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EDMs and Dark Sectors

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[work with Shohei Okawa & Maxim Pospelov: 0905.05219]

Motivating Questions...

Sakharov's criteria for generating a baryon asymmetry are over 50 years old!

VIOLATION OF CP INVARIANCE, C ASYMMETRY, AND BARYON ASYMMETRY OF THE UNIVERSE

A. D. Sakharov Submitted 23 September 1966 ZhETF Pis'ma 5, No. 1, 32-35, 1 January 1967

The theory of the expanding Universe, which presupposes a superdense initial state of matter, apparently excludes the possibility of macroscopic separation of matter from antimatter; it must therefore be assumed that there are no antimatter bodies in nature, i.e., the

• Developed at a time before there was clear evidence for dark matter or neutrino mass, and now there are even more questions...



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- Baryon asymmetry? and
- Identity of Dark Matter?
- DM-genesis?
- Neutrino mass?
- Lepton asymmetry?

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 \Rightarrow Common theme - CP violation required in one sector or another

EDMs as precision probes

EDMs are powerful (amplitude-level) probes for new CP/T violation

$$H = -d\vec{E} \cdot \frac{\vec{S}}{S}$$

| Diamagnetic EDMs | Neutron EDM |
|--|--|
| $ d_{\rm Hg} < 7.4 \times 10^{-30} e {\rm cm}$ | $ d_n < 3 \times 10^{-26} e \mathrm{cm}$ |
| U Washington (Hg) [Graner et al '16] | Sussex/RAL/ILL [Baker et al. '06, Pendlebury et al '15] |
| U Michigan (Xe) [Rosenberry & Chupp '01] | |
| Argonne (Ra) [Bishof et al '16] | |
| | Diamagnetic EDMs $ d_{Hg} < 7.4 \times 10^{-30} \text{ ecm}$ U Washington (Hg) [Graner et al '16] U Michigan (Xe) [Rosenberry & Chupp '01] Argonne (Ra) [Bishof et al '16] |

(+ many experiments in development around the world, *including at TRIUMF*)

Looking back ~ 35 years



Looking back ~ 15-20 years



Multiple probes (e.g. of SUSY) with complementary sensitivity in the LEP era...

 $M_{susy} = 500 \text{ GeV}$



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Looking back ~ 0 years



Looking back ~ 0 years



EDMs in the Standard Model (CKM phase)



New physics in a dark/hidden sector

Empirical evidence for new physics (e.g. neutrino mass, dark matter) arguably points to a hidden/dark sector, but not directly to a specific mass scale



a huge parameter space a priori, so what theoretical guidance is there?

New physics in a dark/hidden sector

Given a very large 'model-space' a priori, its useful to develop an EFT expansion, assuming the hidden sector is *SM-neutral*

Substantial research effort over the past decade....

[See review talk at CAP by D. McKeen]

EFT for a (neutral) hidden sector

Many more UV-sensitive interactions at dim \geq 5

EFT for a (neutral) hidden sector

Naturally incorporates minimal models of neutrino mass, and leptogenesis

EFT for a (neutral) hidden sector - IR mediation

Universal couplings to EM/scalar currents at low energy, so hidden sector models have correlated observable effects 15

Some experimental signatures

- Precision corrections
- Rare (visible) decays
- Astrophys/cosmology

- Rare (invisible) decays/Emiss
- Anomalous NC-like scattering
- Astrophys/cosmology

Significant complementarity between precision & intensity

E.g. precision corrections (vector portal)

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EDMs from dark sectors

EDMs from dark sectors

Neutrino portal

Requires at least 2 singlet fermions

 $\mathcal{L}_{\text{hidden}} = M_N \bar{N}^c N + h.c.$

$$d_e \lesssim 10^{-33} \, e \cdot \mathrm{cm}$$

[Archambault, Czarnecki and Pospelov, 0406089; Le Dall, Pospelov and Ritz, 1505.01865; Ng and Ng, 9510306]

- Vector and scalar portal (Dark Barr-Zee mechanism)
 - $\blacktriangleright \mathcal{L}_{\text{hidden}} = Y_S \, S \bar{\psi} i \gamma_5 \psi \qquad (\psi : \text{dark fermion})$
 - electron EDM induced via "dark EDM"

$$\bar{e}\sigma^{\mu\nu}\gamma_5 eF'_{\mu\nu} \to \bar{e}\sigma^{\mu\nu}\gamma_5 e \frac{\Box F_{\mu\nu}}{m_{A'}^2}$$

$$d_e \sim 4 \times 10^{-33} \, e \cdot \mathrm{cm} \left(\frac{1 \, \mathrm{GeV}}{m_\psi}\right) \left(\frac{\epsilon}{10^{-4}}\right)^2 \left(\frac{\theta_h}{10^{-3}}\right)$$

[Le Dall, Pospelov and Ritz, 1505.01865]

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EDM suppressed by limit on 1-loop ("dark photon") correction to $(g-2)_e$

EDM via the "singlet portal"

$$d_e \sim \frac{e}{(16\pi^2)^2} \cdot \frac{\theta_h \theta_\nu^2}{m_{NP}^2} \cdot \frac{m_e}{m_{NP}^2} \cdot \lambda_N \sim 4 \cdot 10^{-29} \, e \cdot \mathrm{cm} \times \left(\frac{\theta_h \theta_\nu^2}{10^{-3}}\right) \left(\frac{100 \,\mathrm{GeV}}{m_{NP}}\right)^2$$

This is quite close to the current bound, motivating a full computation of the
relevant 2-loop diagrams...[Okawa, Pospelov & AR '19]22

 $\frac{\lambda_N}{1}$

Scaling of the induced electron EDM

Note the weak decoupling for large m_N , similar to FCNCs

$$\int d^4q d^4k f_{\text{scalar}}(q,k) \to \begin{cases} \frac{3}{4} \ln \frac{m_S^2}{m_N^2}, & m_N \ll m_S \\ \frac{1}{4} \frac{m_S^2}{m_N^2} \ln \frac{m_N^2}{m_S^2}, & m_S \ll m_N \end{cases}$$

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Sensitivity to portal couplings

Assuming a maximal dark sector CP phase, existing EDM experiments (e.g. ACME) already have significant (and complementary) sensitivity to portal couplings for masses of ~ GeV and above

Summary/Outlook

- The problem of matter-genesis now goes beyond baryogenesis, but the Sakharov criteria, e.g. CP violation (and thus precision tests such as EDMs), remain central.
- Dark sectors, motivated by the need for dark matter and the generation of neutrino mass, may be probed in a manner that is complementary to colliders and fixed target experiments using precision experiments, e.g. EDMs
 - Highlighted a sizeable 2-loop contribution to the electron EDM that is generic in the presence of the neutrino and Higgs (singlet) portals
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