



Contribution ID: 109

Type: Poster

Implications of a transition in the dark energy equation of state for the H_0 and σ_8 tensions

We explore the implications of a rapid appearance of dark energy between the redshifts (z) of one and two on the expansion rate and growth of perturbations. Using both Gaussian process regression and a parametric model, we show that this is the preferred solution to the current set of low-redshift ($z < 3$) distance measurements if $H_0 = 73 \text{ km s}^{-1} \text{ Mpc}^{-1}$ to within 1% and the high-redshift expansion history is unchanged from the Λ CDM inference by the Planck satellite. Dark energy was effectively non-existent around $z = 2$, but its density is close to the Λ CDM model value today, with an equation of state greater than -1 at $z < 0.5$. If sources of clustering other than matter are negligible, we show that this expansion history leads to slower growth of perturbations at $z < 1$, compared to Λ CDM, that is measurable by upcoming surveys and can alleviate the σ_8 tension between the Planck CMB temperature and low-redshift probes of the large-scale structure.

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Session Classification: Parallel Sessions: Modified Gravity and Dark Energy (C.A.R.L., H03)

Track Classification: Modified Gravity and Dark Energy