

Scale invariance, Stueckelberg breaking of Weyl gravity and inflation

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- “New physics” beyond SM: new symmetry?

- SUSY @ TeV: - nice [in] theory... protected a hierarchy of scales. Not found @ LHC.

- scale invariance (SI); - discrete

- global SI. e.g.: SM with higgs $m_\phi = 0$ has scale invariance. [Bardeen 1995]
- local SI
- gauged local SI = Weyl gauge symmetry \rightarrow Weyl gravity [Weyl 1919]
- quantum scale invariance (QSI): global, local. [Englert, Gastmans, Truffin 1976]

(1). **Global** scale invariance $x'_\mu = \rho x_\mu$; $\phi'(\rho x) = \rho^{-1} \phi(x)$, forbids $\int d^4x m^2 \phi^2$

- no dim-ful coupling; all scales (M_{Planck}, \dots) from vev's; e.g. SM with $m_h = 0$

• I show: QSI in flat spacetime protects a hierarchy of fields vev's $\phi \ll \sigma$.

(2). **Local** scale invariance: $g'_{\mu\nu} = \Omega(x)^2 g_{\mu\nu}$, $\phi' = \frac{1}{\Omega(x)} \phi$

(3). **Gauged** local scale invariance $g'_{\mu\nu} = \Omega(x)^2 g_{\mu\nu}$, $\phi' = \frac{1}{\Omega(x)} \phi$, $\omega'_\mu = \omega_\mu - \partial_\mu \ln \Omega(x)$

- also called Weyl gauge symmetry. Weyl gravity/conformal geometry (1918).

- Non-metricity $\nabla_\mu g_{\alpha\beta} = \omega_\mu g_{\alpha\beta}$. Einstein critique due to ω_μ (massless)

- Not a problem if ω_μ massive/decouples ($\omega_\mu = 0$) at very high scale.

• I show: Weyl gravity spontaneously broken to Einstein-Proca+cc; $m_\omega \sim M_{\text{Planck}}$

M_{Planck} =phase transition emergent scale. Weyl inflation: $0.00257 < r < 0.00303$.

(1). Global Scale Invariance – at Quantum level:

[Englert 1976, Itzykson/Zuber, Shaposhnikov 2008]

DR: $d=4 - 2\epsilon$: $L = \frac{1}{2}(\partial_\mu\phi)^2 - \lambda\mu^{2\epsilon}\phi^4$, explicit ~~SI~~ by UV regulators, to avoid.

replace $\mu \rightarrow$ field σ , spontaneous $\langle\sigma\rangle \neq 0 \Rightarrow$ Goldstone (dilaton): $\sigma = \langle\sigma\rangle e^\tau$, $\tau \rightarrow \tau - \ln\rho$

[nothing new, recall M_{string} moduli dep]

• Action: $d=4$, $S = \int d^4x \left[\underbrace{(\partial_\mu\phi_j)^2 - V(\phi_j)}_{\text{visible}} \right] + \int d^4y \underbrace{L_h(\sigma, \partial\sigma)}_{\text{hidden}}$ spectrum extended by σ !

- each sector SI (shift symmetry) \rightarrow enhanced shift symmetry: $S_h \times S_v \Rightarrow \lambda_m \phi^2 \sigma^2$: λ_m naturally small

[Fubini 1976; Volkas, Kobakhidze, Foot 2013]

$$\mu^{2\epsilon}\phi^4 \rightarrow \tilde{V} \sim (z\sigma)^{2\epsilon}\phi^4 = (z\langle\sigma\rangle)^{2\epsilon} \left[1 + 2\epsilon \left(\frac{\tilde{\sigma}}{\langle\sigma\rangle} - \frac{\tilde{\sigma}^2}{2\langle\sigma\rangle^2} + \dots \right) + \mathcal{O}(\epsilon^2) \right] \phi^4, \quad \sigma = \langle\sigma\rangle + \tilde{\sigma}, \quad (z : \text{dim-less})$$

\Rightarrow scale inv reg=DR+dilaton with ∞ -many ϵ -couplings. $\langle\sigma\rangle$: 'new physics'. $\langle\phi\rangle \ll \langle\sigma\rangle$ quantum stable?

[$\sigma\sigma \rightarrow \sigma\sigma$ at 3 loops! $(\partial_\mu \ln \sigma)^4$]

• One-loop SI potential:

$$L = \frac{1}{2}(\partial_\mu\phi)^2 + \frac{1}{2}(\partial_\mu\sigma)^2 - \overbrace{\frac{1}{4!}\lambda\phi^4}^{=V(\phi)}, \quad V \rightarrow \tilde{V} = z^{2\epsilon}\sigma^{2\epsilon}V(\phi).$$

$$V_1 = -\frac{i}{2}\int\frac{d^d p}{(2\pi)^d}\text{tr}\ln[p^2 - \tilde{V}_{\alpha\beta} + i\epsilon] = \sum_{s=\phi,\sigma}\frac{\tilde{M}_s^4}{4\kappa}\left[\frac{-1}{\epsilon} + \ln\frac{\tilde{M}_s^2}{c_0}\right]\mu(\sigma)^{2\epsilon},$$

$$\partial_{\sigma\dots}^2\tilde{V} \sim \partial_{\sigma\dots}^2V\dots + \epsilon(\dots); \quad \tilde{M}_\phi^4 = M_\phi^4 + \epsilon\dots, \quad \tilde{M}_\sigma^4 \sim \epsilon^2, \quad \kappa = (4\pi)^2$$

$$U = V + \frac{\lambda^2\phi^4}{16\kappa}\left[\overline{\ln}\frac{\lambda\phi^2}{(z\sigma)^2} - \frac{1}{2}\right] = V + \underbrace{\frac{\lambda^2\phi^4}{16\kappa}\left[\overline{\ln}\frac{\lambda\phi^2}{(z\langle\sigma\rangle)^2} - \frac{1}{2}\right]}_{CW} + \underbrace{\frac{\lambda^2\phi^4}{16\kappa}\left[\frac{-\tilde{\sigma}}{\langle\sigma\rangle} + \frac{\tilde{\sigma}^2}{2\langle\sigma\rangle^2} + \dots\right]}_{\rightarrow 0; \text{ small yet maintains SI}}$$

$$\sigma = \langle\sigma\rangle + \tilde{\sigma}, \quad \tilde{\sigma}: \text{fluctuations}$$

$\Rightarrow U$ scale invariant due to dilaton ($\ln\sigma$).

No new poles, same beta function $\beta_\lambda^{(1)}$; $\lambda^B = \lambda Z_\lambda Z_\phi^{-2}$, $d\lambda^B/d(\ln z) = 0$, $\beta_\lambda^{(1)} = d\lambda/d(\ln z) = 3\lambda^2/\kappa$.

QSI with $\beta_\lambda \neq 0$, dilation current $\partial_\mu D^\mu = 0$. $\beta_\lambda = 0$ is sufficient, not necessary for QSI. CS: $dU/d(\ln z) = 0$

• **Two-loop SI potential:** using: $\tilde{V}(\phi, \sigma) \sim z^{2\epsilon} \sigma^{2\epsilon} V(\phi)$ [background field method]

$$\Rightarrow \tilde{V}(\phi + \delta_\phi, \sigma + \delta_\sigma) = \tilde{V}(\phi, \sigma) + \tilde{V}_\alpha \delta_\alpha + \frac{1}{2} \tilde{V}_{\alpha\beta} \delta_\alpha \delta_\beta + \frac{1}{3!} \tilde{V}_{\alpha\beta\gamma} \delta_\alpha \delta_\beta \delta_\gamma + \frac{1}{4!} \tilde{V}_{\alpha\beta\gamma\rho} \delta_\alpha \delta_\beta \delta_\gamma \delta_\rho + \dots \quad \alpha, \beta = \phi, \sigma.$$

$$\tilde{V}_{\alpha\beta\dots} = \partial_\alpha \partial_\beta \dots \tilde{V}$$

$$V_2 = \frac{i}{12} \text{[bubble]} + \frac{i}{8} \text{[figure-eight]} + \frac{i}{2} \text{[triangle with cross]} = \frac{i}{12} \tilde{V}_{\alpha\beta\gamma} \tilde{V}_{\alpha'\beta'\gamma'} \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} (\tilde{D}_p)_{\alpha\alpha'} (\tilde{D}_q)_{\beta\beta'} (\tilde{D}_{p+q})_{\gamma\gamma'} + \dots$$

$$= (z\sigma)^{2\epsilon} \frac{\lambda^3 \phi^4}{32\kappa^2} \left\{ -\frac{3}{\epsilon^2} + \frac{2}{\epsilon} + \mathcal{O}(\epsilon^0) \right\}; \quad (\tilde{D}_p)_{\alpha\beta} = (D_p)_{\alpha\beta} + \epsilon (\dots)_{\alpha\beta} + \epsilon^2 (\dots)_{\alpha\beta}$$

same poles, ϵ -shifts to propagators, vertices:

$$\tilde{V}_{\alpha\beta\gamma\dots} = V_{\alpha\beta\gamma\dots} + \epsilon (\dots)_{\alpha\beta\gamma\dots} + \epsilon^2 (\dots)_{\alpha\beta\gamma\dots}$$

[Z. Lalak, P. Olszewski, DG]

[1712.06024]

Two-loop corrected U

$$U = \frac{\lambda}{4!} \phi^4 \left\{ \underbrace{1}_V + \frac{3\lambda}{2\kappa} \left(\overline{\ln} \frac{V_{\phi\phi}}{(z\sigma)^2} - \frac{1}{2} \right) + \frac{3\lambda^2}{4\kappa^2} \left(4 + A_0 - 4 \overline{\ln} \frac{V_{\phi\phi}}{(z\sigma)^2} + 3 \overline{\ln}^2 \frac{V_{\phi\phi}}{(z\sigma)^2} \right) + \frac{5\lambda^2 \phi^2}{\kappa^2 \sigma^2} + \frac{7\lambda^2 \phi^4}{24\kappa^2 \sigma^4} \right\},$$

V $V^{(1)}$ $V^{(2)}$ **new $V^{(2,n)}$ finite z -independent**

- **Two-loop:** Taylor expand about $\sigma = \langle \sigma \rangle + \tilde{\sigma}$:

$$U = \frac{\lambda}{4!} \phi^4 \left\{ 1 + \frac{3\lambda}{2\kappa} \left(\overline{\ln} \frac{V_{\phi\phi}}{(z\langle\sigma\rangle)^2} - \frac{1}{2} \right) + \frac{3\lambda^2}{4\kappa^2} \left(4 + A_0 - 4\overline{\ln} \frac{V_{\phi\phi}}{(z\langle\sigma\rangle)^2} + 3\overline{\ln}^2 \frac{V_{\phi\phi}}{(z\langle\sigma\rangle)^2} \right) + \mathcal{O}\left(\frac{1}{\langle\sigma\rangle}\right) \right\}$$

- This is the **“usual” CW result** with $\mu = \langle \sigma \rangle$, broken SI, **no dilaton present**. [Cheng, I. Jack, T. Jones, S. Martin]
- new terms comparable/larger than standard two-loop terms

$$\frac{\phi^n}{\sigma^n} \sim 1., \quad n = 1, 2; \quad \frac{\phi}{\sigma} = \frac{\phi}{\langle\sigma\rangle} \left(1 - \frac{\tilde{\sigma}}{\langle\sigma\rangle} + \frac{\tilde{\sigma}^2}{\langle\sigma\rangle^2} + \dots \right), \quad \rightarrow 0, \text{ if } \phi \ll \sigma$$

- ⇒ **Non-polynomial terms:**
- SI, vanish if $\phi \ll \sigma$, only $\log \sigma$ left;
 - **no** $\lambda^n \phi^2 \sigma^2 = \lambda^n \langle \sigma \rangle^2 \phi^2 + \dots \Rightarrow$ **no** tuning of higgs self-coupling λ .
 - finite c-terms, cannot be seen in a scheme that breaks this symmetry.
 - not forbidden by symmetry \rightarrow ops mandatory, quantum generated; non-renorm
 - similar operators at three-loop SI potential (computed), as **counterterms**.

• **SM+dilaton: one-loop SI potential**

[Shaposhnikov et al 2008; DG, Z. Lalak, P. Olszewski]

$$\begin{aligned}
 V &= \frac{\lambda_\phi}{3!} (H^\dagger H)^2 + \frac{\lambda_m}{2} (H^\dagger H) \sigma^2 + \frac{\lambda_\sigma}{4!} \sigma^4 + \frac{4\lambda_6}{3} \frac{(H^\dagger H)^3}{\sigma^2} + \dots \\
 &= \frac{\lambda_\phi}{4!} \phi^4 + \frac{\lambda_m}{4} \phi^2 \sigma^2 + \frac{\lambda_\sigma}{4!} \lambda_\sigma \sigma^4 + \frac{\lambda_6}{6} \frac{\phi^6}{\sigma^2} \dots; (\lambda_m < 0), \\
 &\rightarrow \frac{\lambda_\phi}{4} \left(\phi^2 + \frac{3\lambda_m}{\lambda_\phi} \sigma^2 \right)^2 + \text{loops}.
 \end{aligned}$$

1) $9\lambda_m^2 = \lambda_\phi \lambda_\sigma + \text{loops}$
2) $\frac{\langle \phi \rangle^2}{\langle \sigma \rangle^2} = \frac{-3\lambda_m}{\lambda_\phi} (1 + \text{loops})$

min $\langle \sigma \rangle \neq 0$:

~~SI~~: tuning (i) or $V_{\min} = 0 \Rightarrow$ EWSB; $m_{\tilde{\phi}} \approx \lambda_\phi \langle \phi \rangle^2 = (-3)\lambda_m \langle \sigma \rangle^2 \ll \langle \sigma \rangle^2$ if $|\lambda_m| \ll \lambda_\phi$ one classical tuning

$$V^{(1)} = \text{old CW with } \mu \rightarrow z\sigma$$

$$\begin{aligned}
 V^{(1,n)} &\equiv \frac{1}{48\kappa} \left[(-16\lambda_m \lambda_\phi - 18\lambda_m^2 + \lambda_\phi \lambda_\sigma) \phi^4 - \lambda_m (48\lambda_m + 25\lambda_\sigma) \phi^2 \sigma^2 - 7\lambda_\sigma^2 \sigma^4 \right. \\
 &\quad \left. + (\lambda_\phi \lambda_m + 6\lambda_6 \lambda_\sigma) \frac{\phi^6}{\sigma^2} + 8(4\lambda_\phi - 2\lambda_m) \lambda_6 \frac{\phi^8}{\sigma^4} + (192\lambda_6 + 2\lambda_\phi) \lambda_6 \frac{\phi^{10}}{\sigma^6} + 40\lambda_6^2 \frac{\phi^{12}}{\sigma^8} \right], \quad \text{large } \sigma : S_v \times S_h
 \end{aligned}$$

$\rightarrow 0$ if $\phi \ll \sigma$ and $\lambda_m \rightarrow 0$.

no $\lambda_\phi^2 \phi^2 \sigma^2 \sim \lambda_\phi \phi^2 \langle \sigma \rangle^2$

One-loop corrected potential $U = V + V^{(1)} + V^{(1,n)}$ gives:

$$\Delta m_{\tilde{\phi}}^2 = \frac{-\lambda_m \langle \sigma \rangle^2}{\lambda_{\phi} 16\kappa} \left\{ 27 \left[g^4 \left(\ln \frac{g^2}{4} + \frac{1}{3} \right) + 2g_2^4 \left(\ln \frac{g_2^2}{4} + \frac{1}{3} \right) - 16h_t^4 \left(\ln \frac{h_t^2}{2} - \frac{1}{3} \right) \right] + 4\lambda_{\phi}^2 \left[5 \ln \frac{\lambda_{\phi}^2}{12} - 8 + \ln 27 \right] \right\} = \lambda_m \lambda_{\phi} \langle \sigma \rangle^2 + \dots \ll \langle \sigma \rangle^2, \text{ but no } \lambda_{\phi}^2 \langle \sigma \rangle^2.$$

- $\beta_{\lambda_m} \propto \lambda_m$, so λ_m stays ultraweak, natural once chosen classically small.

- result is **not** a DR artefact, since all scales from fields' vev's. More scalars?

\Rightarrow (classical) hierarchy $\langle \phi \rangle \ll \langle \sigma \rangle$ protected by quantum scale inv $S_v \times S_h$, (all orders, spont ~~S~~)

• **Curved spacetime** - Einstein gravity breaks $S_v \times S_h$ by $(\xi_1 \phi^2 + \xi_2 \sigma^2) R \Rightarrow \beta_{\lambda_m} = \xi_1(\dots) + \xi_2(\dots) + \lambda_m(\dots)$

- hierarchy of vev's (Higgs vs dilaton)? maybe preserved, fixed point for ξ_1, ξ_2 .

- what is the fate of the dilaton? can it decouple?

(2). Local scale invariance:

[t'Hooft 1104.4543; 1410.6675, IJMP 2016]

invariance under $\hat{g}_{\mu\nu} = \Omega(x)^2 g_{\mu\nu}, \quad \hat{\sigma} = \frac{\sigma}{\Omega(x)} \quad (a)$

want to generate spontaneously: $L_{EH} = -\frac{1}{2} \sqrt{g} M_p^2 R \Rightarrow L_1 = -\sqrt{g} \frac{1}{2} \left[\frac{1}{6} \sigma^2 R + g^{\mu\nu} \partial_\mu \sigma \partial_\nu \sigma \right] \quad (\text{inv.})$

a) - has a ghost (σ)!

b) - Fake conformal symmetry? [Jackiw, Pi 2015],

c) - Going to Einstein frame: d.o.f. number not conserved! \Rightarrow something is missing.... if $\langle \sigma \rangle$ somehow fixed, is this really spontaneous breaking? does not look no.

To avoid!

Next:

\Rightarrow solution to (a),(b),(c): gauged local scale invariance: Weyl gravity: dilaton σ eaten by Weyl “photon”;
hence, no ghost.

(3). Gauged local scale invariance - Weyl quadratic gravity:

invariance under $\hat{g}_{\mu\nu}(x) = \Omega(x)^2 g_{\mu\nu}(x), \quad \hat{\sigma}(x) = \frac{\sigma(x)}{\Omega(x)}, \quad \hat{\omega}_\mu(x) = \omega_\mu(x) - \frac{1}{q} \partial_\mu \ln \Omega(x)^2 \quad (b)$

Weyl geometry $\tilde{\nabla}_\mu g_{\alpha\beta} = -q \omega_\mu g_{\alpha\beta};$ with $\tilde{\Gamma}_{\mu\nu}^\rho = \Gamma_{\mu\nu}^\rho + (q/2) [\delta_\mu^\rho \omega_\nu + \delta_\nu^\rho \omega_\mu - g_{\mu\nu} \omega^\rho]$ inv of (b)
 $\Gamma_{\mu\nu}^\rho = \text{Levi-Civita}..$

$$\tilde{R} = R - 3q D_\mu \omega^\mu - \frac{3}{2} q^2 \omega^\mu \omega_\mu. \Rightarrow \hat{R} = \frac{\tilde{R}}{\Omega^2}, \quad \tilde{D}_\mu \sigma = (\partial_\mu - q/2 \omega_\mu) \sigma \Rightarrow \hat{D}_\mu \hat{\sigma} = \frac{\tilde{D}_\mu \sigma}{\Omega}$$

\Rightarrow if $\omega_\mu \rightarrow 0$: $\tilde{\Gamma}_{\mu\nu}^\rho \rightarrow \Gamma_{\mu\nu}^\rho$, **Weyl geometry** \rightarrow Riemann geometry; $\tilde{R} \rightarrow R$, Weyl tensor $\tilde{C}_{\mu\nu\rho\sigma} \rightarrow C_{\mu\nu\rho\sigma}$

\Rightarrow invariants of (b): $\sqrt{g} \tilde{R}^2, \sqrt{g} \sigma^2 \tilde{R}, \sqrt{g} (\tilde{D}_\mu \sigma)^2, \sqrt{g} F_{\mu\nu}^2, \sqrt{g} \tilde{C}_{\mu\nu\alpha\beta}^2$ where $F_{\mu\nu} = \partial_\mu \omega_\nu - \partial_\nu \omega_\mu$

no other invariants like higher dimensional ops (no scale to suppress them!) \tilde{X} denotes a quantity in Weyl geometry

- Weyl gravity = Einstein gravity + c.c. + massive ω_μ

[D.G. arXiv:1812.08613, 1904.06596]

$$L_0 = \sqrt{g} \left\{ \frac{\xi_0}{4!} \tilde{R}^2 - \frac{1}{4} F_{\mu\nu}^2 \right\}, \quad \xi_0 > 0, \quad \text{original Weyl action (1918)}$$

$$= \sqrt{g} \left\{ \frac{\xi_0}{4!} \left(-2\sigma^2 \tilde{R} - \sigma^4 \right) - \frac{1}{4} F_{\mu\nu}^2 \right\} \quad \text{eom: } \sigma^2 = -\tilde{R}; \quad \text{dilaton: } \ln \sigma \rightarrow \ln \sigma - \ln \Omega$$

+ Gauss-Bonnet + Weyl-tensor-squared.

- Weyl gauge transf of “gauge fixing”:

$$\Omega = \frac{\xi_0 \sigma^2}{6M_p^2} \Rightarrow \hat{g}_{\mu\nu} = \Omega^2 g_{\mu\nu}, \quad \hat{\sigma}^2 = \frac{6M_p}{\xi_0}, \quad \hat{\omega}_\mu = \omega_\mu - \frac{1}{q} \partial_\mu \ln \sigma^2; \quad \text{use } \tilde{R} = R - 3q D_\mu \omega^\mu - \frac{3}{2} q^2 \omega_\mu \omega^\mu$$

$$L_0 = \sqrt{\hat{g}} \left\{ -\frac{1}{2} M_p^2 \hat{R} - \frac{3M_p^4}{2\xi_0} + \frac{3}{4} q^2 M_p^2 \hat{\omega}_\mu \hat{\omega}^\mu - \frac{1}{4} \hat{F}_{\mu\nu}^2 \right\}.$$

- \Rightarrow Einstein-Proca action for massive ω_μ which “absorbed” the dilaton $\ln \sigma$ field! mass of $\omega_\mu \propto q M_p$.
- \Rightarrow Stueckelberg mechanism, massless $\sigma + \omega_\mu \rightarrow$ massive ω_μ . no ghost! # dof=3 conserved!
- \Rightarrow massive ω_μ decouples; M_p : emergent scale of Weyl gauge symmetry breaking. \Rightarrow Weyl gravity viable
- $\Rightarrow M_p \sim \langle \sigma \rangle$ dynamically in FRW universe [arxiv:1801.07676 by G. Ross, C. Hill, P. Ferreira] see talk by Graham Ross

- Geometric Stueckelberg* mechanism: A Weyl gauge invariant action:

$$L_1 = \sqrt{g} \left[-\frac{1}{4} F_{\mu\nu}^2 + \frac{1}{2} (\tilde{D}_\mu \sigma)^2 + \dots \right], \quad F_{\mu\nu} = \partial_\mu \omega_\nu - \partial_\nu \omega_\mu$$

$$\tilde{D}_\mu \sigma = (\partial_\mu - q/2 \omega_\mu) \sigma = (-q/2) \sigma \left[\omega_\mu - (1/q) \partial_\mu \ln \sigma^2 \right].$$

“gauge fixing” transformation: $\Omega^2 = \frac{\sigma^2}{M^2}, \quad \hat{g}_{\mu\nu} = \Omega^2 g_{\mu\nu} \quad \hat{\sigma} = \frac{1}{\Omega} \sigma = M, \quad \hat{\omega}_\mu = \omega_\mu - \frac{1}{q} \partial_\mu \ln \Omega^2$

→ Proca action $L_1 = \sqrt{\hat{g}} \left[-\frac{1}{4} \hat{F}_{\mu\nu}^2 + \frac{q^2}{8} M^2 \hat{\omega}_\mu \hat{\omega}^\mu + \dots \right].$

- spontaneous breaking; ω_μ has eaten the scalar $\ln \sigma$ (note $\ln \sigma \rightarrow \ln \sigma - \ln \Omega$)
- number of d.o.f. is conserved (=3).

* Baron Ernst Carl Gerlach Stueckelberg von Breidenbach zu Breidenstein und Melsbach

- Weyl gravity with matter

ϕ - scalar field, higgs-like, etc

$$\begin{aligned}
 L_2 &= \sqrt{g} \left\{ \frac{\xi_0}{4!} \tilde{R}^2 - \frac{1}{4} F_{\mu\nu}^2 \right\} - \frac{\sqrt{g}}{12} \xi \phi^2 \tilde{R} + \sqrt{g} \left\{ \frac{1}{2} (\tilde{D}_\mu \phi)^2 - \frac{\lambda}{12} \phi^4 \right\}, & \tilde{R}^2 &\rightarrow -2\sigma^2 \tilde{R} - \sigma^4 \\
 &= \sqrt{g} \left\{ -\frac{1}{4!} (\xi_0 \sigma^2 + \xi \phi^2) \tilde{R} - \frac{1}{4} F_{\mu\nu}^2 + \frac{1}{2} (\tilde{D}_\mu \phi)^2 - \frac{1}{4!} (\lambda \phi^4 + \xi_0 \sigma^4) \right\}, & \rho^2 &= \frac{1}{6} (\xi_0 \sigma^2 + \xi \phi^2)
 \end{aligned}$$

- Weyl gauge transformation (b): “gauge fixing”: $\Omega = \frac{\rho^2}{M_p^2}, \quad \hat{\rho} = M_p, \quad \hat{\omega}_\mu = \omega_\mu - \frac{1}{q} \ln \rho^2$

Einstein-Proca action for ω_μ :

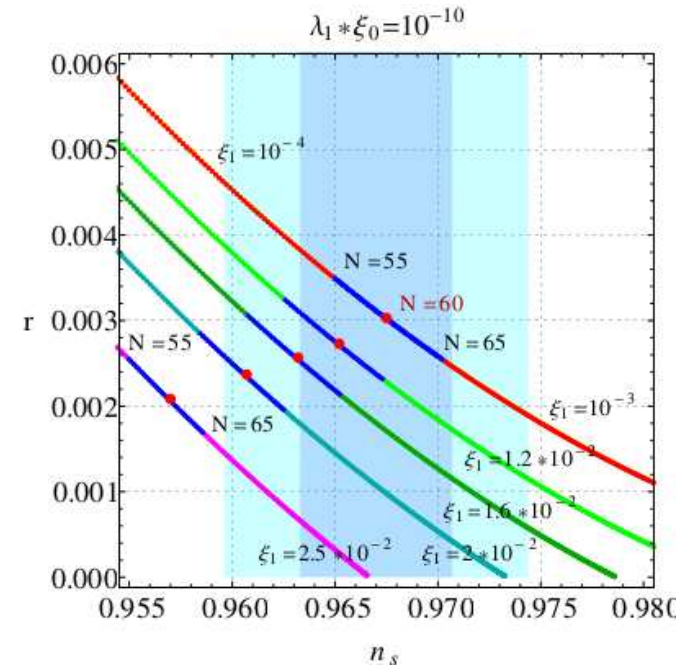
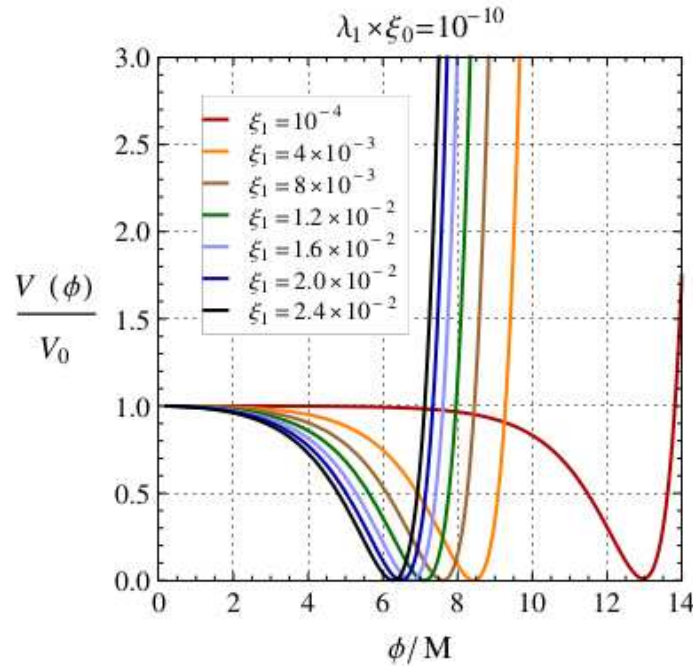
$$L_2 = \sqrt{\hat{g}} \left\{ -\frac{1}{2} M_p^2 \hat{R} + \frac{3}{4} M_p^2 q^2 \hat{\omega}_\mu \hat{\omega}^\mu - \frac{1}{4} \hat{F}_{\mu\nu}^2 + \frac{1}{2} (\hat{D}_\mu \hat{\phi})^2 - V \right\},$$

$$V = \frac{3M_p^4}{2\xi_0} \left[1 - \frac{\xi \hat{\phi}^2}{6 M_p^2} \right]^2 + \frac{\lambda}{4!} \hat{\phi}^4 + \mathcal{O}\left(\frac{\phi^2}{M_p^2}\right), \quad \phi \ll M_p.$$

\Rightarrow No dilaton $\ln \sigma$ left (eaten by ω_μ). Gravitational higgs mechanism. Also $m_\phi^2 = (-\xi/\xi_0) M_p^2$.

\Rightarrow Fixed point for ξ ? The limit $\xi \rightarrow 0 \Rightarrow S_v \times S_h \Rightarrow \beta_\xi \propto \xi?$ not $(\xi + 1/6)!$ Quantum corrections?

- **Weyl inflation:** $V = V_0 \left\{ \left[1 - \xi_1 \sinh^2 \frac{\phi}{M_p \sqrt{6}} \right]^2 + \lambda_1 \xi_0 \sinh^4 \frac{\phi}{M_p \sqrt{6}} \right\}, \quad V_0 = \frac{3}{2} \frac{M_p^4}{\xi_0}; \quad \phi > M_p : \text{OK!}$



$$\lambda_1 \xi_0 \ll \xi_1^2 \ll 1 : \quad \epsilon = \frac{\xi_1^2}{3} \sinh^2 \frac{2\phi}{M_p \sqrt{6}} + \mathcal{O}(\xi_1^3);$$

$$\eta = -\frac{2\xi_1}{3} \cosh \frac{2\phi}{M_p \sqrt{6}} + \mathcal{O}(\xi_1^2)$$

$$n_s = 1 + 2\eta - 6\epsilon = 1 - \frac{4}{3} \xi_1 \cosh \frac{2\phi_*}{M_p \sqrt{6}} + \mathcal{O}(\xi_1^2); \quad \Rightarrow \quad r = 3(1 - n_s)^2 + \mathcal{O}(\xi_1^2)$$

$\Rightarrow 0.002567 \leq r \leq 0.00303$ if $n_s = 0.9670 \pm 0.0037$; $n_s \approx 1 - 2/\bar{N}$, $r = 12/\bar{N}^2$, $\bar{N} \approx N + 4.1$; N e-folds
 r midly smaller than Starobinsky, where: $\bar{N} \rightarrow N$; $r_s \approx 0.0031$, $n_s \approx 0.968$ ($N = 60$)

● Conclusions:

I. Flat space:

- ⇒ Quantum SI can protect a classical hierarchy of: higgs vev \ll dilaton vev
- All scales from fields vev's: hierarchy stability not a DR artefact

II. Curved space:

- Weyl's original gauge symmetry/gravity, viable theory; no ghost dilaton; # dof conserved.
- ⇒ Weyl gravity = Einstein-Proca for $\omega_\mu + \text{cc}$, after Stueckelberg, dilaton eaten by Weyl "photon"
 $m_\omega \sim qM_p$.
- ⇒ Weyl inflation predicts: $0.00257 \leq r \leq 0.00303$ for $N = 60$, $n_s =$ measured 68%CL
(such r values will soon be tested, LiteBIRD, CMB-S4).

• Dilatation current D_μ :
$$D^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_j)} (x^\nu \partial_\nu \phi_j + d_\phi) - x^\mu \mathcal{L}, \quad d_\phi = (d-2)/2 \quad (\text{scalars}),$$

- in $d = 4 - 2\epsilon$, potential \tilde{V} :
$$\begin{aligned} \partial_\mu D^\mu &= (d_\phi + 1) (\partial_\mu \phi_j) \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_j)} + d_\phi \phi_j \frac{\partial \mathcal{L}}{\partial \phi_j} - d \mathcal{L} \\ &= d \tilde{V} - \frac{d-2}{2} \phi_j \frac{\partial \tilde{V}}{\partial \phi_j}, \quad \phi_j = \phi, \sigma; \quad (\text{onshell, canonical k.t.}) \end{aligned}$$

• in SI theory: \tilde{V} homogeneous in d dim's: $\tilde{V}(\rho \phi_j) = \rho^{2d/(d-2)} \tilde{V}(\phi_j) \Rightarrow \partial_\mu D^\mu = 0$, conserved.

• in “usual” reg: $\mu = \text{const}$, no σ : $\tilde{V} = \mu^{2\epsilon} V(\phi)$, and V is scale inv in $d = 4$:

$$\partial_\mu D^\mu = 2\epsilon \mu^{2\epsilon} V \sim 2\epsilon \mu^{2\epsilon} \left[\lambda + \frac{\beta_\lambda}{\epsilon} + \dots \right] \frac{\partial V}{\partial \lambda} \propto \beta_\lambda \frac{\partial V}{\partial \lambda}, \quad [=0 \text{ only if } \beta_\lambda = 0]. \quad (\text{scale anomaly})$$

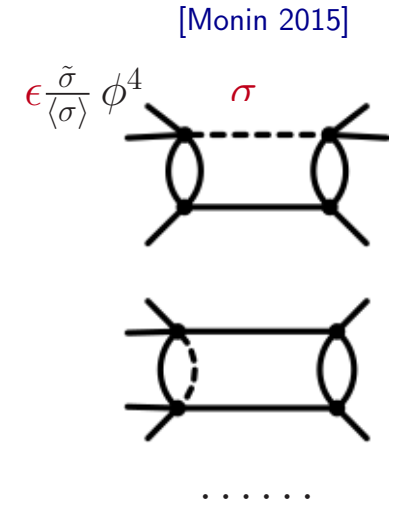
- different field content! [Shaposhnikov et al, Tamarit 2013]

“For scale invariance, though, the situation is hopeless; any cutoff procedure necessarily involves a large mass and a large mass necessarily breaks scale invariance in a large way. This argument does not show that the occurrence of anomalies is inevitable [.....]” (S. Coleman: Aspects of Symmetry, p.82, 1985).

• **Three-loop SI potential: Counterterms!**

$$\delta L_3 = \frac{1}{2} \delta_\phi^{(3)} (\partial_\mu \phi)^2 - \mu^{2\epsilon} \left\{ \frac{1}{4!} \delta_\lambda^{(3)} \lambda \phi^4 + \frac{1}{6} \delta_{\lambda_6}^{(3)} \lambda_6 \frac{\phi^6}{\sigma^2} + \frac{1}{8} \delta_{\lambda_8}^{(3)} \lambda_8 \frac{\phi^8}{\sigma^4} \right\}$$

$$\delta_{\lambda_6}^{(3)} = \frac{3 \lambda^4}{2 \lambda_6 \kappa^3 \epsilon}, \quad \delta_{\lambda_8}^{(3)} = \frac{275 \lambda^4}{864 \lambda_8 \kappa^3 \epsilon} \Rightarrow \gamma_\phi^{(3)}, \beta_{\lambda_6}, \beta_{\lambda_8} = \dots$$



Integrate CS: Three-loop term: $U_3 = \Delta V + V^{(3)} + V^{(3,n)}$. $V^{(3)}$ ='old' [$\mu \rightarrow z\sigma$]; **new: $V^{(3,n)}$ and ΔV**

$$\Delta V = \frac{\lambda_6 \phi^6}{\sigma^2} + \frac{\lambda_8 \phi^8}{\sigma^4}, \quad V^{(3)} = \frac{\lambda^4 \phi^4}{\kappa^3} \left\{ \mathcal{Q} + \left(\frac{97}{128} + \frac{9}{64} A_0 + \frac{\zeta[3]}{4} \right) \overline{\ln} \frac{V_{\phi\phi}}{(z\sigma)^2} - \frac{31}{96} \overline{\ln}^2 \frac{V_{\phi\phi}}{(z\sigma)^2} + \frac{9}{64} \overline{\ln}^3 \frac{V_{\phi\phi}}{(z\sigma)^2} \right\}$$

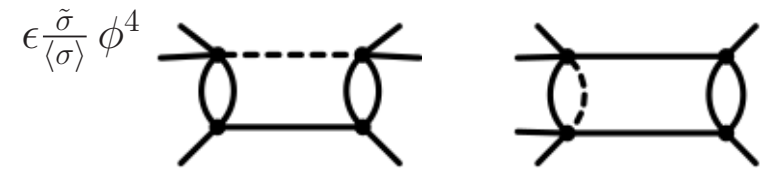
$$V^{(3,n)} = \frac{\lambda^3 \phi^4}{2 \kappa^3} \left\{ \left(27\lambda - \frac{\lambda_6}{2} \right) \frac{\phi^2}{8 \sigma^2} + \left(\frac{401 \lambda}{72} - \lambda_8 \right) \frac{\phi^4}{16 \sigma^4} \right\} \overline{\ln} \frac{V_{\phi\phi}}{(z\sigma)^2}.$$

$$V_{\phi\phi} \equiv \frac{\lambda}{2} \phi^2.$$

[D.G. arXiv:1712.06024]

\Rightarrow SI effective operators always **suppressed by σ** \Rightarrow . At large σ enhanced symmetry $S_v \times S_h$ restored \Rightarrow

\Rightarrow **forbids c-terms** $\lambda^2 \phi^2 \sigma^2 = \lambda^2 \langle \sigma \rangle^2 \phi^2 + \dots$ No tuning of Higgs selfcoupling λ for **large σ** .



• Three-loop SI potential

Counterterm:
$$\delta L_3 = \frac{1}{2} \delta_\phi^{(3)} (\partial_\mu \phi)^2 - \mu^{2\epsilon} \left(\frac{1}{4!} \delta_\lambda^{(3)} \lambda \phi^4 + \frac{1}{6} \delta_{\lambda_6}^{(3)} \lambda_6 \frac{\phi^6}{\sigma^2} + \frac{1}{8} \delta_{\lambda_8}^{(3)} \lambda_8 \frac{\phi^8}{\sigma^4} \right)$$
 [Monin 2015]

$$\delta_\phi^{(3)} = -\frac{\lambda^3}{4\kappa^3} \left(\frac{1}{6\epsilon^2} - \frac{1}{12\epsilon} \right), \quad \delta_{\lambda_6}^{(3)} = \frac{3}{2} \frac{\lambda^4}{\lambda_6 \kappa^3 \epsilon}, \quad \delta_{\lambda_8}^{(3)} = \frac{275}{864} \frac{\lambda^4}{\lambda_8 \kappa^3 \epsilon}.$$

So $\gamma_\phi^{(3)} = \lambda^3/(16\kappa^3)$. With $Z_X = 1 + \delta_X$ and $\lambda_6^B = \mu^{2\epsilon}(\sigma)\lambda_6 Z_{\lambda_6} Z_\phi^{-3} Z_\sigma$ and $(d/d \ln z) \lambda_6^B = 0$,

$$\beta_{\lambda_6} = \frac{\lambda^2 \lambda_6}{2\kappa^2} + \frac{\lambda^3}{\kappa^3} \left(9\lambda - \frac{3}{8}\lambda_6 \right), \quad (\text{similar } \beta_{\lambda_8}).$$

Callan-Symanzik eq: $V^{(3)}$ [“usual” with $\mu \rightarrow \sigma$] + $V^{(3,n)}$ [new]: gives solution shown on previous page:

$$\frac{\partial V^{(3)}}{\partial \ln z} + \beta_\lambda^{(1)} \frac{\partial V^{(2)}}{\partial \lambda} + \beta_\lambda^{(2)} \frac{\partial V^{(1)}}{\partial \lambda} + \beta_\lambda^{(3)} \frac{\partial V}{\partial \lambda} + \gamma_\phi^{(2)} \frac{\partial V^{(1)}}{\partial \ln \phi} + \gamma_\phi^{(3)} \frac{\partial V}{\partial \ln \phi} = \mathcal{O}(\lambda_j^5).$$

$$\frac{\partial V^{(3,n)}}{\partial \ln z} + \beta_{\lambda_j}^{(1)} \frac{\partial V^{(2,n)}}{\partial \lambda_j} + \beta_{\lambda_j}^{(3,n)} \frac{\partial V}{\partial \lambda_j} = \mathcal{O}(\lambda_j^5), \quad \lambda_j = \lambda, \lambda_6, \lambda_8.$$

- **SM+dilaton: one-loop beta functions** (similar to when $\mu = \text{constant}$, same field content):

$$\beta_{\lambda_\phi} = \frac{1}{\kappa} \left[3 \left(\frac{9}{4}g_2^4 + \frac{3}{4}g_1^4 + \frac{3}{2}g_1^2g_2^2 - 12h_t^4 \right) - 4\lambda_\phi \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) + 4\lambda_\phi^2 + 3\lambda_m^2 + 96\lambda_m\lambda_6 \right]$$

$$\beta_{\lambda_m} = \frac{2\lambda_m}{\kappa} \left[\lambda_\phi + 2\lambda_m + \frac{1}{2}\lambda_\sigma - \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) \right] \propto \lambda_m, \quad \Rightarrow \lambda_m \text{ stays small (f.p.).}$$

- for new couplings

$$\beta_{\lambda_6} = \frac{3\lambda_6}{\kappa} \left[6\lambda_\phi - 8\lambda_m + \lambda_\sigma - 2 \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) \right]$$

$$\beta_{\lambda_8} = \frac{2}{\kappa} \left[2\lambda_6 (28\lambda_6 + \lambda_m) - 4\lambda_8 \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) \right]$$

$$\beta_{\lambda_{10}} = 10 \left[4\lambda_6^2 - \lambda_{10} \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) \right]$$

$$\beta_{\lambda_{12}} = 2 \left[3\lambda_6^2 - 6\lambda_{12} \left(\frac{3}{4}g_1^2 + \frac{9}{4}g_2^2 - 3h_t^2 \right) \right]$$

\Rightarrow if one sets $\lambda_{6,8,10,\dots} = 0$ at tree level, then $\beta_{\lambda_{6,8,10,12}} = 0$ at one-loop, but emerge at 2-loops.

- Weyl quadratic gravity with matter \rightarrow Einstein gravity + massive ω_μ

[DG arXiv:1812.08613, 1904.06596]

$$L_1 = \sqrt{g} \left[\frac{\xi_0}{4!} \tilde{R}^2 - \frac{1}{4} F_{\mu\nu}^2 \right], \quad \xi_0 > 0, \quad \text{original Weyl action (1918)}$$

$$= \sqrt{g} \left[\frac{\xi_0}{4!} (-2\sigma^2 \tilde{R} - \sigma^4) - \frac{1}{4} F_{\mu\nu}^2 \right] \quad \text{eom: } \sigma^2 = -\tilde{R}; \quad \text{dilaton: } \ln \sigma \rightarrow \ln \sigma - \ln \Omega$$

Use $\tilde{R} = R - 3q D_\mu \omega^\mu - 3/2 q^2 \omega_\mu \omega^\mu$ then, in Riemann language:

$$L_1 = \sqrt{g} \left\{ -\frac{\xi_0}{2} \left[\frac{1}{6} \sigma^2 R + (\partial_\mu \sigma)^2 \right] - \frac{\xi_0}{4!} \sigma^4 + \frac{q^2}{8} \xi_0 \sigma^2 (\omega_\mu - 1/q \partial_\mu \ln \sigma^2)^2 - \frac{1}{4} F_{\mu\nu}^2 \right\}.$$

- gauge transf (b) “gauge fixing!” $\Omega = \frac{\xi_0 \sigma^2}{6M_p^2} \Rightarrow \hat{g}_{\mu\nu} = \Omega^2 g_{\mu\nu}, \quad \hat{\sigma}^2 = \frac{6M_p}{\xi_0}, \quad \hat{\omega}_\mu = \omega_\mu - \frac{1}{q} \partial_\mu \ln \sigma^2$

$$L_1 = \sqrt{\hat{g}} \left\{ -\frac{1}{2} M_p^2 \hat{R} - \frac{3M_p^4}{2\xi_0} + \frac{3}{4} q^2 M^2 \hat{\omega}_\mu \hat{\omega}^\mu - \frac{1}{4} \hat{F}_{\mu\nu}^2 \right\}.$$

\Rightarrow Einstein-Proca action for (massive) ω_μ which has “eaten” $\ln \sigma$ field! no ghost! # dof conserved!

$$\text{cons. current } \partial^\alpha (F_{\alpha\mu} \sqrt{g}) + \frac{1}{2} \sqrt{g} \xi_0 \phi_0 \left[\partial_\mu - \frac{q}{2} \omega_\mu \right] \phi_0 = 0$$