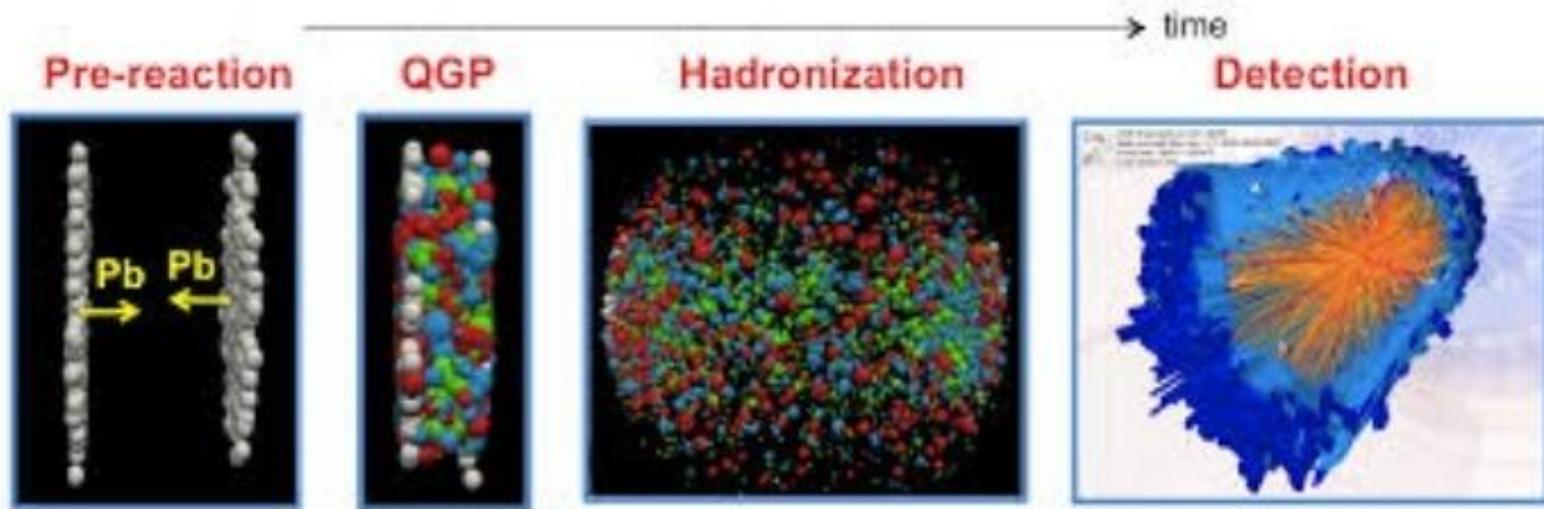


Heavy quarks dynamics in heavy ion collisions



Santosh Kumar Das

**School of Physical Science
Indian Institute of Technology Goa
Goa, India**

**In Collaboration with:
V. Greco, M. Ruggieri, S. Plumari, F. Scardina**

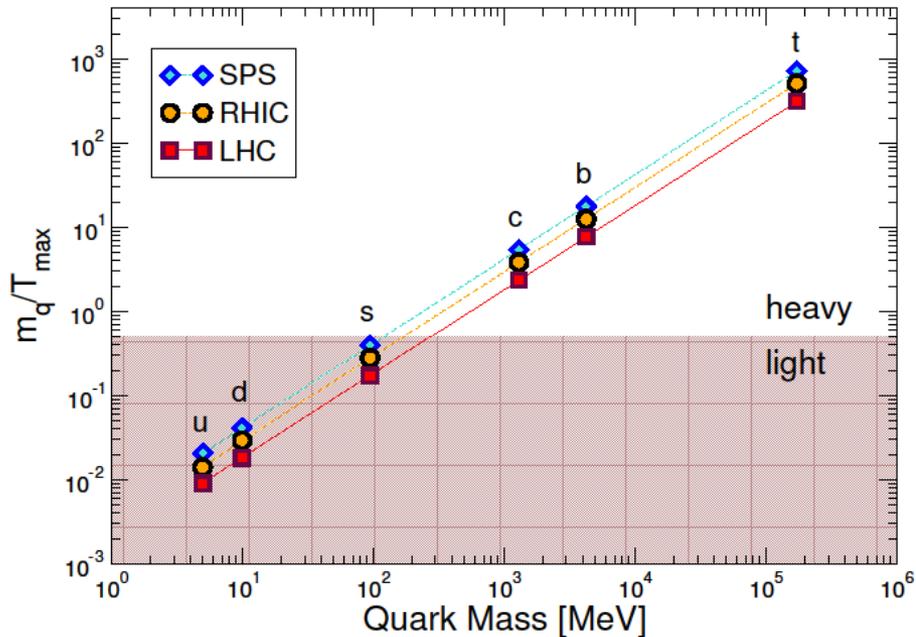


OUTLINE

- ❑ Introduction**
- ❑ Quark Gluon Plasma - the primordial fluid**
- ❑ Heavy quark momentum evolution: Langevin vs Boltzmann**
- ❑ Probing of initial electromagnetic field by heavy quarks**
- ❑ Heavy quark dynamics in small system**
- ❑ Summary and outlook**

Heavy Quark & QGP

At very high density and temperature hadrons melt to a new phase of matter called **Quark Gluon Plasma (QGP)**.



SPS to LHC

$$\sqrt{s} = 17.3 \text{ GeV to } 2.76 \text{ TeV} \sim 100 \text{ times}$$

$$T_i = 200 \text{ MeV to } 600 \text{ MeV} \sim 3 \text{ times}$$

$$M_{c,b} \gg \Lambda_{QCD}$$

Produced by pQCD process (out of Equil.)

$$\tau_{c,b} \gg \tau_{QGP}$$

They go through all the QGP life time

$$M_{c,b} \gg T_0$$

No thermal production

Boltzmann Kinetic equation

$$\left(\frac{\partial}{\partial t} + \frac{P}{E} \frac{\partial}{\partial x} + \mathbf{F} \cdot \frac{\partial}{\partial \mathbf{p}} \right) f(x, p, t) = \left(\frac{\partial f}{\partial t} \right)_{col}$$

➤ The plasma is uniform ,i.e., the distribution function is independent of \mathbf{x} .

➤ In the absence of any external force, $\mathbf{F}=\mathbf{0}$

$$R(p, t) = \left(\frac{\partial f}{\partial t} \right)_{col} = \int d^3k [\omega(p+k, k) f(p+k) - \omega(p, k) f(p)]$$

$$\omega(p, k) = g \int \frac{d^3q}{(2\pi)^3} f'(q) v_{q,p} \sigma_{p,q \rightarrow p-k, q+k} \longrightarrow \text{is rate of collisions which change the momentum of the charmed quark from } p \text{ to } p-k$$

$$\omega(p+k, k) f(p+k) \approx \omega(p, k) f(p) + k \cdot \frac{\partial}{\partial \mathbf{p}} (\omega f) + \frac{1}{2} k_i k_j \frac{\partial^2}{\partial p_i \partial p_j} (\omega f)$$

$$\frac{\partial \mathbf{f}}{\partial t} = \frac{\partial}{\partial \mathbf{p}_i} \left[\mathbf{A}_i(\mathbf{p}) \mathbf{f} + \frac{\partial}{\partial \mathbf{p}_j} [\mathbf{B}_{ij}(\mathbf{p}) \mathbf{f}] \right]$$

B. Svetitsky PRD 37(1987)2484

where we have defined the kernels

$$\mathbf{A}_i = \int d^3k \omega(\mathbf{p}, \mathbf{k}) \mathbf{k}_i \rightarrow \text{Drag Coefficient}$$

$$\mathbf{B}_{ij} = \int d^3k \omega(\mathbf{p}, \mathbf{k}) \mathbf{k}_i \mathbf{k}_j \rightarrow \text{Diffusion Coefficient}$$

Langevin Equation

$$dx_j = \frac{p_j}{E} dt$$

$$dp_j = -\Gamma p_j dt + \sqrt{dt} C_{jk}(t, p + \xi dp) \rho_k$$

where Γ is the deterministic friction (drag) force
 C_{ij} is stochastic force in terms of independent Gaussian-normal distributed random variable

$$\rho = (\rho_x, \rho_y, \rho_z), \quad P(\rho) = \left(\frac{1}{2\pi}\right)^3 \exp\left(-\frac{\rho^2}{2}\right)$$

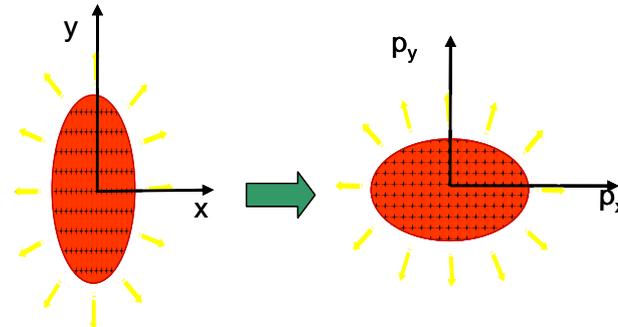
With $\langle \rho_i(t) \rho_k(t') \rangle = \delta(t-t') \delta_{jk}$

$\xi = 0$ the pre-point Ito interpretation of the momentum argument of the covariance matrix.

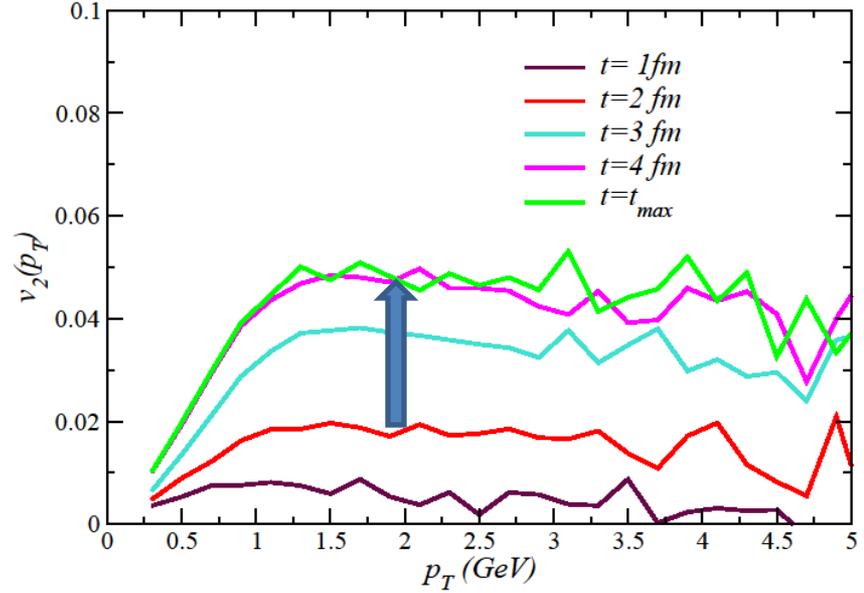
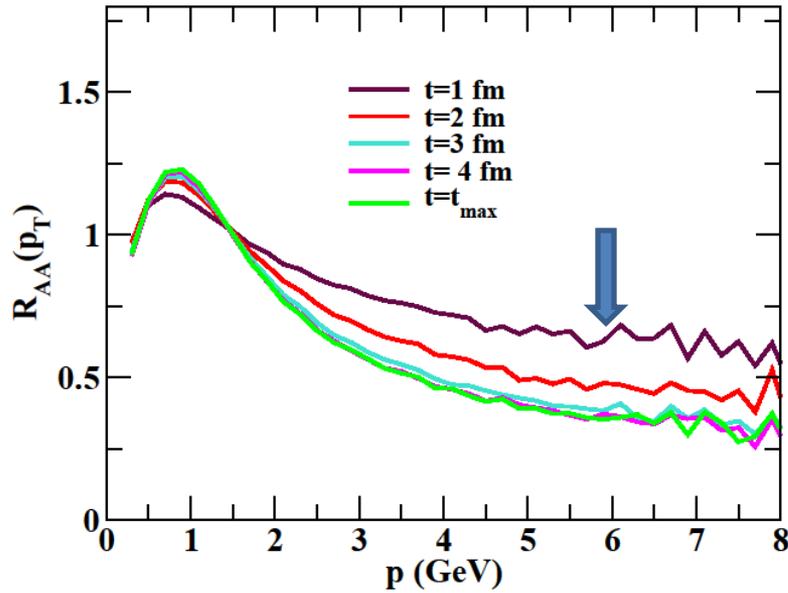
H. v. Hees and R. Rapp
 arXiv:0903.1096

$$R_{AA} = \frac{\left(\frac{dN}{d^2 p_T dy}\right)^{Au+Au}}{N_{coll} \left(\frac{dN}{d^2 p_T dy}\right)^{p+p}}$$

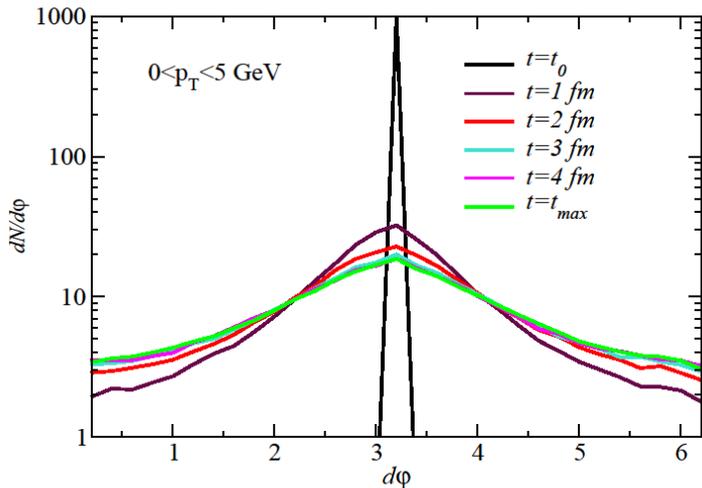
$$v_2(p_T) = \left\langle \frac{p_x^2 - p_y^2}{p_x^2 + p_y^2} \right\rangle$$



Time evolution of Heavy quarks observables



Das, Scardina, Plumari, Greco
J. Phys. Conf. Ser. 668 (2016)012051



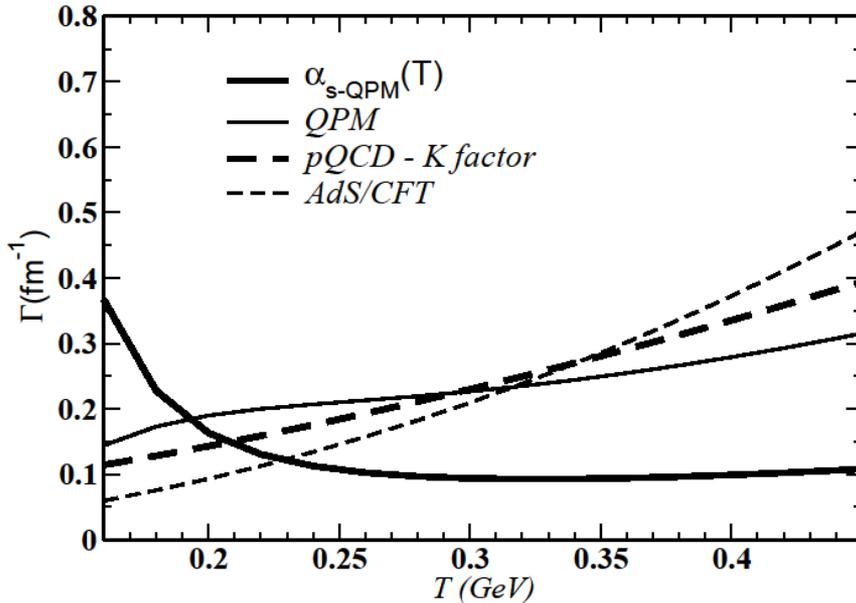
RAA and $dN/d\phi_{c\bar{c}}$ developed during the early stage of the evolution $\rightarrow T_i$

v_2 developed during the later stage of the evolution $\rightarrow T_c$

T dependence of the interaction i.e the transport coefficients are the essential ingredient for the simultaneous description of HQ observables.

T-dependence of the Drag Coefficient

Drag Coefficient



Das, Scardina, Plumari, Greco
Phys. Lett. B 747 (2015)260-264

Chandra and Das
Phys. Rev. D,93 (2016) , 094036

pQCD (Combridge)

$$\alpha_{pQCD} = \frac{4\pi}{11\ln(2\pi T\Lambda^{-1})}, \quad m_D^2 = 4\pi\alpha_{pQCD}(T)T^2$$

AdS/CFT

$$\mathcal{V}_{AdS/CFT} = k \frac{T^2}{M}$$

Gubser
PRD,74,126005 (2006)
Akamatsu, Hatsuda, Hirano
PRC, 79, 054907 (2009)
Das and Davody
PRC, 89,054912 (2014)

Quasi-Particle-Model (fit to IQCD ϵ, P)

$$g_{QP}^2(T) = \frac{48\pi^2}{(11N_c - 2N_f) \ln \left[\lambda \left(\frac{T}{T_c} - \frac{T_s}{T_c} \right) \right]^2} \quad \lambda=2.6, \quad T_s=0.57$$

$$m_g^2 = \frac{1}{6} \left(N_c + \frac{1}{2} N_f \right) g^2 T^2$$

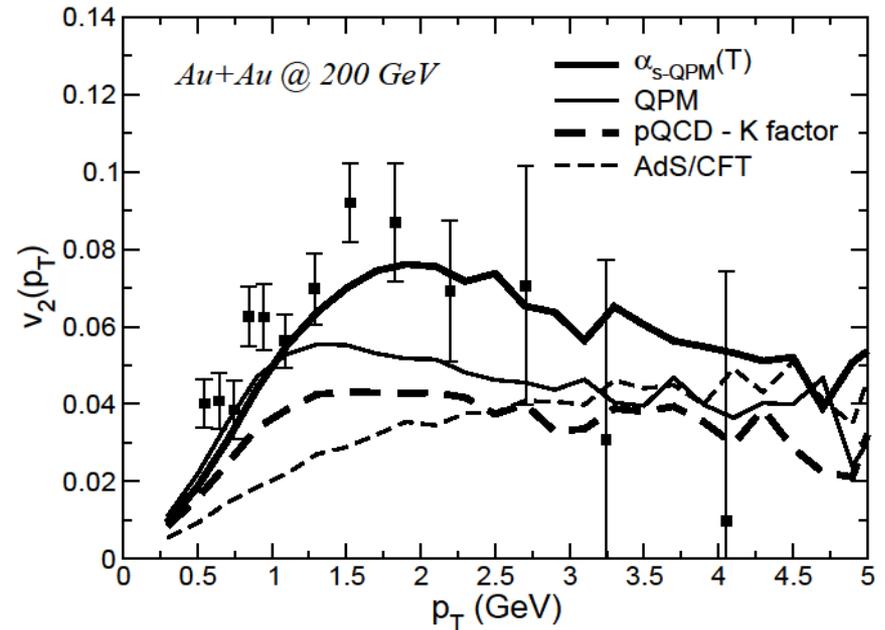
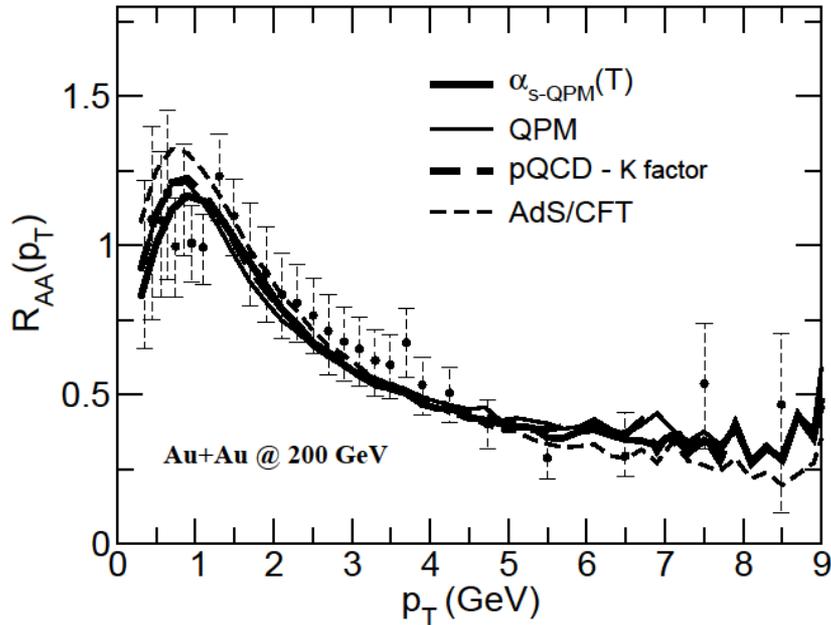
$$m_q^2 = \frac{N_c^2 - 1}{8N_c} g^2 T^2$$

$\alpha_{QPM}(T)$, $m_{q,g}=0$

we mean simply the coupling of the QPM,
but with a bulk of massless q and g .

RAA and v_2 @ RHIC

(Au+Au@200A GeV. $h=8$ fm)

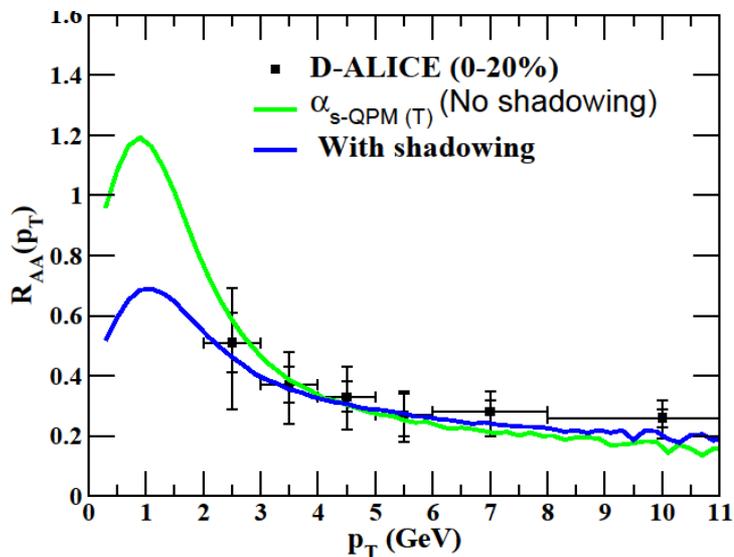
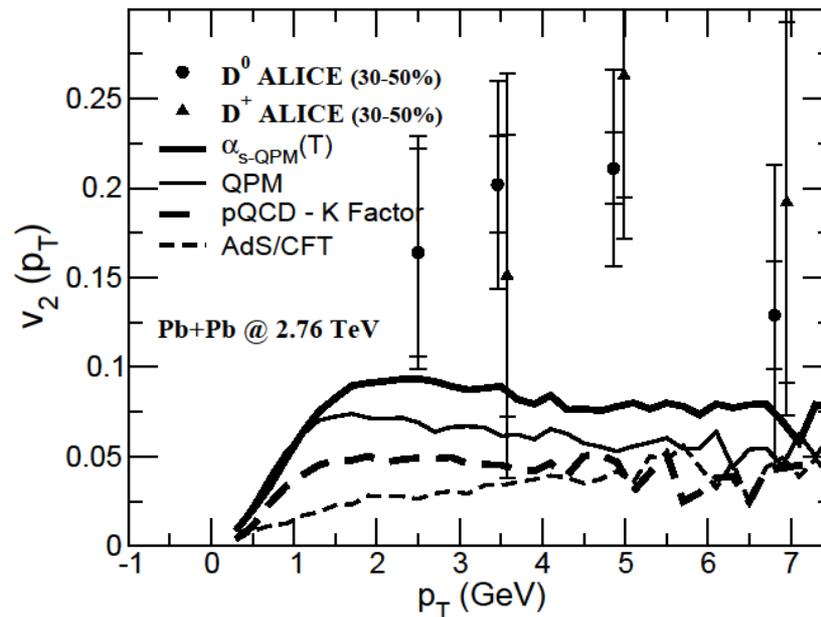
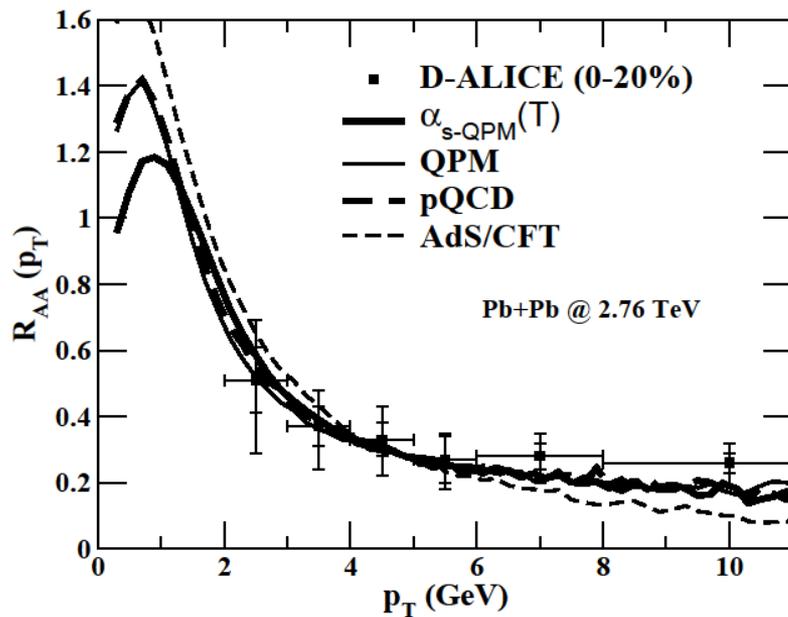


Das, Scardina, Plumari, Greco
Phys. Lett. B 747 (2015)260-264

The Larger the interaction /drag coefficients at T_c , the larger the v_2 for the same RAA.

This indicates we need to go beyond pQCD.

RAA and v_2 @ ALICE



Das, Scardina, Plumari, Greco
Phys. Lett. B 747 (2015) 260-264

At LHC we underestimate the v_2 .

Heavy quark momentum evolution: Langevin vs Boltzmann

$$\omega(p+k, k)f(p+k) \approx \omega(p, k)f(p) + k \cdot \frac{\partial}{\partial p} (\omega f) + \frac{1}{2} k_i k_j \frac{\partial^2}{\partial p_i \partial p_j} (\omega f)$$



Boltzmann Equation



Fokker Planck

It will be interesting to study both the equation in a identical environment to ensure the validity of this assumption at different momentum transfer and their subsequent effects on RAA and v_2 .

Langevin dynamics:

$$dx_j = \frac{p_j}{E} dt$$

$$dp_j = -\Gamma p_j dt + \sqrt{dt} C_{jk}(t, p + \xi dp) \rho_k$$

Das, Scardina, Plumari and Greco
Phys. Rev. C, 90, 044901 (2014)

Γ is the deterministic friction (drag) force

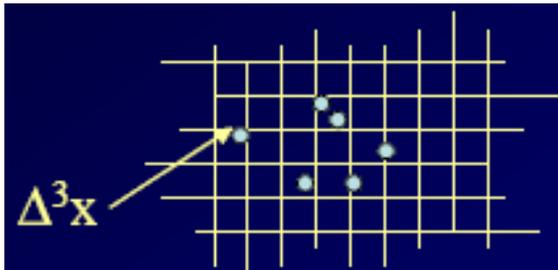
C_{ij} is stochastic force in terms of independent Gaussian-normal distributed random variable.

Transport theory

$$p^\mu \partial_\mu f(x, p) = C_{22}$$

We consider two body collisions

$$C_{22} = \frac{1}{2E_1} \int \frac{d^3 p_2}{(2\pi)^3 2E_2} \frac{1}{\nu} \int \frac{d^3 p'_1}{(2\pi)^3 2E'_1} \frac{d^3 p'_2}{(2\pi)^3 2E'_2} f'_1 f'_2 |\mathcal{M}_{1'2' \rightarrow 12}|^2 (2\pi)^4 \delta^{(4)}(p'_1 + p'_2 - p_1 - p_2) \\ - \frac{1}{2E_1} \int \frac{d^3 p_2}{(2\pi)^3 2E_2} \frac{1}{\nu} \int \frac{d^3 p'_1}{(2\pi)^3 2E'_1} \frac{d^3 p'_2}{(2\pi)^3 2E'_2} f_1 f_2 |\mathcal{M}_{12 \rightarrow 1'2'}|^2 (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p'_1 - p'_2)$$



$$\Delta t \rightarrow 0$$

$$\Delta^3 x \rightarrow 0$$



**Exact
solution**

Collision integral is solved with a **local stochastic sampling**

Das, Scardina, Plumari and Greco
Phys. Rev. C, 90, 044901 (2014)

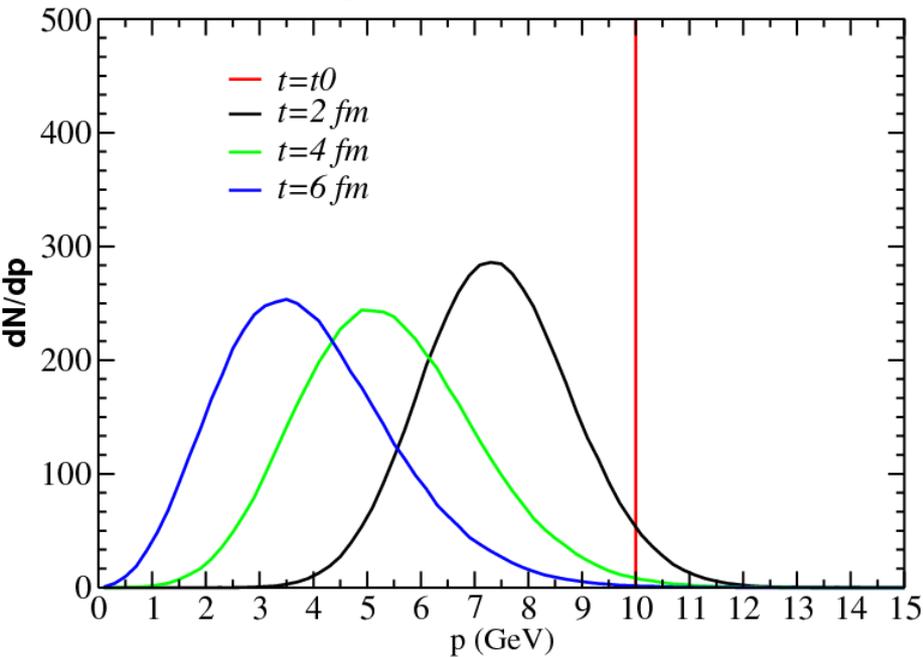
$$P_{22} = \frac{\Delta N_{\text{coll}}^{2 \rightarrow 2}}{\Delta N_1 \Delta N_2} = v_{\text{rel}} \sigma_{22} \frac{\Delta t}{\Delta^3 x}$$

Evolution: Boltzmann vs Langevin (Charm)

Momentum evolution starting from a δ (Charm) in a Box

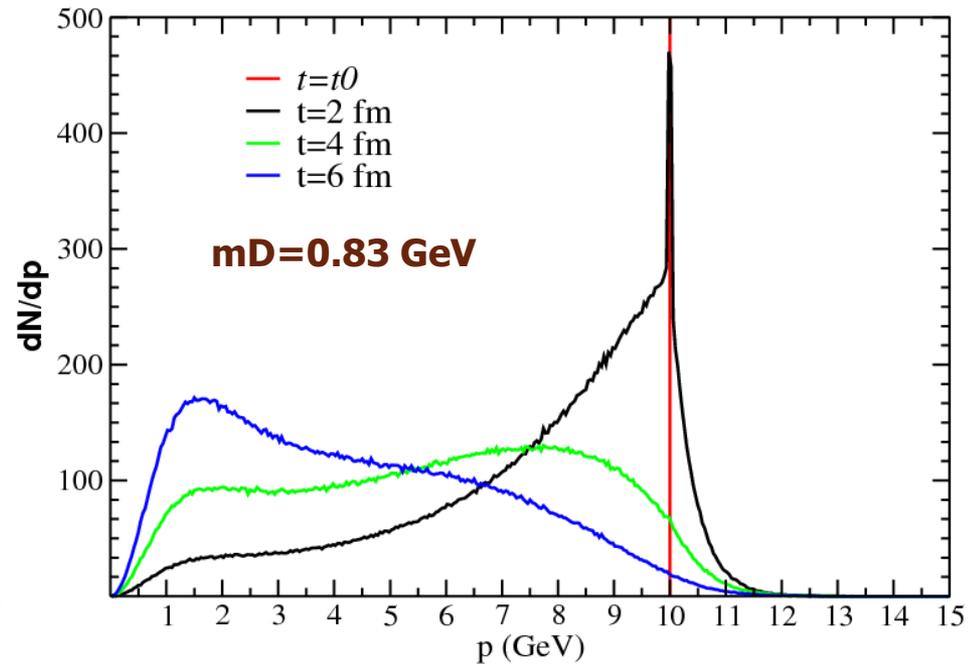
$$\frac{dN}{d^3 p_{initial}} = \delta(p - 10 \text{ GeV})$$

Langevin



In case of Langevin the distributions are Gaussian as expected by construction

Boltzmann



In case of Boltzmann the charm quarks does not follow the Brownian motion

Das, Scardina, Plumari and Greco
Phys. Rev. C,90,044901(2014)

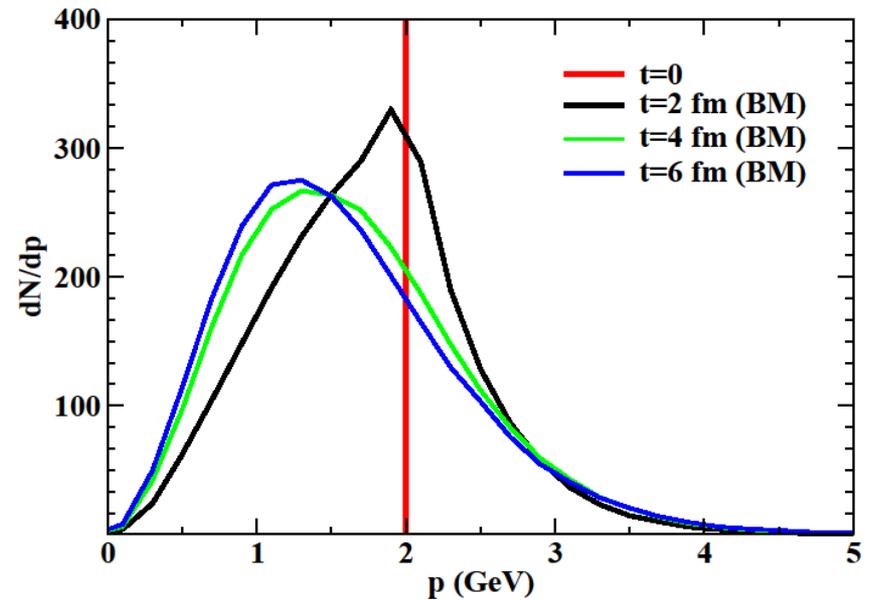
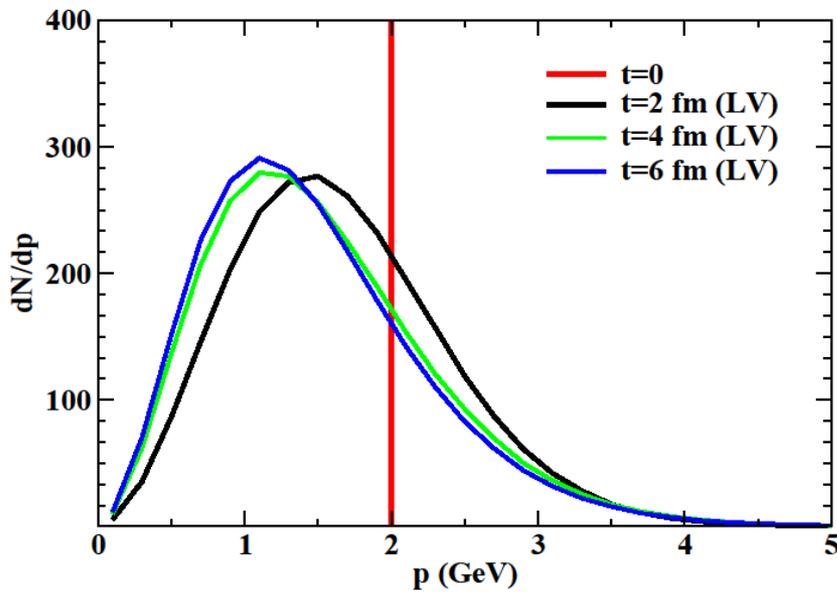
Evolution: Boltzmann vs Langevin (Charm)

Momentum evolution starting from a δ (Charm) in a Box

$$\frac{dN}{d^3 p_{initial}} = \delta(p - 2\text{GeV})$$

Langevin

Boltzmann



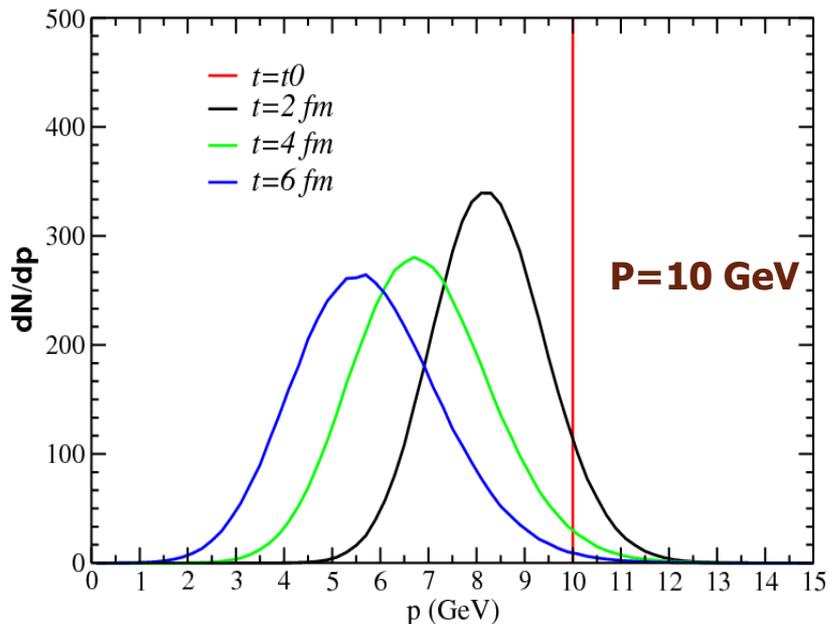
In case of Langevin the distributions are Gaussian as expected by construction

In case of Boltzmann the charm quarks follow the Brownian motion: At Low Momentum.

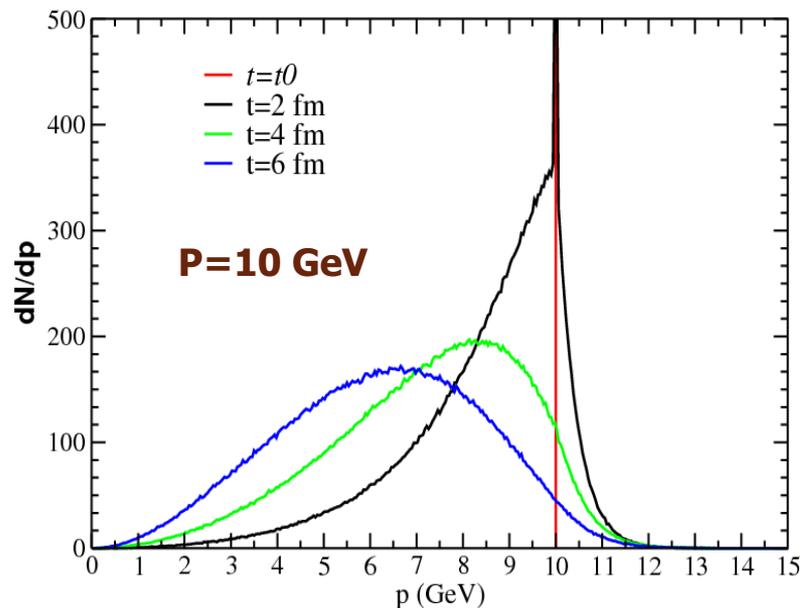
Momentum evolution starting from a δ (Bottom)

In a Box

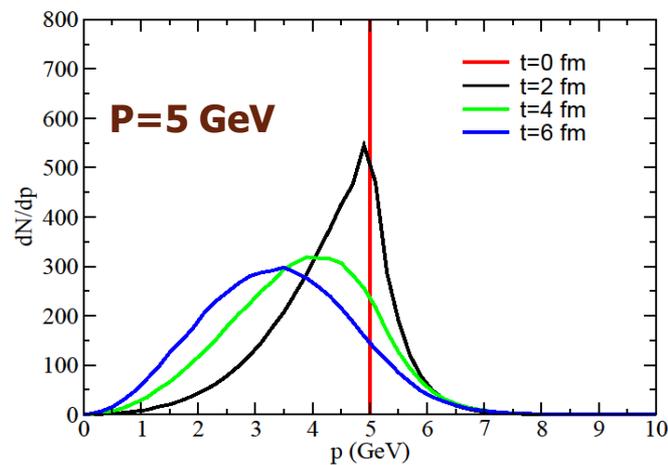
Langevin



Boltzmann

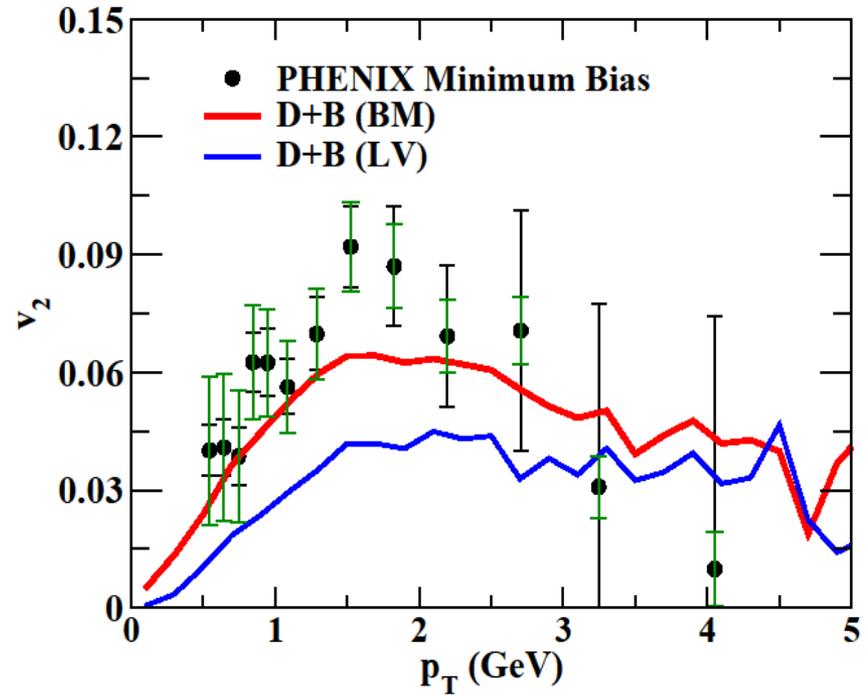
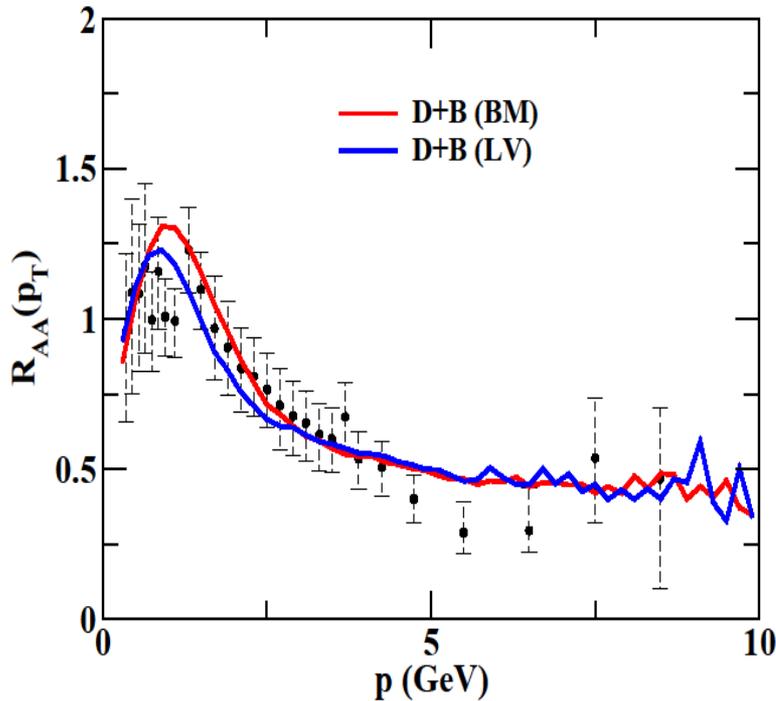


For bottom quarks it works better.



R_{AA} and v_2 at RHIC

(With near isotropic cross-section)



Das, Scardina, Plumari and Greco
Phys. Rev. C,90,044901(2014)

At fixed RAA Boltzmann approach generate larger v_2 .
(depending on mD and M/T)

With isotropic cross section one can describe both RAA and V_2
simultaneously within the Boltzmann approach !

Hadronization: Coalescence plus Fragmentation

Fragmentation function gives the probability to get a hadron from a parton:

$$f_H(p_T) = \sum_p f_p(p_T / z) \otimes D_{p \rightarrow H}(z)$$

$\langle z \rangle \sim 0.9$ for charm quark and $\langle z \rangle \sim 0.5$ for light quark

Coalescence is the convolution of two /three parton distribution folded by a wave function:

$$\frac{dN_{Meson}}{d^2 p_T} = g_M \sum_{i,j} P_q(i) P_q(j) \delta^{(2)}(p_T - p_{iT} - p_{jT}) f_M(x_i, x_j; p_i, p_j)$$

Hadron wave function

$$\frac{dN_{Baryon}}{d^2 p_T} = g_B \sum_{i,j,k} P_q(i) P_q(j) P_q(k) \delta^{(2)}(p_T - p_{iT} - p_{jT} - p_{kT}) f_B(x_i, x_j, x_k; p_i, p_j, p_k)$$

V. Greco, C.M. Ko, and P. L'evai
PRL 90 , 202302 (2003)

Das, Torres-Rincon, Tolos, Minissale, Scardina, Greco
PRD,94,114039,2016

Our Approach:

Charm quark production: NLO pQCD

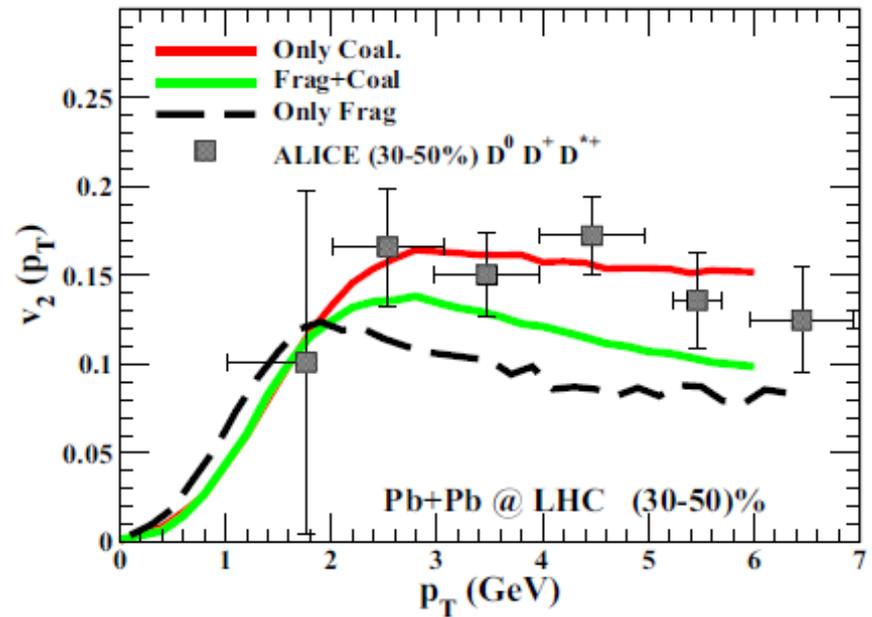
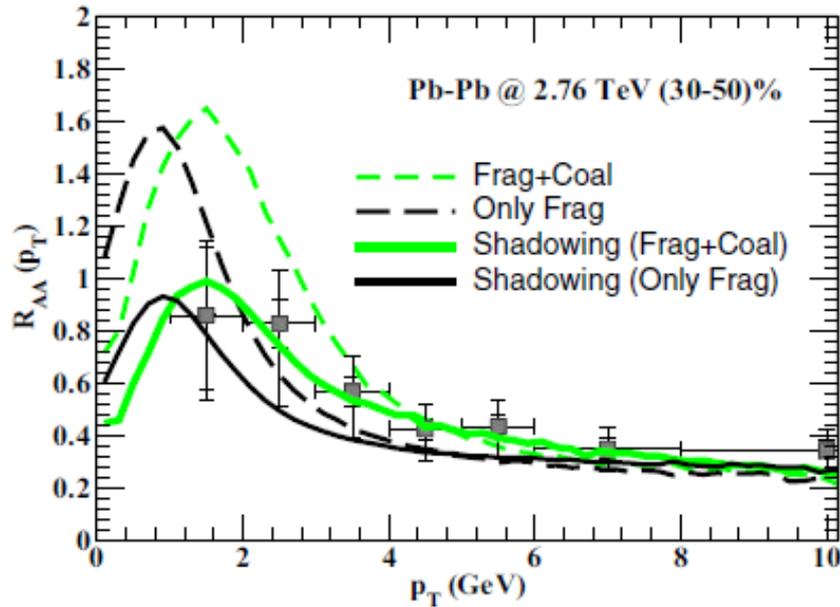
Heavy quark momentum evolution: Boltzmann equation

Heavy quark –bulk interaction: Quasiparticle model

Bulk evolution: Boltzmann equation

Hadronization: Coalescence plus Fragmentation

LHC results: RAA vs v2



T-dependence of interaction: QPM

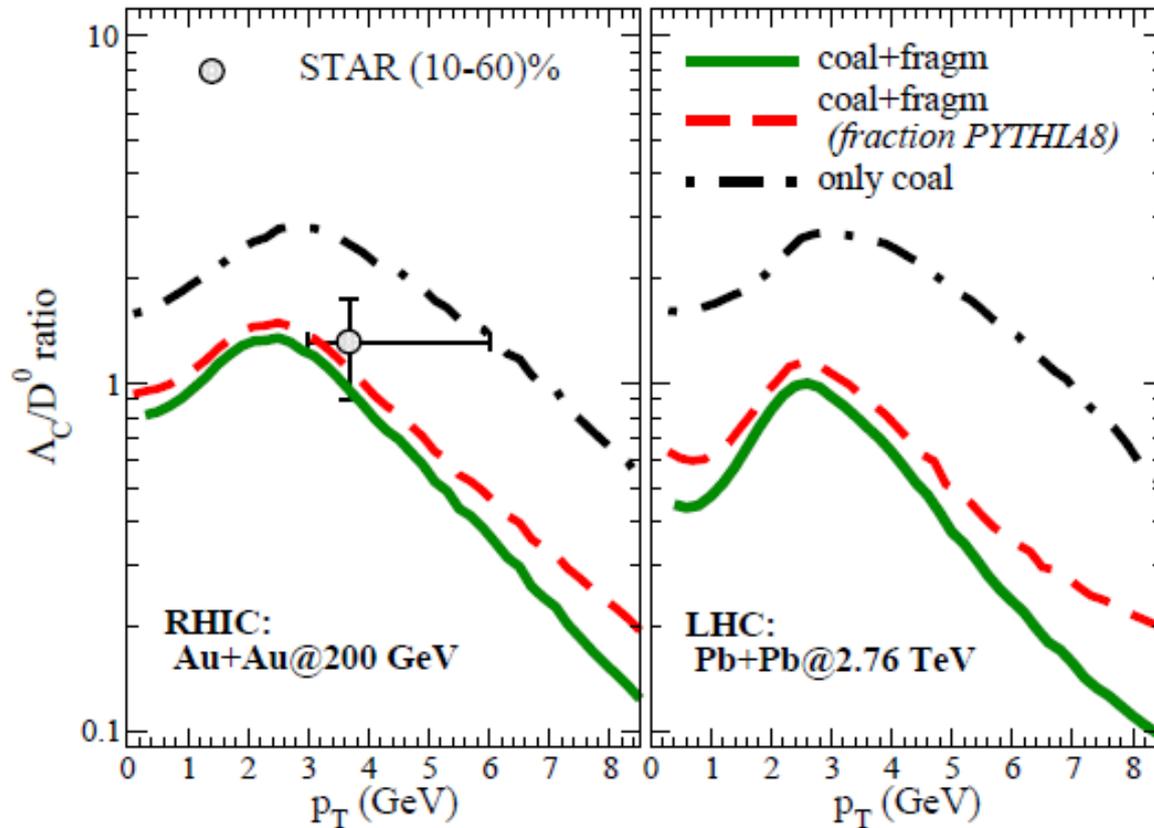
Boltzmann equation for HQ momentum evolution

Hadronization by coalescence plus fragmentation

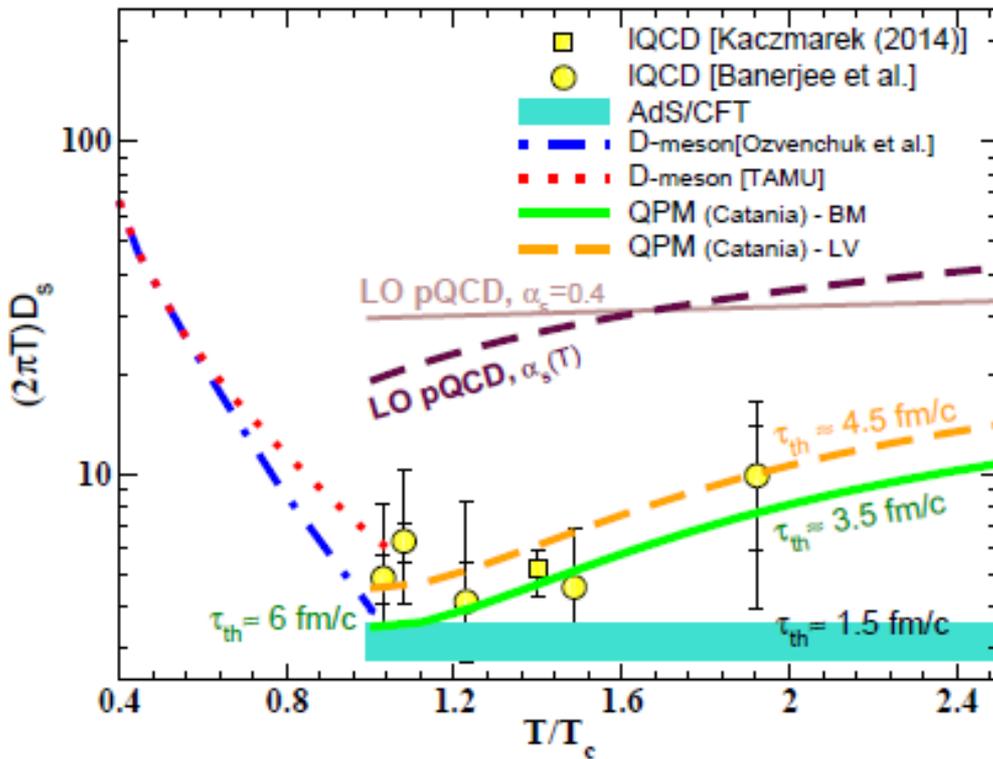
Scardina, Das, Minissale, Plumari, Greco
PRC, 96,044905 (2017)

Heavy Baryon to meson ratio

(Serve as a tool to disentangle different hadronization mechanisms)



Connection to lattice QCD



$$D_s(p=0) = \frac{T}{m_Q \gamma} = T m_Q \tau_{th}$$

$$\tau_{th} = \frac{M}{2\pi T^2} (2\pi T D_s) \cong 1.8 \frac{2\pi T D_s}{(T/T_c)^2} \text{ fm/c}$$

Open Heavy Flavour physics in heavy-ion collisions have the potential to link the phenomenology to lattice QCD and experiments.

**Scardina, Das, Minissale, Plumari, Greco
PRC, 96,044905 (2017)**

Unified attempts to understand heavy quark transport coefficients in QGP (HQ-RRTF)

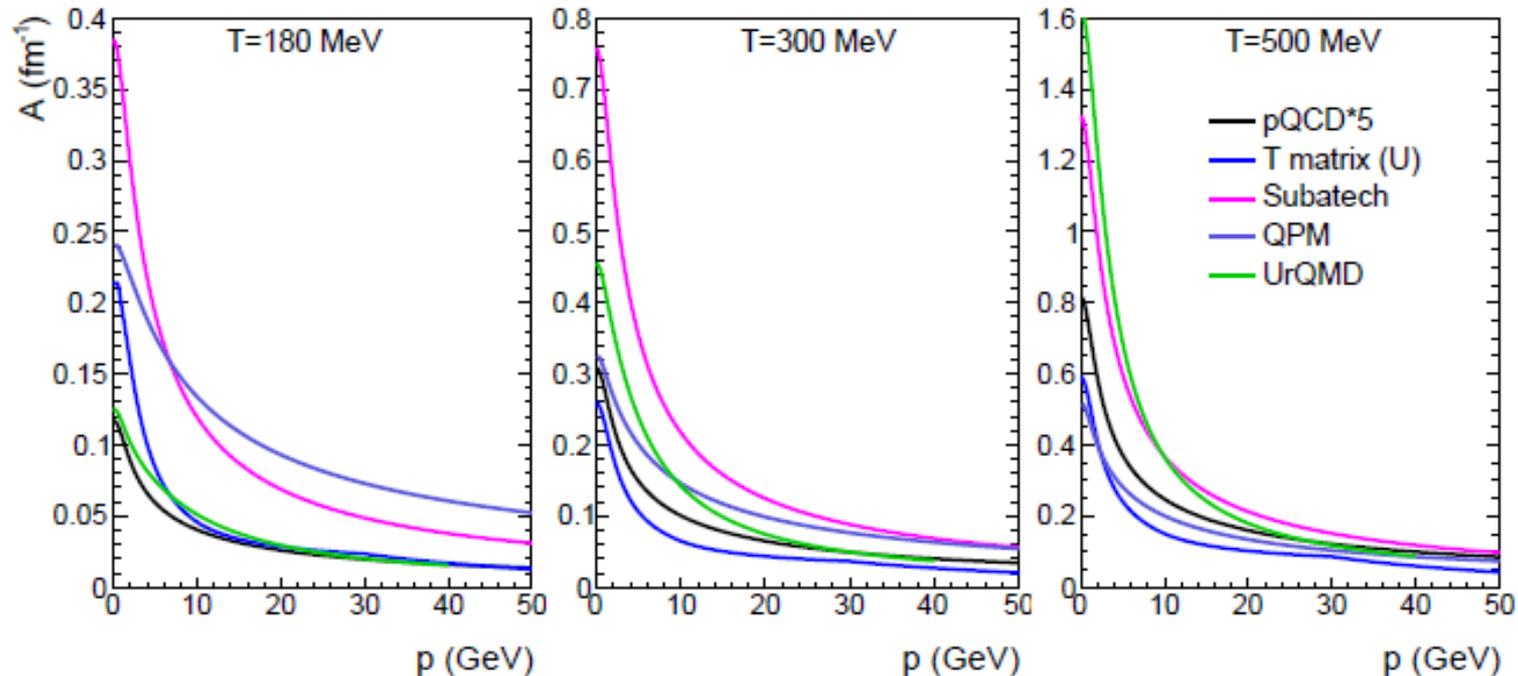
Extraction of Heavy-Flavor Transport Coefficients in QCD Matter

R. Rapp^{*1}, P.B. Gossiaux^{*2}, A. Andronic^{*3,4}, R. Averbeck^{*3}, S. Masciocchi^{*3}, A. Beraudo⁵,
E. Bratkovskaya^{3,6}, P. Braun-Munzinger^{3,7}, S. Cao⁸, A. Dainese⁹, S.K. Das^{10,11},
M. Djordjevic¹², V. Greco^{11,13}, M. He¹⁴, H. van Hees⁶, G. Inghirami^{3,6,15,16}, O. Kaczmarek^{17,18},
Y.-J. Lee¹⁹, J. Liao²⁰, S.Y.F. Liu¹, G. Moore²¹, M. Nahrgang², J. Pawłowski²², P. Petreczky²³,
S. Plumari¹¹, F. Prino⁵, S. Shi²⁰, T. Song²⁴, J. Stachel⁷, I. Vitev²⁵, and X.-N. Wang^{26,18}

¹*Cyclotron Institute and Department of Physics and Astronomy, Texas A&M University, College Station, USA*

²*SUBATECH, IMT Atlantique, Université de Nantes, CNRS-IN2P3, Nantes, France*

³*Research Division and EMMI, GSI Helmholtzzentrum für Schwerionenforschung, Darmstadt, Germany*



JET-HQ

Towards the extraction of heavy-quark transport coefficients in quark-gluon plasma

Shanshan Cao,¹ Gabriele Coci,^{2,3} Santosh Kumar Das,^{4,2} Weiyao Ke,⁵ Shuai Y.F. Liu,⁶ Salvatore Plumari,² Taesoo Song,⁷ Yingru Xu,⁵ Jörg Aichelin,⁸ Steffen Bass,⁵ Elena Bratkovskaya,^{9,10} Xing Dong,¹¹ Pol Bernard Gossiaux,⁸ Vincenzo Greco,^{2,3} Min He,¹² Marlene Nahrgang,⁸ Ralf Rapp,⁶ Francesco Scardina,^{2,3} and Xin-Nian Wang^{13,11,*}

¹*Department of Physics and Astronomy, Wayne State University, Detroit, Michigan 48201, USA*

²*Department of Physics and Astronomy, University of Catania, Via S. Sofia 64, I-95125 Catania, Italy*

³*Laboratori Nazionali del Sud, INFN-LNS, Via S. Sofia 62, I-95123 Catania, Italy*

⁴*School of Physical Science, Indian Institute of Technology Goa, Ponda, Goa, India*

⁵*Department of Physics, Duke University, Durham, North Carolina 27708, USA*

⁶*Columbia Institute and Department of Physics and Astronomy*

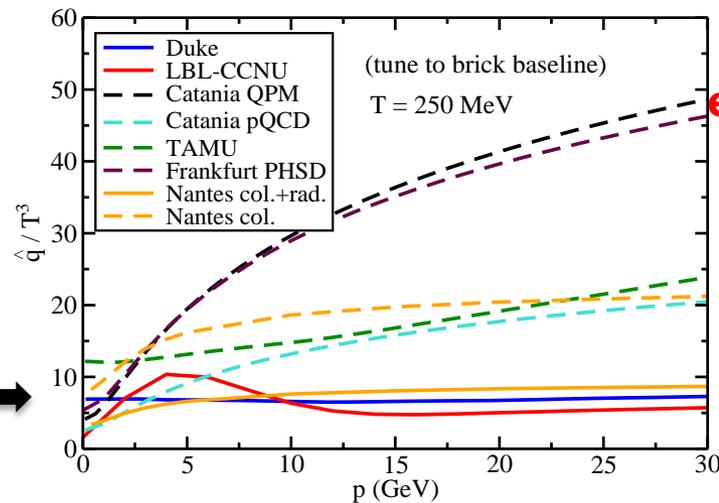
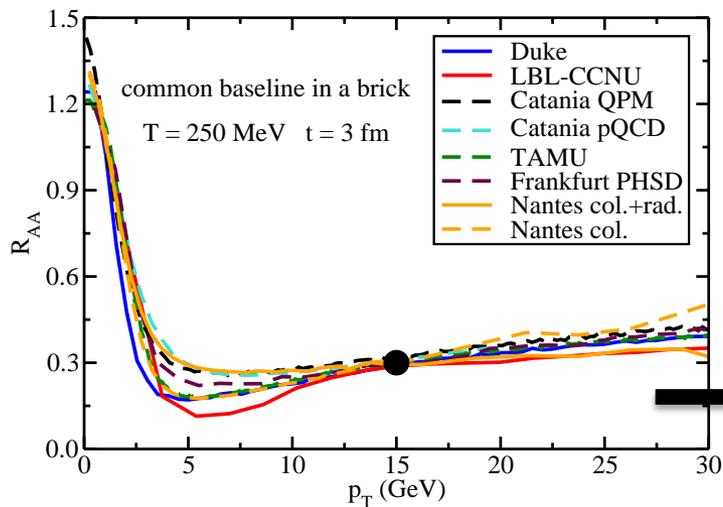
To investigate the origin of the large theoretical uncertainty:

Models	transport schemes	basic	tune 1	tune 2
Duke	Langevin	fixed $\alpha_s = 0.3$ only	$(\alpha, \beta, \gamma) = (1.89, 1.59, 0.26)$	$D_s = 0.77 \times D_s(\text{tune 1})$
LBNL-CCNU	Boltzmann	fixed $\alpha_s = 0.3$ only	$\alpha_s = 0.24$	$\alpha_s = 0.28$
Catania QPM	Boltzmann	running $\alpha_s(T)$	$K = 2.25$	$K = 3.45$
Catania pQCD	Langevin	running $\alpha_s(T)$	$K = 3.4$	$K = 3.1$
TAMU	Langevin	U from lQCD	no tuning	$K = 2.45$
Frankfurt PHSD	Boltzmann	running $\alpha_s(T)$	no tuning	$K = 1.6$
Nantes col. + rad.	Boltzmann	running $\alpha_s(q^2)$	$K = 0.8$	$K = 0.45$
Nantes col. only	Boltzmann	running $\alpha_s(q^2)$	$K = 1.5$	$K = 1.1$

TABLE II. Key inputs and model tunings of different HQ transport formalisms. In LBNL-CCNU model a momentum-dependent K_p factor is applied to the effective fixed coupling constant α_s .

Systematic error ONLY due to energy loss

Common baseline: same initial c spectrum, static medium $T = 250$ MeV, $L = 3$ fm, $R_{AA}(c) = 0.3$ at $p_T = 15$ GeV



elastic: QPM

elastic:
pQCD or T -
matrix

elastic +
inelastic

- Convergence of transport parameter into 3 groups when energy loss in a brick is under control
- Can be further distinguished by future data on 2-particle correlation

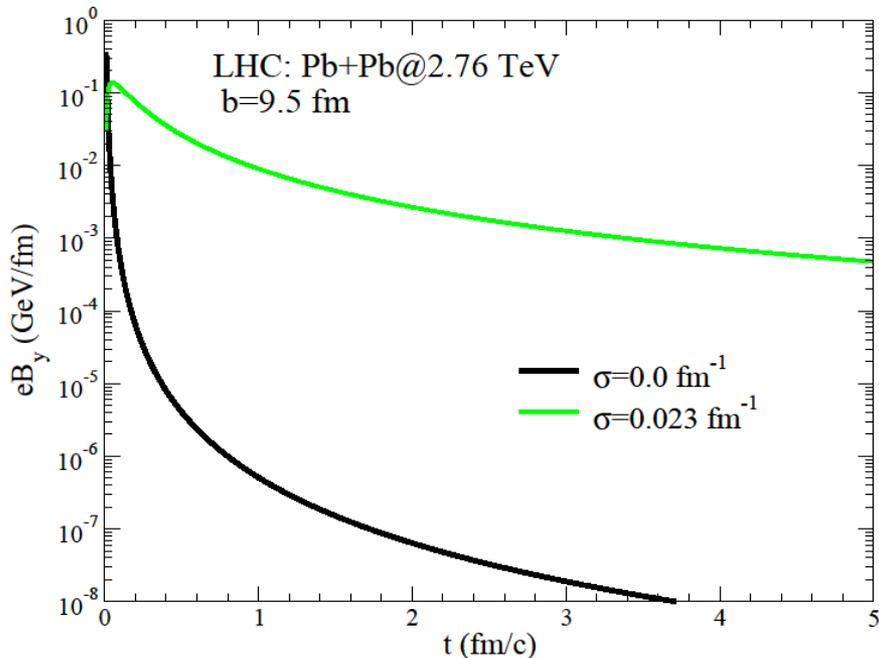
Impact of EM field on heavy quark dynamics at LHC

$$dp_j = -\Gamma p_j dt + \sqrt{dt} C_{jk}(t, p + \xi dp) \rho_k + F_{ext} dt$$

$$F_{ext} = q(E' + v \times B')$$

$$E' = \gamma(E + v \times B) - (\gamma - 1)(E \cdot \hat{v}) \hat{v}$$

$$B' = \gamma(B - v \times E) - (\gamma - 1)(B \cdot \hat{v}) \hat{v}$$



Electromagnetic field has been included in the Langevin equation as an external force.

We consider both E and B.

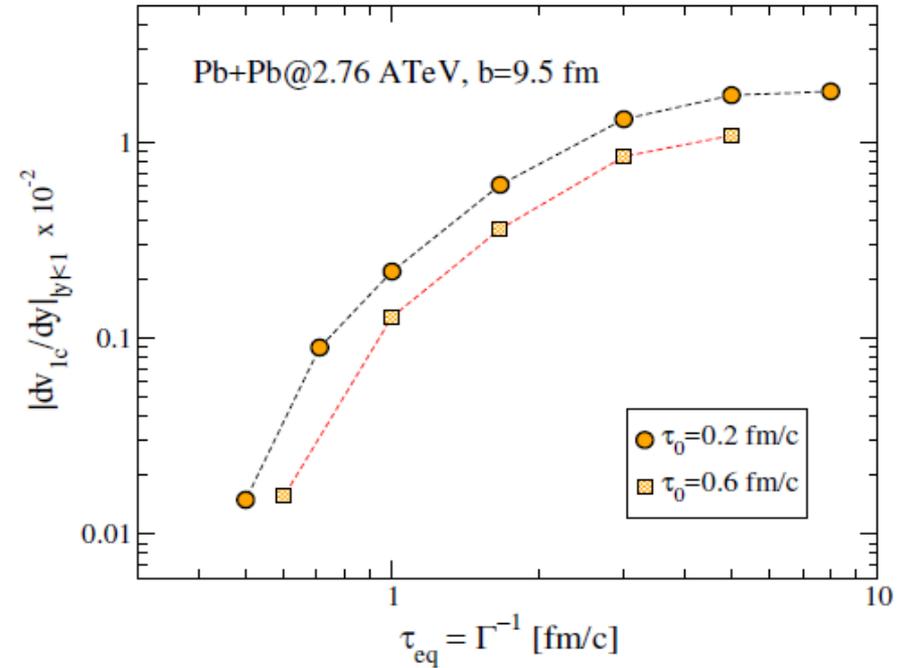
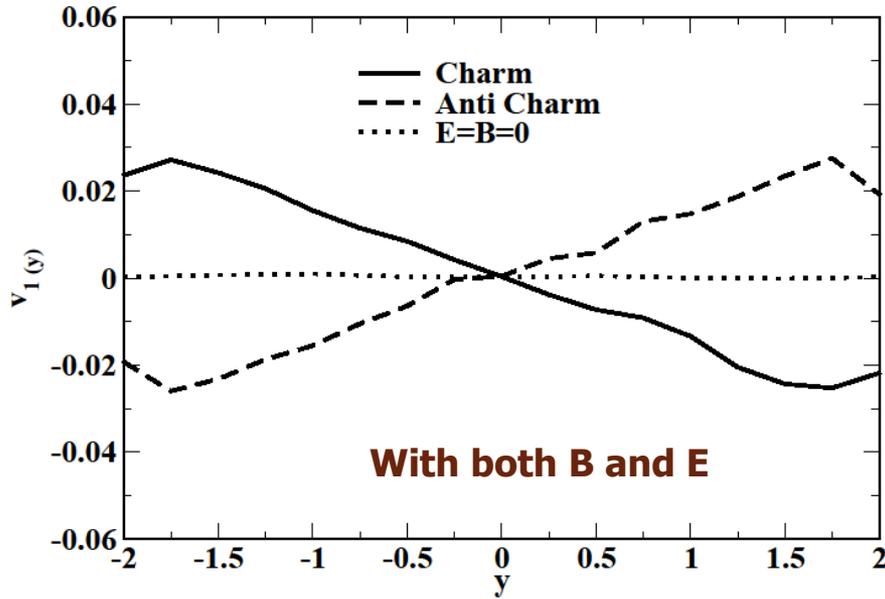
$$\mathbf{B}_x = \mathbf{B}_z = 0$$

$$\text{And } E_y = E_z = 0$$

$$v_1 = \left\langle \frac{p_x}{p_T} \right\rangle$$

**Das, Plumari, Chartarjee, Scardina, Greco, Alam
Phys. Lett. B, 768 (2017) 260**

Heavy quark v1@LHC

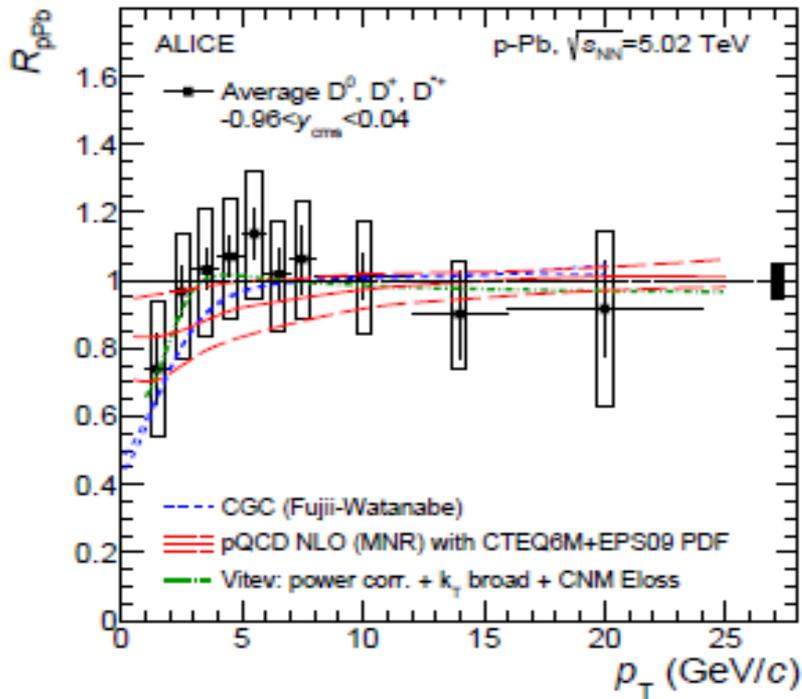


Das, Plumari, Chartarjee, Scardina, Greco, Alam
Phys. Lett. B, 768 (2017) 260

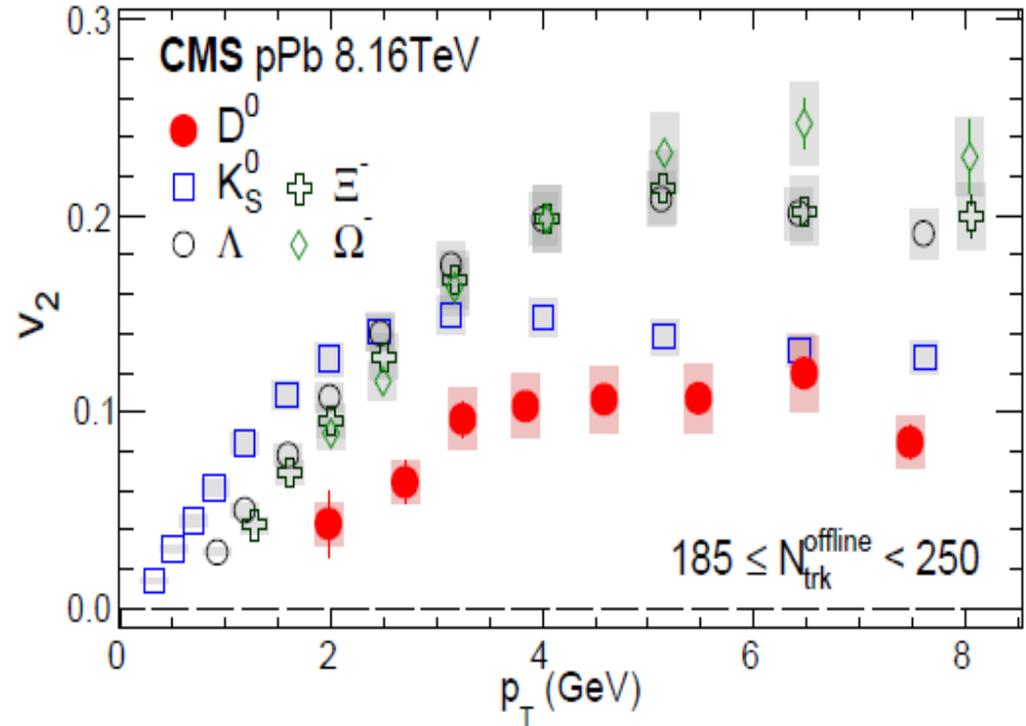
Heavy quark v1 is larger than light quark v1.

**Recent data from ALICE indicates splitting in D and Dbar v1.
(Hard probes 2018)**

Heavy quark in small system (p-nucleus)



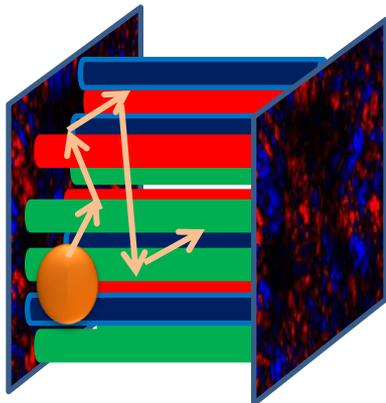
ALICE Collaboration
 Phys. Rev. Lett. 113 (2014) 232301



CMS Collaboration
 arXiv:1804.09767v2

What mechanism could build up v_2 without energy loss?

Heavy quarks as probes of the evolving Glasma



(Adapted from M. Ruggieri)

$$t_{\text{formation}} \approx \frac{1}{2m_c} \approx 0.06 \text{ fm}/c$$



HQs can probe the very early evolution of the Glasma fields

Hamilton equations of motion of c -quarks:

$$\frac{dx_i}{dt} = \frac{p_i}{E} \quad E = \sqrt{\mathbf{p}^2 + m^2}$$

$$\mathbf{v} \equiv \frac{\mathbf{p}}{E} \quad \text{(Relativistic) Velocity}$$

$$E \frac{dp_i}{dt} = gQ_a F_{i\nu}^a p^\nu,$$

$$\frac{d\mathbf{p}}{dt} = q\mathbf{E} + q(\mathbf{v} \times \mathbf{B}) \quad \text{Lorentz force}$$

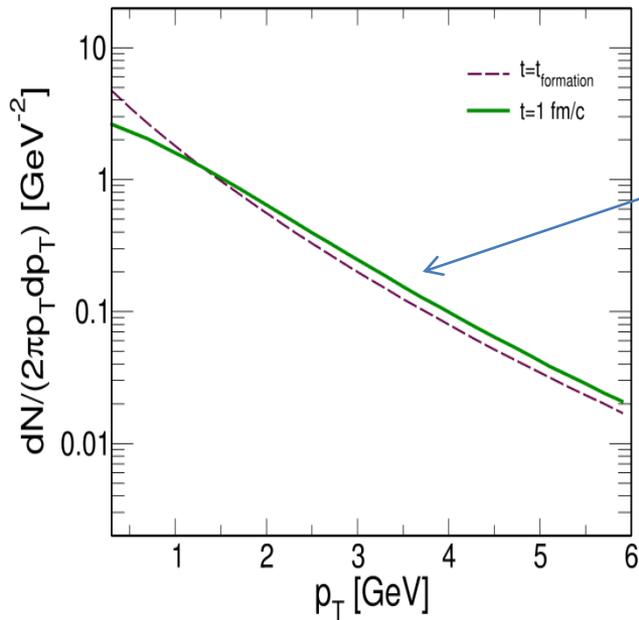
$$E \frac{dQ_a}{dt} = -gQ_c \varepsilon^{cba} \mathbf{A}_b \cdot \mathbf{p} \quad \text{Wong (1979)}$$

$$D_\mu J_a^\mu = 0$$

Gauge-invariant conservation of the color Current carried by charm quarks + gluons

$$J_a^\mu = \bar{c} \gamma^\mu T_a c$$

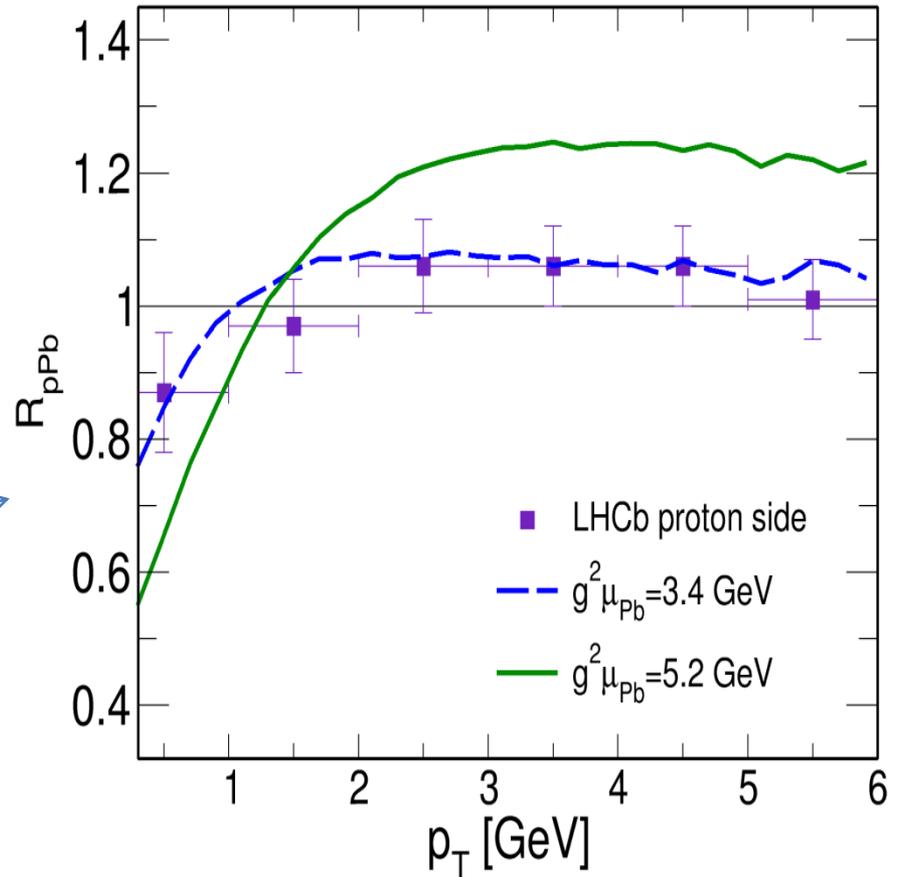
Equations of motion of heavy quarks are solved in the background given by the evolving Glasma fields



Initial distribution: from perturbative QCD
Evolution: interaction with the Glasma

D-mesons R_{pPb}

Standard fragmentation [Peterson et al.(1983)]

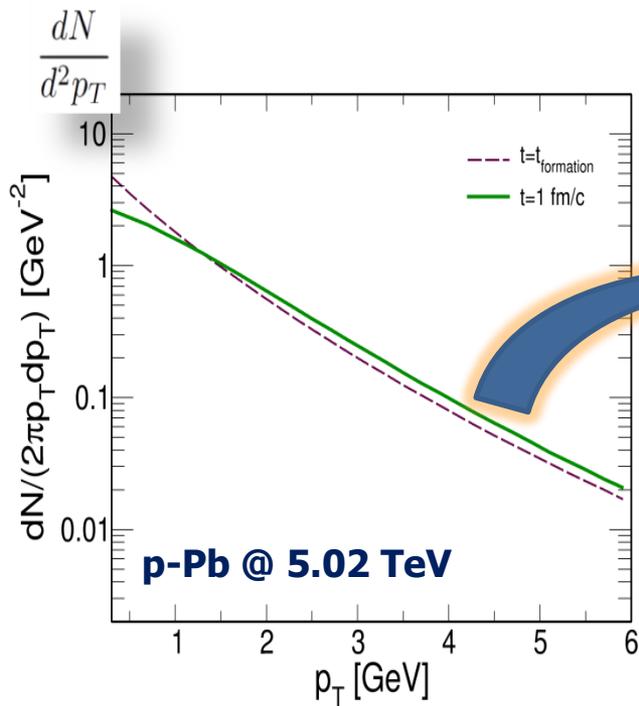


$$R_{pPb} = \frac{(dN/d^2p_T)_{\text{final}}}{(dN/d^2p_T)_{p\text{QCD}}}$$

$R_{pPb} \neq 1$

Interaction with the fields created by the collision

Diffusion results in acceleration

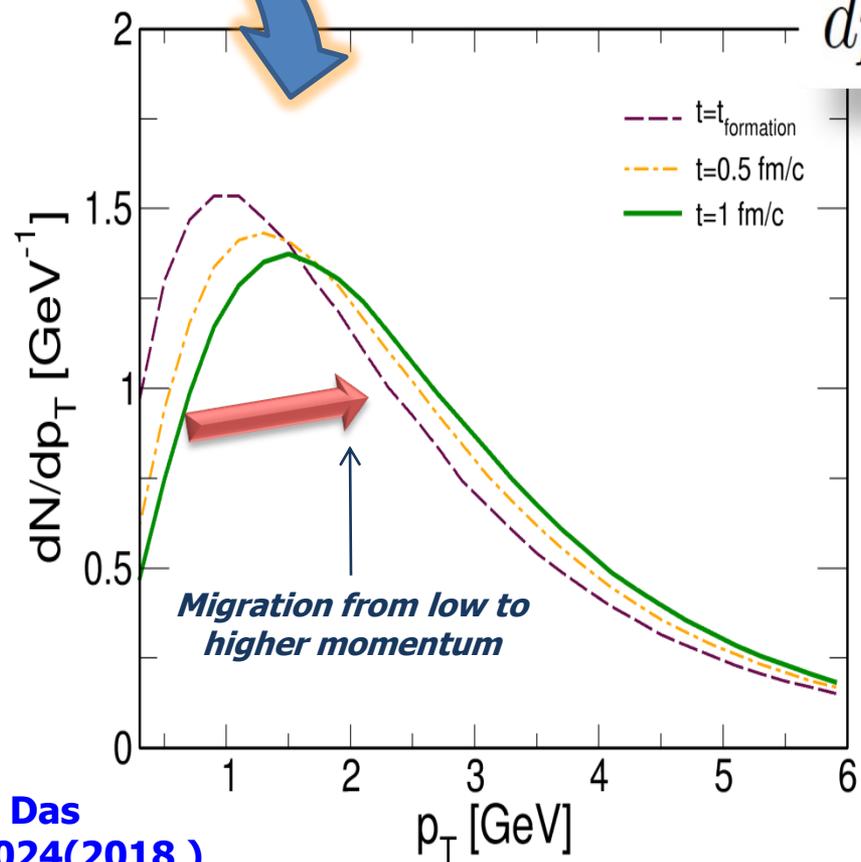


$\times p_T (\times 2\pi)$

Energy gain

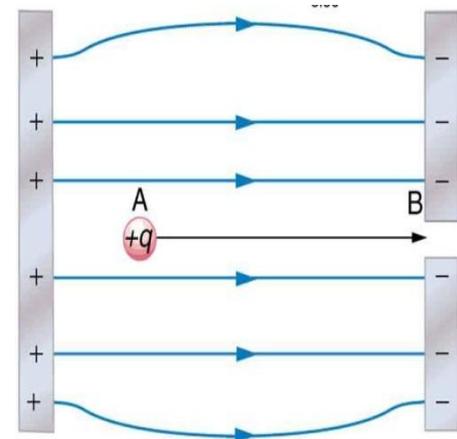
$$\frac{dN}{dp_T}$$

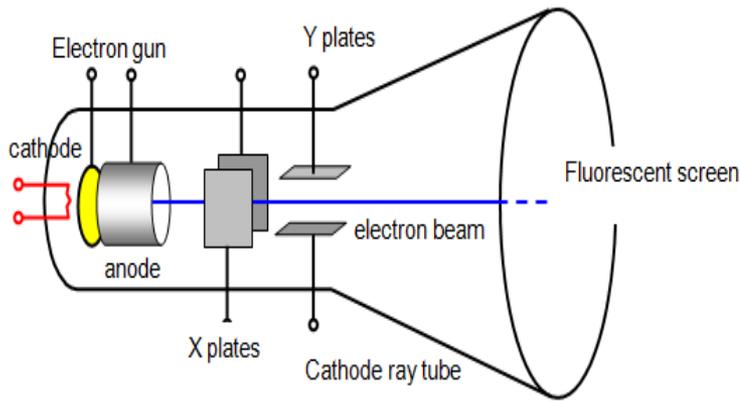
Heavy quarks seem to be accelerated by the transverse (color-)electric field



$$\frac{\Delta p}{\Delta t} = qE$$

**Ruggieri and Das
PRD, 98, 094024(2018)**



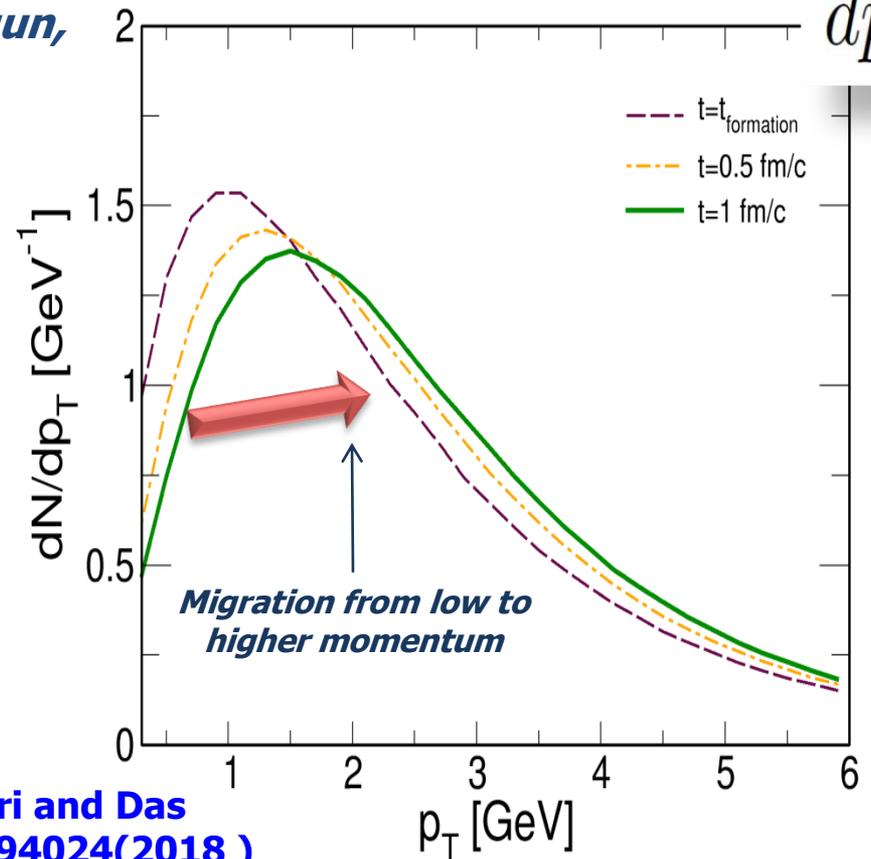


Why cathode tube?

Electrons are produced by the electron gun, then accelerated by the electric field

Heavy quarks seem to be accelerated by the (color-)electric field

$$\frac{dN}{dp_T}$$

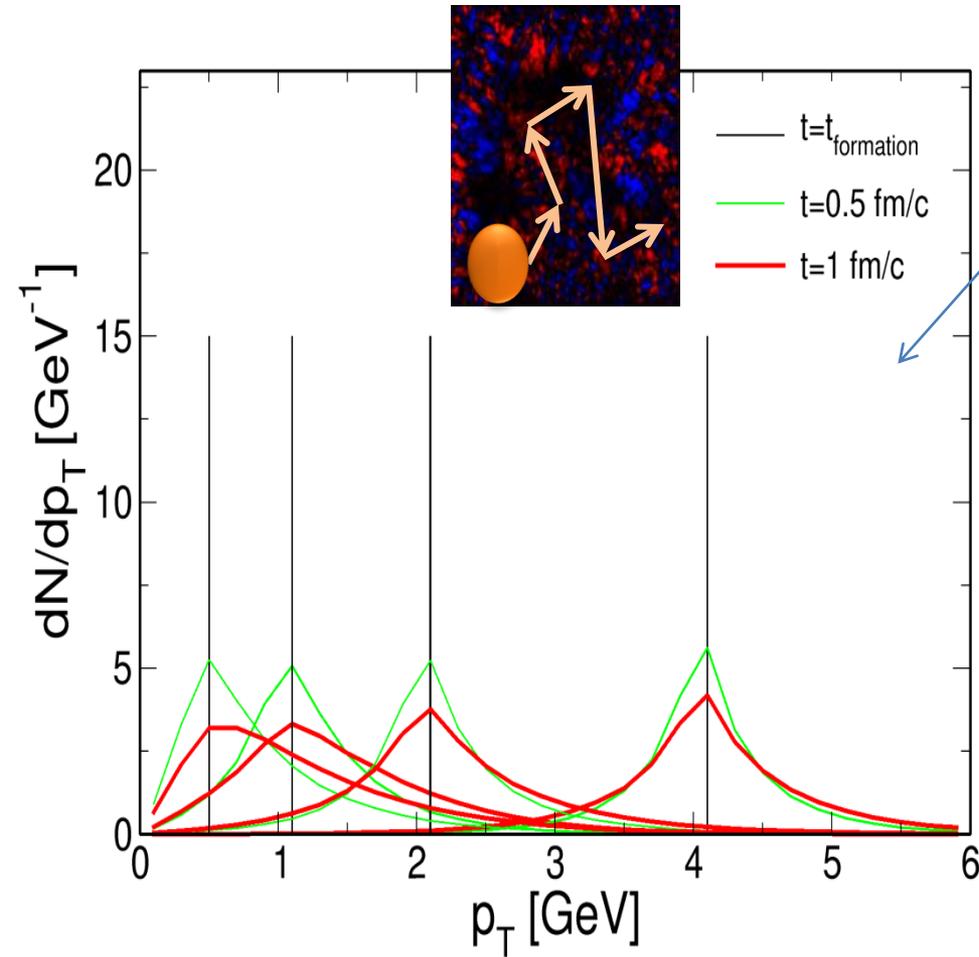


BRIONVEGA



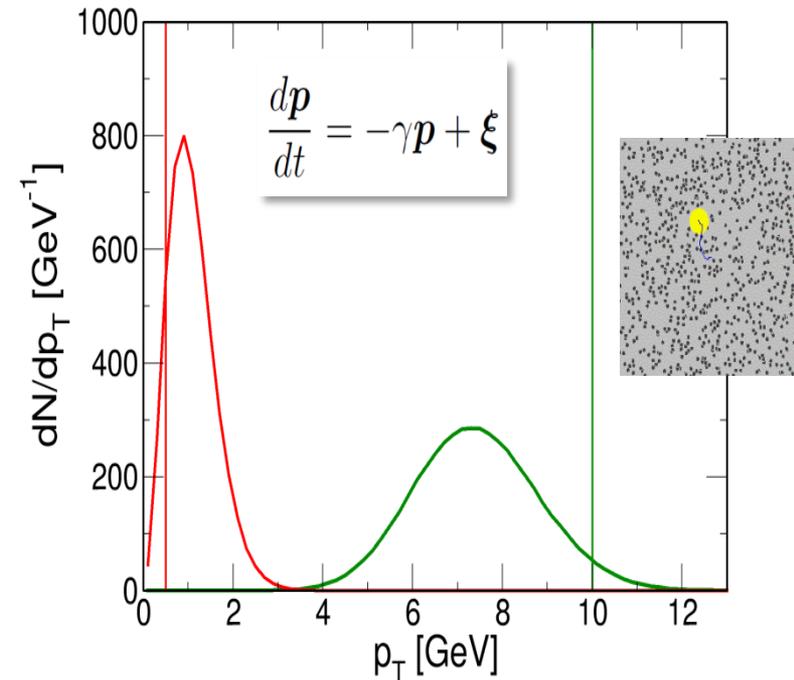
Ruggieri and Das
PRD, 98, 094024(2018)

HQs in Glasma



(Anomalous) Diffusion in momentum space

T=500 MeV HQs in hot plasma



For anomalous diffusion see Havlin and Avraham (1987)

J. H. Liu, G. Coci, S. K. Das, M. Ruggieri
Under preparation

Summary & Outlook

- Heavy quarks are the novel probe to characterized QGP and to probe initial state.
- Our study indicates the temperature of the system produced at RHIC ($T=340$ MeV) and LHC ($T=510$ MeV) energies are much larger than the temperature needed to create the QGP.
- Several new experiments are coming up (FAIR, FCC) and we are looking for new observables which help us to understand several basic issues ...
 - ❖ Heavy quark diffusion coefficient in QGP and in Glasma.
 - ❖ Hadronization.
 - ❖ Einstein relation will be studied.
 - ❖ Heavy quark thermalization.
 - ❖ QGP in small system (p-Au)
 - ❖ To probe the effect of initial magnetic field.

Thank You



Hadronization: Coalescence plus Fragmentation

Fragmentation function gives the probability to get a hadron from a parton:

$$f_H(p_T) = \sum_p f_p(p_T/z) \otimes D_{p \rightarrow H}(z)$$

$\langle z \rangle \sim 0.9$ for charm quark and $\langle z \rangle \sim 0.5$ for light quark

Coalescence is the convolution of two /three parton distribution folded by a wave function:

$$\frac{dN_{Meson}}{d^2 p_T} = g_M \sum_{i,j} P_q(i) P_q(j) \delta^{(2)}(p_T - p_{iT} - p_{jT}) f_M(x_i, x_j; p_i, p_j)$$

Hadron wave function



$$\frac{dN_{Baryon}}{d^2 p_T} = g_B \sum_{i,j,k} P_q(i) P_q(j) P_q(k) \delta^{(2)}(p_T - p_{iT} - p_{jT} - p_{kT}) f_B(x_i, x_j, x_k; p_i, p_j, p_k)$$

V. Greco, C.M. Ko, and P. L'evai
PRL 90 , 202302 (2003)

Das, Torres-Rincon, Tolos, Minissale, Scardina, Greco
PRD,94,114039,2016

Motivation for Transport approach

$$\left\{ p^{*\mu} \partial_{\mu} + \left[p_{\nu}^{*} F^{\mu\nu} + m^{*} \partial^{\mu} m^{*} \right] \partial_{\mu}^{p^{*}} \right\} f(x, p^{*}) = C[f]$$

Free streaming Field Interaction

Collisions $\rightarrow \eta \neq 0$

- Starting from 1-body distribution function $f(x, p)$ and not from $T_{\mu\nu}$:
 - **$f(x, p)$ out-of-equilibrium: CGC-Qs scale or high p_T**
 - Extract viscous correction δf to $f(x, p)$
 - Relevant at LHC due to large amount of minijet production
 - **Freeze-out self-consistently related to $\eta/s(T)$**
 - **HQ dynamics in the same framework**

DISADVANTAGES?!

- Relaxation times fixed by kinetic theory
- Hadronization needed: coal.+frag . under progress

Simulate at fixed shear viscosity

Usually input of a transport approach are *cross-sections and fields*, but here we reverse it and start from η/s with aim of creating a more direct link to viscous hydrodynamics

Chapmann-Enskog

$$\frac{\eta}{s} = \frac{1}{15} \langle p \rangle \cdot \tau_\eta = \frac{1}{15} \frac{\langle p \rangle}{g\left(\frac{m_D}{T}\right) \sigma_{TOT} \rho}$$

$$g(a) = \frac{1}{50} \int dy y^6 \left[\left(y^2 + \frac{1}{3}\right) K_3(2y) - y K_2(2y) \right] h\left(\frac{a^2}{y^2}\right)$$

$g(a=m_D/2T)$ correct function that fix the relaxation time for the shear motion

$$0 < g(m_D/2T) < 2/3$$

forward
peaked

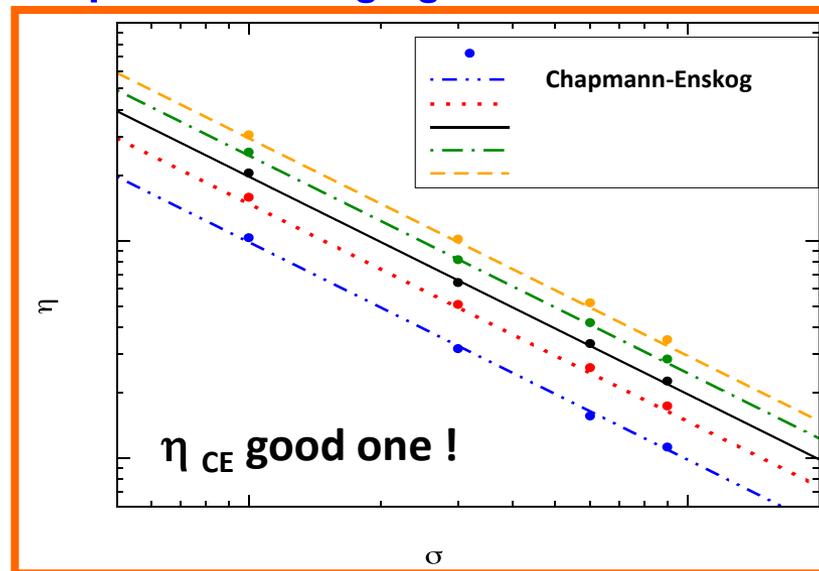
Isotropic
 $m_D \rightarrow \infty$

Transport code

$$\sigma_{tot}(n(\vec{r}), T) = \frac{1}{15} \frac{\langle p_\alpha \rangle}{g(a)n_\alpha} \frac{1}{\eta/s}$$

Space-Time dependent cross section evaluated locally

Chapman-Enskog agrees with Green-Kubo



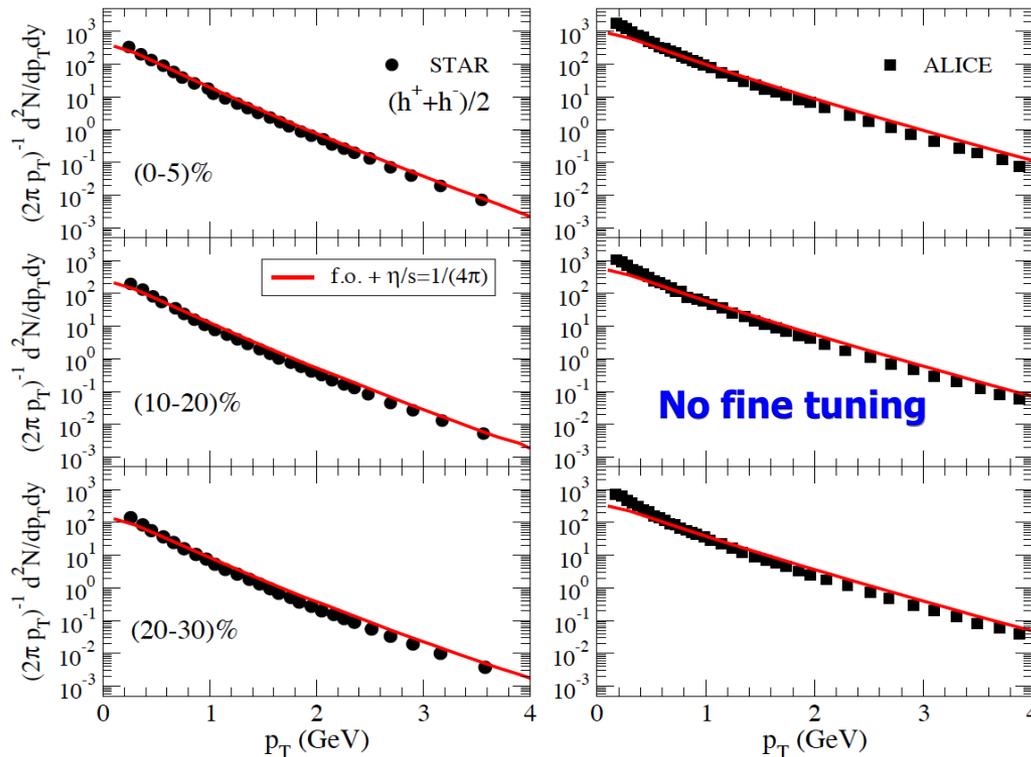
Bulk Initial Conditions

✧ r-space: standard Glauber model

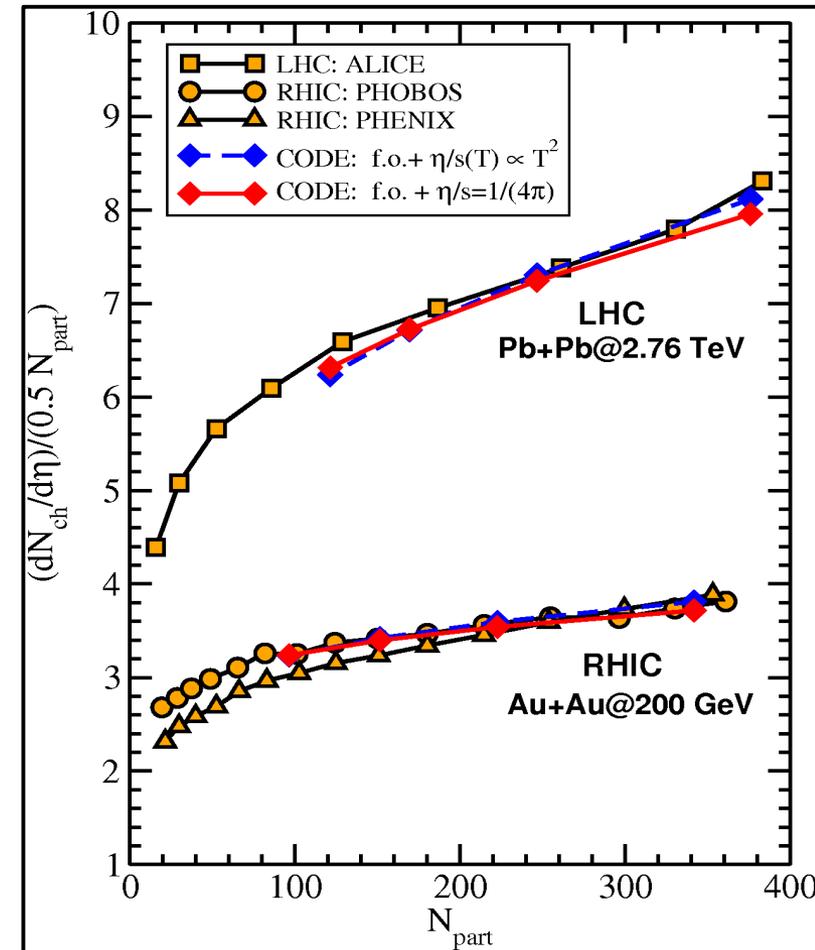
✧ p-space: Boltzmann-Juttner at T+ minijet [$p_T > 3\text{GeV}$]

Typical hydro condition

	62 GeV	200 GeV	2.76 TeV
T_0	290 MeV	340 MeV	510 MeV
τ_0	0.7 fm/c	0.6 fm/c	0.3 fm/c

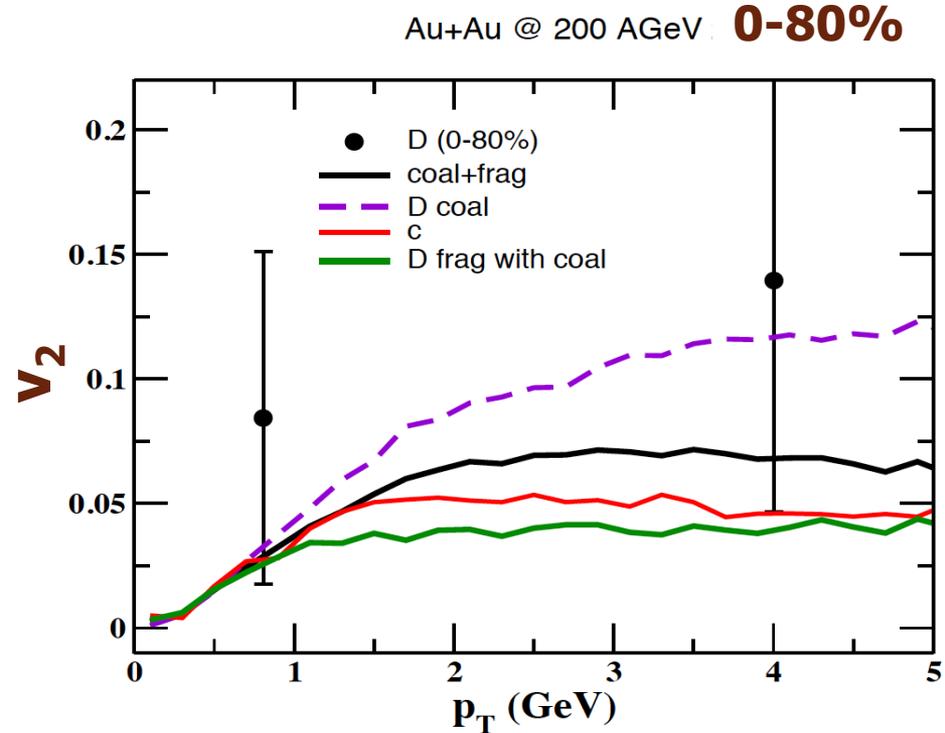
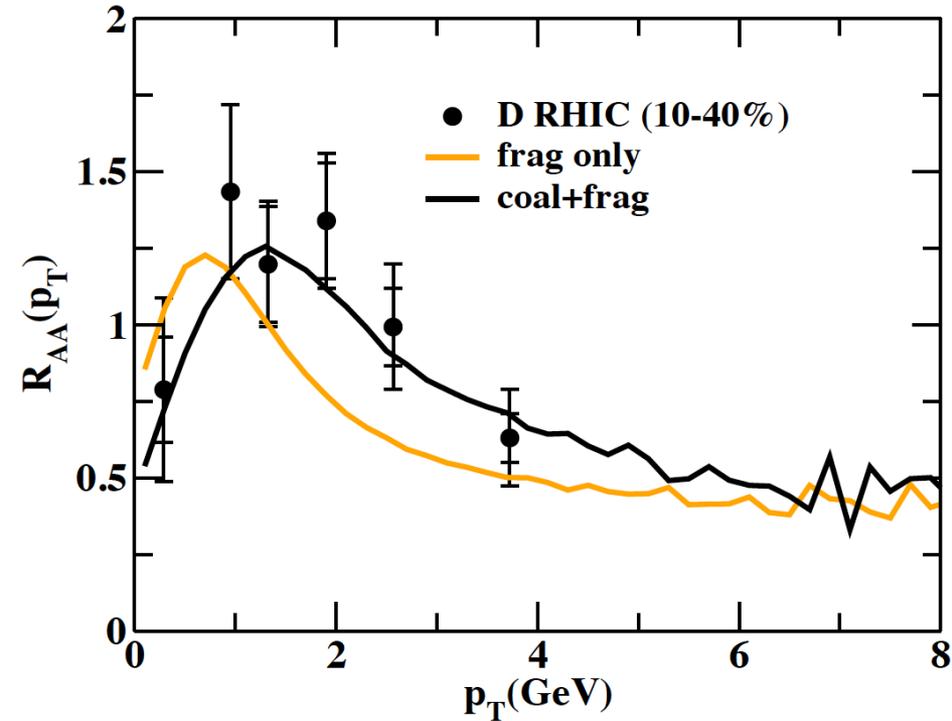


Spectra and multiplicity



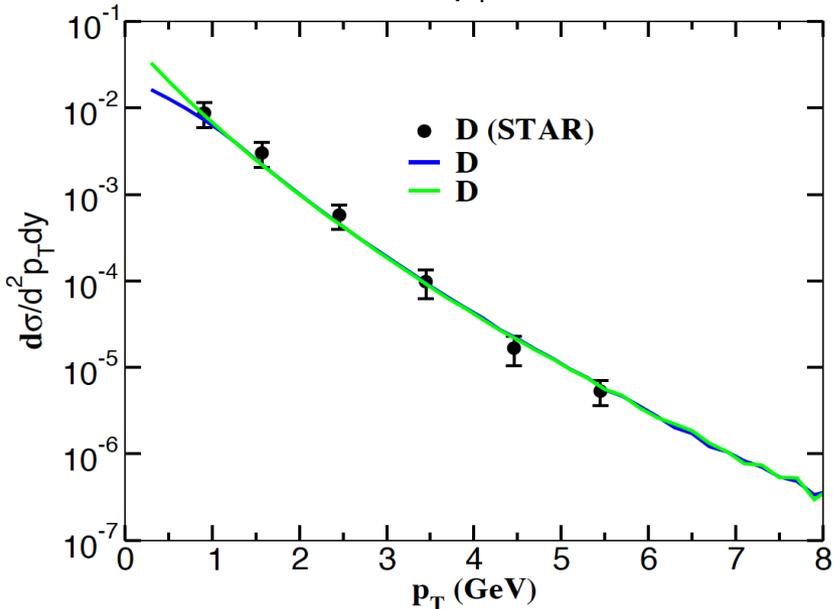
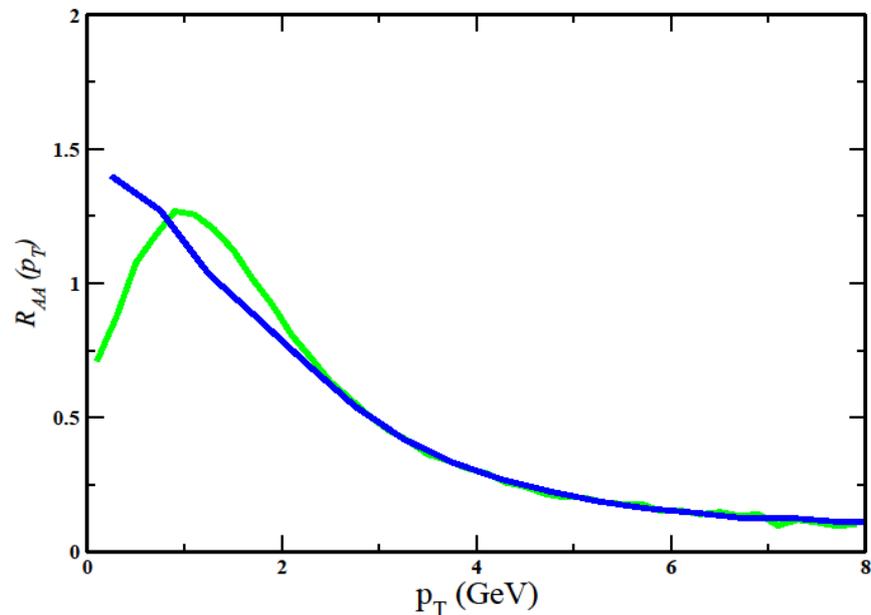
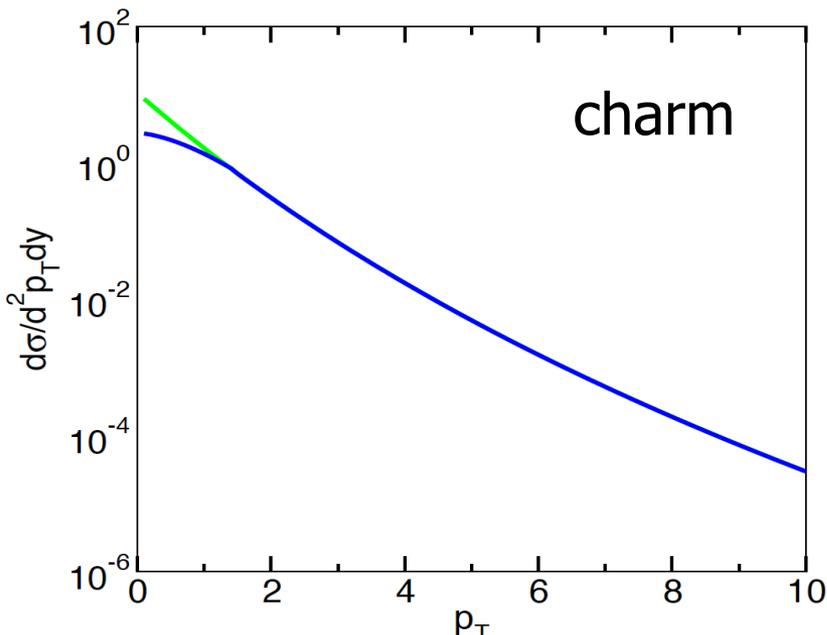
RHIC – D meson

QPM - Boltzmann

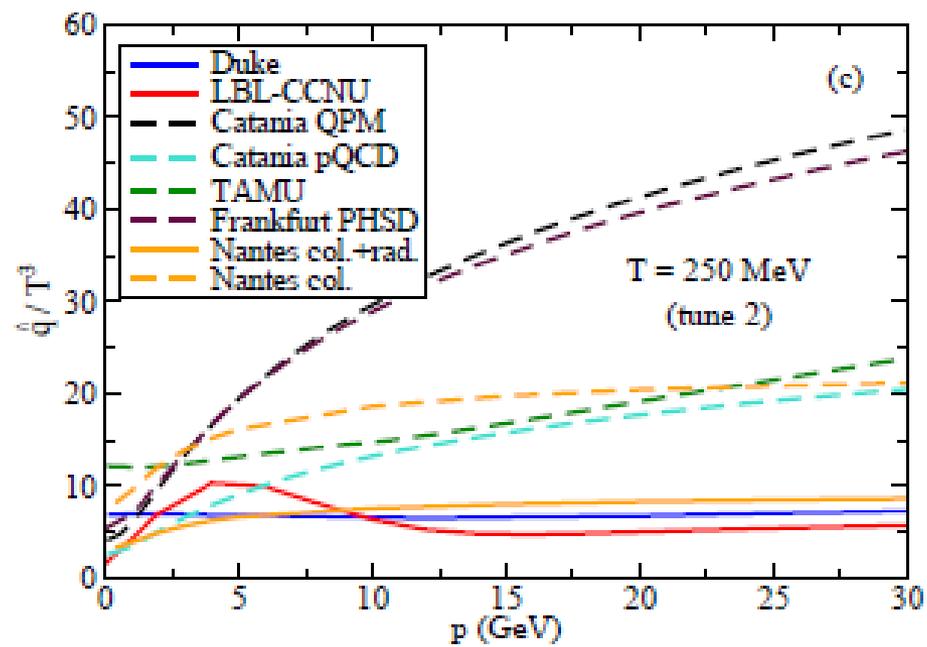


- ✧ No Hadronic Rescattering included
- ✧ Bump can be present also w/o coalescence
- ✧ Coalescence shift the bump

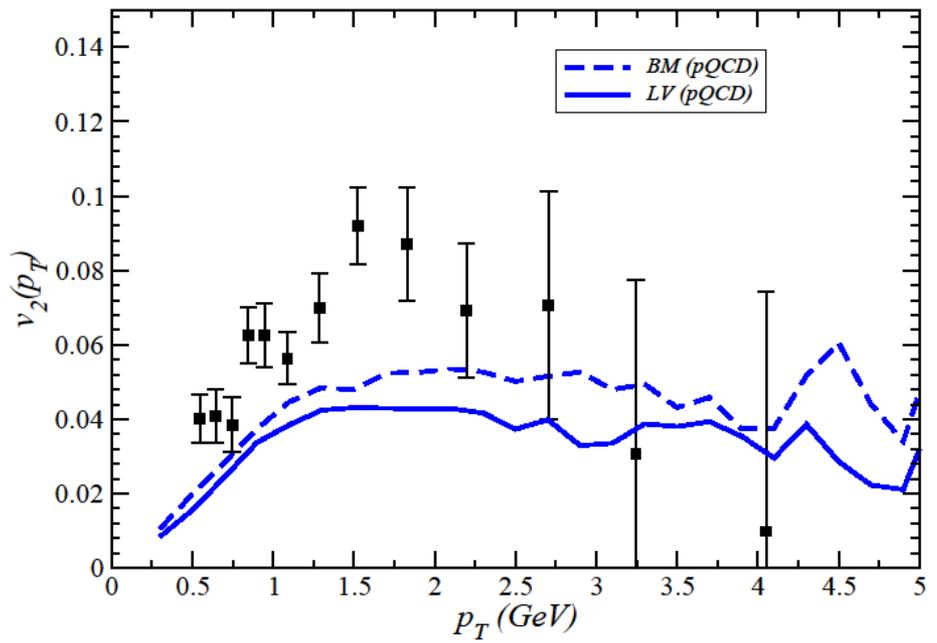
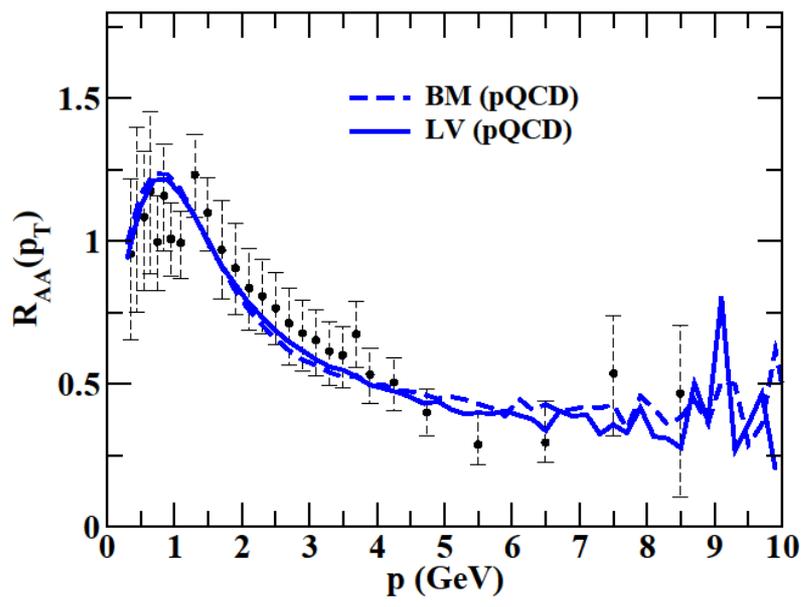
Impact of 2 dN/dp_T well within FONLL



Especially for the bump
Look at distribution and data
down to 0 p_T essential



R_{AA} and v_2 at RHIC at $mD=gT$



Das, Scardina, Plumari and Greco
PRC,90,044901(2014)

**At fixed RAA Boltzmann approach generate larger v_2 .
(depending on mD and M/T)**

I) LPM effect : Suppression of bremsstrahlung and pair production.

Formation length ($l_f = \frac{\eta}{q_\perp}$) : The distance over which interaction is spread out

- 1) It is the distance required for the final state particles to separate enough that they act as separate particles.
- 2) It is also the distance over which the amplitude from several interactions can add coherently to the total cross section.

As q_\perp increase $\rightarrow l_f$ reduce \rightarrow **Radiation drops proportional**

S. Klein, Rev. Mod. Phys 71 (1999)1501

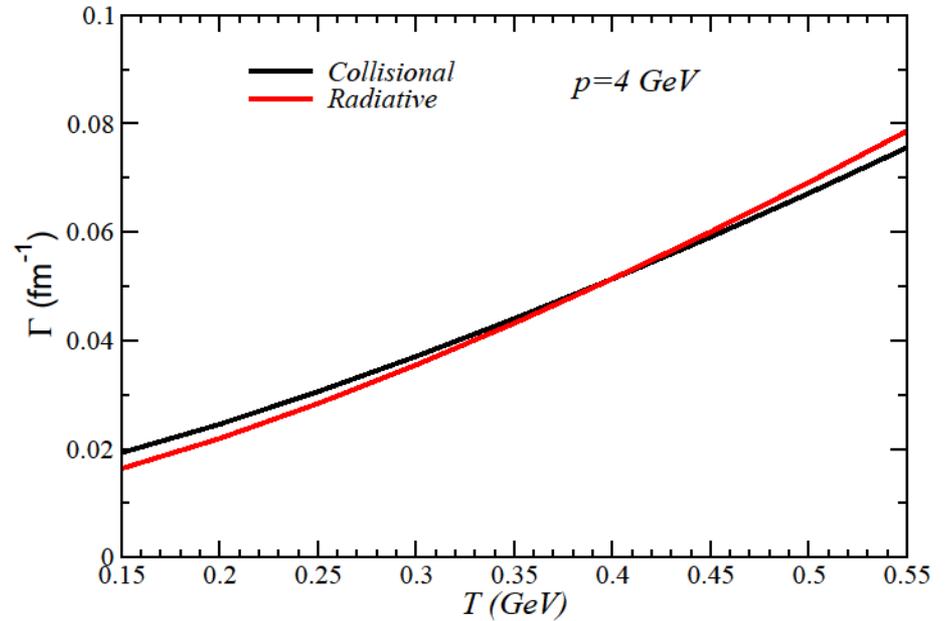
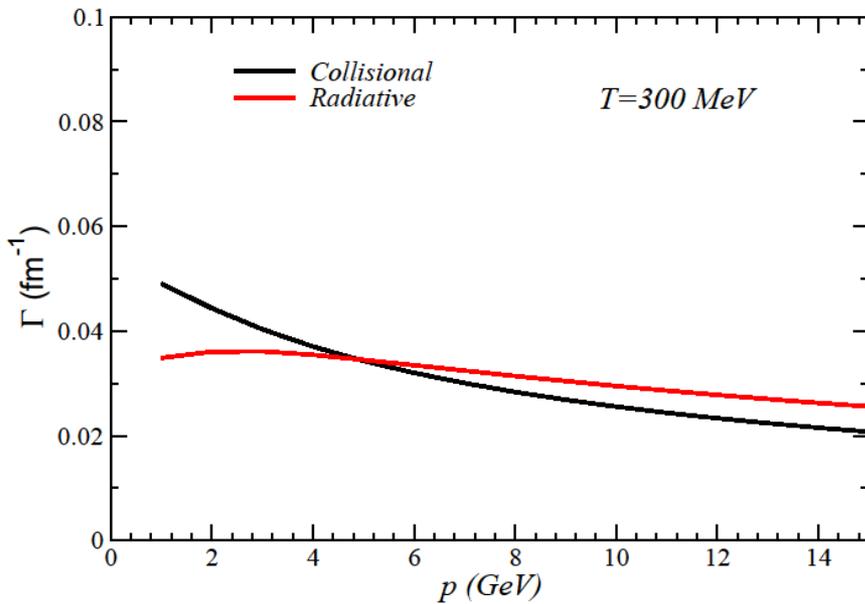
(II) Dead cone Effect : Suppression of radiation due to mass

$$\frac{1}{\sigma} \frac{d^2\sigma}{dzd\theta^2} \sim C_F \frac{\alpha_s}{\pi} \frac{1}{z} \frac{\theta^2}{(\theta^2 + 4\gamma)^2} \quad \text{where } z = 2 - x_1 - x_2 \quad \text{and} \quad \gamma = \frac{m^2}{s}$$

Where $x_1 = 2E_q / \sqrt{s}$ and $x_2 = 2E_{\bar{q}} / \sqrt{s} \rightarrow$ the energy fraction of the final state quark and anti-quark.

Radiation from heavy quarks suppress in the cone from $\theta = 0$ (minima) to $\theta = 2\sqrt{\gamma}$ (maxima)

Radiative vs Collisional



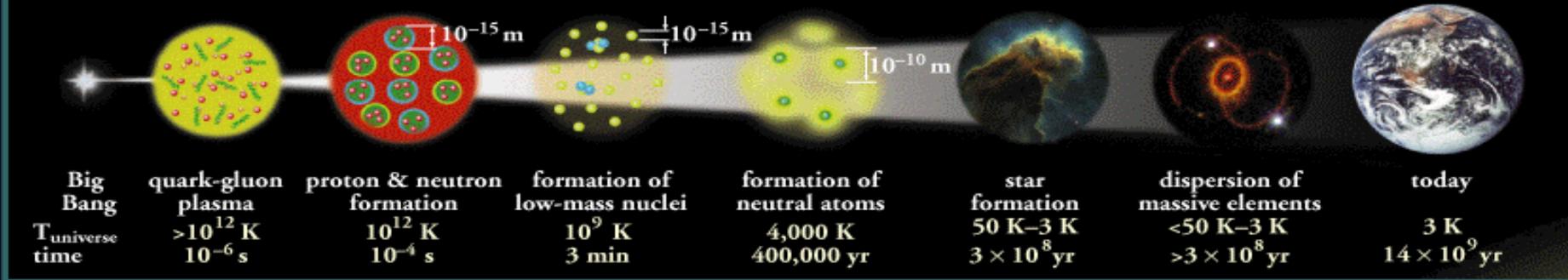
At High momentum
radiative loss dominate
over collisional loss

Das, Alam, Mohanty
PRC, 82,014908,2010

Introduction:

Expansion of the Universe

After the Big Bang, the universe expanded and cooled. At about 10^{-6} second, the universe consisted of a soup of quarks, gluons, electrons, and neutrinos. When the temperature of the Universe, T_{universe} , cooled to about 10^{12} K, this soup coalesced into protons, neutrons, and electrons. As time progressed, some of the protons and neutrons formed deuterium, helium, and lithium nuclei. Still later, electrons combined with protons and these low-mass nuclei to form neutral atoms. Due to gravity, clouds of atoms contracted into stars, where hydrogen and helium fused into more massive chemical elements. Exploding stars (supernovae) form the most massive elements and disperse them into space. Our earth was formed from supernova debris.



No direct way to look at the time before $\sim 400,000$ years

Nuclear collisions at high energy can create the micro second old universe.

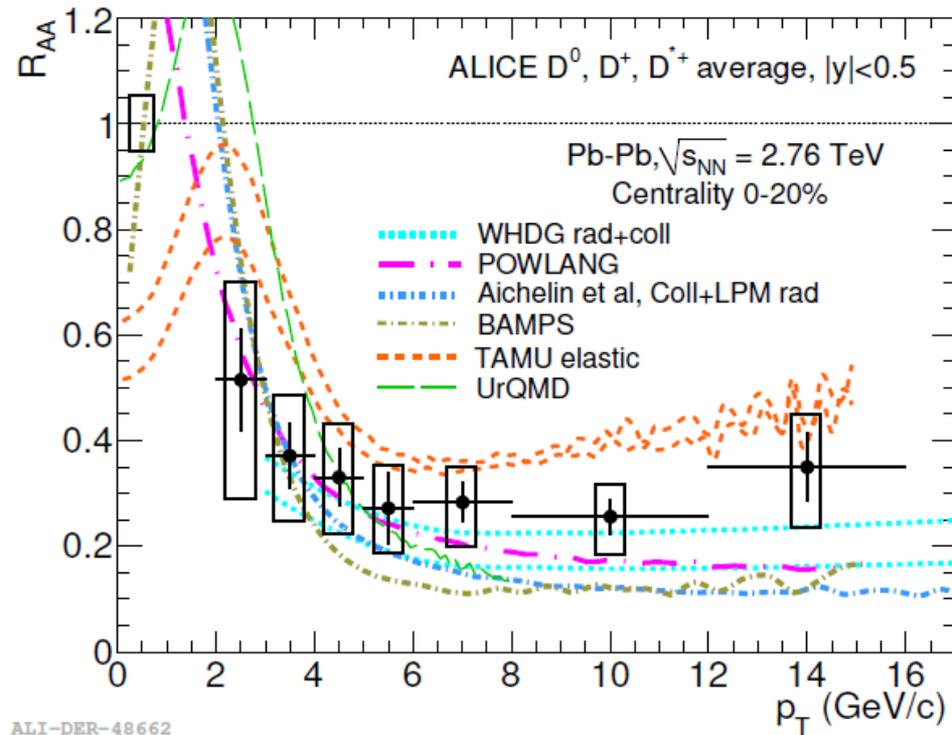
Quark-Hadrons transition occurred at $T \sim 170$ MeV (Lattice-QCD)

Sun Temperature \sim KeV

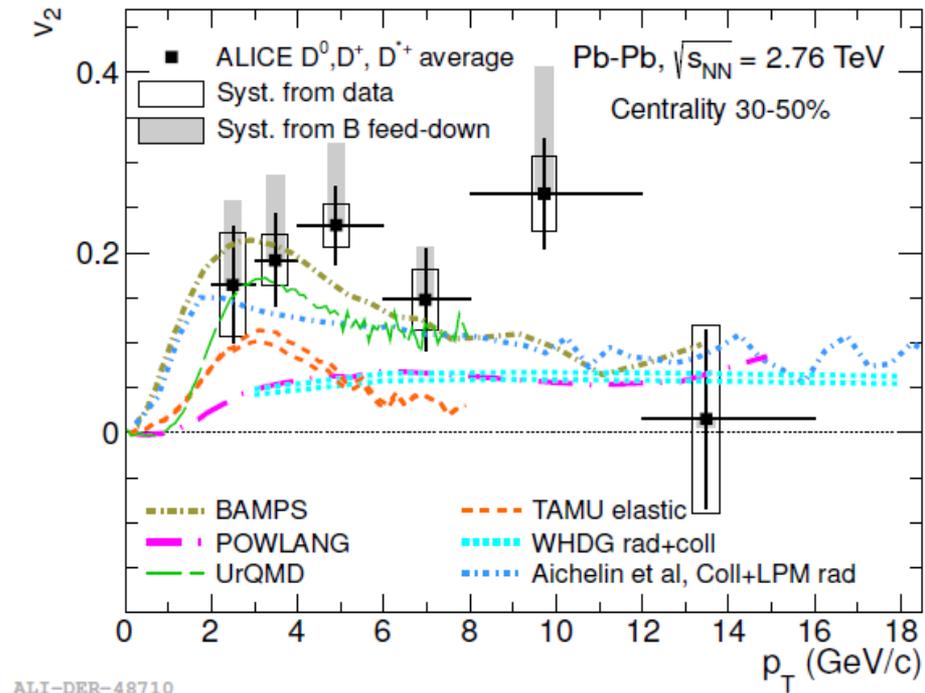
More than 10000 times of the temperature of sun !

Heavy Flavors at LHC

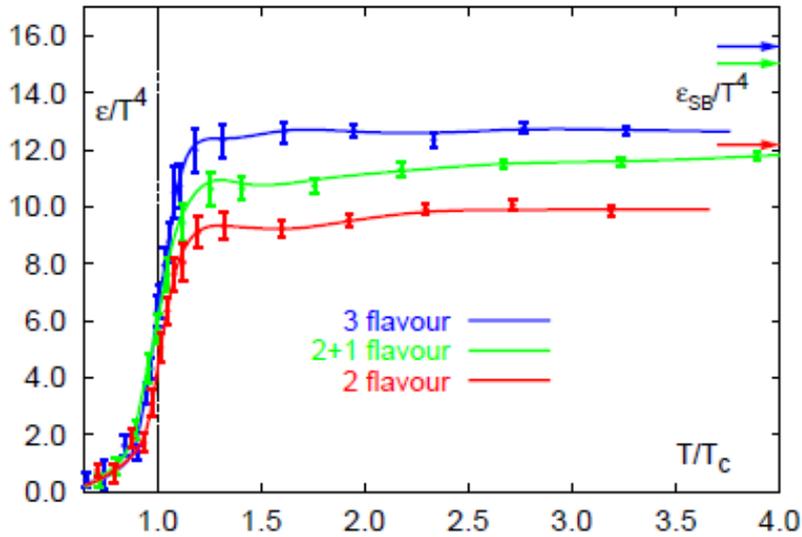
JHEP 1209 (2012) 112



arXiv:1305.2707

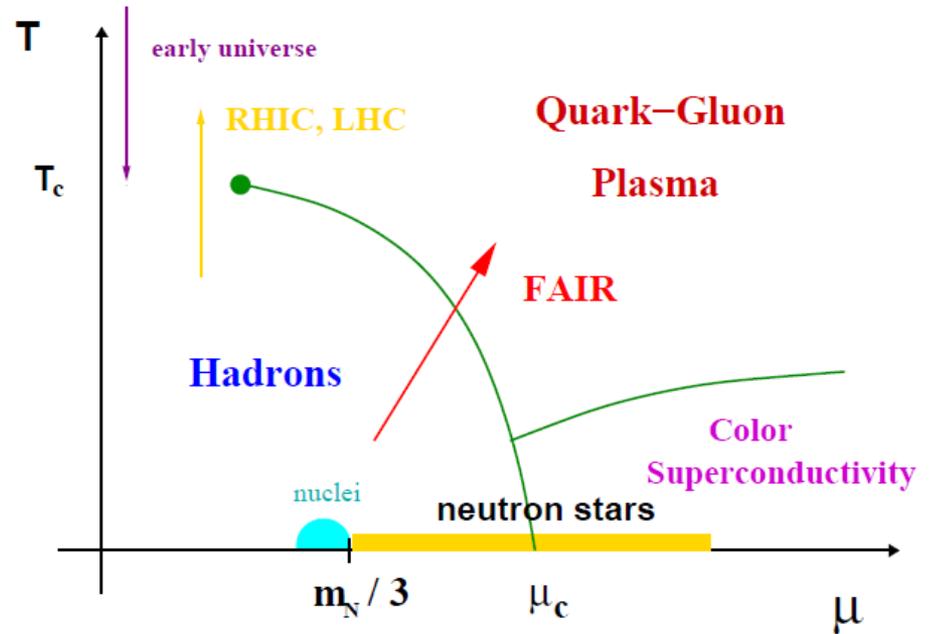


How to probe it



Energy density as a function of temperature scaled by the critical temperature T_c .

F. Karsch, Lect. Notes Phys.583 209 (2002)



Deconfined phases of QCD matter at two extreme conditions.