

# Connection between neutrino flux, muon excess and LHC data

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Karlsruhe, Germany



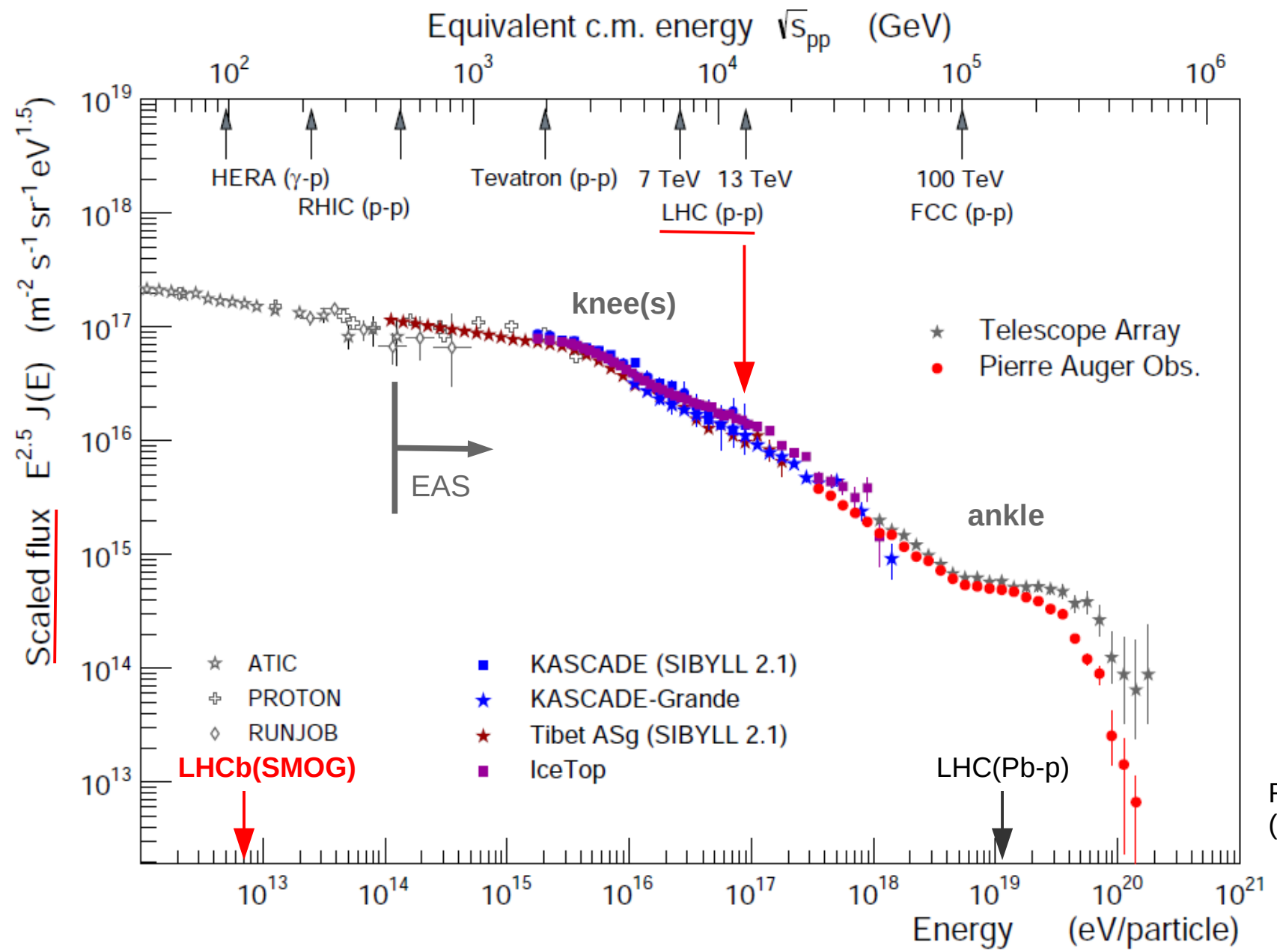
**Workshop for Atmospheric Neutrino Production 2019,  
Nagoya, Japan  
March the 21<sup>st</sup> 2019**

# Outline

- Hadronic interactions (Monte-carlo (MC)) and extended air showers (EAS)
- Muon signal in EAS and simulations
- New input from LHC
- Consequence on inclusive muon and neutrino flux from cosmic rays

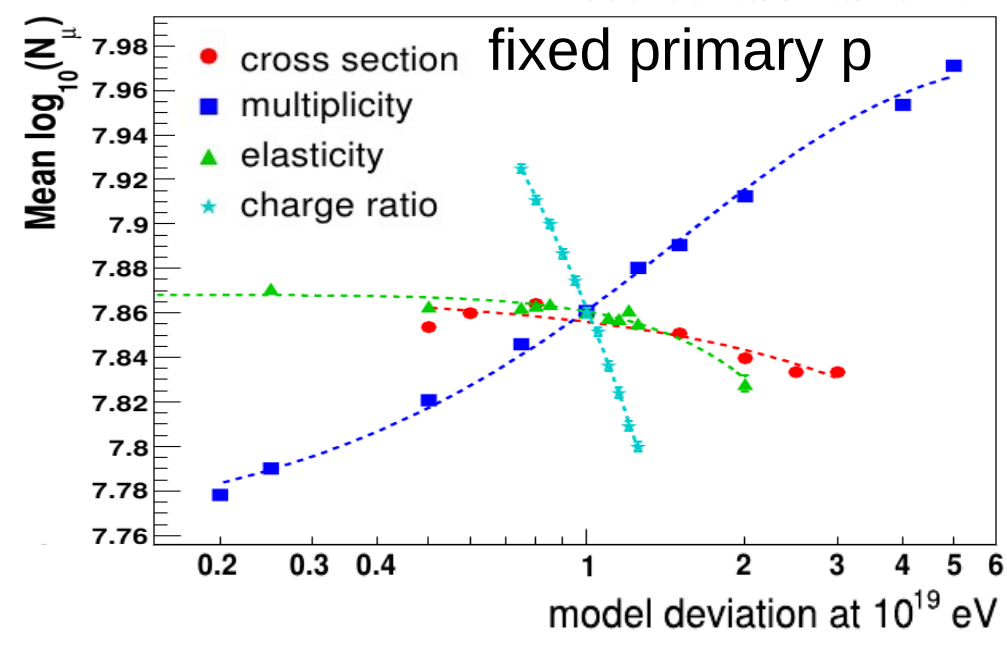
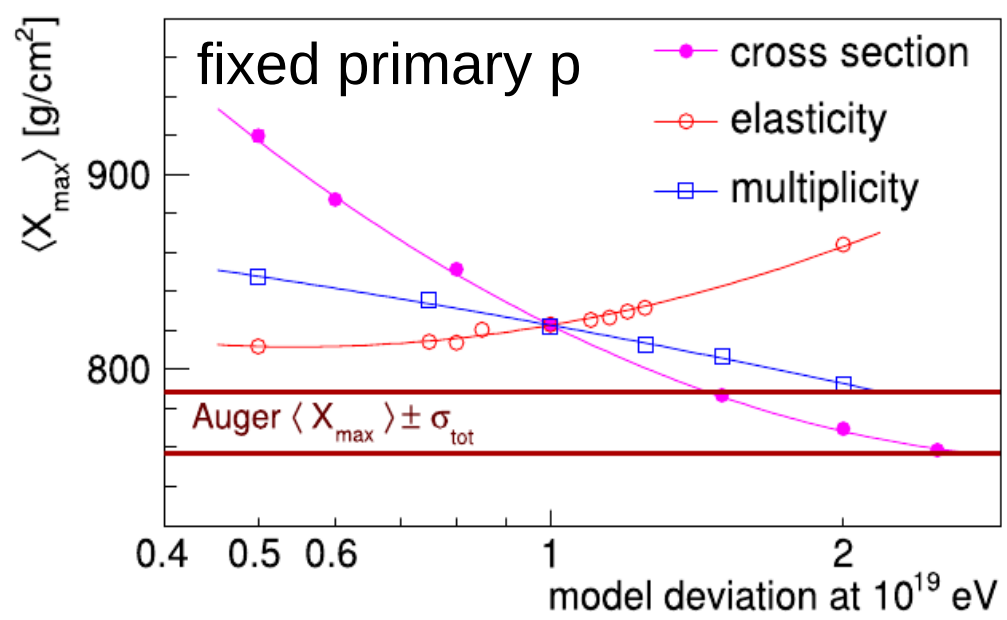
**New input from LHC** which could help to reproduce **EAS data consistently** but solution still under study. Possible consequences on TeV to PeV muons and neutrinos.

# Energy Spectrum



R. Engel (KIT)

# Sensitivity to Hadronic Interactions



- Air shower development dominated by few parameters
  - ➔ mass and energy of primary CR
  - ➔ cross-sections (p-Air and (π-K)-Air)
  - ➔ (in)elasticity
  - ➔ multiplicity
  - ➔ charge ratio and baryon/resonance production
- Change of primary = change of hadronic interaction parameters
  - ➔ cross-section, elasticity, mult. ...

**With unknown mass composition hadronic interactions can only be tested using various observables which should give consistent mass results**

From R. Ulrich (KIT)

# Cosmic Ray Analysis from Air Showers

- EAS simulations necessary to study high energy cosmic rays

- complex problem: identification of the primary particle from the secondaries



- Hadronic models are the key ingredient !

- follow the standard model (QCD)

- but mostly non-perturbative regime (phenomenology needed)

- main source of uncertainties

- Which model for CR ? (alphabetical order)

- **DPMJETIII.17-1** by S. Roesler, A. Fedynitch, R. Engel and J. Ranft

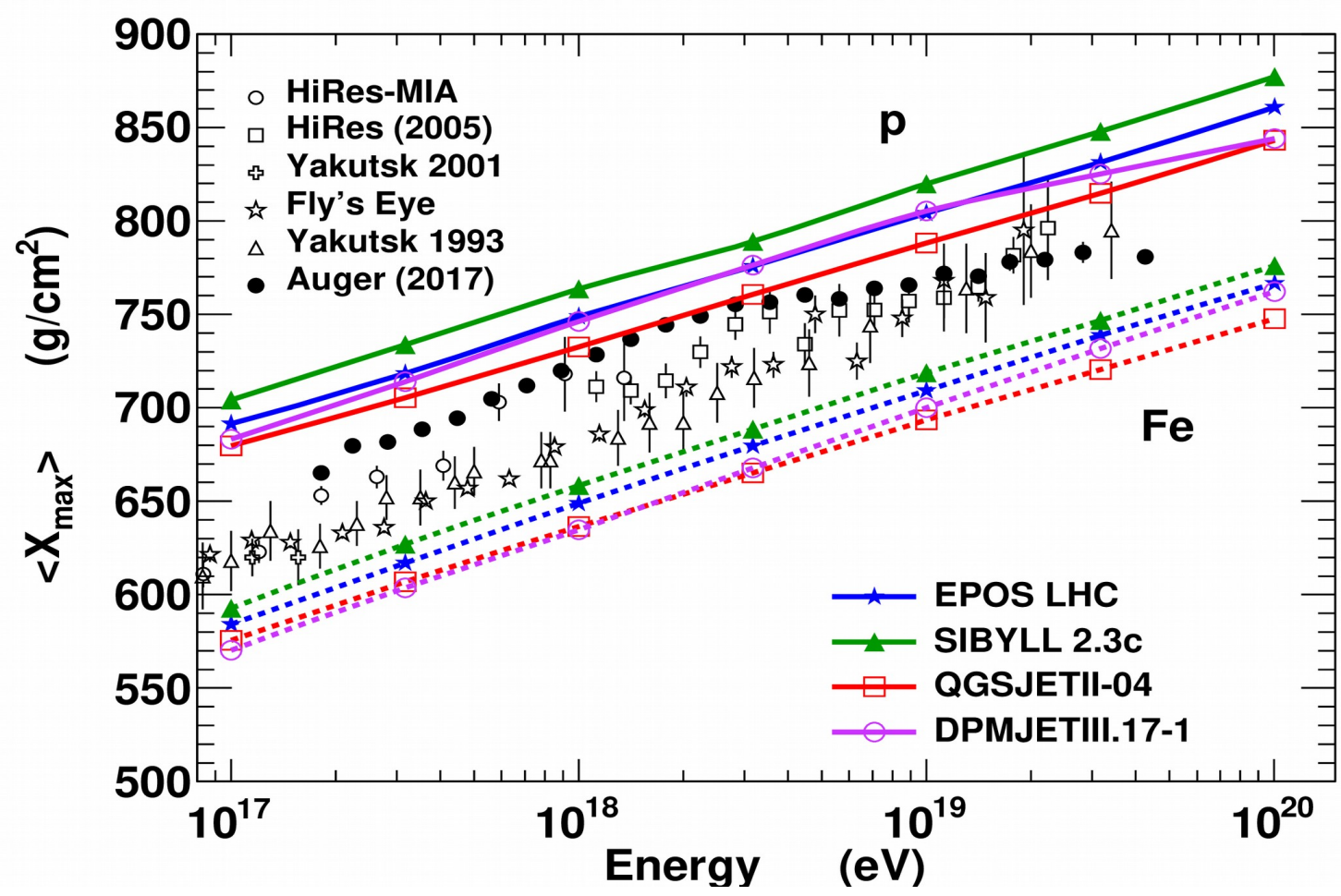
- **EPOS (1.99/LHC/3)** (from VENUS/NEXUS before) by H.J. Drescher, F. Liu, T. Pierog and K.Werner.

- **QGSJET** (01/II-03/II-04/III) by S. Ostapchenko (starting with N. Kalmykov)

- **Sibyll (2.1/2.3c)** by E-J Ahn, R. Engel, R.S. Fletcher, T.K. Gaisser, P. Lipari, F. Riehn, T. Stanev

$\langle X_{max} \rangle$

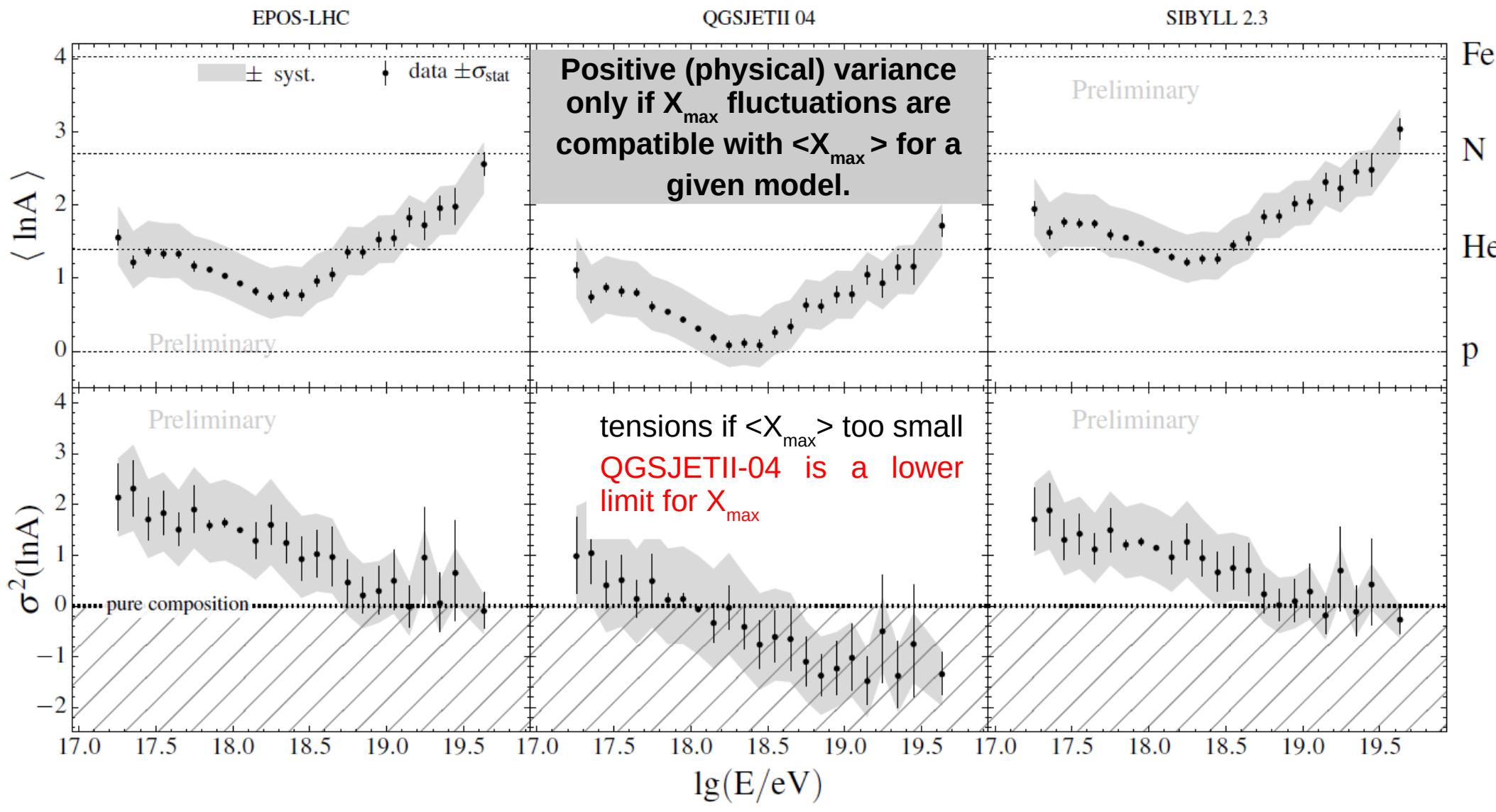
- very similar elongation rate (slope) for all models
  - same mass composition evolution
  - still differences in absolute values
- ➔ +/- 20g/cm<sup>2</sup> is a realistic uncertainty band



# Model Consistency using Electromagnetic Component

## Study by Pierre Auger Collaboration

➔ std deviation of  $\ln A$  allows to test model consistency.



# Cross-Section

For all models cross-section calculation based on optical theorem

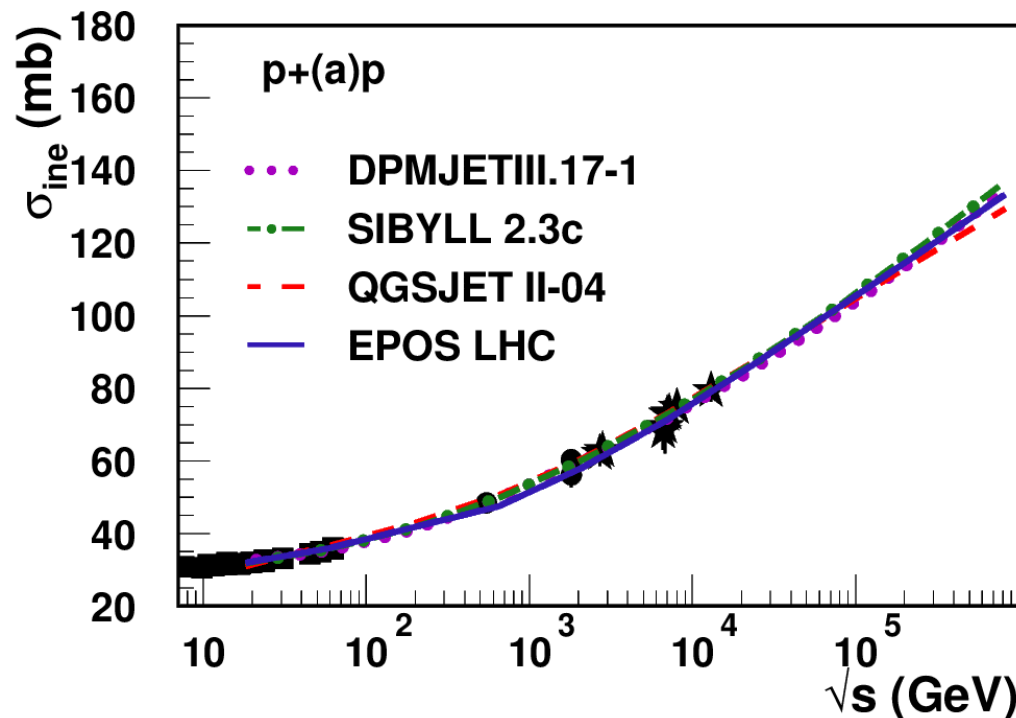
→ total cross-section given by elastic amplitude

$$\sigma_{\text{tot}} = \frac{1}{s} \Im m(A(s, t \rightarrow 0))$$

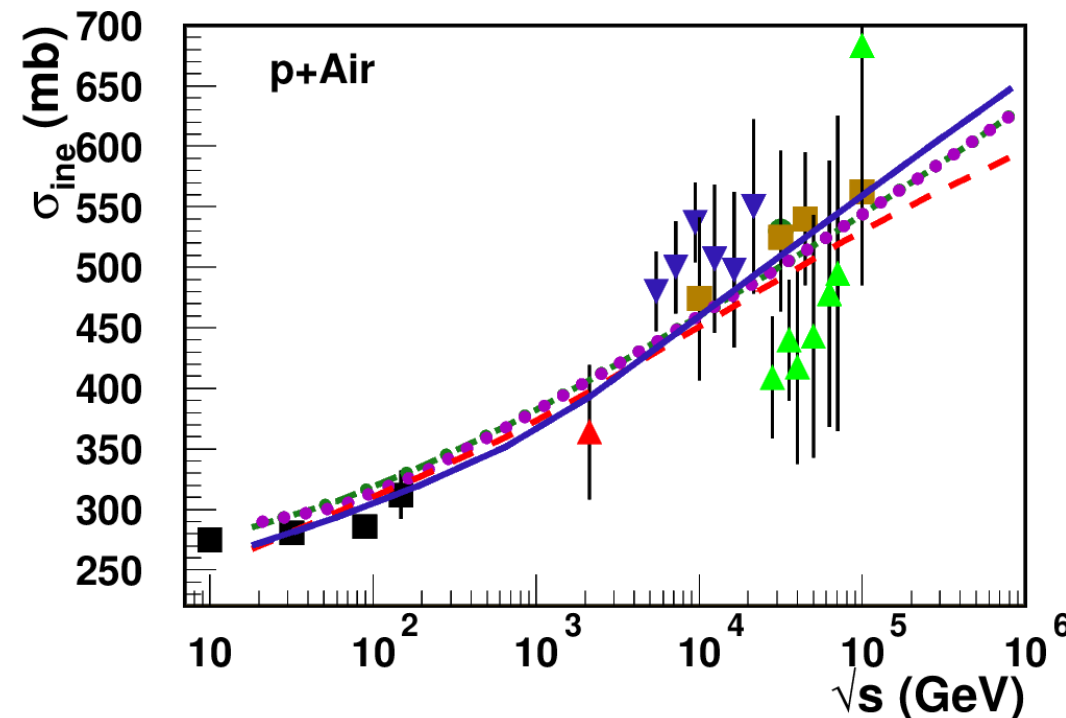
→ different amplitudes in the models but free parameters set to reproduce all p-p cross-sections

→ basic principles + high quality LHC data = same extrapolation

pp



p-Air

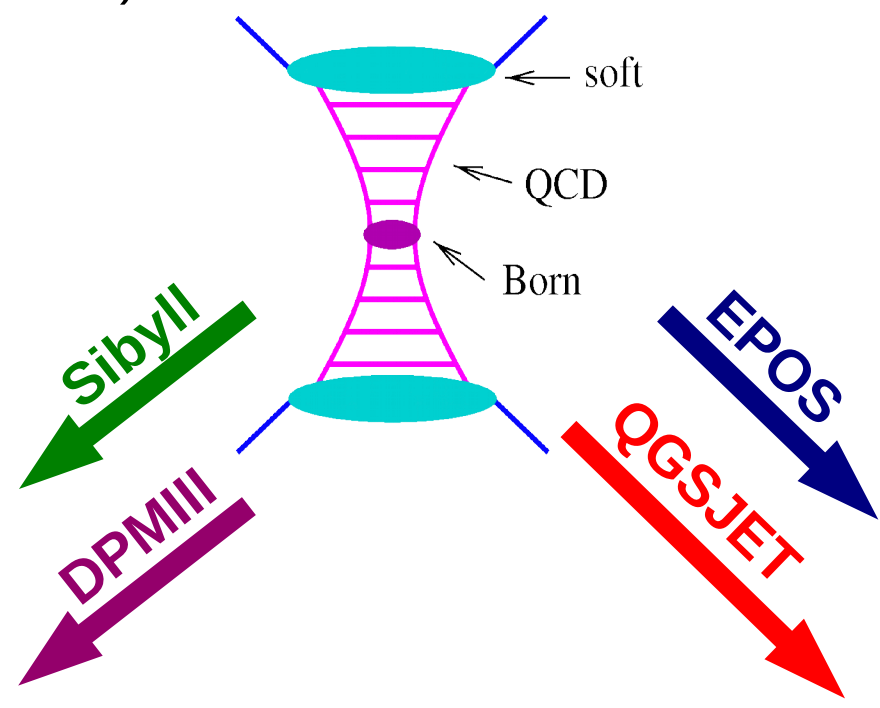




# Pseudorapidity

- **Field theory : scattering via the exchange of an excited field**
  - ➔ parton, hadron, quasi-particle = Reggeon or Pomeron (vacuum excitation)
- **QCD based theory so at high energy, perturbative QCD can be used to build the field amplitude (amplitude used for the cross-section)**
  - ➔ all minijet based (parton cascade and pQCD born process hadronized using string fragmentation) but different definitions

- ➔ soft+hard in different components
- ➔ external parton distribution functions (GRV98, cteq14)
- ➔ connection to projectile/target with small "x"

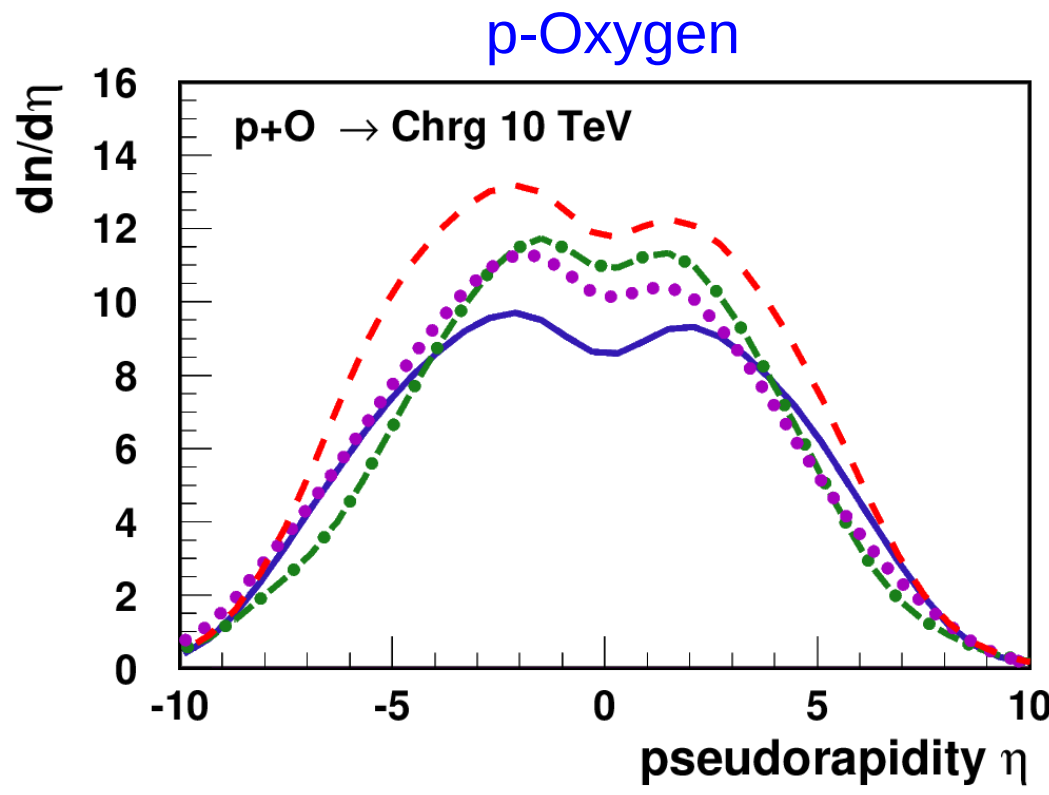
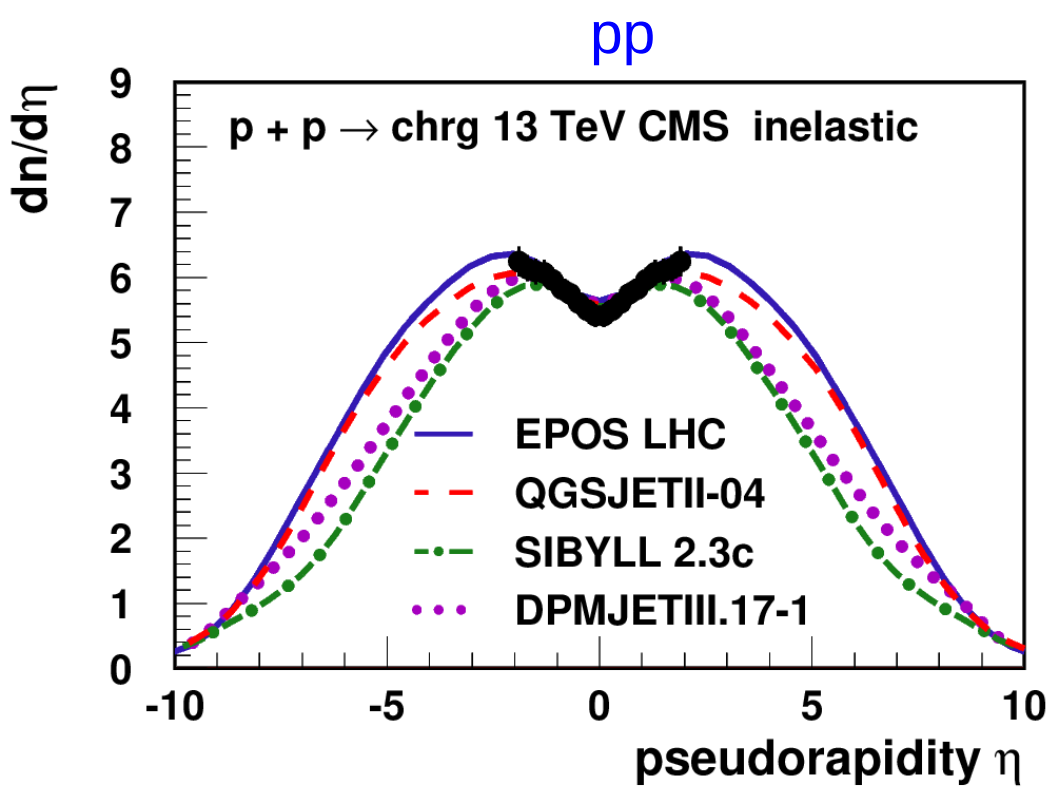


- ➔ soft+hard in the same amplitude
- ➔ own parton distribution function compatible with HERA data (not for QGSJET01: pre-HERA time)
- ➔ connection to projectile/target with large "x"

Ostapchenko et al. Phys.Rev. D94 (2016) no.11, 114026

# Pseudorapidity

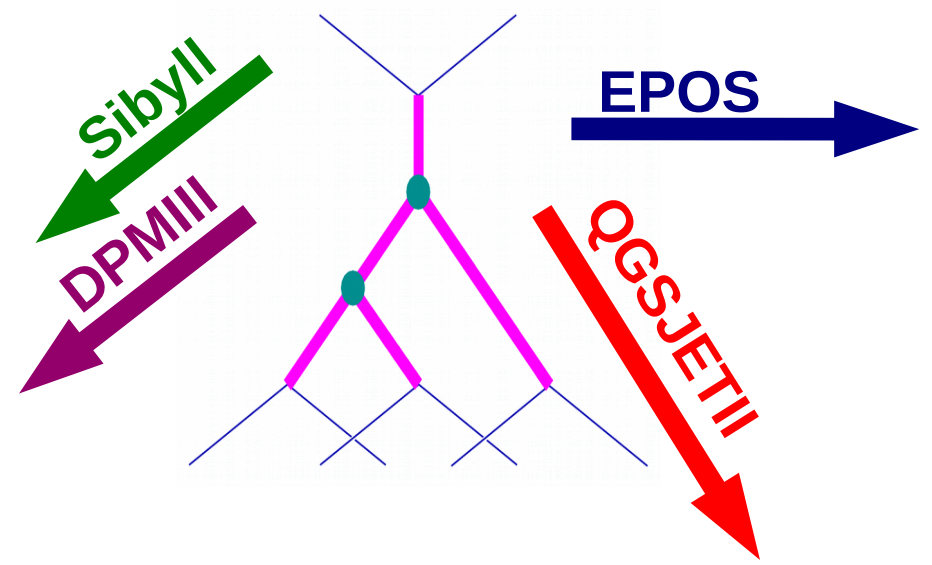
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  - ➔ all minijet based (parton cascade and pQCD born process **hadronized using string fragmentation**) but different definitions



# Energy Evolution

- Multiple scattering not enough to reconcile pQCD minijet cross-section and total cross-section
  - ➔ non-linear effect should be taken into account (interaction between scatterings)
- Solution depends on amplitude definition

- ➔ hard amplitude depend on minimum  $p_t$
- ➔ parametrize minimum  $p_t$  as a function of energy (and impact parameter for DPMJETIII)
- ➔ fit to data (multiplicity and cross-section)

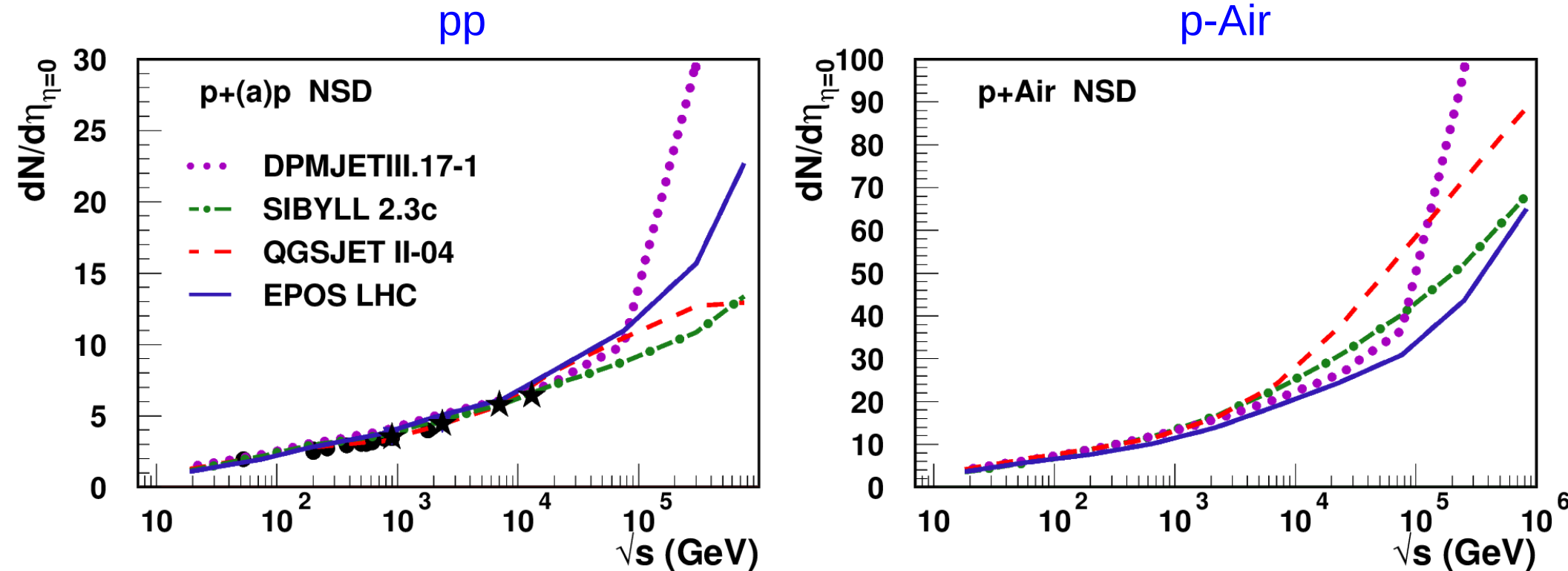


- ➔ fixed minimum  $p_t$  in hard part
- ➔ theory based “fan diagrams” re-summed to infinity without energy sharing

- ➔ fixed minimum  $p_t$  in hard part
- ➔ enhanced diagrams not compatible with energy sharing
- ➔ modification of vertex function to take into account non linear effects (data driven phenomenological approach)

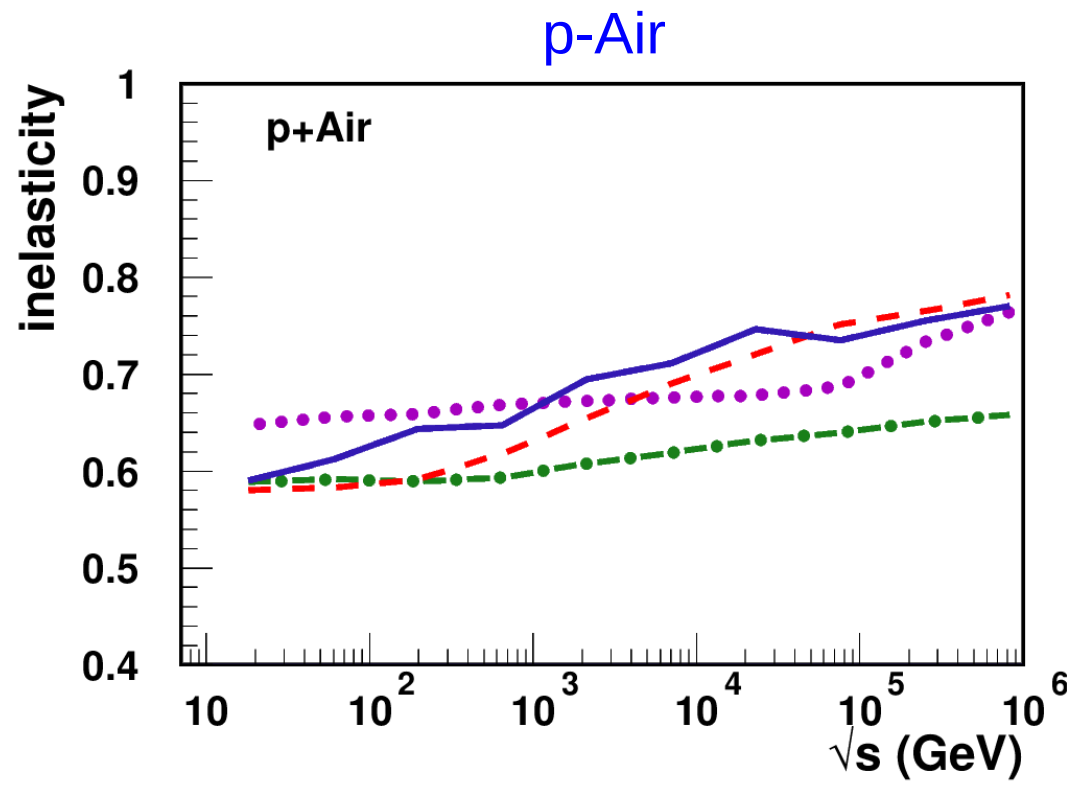
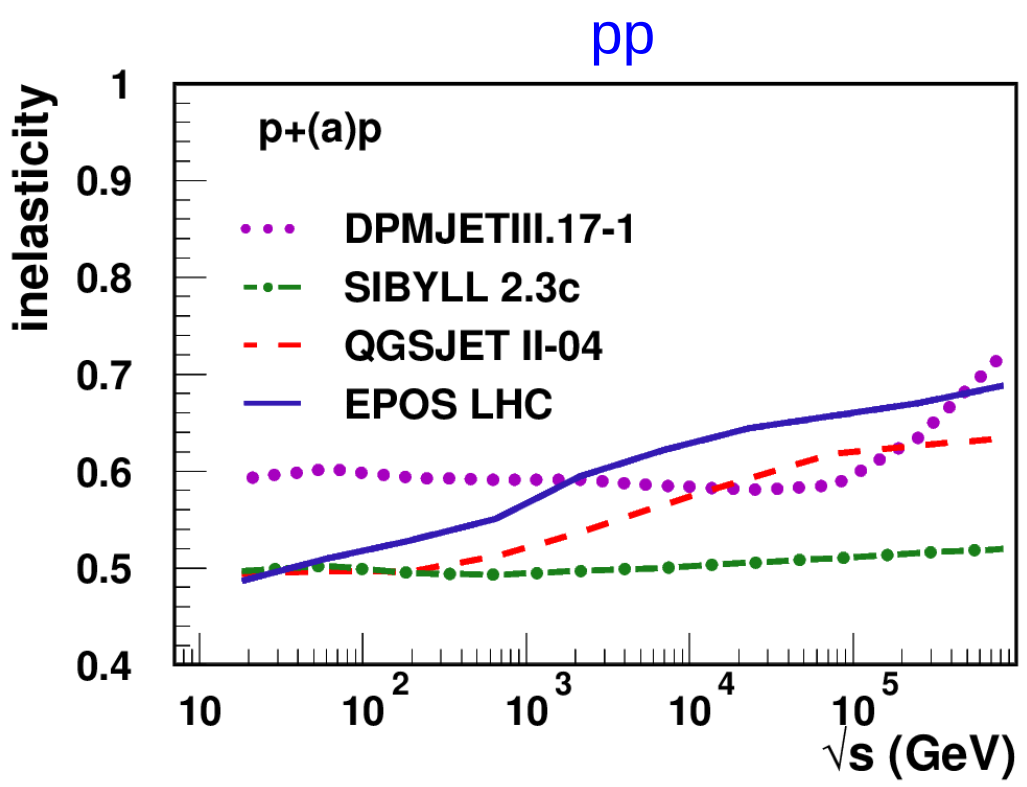
# Energy Evolution

- Multiple scattering not enough to reconcile pQCD minijet cross-section and total cross-section
  - ➔ non-linear effect should be taken into account (interaction between scatterings)
- Solution depends on amplitude definition
  - ➔ still large uncertainties at high energy (but reduced after LHC)



# Inelasticity

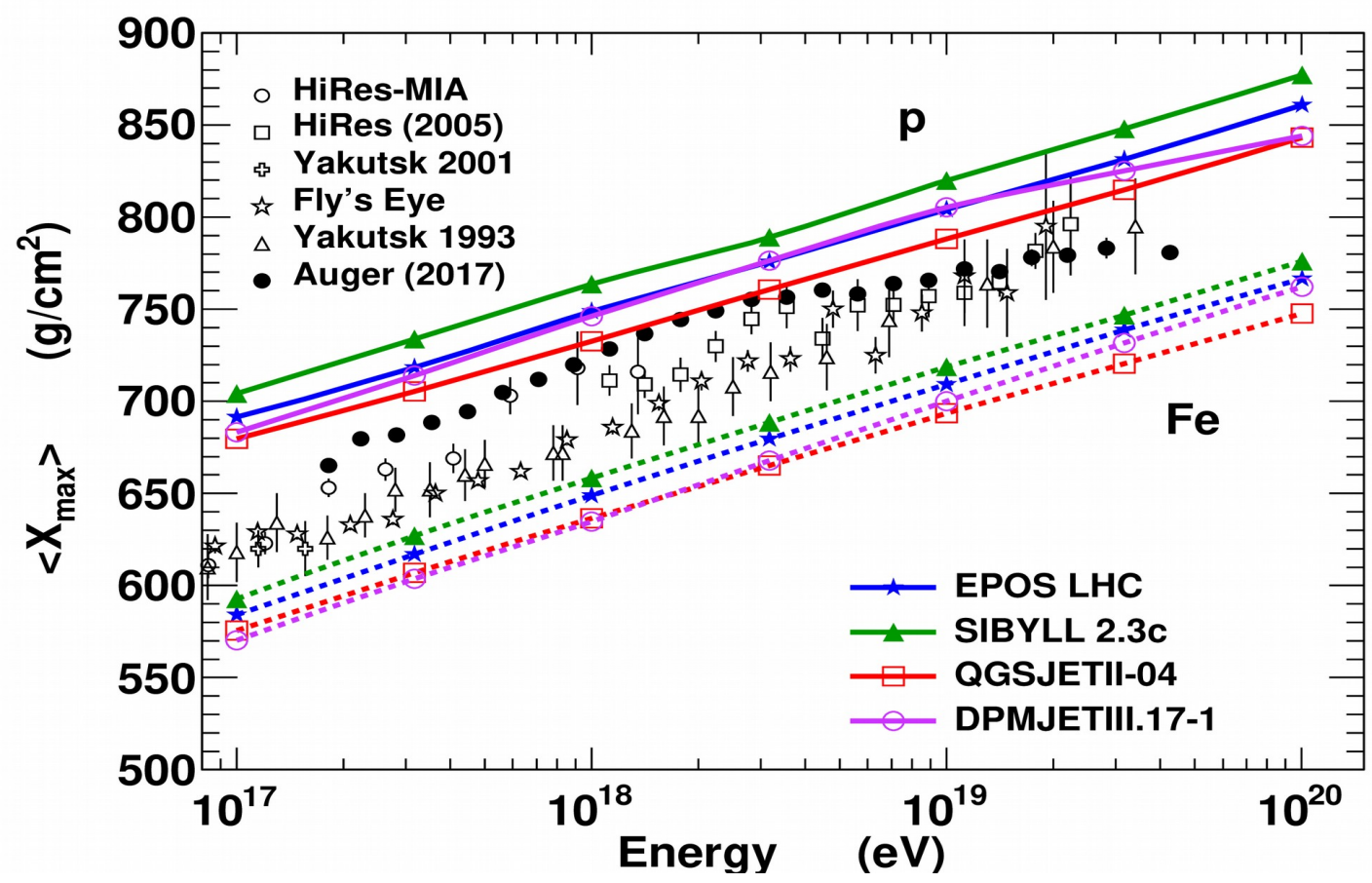
- In most of the cases, the projectile is destroyed by the collision
  - ➔ non-diffractive scattering : high energy loss for leading particle, high multiplicity
- In 10-20% of the time, the projectile have a small energy loss (high elasticity) and is unchanged
  - ➔ diffractive scattering : low energy loss, low multiplicity on target side
- Model difference mostly at technical level (and choice of data)



$$X_{\max}$$

**+/- 20g/cm<sup>2</sup> is a realistic uncertainty band but :**

- ➔ minimum given by QGSJETII-04 (high multiplicity, low elasticity)
- ➔ maximum given by Sibyll 2.3c (low multiplicity, high elasticity)
- ➔ anything below or above won't be compatible with LHC data



# WHISP Working Group

- **Much more measurement available**

- ➔ Auger, EAS-MSU, KASCADE-Grande, IceCube/IceTop, HiRes-MIA, NEMOD/DECOR, SUGAR, TA, Yukutsk

- **Working group (WHISP) created to compile all results together. Analysis led and presented on behalf of all collaborations**

- by **H. Dembinski** at **UHECR 2018** :

- H. Dembinski** (LHCb, Germany),

- L. Cazon** (Auger, Portugal), **R. Conceicao** (AUGER, Portugal),  
**F. Riehn** (Auger, Portugal), **T. Pierog** (Auger, Germany),

- Y. Zhezher** (TA, Russia), **G. Thomson** (TA, USA) , **S. Troitsky** (TA, Russia), **R. Takeishi** (TA, USA),

- T. Sako** (LHCf & TA, Japan), **Y. Itow** (LHCf, Japan),

- J. Gonzales** (IceTop, USA), **D. Soldin** (IceCube, USA),

- J.C. Arteaga** (KASCADE-Grande, Mexico),

- I. Yashin** (NEMOD/DECOR, Russia). **E. Zadeba**  
(NEMOD/DECOR, Russia)

- N. Kalmykov** (EAS-MSU, Russia) and **I.S. Karpikov** (EAS-MSU, Russia)

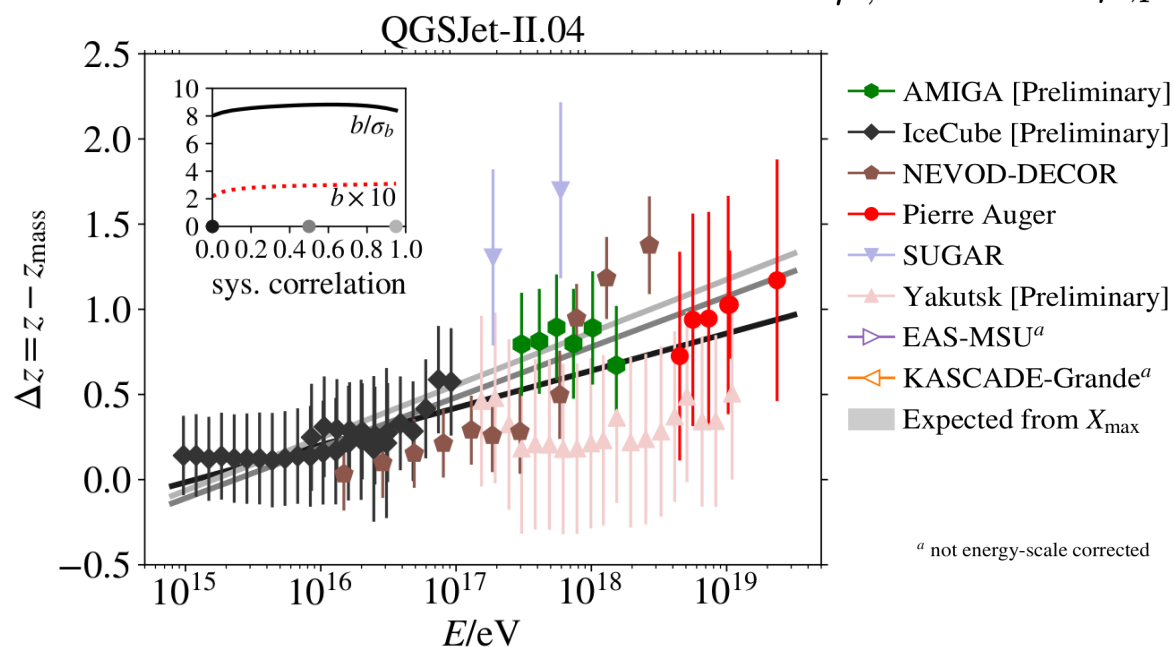
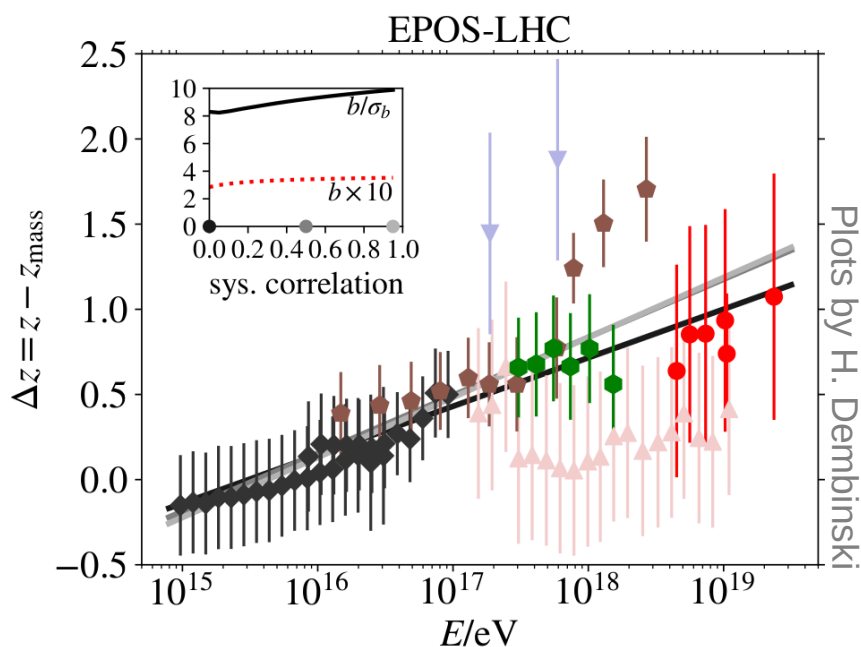
# Global Behavior

## ● Clear muon excess in data compared to simulation

➔ Different energy evolution between data and simulations

➔ Significant non-zero slope ( $>8\sigma$ )

$$z = \frac{\ln N_{\mu}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}{\ln N_{\mu,\text{Fe}}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}$$



## ● Different energy or mass scale cannot change the slope

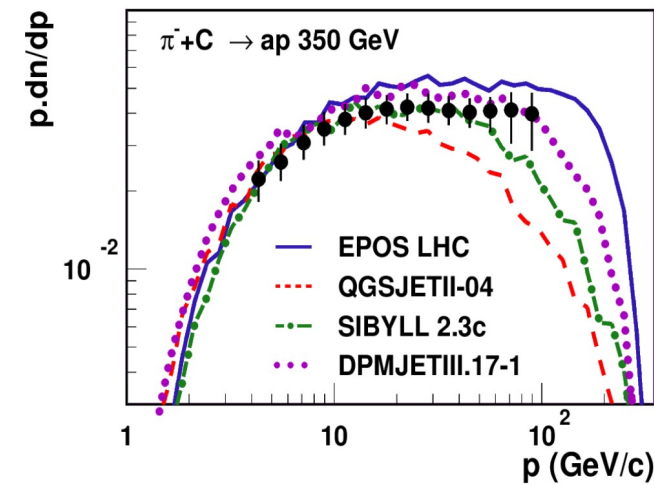
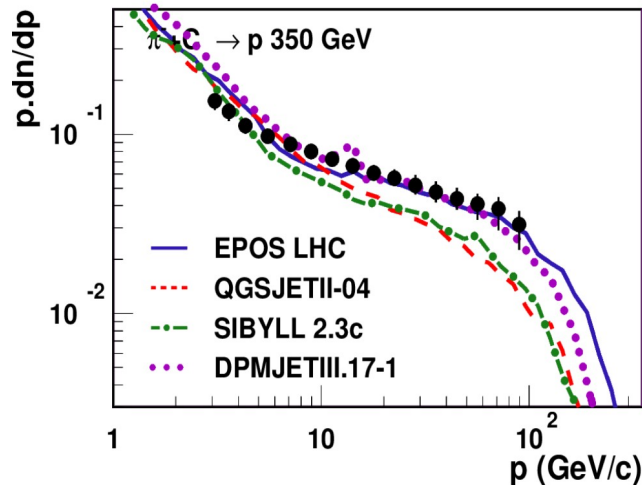
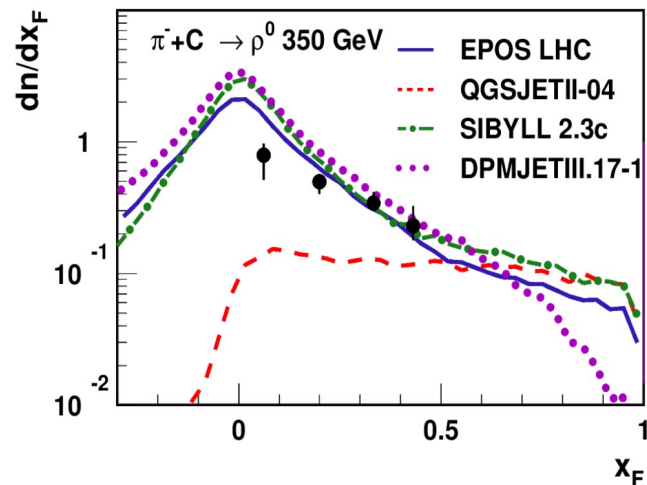
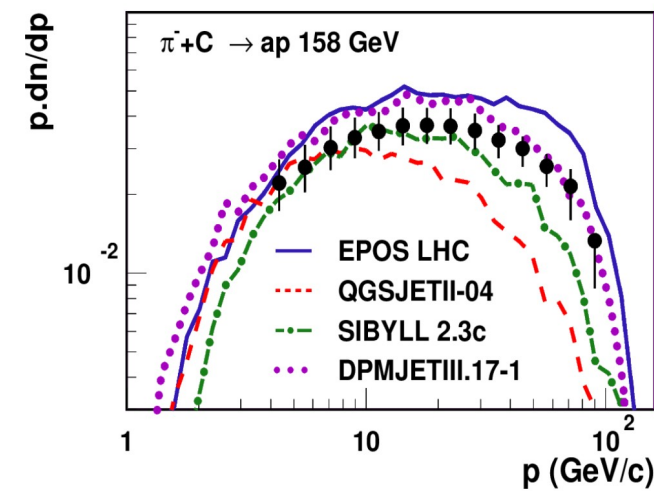
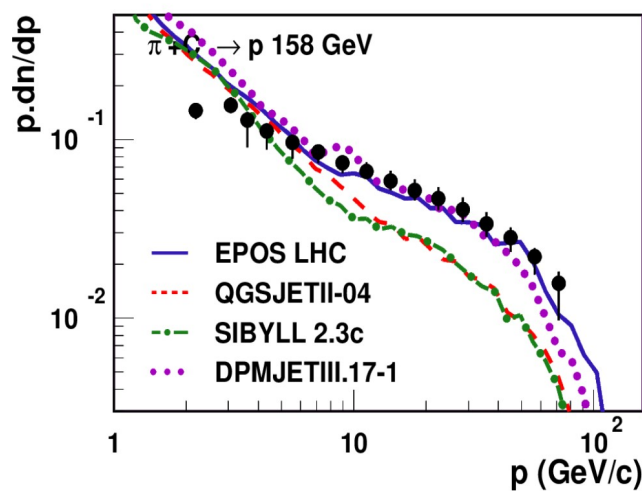
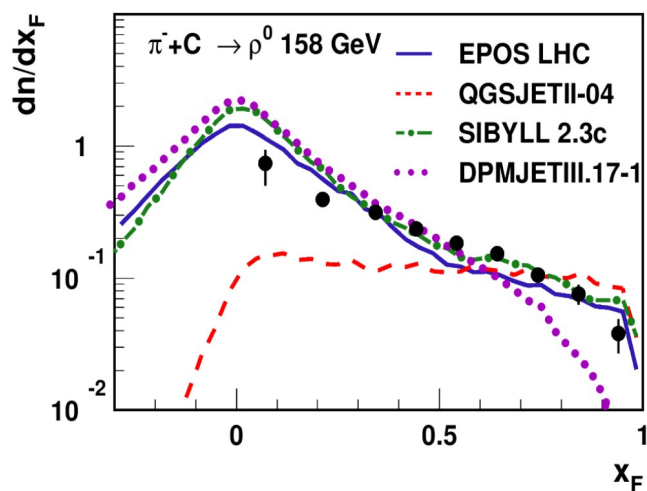
➔ Different property of hadronic interactions at least above  $10^{16}$  eV



# NA61 Pion-Carbon Data

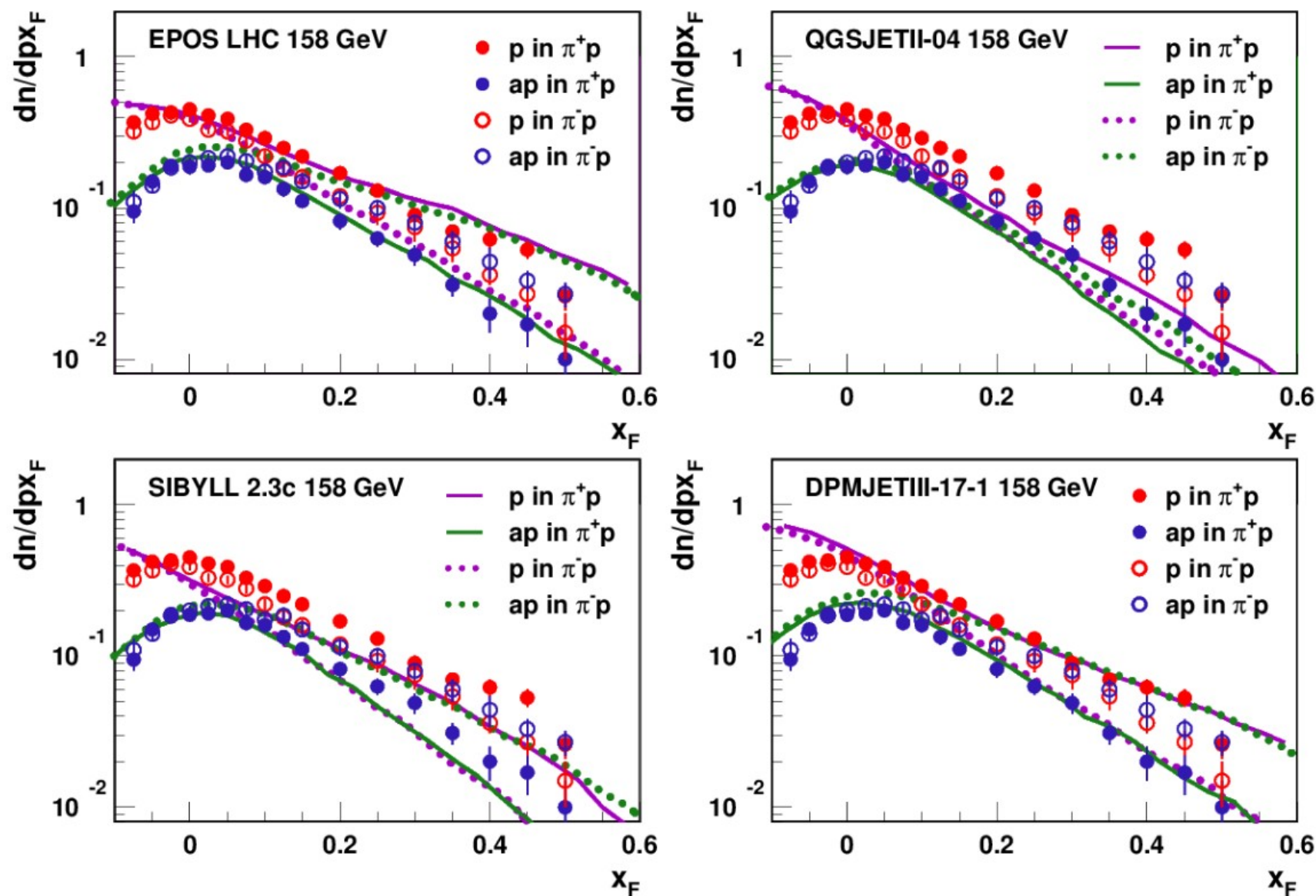
## New data from NA61 : wrong old data interpretation

- ➔ over production of anti-baryons in EPOS LHC : problem in air showers
- ➔ confirmation that QGSJETII-04 underestimate forward baryon production



# Baryons in Pion Interactions

Data from NA49 (Gabor Veres PhD) : full picture



# Constraints from Correlated Change

- One needs to change energy dependence of muon production by  $\sim +4\%$

$$N_{\mu} = A^{1-\beta} \left( \frac{E}{E_0} \right)^{\beta}$$

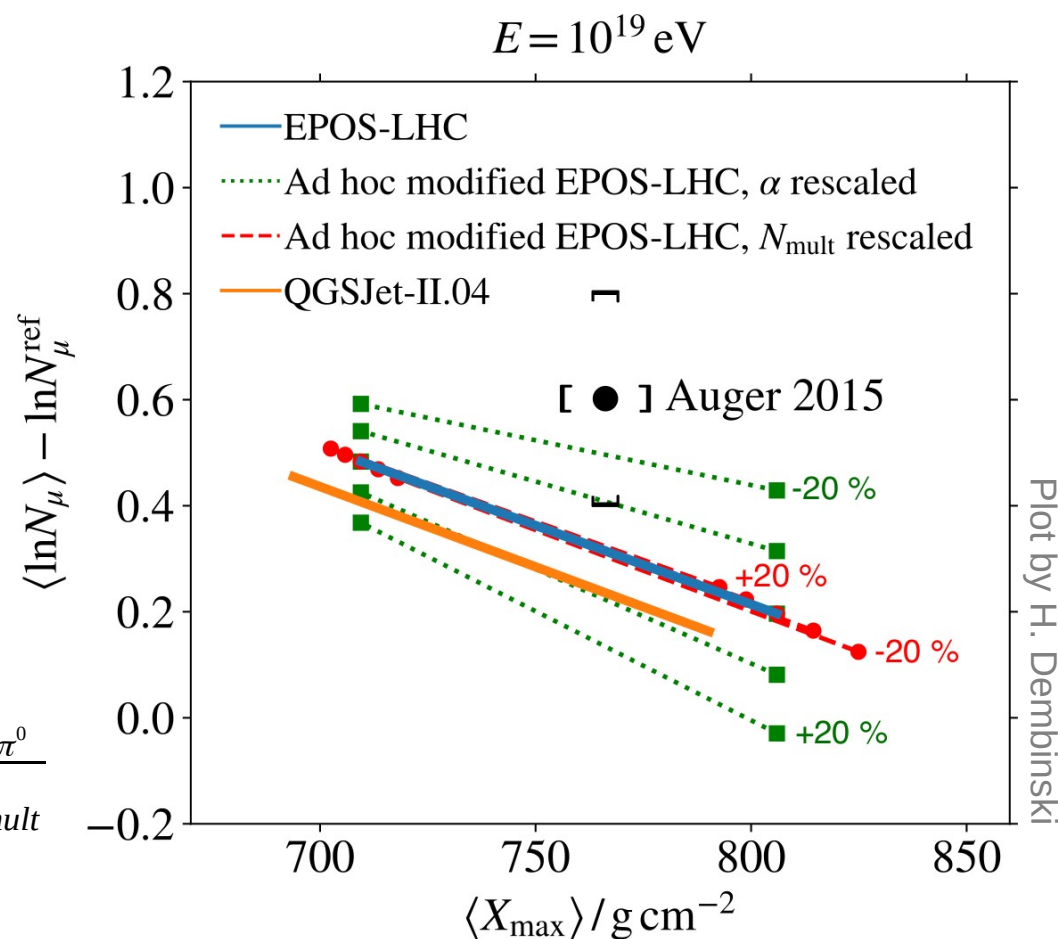
- To reduce muon discrepancy  $\beta$  has to be change

→  $X_{\max}$  alone (composition) will not change the energy evolution

→  $\beta$  changes the muon energy evolution but not  $X_{\max}$

$$\beta = \frac{\ln(N_{\text{mult}} - N_{\pi^0})}{\ln(N_{\text{mult}})} = 1 + \frac{\ln(1 - \alpha)}{\ln(N_{\text{mult}})}$$

→  $+4\%$  for  $\beta$  →  $-30\%$  for  $\alpha = \frac{N_{\pi^0}}{N_{\text{mult}}}$



# Possible Particle Physics Explanations

A 30% change in particle charge ratio ( $\alpha = \frac{N_{\pi^0}}{N_{mult}}$ ) is huge !

→ Possibility to increase  $N_{mult}$  limited by  $X_{max}$

→ New Physics ?

- Chiral symmetry restoration (Farrar et al.) ?

- Strange fireball (Anchordoqui et al.) ?

- String Fusion (Alvarez-Muniz et al.) ?

→ Problem : no strong effect observed at LHC ( $\sim 10^{17}$  eV)

→ Unexpected production of Quark Gluon Plasma (QGP) in light systems observed at the LHC (at least modified hadronization)

- Reduced  $\alpha$  is a sign of QGP formation (Baur et al.) !

- Not properly done in EPOS LHC (QGP only in extreme conditions)

→ Problem :  $\alpha$  changed at most by 20-25%

# Should Everything Be Taken into Account in CR Models ?

## ● Models have different philosophies !

- ➔ number of parameters increase with data set to reproduce
- ➔ predictive power may decrease with number of parameters
- ➔ predictive power increase if we are sure not to neglect something

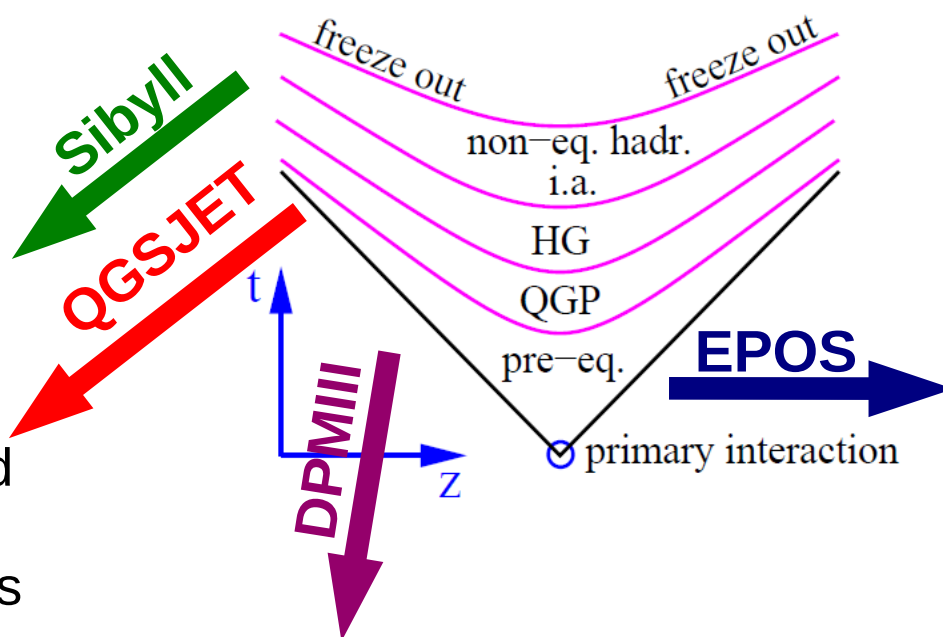
➔ models for CR only

➔ fast and not suppose to describe everything

➔ no detailed hard scattering or collective effects

➔ heavy ion model intended to be used for high energy physics

➔ limited development for collective effects but correct hard scattering



➔ developed first for heavy ion interactions

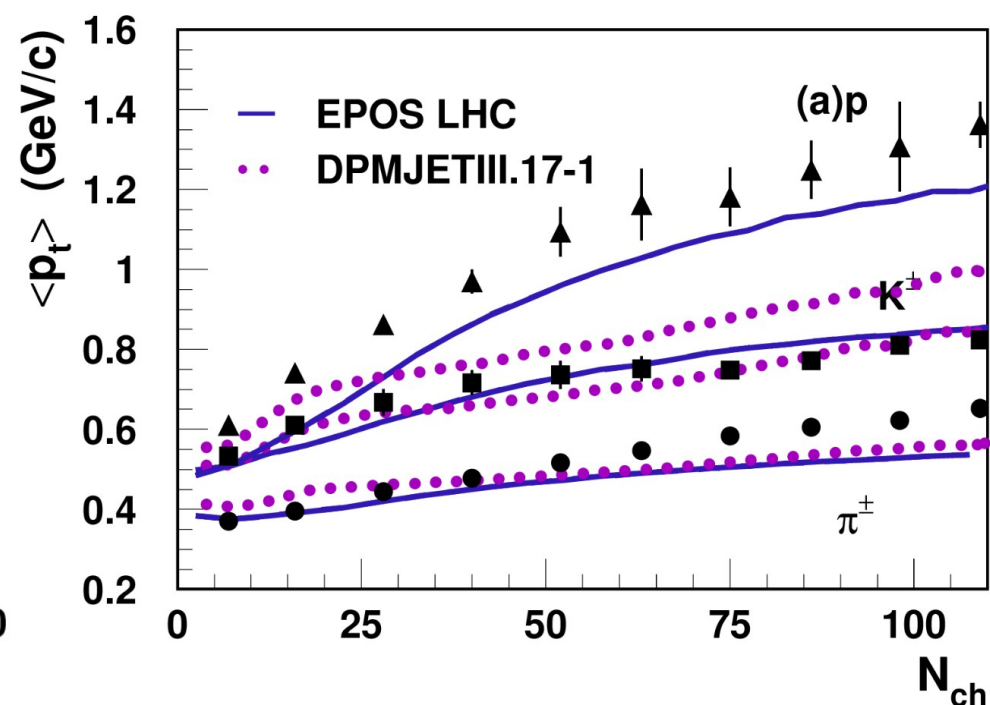
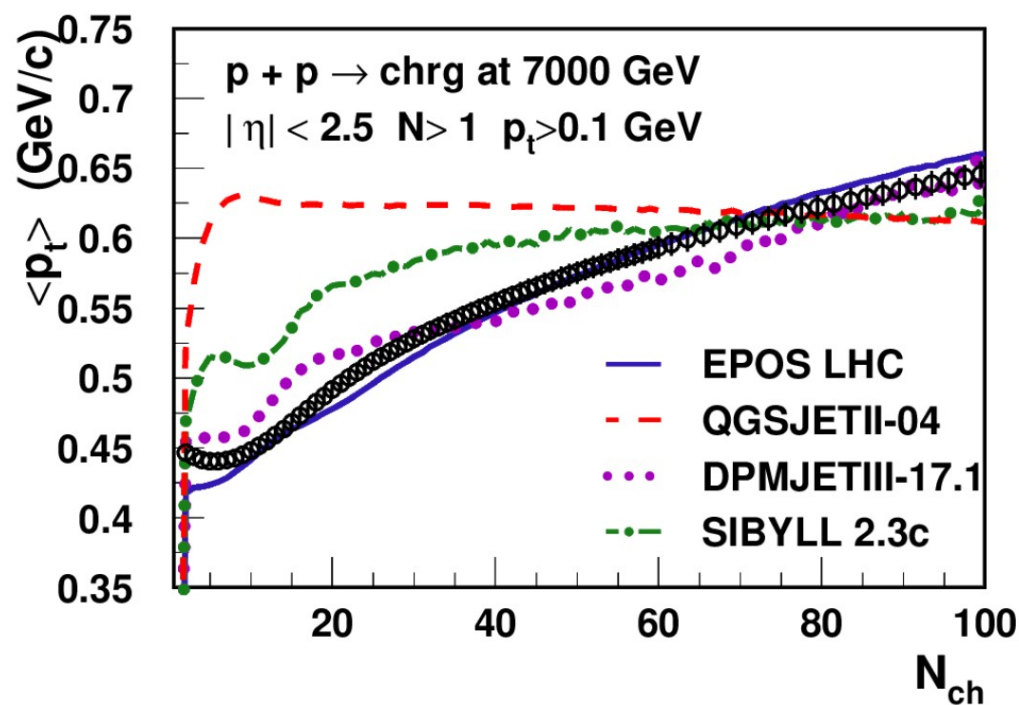
➔ detailed description of every possible "soft" observable (not good for hard scattering yet)

➔ sophisticated collective effect treatment (real hydro for EPOS 2 and 3)

➔ very large complete data set (LEP, HERA, SPS, RHIC, LHC)

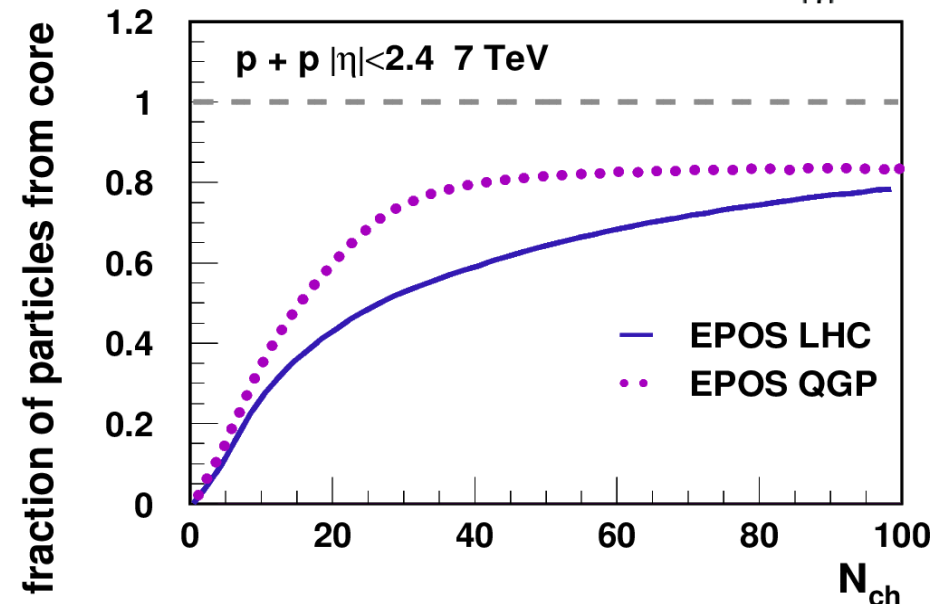
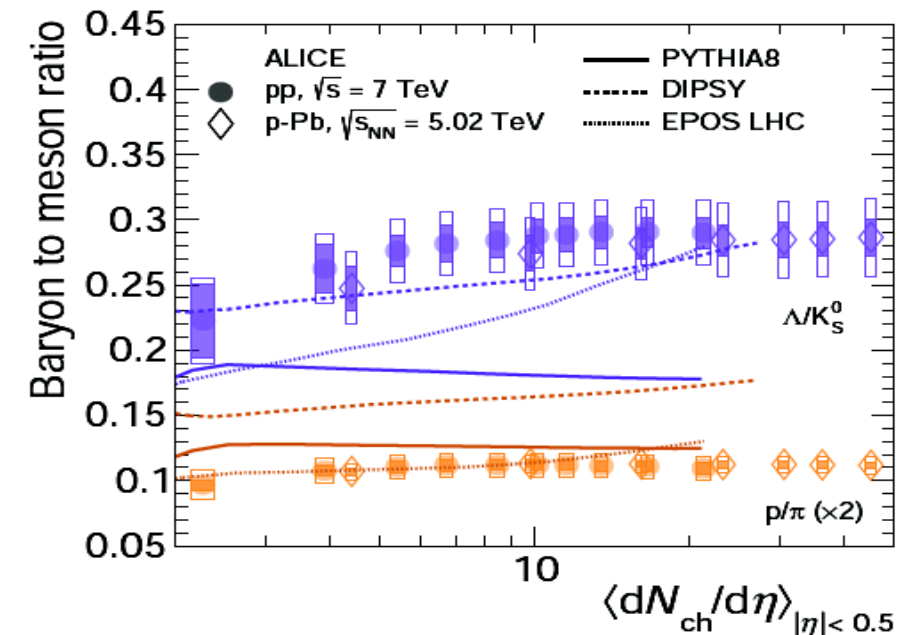
# Should Everything Be Taken into Account in CR Models ?

- **Models have different philosophies !**
  - ➔ number of parameters increase with data set to reproduce
  - ➔ predictive power may decrease with number of parameters
  - ➔ predictive power increase if we are sure not to neglect something
- **No direct influence on air showers but different parameters and extrapolations ?**

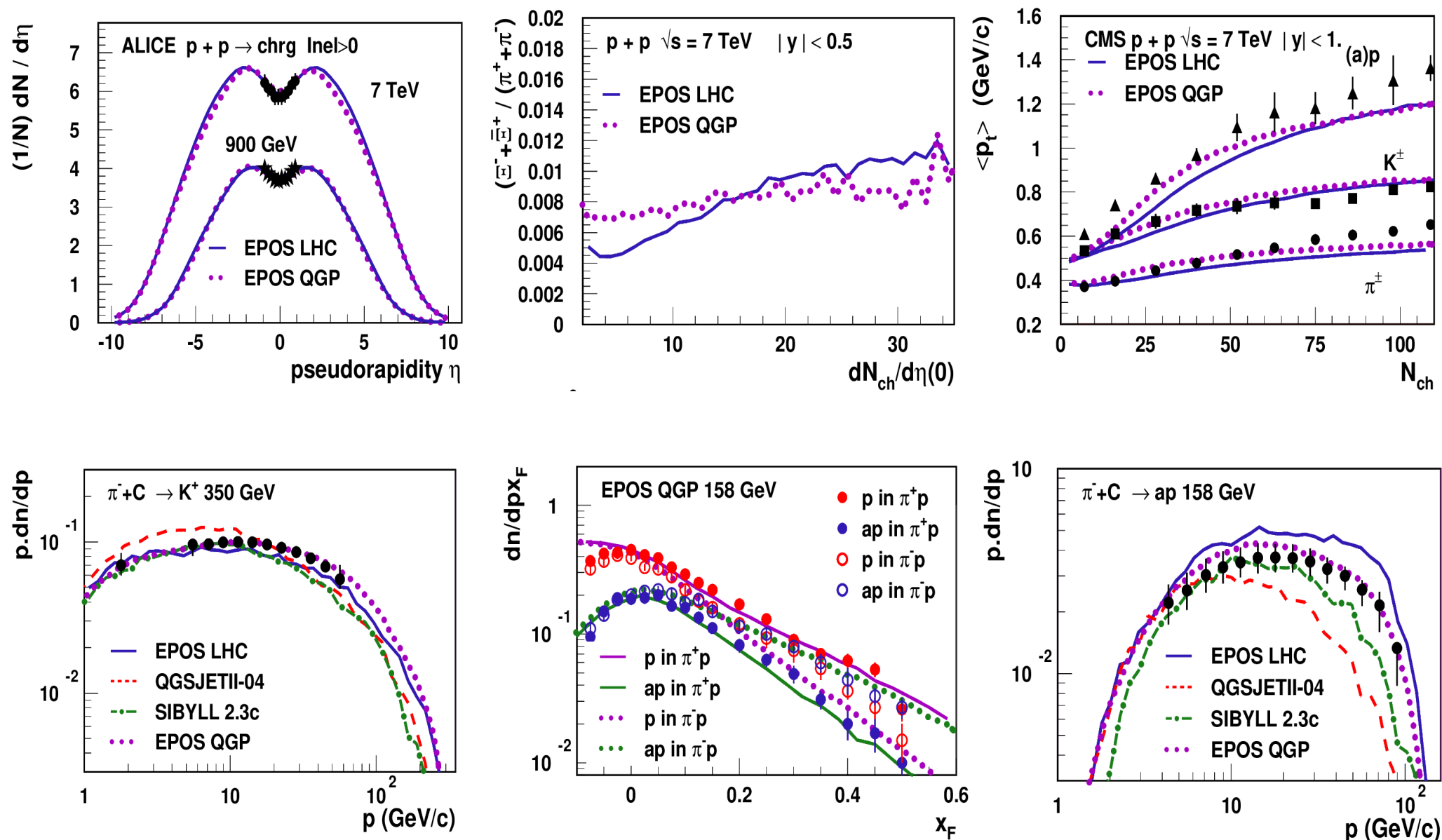


# Modified EPOS with Extended Core

- **Core in EPOS LHC appear too late**
  - ➔ Recent publication show the evolution of chemical composition as a function of multiplicity
  - ➔ Large amount of (multi)strange baryons produced at lower multiplicity than predicted by EPOS LHC
- **Create a new version EPOS QGP with more collective hadronization**
  - ➔ Core created at lower energy density
  - ➔ More remnant hadronized with collective hadronization
  - ➔ Collective hadronization using grand canonical ensemble instead of microcanonical (closer to statistical decay)



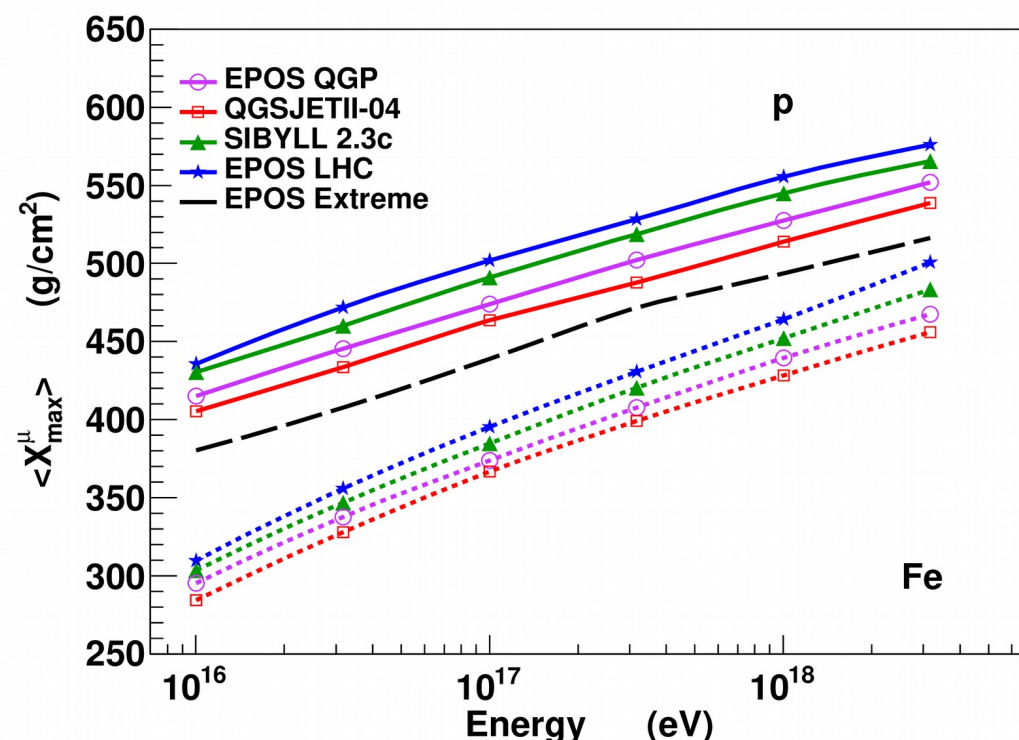
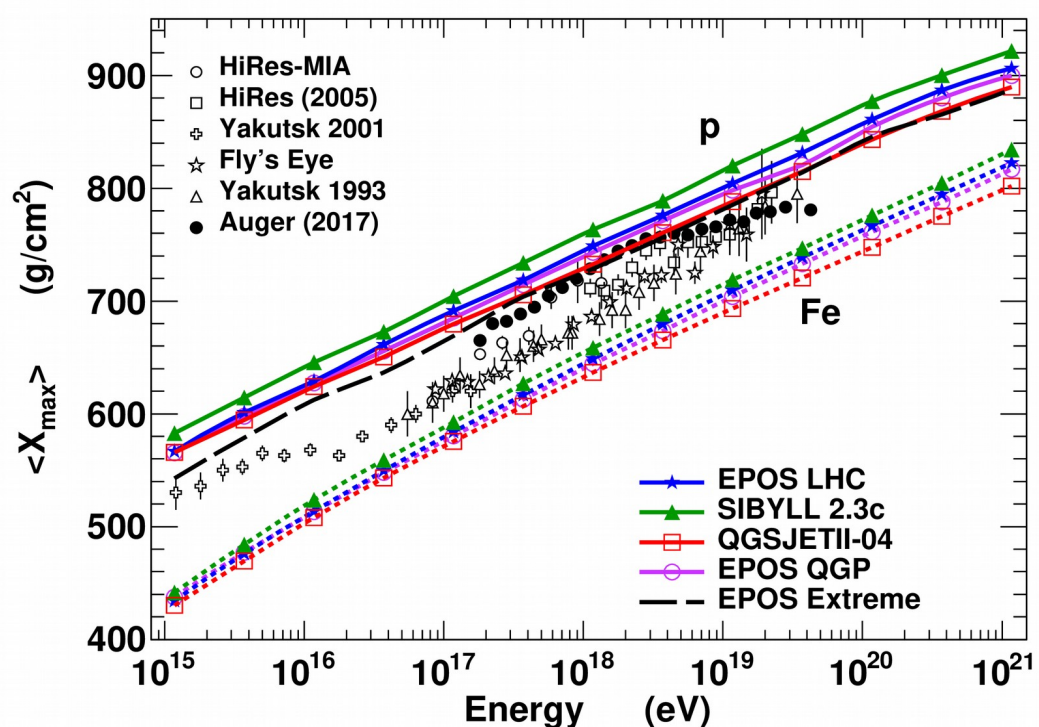
# Preliminary Version with Minimum Constraints





# Results for Air Showers (1)

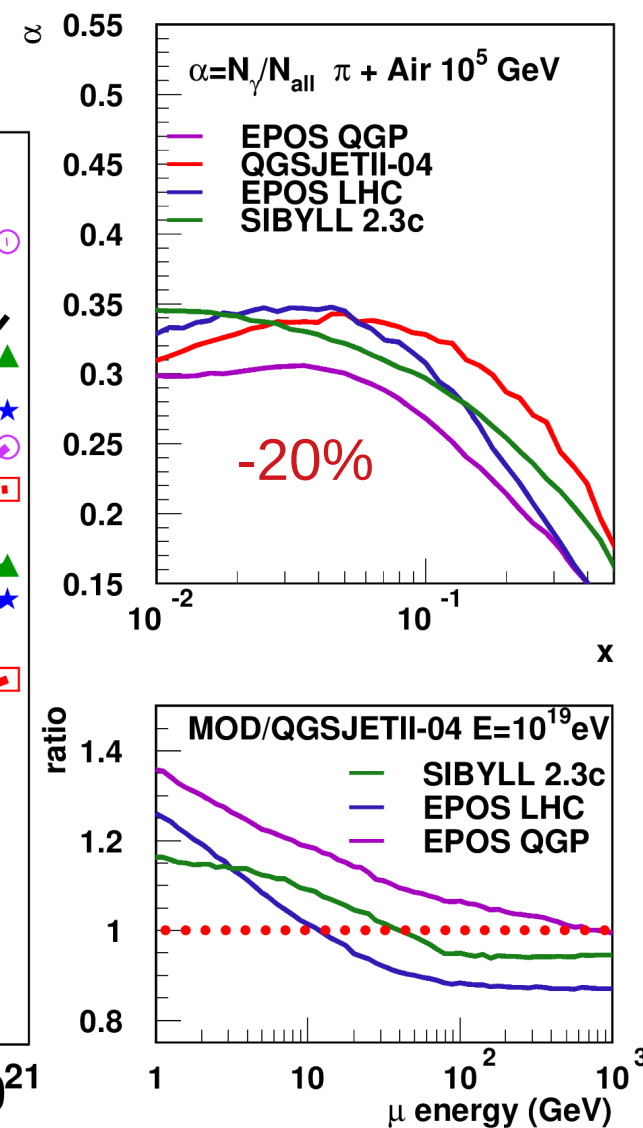
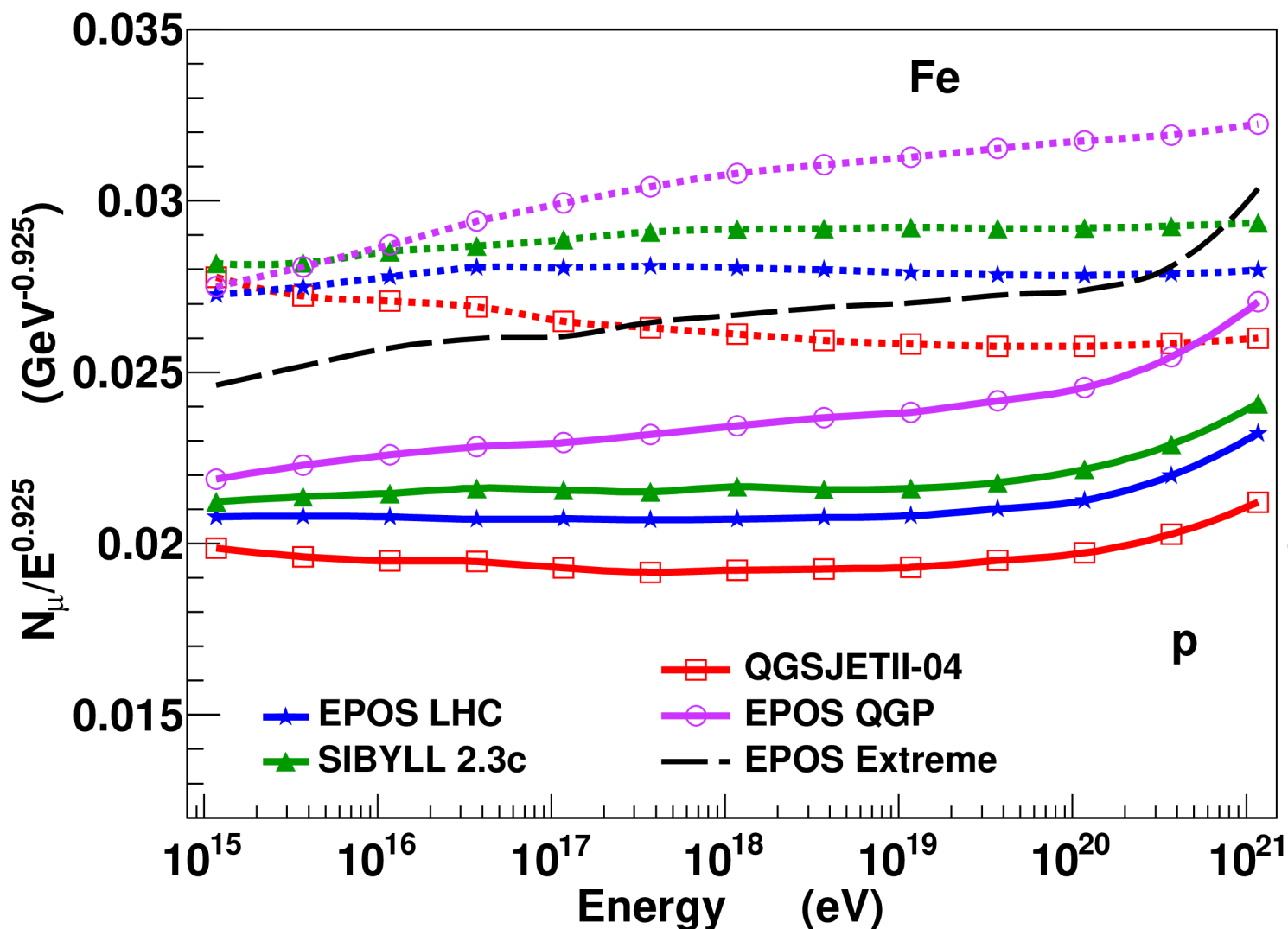
- Small change for  $\langle X_{\max} \rangle$  as expected
- Significant change of  $\langle X_{\max}^{\mu} \rangle$
- Comparison with extreme case (almost only grand canonical hadron.)
  - ➔ maximum effect using this approach
  - ➔ not compatible with accelerator data



# Results for Air Showers (2)

- Large change of the number of muons at ground

➔ Different slope as expected from the change in  $\alpha$



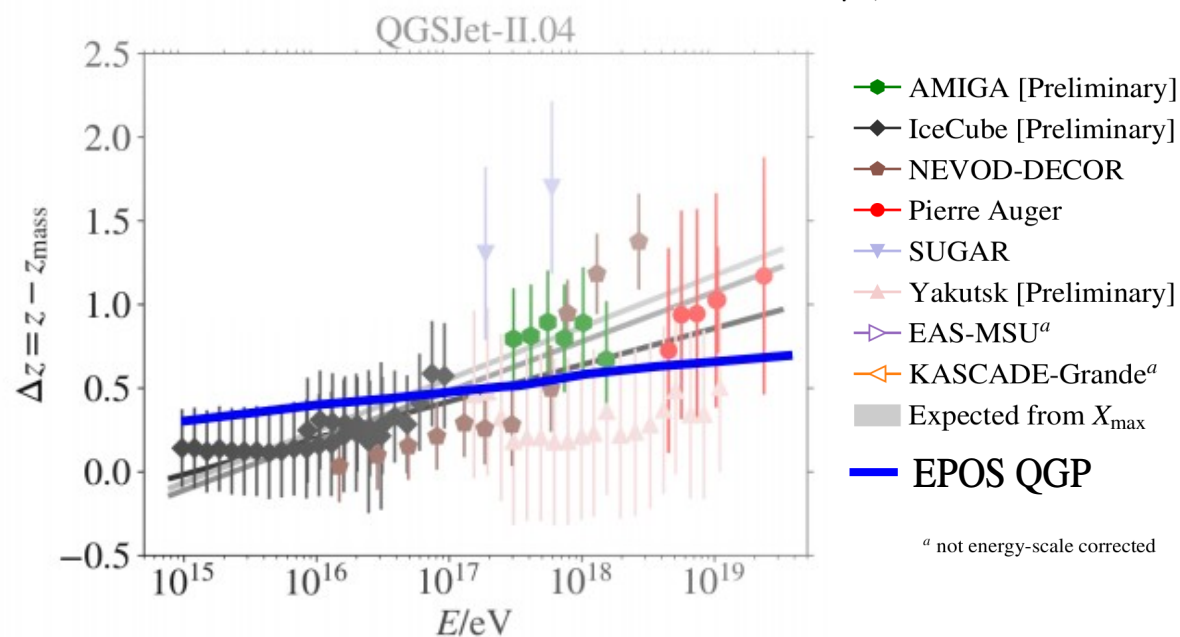
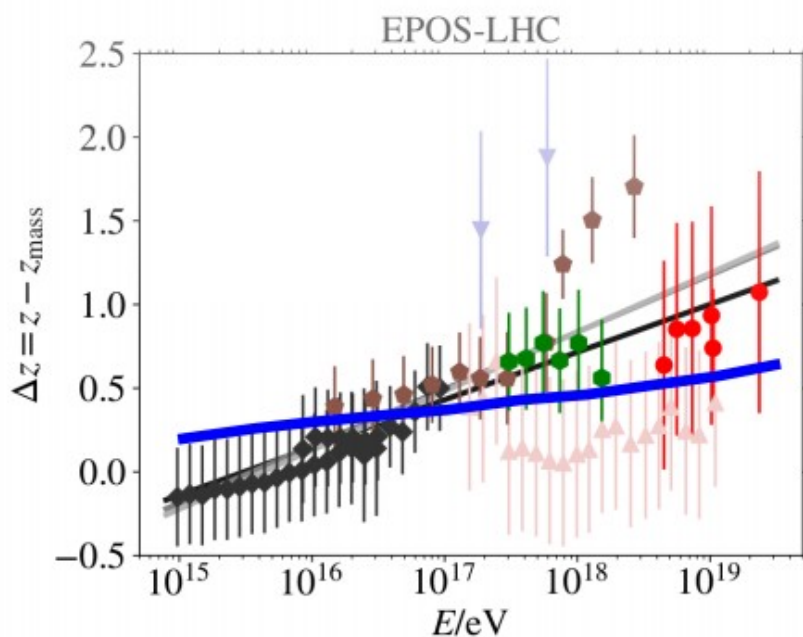
# Comparison with Data

- **Collective hadronization gives a result compatible with data**

➔ Still different energy evolution between data and simulation

➔ **Significance to be tested**

$$z = \frac{\ln N_{\mu}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}{\ln N_{\mu,\text{Fe}}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}$$



- AMIGA [Preliminary]
- IceCube [Preliminary]
- NEVOD-DECOR
- Pierre Auger
- SUGAR
- Yakutsk [Preliminary]
- EAS-MSU<sup>a</sup>
- KASCADE-Grande<sup>a</sup>
- Expected from  $X_{\text{max}}$
- EPOS QGP

<sup>a</sup> not energy-scale corrected

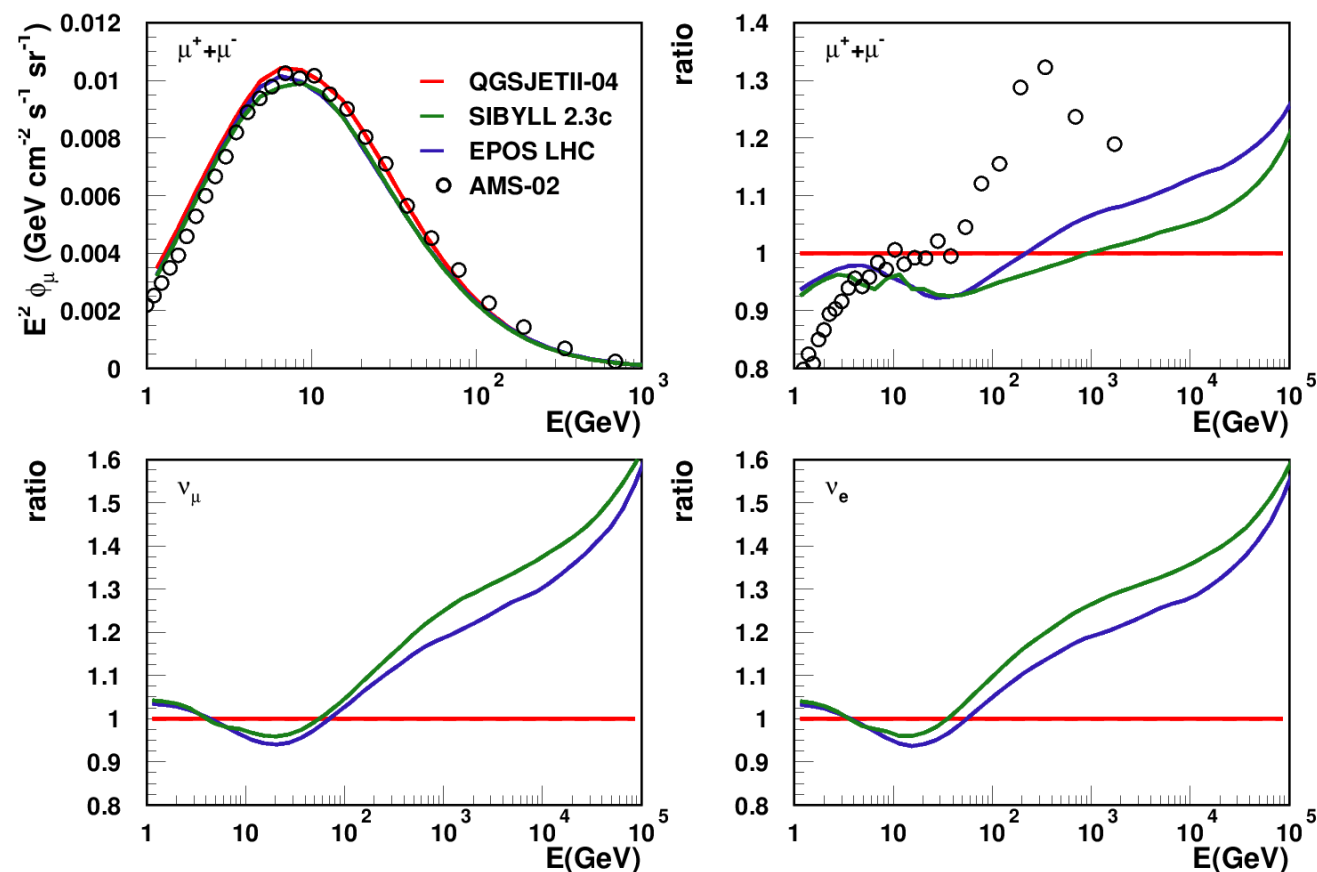
- **Probably tension at low energy (too many muons)**

➔ Ideally a larger slope would be needed ... what kind of hadronization possible ?

# Muon and Neutrino Fluxes

Low energy inclusive muon flux compared to predictions from different models (MCEq)

- ➔ Reasonable agreement below 100 GeV.
- ➔ Uncertainties due to primary CR flux/mass choice (H3a)



# Inclusive Spectra and First Interaction

For inclusive spectra, particles from first interaction dominate

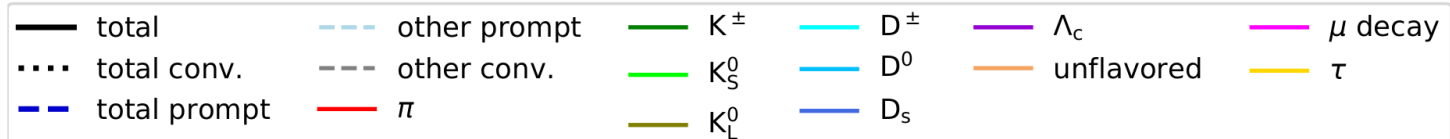
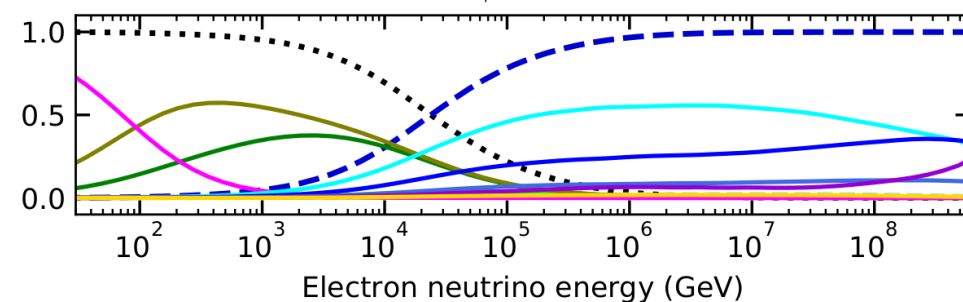
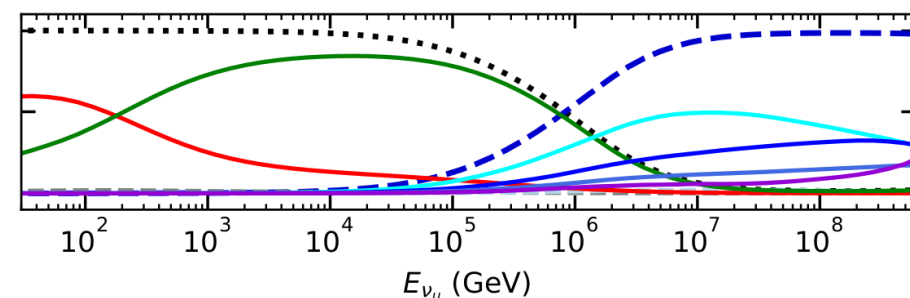
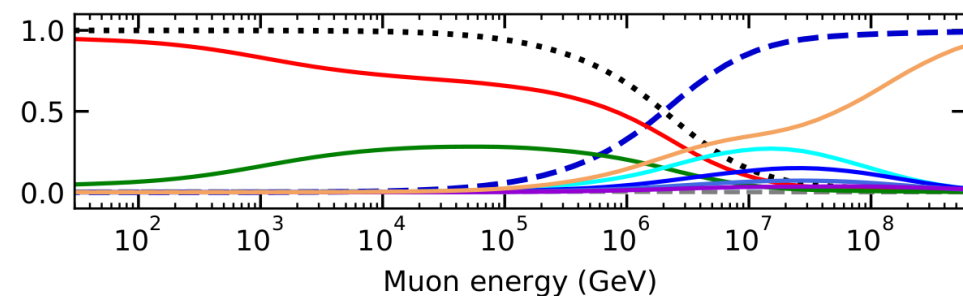
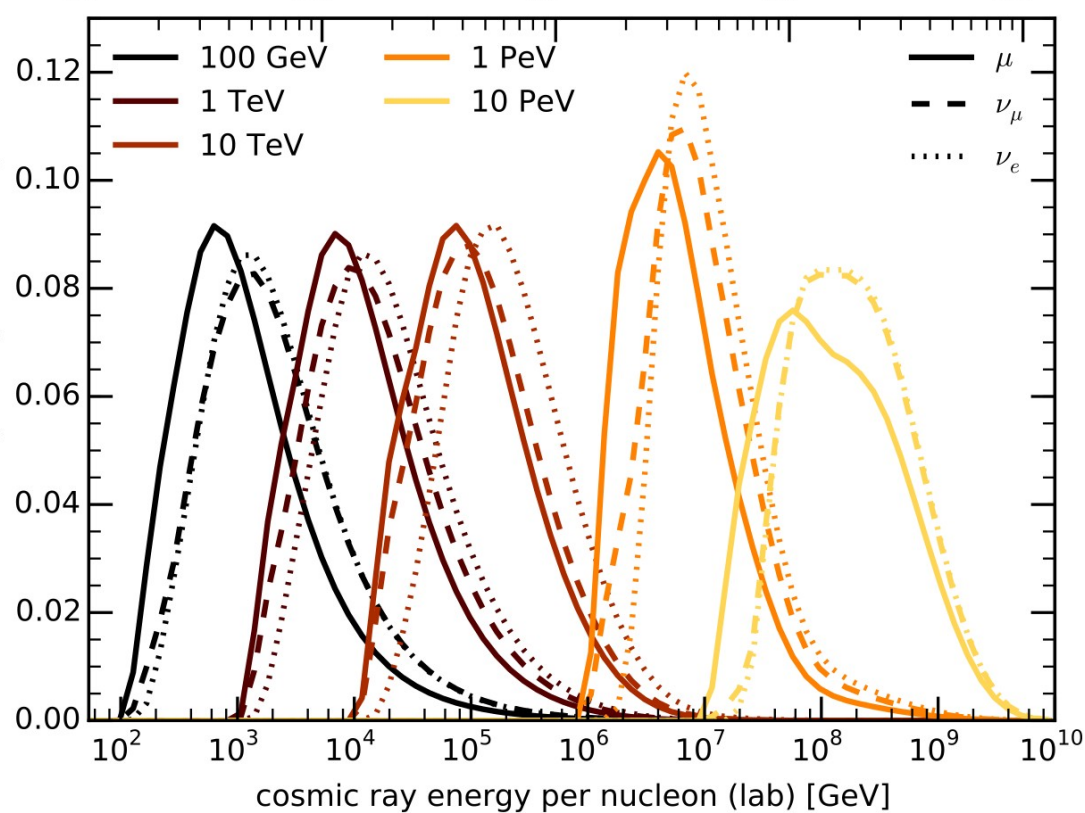
Fixed-target

RHIC

Tevatron

LHC

center of mass energy [GeV]

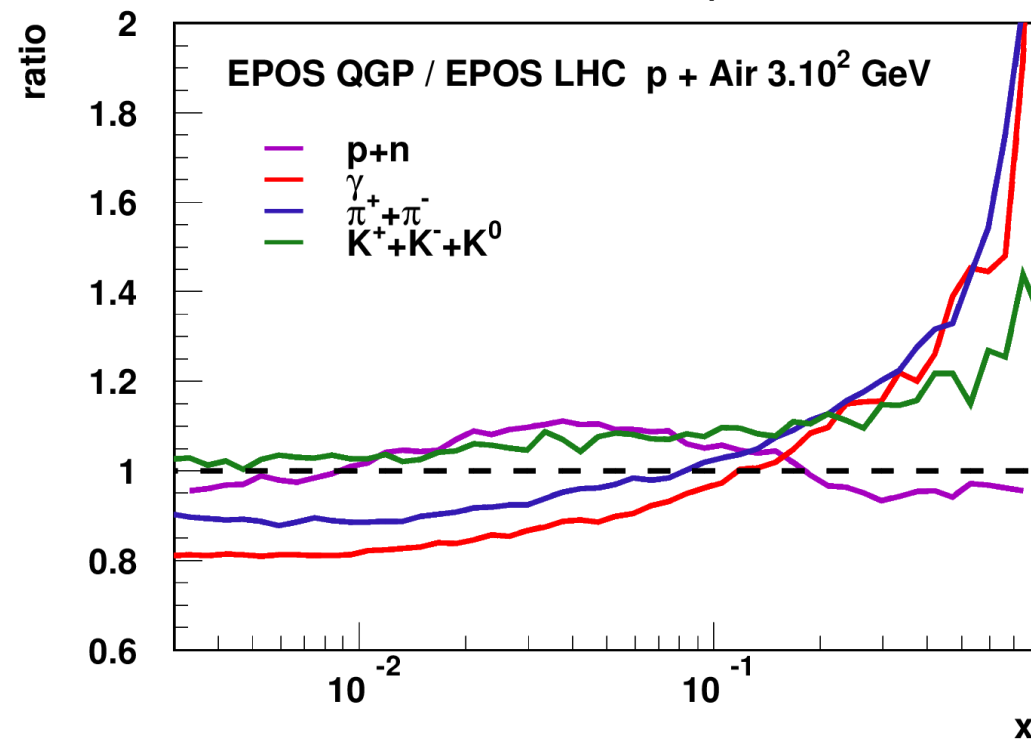
 $10^1$        $10^2$        $10^3$        $10^4$        $10^5$ 


Plots from A. Fedynytsch

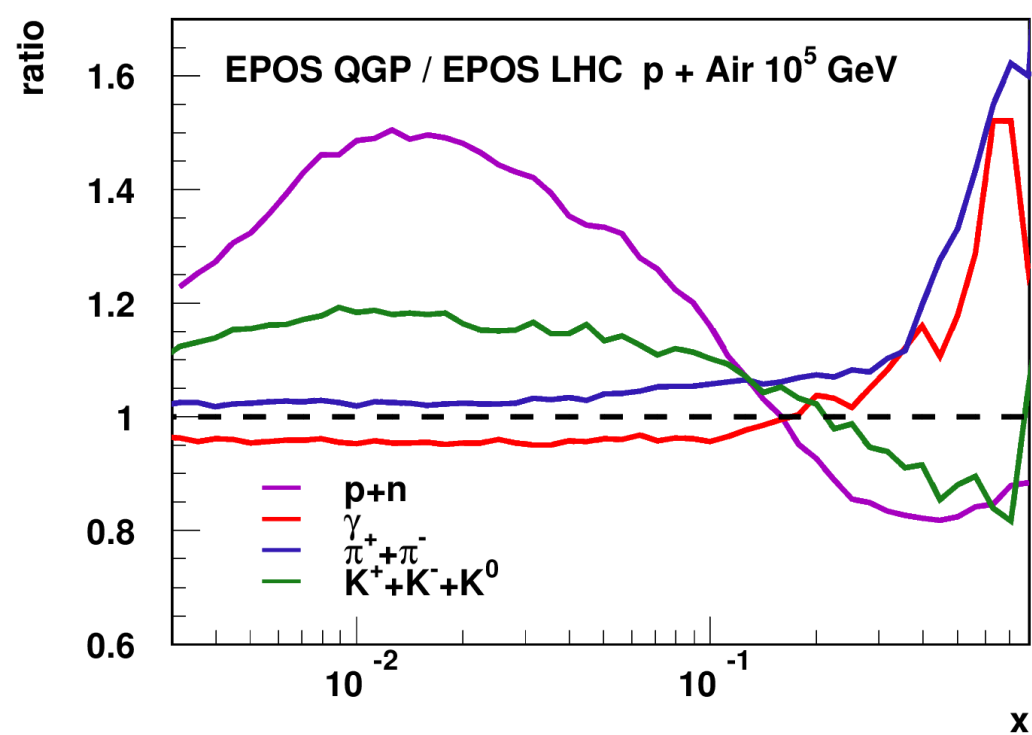
# Modified Spectra with EPOS QGP

- Muons above 100 GeV and neutrinos very sensitive to kaon production
  - ➔ Kaon production increased by up to 20% in EPOS QGP
- Collective hadronization will change inclusive fluxes
  - ➔ **Additional constrain to take into account !**

Source of TeV leptons



Source of PeV leptons

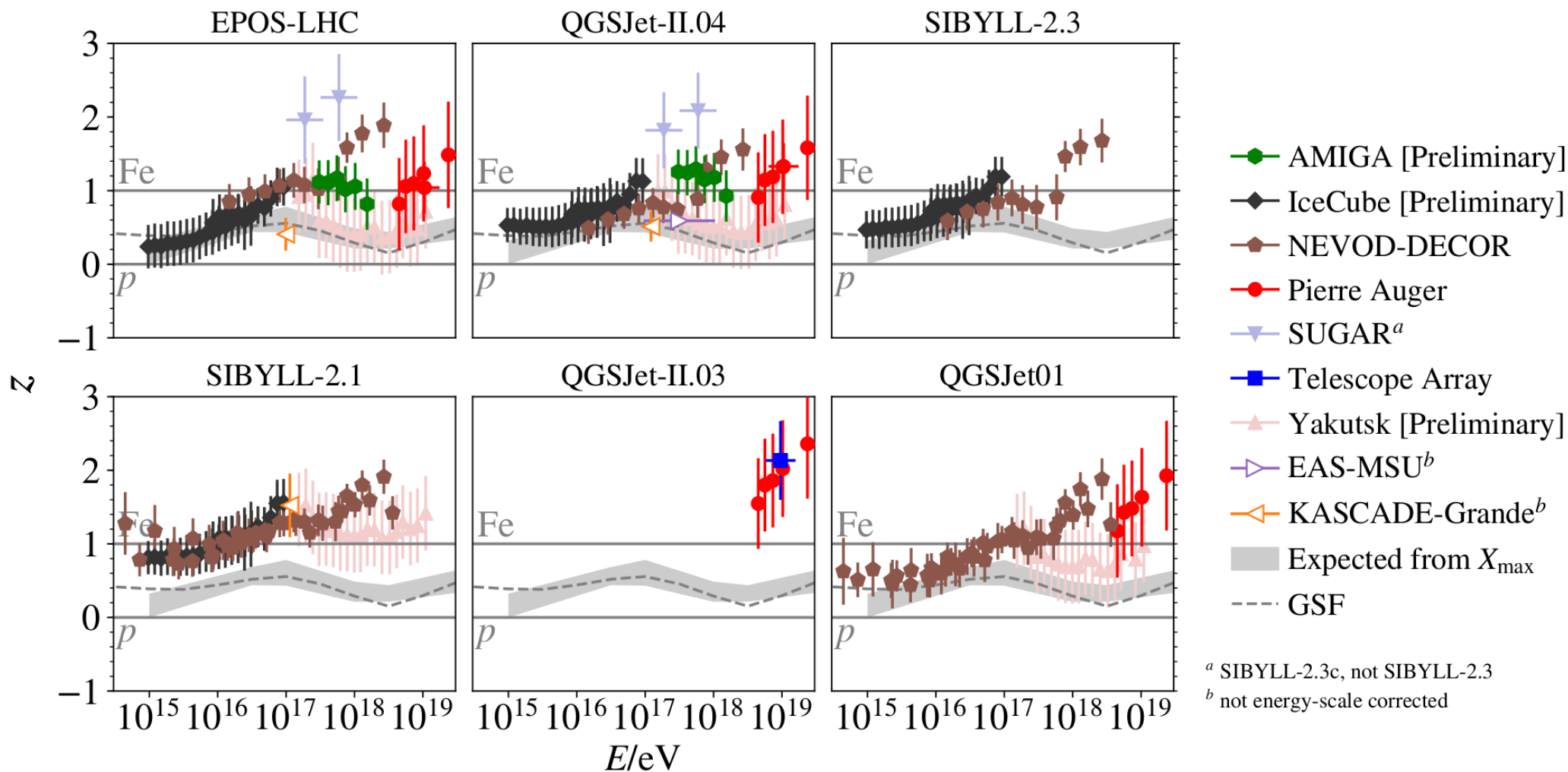


# Summary

**New input from LHC** which could help to reproduce **EAS data consistently** but solution still under study. Possible consequences on TeV to PeV muons and neutrinos.

- WHISP working group clearly established a muon production deficit in air shower simulations.
  - ➔ Exact scale not known (dependent on energy and mass)
- Most “natural” explanation given by a **change in pion charge ratio**.
  - ➔ Other possibilities limited by  $X_{\max}$  (multiplicity, inelasticity)
- Large change needed for a well constrained observable.
  - ➔ Different type of hadronization (**extended range for QGP**)
  - ➔ Possible to approach data but energy slope at the lower limit (preliminary)
- Same effect will increase the inclusive lepton flux above 100 GeV.
  - ➔ Inclusive flux will constrain solution for “muon puzzle” of EAS
- Not all relevant CERN data taken into account in model yet.
  - ➔ NA61, LHCf (+RHICf) ...

## Data Rescaled





# Pion Interactions

MPD measurement helped to understand the importance of pion interactions (lack of accelerator data until NA61) and baryon effect on propagation

- ➔ low pion elasticity in DPMJETIII
- ➔ high pion elasticity (diffraction) in EPOS and Sibyll driven by LHC data (and high baryon number (Ostapchenko et al. Phys.Rev. D93 (2016) no.5, 051501))
- ➔ diffraction with pion projectile or proton projectile are different

