

# Current Status of Smooth Quantum Gravity

*Jerzy Król<sup>a,b</sup> & Torsten Asselmeyer-Maluga<sup>c</sup>*

<sup>a</sup>University of Silesia, Institute of Physics, Katowice

<sup>b</sup>University of Information Technology and Management  
Rzeszów, Poland

<sup>c</sup>German Aerospace Center (DLR), Berlin, Germany

*BEYOND General Relativity*

Warszawa, 01-05. 07. 2019

# smooth quantum gravity - motivations

In dim. 4 there are (typically) continuum infinite many different smoothness structures on open 4-manifolds like  $\mathbb{R}^4$  or  $S^3 \times \mathbb{R}$ . These standard manifolds are extensively used in physics. Do their exotic smoothness is physically (especially QG) valid)?

Or: Given exotic  $R^4$  it is Riemannian smooth 4-manifold homeomorphic to  $\mathbb{R}^4$ . Its Riemannian curvature tensor can not vanish! So exotic  $R^4$  has to have non-zero curvature and density of gravitational energy is non-zero as well. Is this curvature physical?

- i) YES, the Riemannian curvature of exotic  $R^4$ 's leads directly to QG.
- ii) Embeddings of exotic  $R^4$ 's determine topological 3-D invariants with cosmological meaning.

# smooth quantum gravity - motivations

In dim. 4 there are (typically) continuum infinite many different smoothness structures on open 4-manifolds like  $\mathbb{R}^4$  or  $S^3 \times \mathbb{R}$ . These standard manifolds are extensively used in physics. Do their exotic smoothness is physically (especially QG) valid)?

Or: Given exotic  $R^4$  it is Riemannian smooth 4-manifold homeomorphic to  $\mathbb{R}^4$ . Its Riemannian curvature tensor can not vanish! So exotic  $R^4$  has to have non-zero curvature and density of gravitational energy is non-zero as well. Is this curvature physical?

- i) YES, the Riemannian curvature of exotic  $R^4$ 's leads directly to QG.
- ii) Embeddings of exotic  $R^4$ 's determine topological 3-D invariants with cosmological meaning.

SQG IS A THEORETICAL ATTEMPT TO UNDERSTAND AND DERIVE SOME FREE PARAMETERS IN PHYSICS BASED ON TOPOLOGY UNDERLYING SMOOTH EXOTIC  $\mathbb{R}^4$

## smooth quantum gravity - some results

- a. LARGE EXOTIC  $R^4$ 'S ARE GRAVITATIONAL INSTANTONS. ANY THEORY OF QG HAS TO DEAL WITH THEM [G. ETESI, 2019].
- b. THE PATH INTEGRAL OF GR IS DOMINATED BY CONTRIBUTIONS FROM EXOTIC  $R^4$ 'S [TAM, 2016].
- c. THE CURVATURE OF  $R^4$  EMBEDDED IN  $K3\#\overline{CP^2}$  DETERMINES THE REALISTIC (SMALL) VALUE OF THE COSM. CONST. [TAM, JK, 2018]
- d. THIS CC IS A TOPOLOGICAL INVARIANT [TAM, JK, 2018].
- e. THE COSMOLOGICAL FRW MODEL ON EXOTIC  $S^3 \times \mathbb{R}$  (FROM  $R^4 \hookrightarrow K3\#\overline{CP^2}$ ) PREDICTS AND EXPLAINS THE REALISTIC VALUES OF INFLATION PARAMETERS [TAM, JK, 2019].
- f. THE ELECTROWEAK AND GUT'S SCALES AND THE BOUND ON NEUTRINO MASSES ARE ALSO PREDICTED [TAM, JK, 2019].
- g. EXOTIC  $R^4$ 'S DETERMINE VON NEUMAN ALGEBRAS CONTAINING FACTOR  $III_1$  [G. ETESI, 2018, TAM 2016].

# The realistic (small) value of the Cosm. Constant

- ▶ THE FORMULA FOR CC (curvature of exotic  $R^4$  embedded in  $K3\#\overline{CP^2}$ )

$$\Omega_\Lambda = \frac{c^5}{24hGH_0^2} \cdot \exp\left(-\frac{3}{CS(\Sigma(2, 5, 7))} - \frac{3}{CS(P\#P)} - \frac{\chi(A)}{4}\right)$$

here  $\chi(A) = 1$  is the Euler characteristic of the Akbulut cork  $A$ ,  $\Sigma(2, 5, 7)$  is the Brieskorn homology 3-sphere,  $CS(\Sigma(2, 5, 7))$  - the Chern-Simons invariant of  $\Sigma$ ,  $P\#P$  - the connected sum of two copies of the Poincaré 3-spheres. The CC value follows

$$\Omega_\Lambda \approx 0,7029$$

which agrees with PLANCK

THIS CC IS A TOPOLOGICAL INVARIANT!

- ▶ Why it is so? The derivation follows from hyperbolic geometry of 3 and 4-manifolds (cobordisms) in the embedding  $R^4 \hookrightarrow K3\#\overline{CP^2}$  [T.Asselmeyer-Maluga, JK, Phys.Dark.Univ.2018].

## Two topology changes

Exotic  $R^4$ 's are topologically trivial, where does the nontrivial topology come from?

## Two topology changes

Exotic  $R^4$ 's are topologically trivial, where does the nontrivial topology come from?

The embedding  $R^4 \hookrightarrow K3 \# \overline{CP^2}$  determines two topology changes (3-dimensional):

$$S^3 \xrightarrow{1} \Sigma(2, 5, 7) \xrightarrow{2} P \# P$$

- ▶ THE EVOLUTION IN  $K3$  MUST BE SMOOTH THROUGH BOTH TOPOLOGY CHANGES, SO WE NEED TO GLUE IN 3 CASSON HANDLES AT THE SECOND STEP AND ONE IN THE FIRST.

Based on this we can determine energy scales of both transitions...

# 1<sup>st</sup> TOPOLOGY CHANGE

Inside  $K3$  there is a cork (Akbulut) - a compact, contractible smooth submanifold  $A \subset K3$ .

- ▶ THE BOUNDARY  $\partial A$  IS A HOMOLOGY 3-SPHERE  $\Sigma(2, 5, 7)$  (HYPERBOLIC).
- ▶  $\Sigma(2, 5, 7)$  can not be replaced by  $S^3$ : inside  $A$  there is smoothly embedded  $S^3$  but there is no smooth  $S^3 \subset K3$  such that  $A \subset S^3$ .

But

**There exists smooth 4-cobordism in  $K3$  between  $S^3 \subset A$  and  $\Sigma(2, 5, 7)$**



## SMOOTH COB $S^3 \rightarrow \Sigma(2, 5, 7)$

To go smoothly from  $S^3$  to  $\Sigma(2, 5, 7)$  we need to glue in flexible handles – Casson handles (CH). Then we get smooth 4-cobordism  $W(S^3, \Sigma) \subset A \subset K3$ .

- ▶  $\Sigma(2, 5, 7)$  IS HYPERBOLIC AND IT IS RIGID – ONE CAN NOT SCALE IT AND ITS VOLUME  $V$  IS AN INVARIANT.
- ▶ THUS  $\Sigma$  DETERMINES CHARACTERISTIC LENGTH  $\sqrt[3]{V} = L$ .
- ▶ EXPRESSING VIA CS INVARIANT AND TAKING THE RADIUS OF 3-SPHERE AS  $r_{S^3}$  WE OBTAIN THE RESCALING

$$a = r_{S^3} \cdot \exp\left(\frac{3}{2 \cdot CS(\Sigma(2, 5, 7))}\right).$$

# ENERGY AND TIME SCALES OF 1

- ▶ FREEDMAN: EVERY CH IS EMBEDDABLE IN ITS 1ST 3-STAGES.

Then (for  $\theta = \frac{3}{2 \cdot CS(\Sigma(2,5,7))}$ )

$$E_1 = \frac{E_P}{1 + \theta + \frac{\theta^2}{2} + \frac{\theta^3}{6}} \simeq 10^{15} \text{ GeV} \quad t_1 = t_P \left( 1 + \theta + \frac{\theta^2}{2} + \frac{\theta^3}{6} \right) \simeq 10^{-39}$$

**$E_1$  ( $\sim$  GUT energy) – topologically supported**

## ENERGY AND TIME SCALES OF 2

- ▶  $K3$  is decomposed as

$$K3 = |E_8 \oplus E_8| \# (S^2 \times S^2) \# (S^2 \times S^2) \# (S^2 \times S^2).$$

- ▶  $|E_8|$  HAS BOUNDARY  $P$  – THE POINCARÉ SPHERE.  $|E_8 \oplus E_8|$  CAN NOT BE REALIZED AS SMOOTH CLOSED 4-MANIFOLD (DONALDSON). IT HAS BOUNDARY  $P \# P$  (3-SUBMANIFOLD OF  $K3$ ).
- ▶  $\Sigma(2, 5, 7) \xrightarrow{2} P \# P$  HERE WE NEED FULL INFINITE 3CH'S SINCE  $|E_8 \oplus E_8|$  IS NOT SMOOTH (DONALDSON). AFTER GLUING CH'S WE HAVE SMOOTH EVOLUTION 2 WITHIN  $K3 \# \overline{CP}(2)$ .

## ENERGY AND TIME SCALES OF 2

Again  $\theta = \frac{3}{2 \cdot CS(\Sigma(2,5,7))} = \frac{140}{3}$  and taking  $\Delta_0 t$  as the time for 1-level CH we have for the total time and energy after 2<sup>nd</sup> change ( $CS(P\#P) = \frac{1}{60}$ ):

$$\Delta t = \frac{t_{PI} \cdot \exp\left(\frac{-1}{2CS(P\#P)}\right)}{1 + \theta_2 + \frac{\theta_2^2}{2} + \frac{\theta_2^3}{6}}, \quad \Delta E = \frac{E_{PI} \cdot \exp\left(\frac{-1}{2CS(P\#P)}\right)}{1 + \theta_2 + \frac{\theta_2^2}{2} + \frac{\theta_2^3}{6}} \simeq 63 \text{ GeV}.$$

**2<sup>nd</sup> energy scale is close to the electroweak scale and it is topologically supported**

## MASSIVE NEUTRINOS

Seesaw mechanism for generating the masses of neutrinos:

$$\begin{pmatrix} 0 & M \\ M & B \end{pmatrix}, \quad M \ll B, \quad \lambda_1 = B, \quad \lambda_2 = -\frac{M^2}{B}.$$

Let us fix the energy scales  $E_1 = 0,67 \cdot 10^{15} \text{ GeV}$  as  $B$  and  $E_2 = 63 \text{ GeV}$  as  $M$ . Then  $m_n = \frac{M^2}{B} \simeq 0,006 \text{ eV}$  – agrees with current limitation for the sum of 3 neutrino masses from PLANCK  $\sim 0,12 \text{ eV}$ .

**Topology determines the realistic neutrino masses.**

## LEFT-RIGHT NEUTRINOS

- ▶ 1<sup>st</sup> topology change:  $S^3 \rightarrow \Sigma(2, 5, 7)$  with the GUT scale gives rise to 2 Dirac operators on  $\Sigma$ , since the mapping class group of  $\Sigma$  is nontrivial

$$\pi_0(\text{Diff}(\Sigma(2, 5, 7))) = \mathbb{Z}_2.$$

hence left and right-handed neutrinos are there.

- ▶ The 2<sup>nd</sup> topology change  $\Sigma(2, 5, 7) \rightarrow P\#P$  gives rise only to left-handed neutrinos (EW energy scale) since

$$\pi_0(\text{Diff}(P)) = 1.$$

**We found 'first principles' allowing for derivation of certain free parameters in Cosmology and Particle Physics. This is  $R^4 \hookrightarrow K3\#\overline{CP^2}$ ,  $R^4$  exotic.**

## AGAIN 1<sup>st</sup> TOPOLOGY CHANGE

Let  $r_{S^3} \sim P_L$  to be of Planck length thus (with  $CS(\Sigma) = 9/280$ )

$$10^{-34}[m] \rightarrow 10^{-15}[m], \quad N = \frac{3}{2 \cdot CS(\Sigma(2,5,7))} + \ln 8\pi^2 \simeq 51.$$

- ▶ THE TOPOLOGY OF THE AKBULUT CORK FOR EXOTIC K3 CODES THE INFLATIONARY EXPANSION OF THE UNIVERSE.

# STAROBINSKY MODEL FOR INFLATION

$$S = \int_{M^4} d^4x \sqrt{-g} (R + \alpha \cdot R^2), \quad \alpha \text{ free param.}$$

$\alpha \cdot M_{Pl}^{-2} = \frac{\Delta E_{\text{infl}}}{E_{Pl}}$  SO WE GET FROM 1<sup>st</sup> TOPOLOGY CHANGE:

$$\alpha \sim 10^{-5}; \quad \left( \alpha \cdot M_{Pl}^{-2} = \frac{1}{1 + \theta + \frac{\theta^2}{2} + \frac{\theta^3}{6}}, \quad \theta = \frac{3}{2 \cdot CS(\Sigma(2, 5, 7))} \right)$$

the spectral tilt  $n_s$  and the tensor-scalar ratio  $r$  follow

$$n_s = 1 - \frac{2}{\theta + \ln(8\pi^2)} \approx 0,961, \quad r = \frac{12}{(\theta + \ln(8\pi^2))^2} \approx 0,0046.$$

**$\alpha, n_s, r$  are topologically supported due to the K3 smoothness structure**



# SMOOTH QUANTUM GRAVITY

Exclusively in dimension 4:

- ▶ NONSTANDARD SMOOTHNESS  $\Rightarrow$  NONZERO CURVATURE OF  $R^4$
- ▶ NONSTANDARD SMOOTHNESS  $\Rightarrow$  QUANTIZATION IN SPACETIME

Large exotic  $R^4$  embedded in  $K3\#\overline{CP^2}$  is Ricci flat, hyperkähler, self-dual, hence gravitational instanton [Etesi, 2019]. Any QG theory has to deal with them.

# SMOOTH QUANTUM GRAVITY

- ▶ THE 3-SPHERE  $S^3 \leftrightarrow A \leftrightarrow K3$  IS WIDELY EMBEDDED – REPRESENTS QUANTUM STATE.
- ▶ CONNES: WILD  $S^3$  GENERATES THE FACTOR  $III_1$  VON NEUMAN ALGEBRA AND THE FOCK SPACE OF CERTAIN QFT.
- ▶ WHEN SMOOTHNESS OF  $K3$  AND  $R^4$  ARE STANDARD THE SPHERE  $S^3$  IS TAME AND NO QUANTUM ALGEBRAS RESULT.

**SQG explores the overlapping contact space of QM and GR via exotic differentiable structure on  $K3$  and  $R^4$ .**

THANK YOU  
FOR YOUR ATTENTION !!

- ▶ Asselmeyer-Maluga, T.; Król, J. How to obtain a cosmological constant from small exotic  $\mathbb{R}^4$ . *Phys. Dark Universe* 2018, 19, p. 66-77.
- ▶ Etesi, G. The von Neumann algebra of smooth four-manifolds and a quantum theory of space-time and gravity, arXiv:1712.01828
- ▶ Król, J.; Asselmeyer-Maluga, T.; Bielas, K.; Klimasara, P. From Quantum to Cosmological Regime. The Role of Forcing and Exotic 4-Smoothness. *Universe* 2017, 3, 31.
- ▶ Asselmeyer-Maluga, T.; Król, J. A topological approach to Neutrino masses by using exotic smoothness, *Mod. Phys. Lett.* vol 33 No 1, p. 1950097-1 arXiv:1801.10419v3
- ▶ Asselmeyer-Maluga, T., Smooth quantum gravity: Exotic smoothness and quantum gravity, in *At the Frontier of Spacetime*, ed. T. Asselmeyer-Maluga, *Fundamental Theories of Physics* vol 183, pp. 247-308, Springer: Switzerland, 2016.