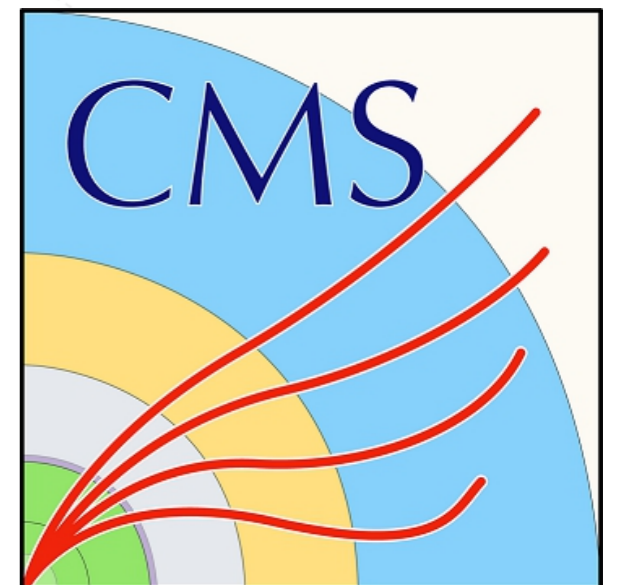


# BSM Results at the LHC

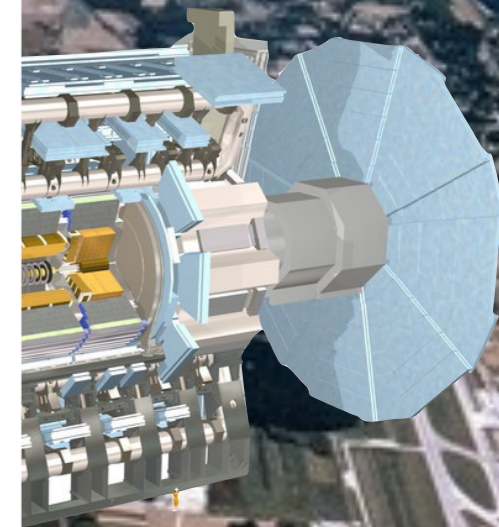
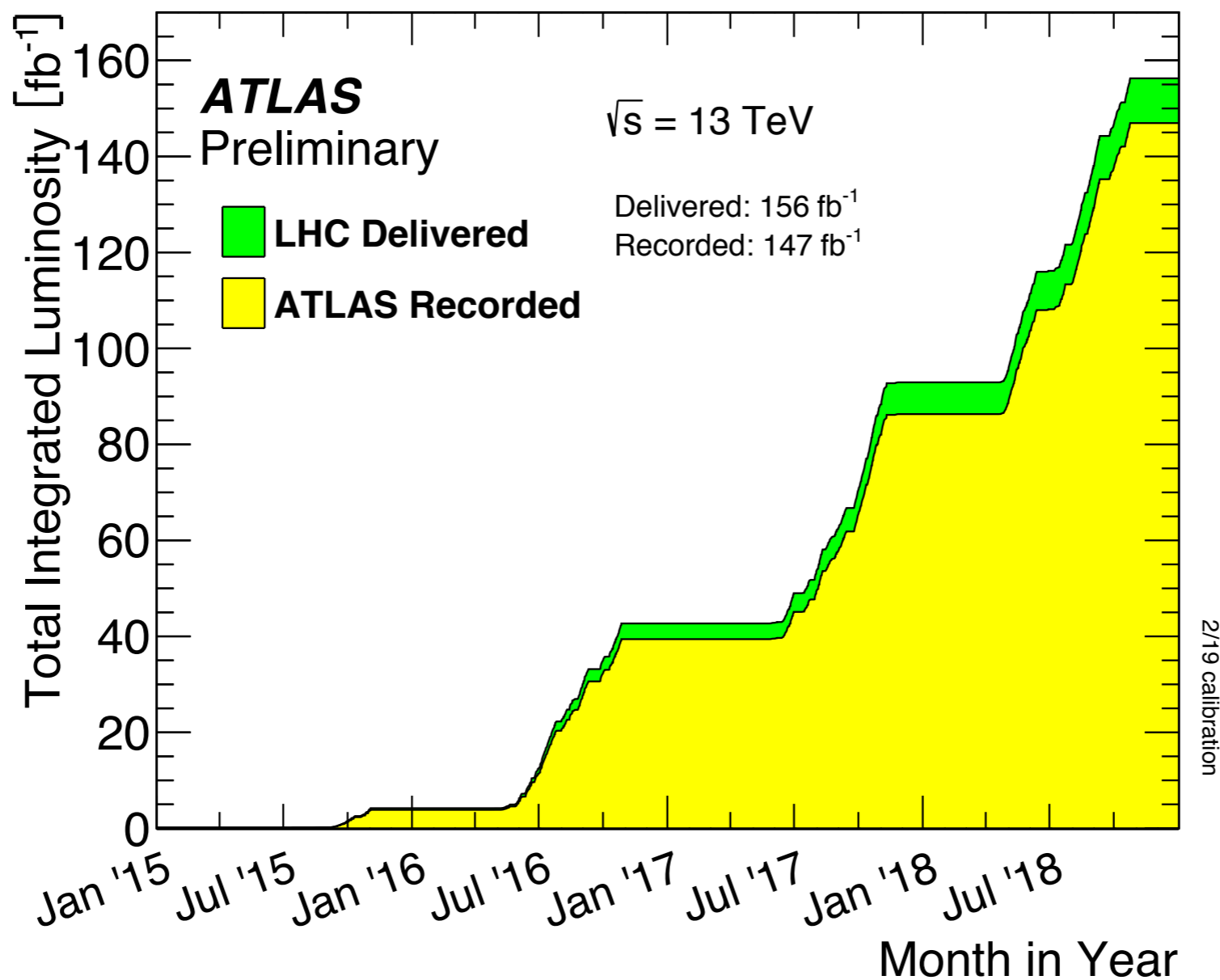
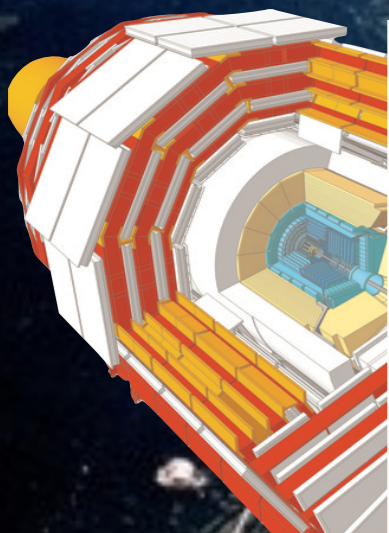
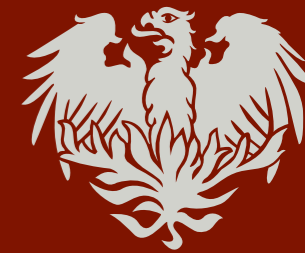
*APS DPF 2019*

Maximilian Swiatlowski,  
for the ATLAS and CMS Collaborations

Enrico Fermi Institute, University of Chicago

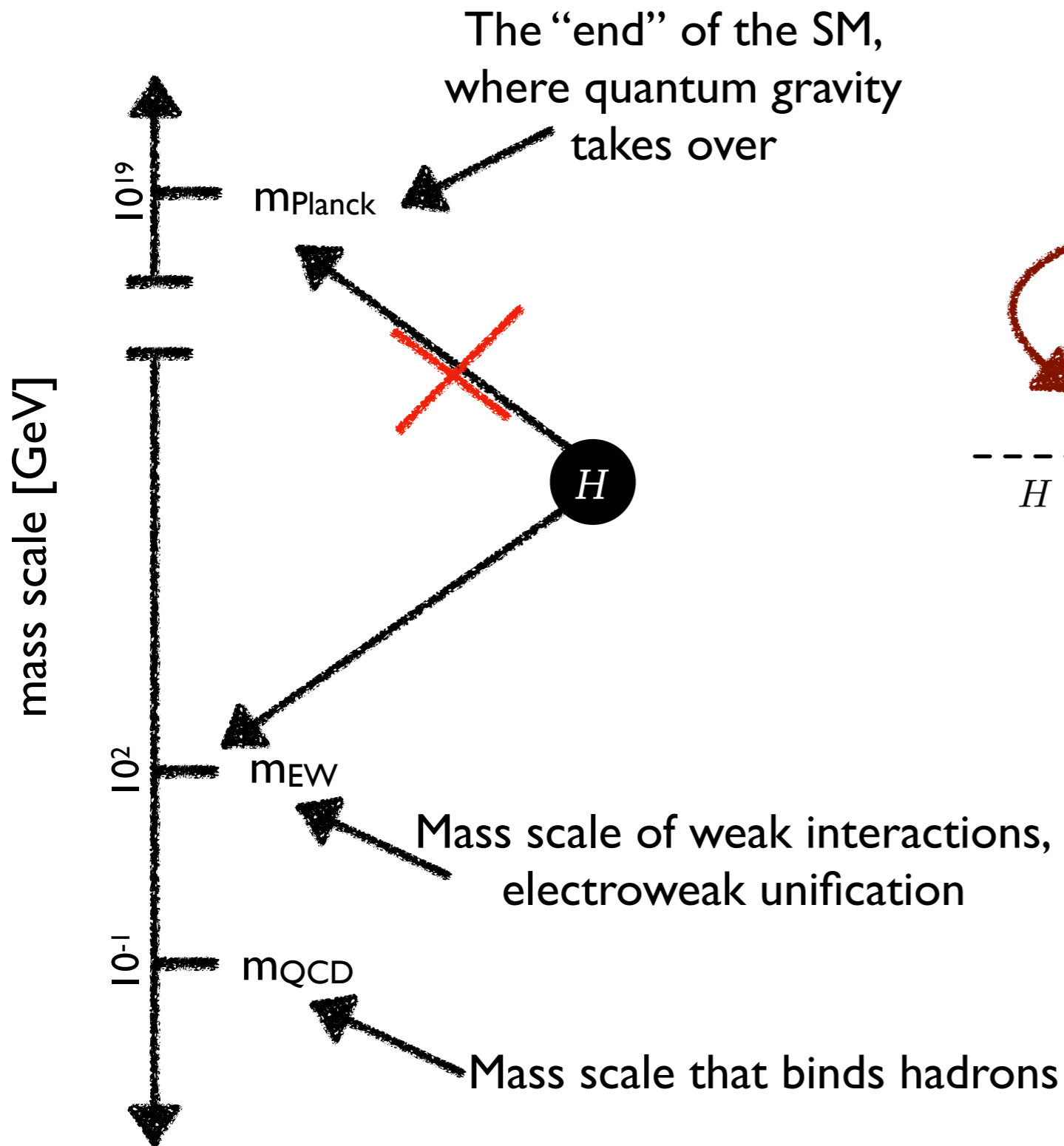


# The LHC



Huge recorded dataset from Run 2:  
Thank you to the LHC team and machine!

# Why BSM?



To first order, the mass of the Higgs is a free parameter, and can be anywhere

But the Higgs interacts with other particles, which affects its mass

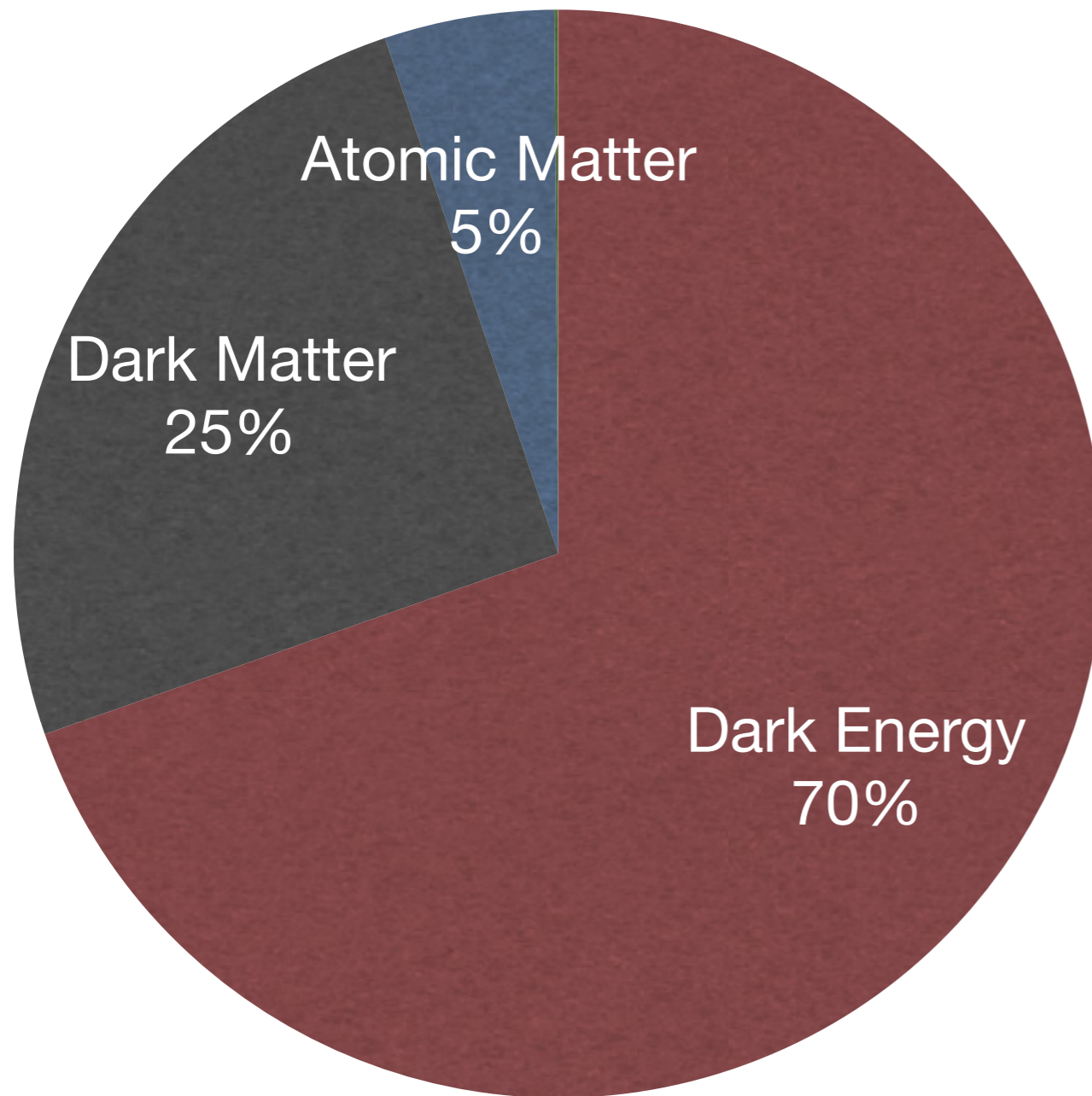
$$\text{---} H \text{---} \text{---} \text{---} \approx \int_0^\Lambda dE \approx m_{\text{Planck}}$$

In the SM, this correction should set the Higgs mass to  $\sim m_{\text{Planck}}$

But we observe it at 125 GeV!

**BSM can stabilize the Higgs mass**

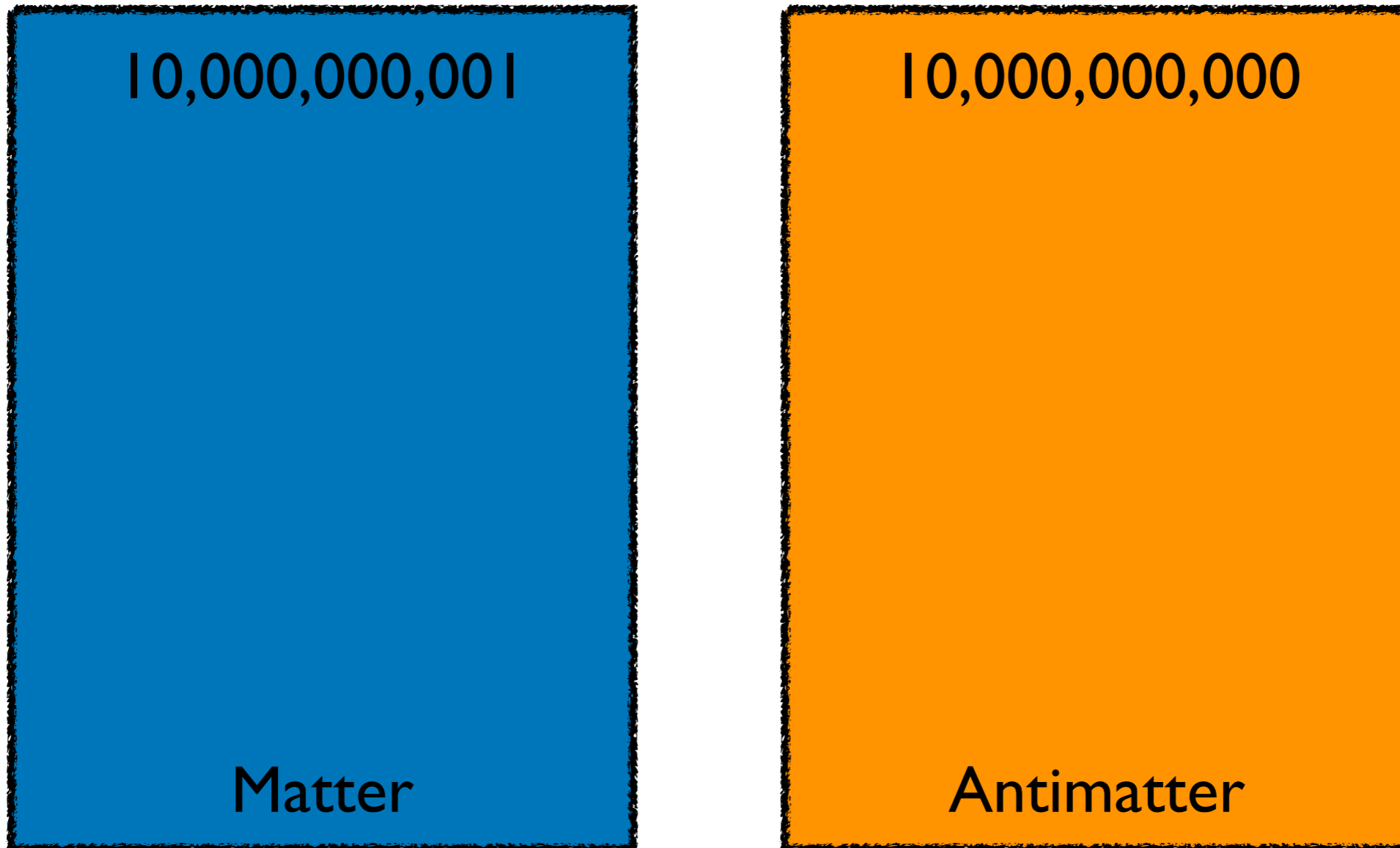
# Why BSM?



We don't understand the majority of the universe!

**BSM can provide Dark Matter candidates, which we could produce at the LHC**

# Why BSM?

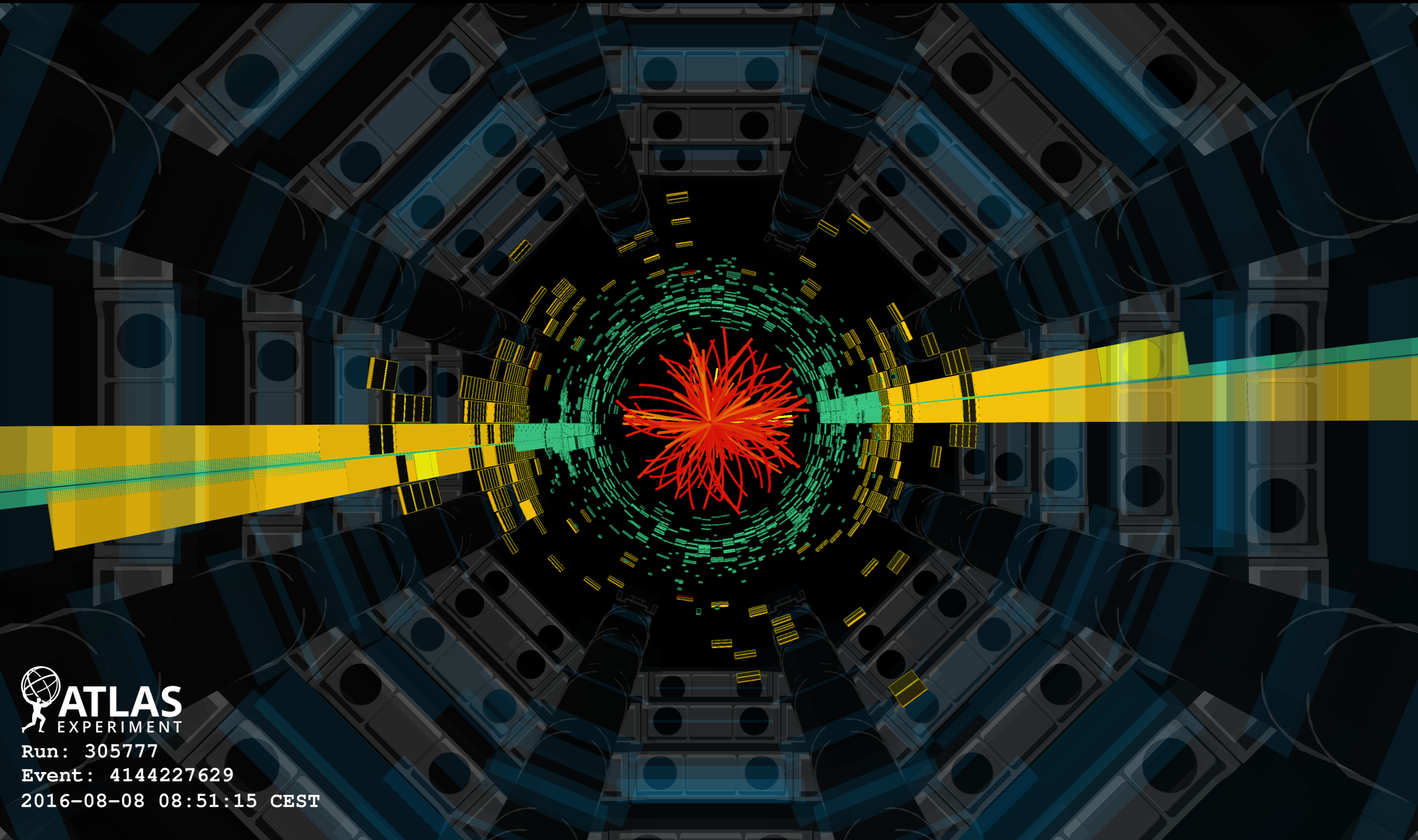


Where did the tiny matter/antimatter asymmetry that leads to a matter dominated universe arise?

CP violation in the SM is not enough

**BSM can explain the matter dominated universe**

# How do we search for BSM physics at the LHC?



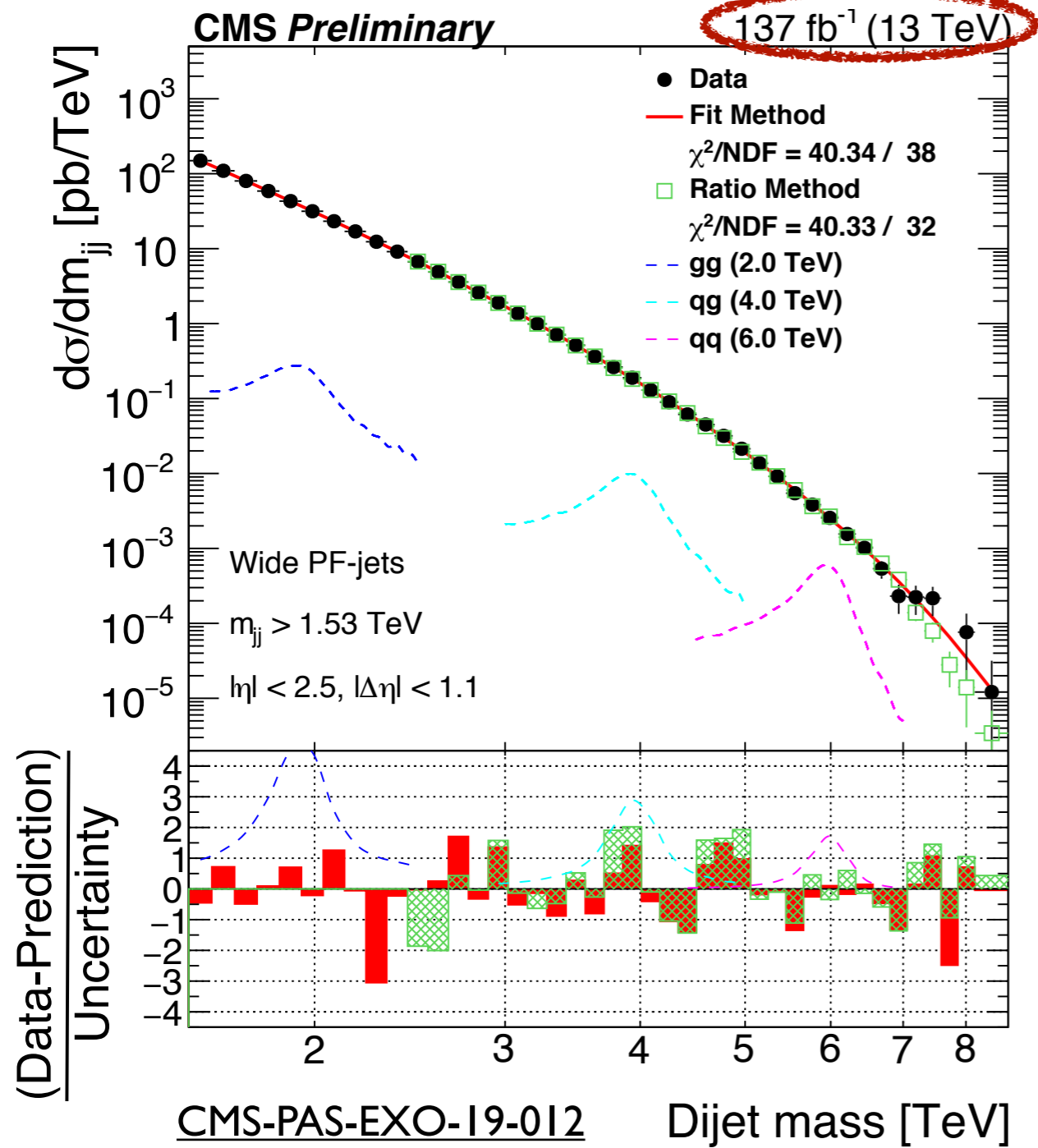
 **ATLAS**  
EXPERIMENT

Run: 305777

Event: 4144227629

2016-08-08 08:51:15 CEST

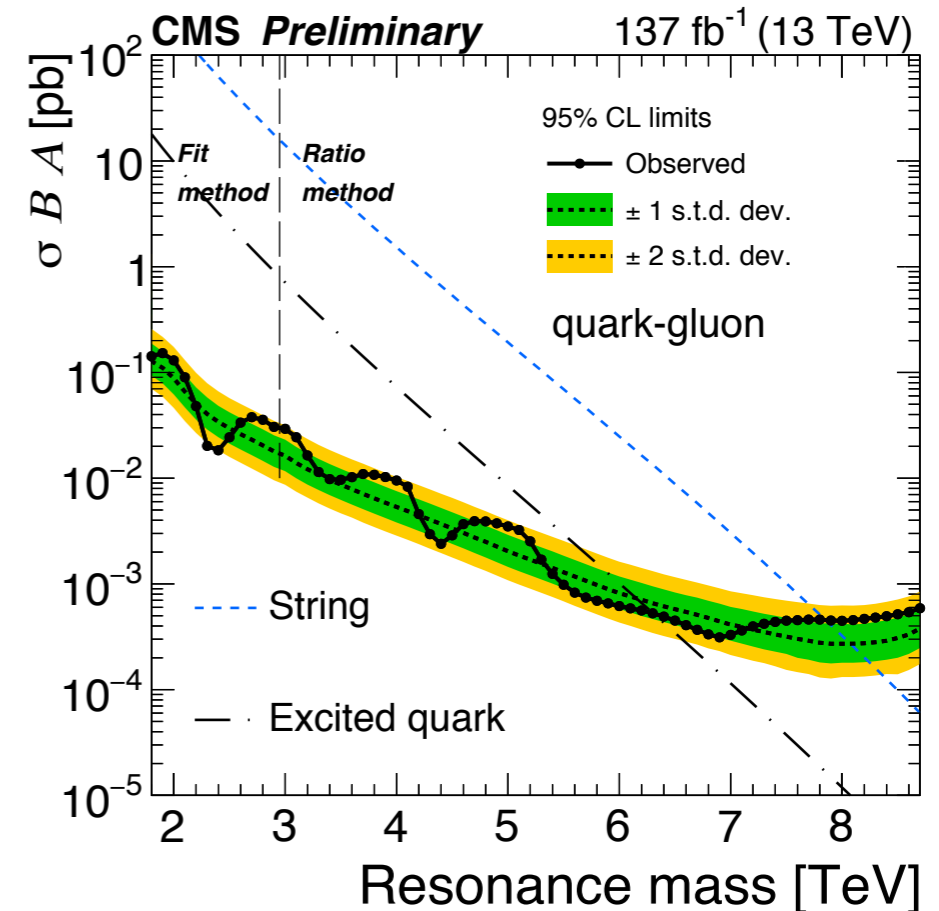
# Searching with Dijets



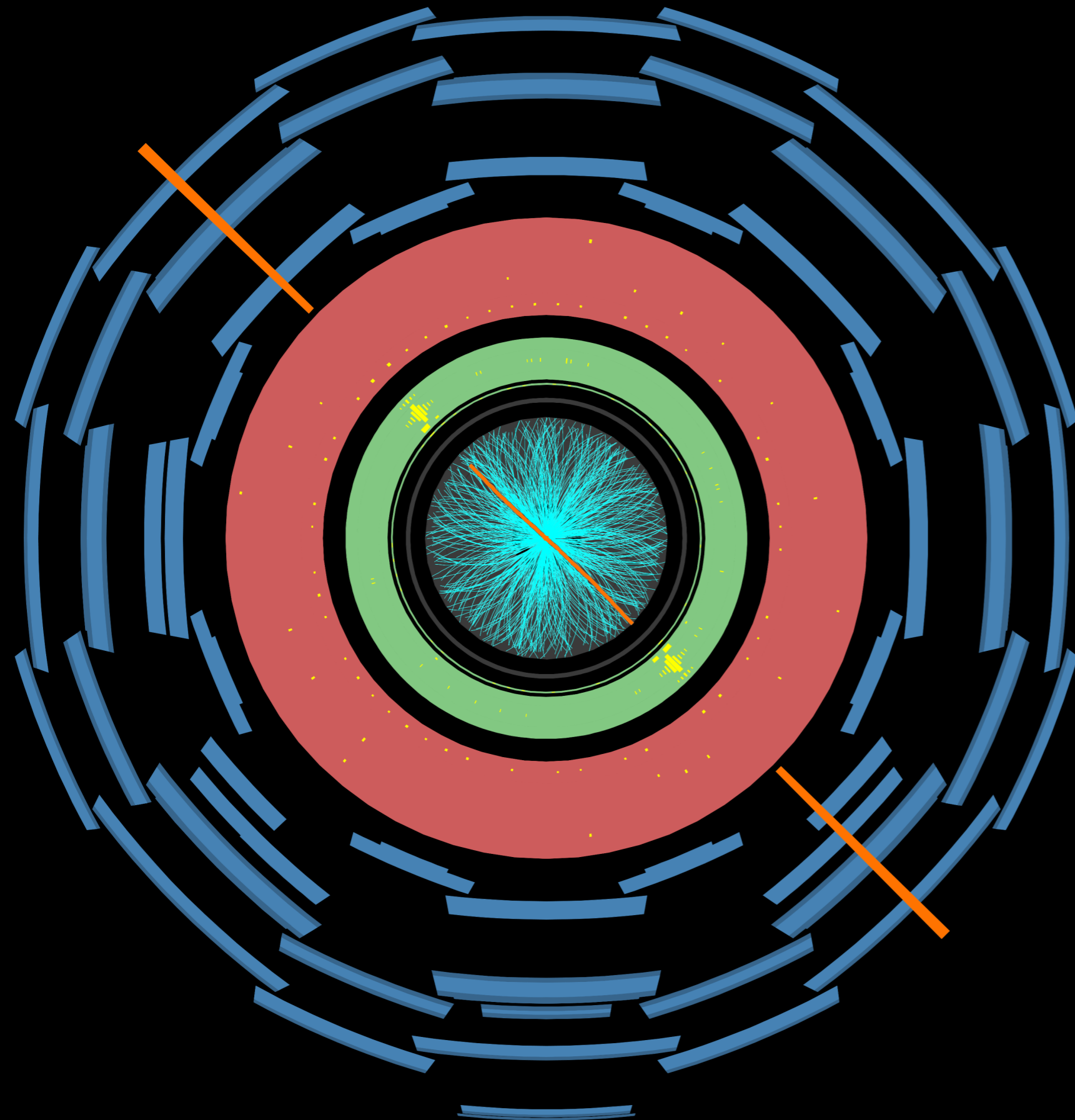
One of the best BSM signatures:  
 pairs of jets

Huge dataset enables some of  
 the best sensitivity yet!

But no hints of new physics...



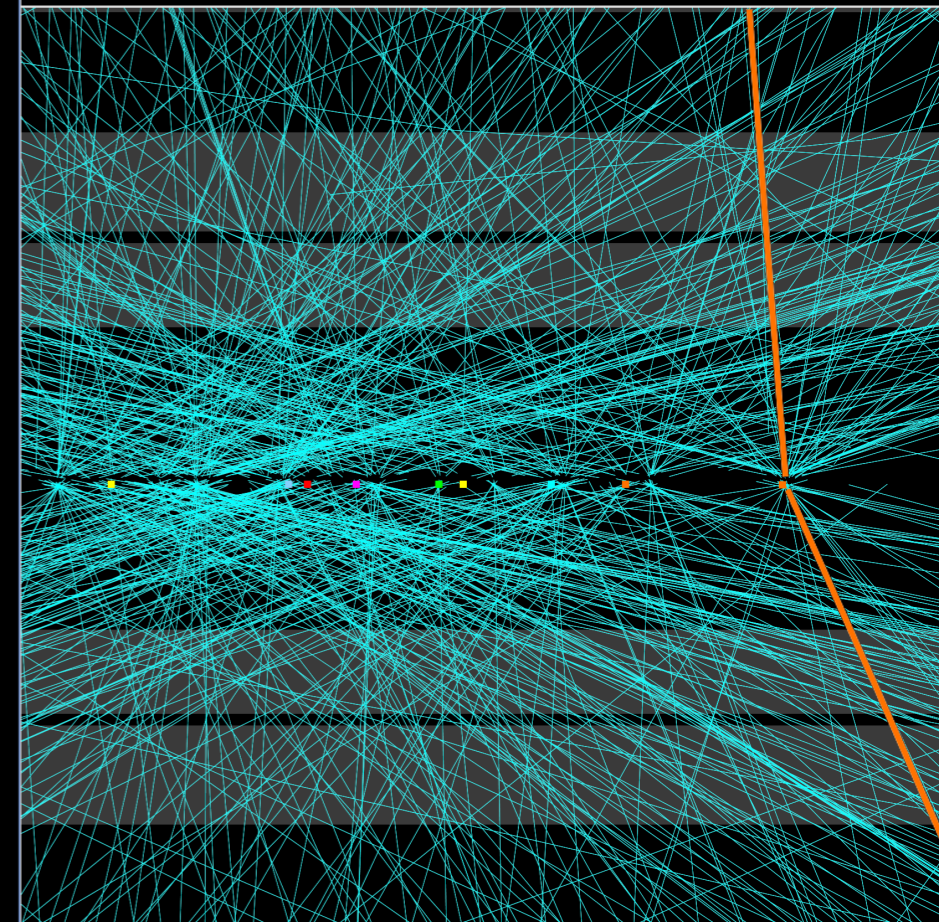
# What about other particles?



**ATLAS**  
EXPERIMENT

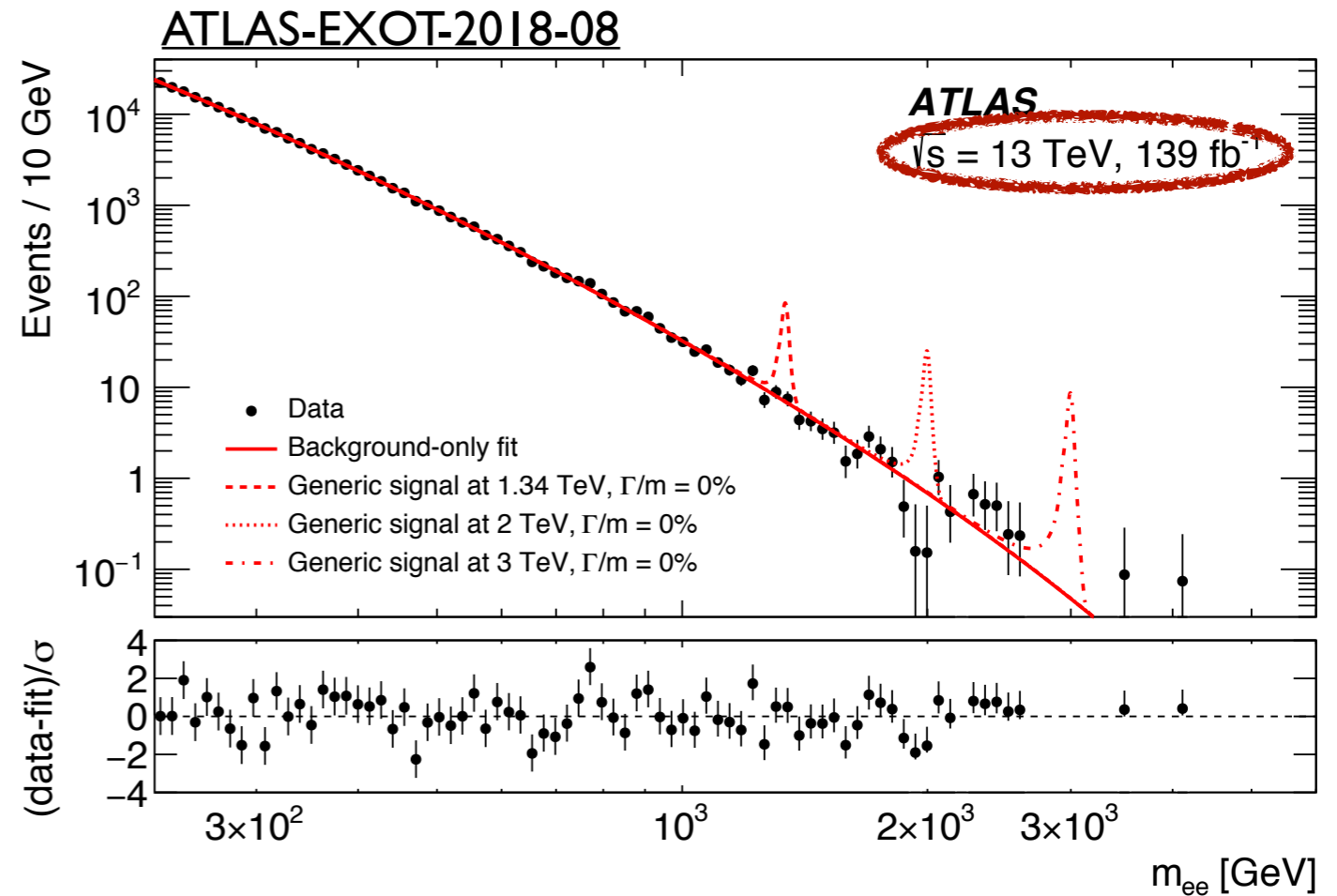
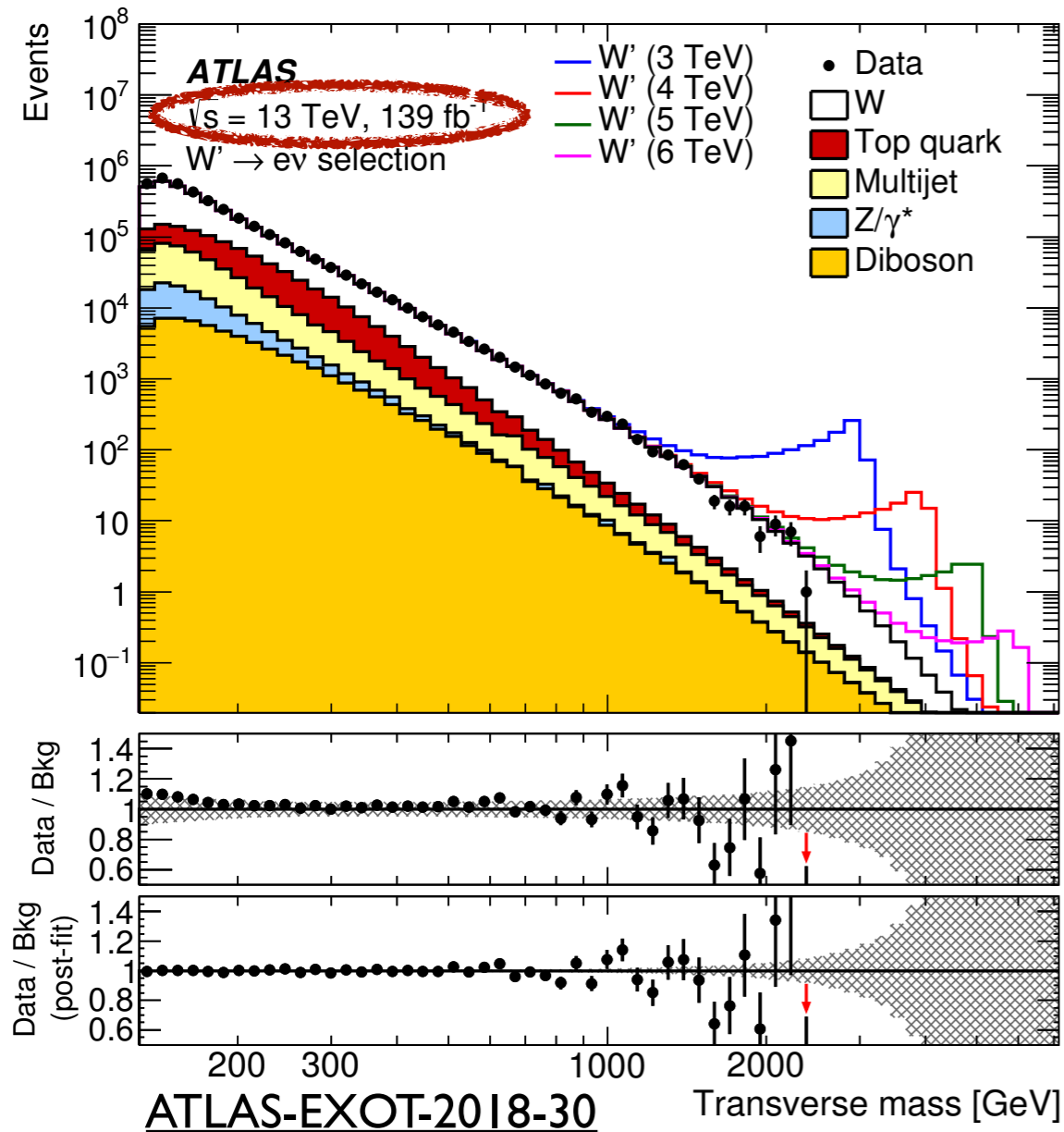
Run Number: 336852, Event Number: 1440436043

Date: 2017-09-29 11:44:35 CEST





# Searching with Leptons



You can also search for BSM with lepton resonances!

Not better luck here, even with the full datasets

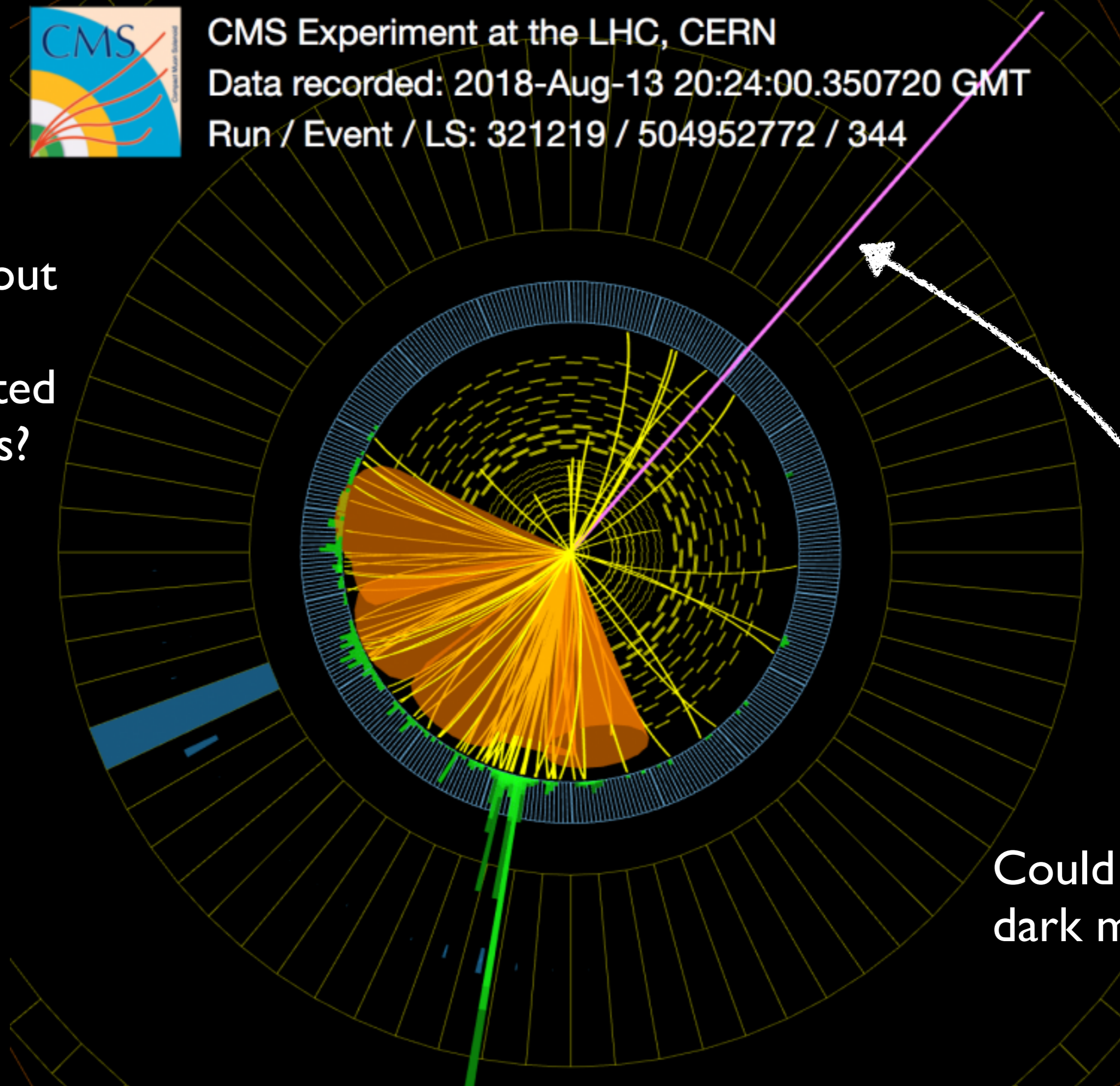


CMS Experiment at the LHC, CERN

Data recorded: 2018-Aug-13 20:24:00.350720 GMT

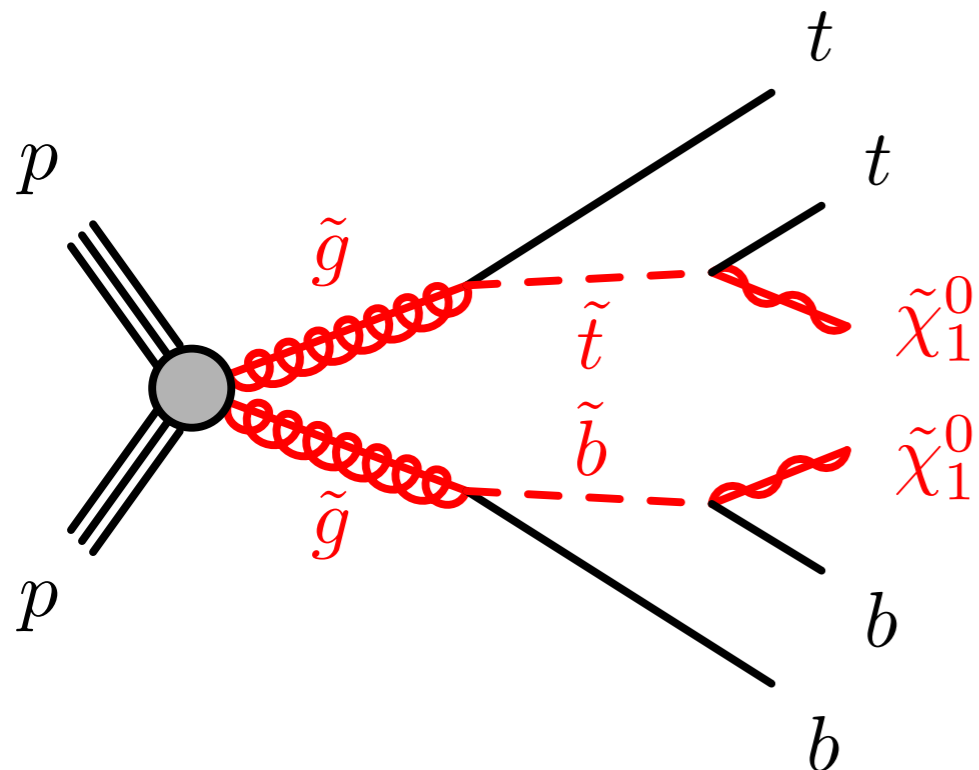
Run / Event / LS: 321219 / 504952772 / 344

What about  
more  
complicated  
signatures?

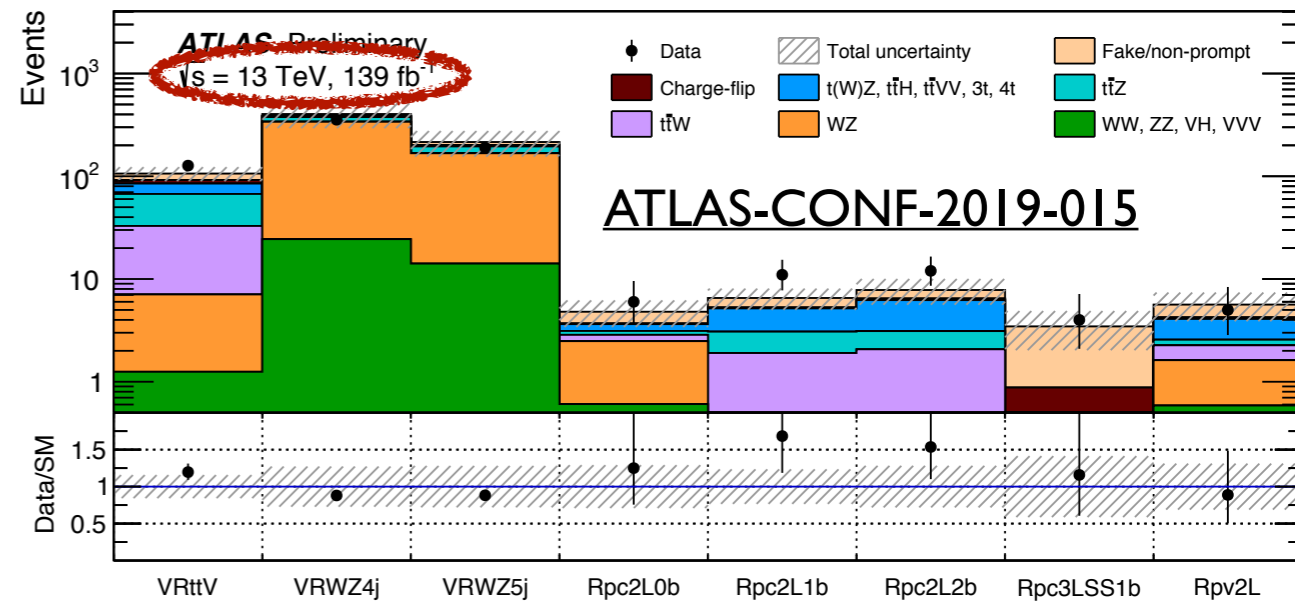
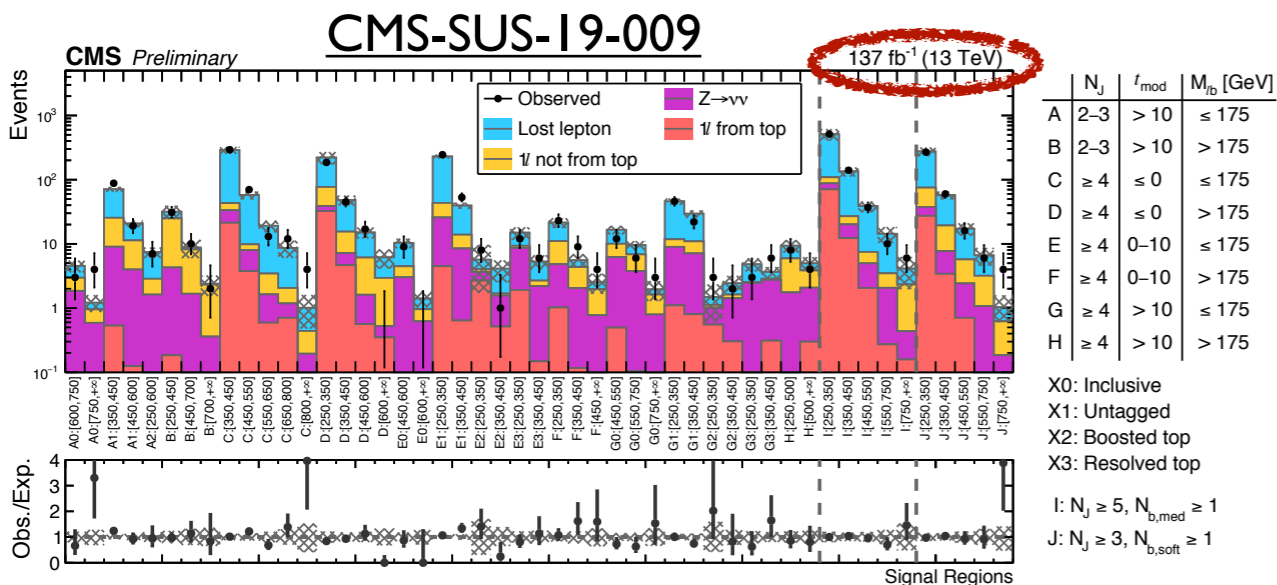
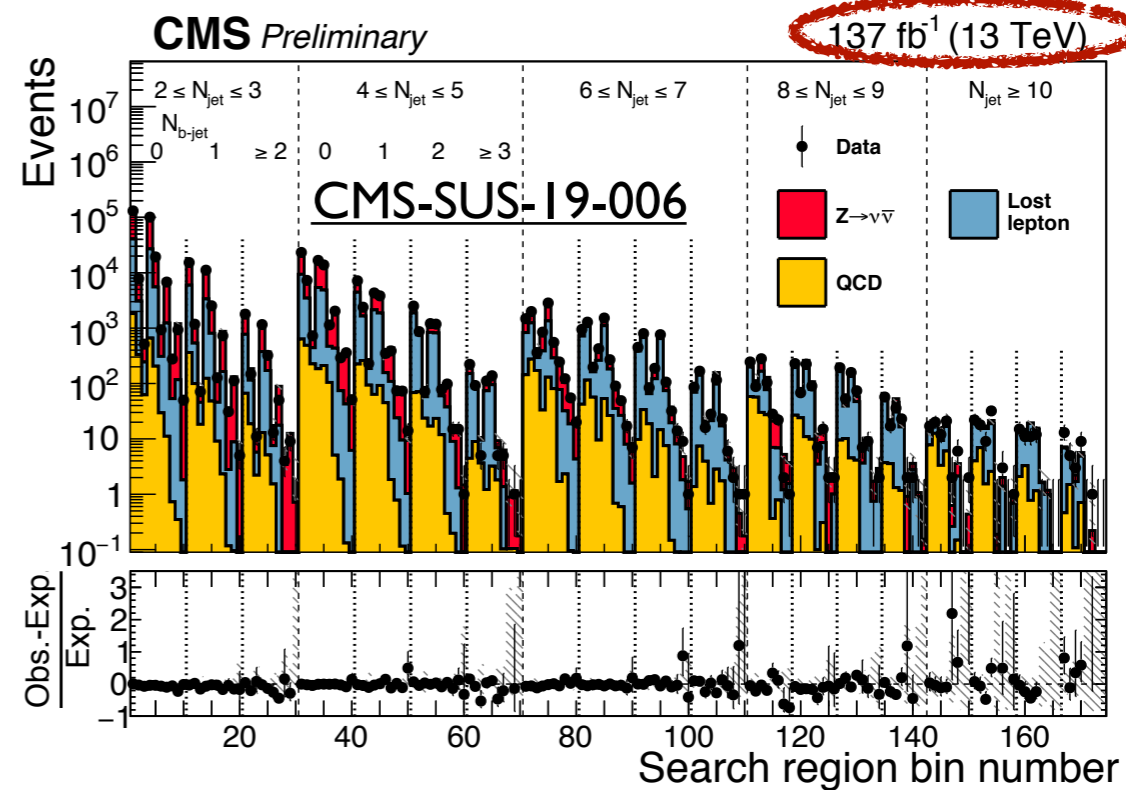


Could this be  
dark matter?

# SUSY Signatures



Supersymmetry predicts more complicated final states...



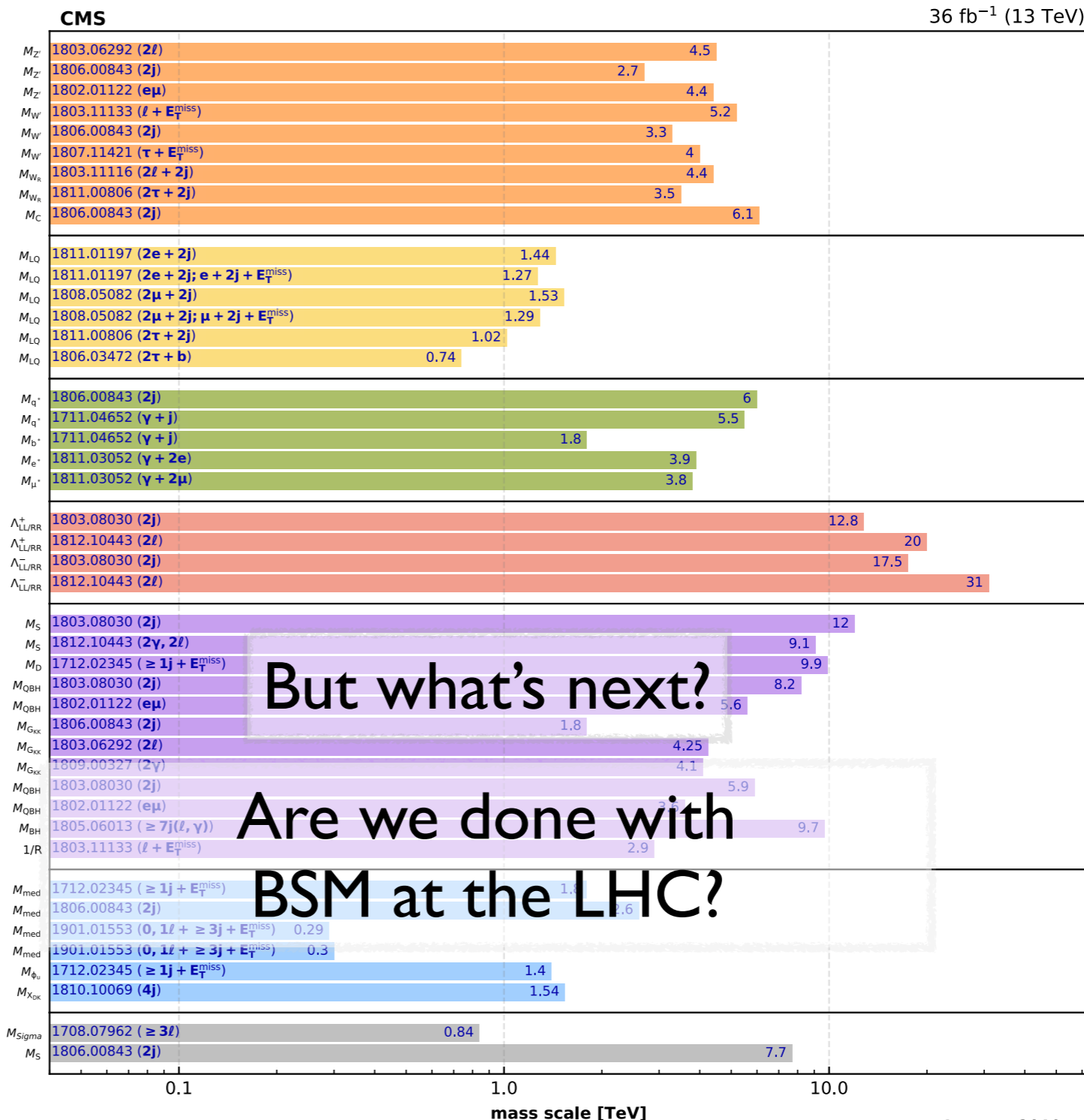
But still no hints, even with full data!

# Our Sensitivity is Better Than Ever



## Overview of CMS EXO results

- Heavy Gauge Bosons**
  - SSM  $Z'(\ell\ell)$
  - SSM  $Z'(q\bar{q})$
  - LFV  $Z'$ ,  $BR(e\mu) = 10\%$
  - SSM  $W'(\ell\nu)$
  - SSM  $W'(q\bar{q})$
  - SSM  $W'(\tau\nu)$
  - LRSM  $W_R(\ell N_R)$ ,  $M_{N_R} = 0.5M_{W_R}$
  - LRSM  $W_R(\tau N_R)$ ,  $M_{N_R} = 0.5M_{W_R}$
  - Axigluon, Coloron,  $\cot\theta = 1$
- Leptoquarks**
  - scalar LQ (pair prod.), coupling to 1<sup>st</sup> gen. fermions,  $\beta = 1$
  - scalar LQ (pair prod.), coupling to 1<sup>st</sup> gen. fermions,  $\beta = 0.5$
  - scalar LQ (pair prod.), coupling to 2<sup>nd</sup> gen. fermions,  $\beta = 1$
  - scalar LQ (pair prod.), coupling to 2<sup>nd</sup> gen. fermions,  $\beta = 0.5$
  - scalar LQ (pair prod.), coupling to 3<sup>rd</sup> gen. fermions,  $\beta = 1$
  - scalar LQ (single prod.), coup. to 3<sup>rd</sup> gen. ferm.,  $\beta = 1, \lambda = 1$
- Excited Fermions**
  - excited light quark ( $qg$ ),  $\Lambda = m_q^*$
  - excited light quark ( $q\gamma$ ),  $f_S = f = f' = 1, \Lambda = m_q^*$
  - excited b quark,  $f_S = f = f' = 1, \Lambda = m_q^*$
  - excited electron,  $f_S = f = f' = 1, \Lambda = m_e^*$
  - excited muon,  $f_S = f = f' = 1, \Lambda = m_\mu^*$
- Contact Interactions**
  - quark compositeness ( $q\bar{q}$ ),  $\eta_{LL/RR} = 1$
  - quark compositeness ( $\ell\ell$ ),  $\eta_{LL/RR} = 1$
  - quark compositeness ( $q\bar{q}$ ),  $\eta_{LL/RR} = -1$
  - quark compositeness ( $\ell\ell$ ),  $\eta_{LL/RR} = -1$
- Extra Dimensions**
  - ADD ( $jj$ ) HLZ,  $n_{ED} = 3$
  - ADD ( $\gamma\gamma, \ell\ell$ ) HLZ,  $n_{ED} = 3$
  - ADD  $G_{KK}$  emission,  $n = 2$
  - ADD QBH ( $jj$ ),  $n_{ED} = 6$
  - ADD QBH ( $e\mu$ ),  $n_{ED} = 6$
  - RS  $G_{KK}(q\bar{q}, gg)$ ,  $k/\overline{M}_{Pl} = 0.1$
  - RS  $G_{KK}(\ell\ell)$ ,  $k/\overline{M}_{Pl} = 0.1$
  - RS  $G_{KK}(\gamma\gamma)$ ,  $k/\overline{M}_{Pl} = 0.1$
  - RS QBH ( $jj$ ),  $n_{ED} = 1$
  - RS QBH ( $e\mu$ ),  $n_{ED} = 1$
  - non-rotating BH,  $M_D = 4$  TeV,  $n_{ED} = 6$
  - split-UED,  $\mu \geq 4$  TeV
- Dark Matter**
  - (axial-)vector mediator ( $\chi\chi$ ),  $g_q = 0.25, g_{DM} = 1, m_\chi = 1$  GeV
  - (axial-)vector mediator ( $q\bar{q}$ ),  $g_q = 0.25, g_{DM} = 1, m_\chi = 1$  GeV
  - scalar mediator ( $+t/t\bar{t}$ ),  $g_q = 1, g_{DM} = 1, m_\chi = 1$  GeV
  - pseudoscalar mediator ( $+t/t\bar{t}$ ),  $g_q = 1, g_{DM} = 1, m_\chi = 1$  GeV
  - scalar mediator (fermion portal),  $\lambda_u = 1, m_\chi = 1$  GeV
  - complex sc. med. (dark QCD),  $m_{\text{tok}} = 5$  GeV,  $c\tau_{\text{tok}} = 25$  mm
- Other**
  - Type III Seesaw,  $B_e = B_\mu = B_\tau$
  - string resonance



Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included).

January 2019



Low Hanging Fruit Records

# BSM is More Vibrant Than Ever!

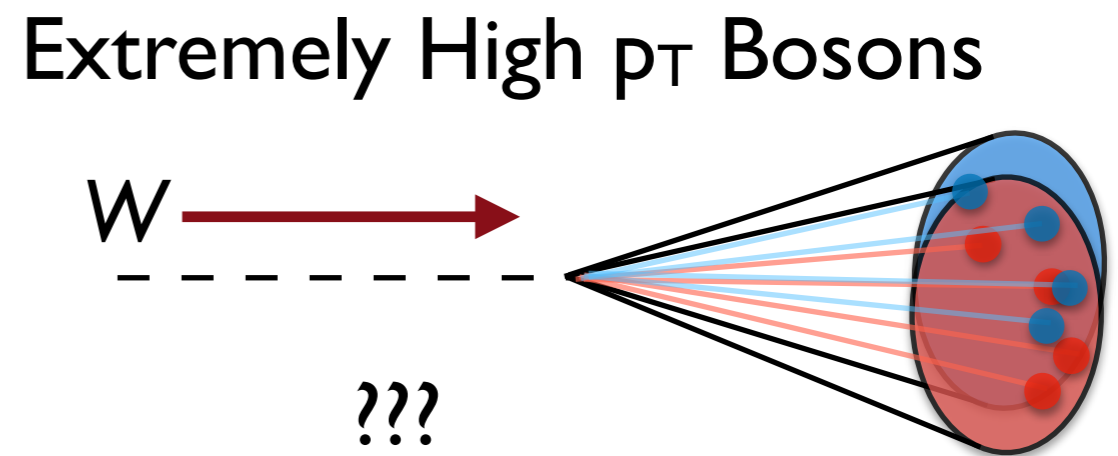
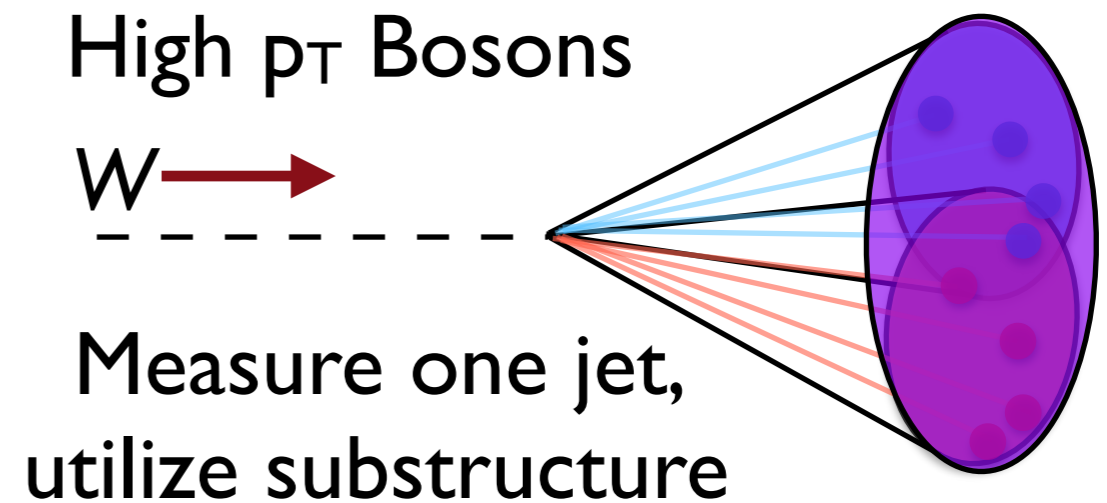
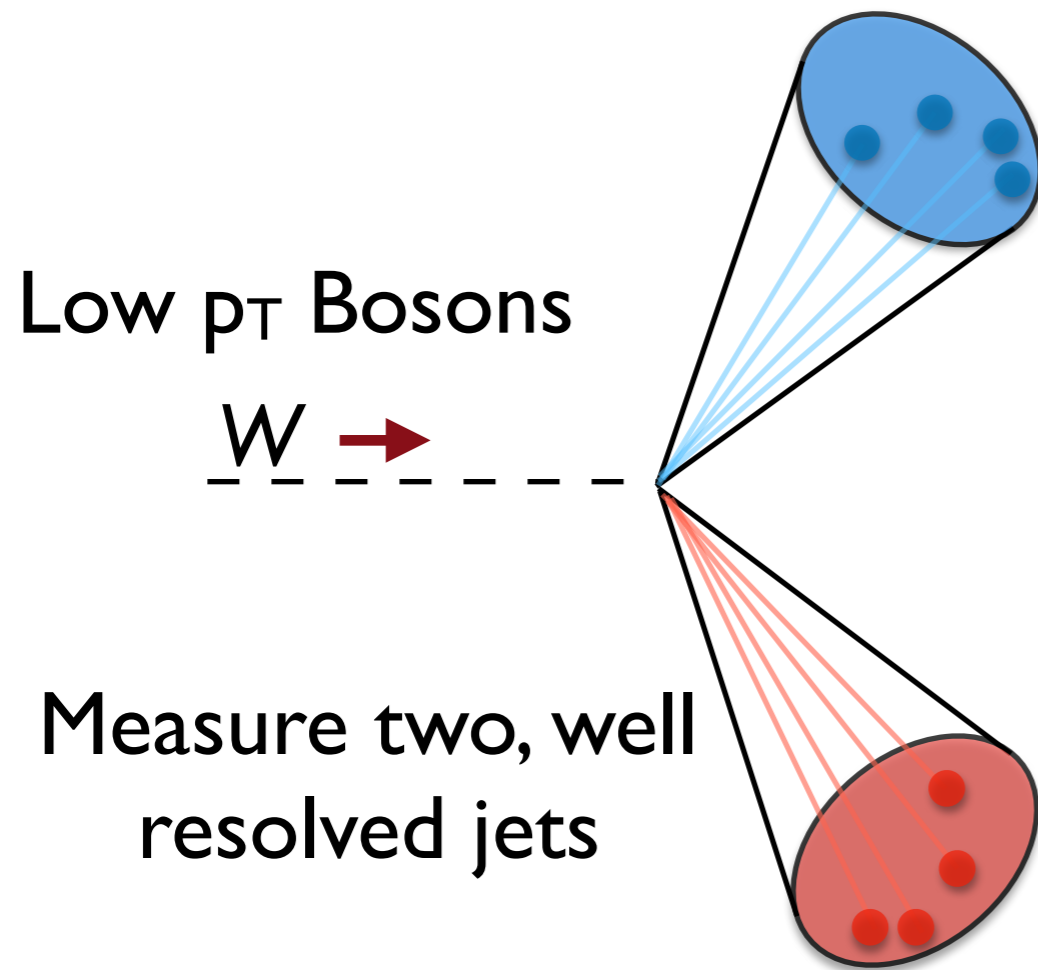


- The low-hanging fruit is mostly gone, but that just means we have to **work harder!**
- Let's take a look at the new methods which are enabling discovery today:
  - **Squeeze every last drop:** reconstruction and machine learning makes the most of our data
  - **Rarer than rare:** use huge datasets to access tiny signals
  - **Not your advisor's signals:** searching for things ATLAS and CMS weren't designed for
- My apologies for omitting many interesting and exciting results!

# Squeeze Every Drop

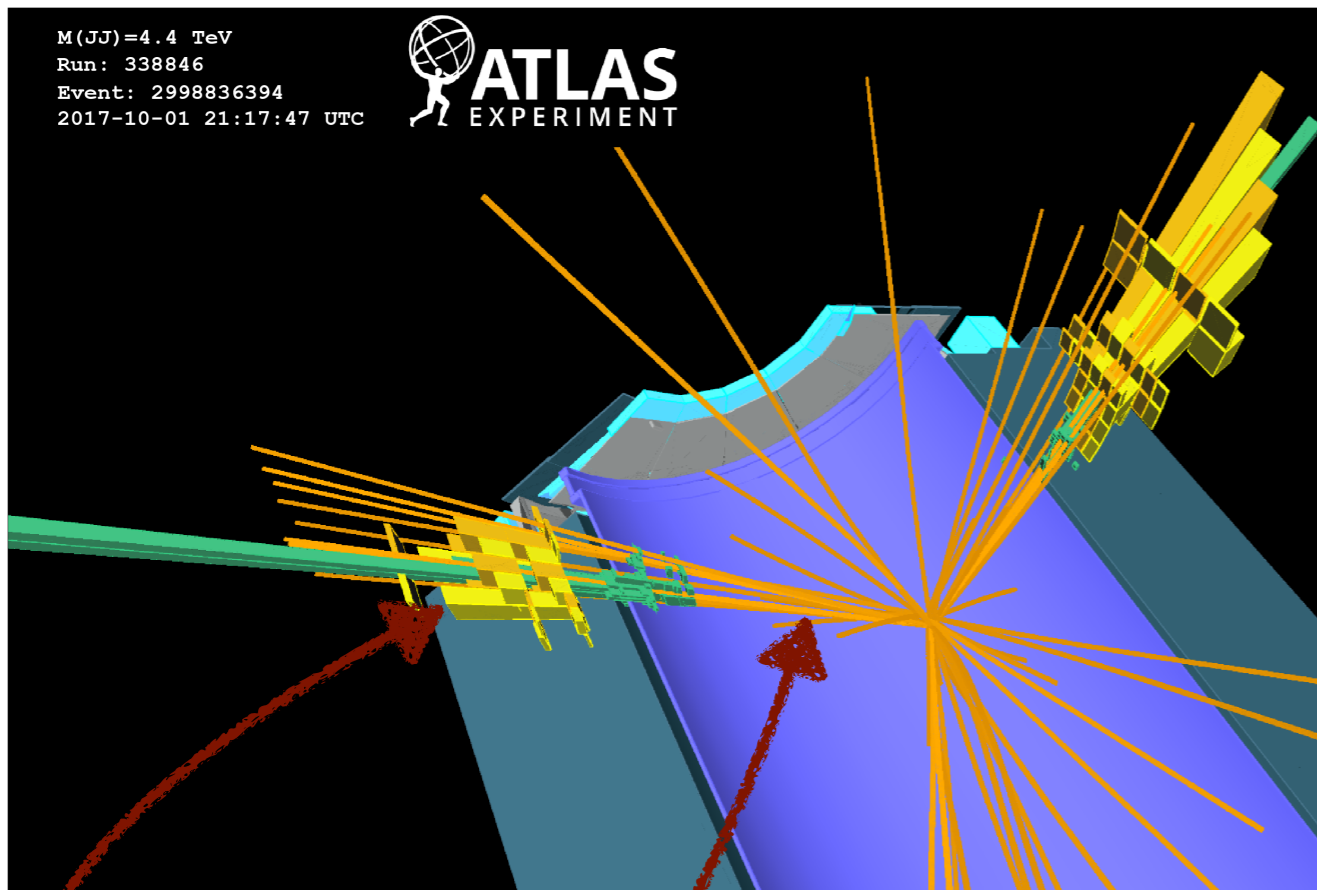
*Or: how advances in reconstruction,  
triggering, and machine learning are unlocking  
new insights into BSM*

# The Challenge of High $p_T$



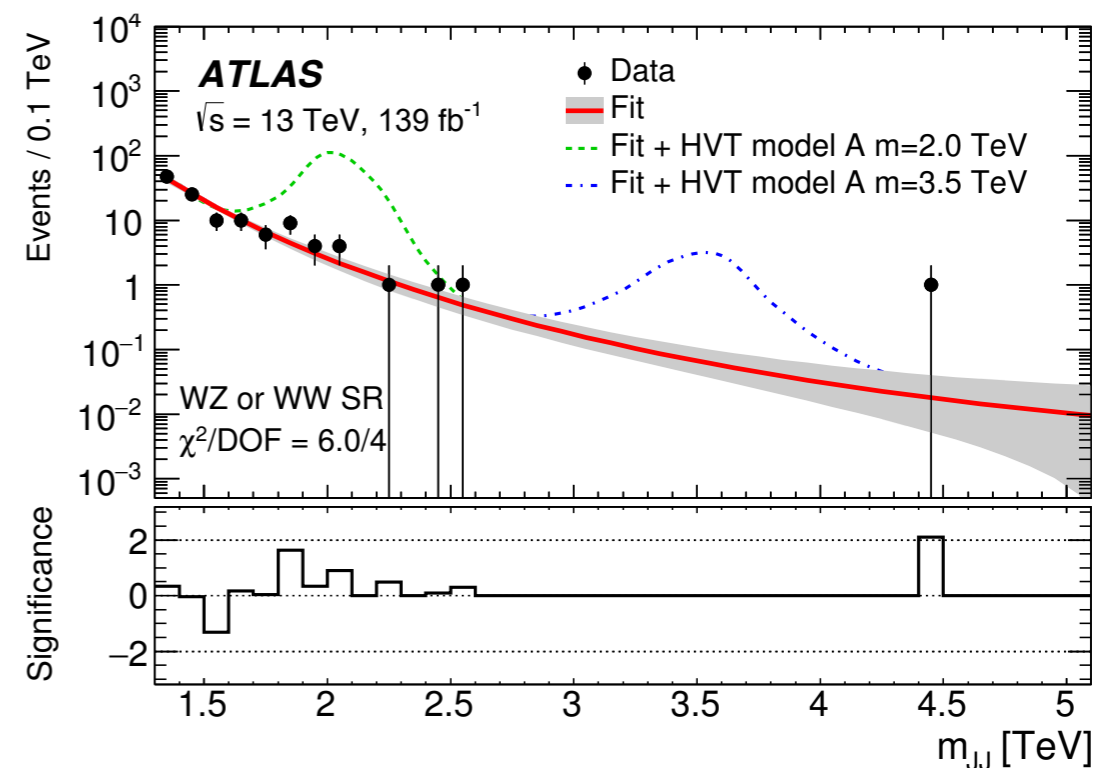
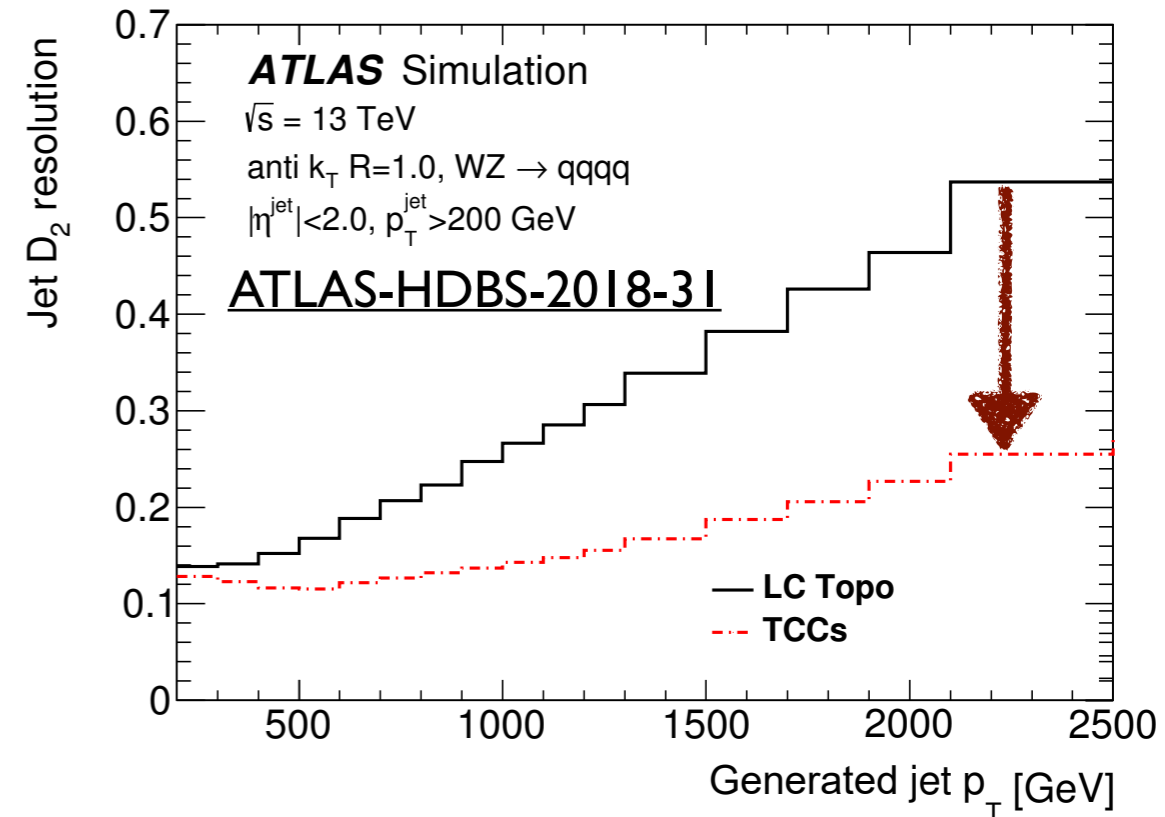


# Tagging at High $p_T$

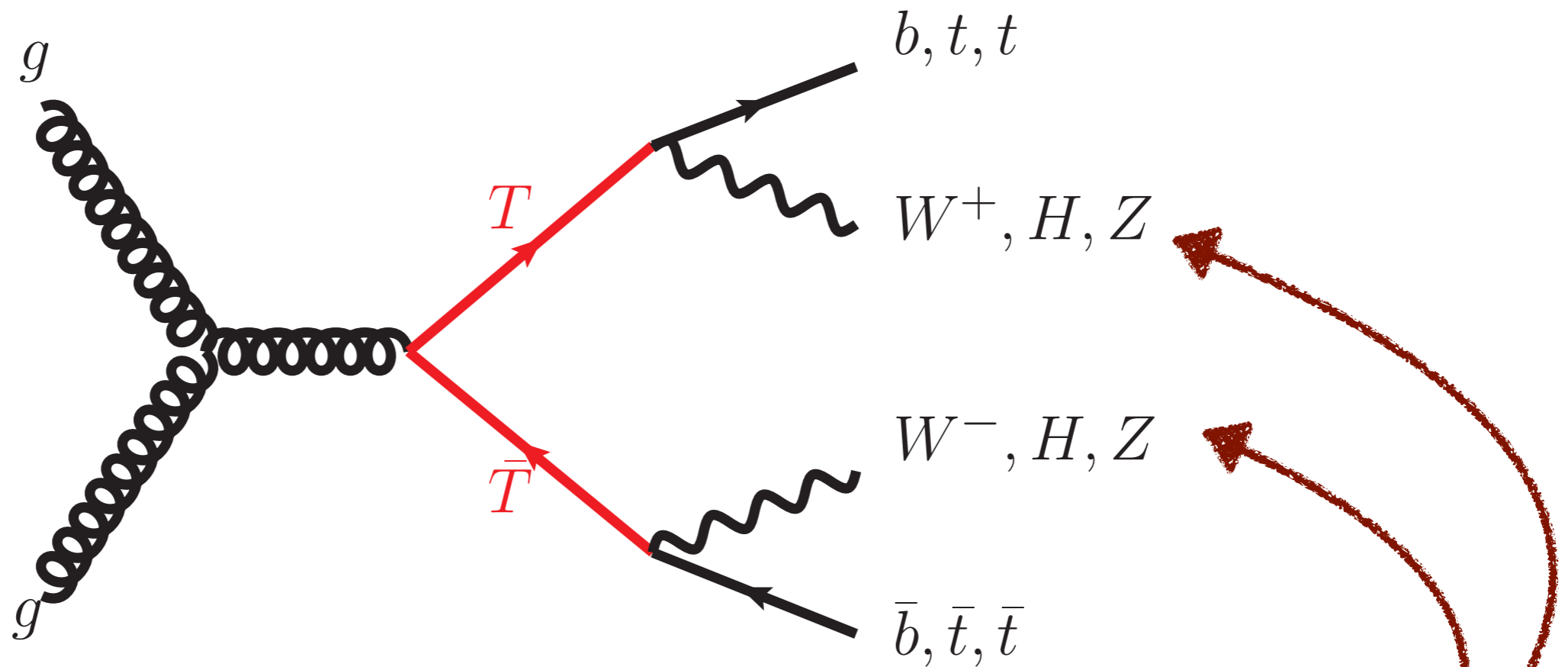


Utilize better spatial resolution from tracker to separate energy deposits in the calorimeter!

Enables strongest sensitivity yet to boosted all-hadronic final states



# The Challenge of Branching

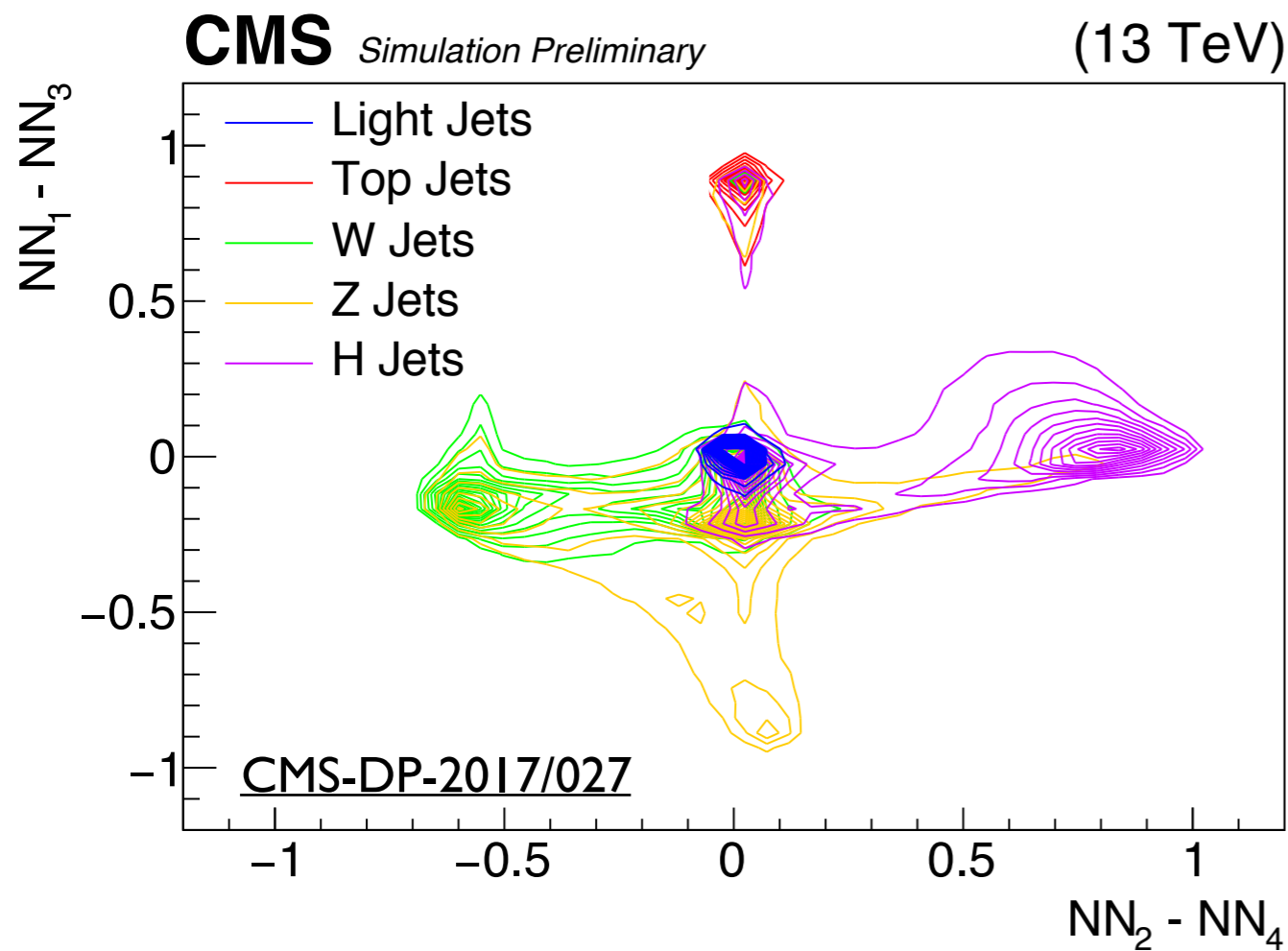


Vector-like tops are a consequence of many BSM models:  
can explain Higgs mass, etc.

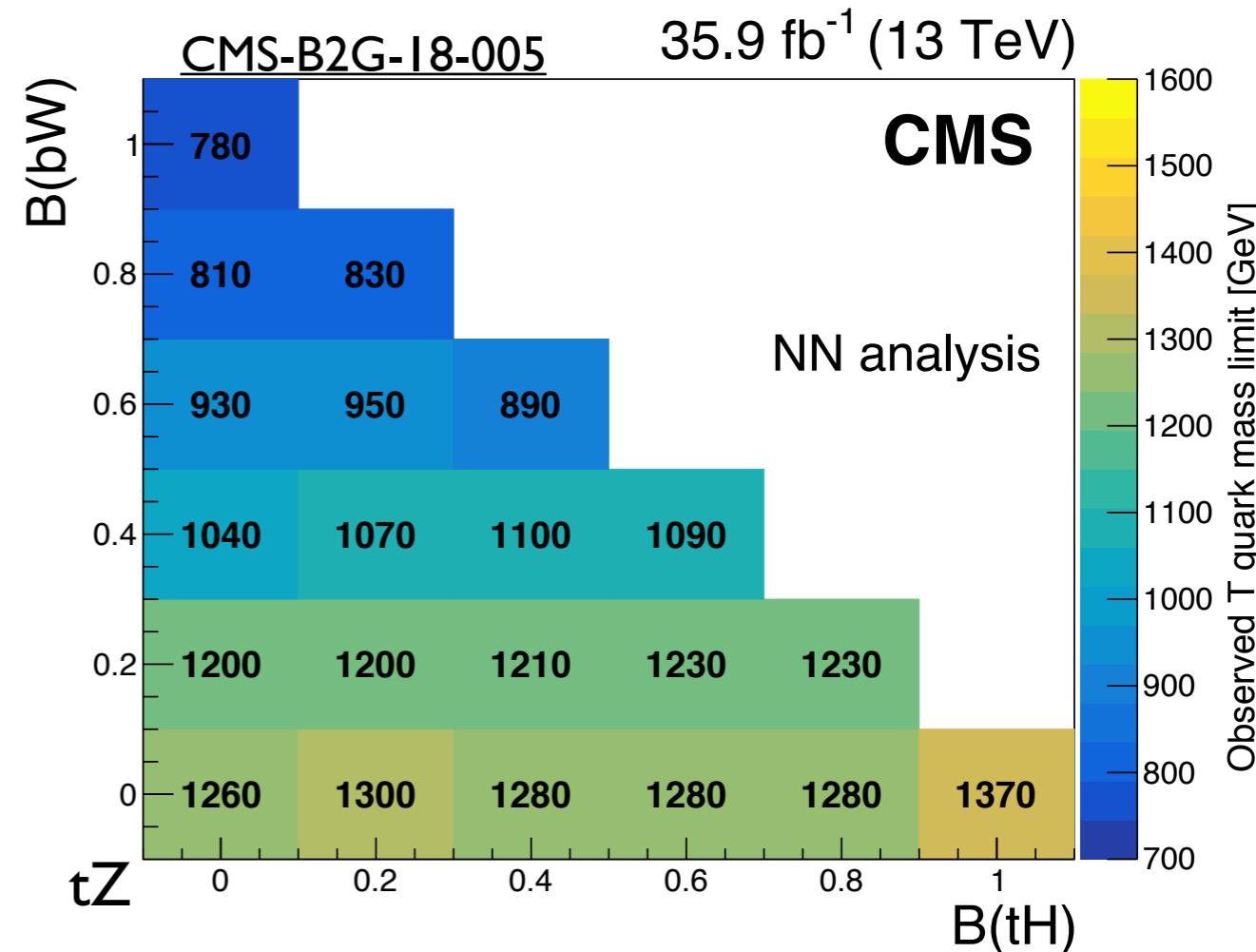
But there are a huge number of final states!

How do you search for all of these efficiently?

# Multi-Node Tagging

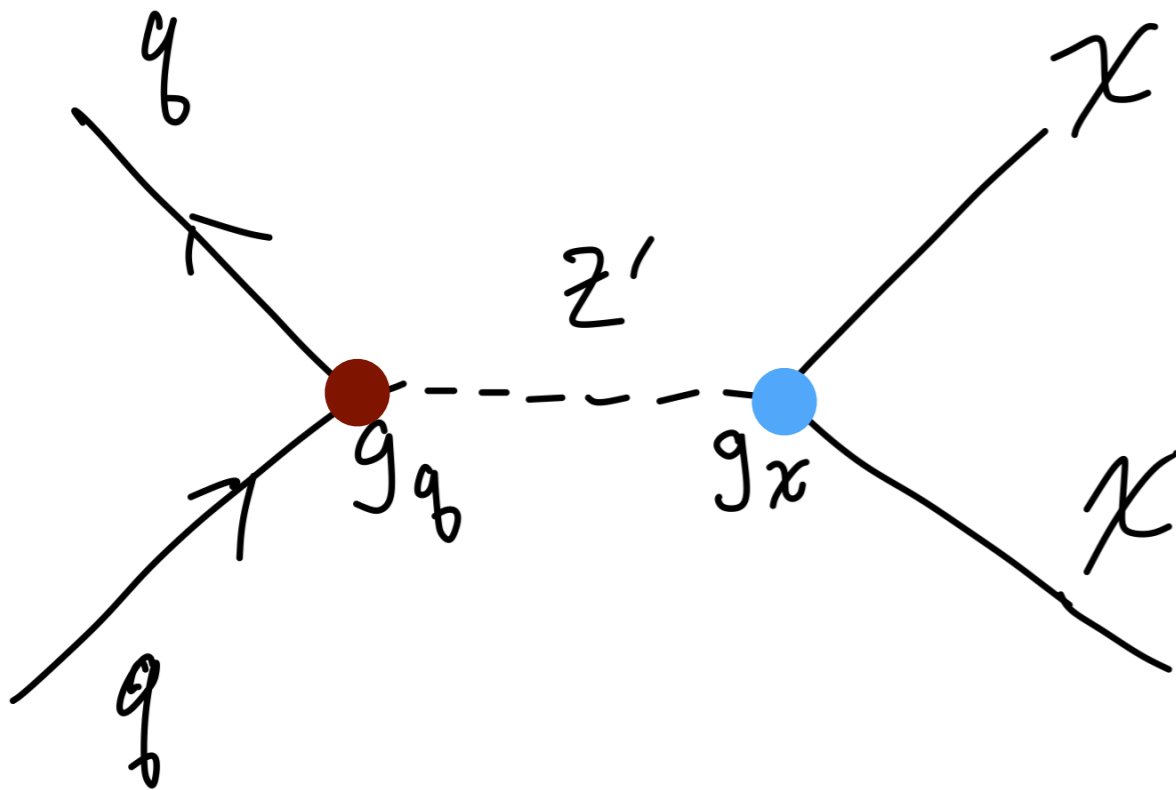


Train a neural network to distinguish all signal classes from background

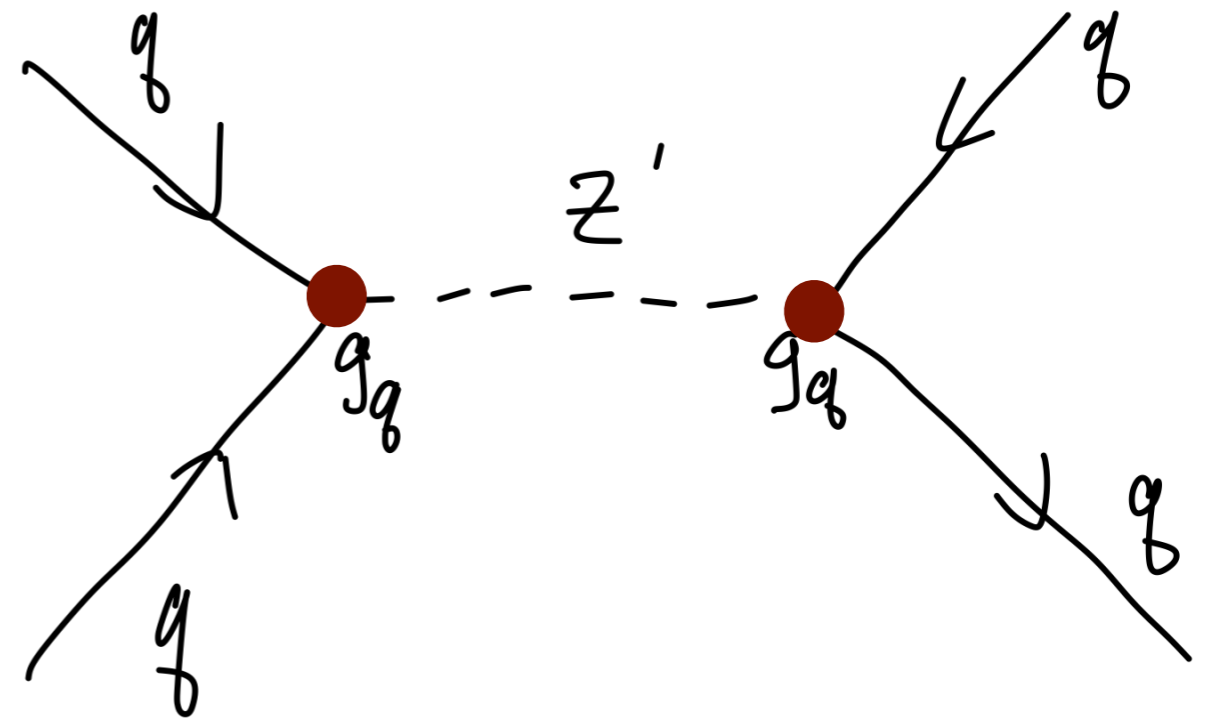


Use the NN to search for new physics in many final states at once

# The Challenges of Dijets



If we can produce DM at the LHC like this...



Then the mediator should decay back to jets!

But if the  $Z'$  is low enough mass, we won't trigger the event from the jets!

# Data Scouting



Trigger limitations  
come from  
total bandwidth

$$\text{Event Size} \times \text{Event Rate} = \text{Total Bandwidth}$$

Normally, we record  
“full events,” so the  
rate gets limited

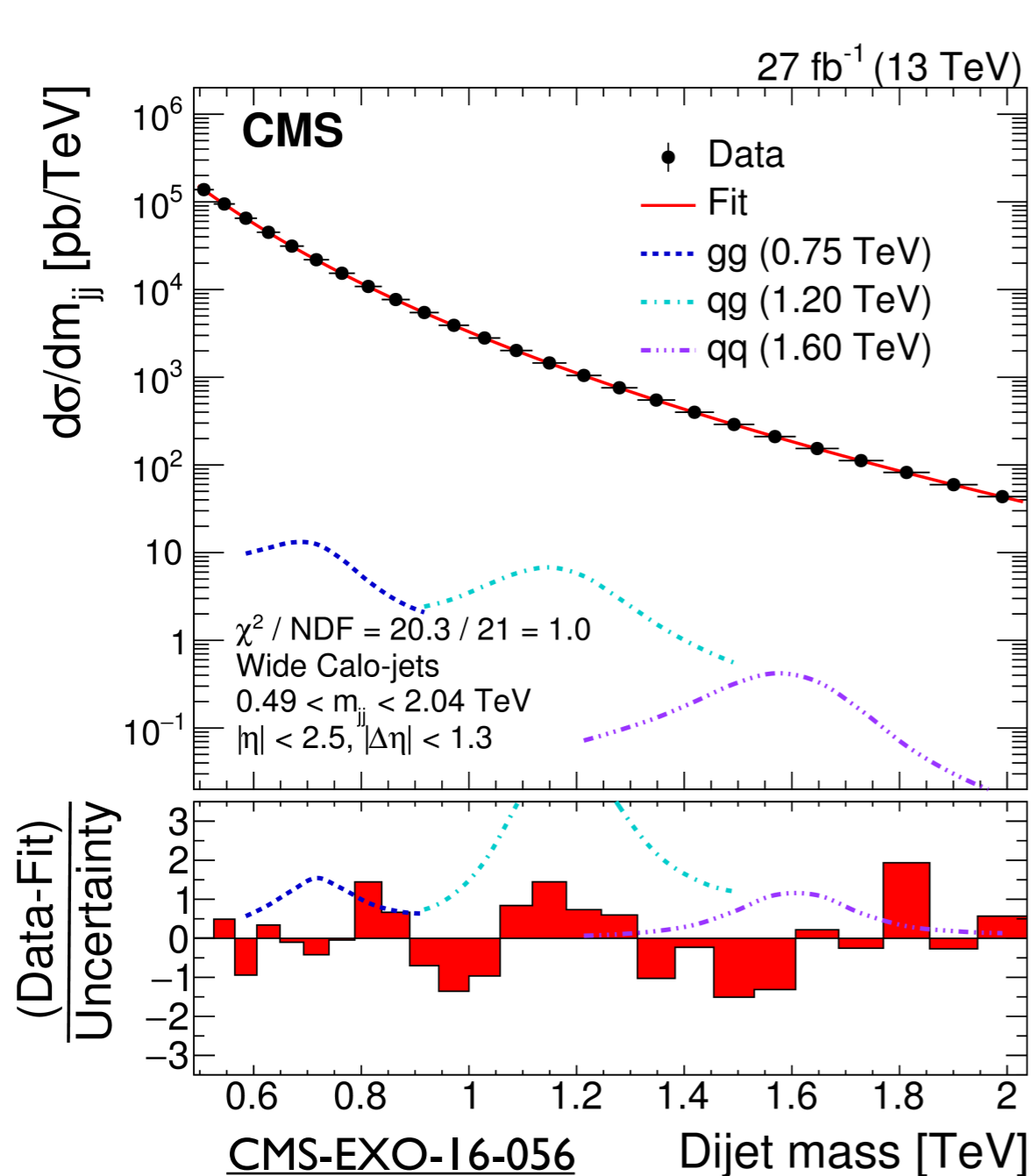
$$\text{Event Size} \times \text{Rate} = \text{Total Bandwidth}$$

But we can reco  
events in the trigger,  
and save smaller data

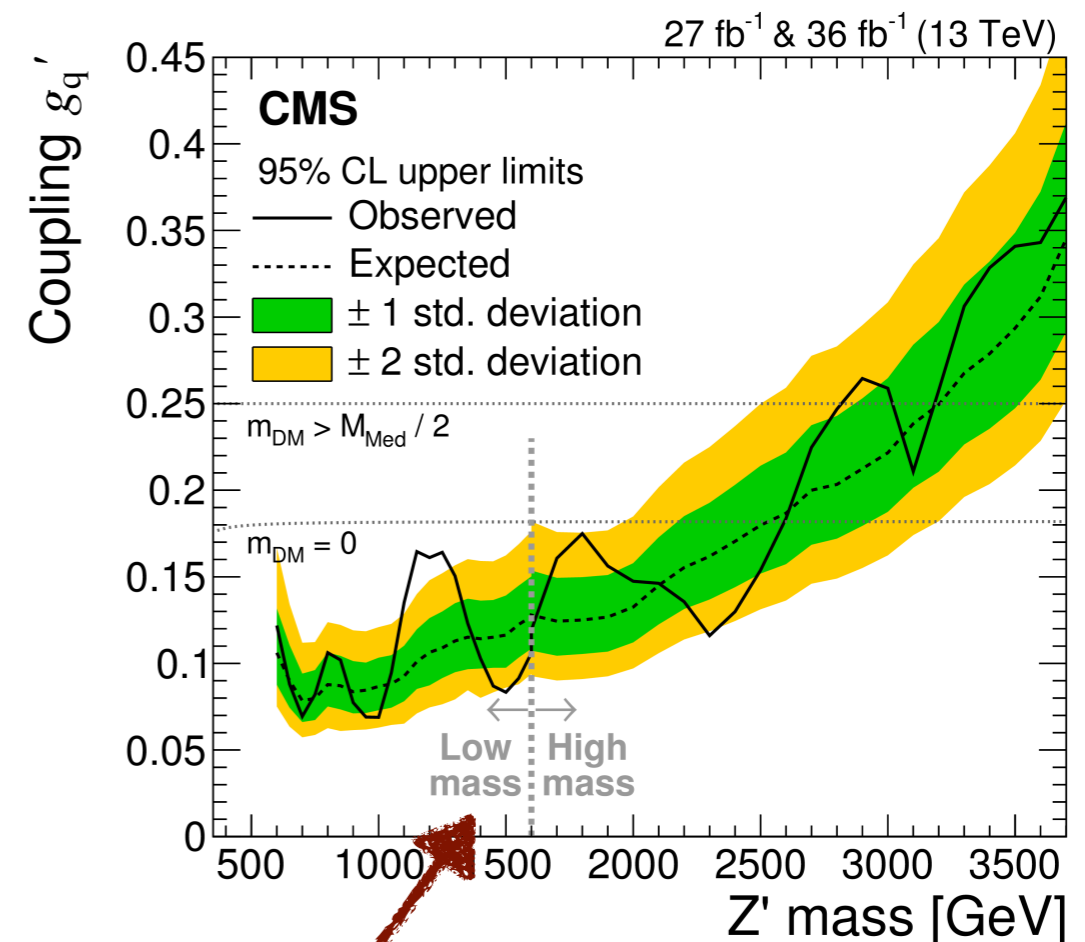
$$\text{Size} \times \text{Event Rate} = \text{Total Bandwidth}$$

If you save smaller events, you can save more of them!

# What You Can Accomplish



Can collect huge datasets  
 at low mass using this technique!

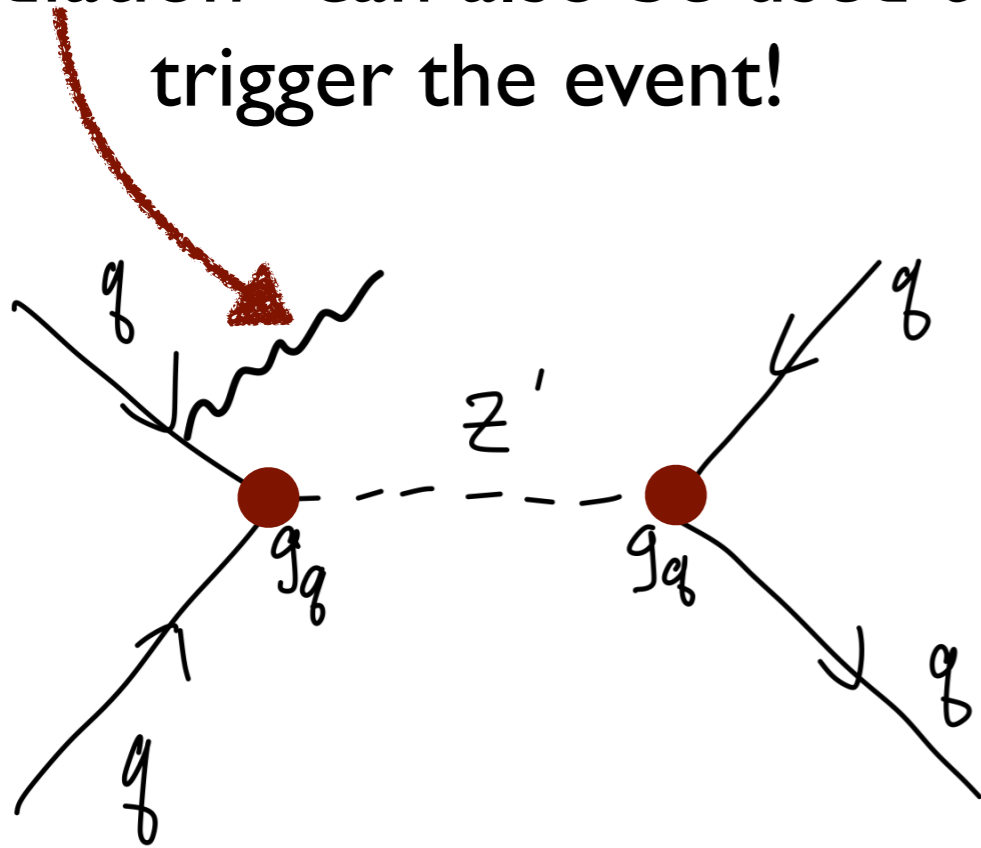


Can probe unique physics  
 phase space by recording  
 only a portion of the events!

# Extra Radiation

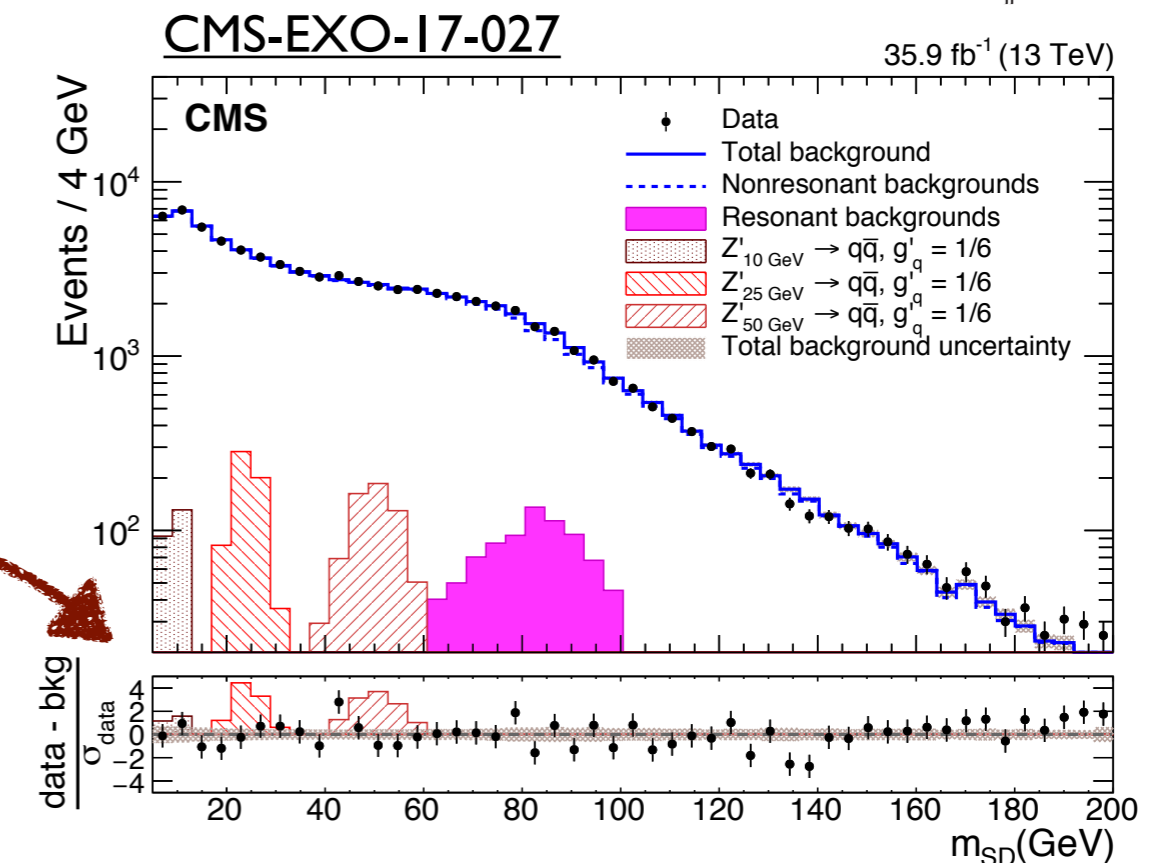
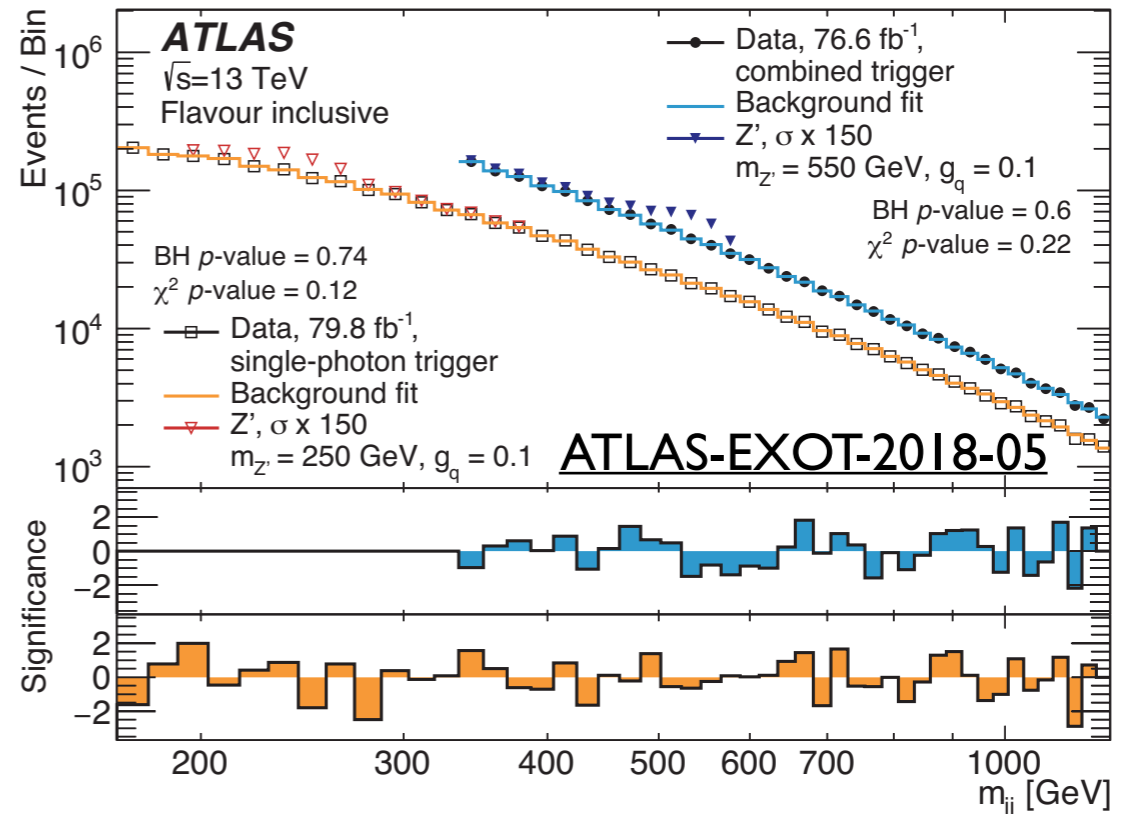


Photon or jet “initial state radiation” can also be used to trigger the event!

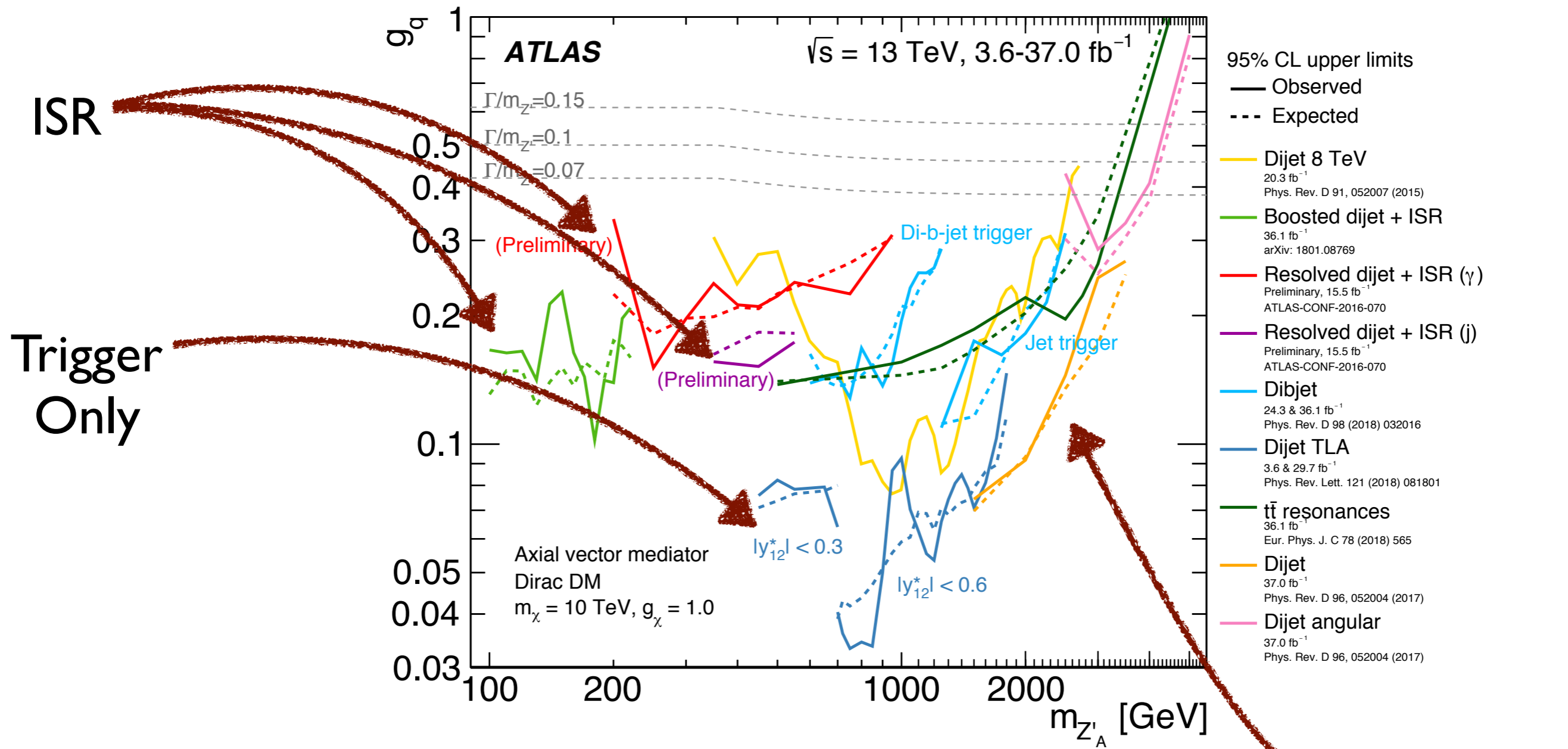


Even lower regions of mediator mass can be explored with these techniques

Can even search as low as 10 GeV!



# The Big Picture for Z'



In the hunt for Z', new techniques are taking center stage!



# Rarer than Rare

*Or: how huge datasets and clever strategies  
are enabling searches for vanishingly  
small signals*

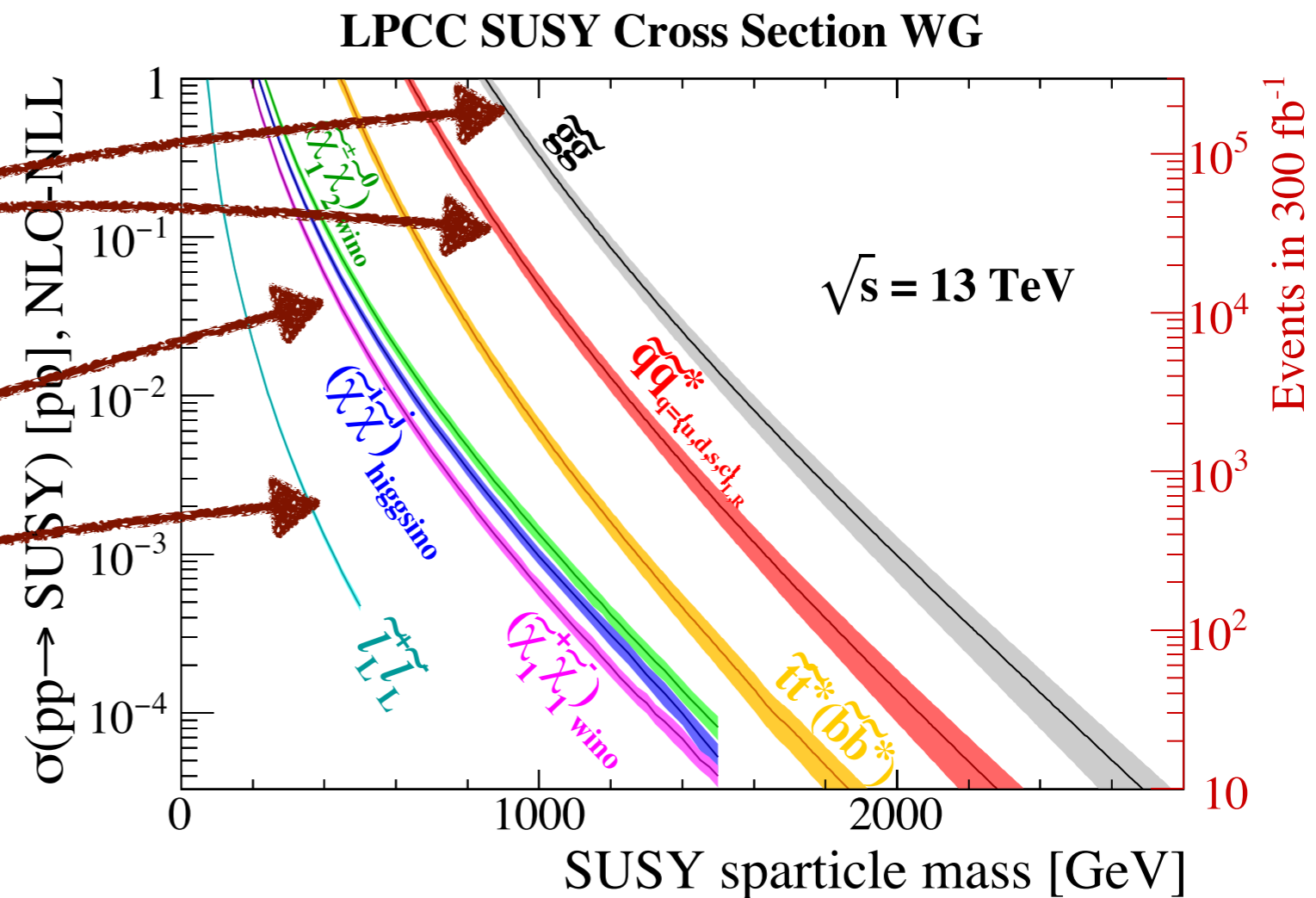
# Why Search for Rare Signals?



Because many theories predict low cross-section signals!

If you've already excluded  
gluinos and squarks  
and stops...

Higgsinos and sleptons  
might still be in sight!



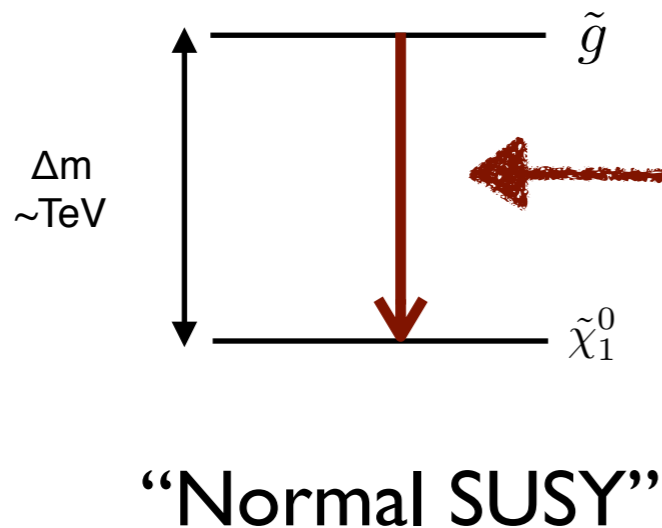
<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/SUSYCrossSections>

arXiv:1407.5066

# Higgsinos: The Challenge



mass  
↑  
~0

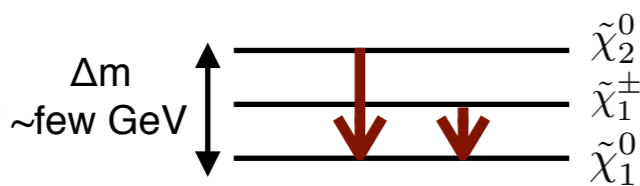


"Normal SUSY"

"Classic" SUSY searches exploit large mass splittings

Cascades from initial sparticles will be high energy, lots of missing energy

mass  
↑  
 $\mu$   
~0

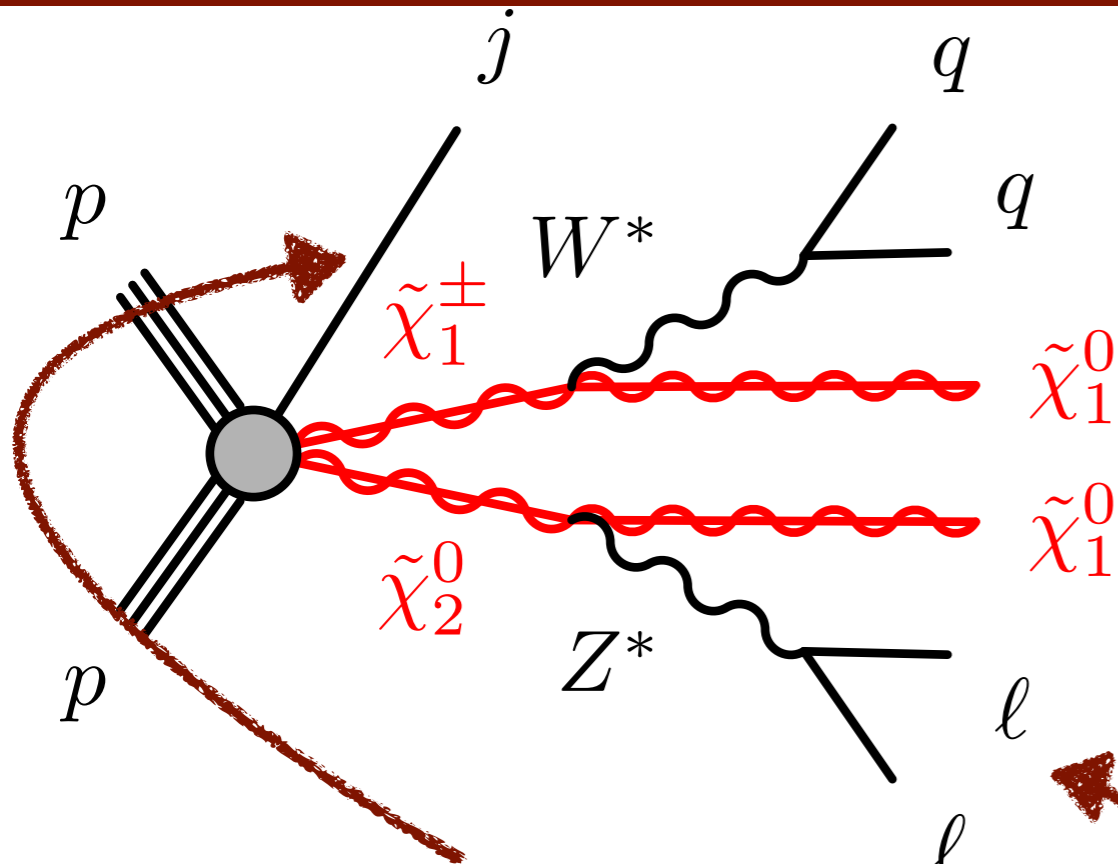


Higgsinos

Higgsinos are another story...

No large mass splitting: very soft visible particles, no obvious missing energy

# Searching for Rare Higgsinos

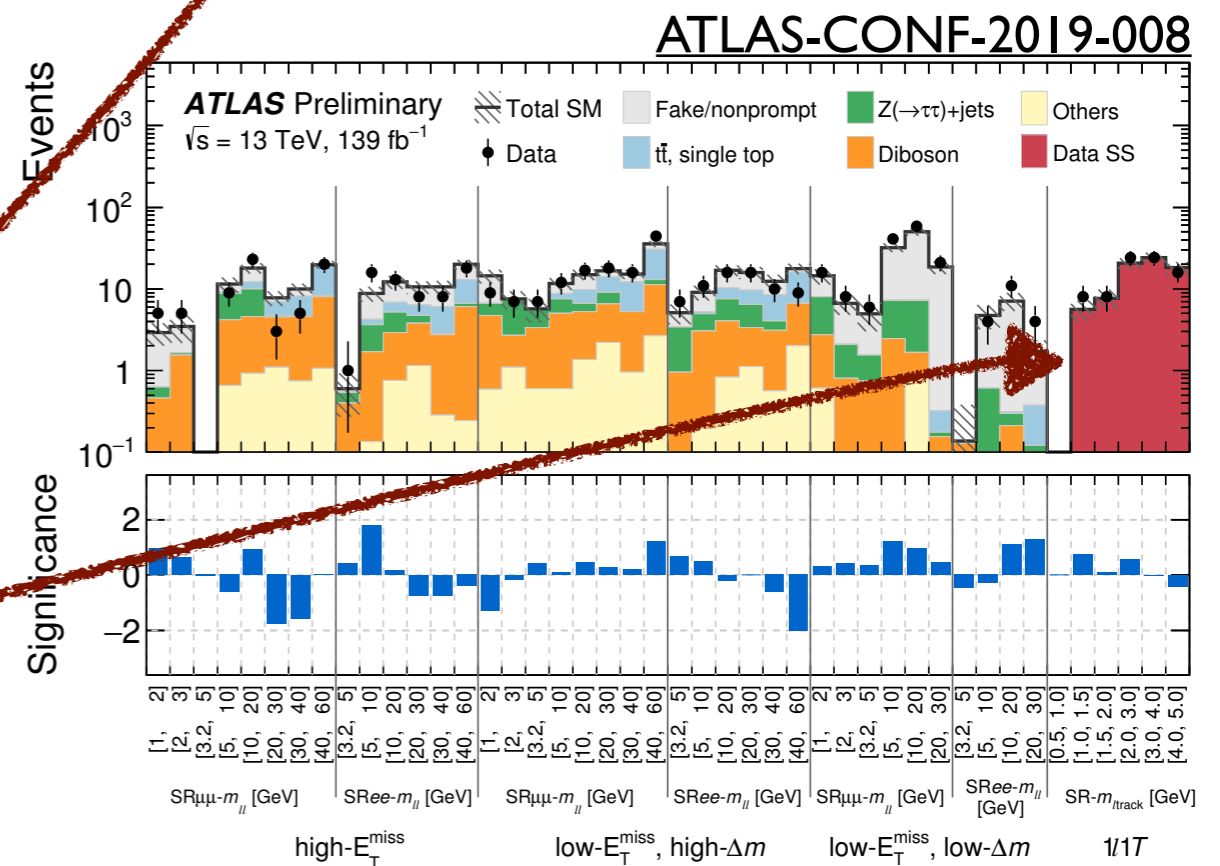
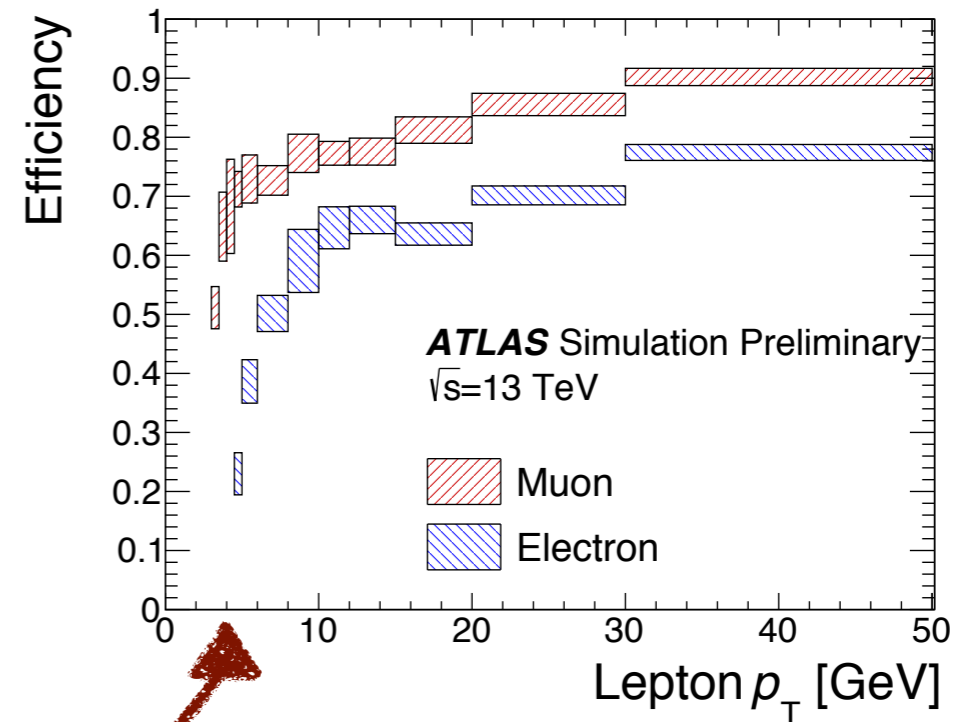


Use an ISR jet to boost the system and trigger

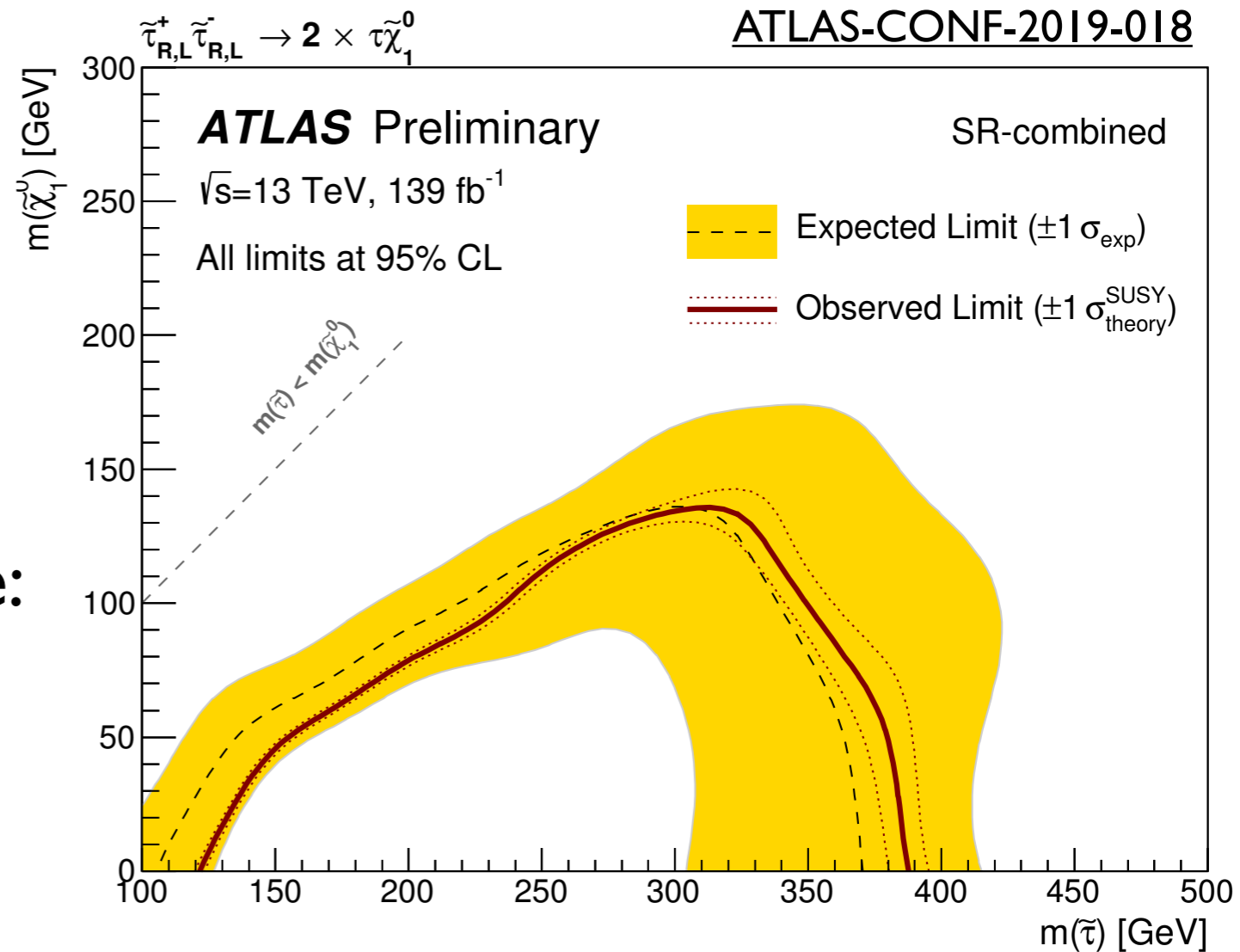
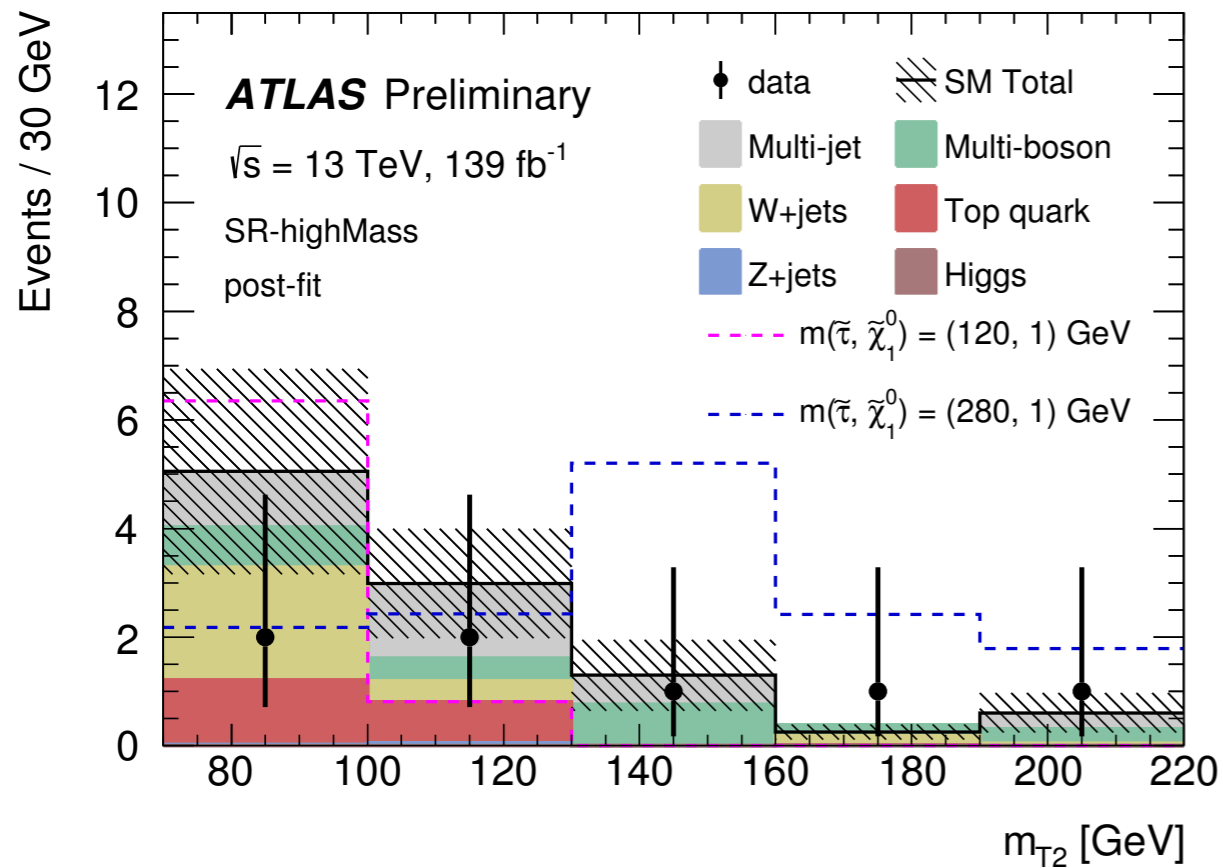
Then, look for low-mass  $Z^*$  decays

Critical to reconstruct leptons at extremely low  $p_T$

New lepton + track region extends sensitivity to even lower masses!



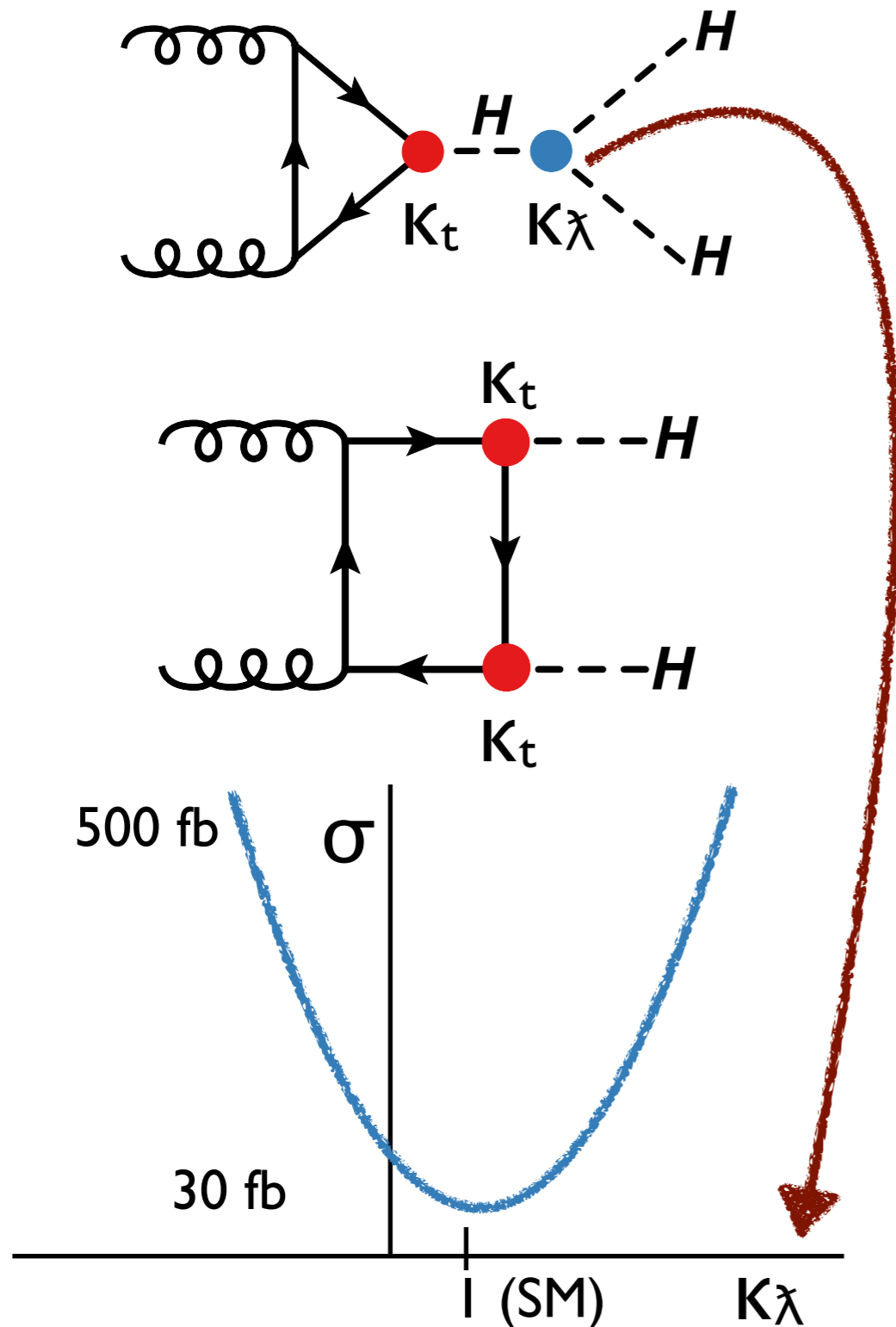
# Even Rarer: Staus



Staus are an even larger challenge:  
 incredibly low cross-section!

Effective triggers and the large  
 dataset enable first substantial  
 sensitivity at the LHC!

# Di-Higgs at the LHC



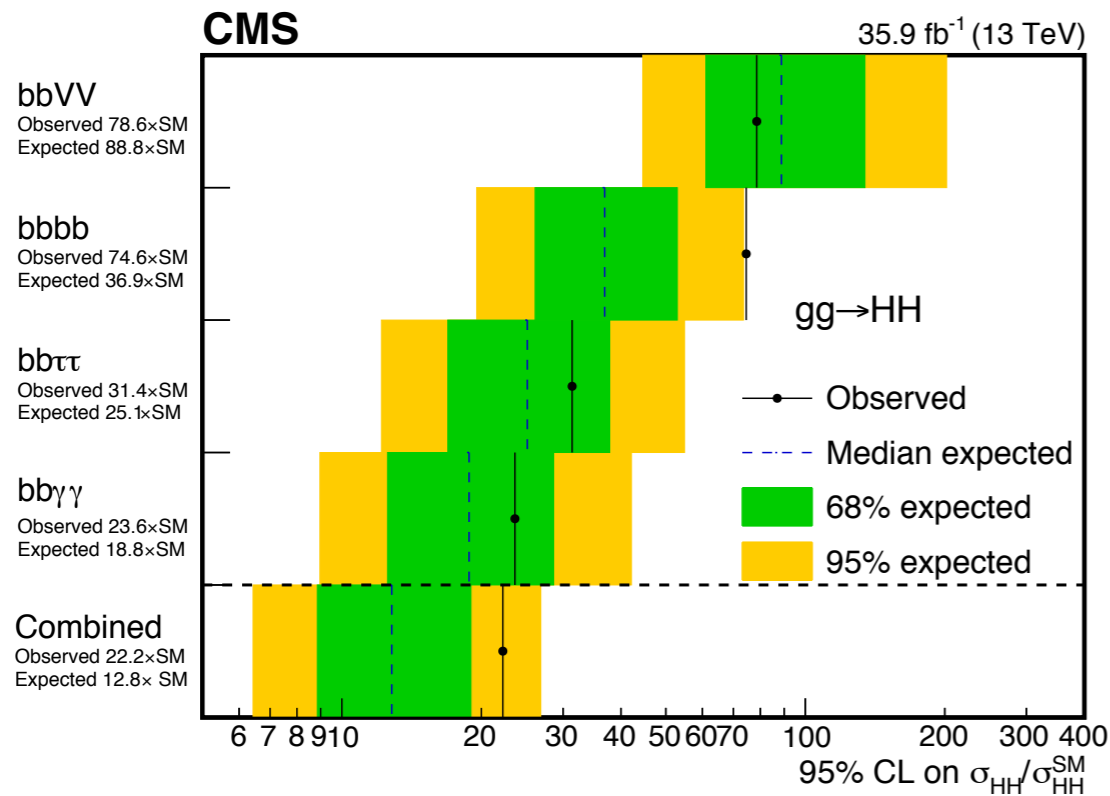
Di-Higgs is another exciting new target for the LHC

Exciting signal: can reveal the shape of the Higgs potential!

Interference between SM diagrams leads to very low x-sec

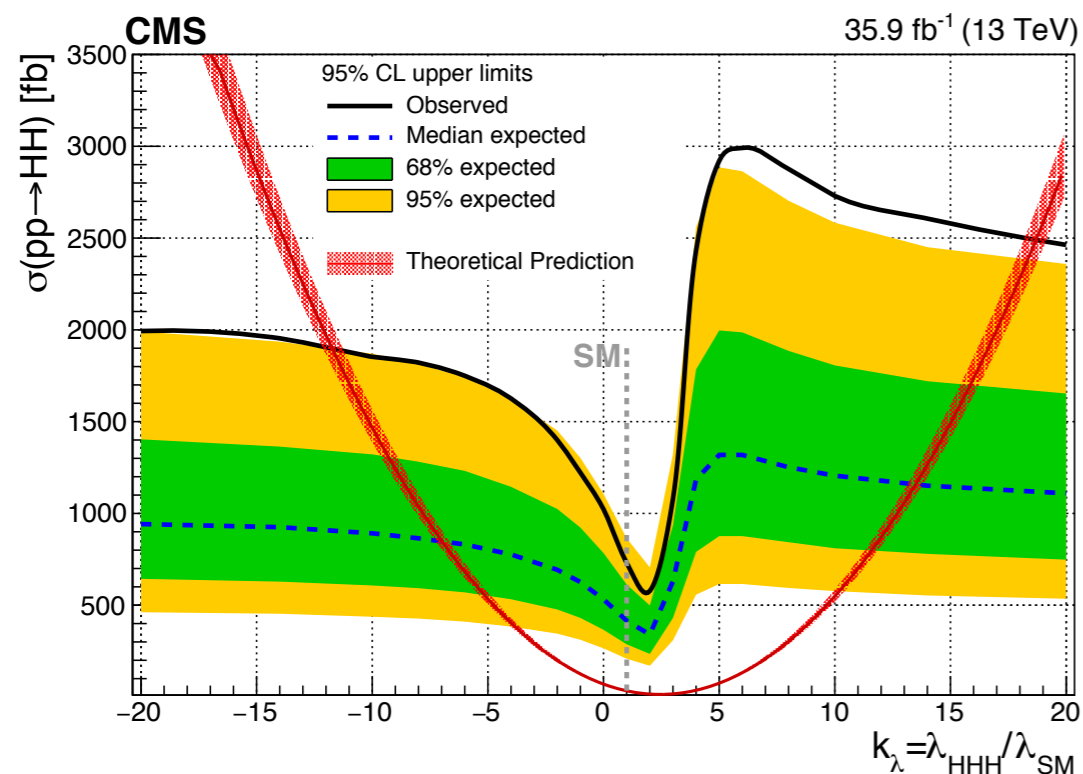
But small deviations from the SM can lead to huge x-sec increases!

# Results on Di-Higgs



Many orthogonal channels can be combined to increase sensitivity!

Note the use of b-jets: rare signals mean that high-BR decays of the Higgs need to be used

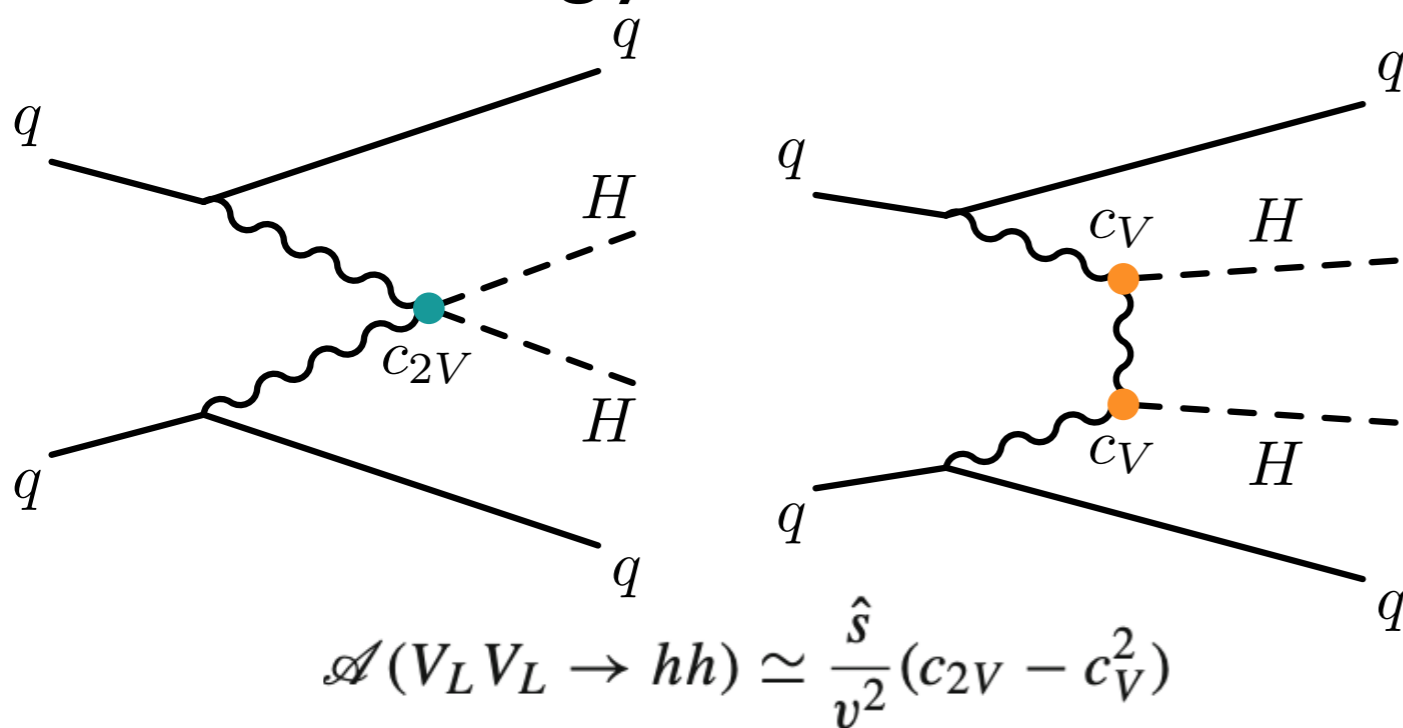


Wide range of couplings still allowed!  
Even more data still necessary

# Hunting Rare SM Deviations

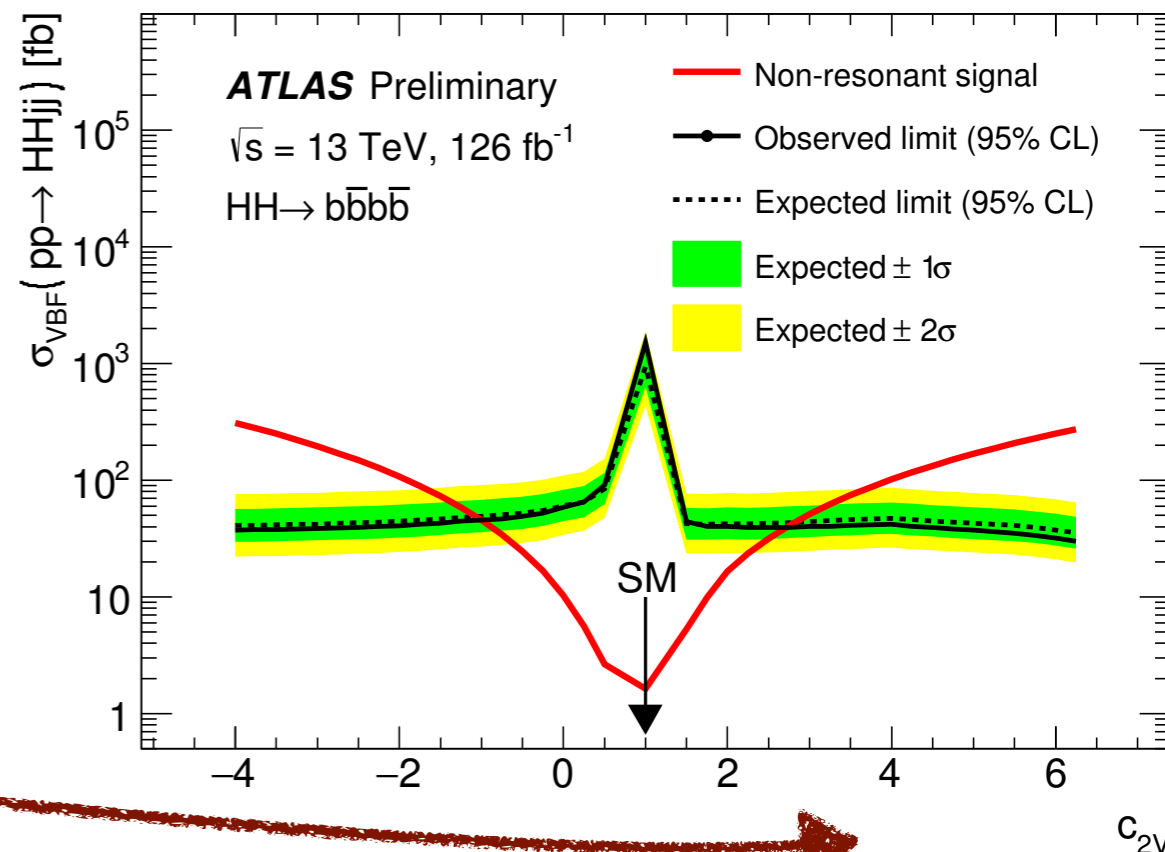
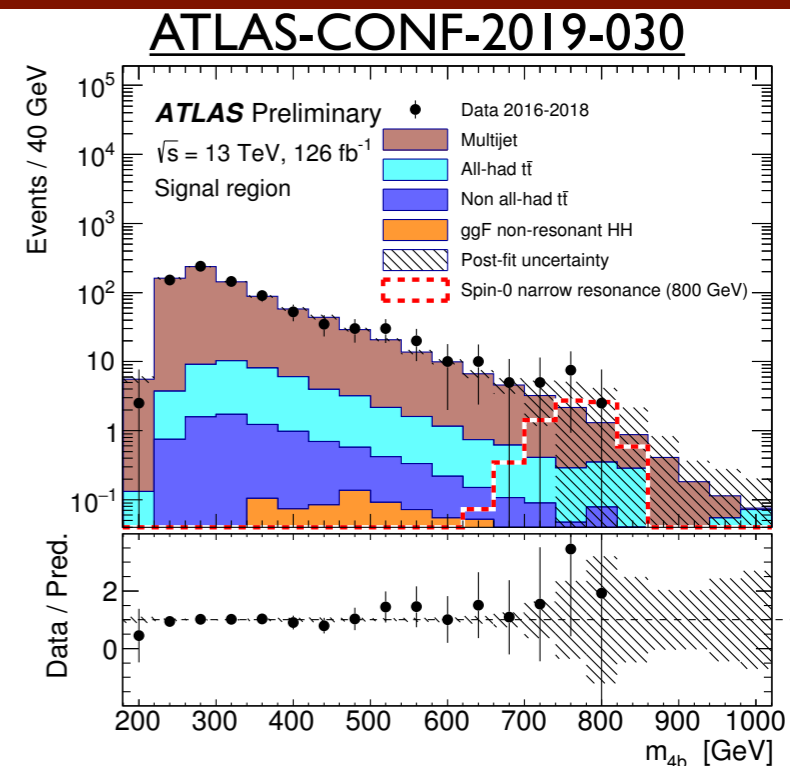


Interference in other SM diagrams lead to similar vanishingly small x-sec



VBF di-Higgs production is one example!

First limits on the 4-point VVHH coupling

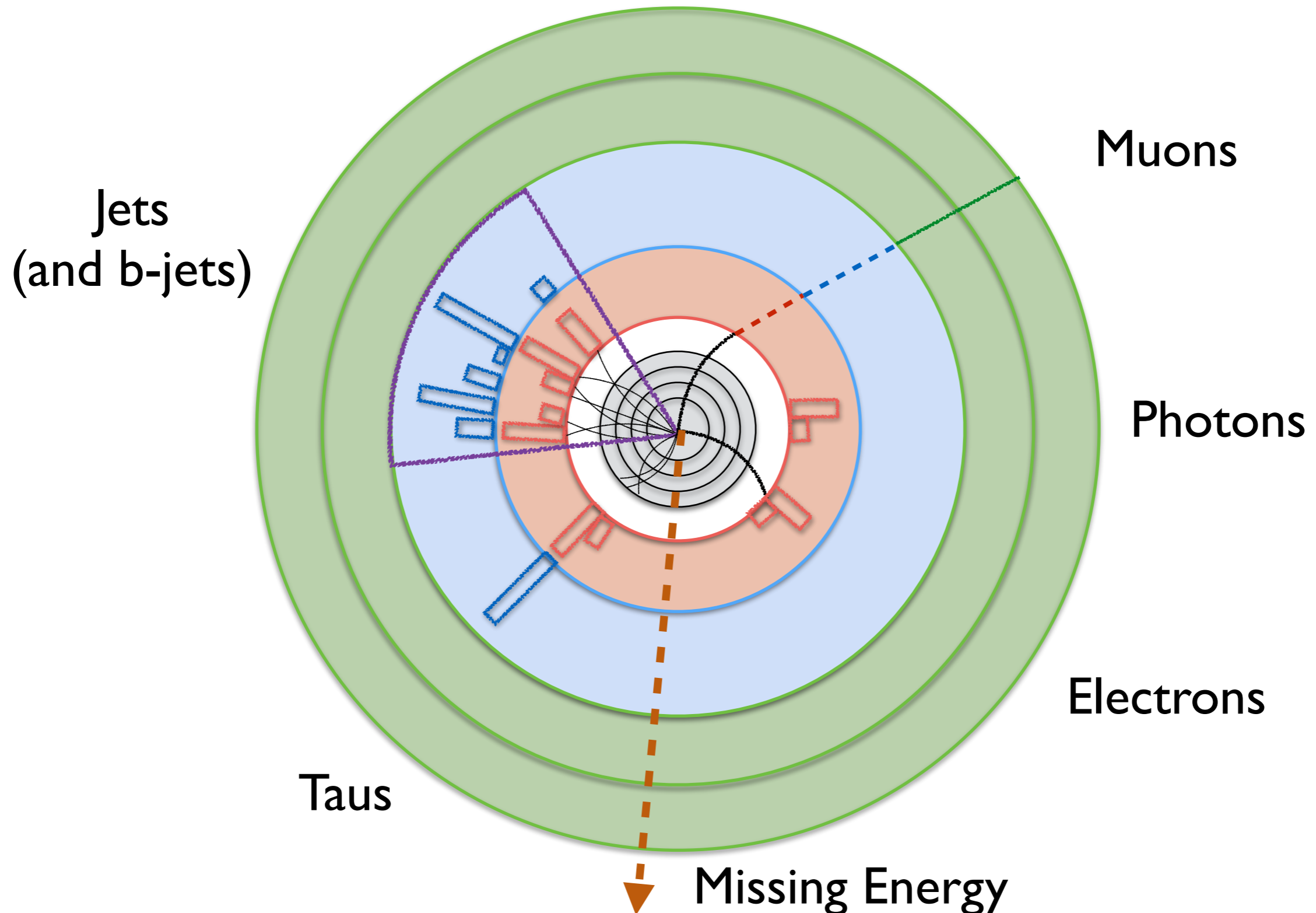




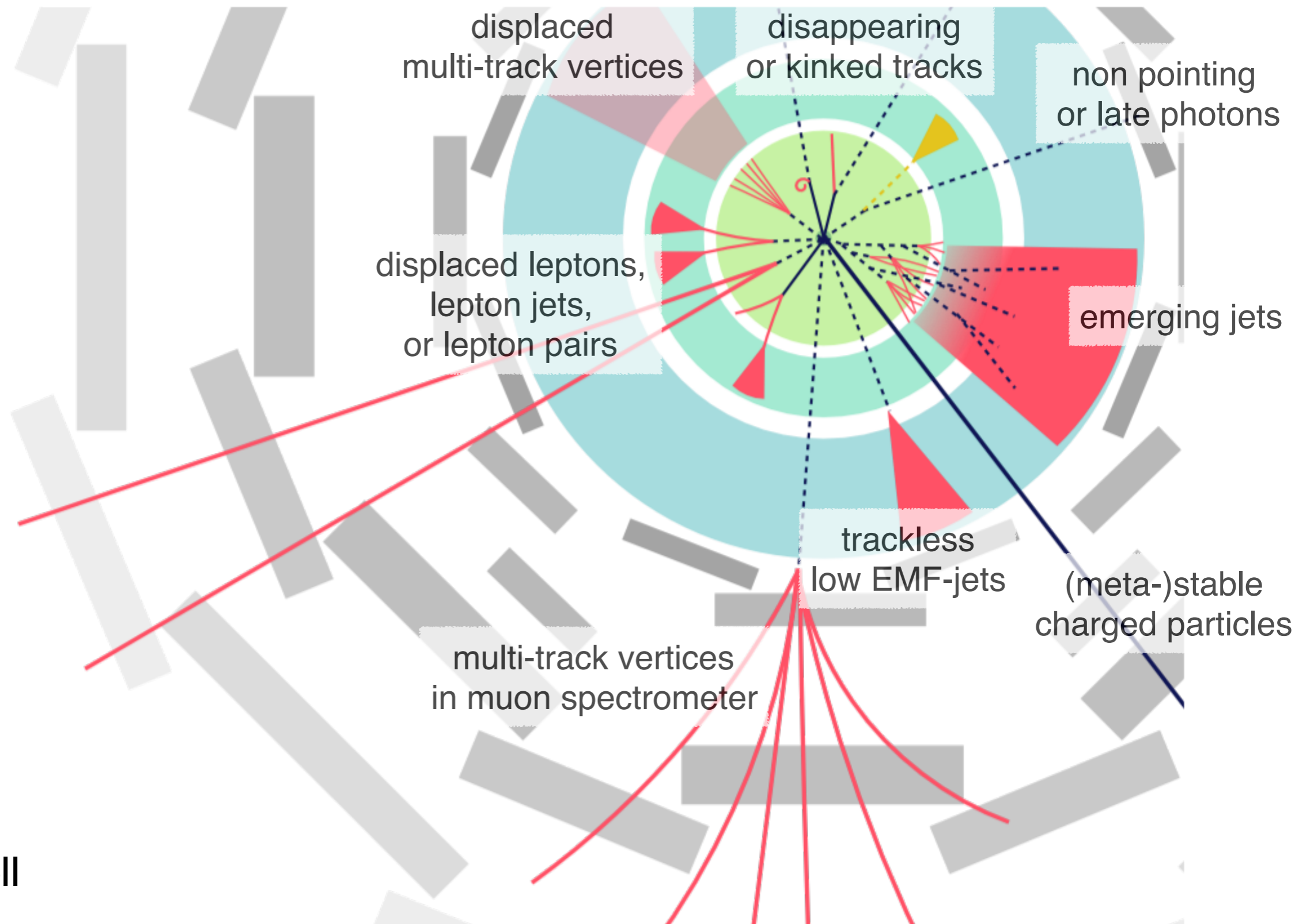
# Not Your Advisor's Signals

*Or: how new reconstruction strategies are expanding  
our sensitivity to signals our detectors weren't  
built to measure*

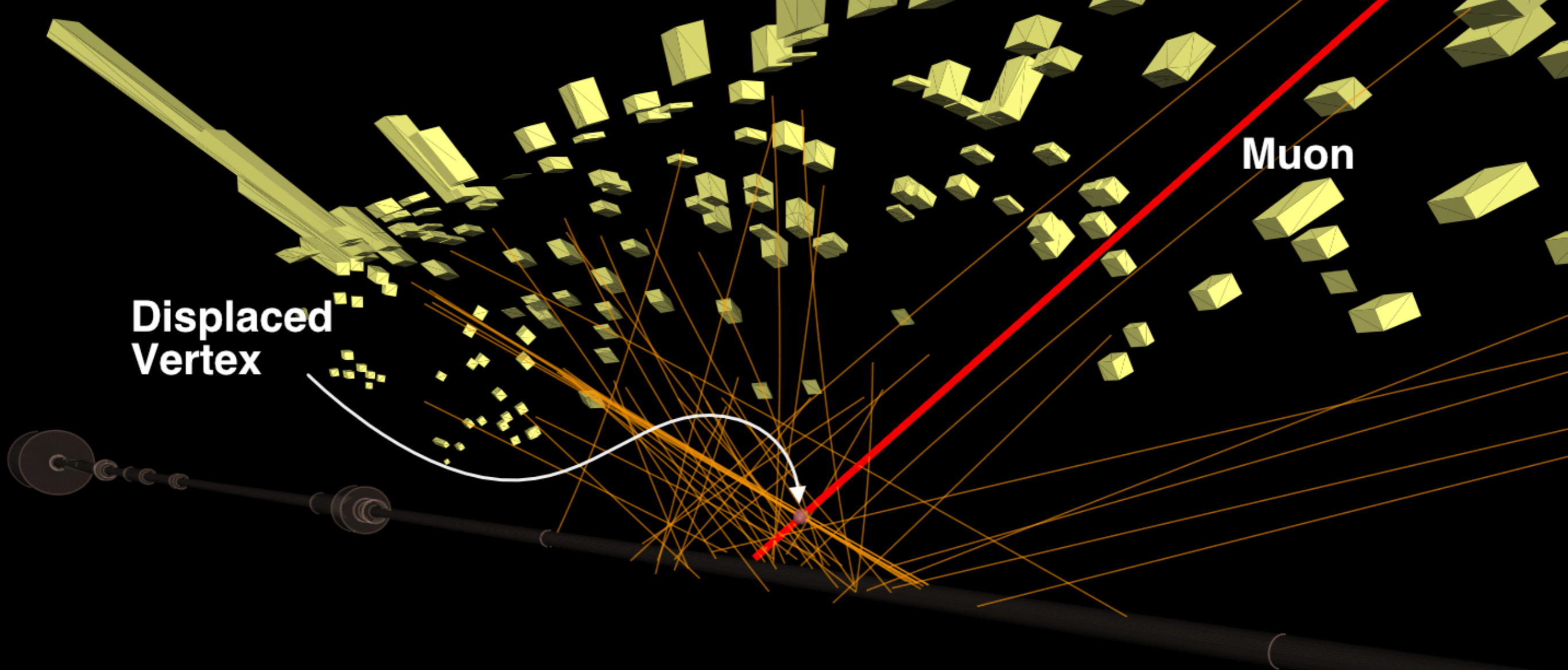
# Your Advisor's Signals



# Not Your Advisor's Signals



Heather Russell



Displaced Vertex

Muon

Long-lived particles will travel an appreciable distance before they decay

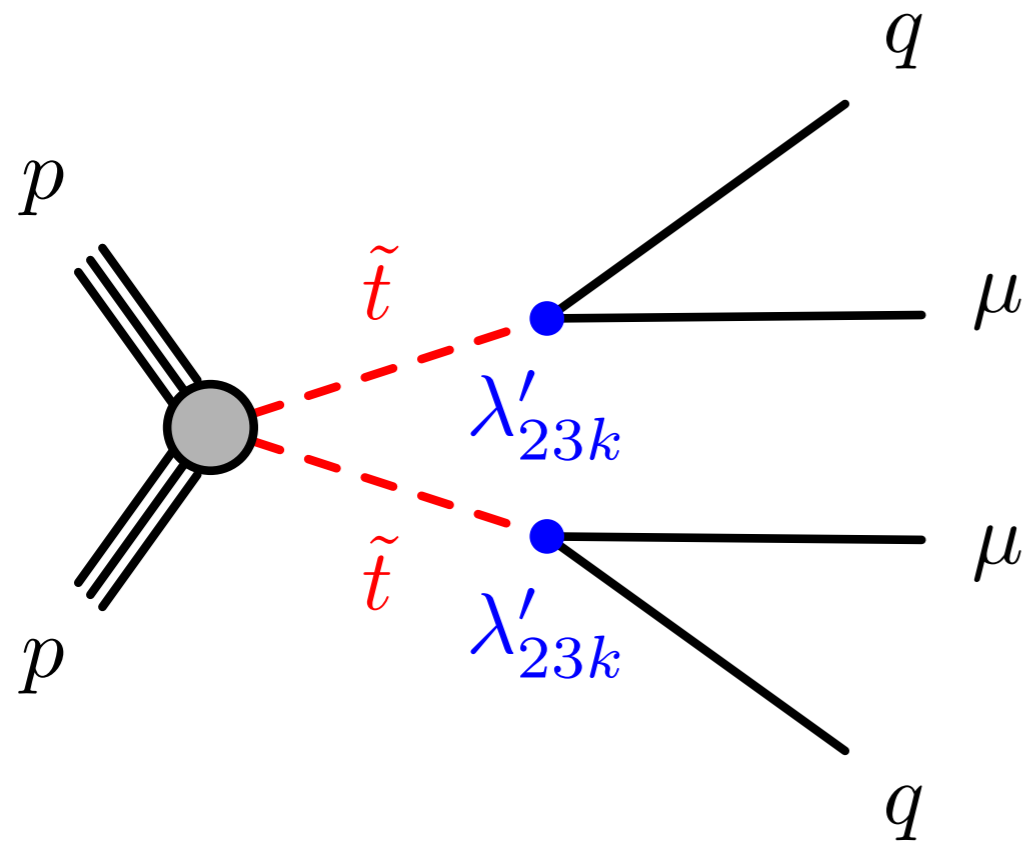
Simulated Signal Event  
Top Squark Pair Production

$$m(\tilde{t}) = 1.5 \text{ TeV}, \tau(\tilde{t}) = 1 \text{ ns}$$
$$\tilde{t} \rightarrow \mu j$$

Special algorithms can reconstruct these “displaced tracks” and look for displaced vertices



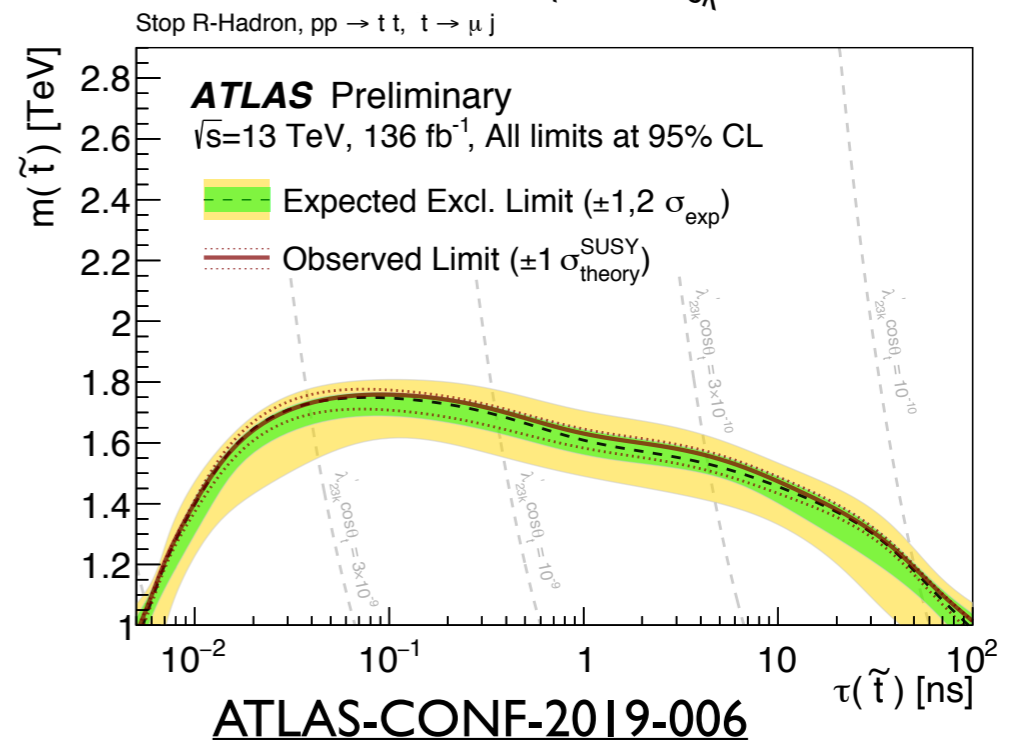
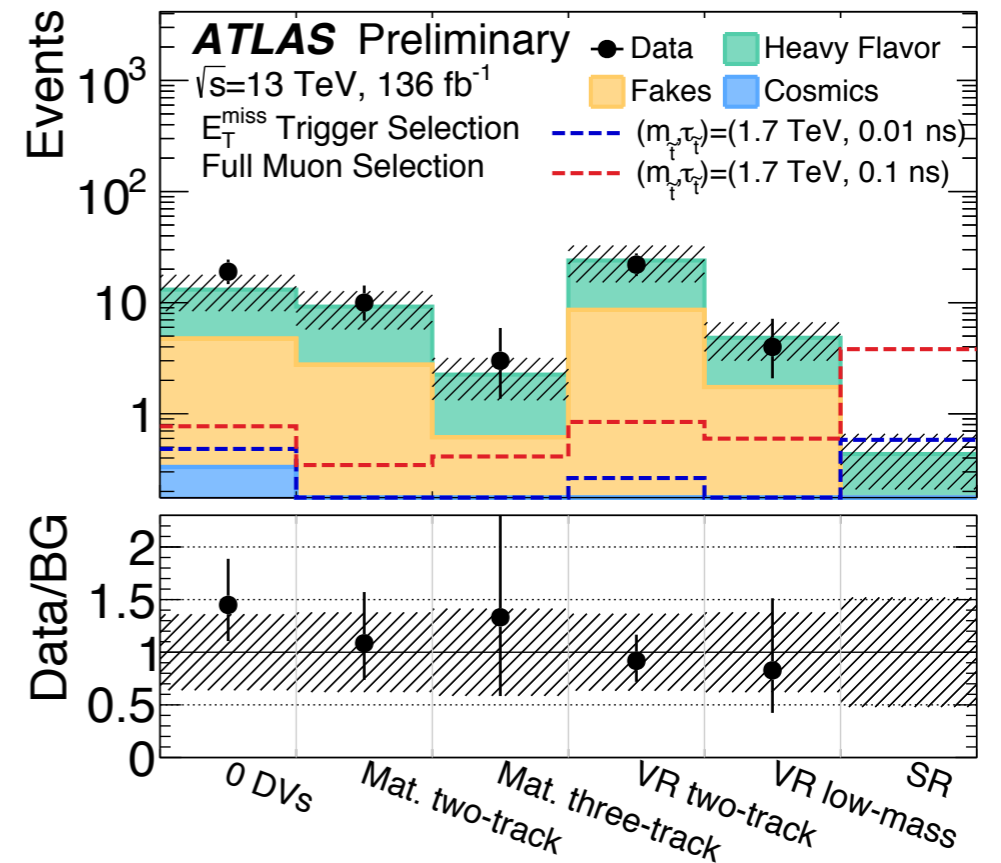
# Displaced Vertex + $\mu$



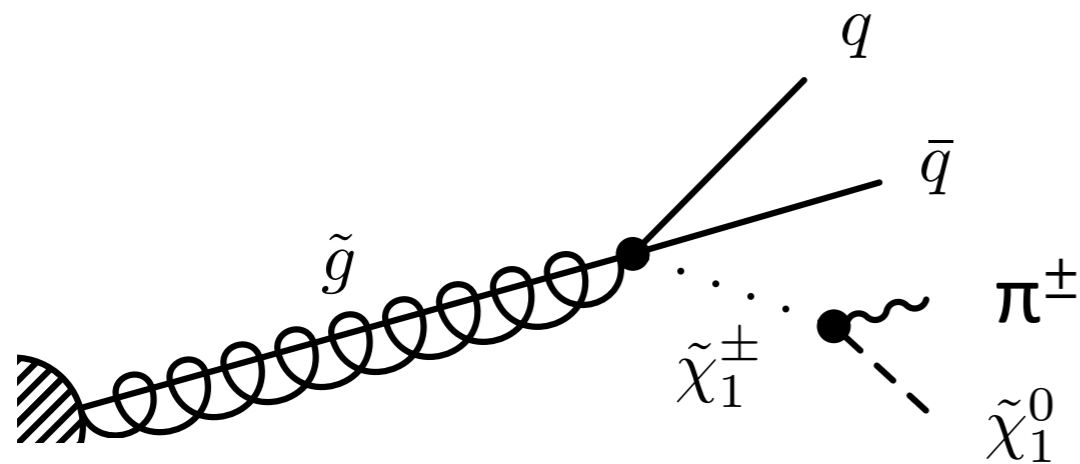
Stops can be long-lived in RPV models with small couplings

Data-driven backgrounds estimate  $\sim 0$  background, with no signal observed

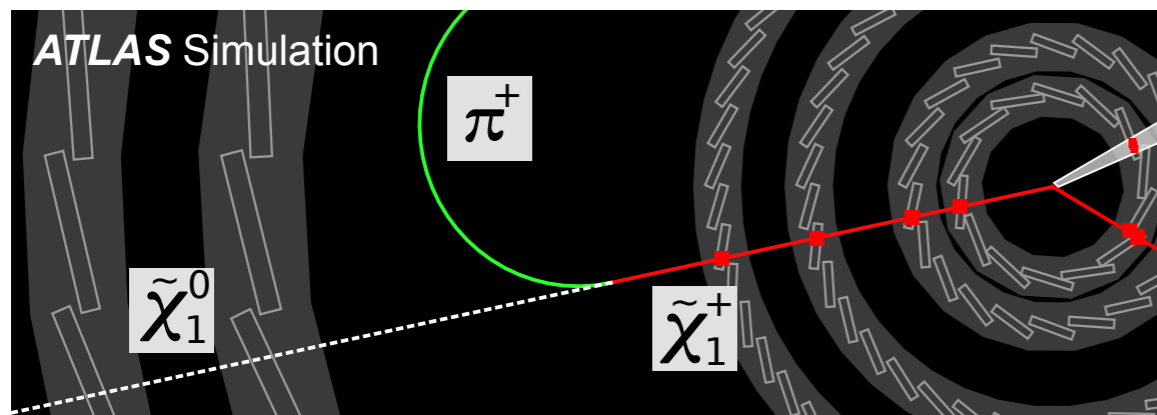
Strongest limits yet on stops!



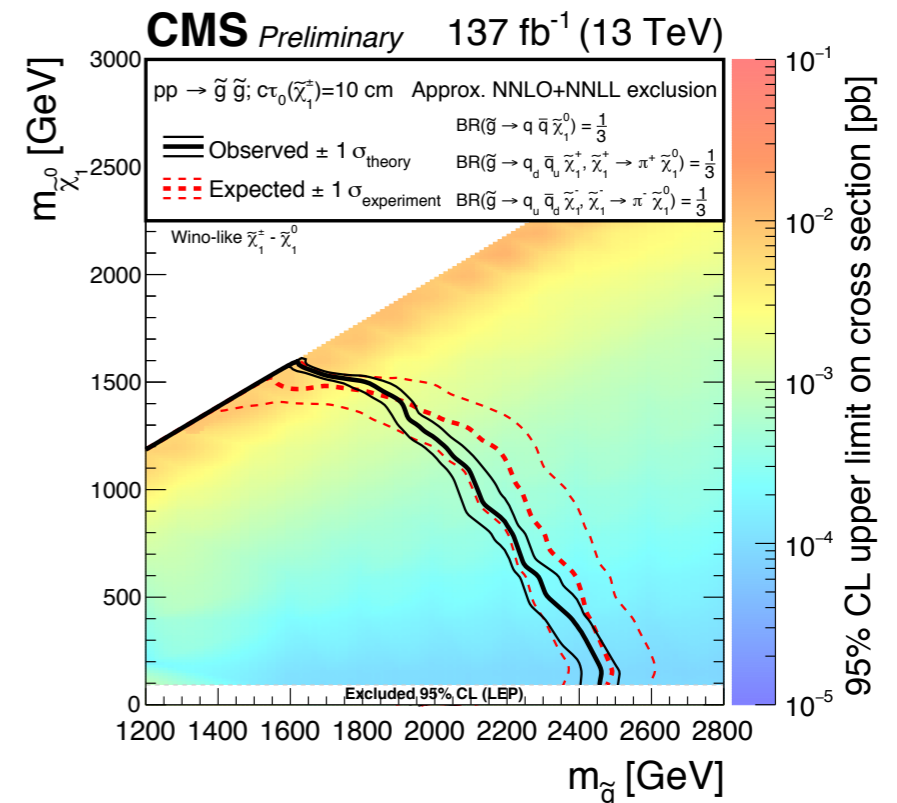
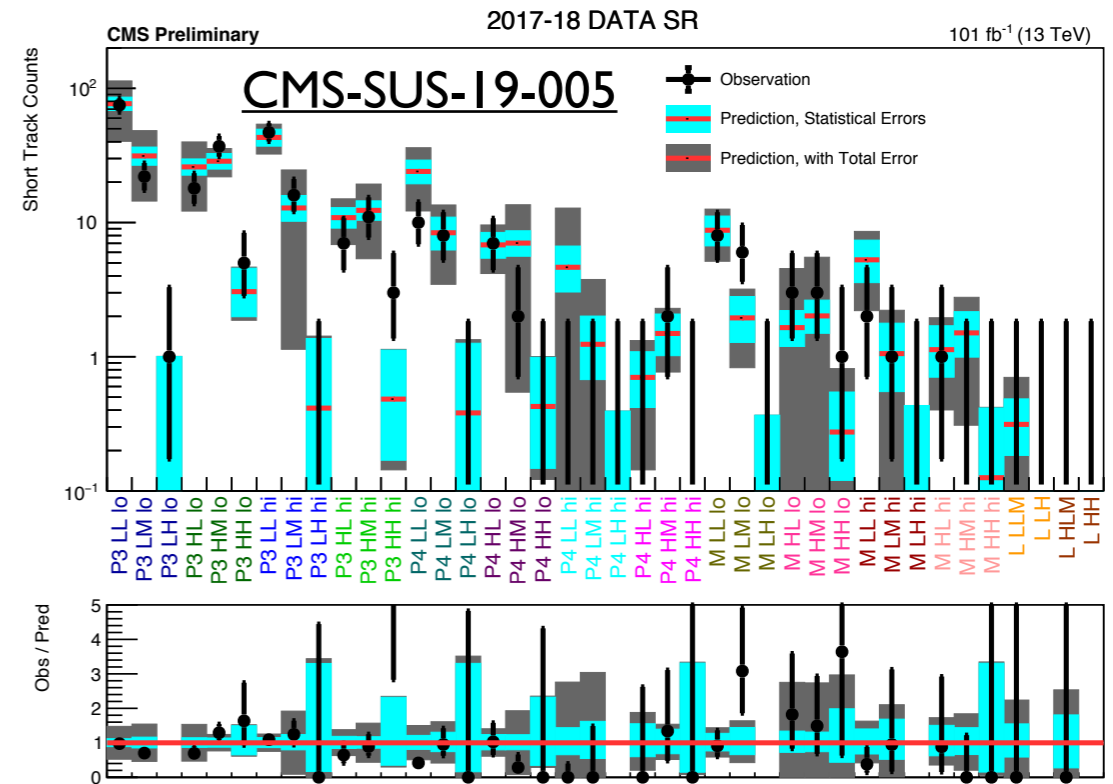
# Disappearing Tracks



With extremely small mass splittings between SUSY particles, charginos can become long-lived!



Can interact with the first few layers of the tracker, and then “disappear”



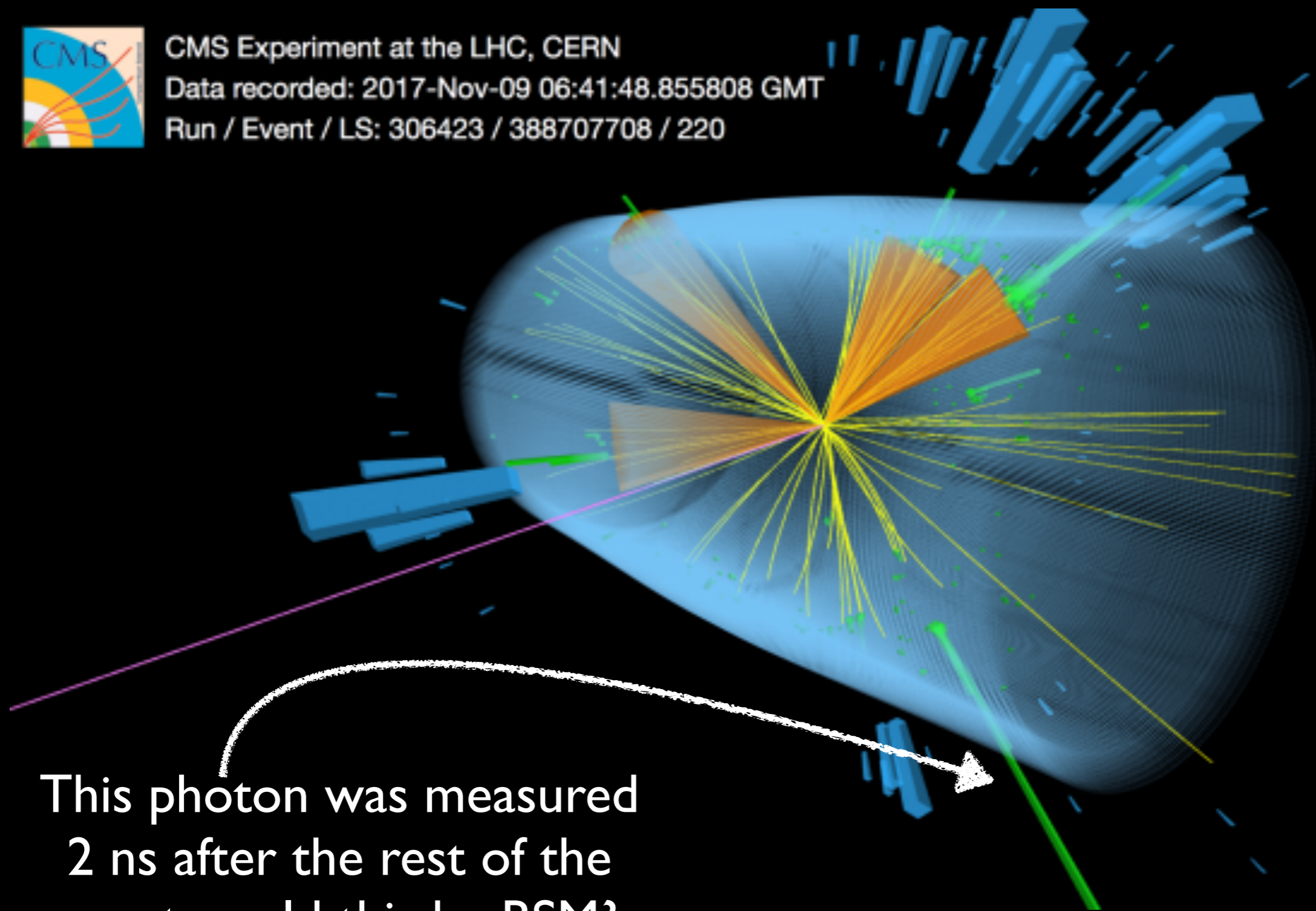
# Can also extend reconstruction using additional information: **timing**



CMS Experiment at the LHC, CERN

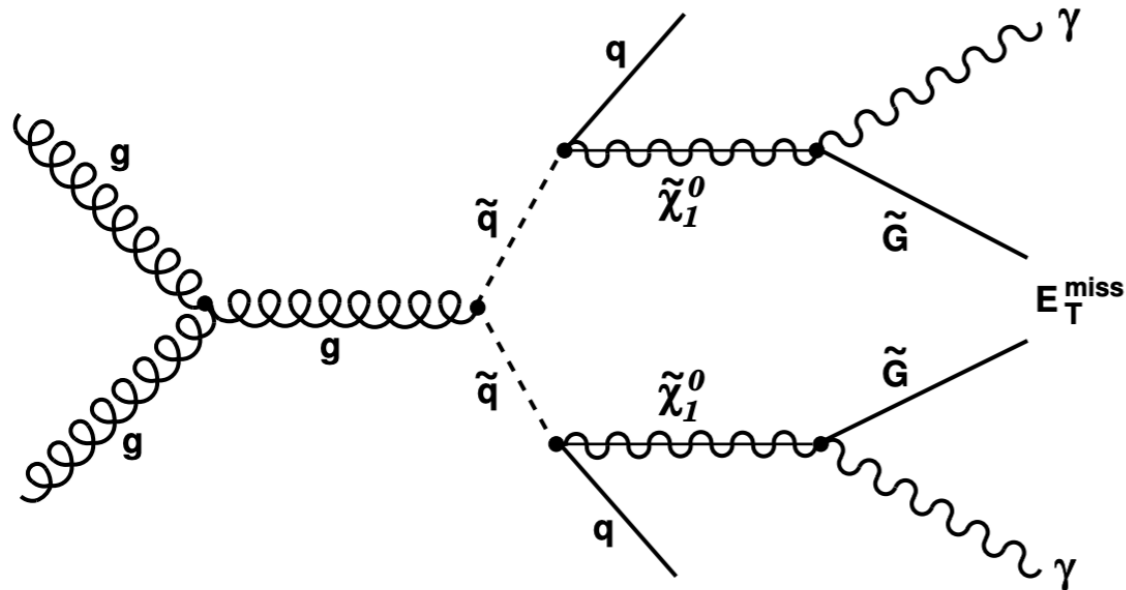
Data recorded: 2017-Nov-09 06:41:48.855808 GMT

Run / Event / LS: 306423 / 388707708 / 220



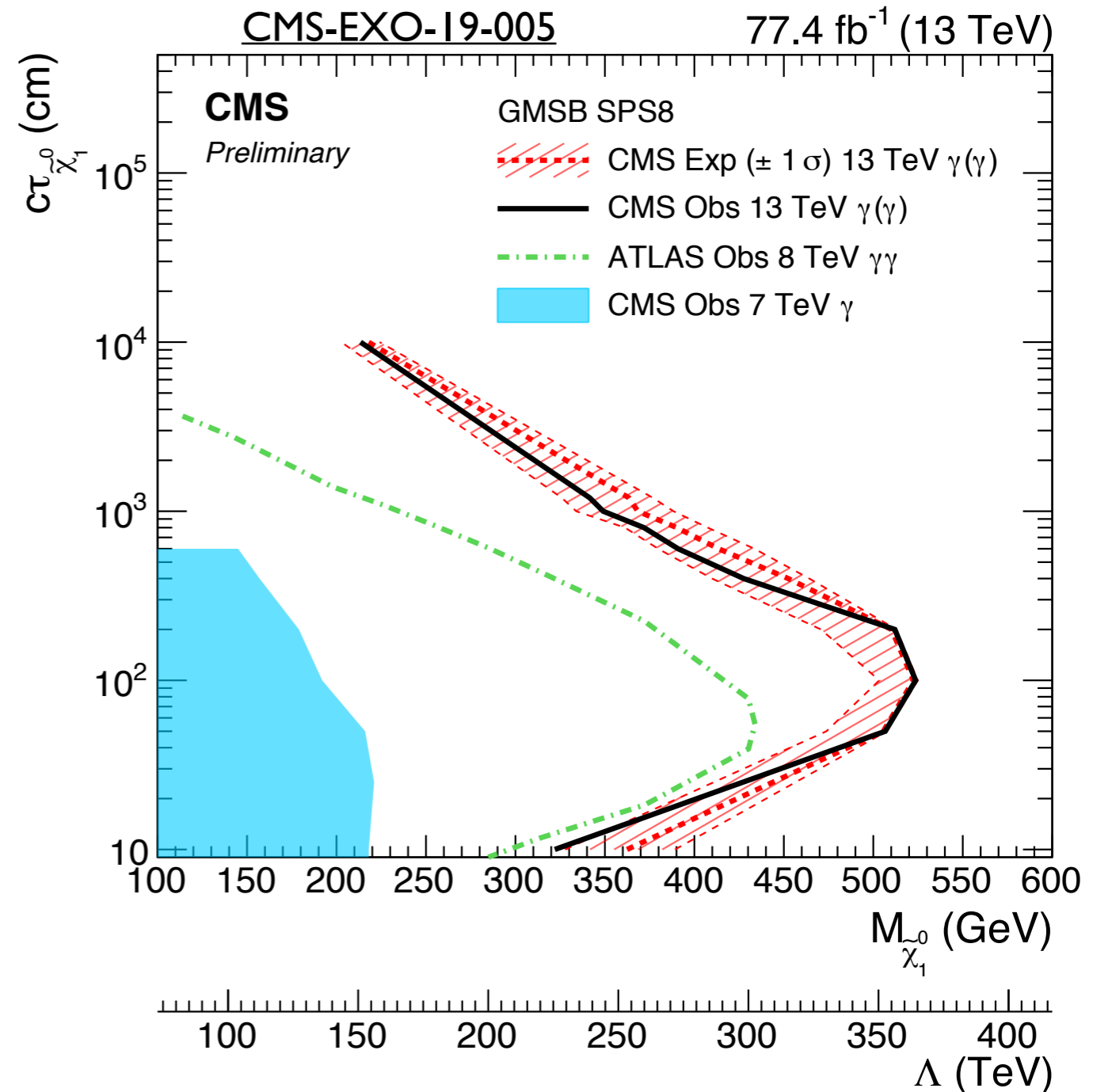
This photon was measured  
2 ns after the rest of the  
event: could this be BSM?

# Delayed Photons



Many models (GMSB SUSY, etc.) have long-lived particles decaying to photons

These appear “late” compared to other energy in the event: use crystal timing in CMS to search





# Conclusions

# Keep Looking!



BSM is extremely well motivated  
at the LHC!

Naturalness, dark matter, and  
baryogenesis all need answers,  
and the LHC could find them

Searches are moving  
into a new era: the low-  
hanging fruit has been  
picked, and the  
challenge is increasing



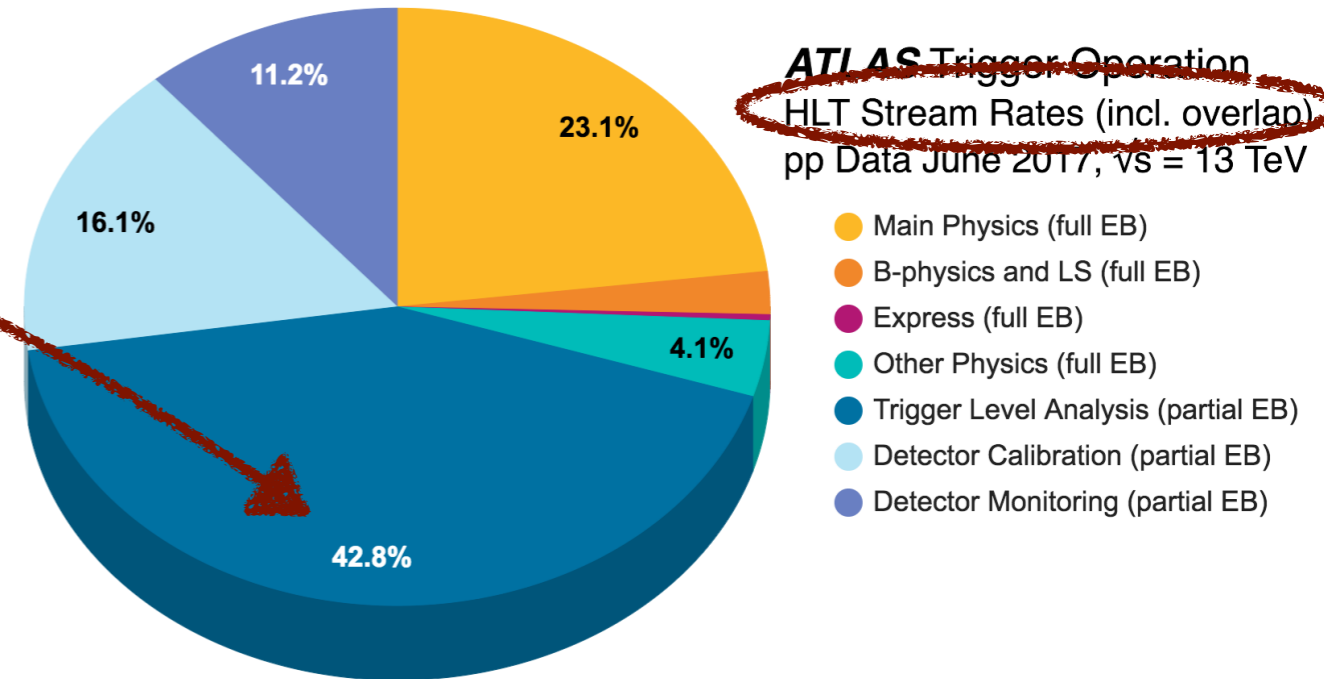
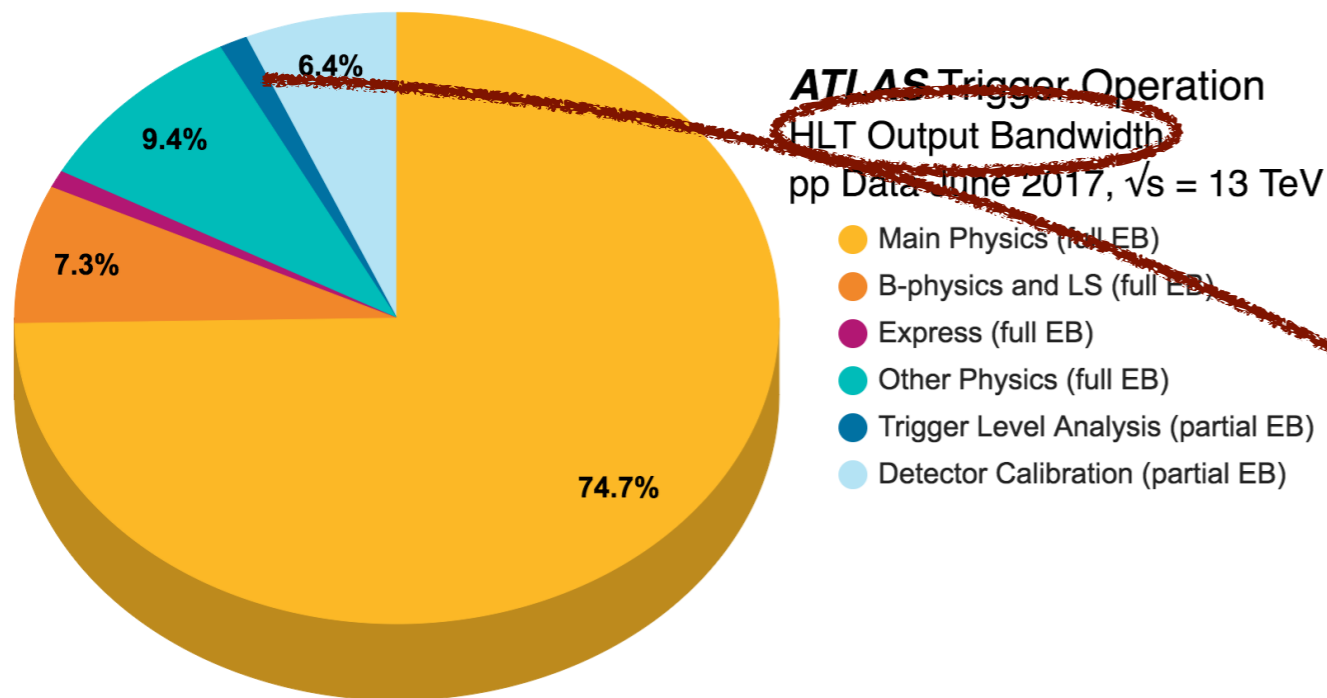
Low Hanging Fruit Records

But new tools, large datasets, and creative reconstruction  
mean we can rise to the challenge!

**Thank you!**

# Backup

# Data Scouting



Most of the time, record the “full event” when triggered

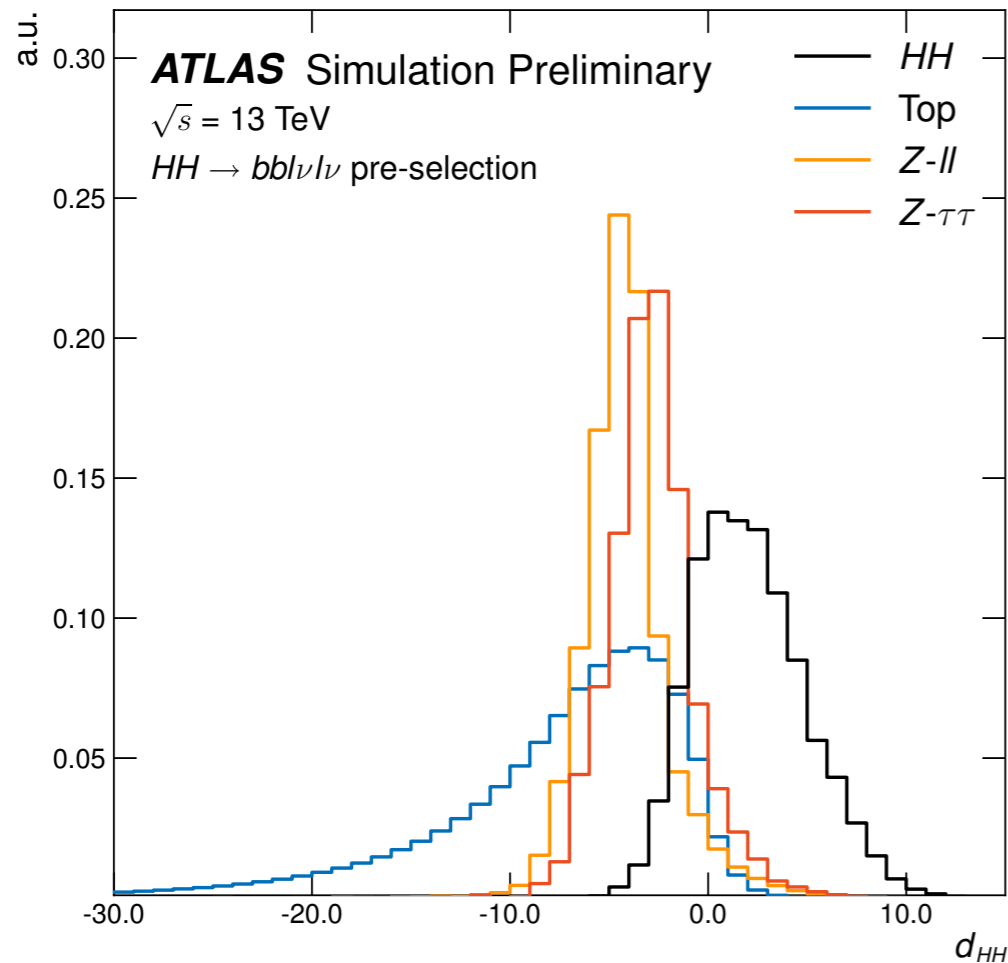
But for some analyses, we don't need the full event:  
just the jets can be enough to do physics!

If you make the event small, then  
you can record a lot more data!

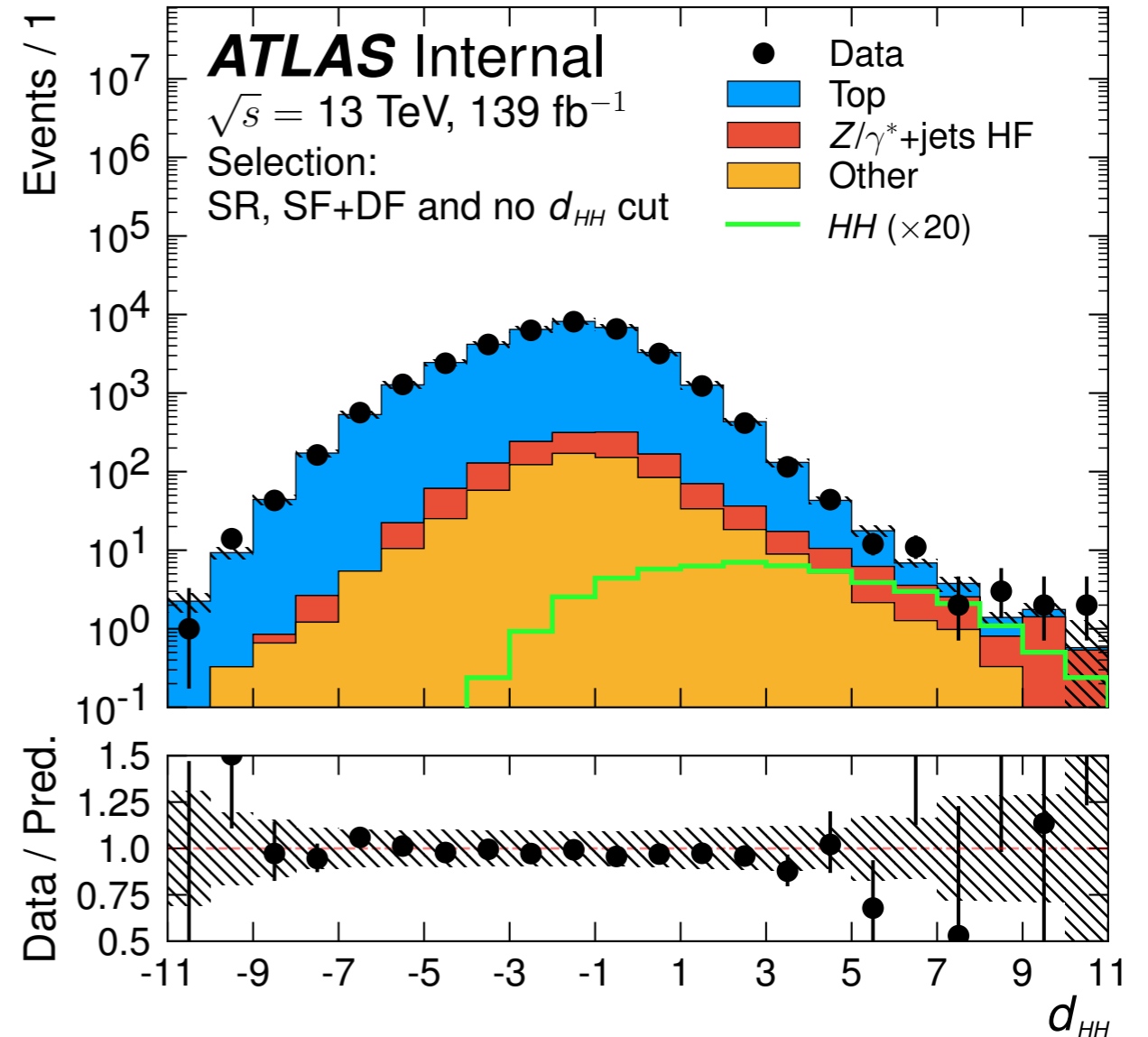
# Machine Learning for di-Higgs



Rare signals are an excellent target for machine learning!



ATLAS uses a DNN to search for di-Higgs in  $bbl$  final state



Machine learning can make this rarer channel competitive in the hunt for di-Higgs!

# CMS Jet Timing



# ATLAS WH results

