrino, direct detection experiments.
- We studied interesting collider signatures, displaced vertex signature and disappearing charged track signature.

*Thank You*



$M_{N_1}$ (GeV)	$M_{\eta^\pm}$ , $M_{\eta^{0R}}$ , $M_{\eta^{0I}}$ (GeV)	$\sigma_{pp \rightarrow \eta^+\eta^-}$ (pb)
100	105, 120, 120	0.189
200	205, 220, 220	$1.65 \times 10^{-2}$
300	305, 320, 320	$3.46 \times 10^{-3}$
400	405, 420, 420	$1.04 \times 10^{-3}$
500	505, 520, 520	$3.817 \times 10^{-4}$
600	605, 620, 620	$1.593 \times 10^{-4}$
700	705, 720, 720	$7.286 \times 10^{-5}$
800	805, 820, 820	$3.568 \times 10^{-5}$
900	905, 920, 920	$1.828 \times 10^{-5}$
1000	1005, 1020, 1020	$9.794 \times 10^{-6}$

**Table:** Production cross sections of  $\eta^+\eta^-$  from  $pp$  collisions at  $\sqrt{s} = 14$  TeV LHC. Here we have kept fixed the mass splittings as  $M_{\eta^\pm} - M_{N_1} = 5$  GeV and  $M_{\eta^{0R}} - M_{\eta^\pm} = M_{\eta^{0I}} - M_{\eta^\pm} = 15$  GeV

$M_{N_1}$ (GeV)	$M_{\eta^\pm}$ , $M_{\eta^{0R}}$ , $M_{\eta^{0I}}$ (GeV)	$\sigma_{pp \rightarrow \eta^+ \eta^-}$ (pb)
100	101, 120, 120	0.2176
200	201, 220, 220	$1.782 \times 10^{-2}$
300	301, 320, 320	$3.65 \times 10^{-3}$
400	401, 420, 420	$1.087 \times 10^{-3}$
500	501, 520, 520	$3.957 \times 10^{-4}$
600	601, 620, 620	$1.647 \times 10^{-4}$
700	701, 720, 720	$7.523 \times 10^{-5}$
800	801, 820, 820	$3.656 \times 10^{-5}$
900	901, 920, 920	$1.879 \times 10^{-5}$
1000	1001, 1020, 1020	$1.004 \times 10^{-5}$

**Table:** Production cross sections of  $\eta^+ \eta^-$  from  $pp$  collisions at  $\sqrt{s} = 14$  TeV LHC. Here we have kept fixed the mass splittings as  $M_{\eta^\pm} - M_{N_1} = 1$  GeV and  $M_{\eta^{0R}} - M_{\eta^\pm} = M_{\eta^{0I}} - M_{\eta^\pm} = 19$  GeV

$M_{N_1}$ (GeV)	$M_{\eta^\pm}, M_{\eta^{0R}}, M_{\eta^{0I}}$ (GeV)	$\sigma_{pp \rightarrow \eta^\pm \eta^0}$ (pb)
100	101.2, 101, 101.2	0.2473
200	201.2, 201, 201.2	$2.057 \times 10^{-2}$
300	301.2, 301, 301.2	$4.359 \times 10^{-3}$
400	401.2, 401, 401.2	$1.341 \times 10^{-3}$
500	501.2, 501, 501.2	$5.001 \times 10^{-4}$
600	601.2, 601, 601.2	$2.141 \times 10^{-4}$
700	701.2, 701, 701.2	$9.938 \times 10^{-5}$
800	801.2, 801, 801.2	$4.91 \times 10^{-5}$
900	901.2, 901, 901.2	$2.546 \times 10^{-5}$
1000	1001.2, 1001, 1001.2	$1.367 \times 10^{-5}$

**Table:** Production cross sections of  $\eta^\pm \eta^0$  from  $pp$  collisions at  $\sqrt{s} = 14$  TeV LHC. Here we have kept fixed the mass splittings as  $M_{\eta^{0R}} - M_{N_1} = 1$  GeV and  $M_{\eta^\pm} - M_{\eta^{0R}} = M_{\eta^{0I}} - M_{\eta^{0R}} = 200$  MeV