

\mathcal{A} im of our \mathcal{W} ork

- ▲ Propose a solution to the Moduli Stabilisation problem
- \blacktriangle Examine whether a dS vacuum exists in String Theory
- $(\cdots \text{ based on$ **perturbative** $quantum corrections only!})$
- ▲ If yes,

examine if slow roll inflation can be accommodated.

Outline of the Talk

- ▲ II-B/ \mathcal{F} -Theory bosonic spectrum and Moduli fields
- ▲ Effective Supergravity from type II-B & effective potential
- \blacktriangle D7 branes and logarithmic corrections
- ▲ Inflation
- ▲ Concluding Remarks

Type II-B/F-theory ★ Bosonic Spectrum - Moduli Space ▲ Graviton, dilaton and Kalb-Ramond (KR)-field $g_{\mu\nu}, \phi, B_{\mu\nu} \to B_2$ \blacktriangle Scalar, 2- and 4-index fields (*p*-form potentials) $\mathbf{C}_{\mathbf{0}}, C_{\mu\nu}, C_{\kappa\lambda\mu\nu} \rightarrow C_{p}, \ p = 0, 2, 4$ 1. $\land C_0, \phi \rightarrow combined to axion-dilaton modulus:$

$$S = C_0 + i e^{\phi} \rightarrow C_0 + rac{i}{g_s}$$

2. z_a : Complex Structure (CS) moduli (shape)

3. T_i : Kähler (*size*) moduli ($\tau_i = \text{Re}T$).

▲ ∃ plethora of moduli fields in CY compactifications → ... if massless → problems with fifth forces and other cosmological issues...

▲ Important Task ▲

▲ Generate a potential and assure positive mass-squared for all moduli fields \Rightarrow

 $\Rightarrow Moduli Stabilisation \leftarrow$

Type II-B effective Supergravity

Basic 'ingredients': Superpotential \mathcal{W} and Kähler potential \mathcal{K} \blacktriangle The Superpotential \mathcal{W}

▲ *Field strengths:*

 $F_p := d C_{p-1}, \ H_3 := d B_2, \ \Rightarrow G_3 := F_3 - SH_3$

▲ Holomorphic (3,0)-form: $\Omega(z_a)$

Flux-induced superpotential (*G.V.W.* hep-th/9906070):

$$\mathcal{W}_{\mathbf{0}} = \int \, \mathbf{G}_{\mathbf{3}} \wedge \mathbf{\Omega}(z_a)$$

▲ Supersymmetric conditions:

$$\mathcal{D}_{\boldsymbol{z}_{\boldsymbol{a}}} \mathcal{W} = 0, \quad \mathcal{D}_{\boldsymbol{S}} \mathcal{W} = 0 :$$

 $\Rightarrow z_a \ and \ S \ {f stabilised} \leftarrow {f but!}$

∧ Kähler moduli $\notin \mathcal{W}_0 \Rightarrow$ remain unfixed! ∧

▲ The Kähler potential ▲

$$\mathcal{K}_0 = -\sum_{i=1}^3 \ln(-i(T_i - \bar{T}_i)) - \ln(-i(S - \bar{S})) - \ln(i\int \Omega \wedge \bar{\Omega}) + \frac{1}{2} \ln(-i(T_i - \bar{T}_i)) - \ln(i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(T_i - \bar{T}_i))) - \ln(i(S - \bar{S})) - \ln(i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) - \ln(i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S - \bar{S}))) + \frac{1}{2} \ln(-i(S - \bar{S})) + \frac{1}{2} \ln(-i(S$$

 \blacktriangle The scalar potential \blacktriangle

$$V = e^{\mathcal{K}} \left(\sum_{I,J} \mathcal{D}_{I} \mathcal{W}_{0} \mathcal{K}^{I\bar{J}} \mathcal{D}_{\bar{J}} \mathcal{W}_{0} - 3 |\mathcal{W}_{0}|^{2} \right)$$
$$= e^{\mathcal{K}} \sum_{I,J=z_{a},\neq T_{i}} D_{I} \mathcal{W}_{0} \mathcal{K}^{-1}_{I\bar{J}} D_{\bar{J}} \mathcal{W}_{0} \quad (D_{I} \mathcal{W}_{0} = 0, \text{ flatness})$$
$$+ e^{\mathcal{K}} \left(\sum_{I,J=T_{i}} \mathcal{K}^{I\bar{J}}_{0} D_{I} \mathcal{W}_{0} D_{\bar{J}} \mathcal{W}_{0} - 3 |\mathcal{W}_{0}|^{2} \right) \quad (= 0, \text{ no scale})$$

Kähler moduli completely undetermined!

 \Rightarrow ...need to include **Quantum corrections ...** $f(\tau)$... "breaking" no-scale structure:

$$\mathcal{K} = -2\log\left(au^{rac{3}{2}} + \gamma f(au)
ight), \ \ \mathcal{V} = au^{rac{3}{2}}$$

Resulting F-term potential (γ -expansion):

$$V_F \propto \gamma \tau^{-\frac{9}{2}} \left(3f(\tau) - 4\tau f'(\tau) + 4\tau^2 f''(\tau)\right)$$

Some possible $f(\tau)$ functions:

 $\Delta \Delta \alpha$) power-law corrections (V_F: homogeneous) $f(\lambda \tau) = \lambda^n f(\tau)$)

$$f(au) \propto au^n \Rightarrow \left[V_F \propto au^{n-rac{9}{2}}
ight] \Rightarrow \nexists \ (V_F)_{min}$$

 $\Delta \beta$) logarithmic $f(\tau) \propto \log \tau :$

$$V_F \propto \gamma \tau^{-\frac{9}{2}} \left(\log(\tau) - \frac{8}{3} \right) + \cdots$$
 $\Rightarrow \exists (V_F)_{min} \forall \gamma < 0$

▲ Moreover, adding a constant $\xi = \gamma \log(\mu)$:

$$\mathcal{K} = -2\log\left(\mathcal{V} + \gamma\log(\mathcal{V}) + \boldsymbol{\xi} + \mathcal{O}(\frac{1}{\mathcal{V}})\right)$$
$$= -2\log\left(\mathcal{V} + \gamma\log(\mu\mathcal{V}) + \mathcal{O}(\frac{1}{\mathcal{V}})\right)$$

At $(V_F)_{min}$ volume size controlled by μ :

$$\mathcal{V}_{min} = \frac{1}{\mu} e^{\frac{13}{3}}$$

 \Rightarrow large volume expansion for $\mu = e^{-\frac{\xi}{|\gamma|}} \ll 1$

Origin of ξ and $f(\tau)$ from $\mathcal{PERTURBATIVE}$ String Loop Corrections

Quantum Corrections

1. α' -corrections

shifting volume by a constant (Becker et al, hep-th:0204254)

$$\mathcal{V} \to \mathcal{V} + \boldsymbol{\xi}, \quad \boldsymbol{\xi} = -\frac{\zeta(3)}{4(2\pi)^3}\boldsymbol{\chi}$$

2. Perturbative string loop-corrections important in the presence of *D*-branes (Antoniadis, Chen, G.K.L.: hep-th/1803.08941 & to appear)

▲ ▲ String Coupling Loop Corrections ▲ ▲

In String Theory:

Multigraviton scattering generates higher derivative couplings in curvature

Leading term in type II-B:

 $\propto R^4$

Reduction on $\mathcal{M}_4 \times \mathcal{X}_6$, (with \mathcal{M}_4 4-d Minkowski) induces:

localised Einstein Hilbert (\mathcal{EH}) term \propto Euler characteristic

$$\frac{1}{3!(2\pi)^3} \chi = \int R \wedge R \wedge R$$

 \blacktriangle this \mathcal{EH} term possible in 4-dimensions only!

Localised vertices can emit gravitons and KK-excitations in 6d \Rightarrow KK-exchange between graviton vertices and D7-branes



Non-zero contribution from 1-loop; 3-graviton scattering amplitude 2 massless 1 KK:

Correction
$$\propto \chi \int_{M_4} (1 - \sum_{i=1,2,3} e^{\phi} T_i \log(\mathbf{R}^i_{\perp})) \mathcal{R}_{(4)}$$

▲ Analogously, the corrections to the Kähler potential involve the corresponding Kähler moduli and have the form:

$$\boldsymbol{\delta} = \sum_{i=1}^{3} \gamma_i \log(\tau_i)$$

$$\mathcal{C} = -2\ln\left[e^{-2\phi}(\mathcal{V} + \boldsymbol{\xi}) + \delta\right]$$
$$= -\ln\left[-i(S - \bar{S})\right] - 2\ln\left(\hat{\mathcal{V}} + \hat{\boldsymbol{\xi}} + \hat{\delta}\right)$$

with $S = b + i e^{-\phi}$, $\hat{\delta} = \delta g_s^{1/2}$

 \blacktriangle ξ and δ break **no-scale** structure of Kähler potential

Stabilisation and D7 Branes

Looking for the minimum number of D7 branes required to stabilise the T_i -fields and lead to a dS minimum.







▲ Kähler potential including loop corrections: from $3 \times D7$:

$$\mathcal{K} = -2\ln\{\mathcal{V} + \sum_{k=1}^{3} \gamma_k \ln(\mathcal{V}/\tau_k)\}$$

F-term potential (assuming $\gamma_k \to \gamma, \xi = 0$):

$$V_F \approx 3\gamma rac{\ln \mathcal{V} - 4}{\mathcal{V}^3} + \mathcal{O}(\gamma^3)$$

A Inclusion of \mathcal{D} -terms:

$$\mathcal{V}_{\mathcal{D}} = \sum_{a} \frac{d_{a}}{\tau_{a}} \left(\frac{\partial \mathcal{K}}{\partial \tau_{a}}\right)^{2} \approx \sum_{a} \frac{d_{a}}{\tau_{a}^{3}} + \cdots$$

Minimisation of $V_{\text{eff}} = V_F + V_D$ w.r.t.:

$$au_1, au_2 \ and \ \mathcal{V} = \sqrt{ au_1 au_2 au_3} \Rightarrow V_{ ext{eff}} \propto \gamma rac{\ln \mathcal{V} - 4}{\mathcal{V}^3} + rac{d}{\mathcal{V}^2}$$



vertical line represents any value of $z_m = \frac{2d}{3\gamma}e^{\frac{13}{3}}$ between $(-e^{-1}, 0)$ where min and max can coexist.



but! requirement for de Sitter vacua puts additional restrictions

\blacktriangle de Sitter vacua

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minimum $V_{\text{eff}} = V_F + V_D$ at \mathcal{V}_0 must be positive:

$$V_{\text{eff}}^{min} = \frac{\gamma}{\mathcal{V}_0^3} + \frac{d}{\mathcal{V}_0^2} > 0 \rightarrow -7.24 < 10^3 \frac{d}{\gamma} < -6.74$$



INFLATION I. Antoniadis, Yifan Chen, GKL (hep-th/1809.05060) Although a dS minimum exists, allowed region too restrictive. Additional requirements for slow-roll inflation hard to be met.

two ways out of this impasse

 \mathcal{A}) Fayet-Iliopoulos (FI) contributions \rightarrow uplift of V_{eff}

 \land Novel ways to construct FI-terms suggested in 1712.08601

 \mathcal{B}) introducing a nilpotent field $X, X^2 = 0$:

$$\Delta \mathcal{W} = fX, \ \Delta \mathcal{K} = \frac{X\bar{X}}{\tau_1^{n_1}\tau_2^{n_2}\tau_3^{n_3}},$$

for
$$n_i = n \to V_{up} = \frac{f^2}{\mathcal{V}^{2-2n}}$$

▲ Total scalar potential

$$V_{total} = V_F + V_{\mathcal{D}} + V_{up}$$

 $normalised\ kinetic\ terms\ require\ redefinition\ of\ moduli\ fields:$

$$t_i = \frac{1}{\sqrt{2}} \ln \tau_i$$

Inflaton is associated with volume \mathcal{V} :

$$t = \frac{t_1 + t_2 + t_3}{\sqrt{3}} = \sqrt{\frac{2}{3}}\ln(\mathcal{V}),$$

minimum and maximum w.r.t t:

$$t_{min/max} = \sqrt{\frac{2}{3}} \left(\frac{13}{3} - \mathcal{W}_{0/-1} \left(\frac{2}{3} \frac{d}{\gamma} e^{\frac{13}{3}} \right) \right)$$
(1)

Redefining remaining two moduli fields to be "orthogonal" to *t*:

$$u = \frac{t_1 - t_2}{\sqrt{2}}, v = \frac{t_1 + t_2 - 2t_3}{\sqrt{6}}$$

Inflation period should be constrained between t_{max}, t_{min}

Constraints from slow-roll conditions

Ending point of inflation $t_{min} < t_{end} < t_{max}$:

$$t_{end} = max\{t|_{\frac{1}{2}(\frac{V_t}{V})^2 \simeq 1}, t|_{\frac{|V_{tt}|}{V} \simeq 1}\},\$$

Finding point t_* such that $t_{min} < t_{end} < t_* < t_{max}$ where e-folds:

$$N_* = \int_{t_{end}} t_* \frac{V}{V_t} dt \sim 50 - 60$$

Spectral index at t_* :

$$n_s = 1 - 3\left(\frac{V_t}{V}\right)^2 + 2\frac{V_{tt}}{V}\bigg|_{t_*} \approx 0.96$$



Plot of V_{eff} vs \mathcal{V} . Inflation occurs between t_* and t_{end}



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\bigstar IIB/F-theory:

• Stabilisation of Kähler MF possible with Perturbative Corrections:

 $\mathcal{K} = -2\ln\left(\mathcal{V} + \boldsymbol{\xi} + \gamma_i \ln \tau_i\right)$

Origin of log-corrections:

Induced Einstein-Hilbert terms from R^4 -couplings in 10-d theory. This \mathcal{EH} -term \exists in 4d only!

\downarrow

In the present context, *induced* EH-term ... indispensable element for a 4d de Sitter Universe





Type II 10-d effective action with $\mathcal{EH} \& \mathbb{R}^4$ terms: (see hep-th/9704145; hep-th/9707013, hep-th/9707018)

$$S \supset \frac{c}{l_s^8} \int_{M_{10}} e^{-2\phi} \mathcal{R}_{(10)} + \frac{d}{l_s^2} \int_{M_{10}} (-2\zeta(3)e^{-2\phi} + 4\zeta(2)) R^4 \wedge e^2$$

Compactifying on CY they give rise to localised EH-term in 4-d:

A few details on localised graviton vertices

• Tree-level contribution to \mathcal{EH} -term vanishes in orbifold background $CY = T^6/Z_N$.

• 1-loop contribution from SUSY preserving $\mathcal{N} = (1, 1)$ odd-odd spin structure of partition function

$$Z_{odd}^{(1,1)} \to \sum_{f=0,\dots,n_f} \chi_f \equiv \chi$$

 $(0,\ldots,n_f = fixed points)$

1-loop amplitude: two massless gravitons and one KK-excitation. (in odd-odd spin structure \rightarrow one graviton vertex must be taken in (-1, -1)-ghost picture.) localisation property of w/f of $R_{(4)}$. Amplitude:

$$\langle V_{(0,0)}^2 V_{(-1,-1)} \rangle = -C_{\mathcal{R}} \frac{1}{N^2}$$

$$\sum_{\substack{f=0,\dots,n_f\\k=0,\dots,N-1}} e^{i\gamma^k q \cdot x_f} \int_{\mathcal{F}} \frac{d^2 \tau}{\tau_2^2}$$

$$\int \prod_{i=1,2,3} \frac{d^2 z_i}{\tau_2} \sum_{(h,g)} e^{\alpha' q^2 F_{(h,g)}(\tau,z_i)}$$

- $\mathcal{C}_{\mathcal{R}}$ tensor structure,
- $F_{(h,g)}(\tau, z_i) \rightarrow \text{twisted sectors } (f,g) = (l,m)\frac{v}{N}$ Fourier transform \Rightarrow Gaussian profile:

$$\rightarrow \frac{N}{w^6} e^{-\frac{y^2}{2w^2}}, \ w^2 \sim \alpha' F_{(f,g)} \sim \frac{\ell_s^2}{N} \rightarrow \text{eff.width}$$