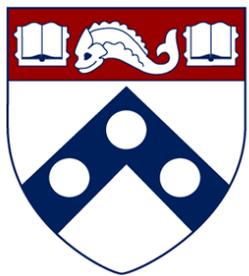


Dichroic Winston Cones (‘Dichroicons’) for Cherenkov and Scintillation Light Separation in Large-Scale Neutrino Detectors

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University of Pennsylvania

July 2019



Penn
UNIVERSITY of PENNSYLVANIA

APS | DIVISION OF
PARTICLES & FIELDS

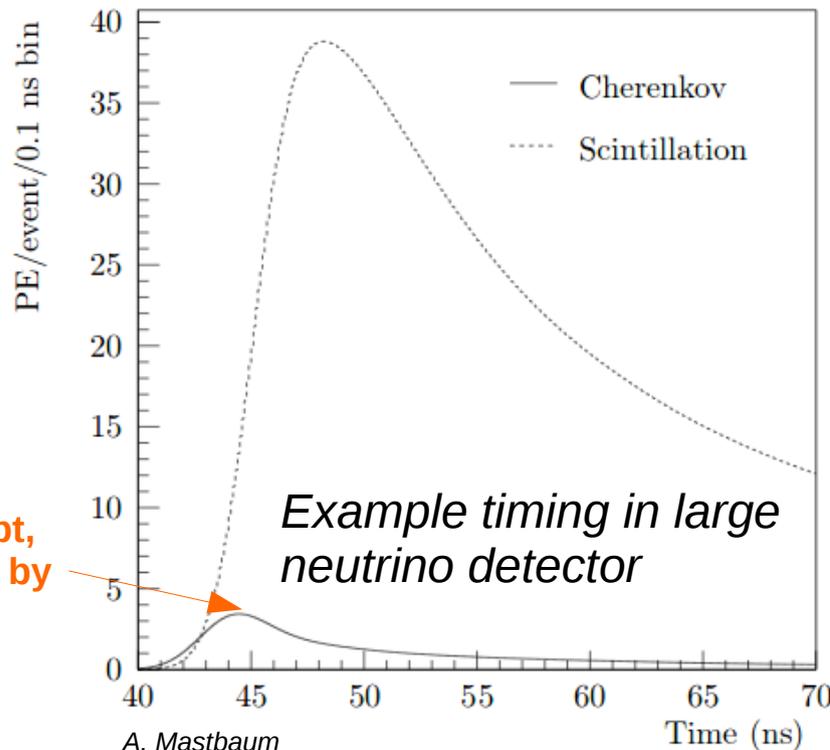
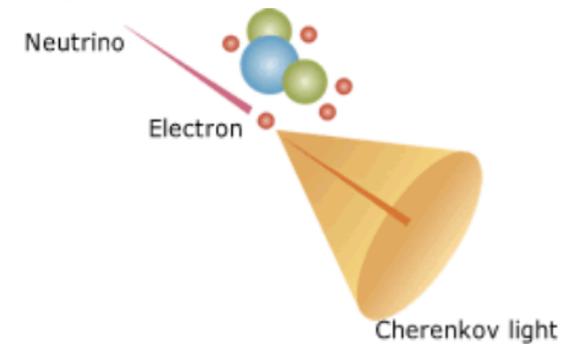


Detecting Cherenkov Light in a Scintillation Detector

- Charged particle traveling through liquid scintillator creates *both* scintillation (~10,000 photons/MeV) *and* Cherenkov light (~100 photons/MeV)
- Challenge is to detect the Cherenkov light, which provides the direction of the traveling particle
- Cherenkov light is emitted promptly, scintillation light delayed and emitted with slower time-profile

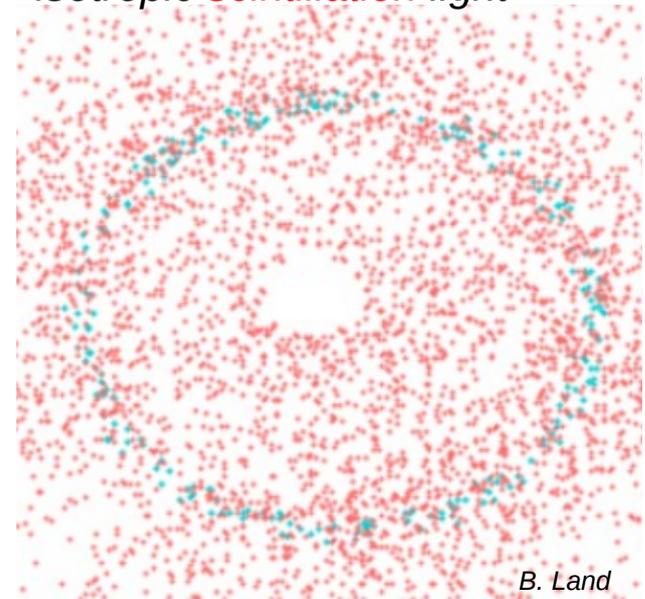


Cherenkov light emitted in a cone around particle direction



Cherenkov light prompt, but spread by detector response

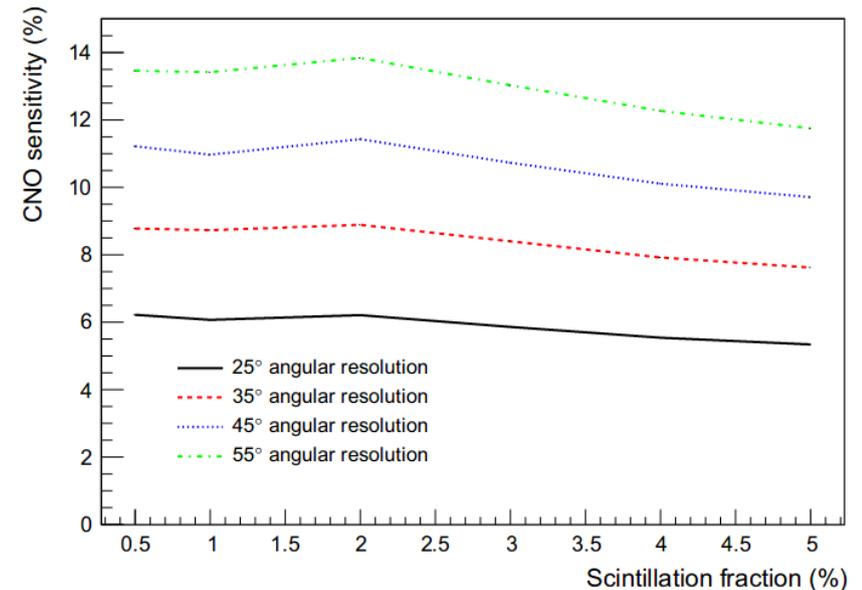
Cherenkov ring on top of isotropic scintillation light



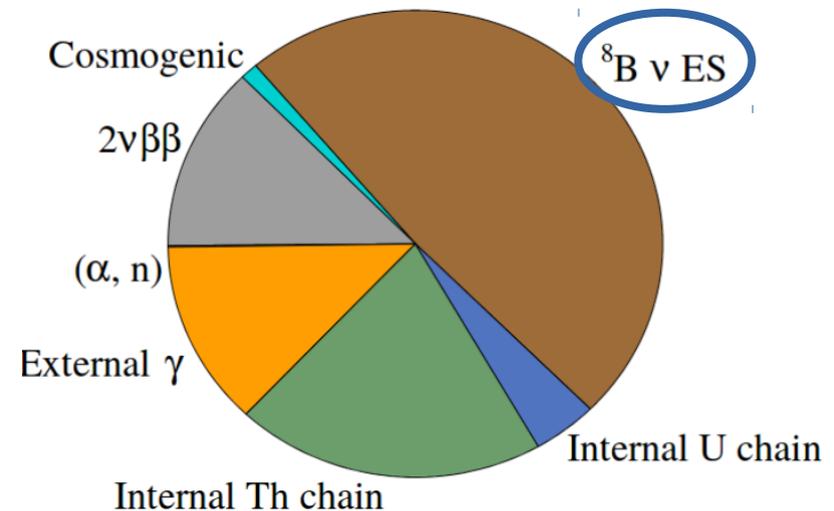
Why Cherenkov / Scintillation Separation?

- Recent trend in large monolithic neutrino detectors to move toward scintillator targets (Borexino, kamLAND-Zen, SNO+, JUNO, etc.)
- High light yield from scintillator provides excellent energy / position resolution, low energy thresholds
Neutrinoless double beta decay, low energy solar neutrinos, reactor & geo antineutrinos
- Cherenkov light allows one to reconstruct direction, improve particle ID
- Allow scintillator-based solar neutrino experiments to suppress backgrounds
- Largest expected backgrounds for SNO+ is solar neutrinos, which can be rejected by detecting the Cherenkov light
- Has not been demonstrated in an existing liquid scintillator detector
- R&D work will be focus of this talk

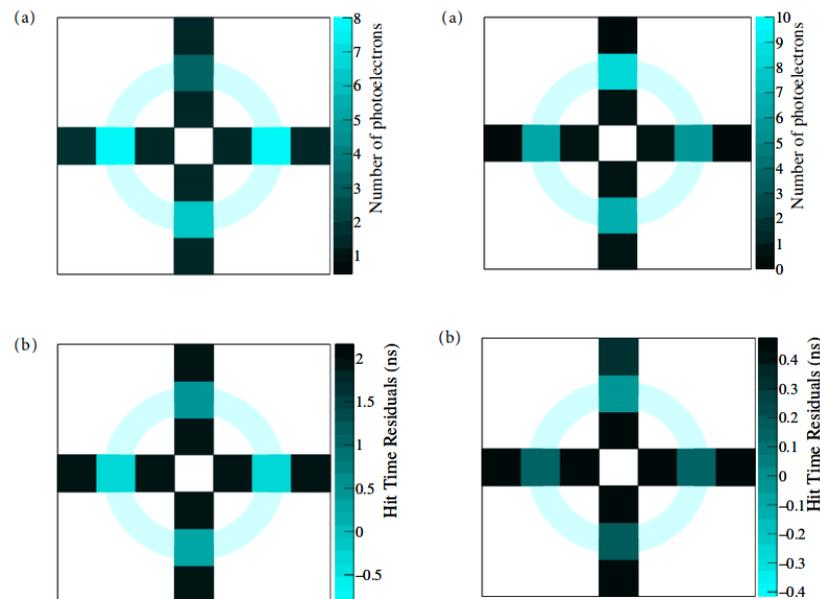
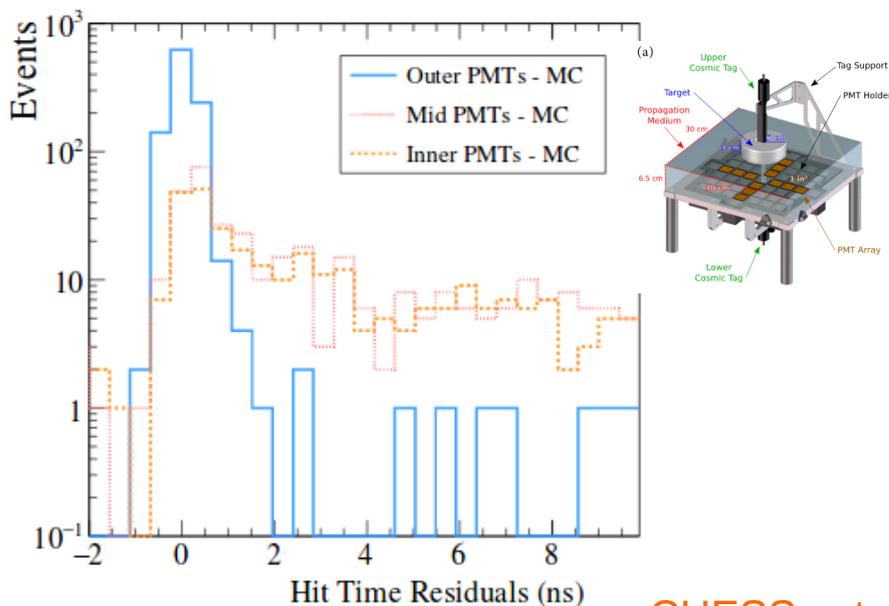
R. Bonventre, G.D. Orebi Gann, Eur. Phys. J. C (2018) 78:435



Expected background for SNO+ $0\nu\beta\beta$

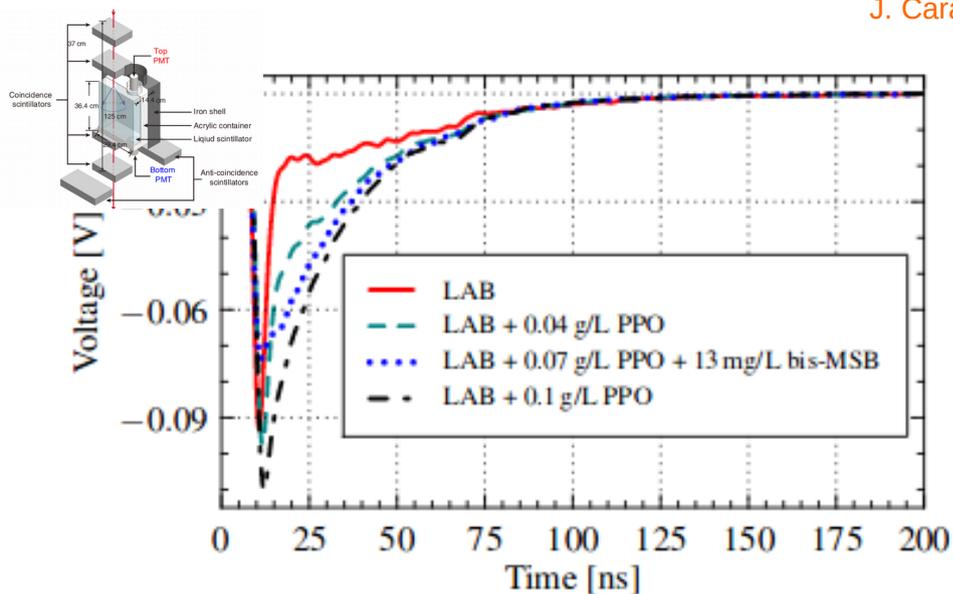


Ongoing R&D For Cherenkov / Scintillation Separation

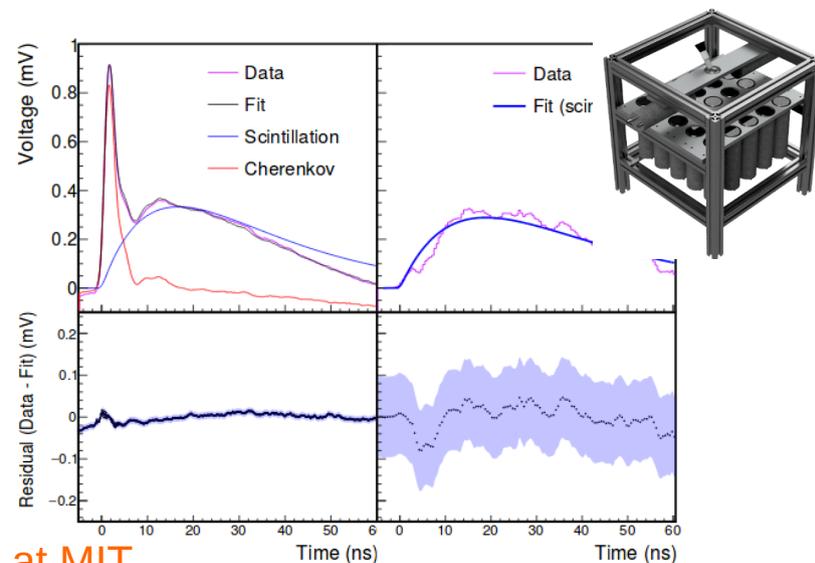


CHES setup at LBNL

J. Caravaca et. al, [10.1103/PhysRevC.95.055801](https://arxiv.org/abs/10.1103/PhysRevC.95.055801)



Slow scintillator characterization for Jinping
 Z. Guo et. al, [10.1016/j.astropartphys.2019.02.001](https://arxiv.org/abs/10.1016/j.astropartphys.2019.02.001)

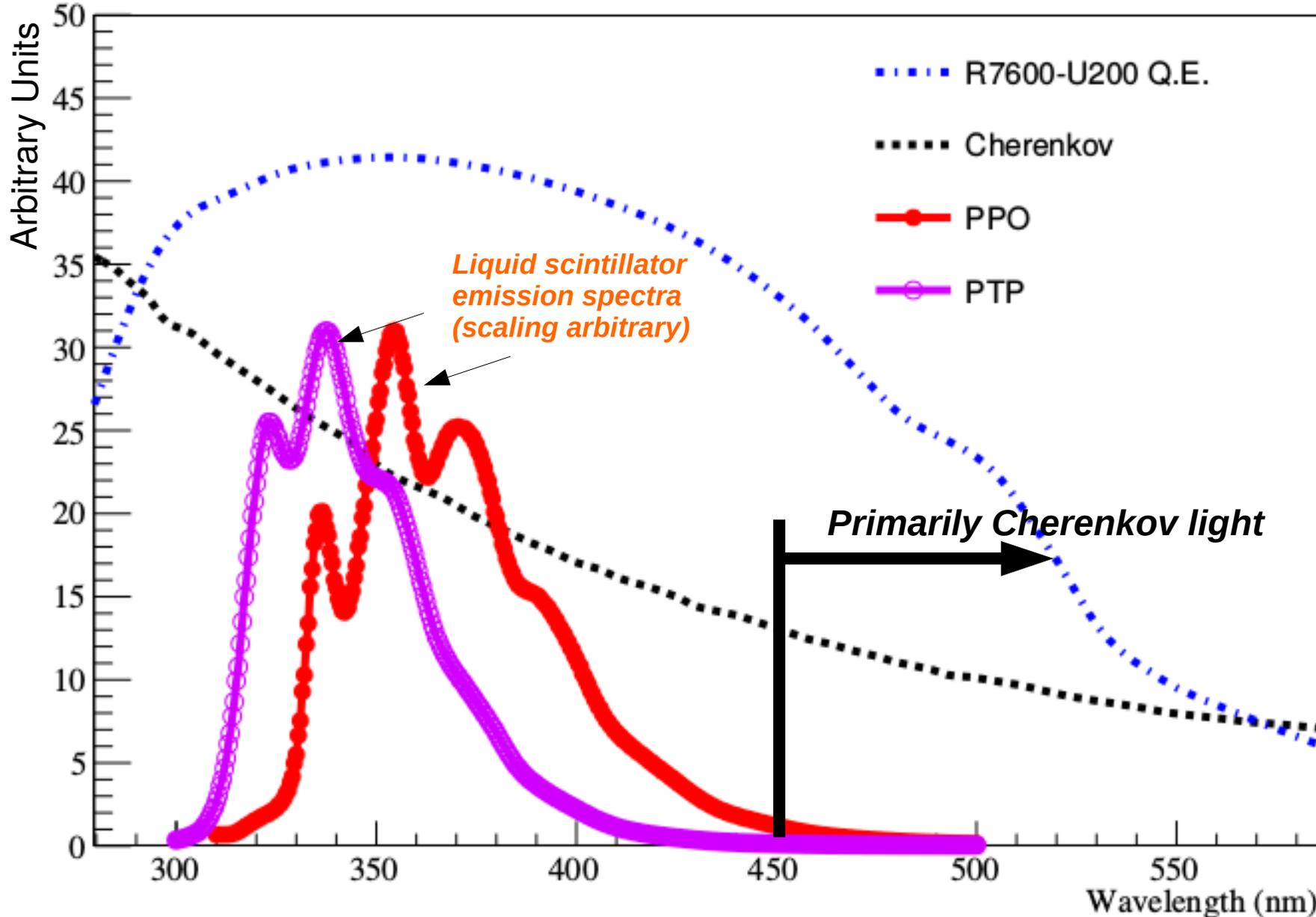


FlatDot at MIT

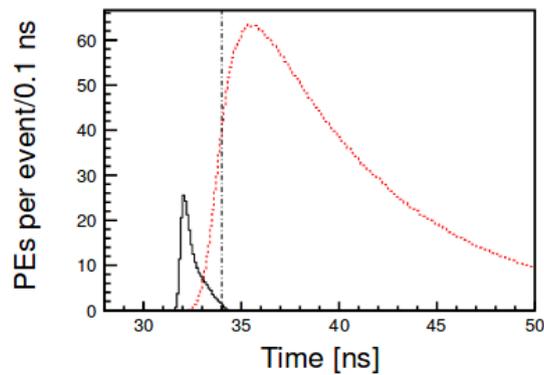
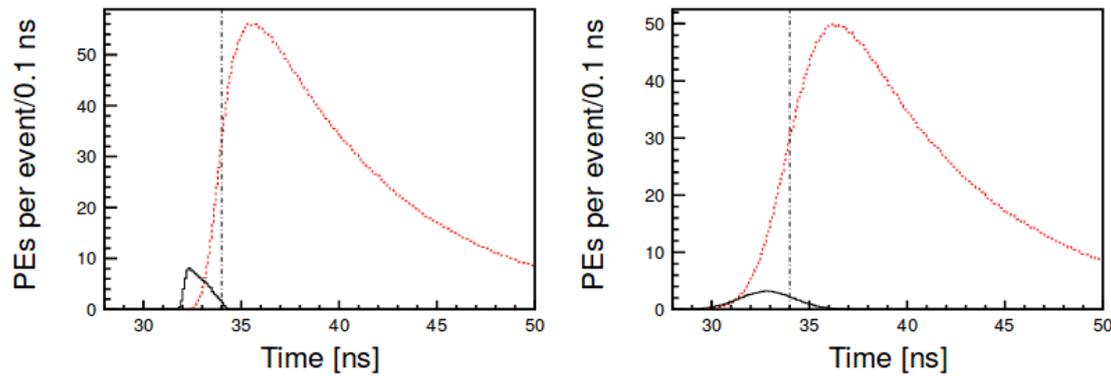
J. Gruszko, et. al, [10.1088/1748-0221/14/02/P02005](https://arxiv.org/abs/10.1088/1748-0221/14/02/P02005)

Only timing and isotropy used to identify the Cherenkov light.

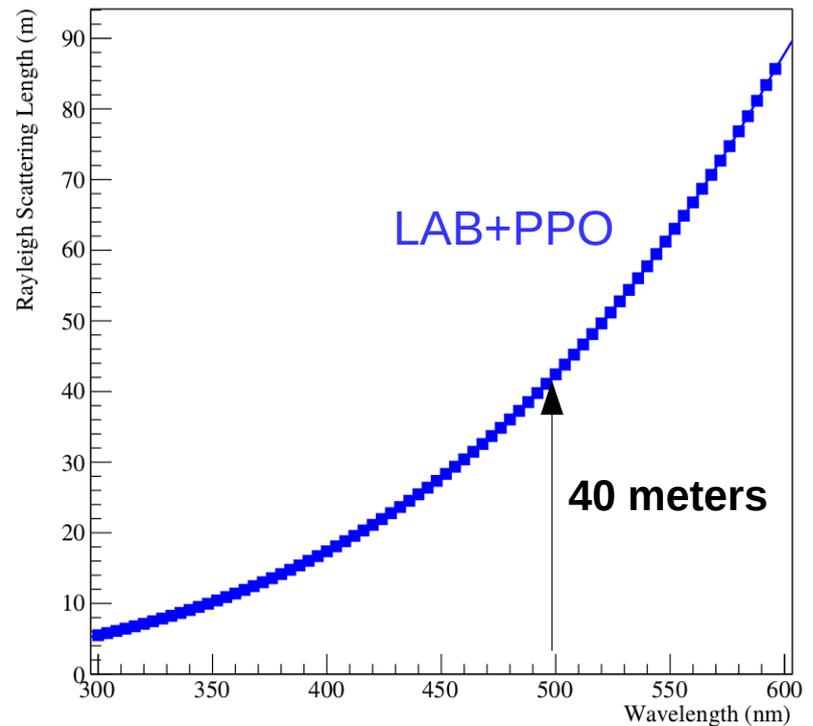
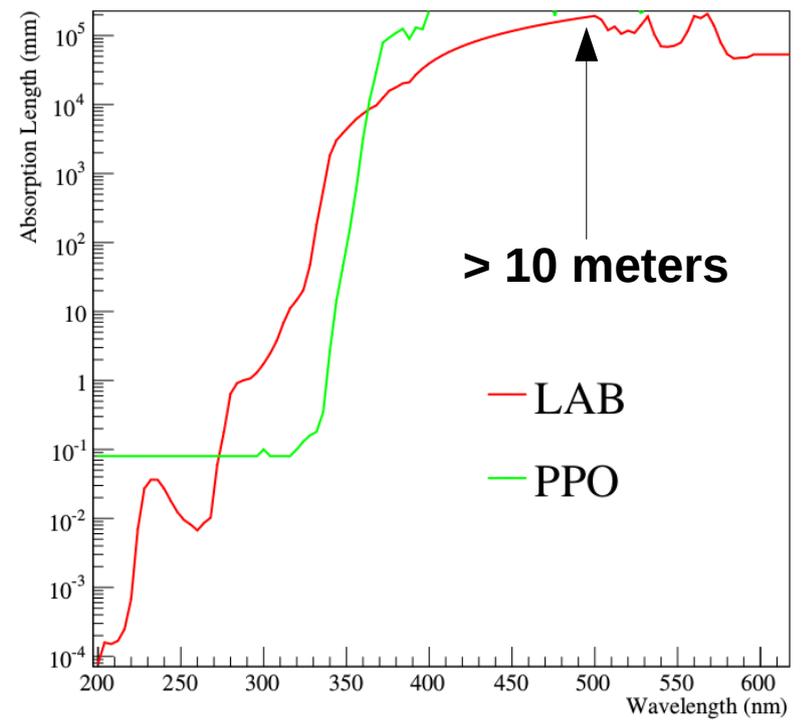
Can we additionally use wavelength?



Ultimately we want to achieve Cherenkov and scintillation separation while losing as few total photons as possible.

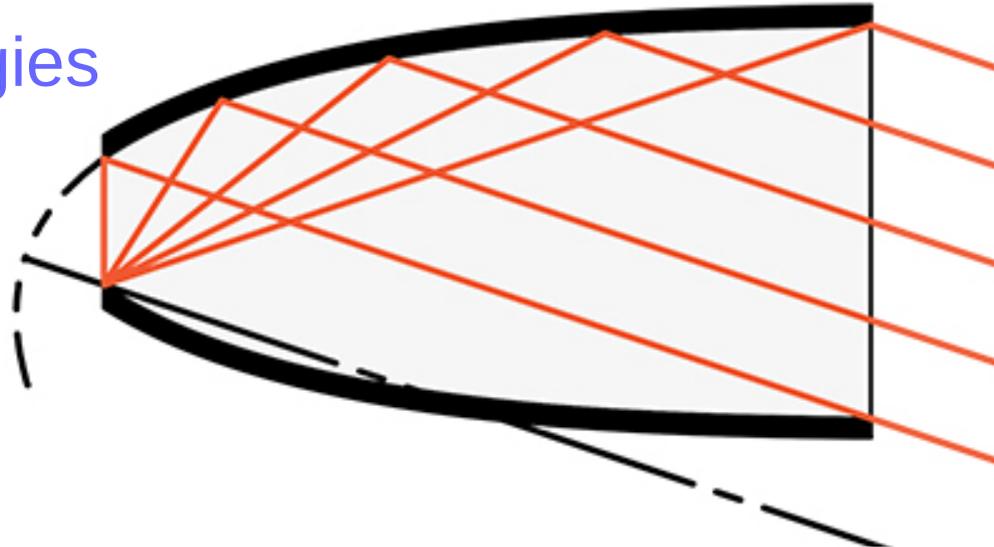


- Simulations of kamLAND-like detector show red-sensitive photocathodes improve Ch/Sc separation
- Absorption and scattering length increases with wavelength
- Cherenkov light that reaches the PMTs with its original direction is primarily long-wavelength



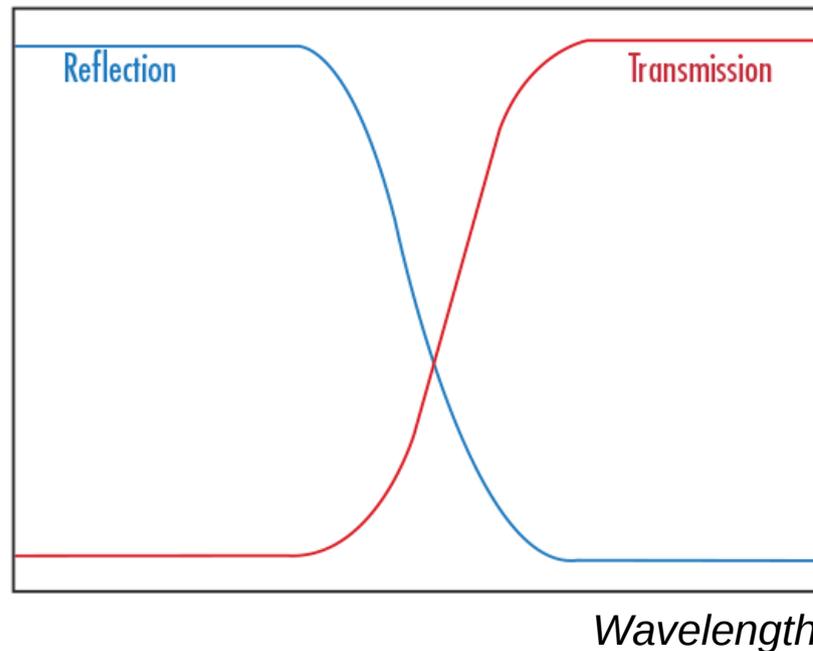
Dichroic Winston Cones: *The Dichroicon*

Combining Two Technologies



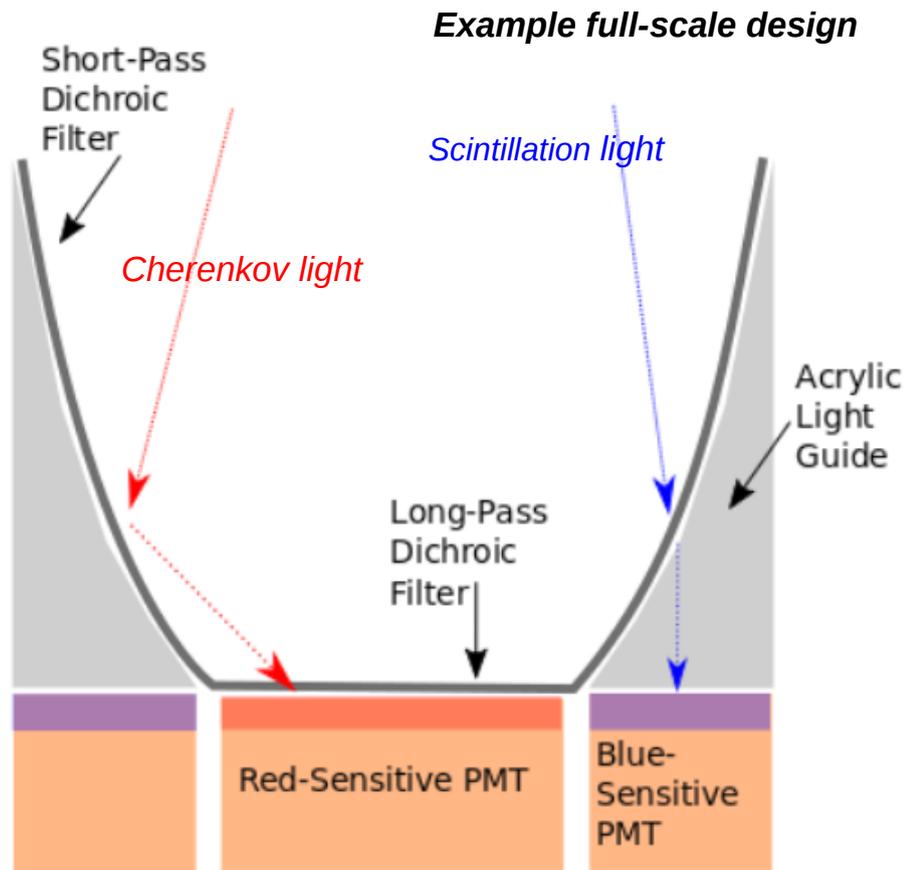
Winston Cones

And Dichroic Filters

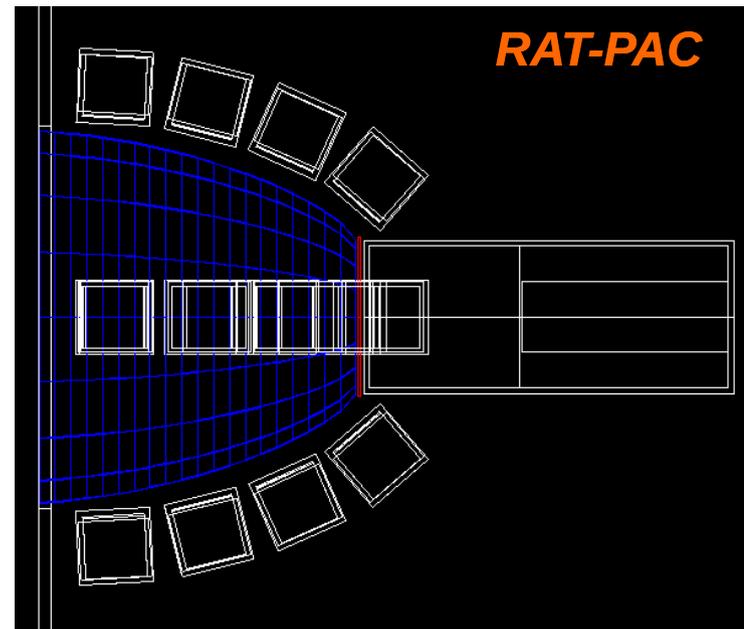
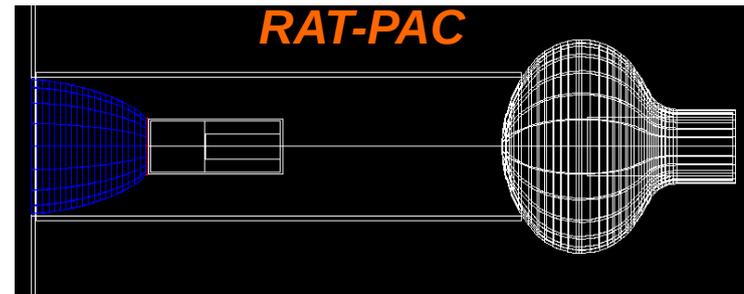


Dichroic Winston Cones: *The Dichroicon*

Sorting the scintillation and Cherenkov light towards different PMTs in order to achieve separation while maintaining a high detection efficiency for the scintillation light



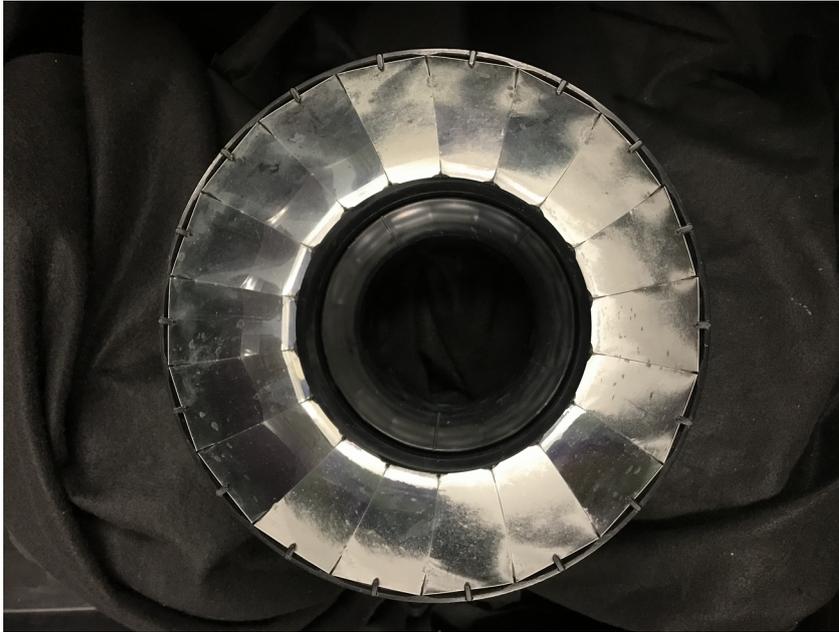
Visualizations from M. Luo



Complimentary to WbLS, slow scintillator, etc.

Familiar Technologies

SNO Winston Concentrator

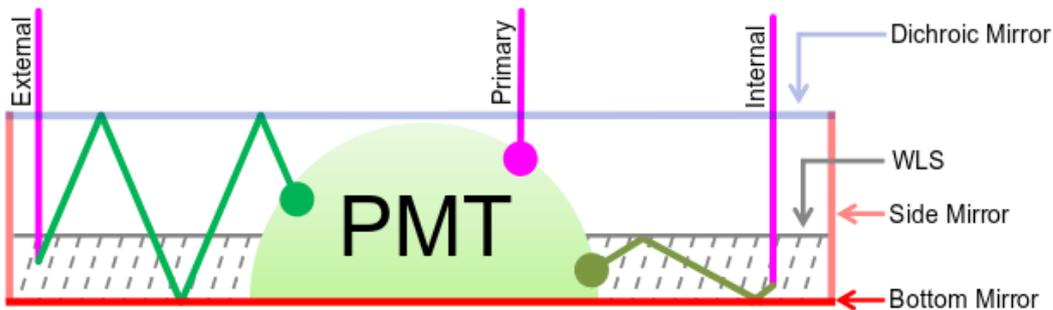


Borexino Winston Concentrator



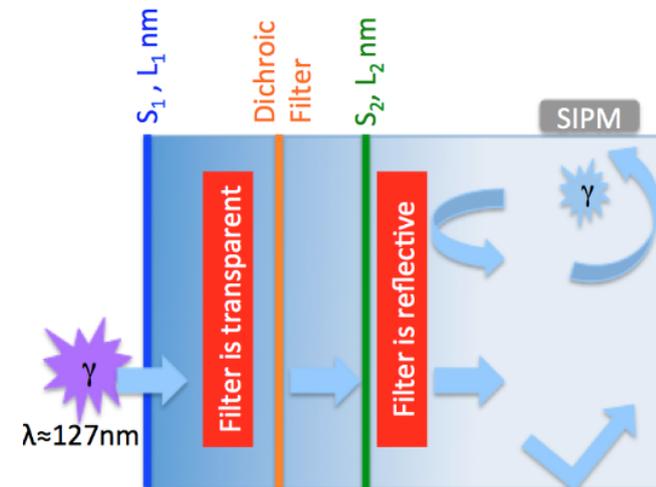
<https://arxiv.org/pdf/physics/0310076.pdf>

Photon trap using dichroic filters proposed for Hyper-K



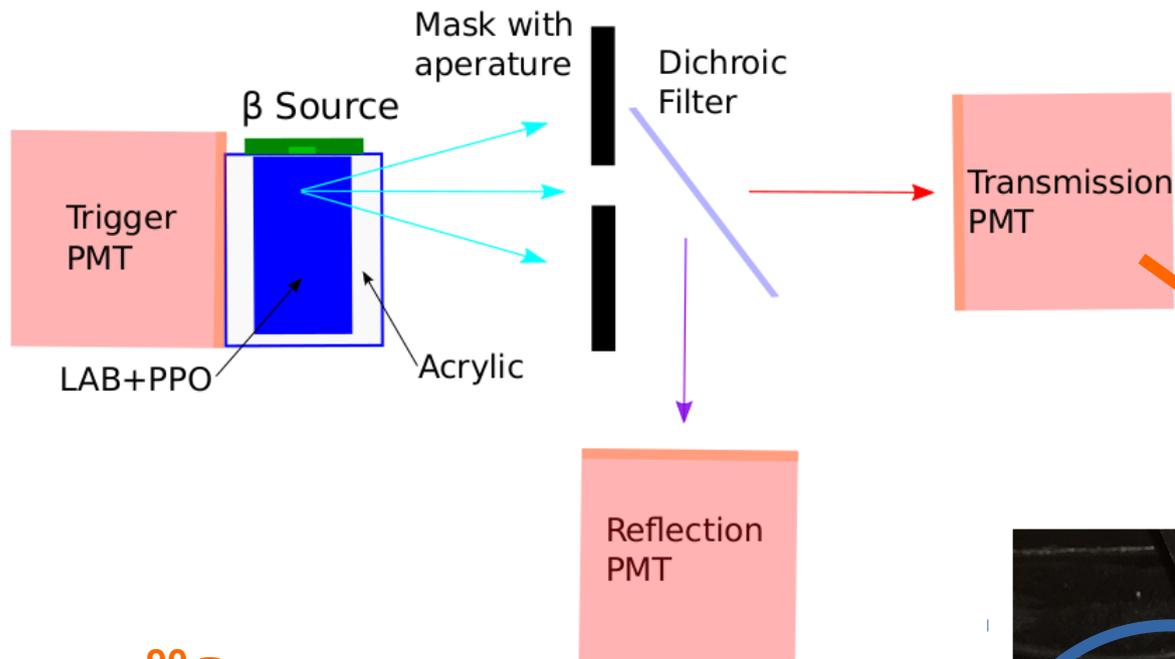
C. Rott et al. JINST 12 (2017)

ARAPUCA design for DUNE/protoDUNE



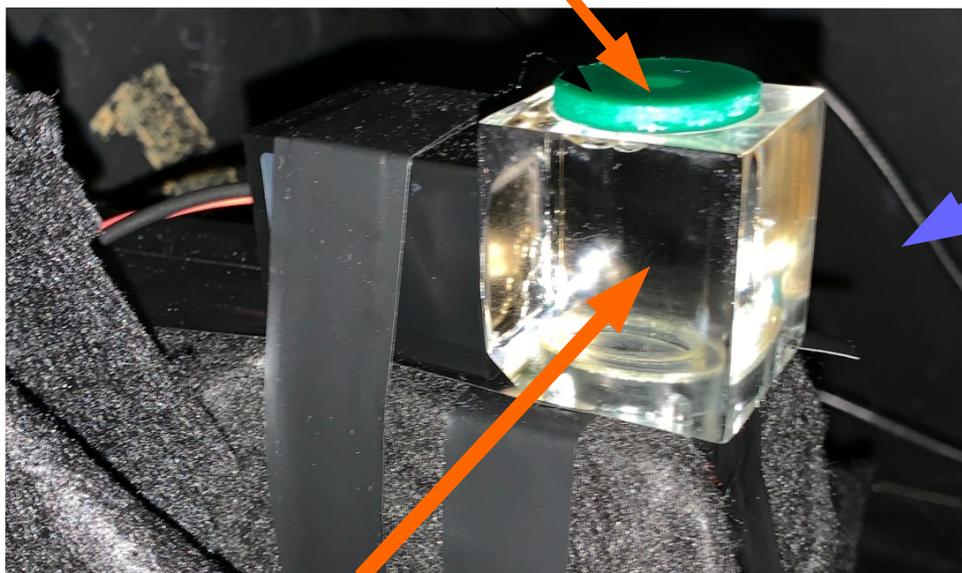
E. Segreto et al., JINST 13 (2018)

Cherenkov and Scintillation Separation With Dichroic Filters

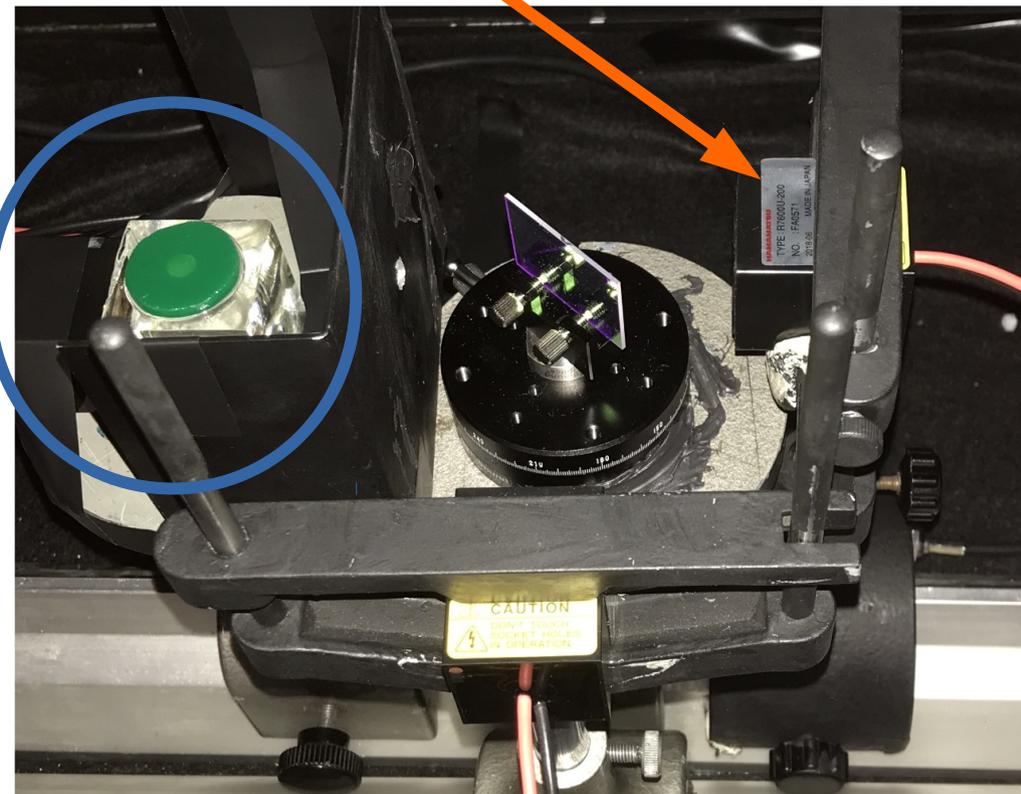


“Photon Sorting”

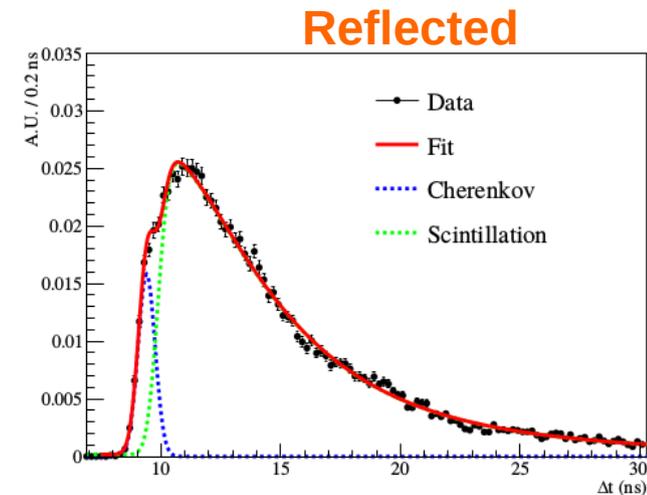
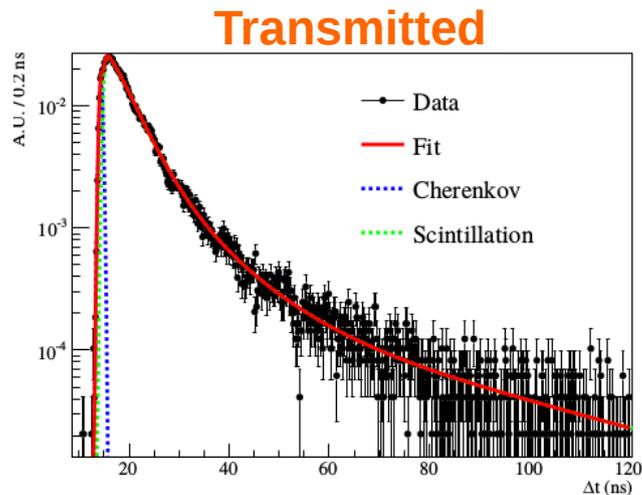
^{90}Sr source



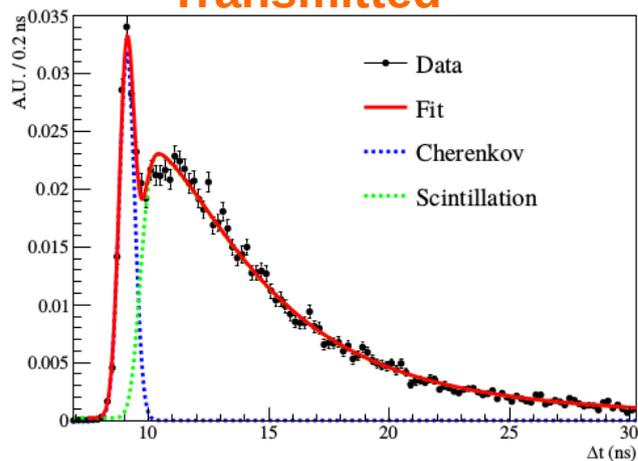
LAB+PPO inside UVT acrylic



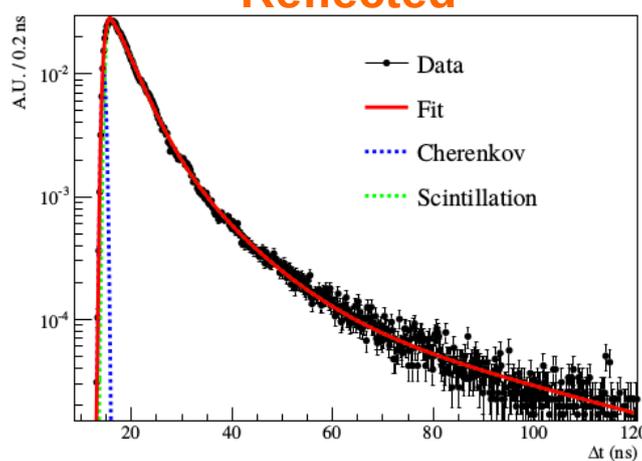
LAB+PPO, 500 nm short-pass dichroic filter



Transmitted



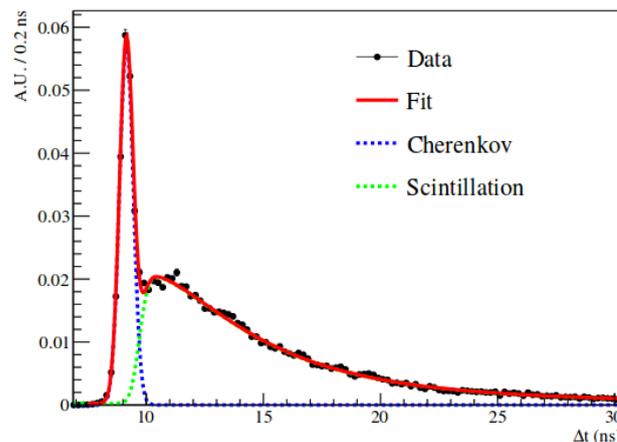
Reflected



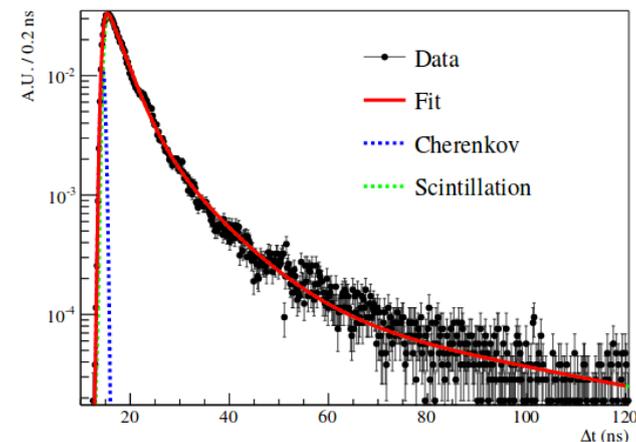
LAB+PPO, 500 nm long-pass dichroic filter

LAB+PTP, 450 nm long-pass dichroic filter

Transmitted

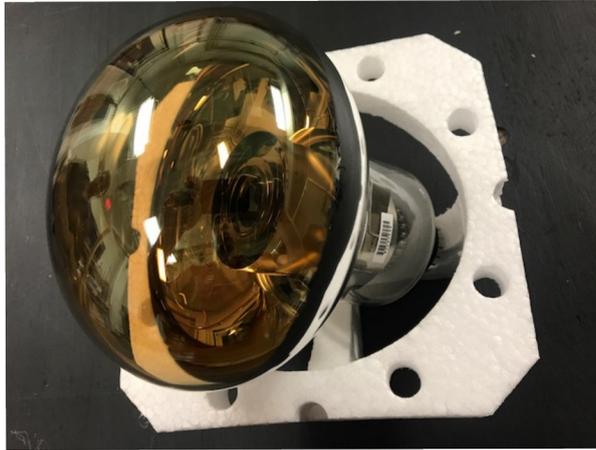


Reflected

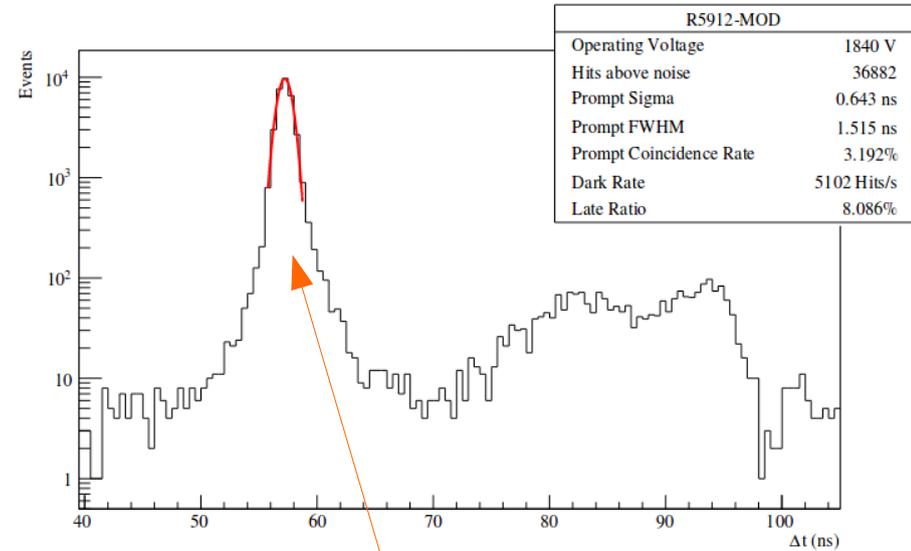
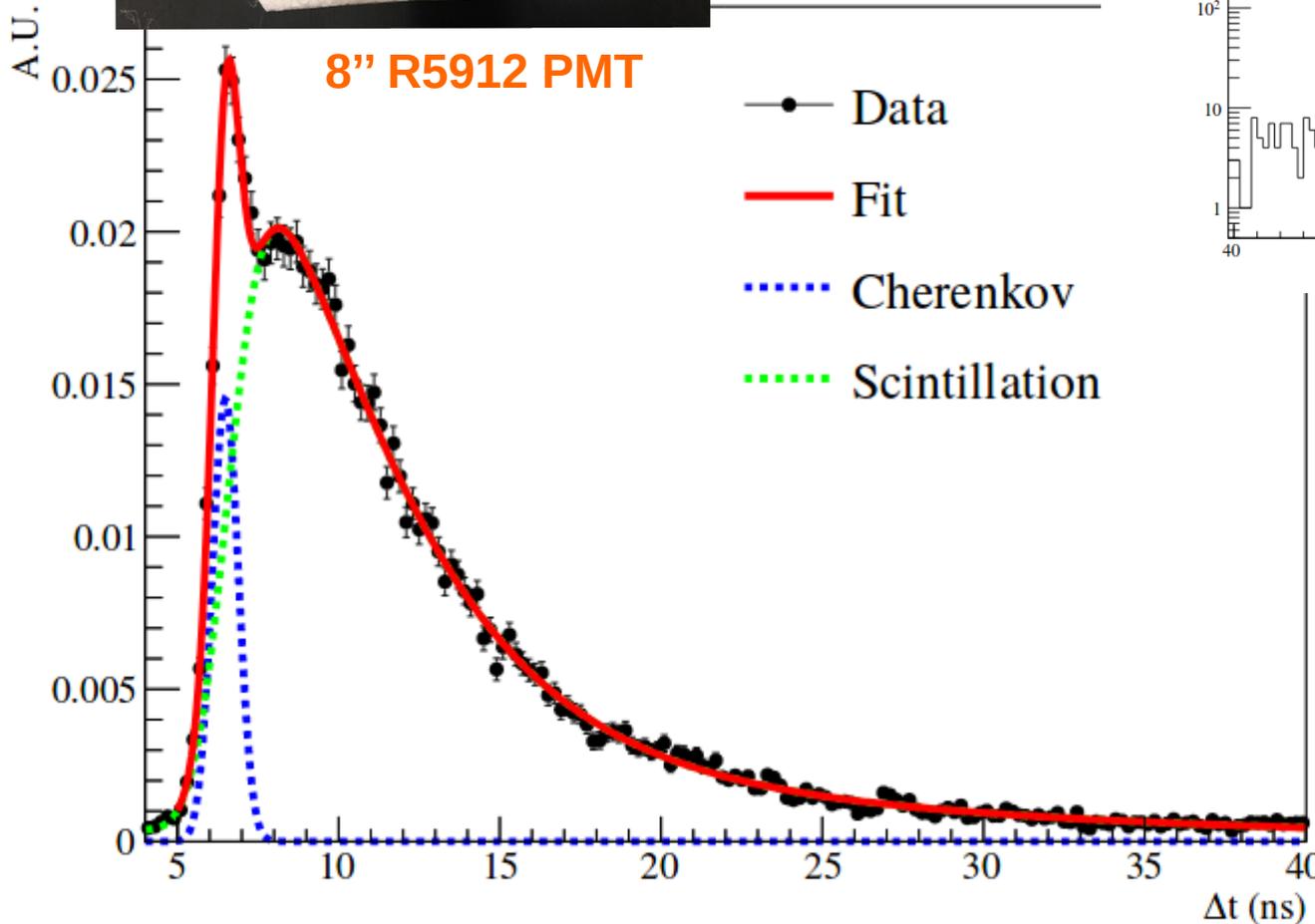


Cherenkov/Scintillation Separation With a Large-Area PMT

T. Kaptanoglu, Nucl.Instrum.Meth. A889 (2018) 69-77



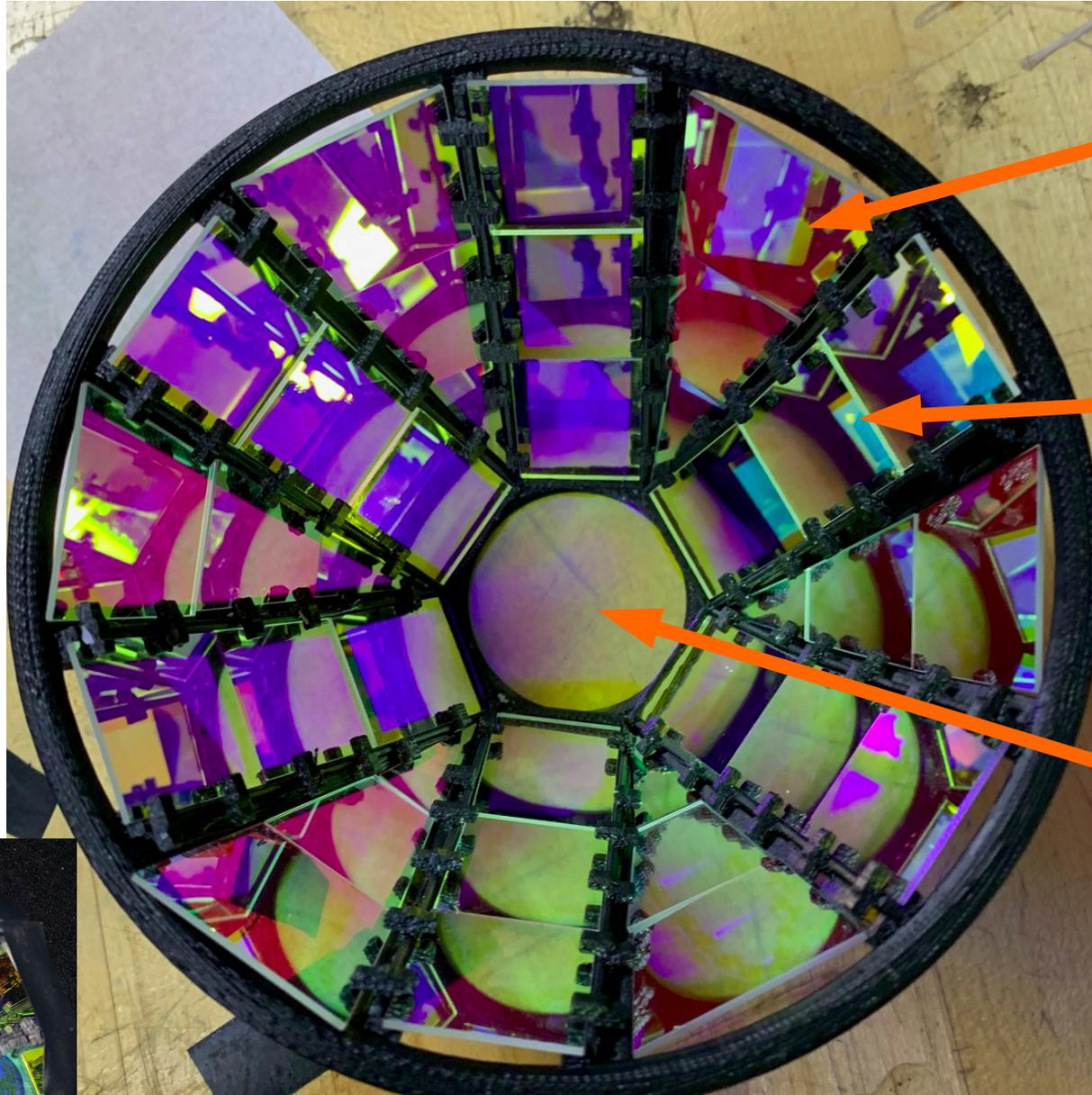
8" R5912 PMT



Transit time spread ~ 640ps

First demonstration of Cherenkov and scintillation separation using a large-area PMT!

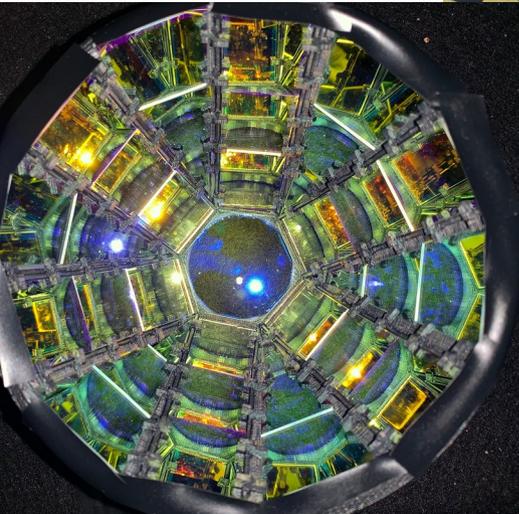
3D Printed Prototype



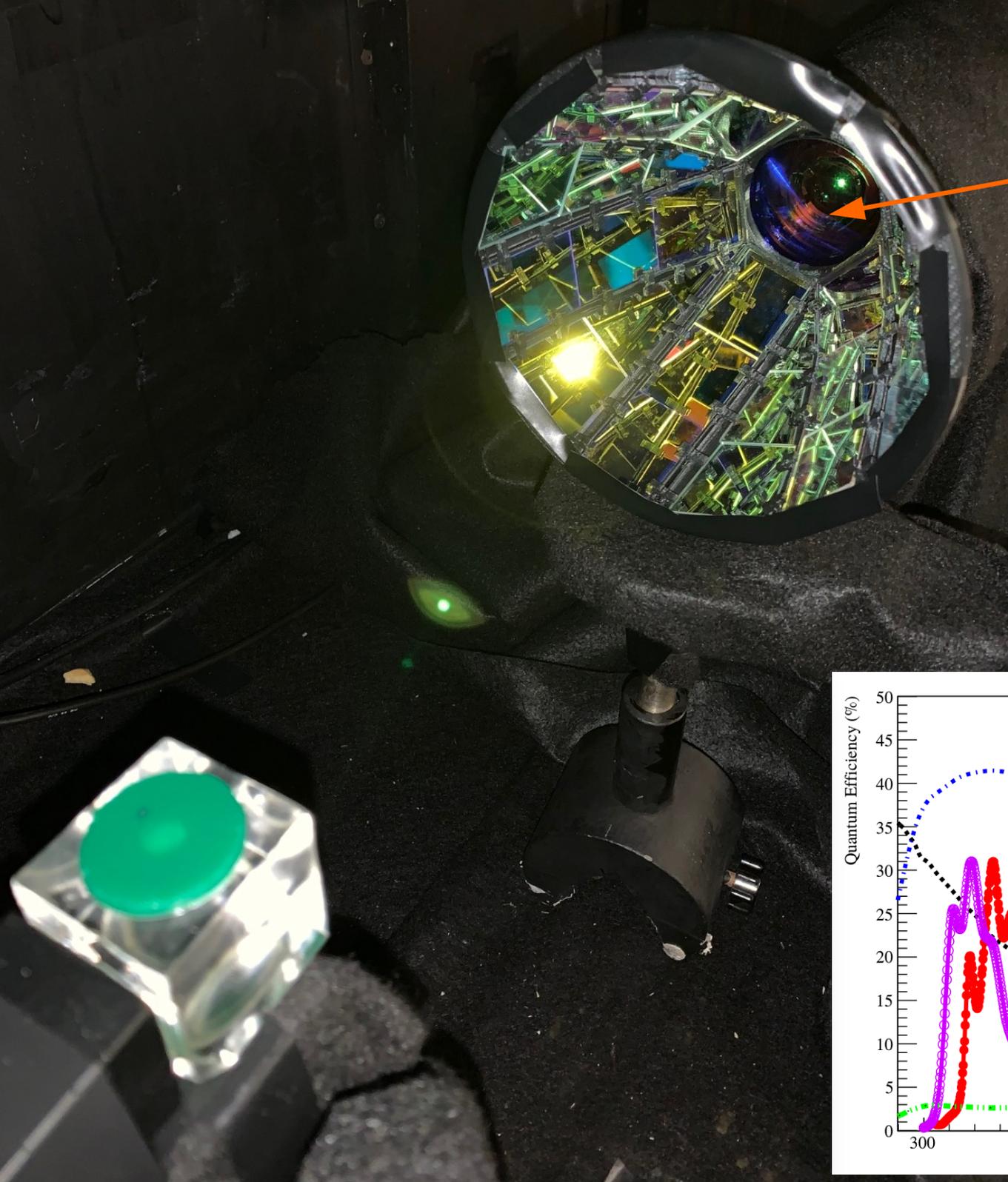
Custom cut short-pass filters from Knight Optical to fill out full 3D printed design

High performance short-pass dichroic filters from Edmund Optics

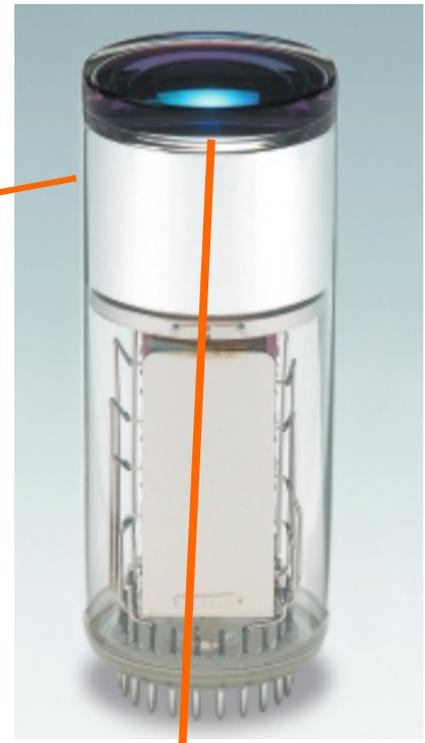
Custom cut long-pass filter from Knight Optical to fit the aperture



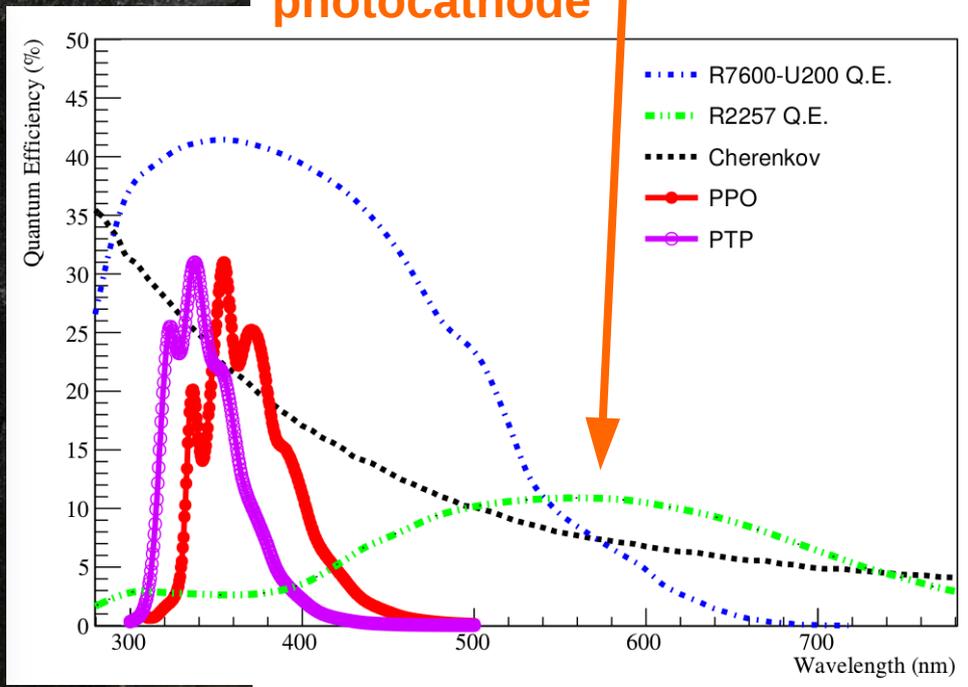
Measurements shown made with this prototype



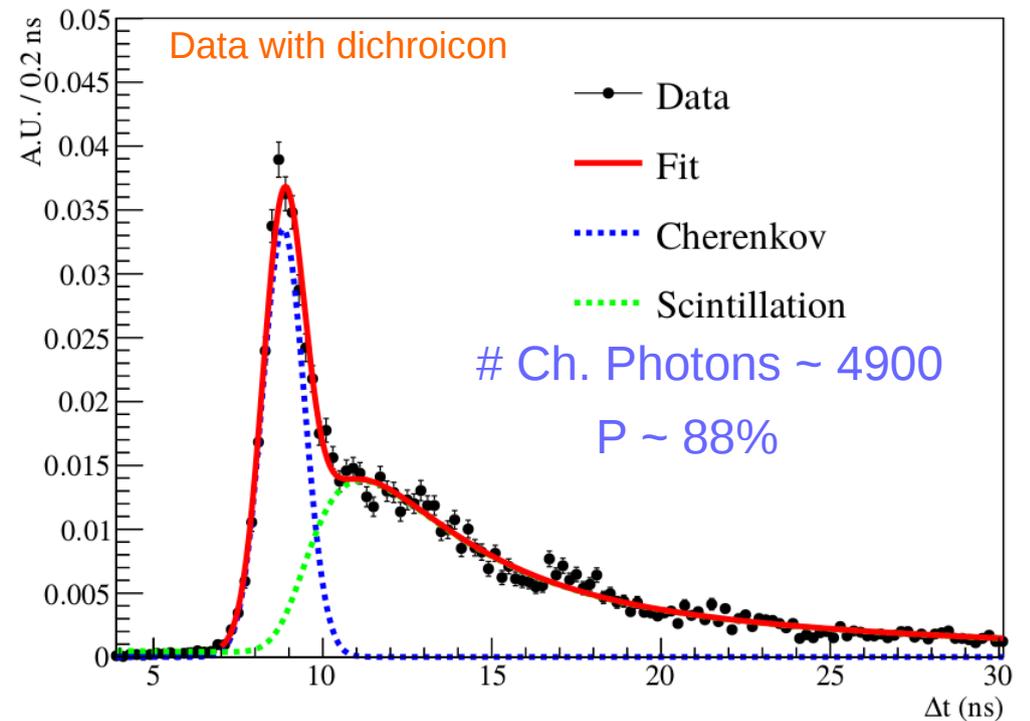
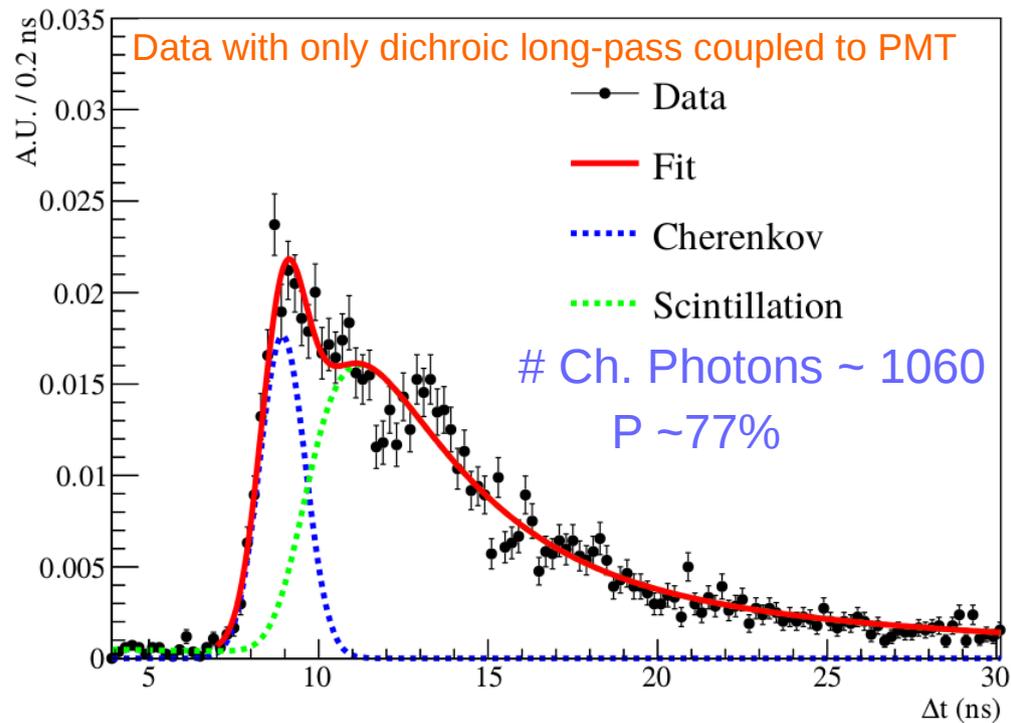
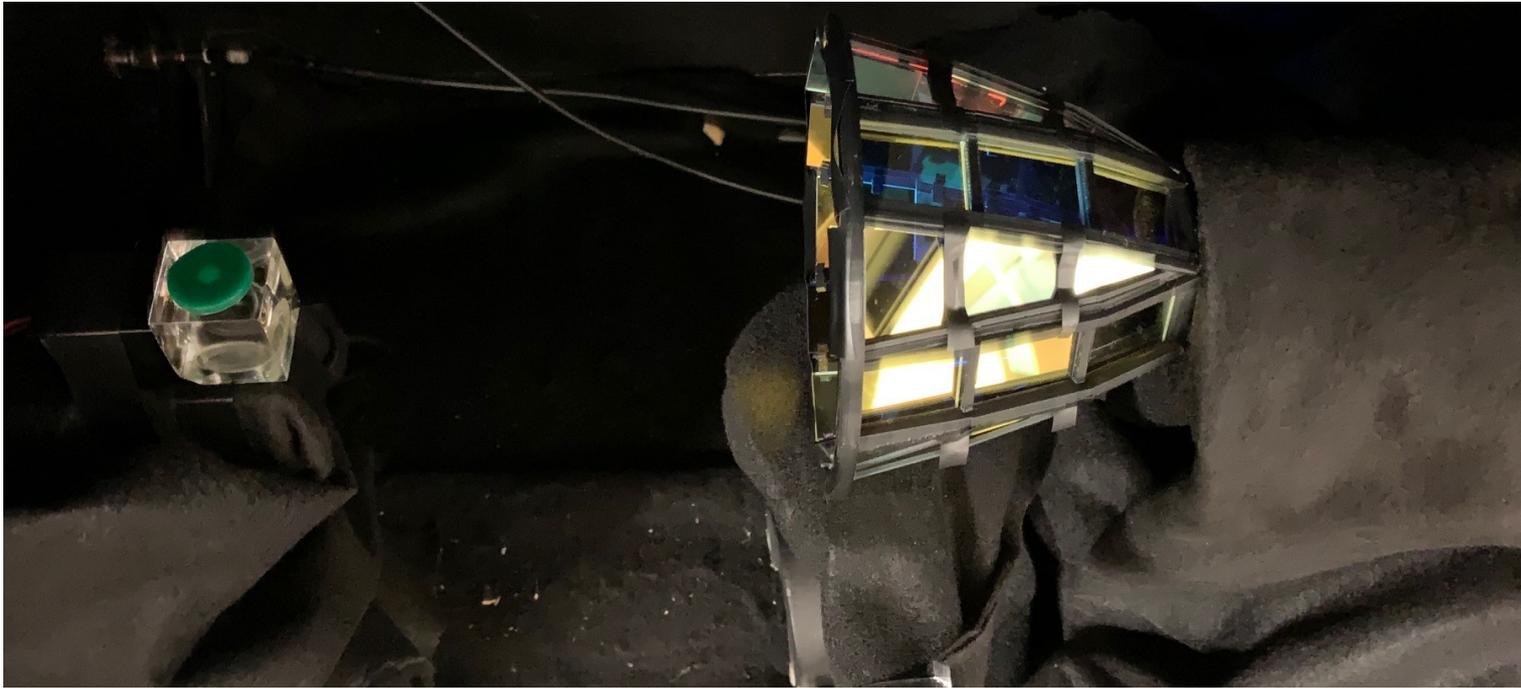
R2257



Red sensitive photocathode

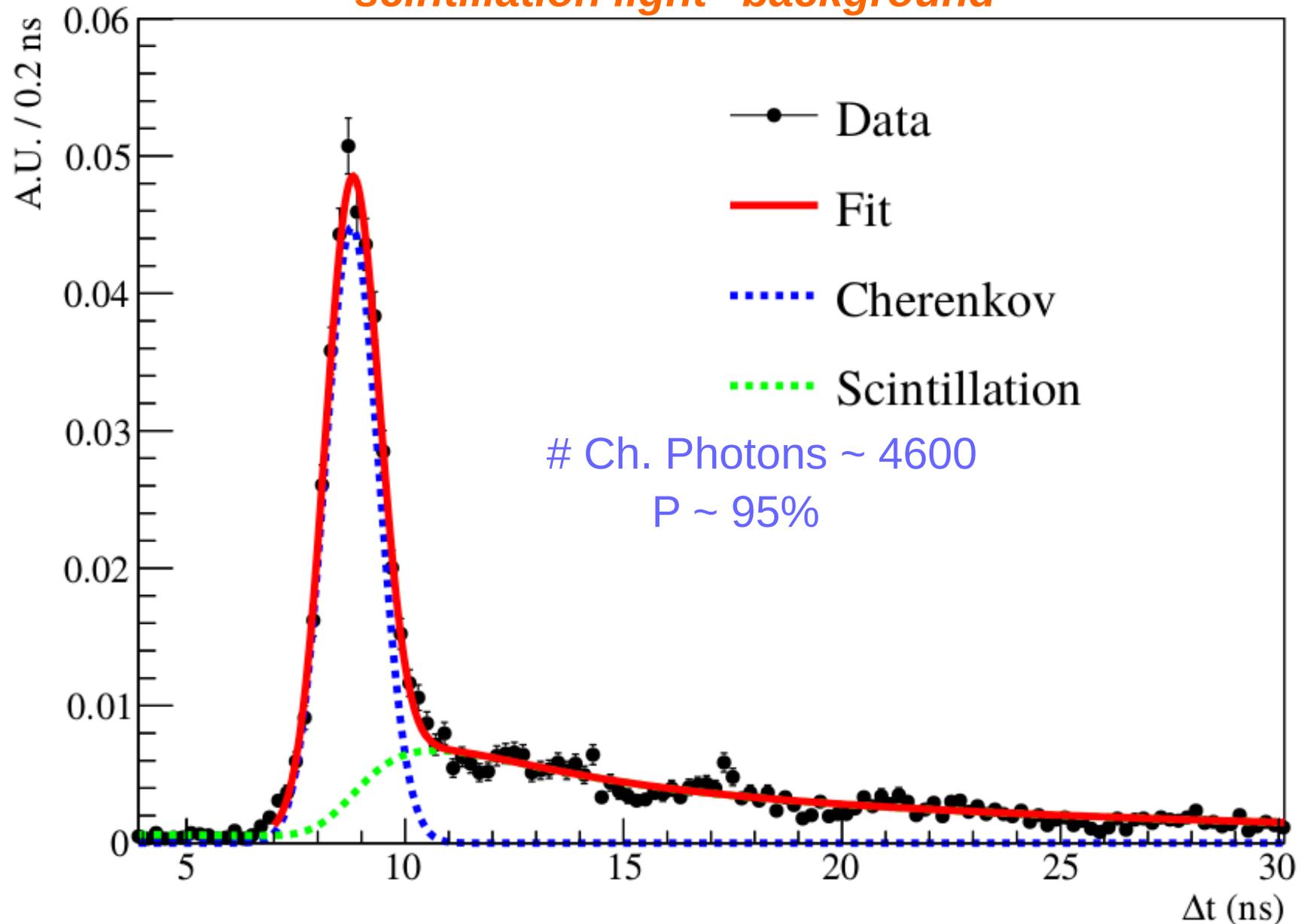


Dichroicon Data

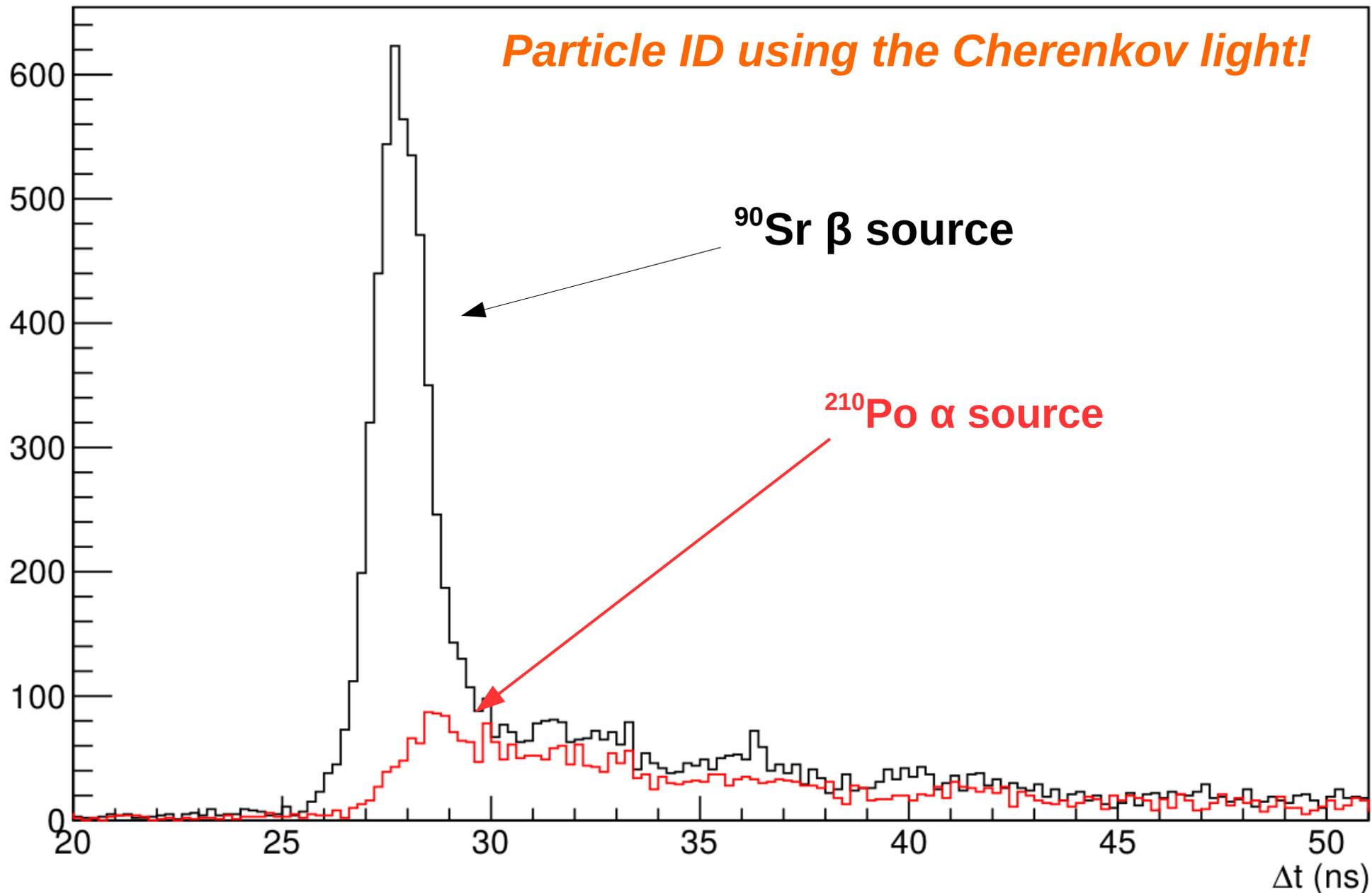


Dichroicon Data

Introducing additional long-pass filter reduces scintillation light “background”



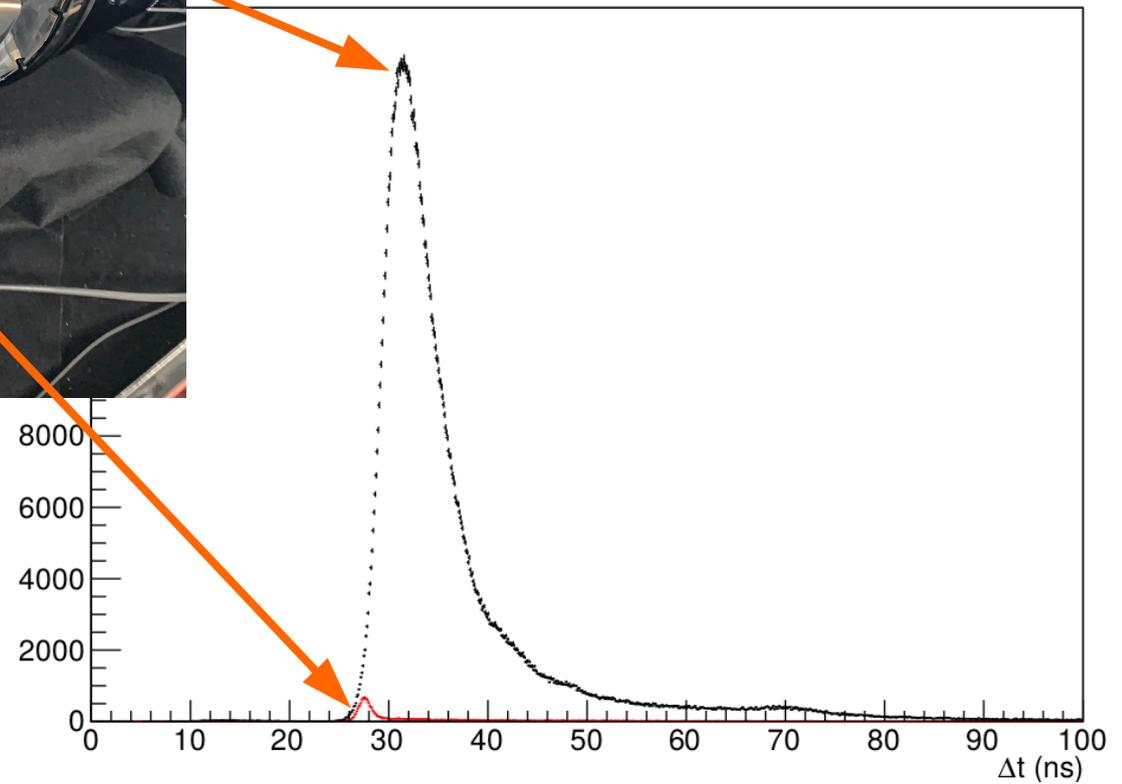
Dichroicon Data



Dichroicon Data

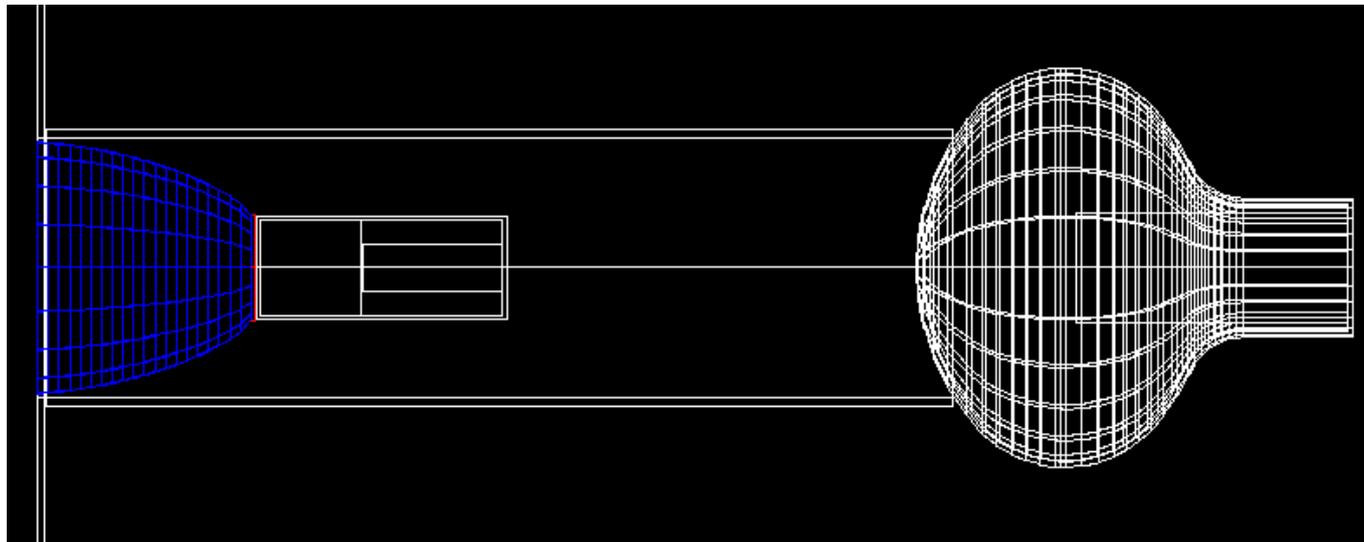
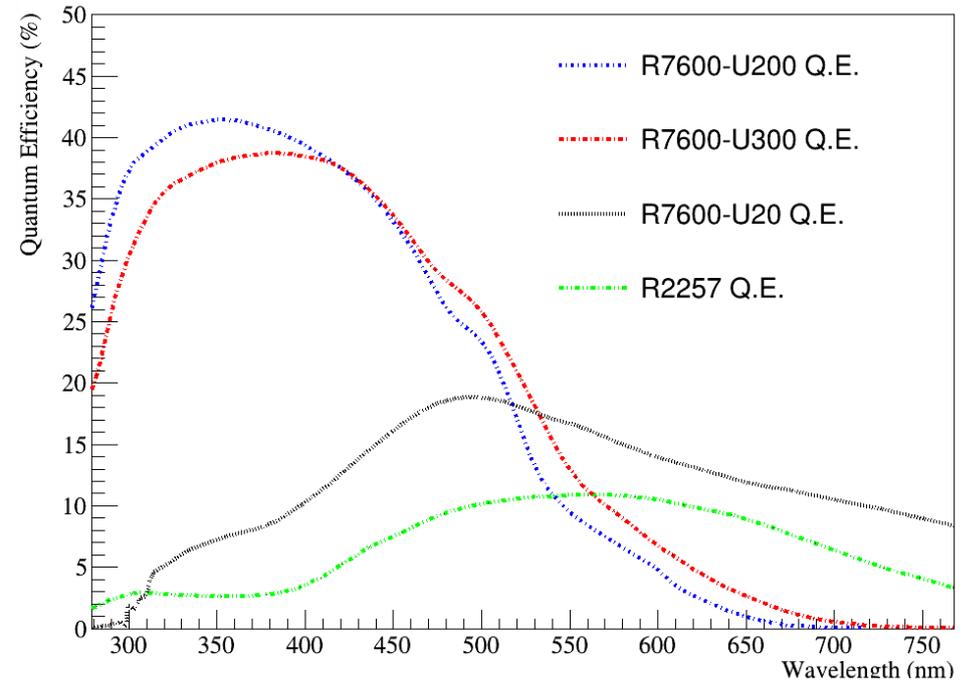


Scintillation light drowns the Cherenkov, but we've identified it by sorting the photons!



Future Measurements

- Transmission and reflections in water/oil
- Measurements with more PMTs
- Simulation studies using RAT-PAC
- Engineering of “monolithic” design
- Scintillation readout design
- Measurements with different fluors and filters



Conclusions

- Cherenkov / scintillation separation at the center of a lot of interesting R&D work
- Applications for many future experiments: THEIA, ANNIE, WATCHMAN, Jinping, JUNO
- Bench-top measurements of single dichroic filter demonstrated the potential for “photon-sorting”
- Cherenkov / scintillation separation demonstrated with first prototype dichroicon
- Lots of interesting measurements and simulations forthcoming with dichroicons

Work supported by Department of Energy Office of High Energy Physics Advanced Detector R&D

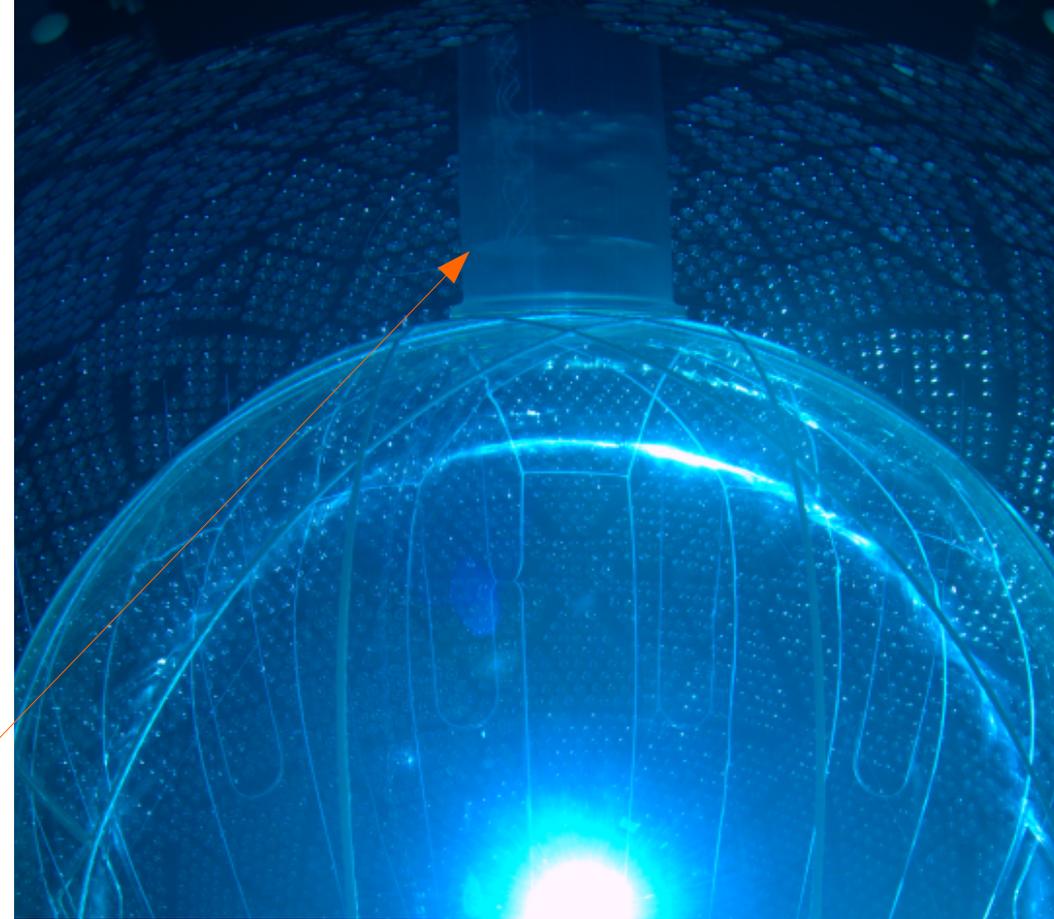
Backup

Interest in this topic has developed from our experience working on SNO+

SNO+ in a unique position of having phases as a water Cherenkov detector and liquid scintillator detector.

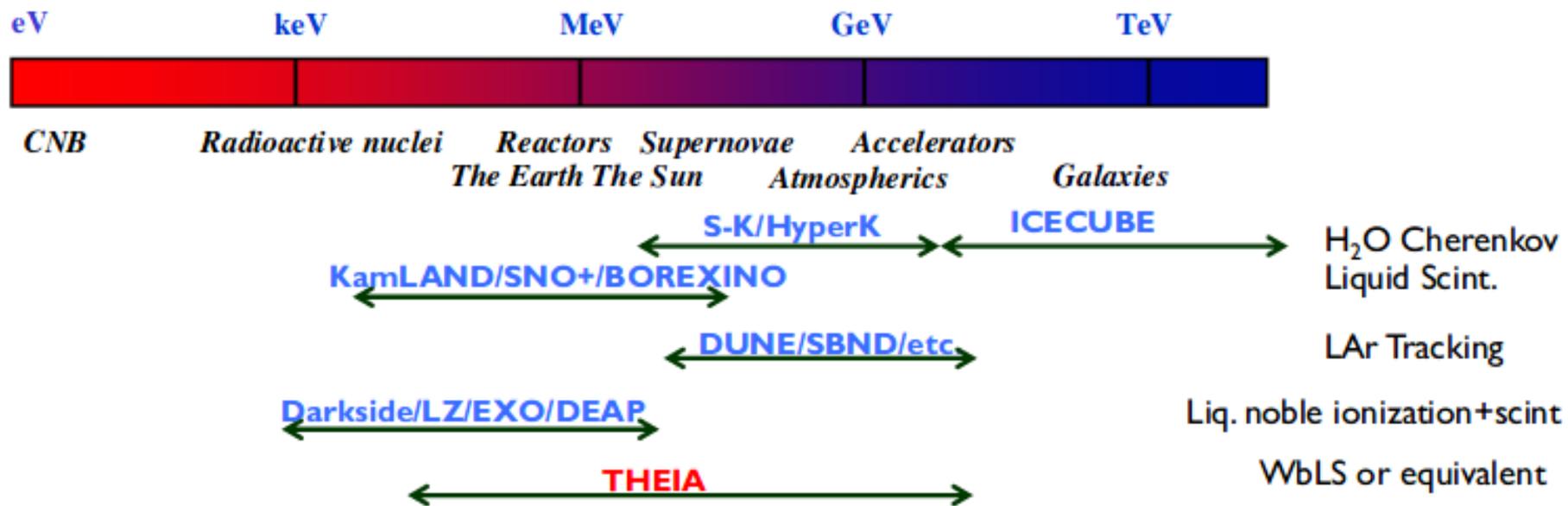
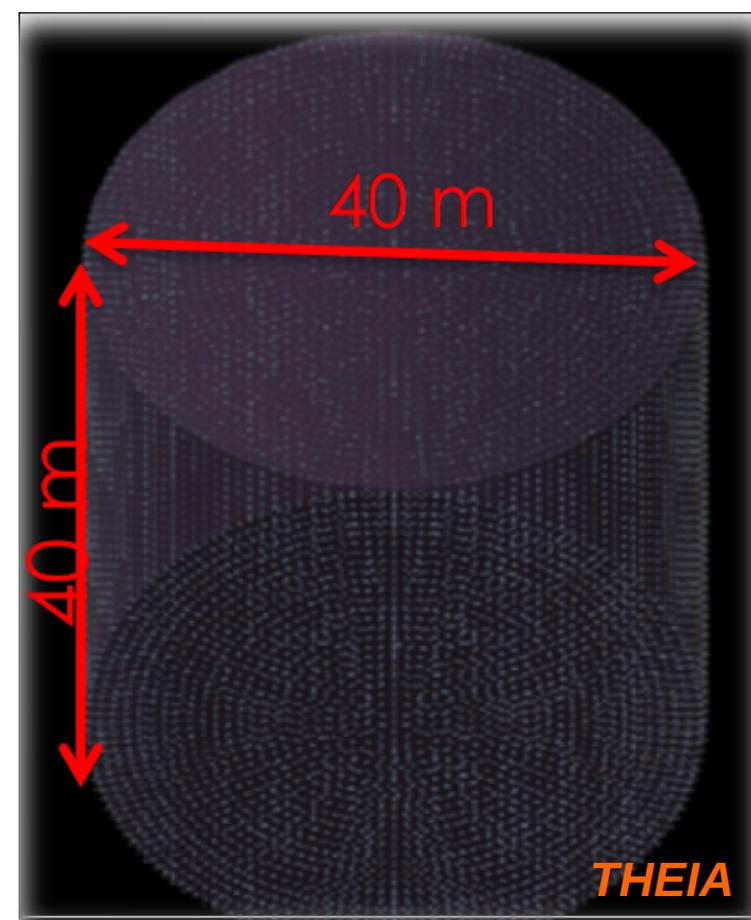
Expertise in running a water Cherenkov detector, but also lots of experience with scintillator (bench-top measurements, modeling in simulation, etc.)

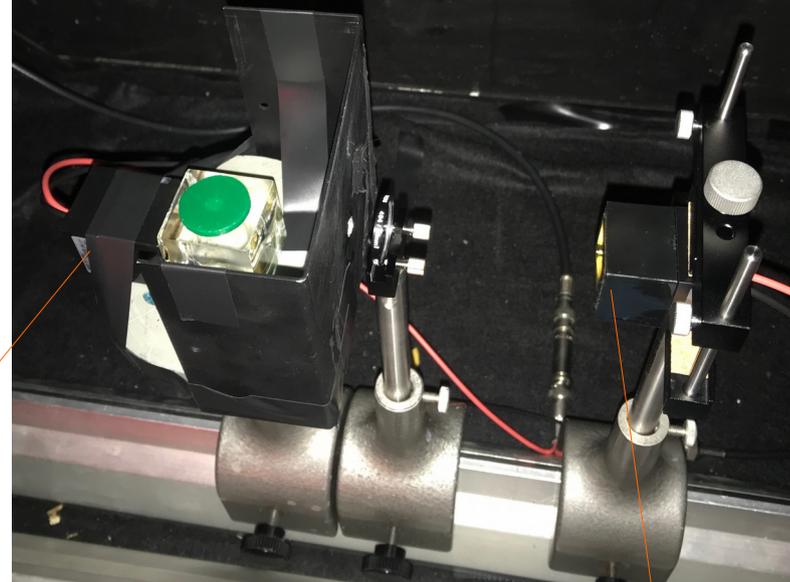
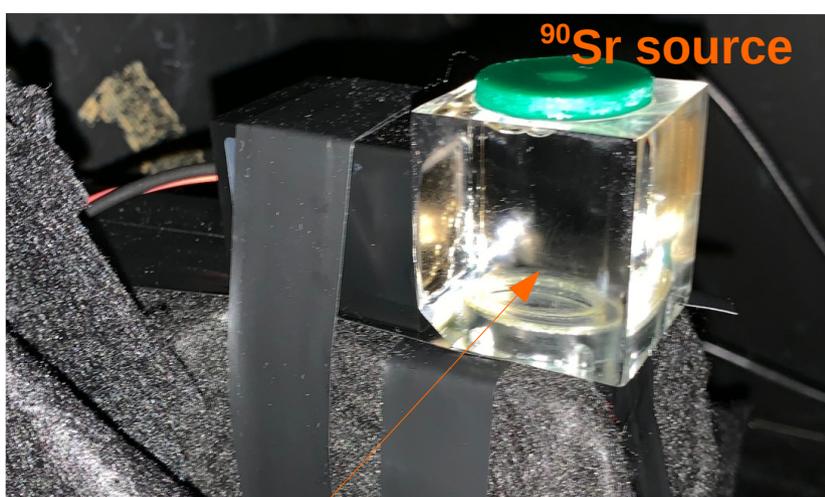
Neck fill with LAB+PPO ongoing!



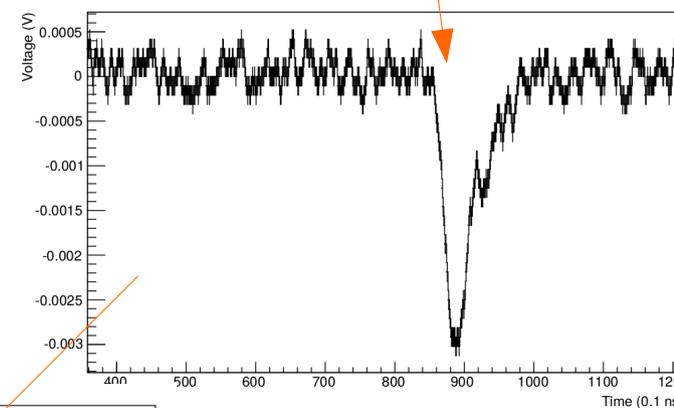
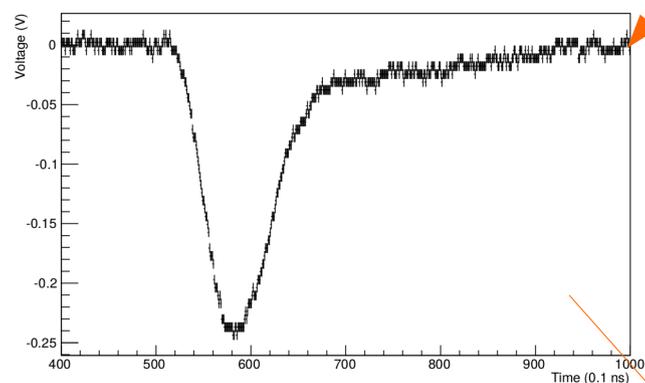
Future Experiments

- Several proposed WbLS detectors hoping to achieve Cherenkov and scintillation separation
- THEIA is a proposed 50kT WbLS (or equivalent technology) detector, potentially complimentary to DUNE
- ANNIE is 26-ton water-based detector measuring neutrino-nucleus interactions. Future phases will likely include LAPPDs and WbLS
- WATCHMAN hot-bed for future technologies – WbLS, LAPPDs, fast PMTs, *dichroicons*



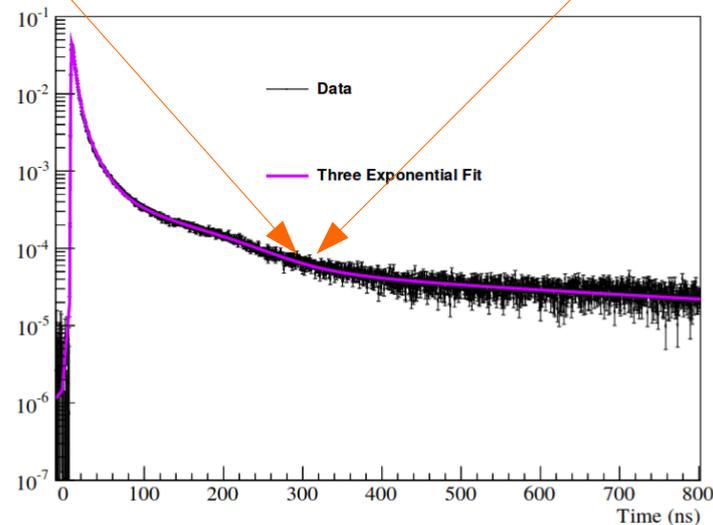


LAB+PPO inside UVT acrylic



Calculate Δt between the two waveforms

Data with no bandpass filter shows typical scintillation spectrum

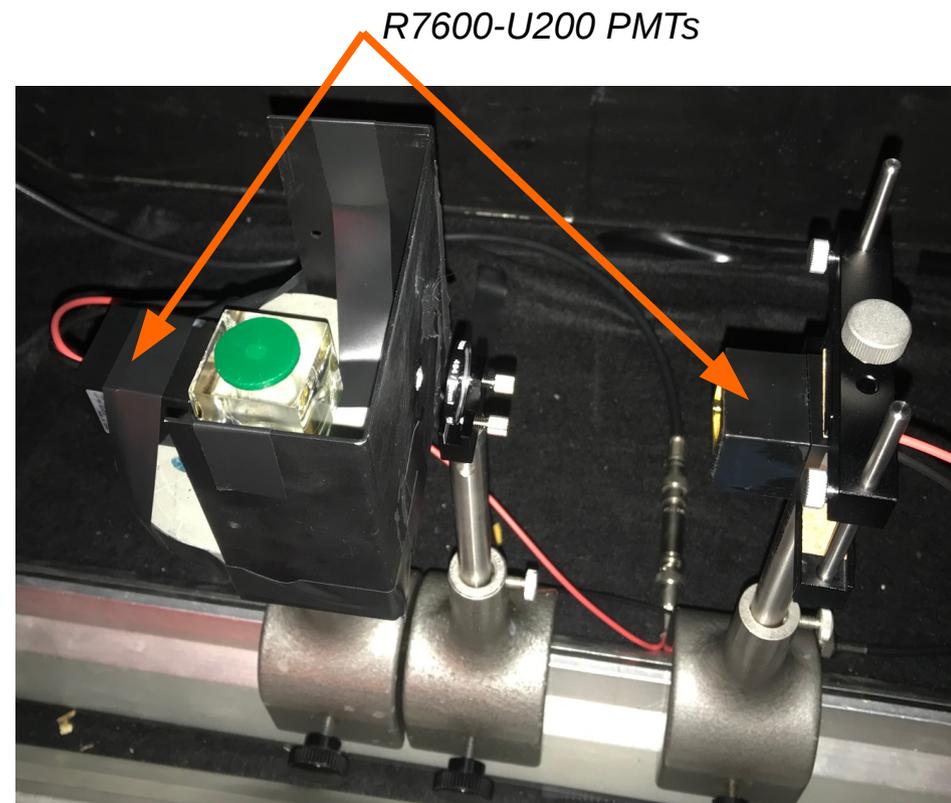
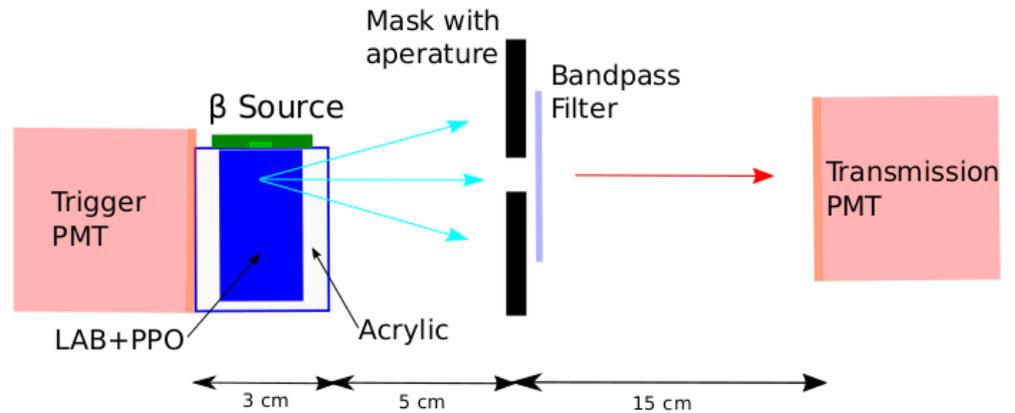


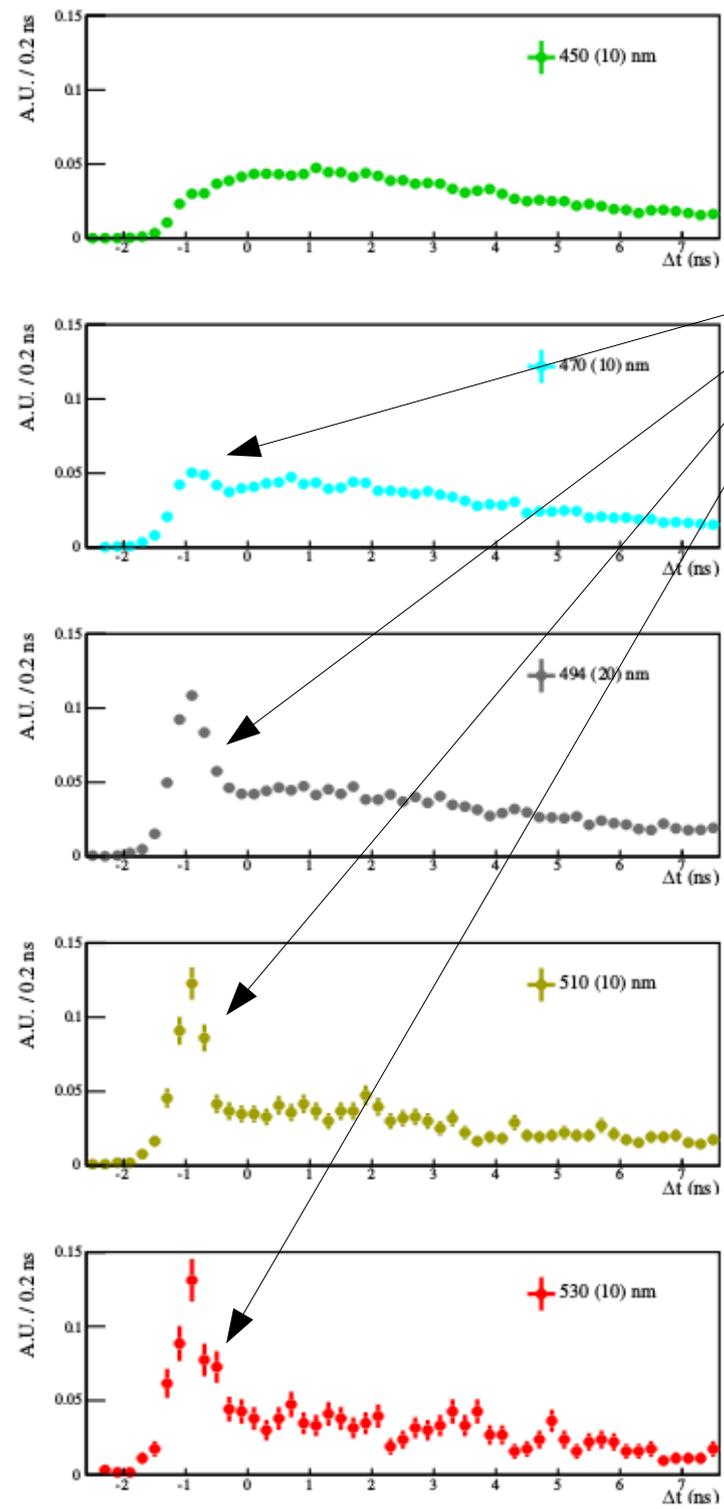
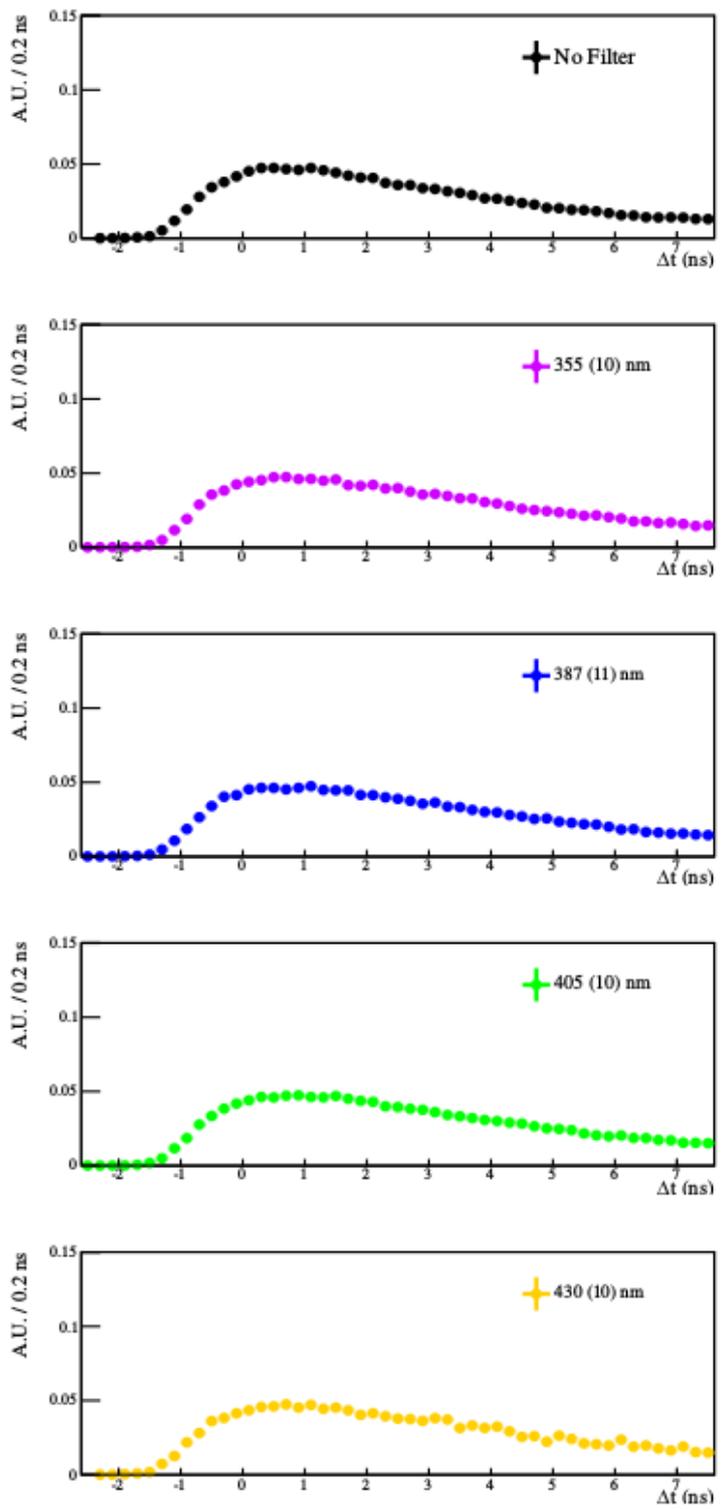
Characterized by intrinsic rise $\tau_r \sim 1$ ns followed by exponential decay with $\tau_{1,2,3} \sim 5$ ns, ~ 20 ns, ~ 400 ns

Cherenkov / Scintillation Separation With Bandpass Filters

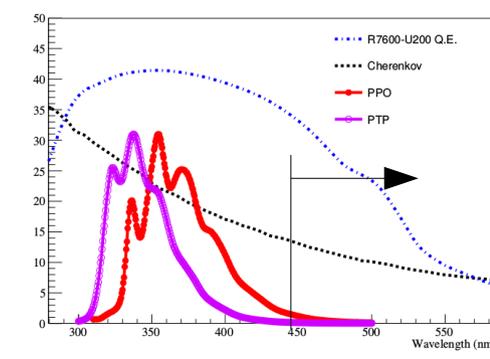
Using a set of bandpass filters to span emission spectrum of LAB+PPO

Center (nm)	FWHM (nm)	Peak Transmission (%)
355	10	95
387	11	95
405	10	96
430	10	46
450	10	98
470	10	53
494	20	95
510	10	60
530	10	54





Clear Cherenkov peak emerges at long wavelengths



Fitting the Spectrum

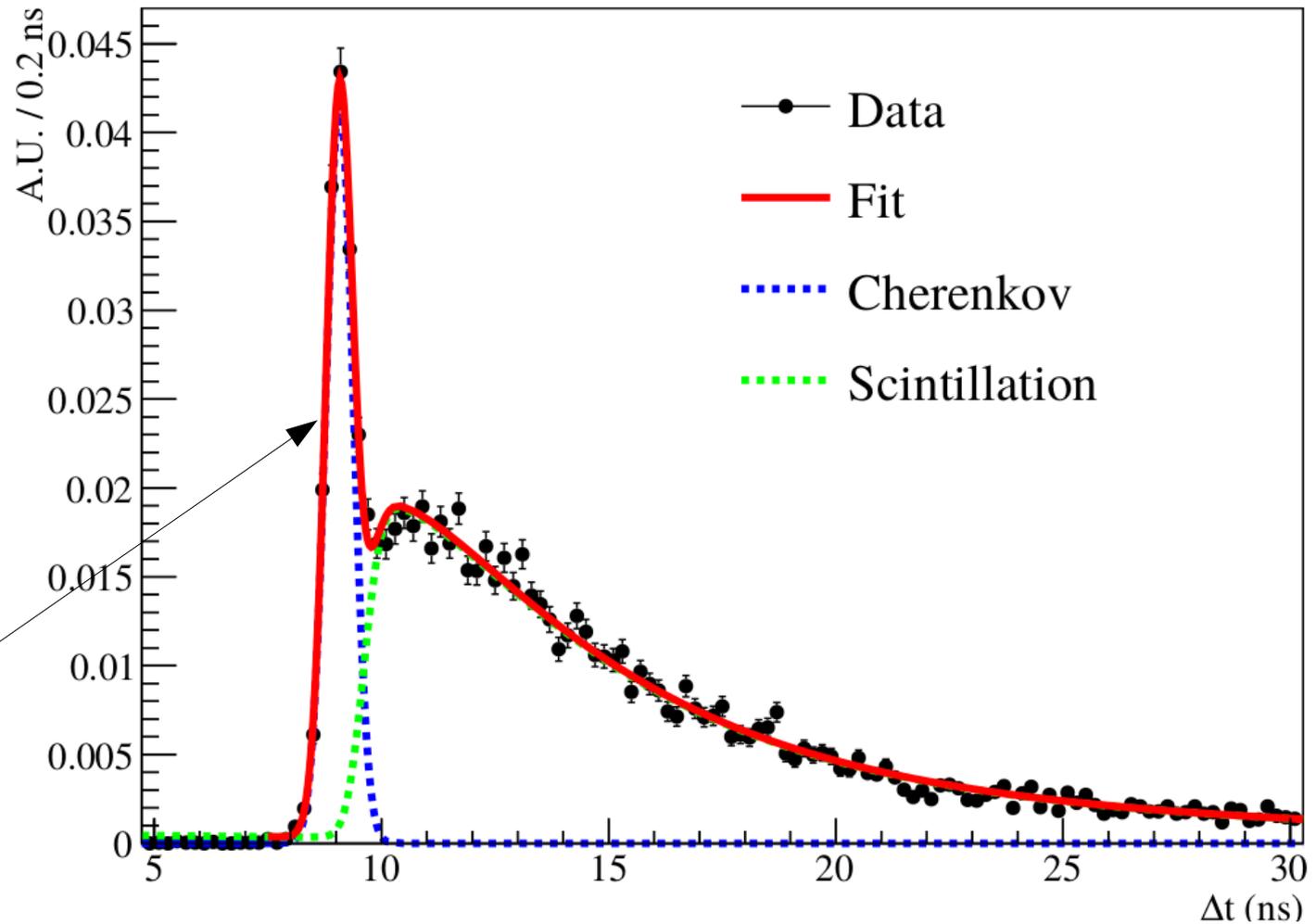
$$F = C \times f_{PMT}(t - t') + (1 - C) \times \sum_{i=1}^2 \frac{A_i \times (e^{-t/\tau_i} - e^{-t/\tau_R})}{(\tau_i - \tau_R)} * f_{PMT}(t - t')$$

$$P = \int_{8.0}^{9.5} \frac{F_C}{F} dt$$

Simultaneously fit both the Cherenkov and scintillation components of the timing profile

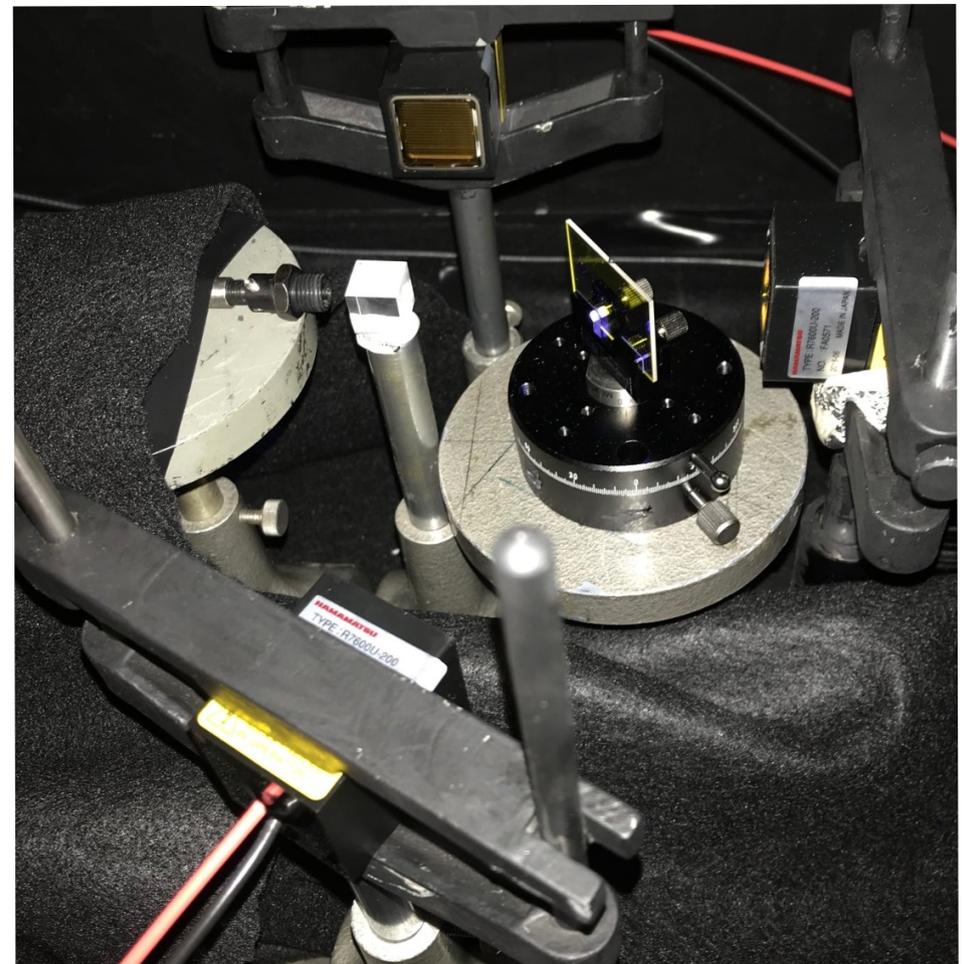
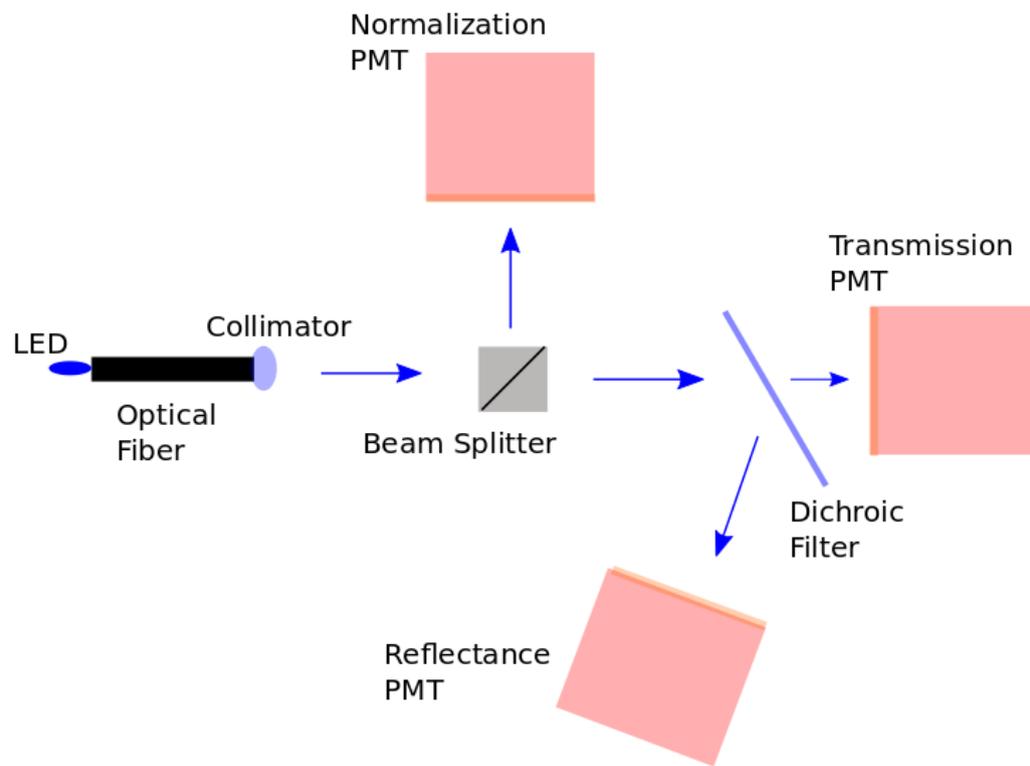
Purity, P , of the Cherenkov light in a prompt window

> 90% of prompt light is Cherenkov light!

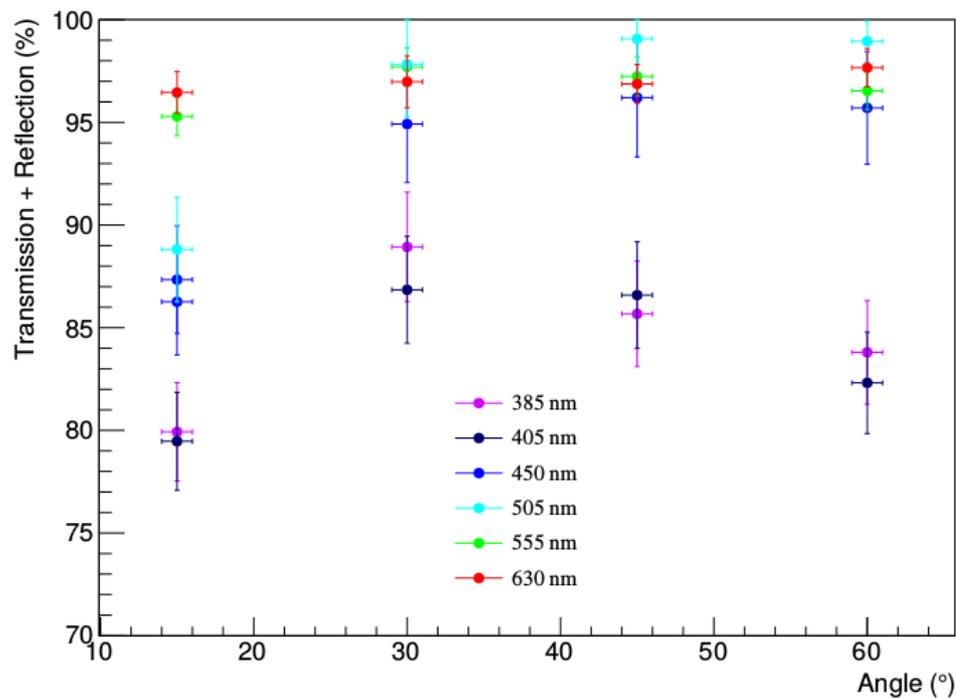
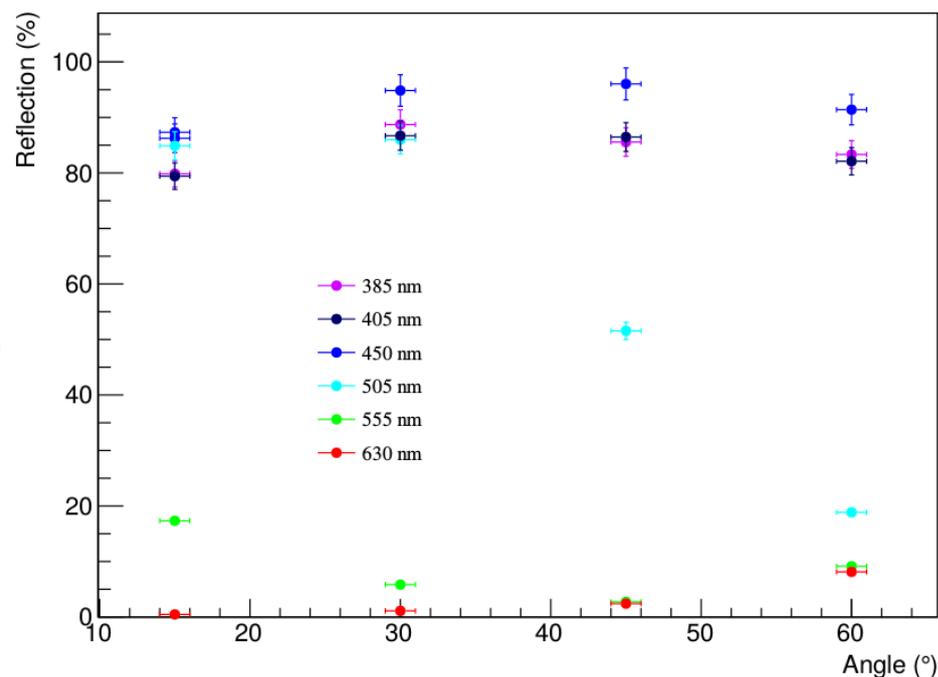
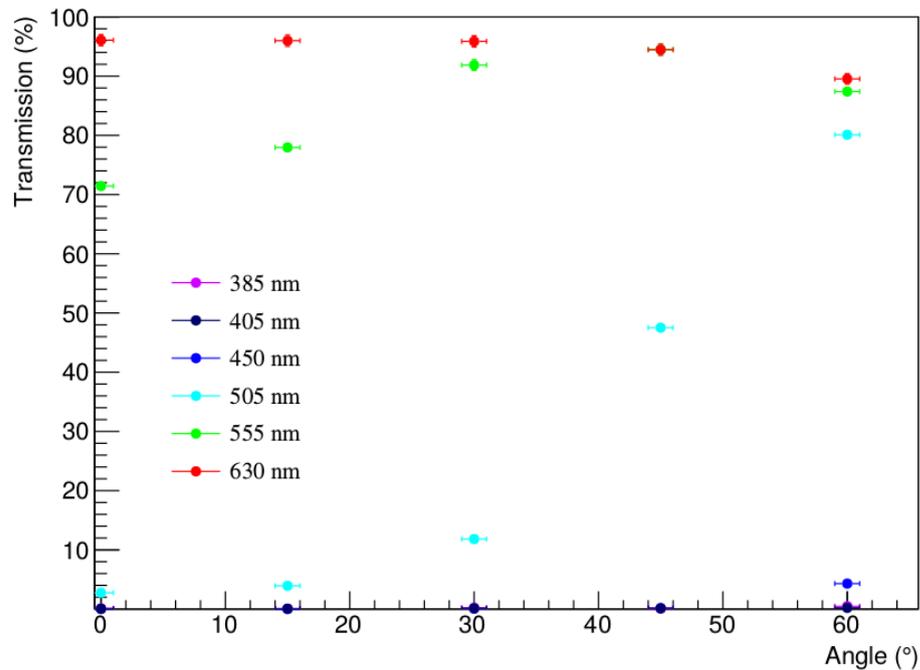


Measuring $T(\lambda, \theta)$ and $R(\lambda, \theta)$

Characterize the transmission and reflection of the dichroic filters as a function of wave and incident angle



Measurements for a 500 nm long-pass dichroic filter

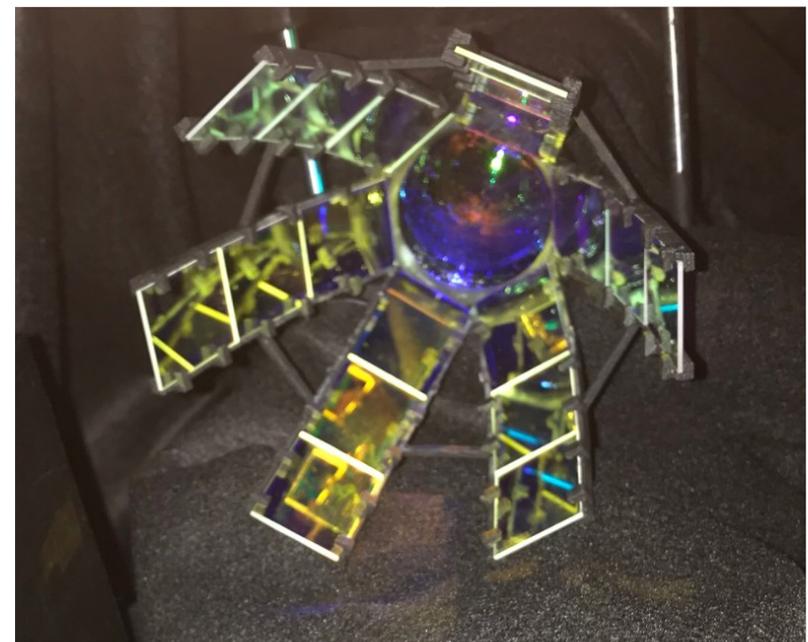
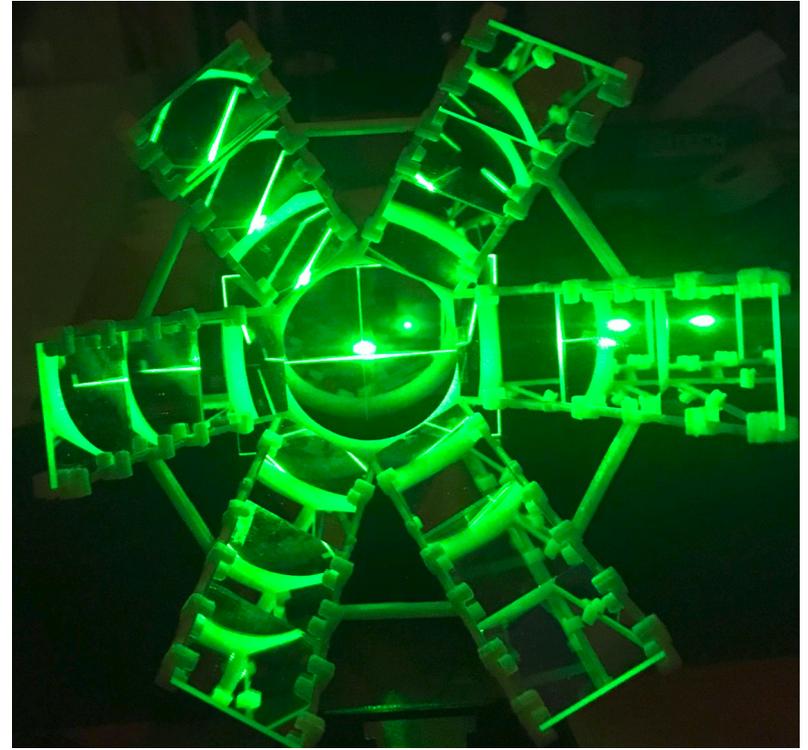


Very little light lost to the dichroic filter over range of wavelengths and incident angles

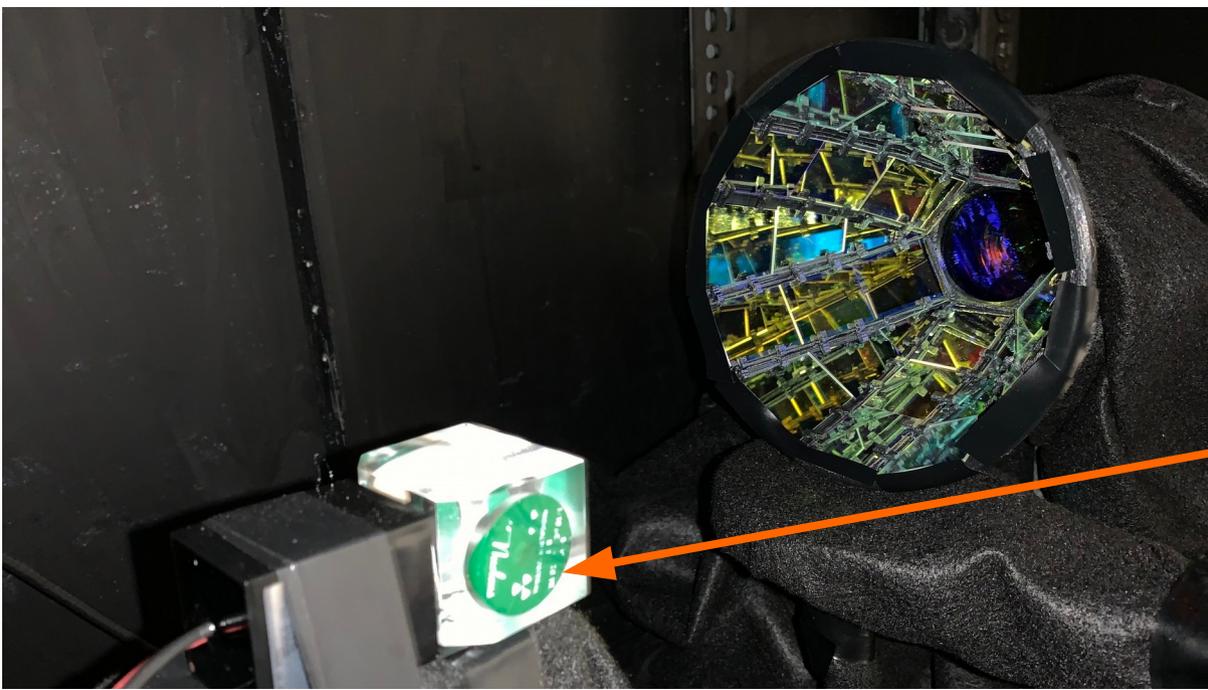
Used for input into our simulation model

Prototype 1

3D printed holder for dichroic filters

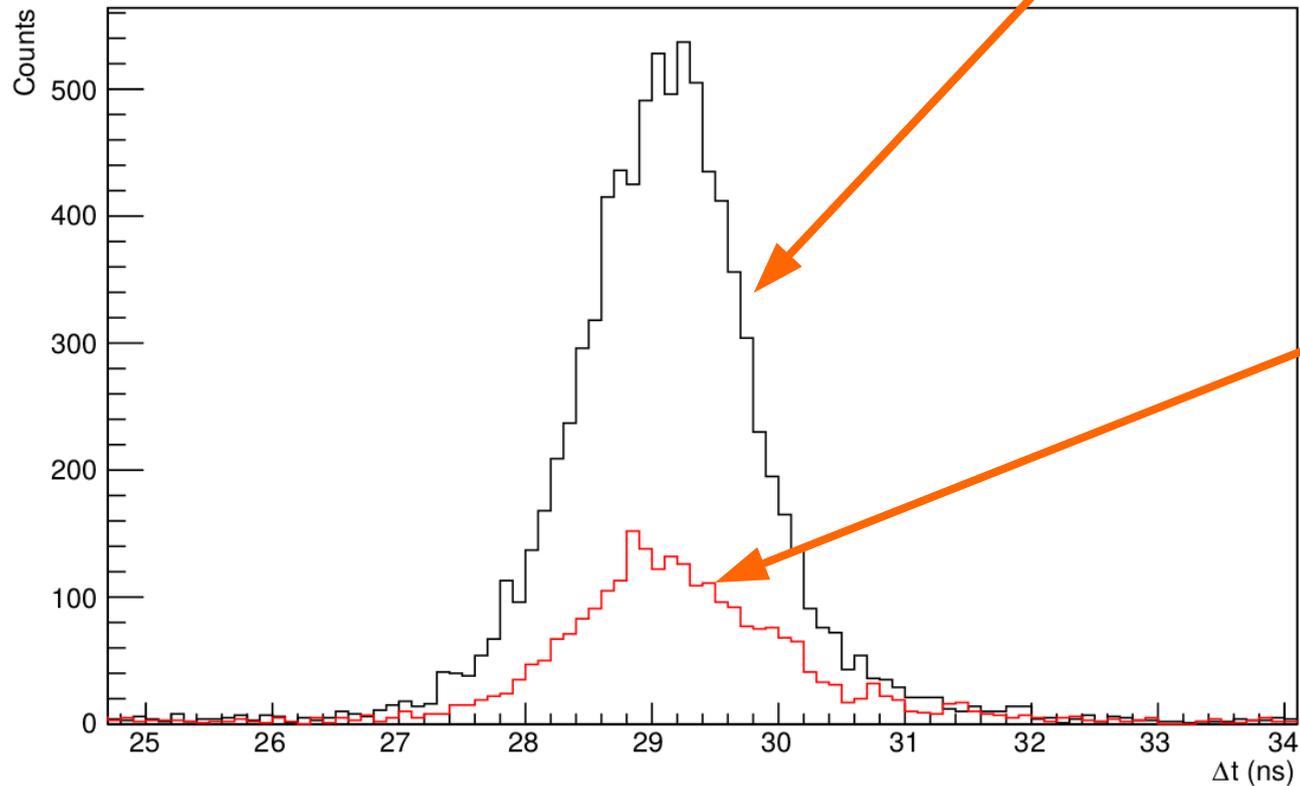


Dichroicon Data



Acrylic Cherenkov source

With dichroic Winston cone!



No Winston cone

Dichroicon Data



