

INVISIBLE WIDTHS OF HEAVY MESONS

Cody Grant

codygrant@wayne.edu



WAYNE STATE
UNIVERSITY

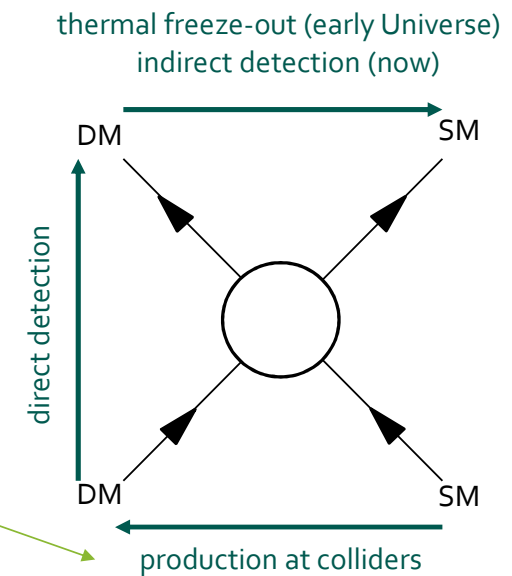
arXiv: 1809.04606 [hep-ph] (accepted to PRD)

with: Bhubanjyoti Bhattacharya (LTU)

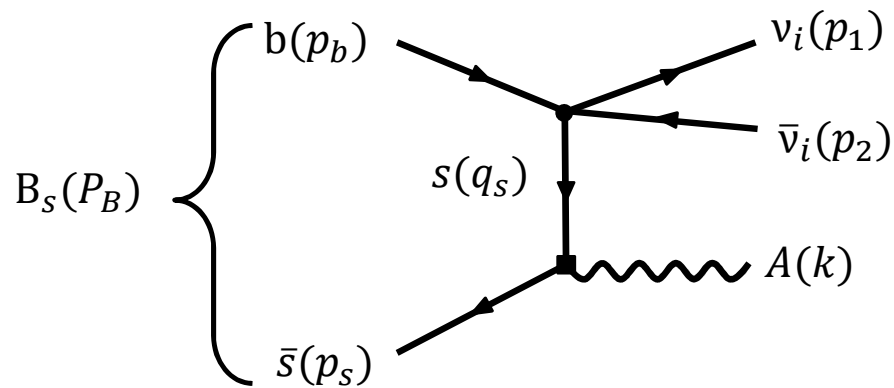
and Alexey A. Petrov (WSU)

Motivation – light Dark Matter

- Only gravitational evidence of Dark Matter
 - Can we see it elsewhere?
- If DM couples to quarks, can find:
 - Final states with other particles
 - Final states by itself: i.e. “Invisible widths”
- Light DM: $M_{\text{DM}} < 5 \text{ GeV}$
 - Can use heavy meson decays to probe for it



Motivation – invisible background



- Main background in SM for invisible width is neutrinos
 - Lowest order in G_F is $M \rightarrow \nu\bar{\nu}$

- Experimental $M \rightarrow$ Invisible:
 - $\mathcal{B}(B_d \rightarrow inv) < 1.3 \times 10^{-4}$ Belle (2012)
 - $\mathcal{B}(B_d \rightarrow inv) < 2.4 \times 10^{-5}$ BaBar (2012)
 - $\mathcal{B}(D^0 \rightarrow inv) < 9.4 \times 10^{-5}$ Belle (2017)

Two body decay: $M \rightarrow \nu \bar{\nu}$

$$\mathcal{L}_{eff} = -\frac{4G_F}{\sqrt{2}} \frac{\alpha}{2\pi \sin^2 \theta_\omega} \sum_{l=e,\mu,\tau} \sum_k \lambda_k X^l(x_k) J_{Qq}^\mu (\bar{\nu}_L^l \gamma_\mu \nu_L^l)$$

where $\begin{cases} \text{for B-decays: } J_{Qq}^\mu = \bar{q}_L \gamma^\mu b_L \\ \text{for D-decays: } J_{Qq}^\mu = \bar{u}_L \gamma^\mu c_L \end{cases}$

also for B-decays: $\lambda_k X^l(x_k) = |V_{tq}^* V_{tb}| \cdot \frac{x_t}{8} \left[\frac{x_t+2}{x_t-1} + \frac{3(x_t-2)}{(x_t-1)^2} \ln x_t \right]$

Branching ratio are tiny:

$$\mathcal{B}(B_s \rightarrow \nu \bar{\nu}) = \frac{G_F^2 \alpha^2 f_B^2 M_B^3}{16\pi^3 \sin^4 \theta_\omega \Gamma_{B_s}} |V_{tq}^* V_{tb}|^2 X^2(x_t) x_\nu^2$$

- Helicity suppression factor of $x_\nu^2 = \left(\frac{m_\nu}{M_B}\right)^2$
- Can it be used to measure neutrino mass?
 - PDG

$$\mathcal{B}(B_s \rightarrow \nu \bar{\nu}) = 3.07 \times 10^{-2} \left(\frac{m_\nu}{0.1 \text{ eV}}\right)^2$$

$$\mathcal{B}(B_d \rightarrow \nu \bar{\nu}) = 1.24 \times 10^{-25} \left(\frac{m_\nu}{0.1 \text{ eV}}\right)^2$$

$$\mathcal{B}(D^0 \rightarrow \nu \bar{\nu}) = 1.1 \times 10^{-30} \left(\frac{m_\nu}{0.1 \text{ eV}}\right)^2$$

Badin,
Petrov
(2010)



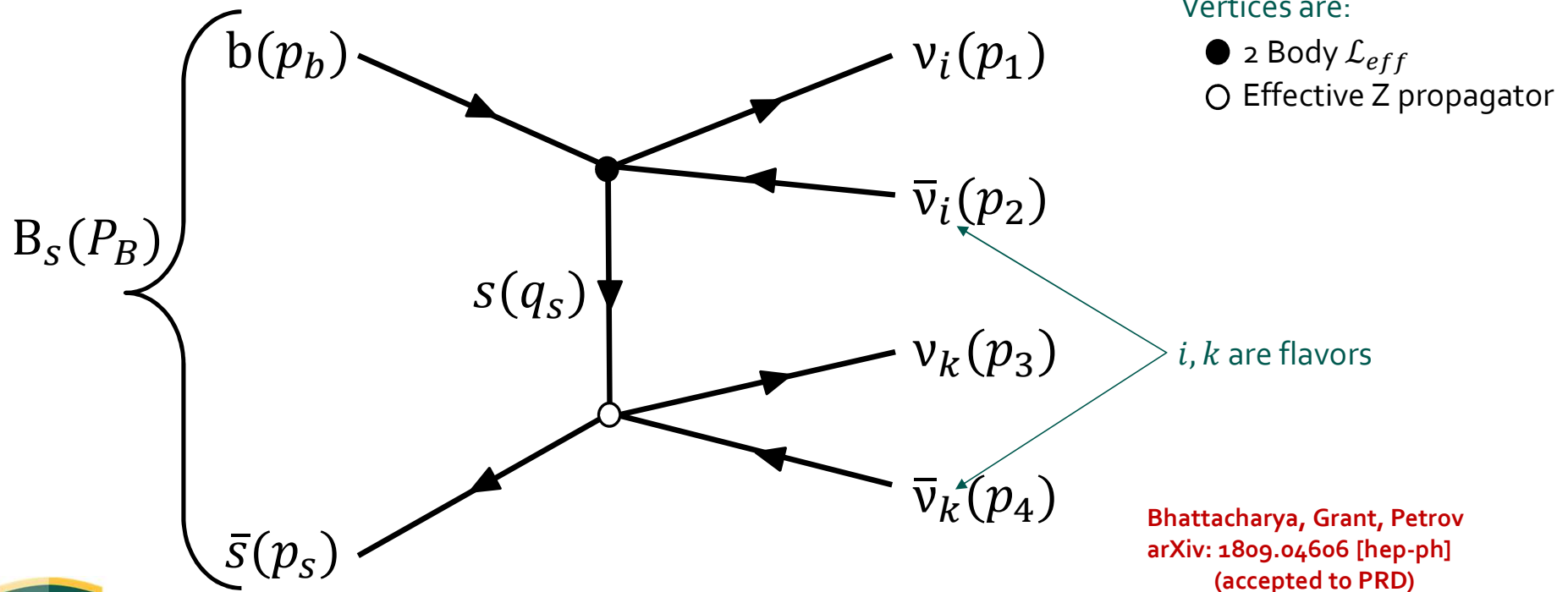
However, SM Invisible decay has more terms!!

$$\mathcal{B}(M \rightarrow \cancel{E}) = \mathcal{B}(M \rightarrow \nu\bar{\nu}) + \boxed{\mathcal{B}(M \rightarrow \nu\bar{\nu}\nu\bar{\nu})} + \dots$$

Bhattacharya, Grant, Petrov
arXiv: 1809.04606 [hep-ph]
(accepted to PRD)

$$\frac{\mathcal{B}(M \rightarrow \nu\bar{\nu}\nu\bar{\nu})}{\mathcal{B}(M \rightarrow \nu\bar{\nu})} \sim \frac{G_F^2 M_M^4}{16\pi^2 x_\nu^2} \gg 1$$

Four body decay to neutrinos



Bhattacharya, Grant, Petrov
arXiv: 1809.04606 [hep-ph]
(accepted to PRD)



Details of the calculation

Amplitude:

$$\mathcal{A}_S = -\frac{G_F^2 \alpha V_{ts}^* V_{tb} X(x_t)}{4\pi \sin^2 \theta_\omega} \sum_{i,k} L_{l_i}^\mu L_{l_k}^\nu \langle 0 | \bar{s} \Gamma^{\mu\nu} b | B_S \rangle$$

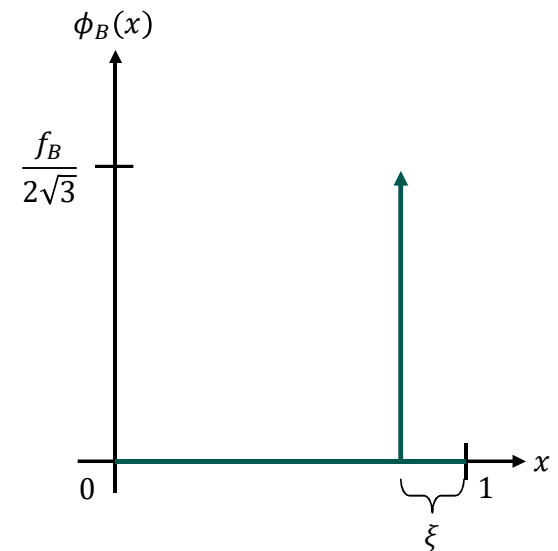
Simple quark model:

$$\langle 0 | \bar{s} \Gamma^{\mu\nu} b | B_S \rangle = \int_0^1 dx \text{Tr}[\Gamma^{\mu\nu} \psi_{B_S}] \quad \text{where } x = \frac{p_b}{P_B}$$

Bound state wave function:

$$\psi_B = \frac{I_c}{\sqrt{6}} \phi_B(x) \gamma^5 (\gamma^\mu P_{B\mu} + M_B g_B(x))$$

$$\phi_B(x) = \frac{f_B}{2\sqrt{3}} \delta(1 - x - \xi)$$



Results: Branching Ratios

$$\mathcal{B}(M \rightarrow \bar{K}) = \mathcal{B}(M \rightarrow \nu\bar{\nu}) + \mathcal{B}(M \rightarrow \nu\bar{\nu}\nu\bar{\nu}) + \dots$$

$$\begin{aligned} B_s &: 3.07 \times 10^{-24} \\ B_d &: 1.24 \times 10^{-25} \\ D^0 &: 1.1 \times 10^{-30} \end{aligned}$$

Badin, Petrov (2010)

$$\begin{aligned} B_s &: (5.48 \pm 0.89) \times 10^{-15} \\ B_d &: (1.51 \pm 0.28) \times 10^{-16} \\ D^0 &: (2.96 \pm 0.39) \times 10^{-27} \end{aligned}$$

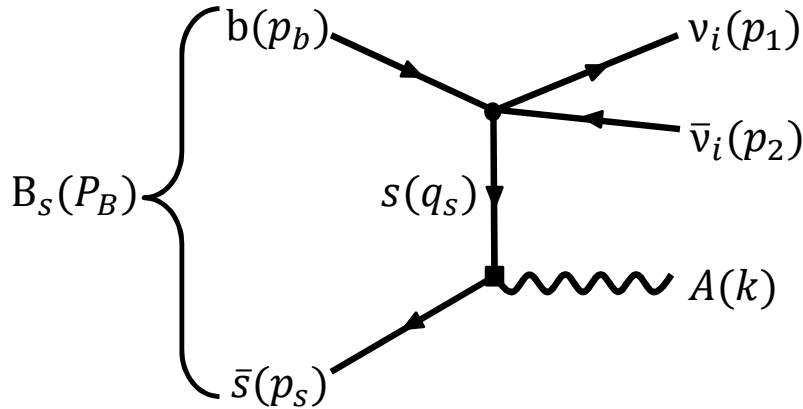
Bhattacharya, Grant, Petrov
arXiv: 1809.04606 [hep-ph]
(accepted to PRD)

**That's 9 Orders
Of Magnitude!**



Example of BSM: Dark Photon

Lowest order diagram: $\mathcal{O}(\epsilon^2)$



$$\mathcal{L} = -\frac{1}{4}W_{3\mu\nu}W^{3\mu\nu} - \frac{1}{4}B_{\mu\nu}B^{\mu\nu} + \frac{\epsilon}{2}B_{\mu\nu}V^{\mu\nu} - \frac{1}{4}V_{\mu\nu}V^{\mu\nu} + \frac{m_V^2}{2}V_\mu V^\mu$$

mixing parameter

Strongest experimental bounds:
 $\mathcal{B}(B_d \rightarrow inv) < 2.4 \times 10^{-5}$
 yields $|\epsilon| < 125$

BaBar (2012)

Strongest SM bounds:
 $\mathcal{B}(B_d \rightarrow 4\nu) = (1.51 \pm 0.28) \times 10^{-16}$
 yields $|\epsilon| \sim 3.1 \times 10^{-4}$

Still have room for DM,
 but need more
 experimental data!

Bhattacharya, Grant, Petrov
 arXiv: 1809.04606 [hep-ph]
 (accepted to PRD)



Conclusions

- SM contribution to the invisible width of heavy mesons:
 - Branching ratio is dominated by a four neutrino final state
 - 9 orders of magnitude for B-decays over $\mathcal{B}(B \rightarrow \nu\bar{\nu})$
 - 3 orders of magnitude for D-decays over $\mathcal{B}(D \rightarrow \nu\bar{\nu})$
- Can we use invisible widths to constrain neutrino mass?
 - Since $\mathcal{B}(M \rightarrow \nu\bar{\nu}\nu\bar{\nu}) \gg \mathcal{B}(M \rightarrow \nu\bar{\nu})$
and $\mathcal{B}(M \rightarrow \nu\bar{\nu}\nu\bar{\nu})$ is **NOT** proportional to m_ν
 \Rightarrow NO!
- Can we use invisible widths to search for light DM?
 - Perhaps, but need more experimental data



Questions?



5/7/2019

Cody Grant - WSU - Pheno 2019

11

References

- [1] N. Craig, A. Katz, M. Strassler and R. Sundrum, JHEP 1507, 105 (2015)
- [2] P. Meade, S. Nussinov, M. Papucci and T. Volansky, JHEP 1006, 029 (2010)
- [3] G. Aad et al. [ATLAS Collaboration], Phys. Rev. Lett. 108, 251801 (2012)
- [4] G. Aad et al. [ATLAS Collaboration], Phys. Rev. D 92, no. 1, 012010 (2015)
- [5] R. Aaij et al. [LHCb Collaboration], Eur. Phys. J. C 77, no. 12, 812 (2017)
- [6] R. Essig et al., "Working Group Report: New Light Weakly Coupled Particles," arXiv:1311.0029 [hep-ph].
- [7] C. Patrignani et al. [Particle Data Group], Chin. Phys. C 40, no. 10, 100001 (2016).
- [8] A. Badin and A. A. Petrov, Phys. Rev. D 82, 034005 (2010)
- [9] C. L. Hsu et al. [Belle Collaboration], Phys. Rev. D 86, 032002 (2012)
- [10] J. P. Lees et al. [BaBar Collaboration], Phys. Rev. D 86, 051105 (2012)
- [11] Y.-T. Lai et al. [Belle Collaboration], Phys. Rev. D 95, no. 1, 011102 (2017)
- [12] S. de Boer, arXiv:1807.05845 [hep-ph].
- [13] G. Buchalla and A. Buras, Nucl. Phys. B 400, 225 (1993)
- [14] T. Inami and C. S. Lim, Prog. Theor. Phys. 65, 297 (1981) [Erratum-ibid. 65, 1772 (1981)]
- [15] A. Goobar, S. Hannestad, E. Mortsell and H. Tu, JCAP 0606, 019 (2006)
- [16] C. D. Lu and D. X. Zhang, Phys. Lett. B 381, 348 (1996)
- [17] Y. G. Aditya, K. J. Healey and A. A. Petrov, Phys. Lett. B 710, 118 (2012)
- [18] W. Altmannshofer, P. Ball, A. Bharucha, A. J. Buras, D. M. Straub and M. Wick, JHEP 0901, 019 (2009)
- [19] B. Holdom, Phys. Lett. 166B, 196 (1986).
- [20] M. Baumgart, C. Cheung, J. T. Ruderman, L. T. Wang and I. Yavin, JHEP 0904, 014 (2009)
- [21] T. M. Aliev, A. Ozpineci and M. Savci, Phys. Lett. B 393, 143 (1997) [22] C. D. Lu and D. X. Zhang, Phys. Lett. B 381, 348 (1996)