



Istanbul, Iran-Turkey Conference
on LHC and related topics
June 10-15th, 2019

The long way from the discovery of W and Z to the Brout-Englert-Higgs boson..... and what next?

D. Denegri,
CEN Saclay/IRFU/SPP and CERN/Ph

key steps in establishing the Standard Model
from W, Z discovery
and the beginnings of the LHC and CMS
to the discovery of the Brout-Englert-Higgs boson
what next?



Towards electroweak unification (~1964 - 1967)

In 1965/67 S. Weinberg, A. Salam, J. Ward, S.L. Glashow... proposed a theory providing a unified description of electromagnetic and weak interactions in terms of a spontaneously broken local gauge symmetry $SU(2) \times U(1)$ - i.e. with a Brout-Englert-Higgs mechanism (proposed in 1964) to give mass to the weak field quanta; the short range of weak interactions ($< 10^{-15}\text{cm}$) implied large masses for the W^+ and W^- force carriers (intermediate vector bosons) as $R_{\text{force}} \sim 1/M_{\text{field quantum}}$

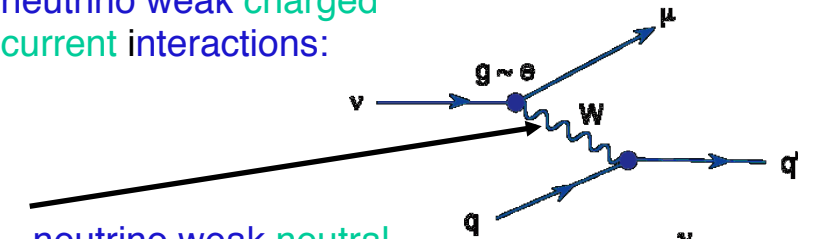
- This unified scheme however implied the existence of four field quanta:

γ - which remains massless – for electromagnetism
 W^+ and W^- - for charged weak interactions (“weak charged currents”),

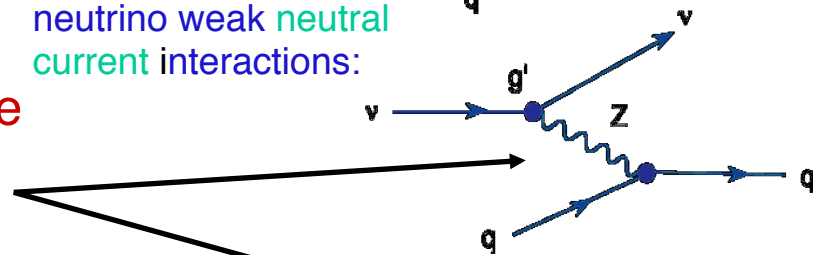
Z^0 - a sort of “heavy photon” - implying the existence of “neutral current” weak interactions i.e. of a new type of weak interactions, unknown at that time!

- The theory had a single parameter, the weak mixing angle θ_w , the W, Z and their masses and couplings to fermions depending on this angle

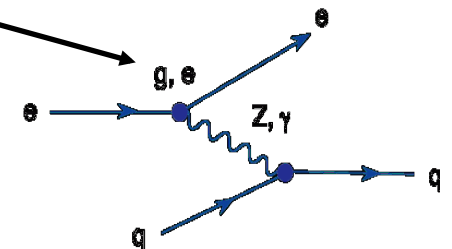
neutrino weak charged current interactions:



neutrino weak neutral current interactions:




electron weak neutral current interactions





Towards the Standard Model (period ~1970 - 1979)

- In 1971 G.'t Hooft, M. Veltman, B. Lee and J. Zinn-Justin show that electroweak theory is renormalizable
 at this point this theory began to be taken seriously by theorists!
- In 1973 discovery of neutral-current neutrino interactions at CERN (Gargamelle): first experimental evidence in favour of the unified electro-weak theory!
- In 1970 the GIM mechanism was invented, the charm quark proposed and found in 1974 as J/ψ - the most “tangible” evidence for existence of quarks, then in 1977 - the Y system with b-quarks, a still more spectacular quark-antiquark system;
- In 1973 QCD was introduced with SU(3) as the gauge symmetry - for the three colors of quarks (Nobels of 2004, Gross, Wilczek, Politzer), with “confinement” and its prediction of “asymptotic freedom” - our present day SM was falling in place!
- parity violation in electron-nucleon scattering as expected from electro-weak theory was found at SLAC in 1978 and angular asymmetry in e^+e^- to $\mu^+\mu^-$ as expected at DESY in 1979
- in 1979 Weinberg, Salam and Glashow got the Nobel Prize for the electro-weak unification and the prediction of the existence of the Z



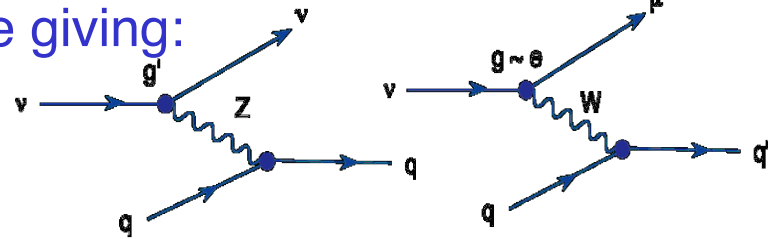
.....things were getting serious!



W, Z, H how it all really started....I

- First measurements (in 70's) of charged and neutral current neutrino interactions interpreted in this unified electroweak scheme were giving:

$$\sin^2\theta_w \sim 0.3 - 0.5$$



with:



$$m_W = [\pi\alpha_{em}/(\sqrt{2}G_F)]^{1/2}/\sin\theta_w = 37.4 \text{ GeV}/\sin\theta_w$$

$$m_Z = m_W/\cos\theta_w$$

implying that:

$$m_{W,Z} \sim 50 - 100 \text{ GeV}$$

but existing machines, for ex. ISR or the CERN SPS in a fixed target mode, could not give more than $\sqrt{s} \sim 30 - 40 \text{ GeV} !!$ and the LEP was still far in the future, at least ten years away.....(ISABELLE.....)



here came the suggestion of Cline, McIntyre, Rubbia to convert an existing proton -synchrotron into an antiproton-proton collider!

$$\text{Event rate} = L \sigma [\text{s}^{-1}] \quad (L \equiv \text{luminosity})$$

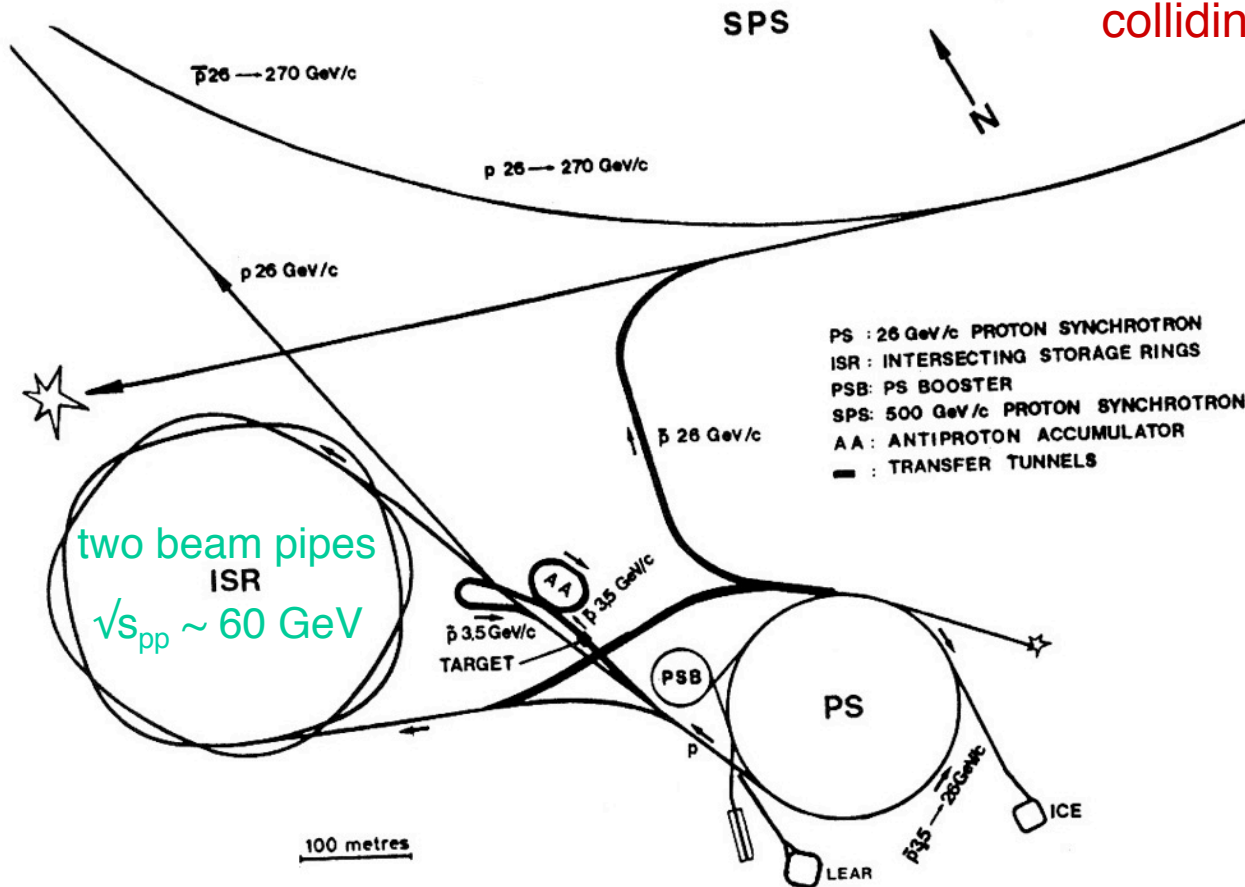


requiring ~ 1 (W, Z leptonic) event / day $L \gg 2.5 \times 10^{29} \text{ cm}^2 \text{ s}^{-1}$ was needed!



The CERN antiproton-proton collider complex (1980-90)

protons and antiprotons counter-rotating in same beam pipe - as in an e+e- machine!
colliding head-on at $\sqrt{s} \sim 500 - 600 \text{ GeV}$



Luminosity:

$$L = n(p)n(\text{anti-p})f/(4\pi\rho^2)$$

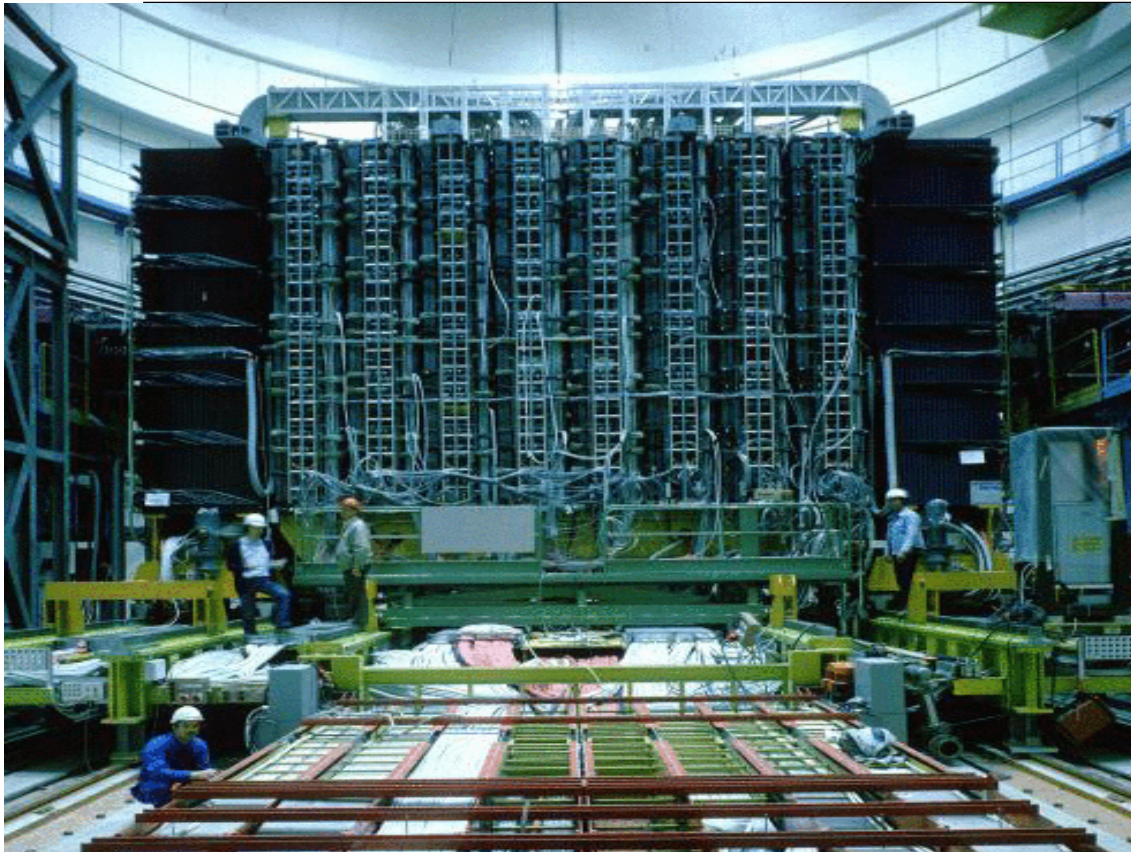
here comes S. Van der Meer and stochastic cooling....



The transformation of the SPS into a collider at C. Rubbia's initiative was accomplished by the summer of 1981 - in $\sim < 3$ years!



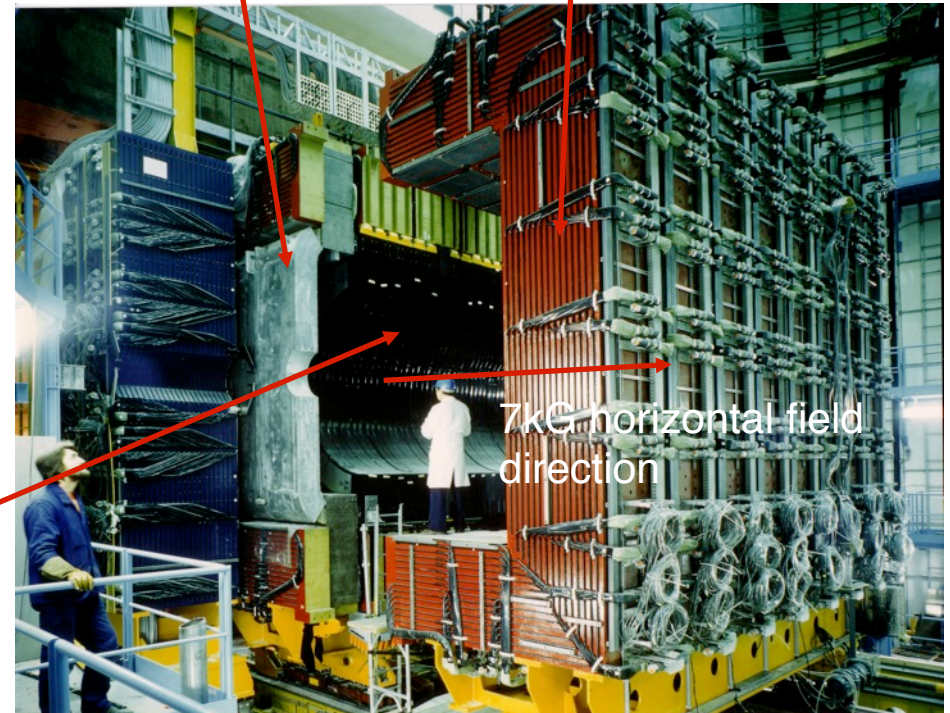
UA1 detector under construction (1979/81!!!)



UA1 detector: $\sim 10 \times 6 \times 6 \text{ m}^3$,
 $\sim 2000 \text{ tons}$ ~ 130 physicists
calorimetric coverage $|\eta| < 5.0$
tracker coverage $|\eta| < 3.0$
muon system coverage $|\eta| < 2.3$

warm Al coil,
7kG horizontal field,

HCAL
5cm iron/1cmScint.
 $3.5 \lambda_{\text{int}}$ deep



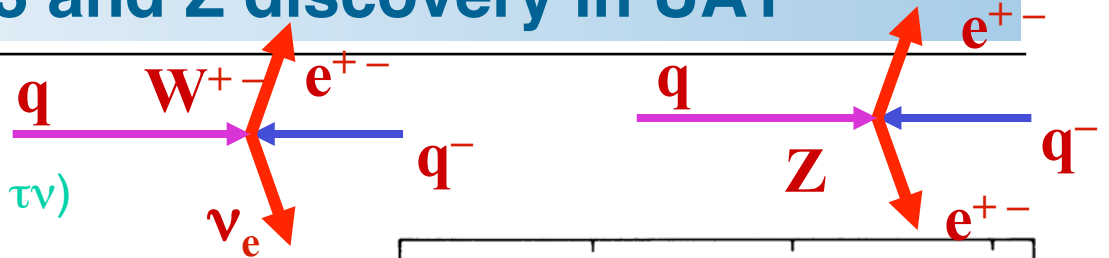
UA1 designed and built in < 3 years!!

ECAL Scint.-Pb sandwich



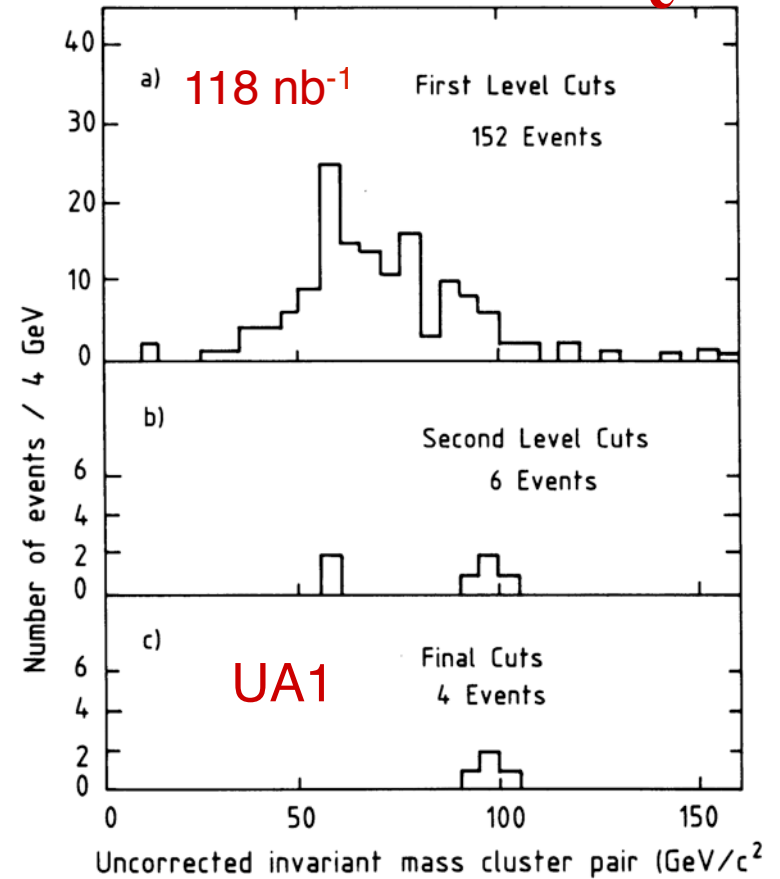
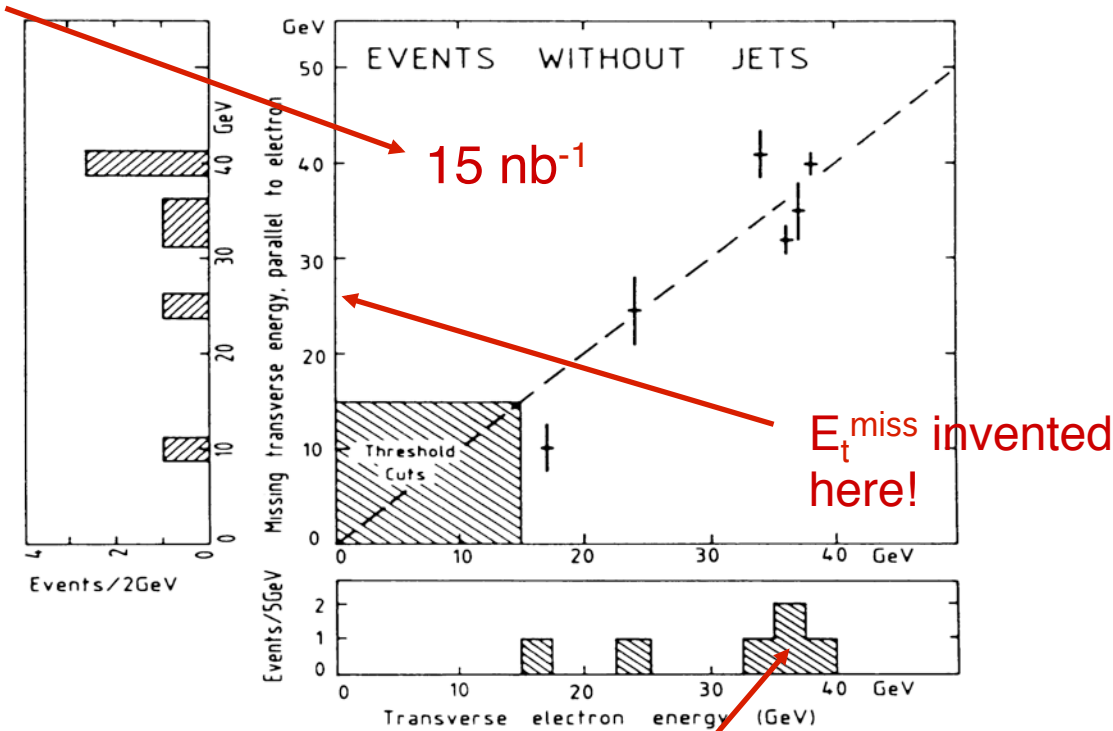
Run of winter 1982, W discovery, followed by run of spring 1983 and Z discovery in UA1

Search for leptonic decays:



6 events selected (5 $W \rightarrow e\nu$ + 1 $W \rightarrow \tau\nu$)

Correlation between missing transverse energy and e^+ transverse energy for the first W events

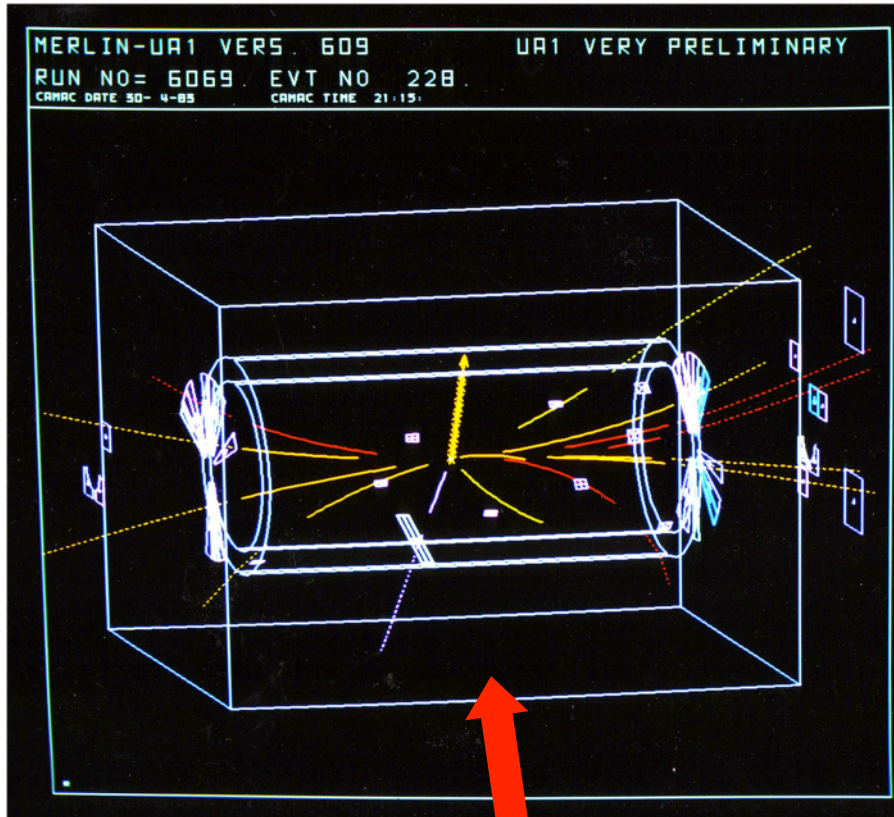


$m_W = 81 \pm 5 \text{ GeV}$ (UA1)
from first "Jacobian peak"

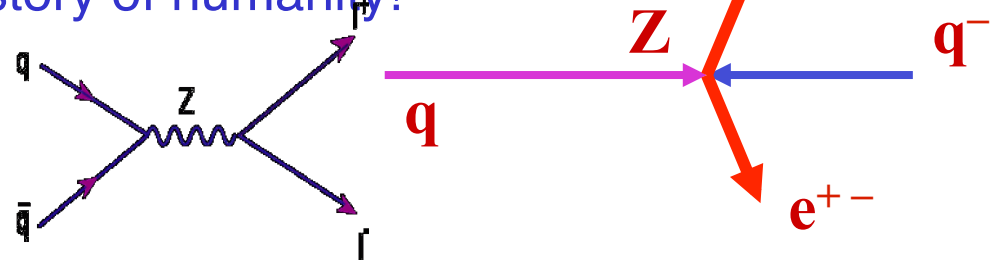
$m_Z = 95.5 \pm 2.5 \pm (3.0) \text{ GeV}$ (UA1)
 $\sigma_Z \text{BR}(Z \rightarrow ll) = 41 \pm 21 (\pm 7) \text{ pb}$



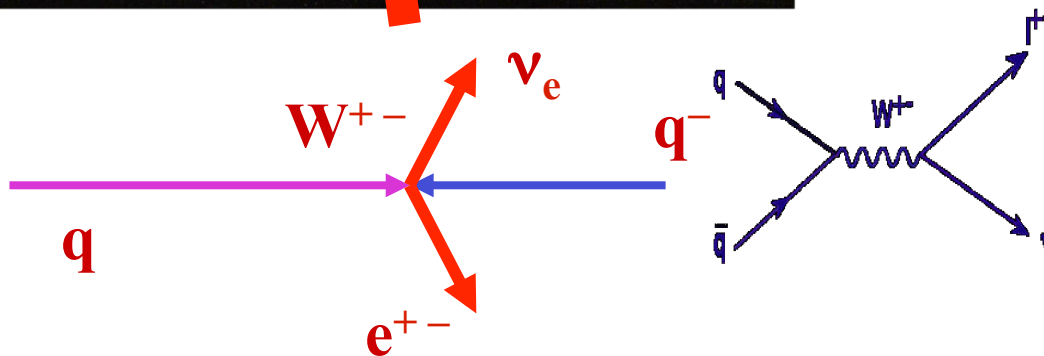
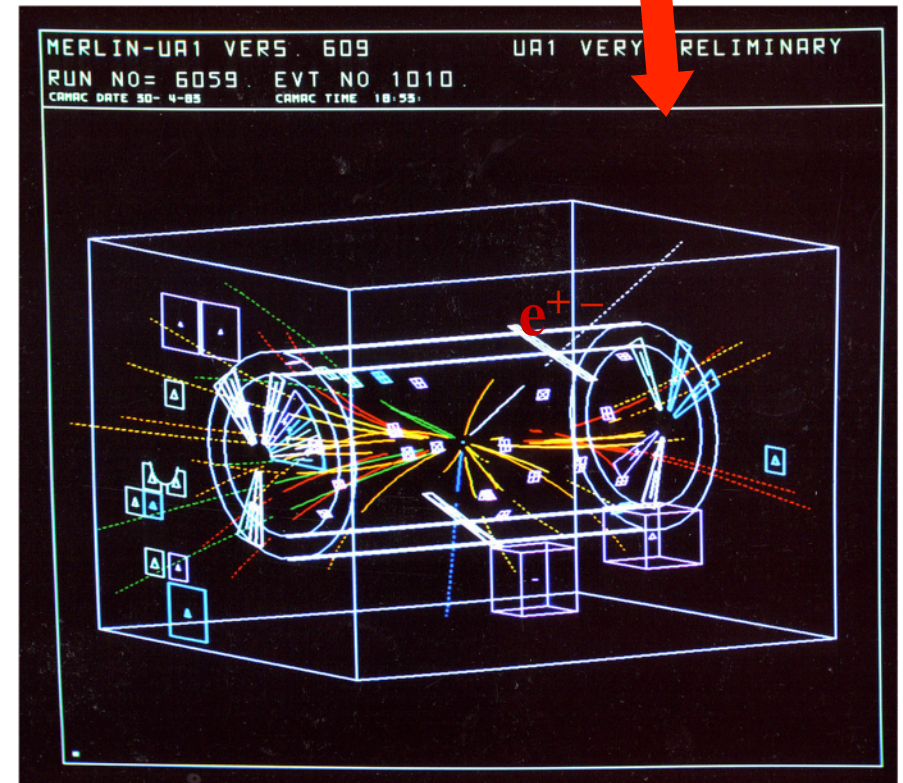
First $W \rightarrow e\nu$ events in UA1 (Dec.1982, Jan.1983) and first $Z \rightarrow e^+e^-$ events in UA1 (May 1983)



First W 's and Z 's in the history of humanity!



Megatek interactive graphic displays





...

.....after W, Z discovery the obvious next goal was elucidating the origin of electroweak symmetry breaking, the masses of W and Z, through the Higgs (BEH) or other alternative mechanism.....

.....there was a possibility to discover the Higgs boson at LEP - provided it was light enough - but theoretical upper limits on m_H in ~ 1990 were $m_H < \sim 1000$ GeV, thus much stronger machines, SSC or LHC were required.....



The Large Hadron Collider (LHC) - the genesis



The LHC project started at the initiative (and with the daring!!) of C. Rubbia

The Aachen Conference marked the real start-up, since then work on the collider and magnets, various detector designs and understanding physics issues went on without let-up

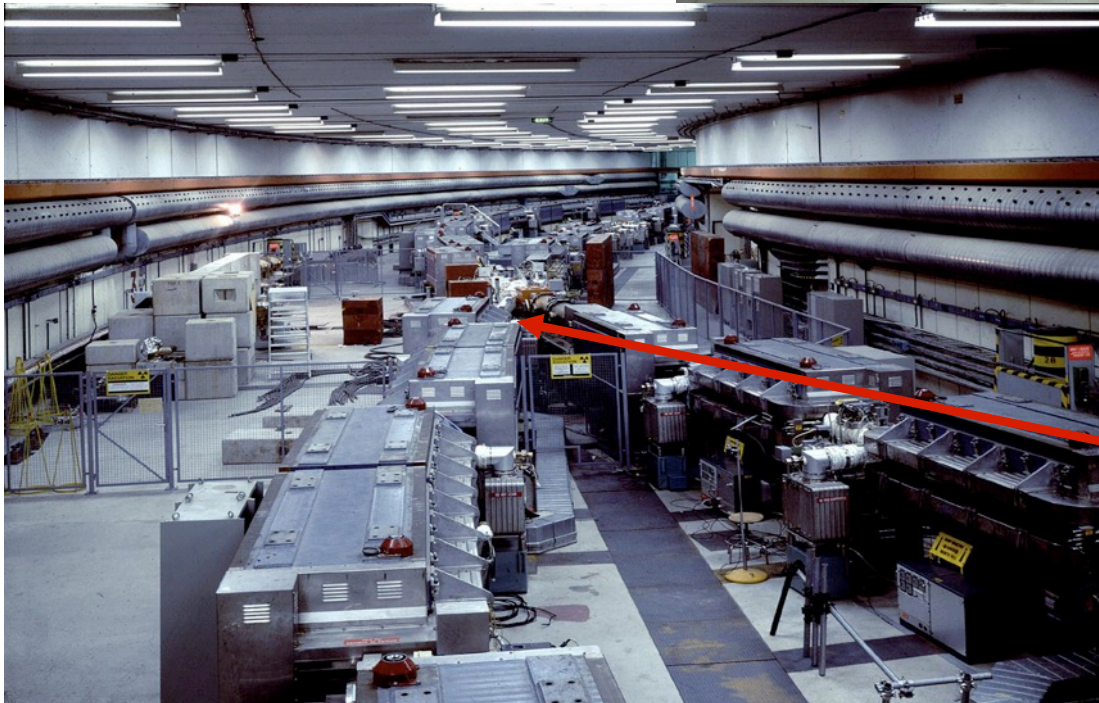
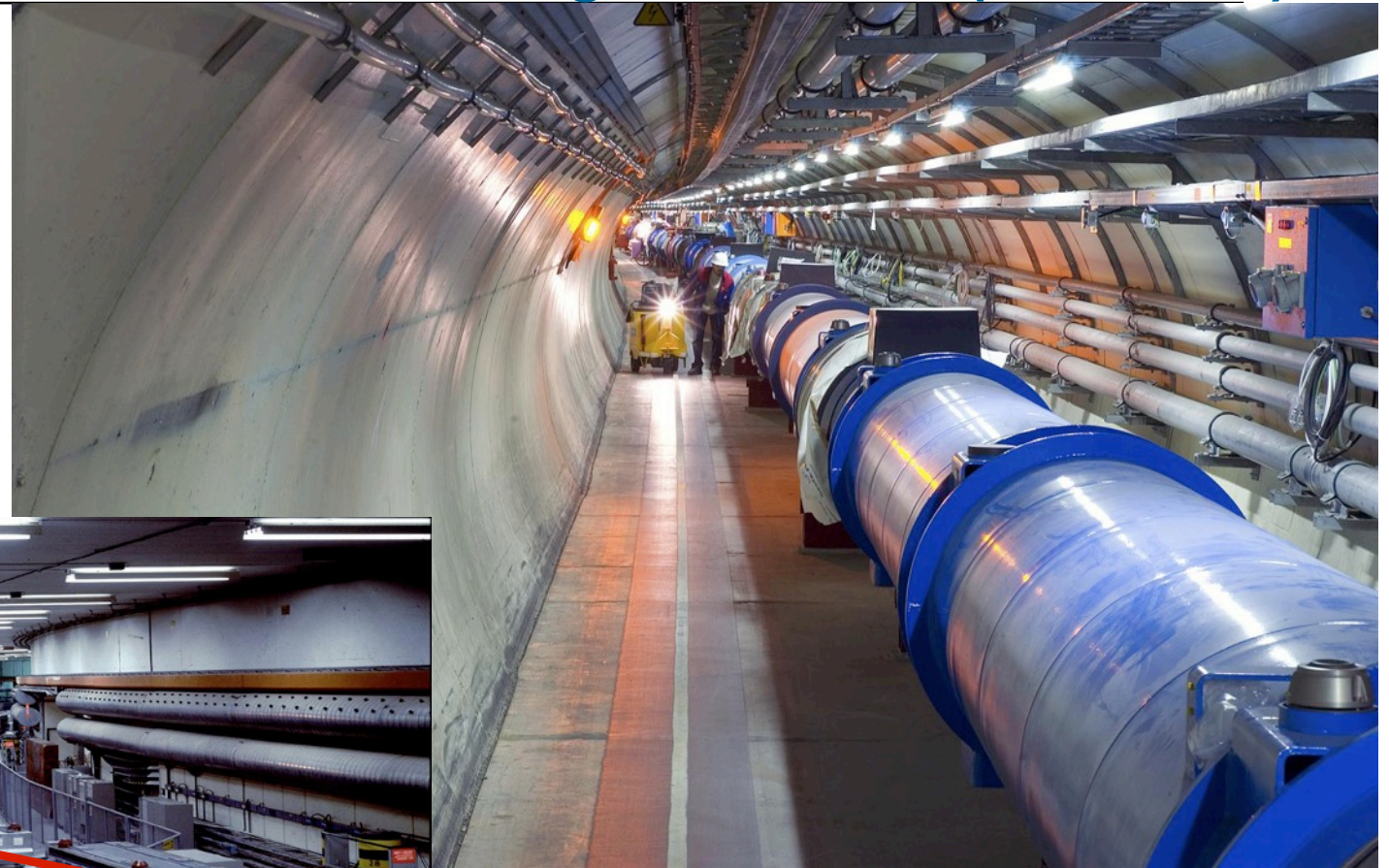
Scientifico-diplomatic trips in 1990/91/92 to Japan, India, Russia, USA, Canada etc

LHC vs SSC: Rubbia's arguments: savings!

- existing LEP tunnel ~1 GCHF
 - existing infrastructure at CERN (PS, SPS, etc) ~ 1 GCHF
 - "two-in-one" scheme for dipoles saves ~ half the cost of magnet ~ 0.7 to 1 GCHF
- thus overall LHC cost ~ 3 GCHF
- will be ready by 1998 - 2000 !!



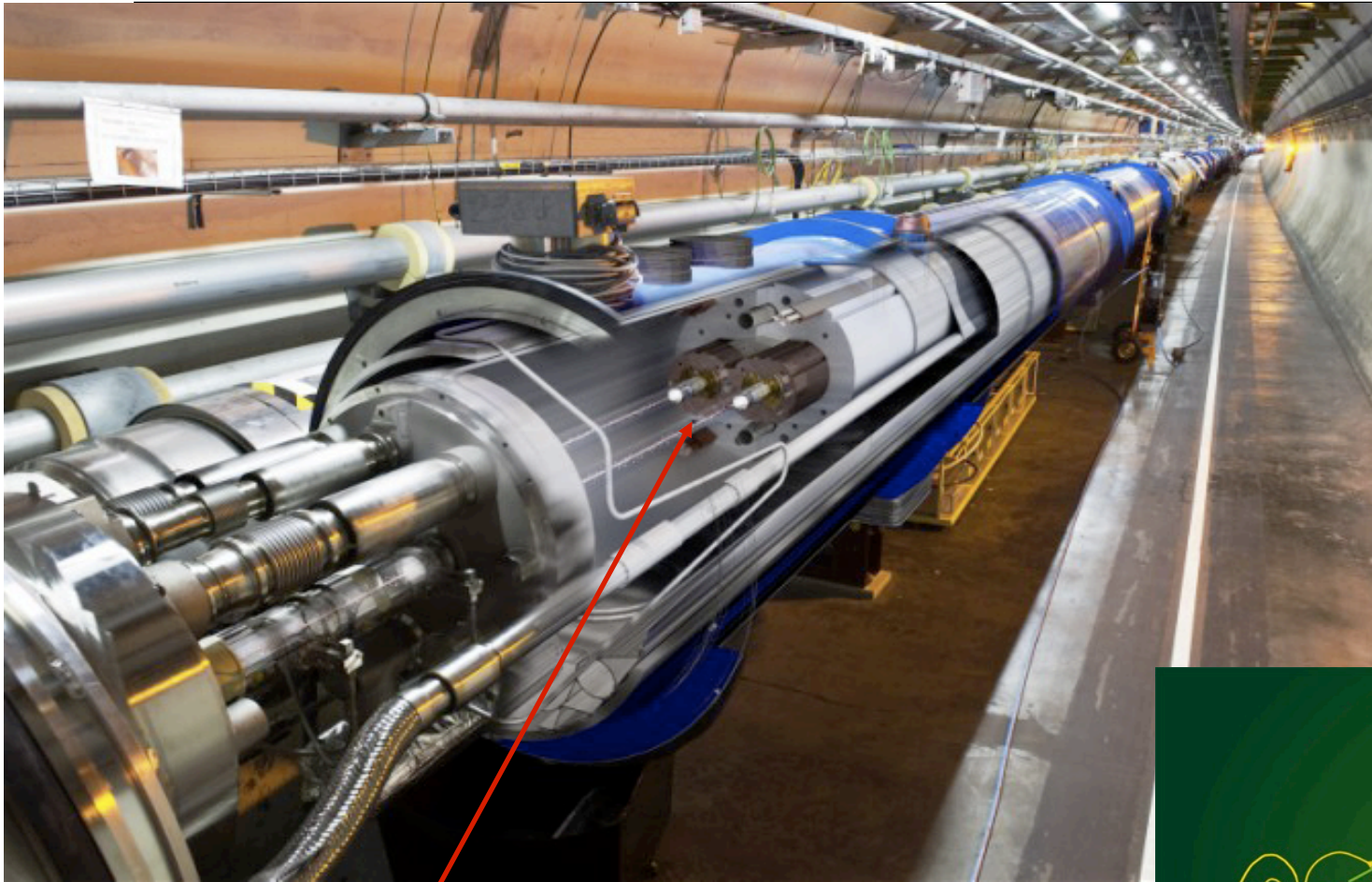
The LHC - a single set of magnets in the tunnel thanks to the “two-in-one” magnet scheme (R. Palmer)



the ISR - a typical proton-proton collider;
ISABELLE and the SSC were of same design



The LHC tunnel - a single set of magnets thanks to the “two-in-one” magnet scheme - II

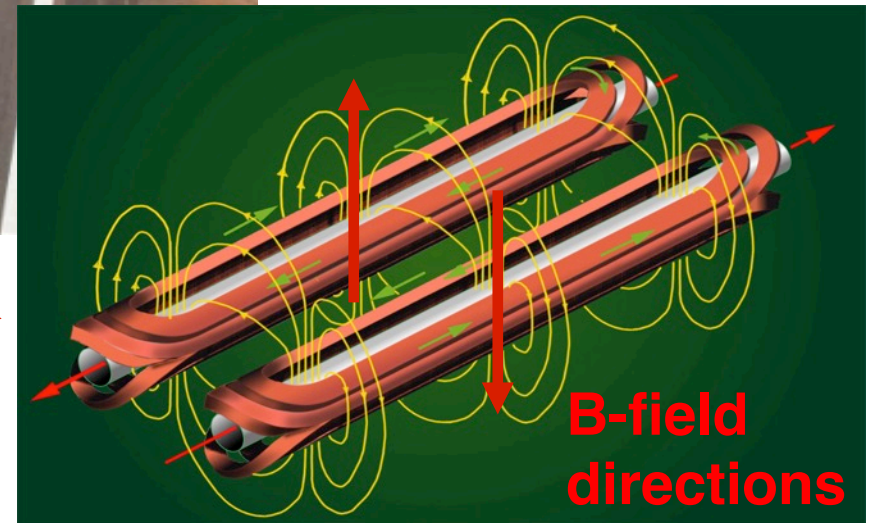


The magnets have to be aligned with $100\mu\text{m}$ precision over the 27 km circumference!

Vacuum in the beam pipe as on the Moon!

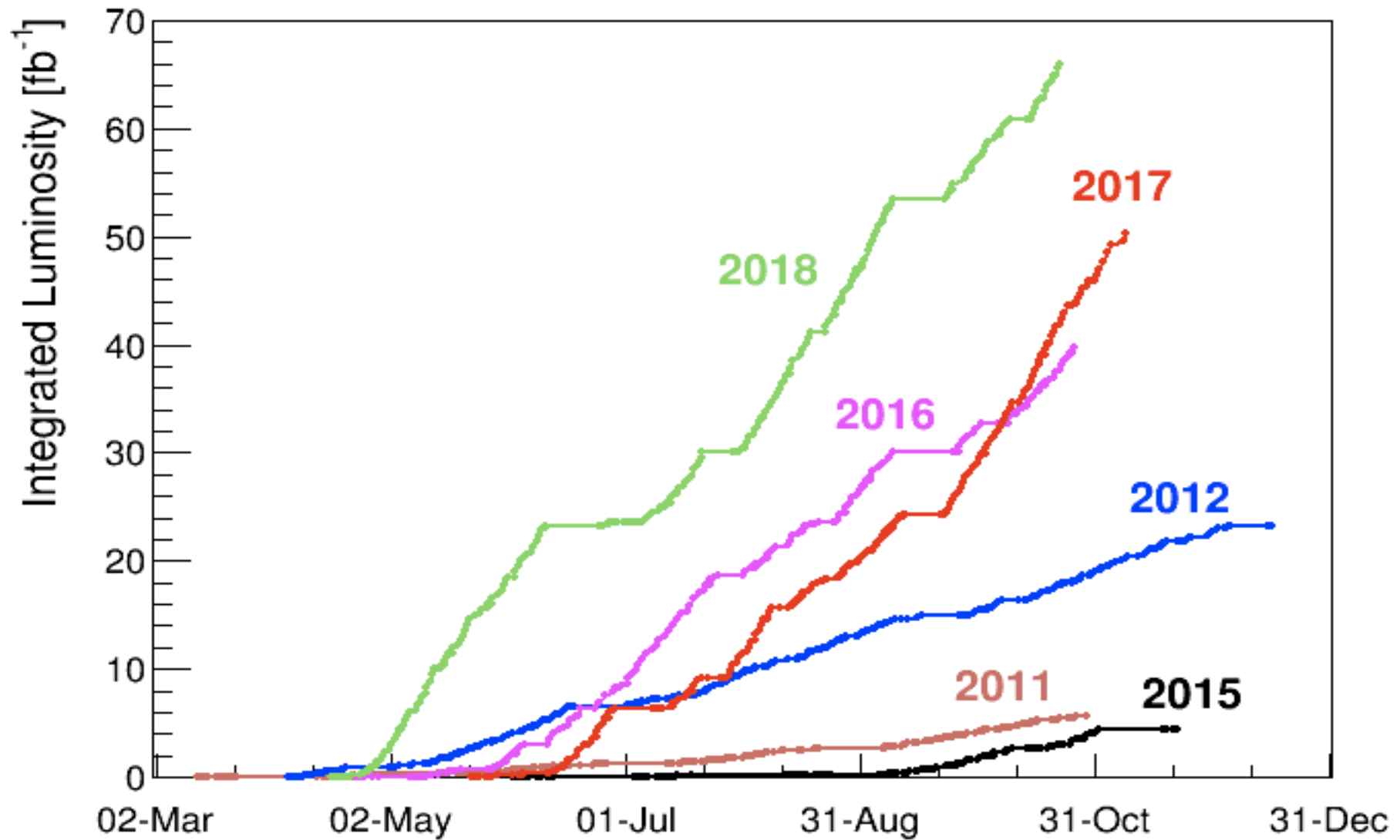
Magnetic field lines

the two-in-one magnet
40.000 tons of material at $-271\text{ }^{\circ}\text{C}$
The coldest place in the Univers!





LHC luminosities over years, including 2018





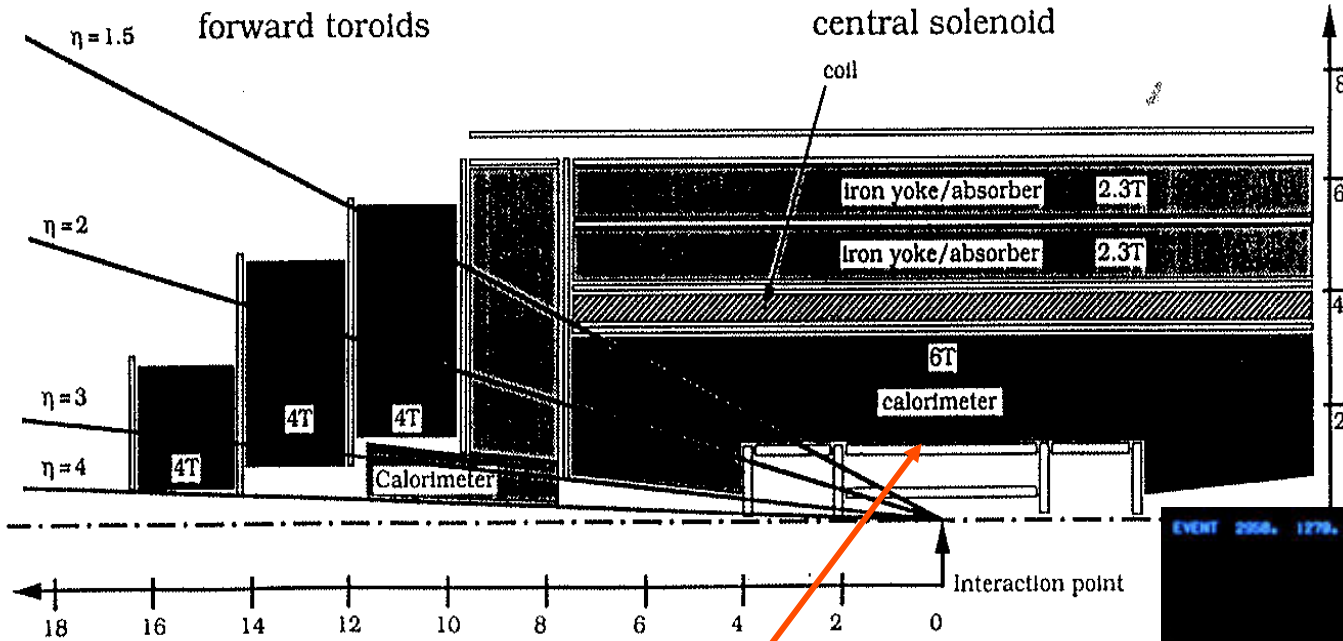
How CMS started in 1990!

Compact Muon Solenoid (CMS)

weights :	2 very forward toroids	750 tons
	4 forward toroids	5000 tons
	solenoid	11500 tons
	calorimetry	4000 tons
	Total	20750 tons

muon chambers

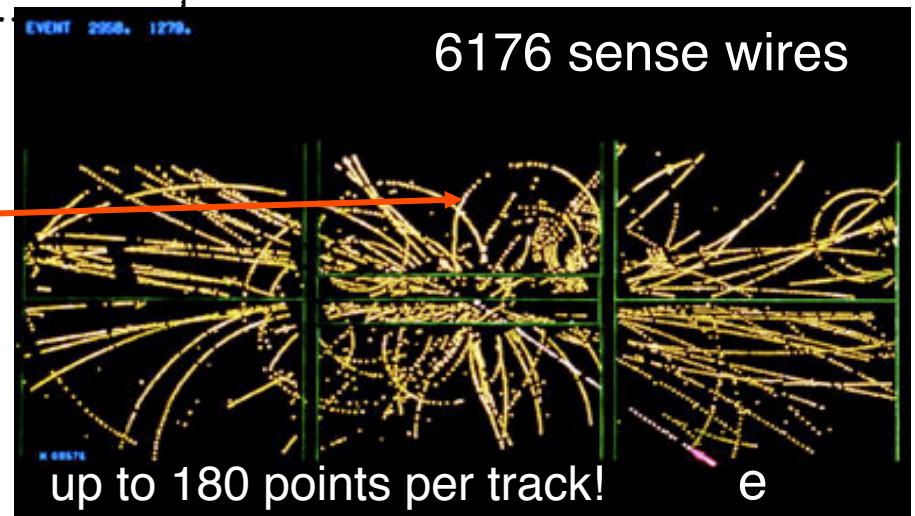
Proto-CMS in 1990.....



The UA1 tracker was by far the most sophisticated in its days, but the p-pbar collider did not exceed $10^{30} \text{cm}^{-2} \text{s}^{-1}$ and now for the LHC we need $10^{34} \text{cm}^{-2} \text{s}^{-1}$!! **The answer:** granularity and fast detector response

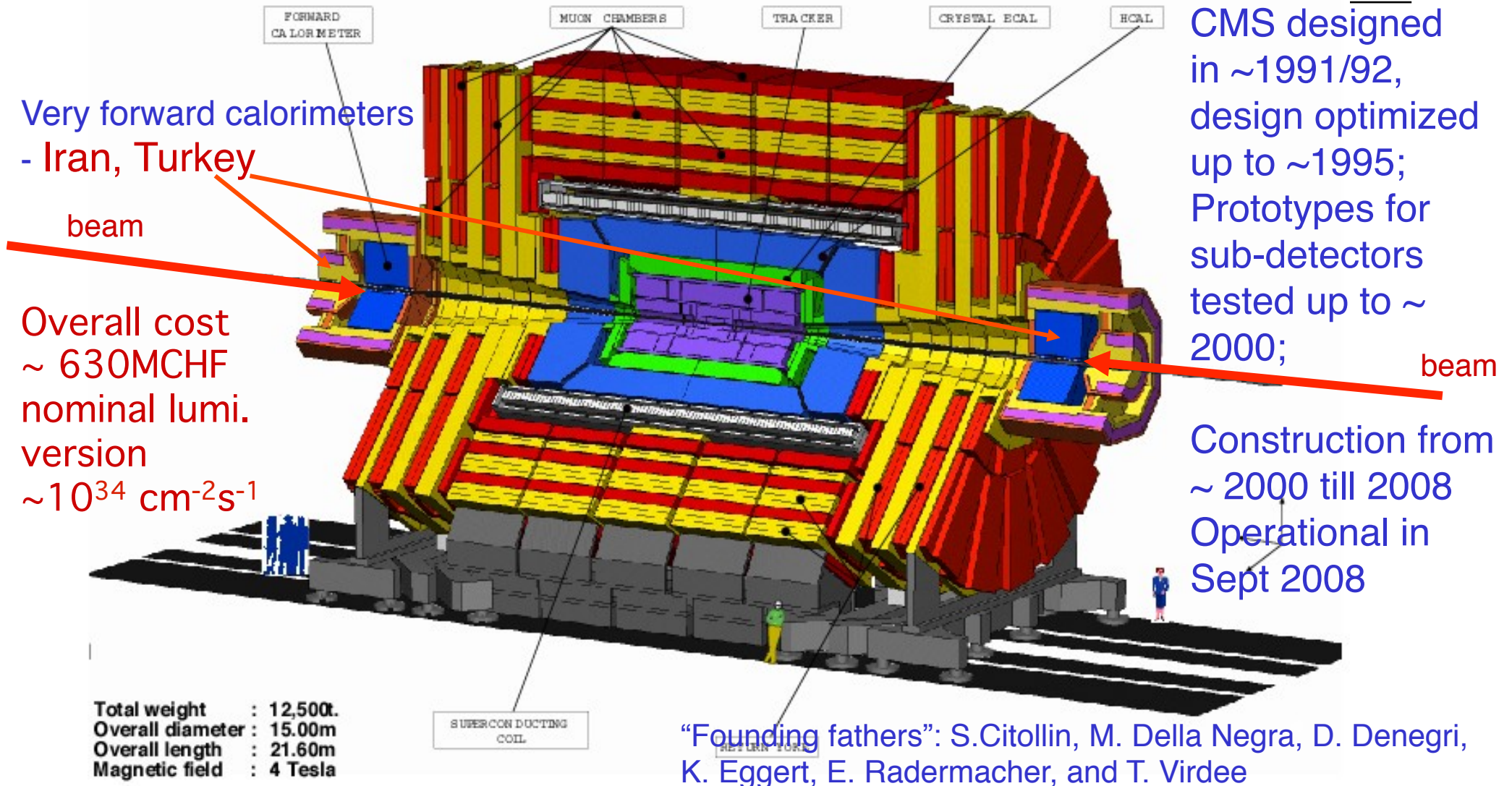
- Tracking? ...maybe in outermost regions...

- Our test beam and MC studies in 1991/92 led to rapid progress in the detector design.





The CMS (Compact Muon Solenoid) detector

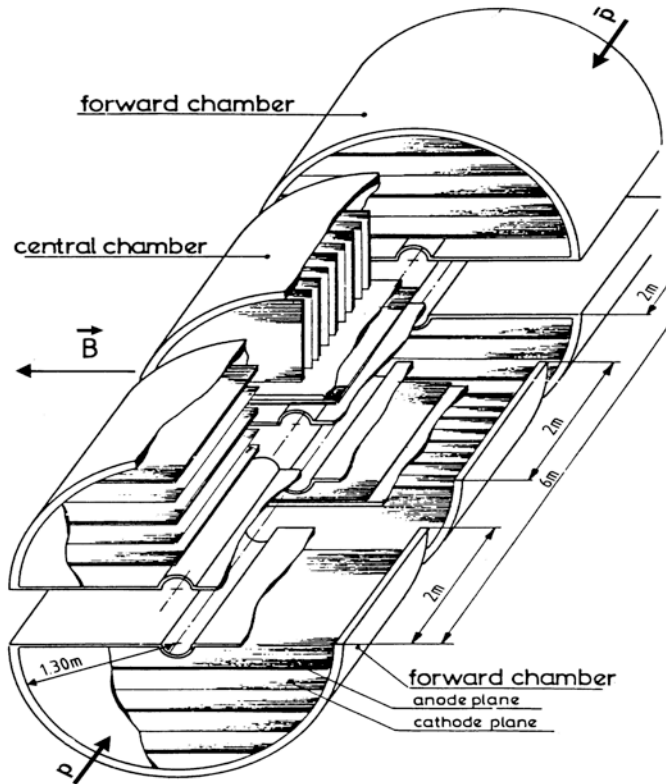


At present undergoing upgrades in several phases to face conditions of HL-LHC running ($\sim 10^{35} \text{ cm}^{-2}\text{s}^{-1}$)

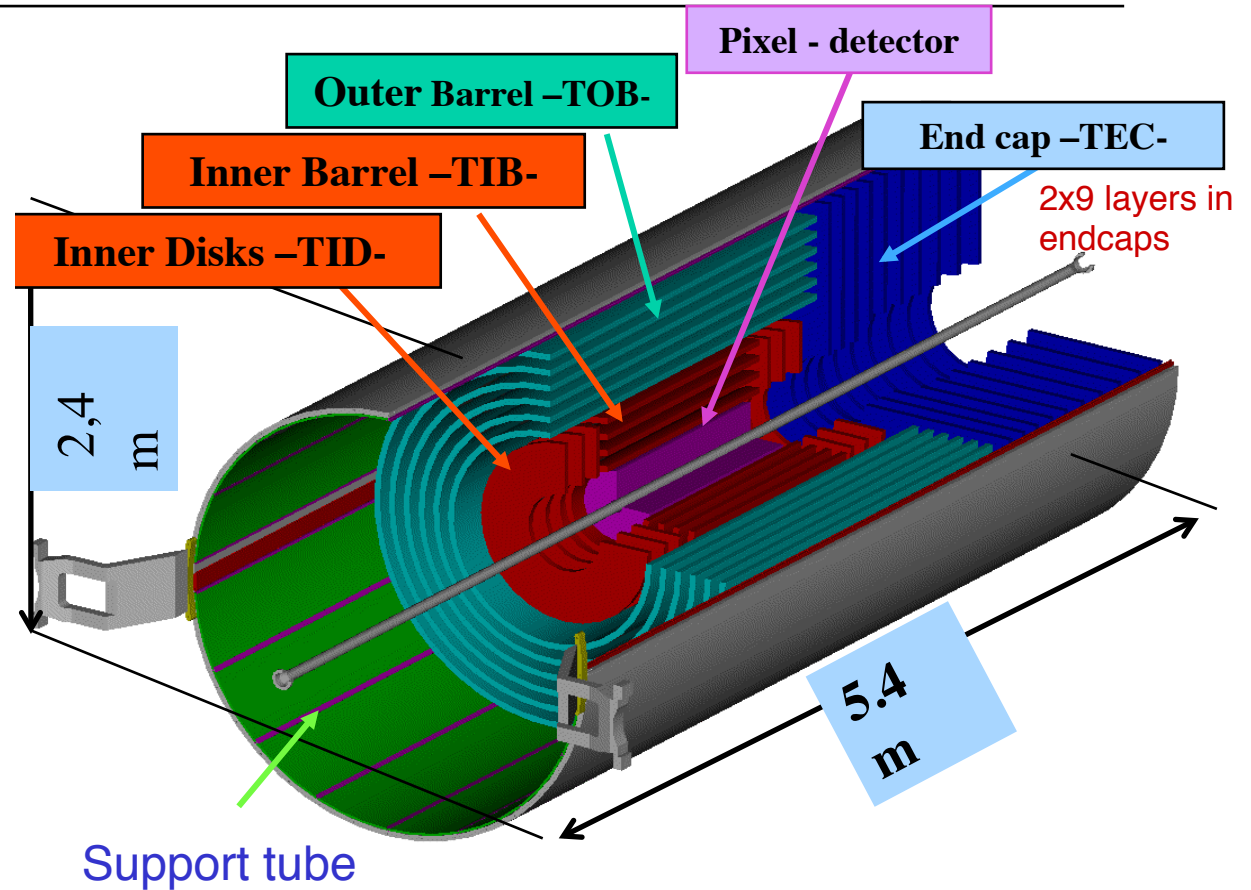


The UA1 tracker and the all-Silicon CMS tracker - high granularity for pattern recognition and fast detectors

The UA1 tracker was by far the most sophisticated in its days !!



UA1 tracker: Imaging drift chamber, 6m long, 2.3m in diameter, 6176 sense wires, up to 180 hits per track, Acceptance $|\eta| < 3.0$

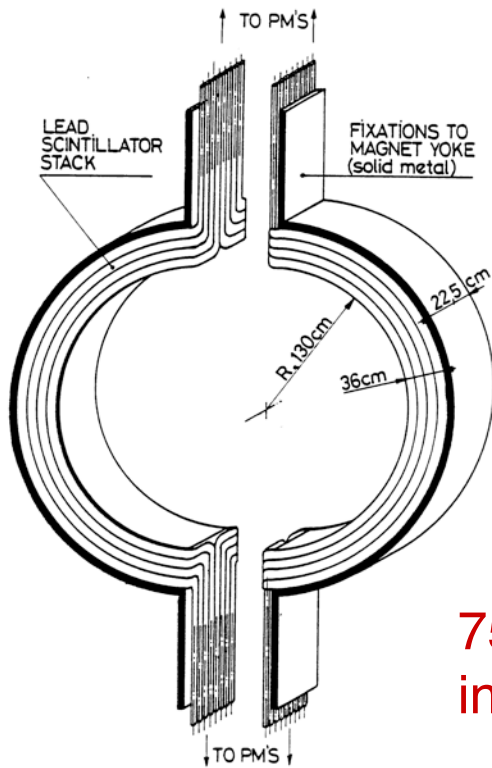


Factor $\sim 10^4$ in granularity!

210 m² of silicon sensors
 $\sim 6,000$ thin detectors (1 sensor)
 $\sim 9,000$ thick detectors (2 sensors)
 10 million microstrips and
 70 Million pixels (~ 2 m²)



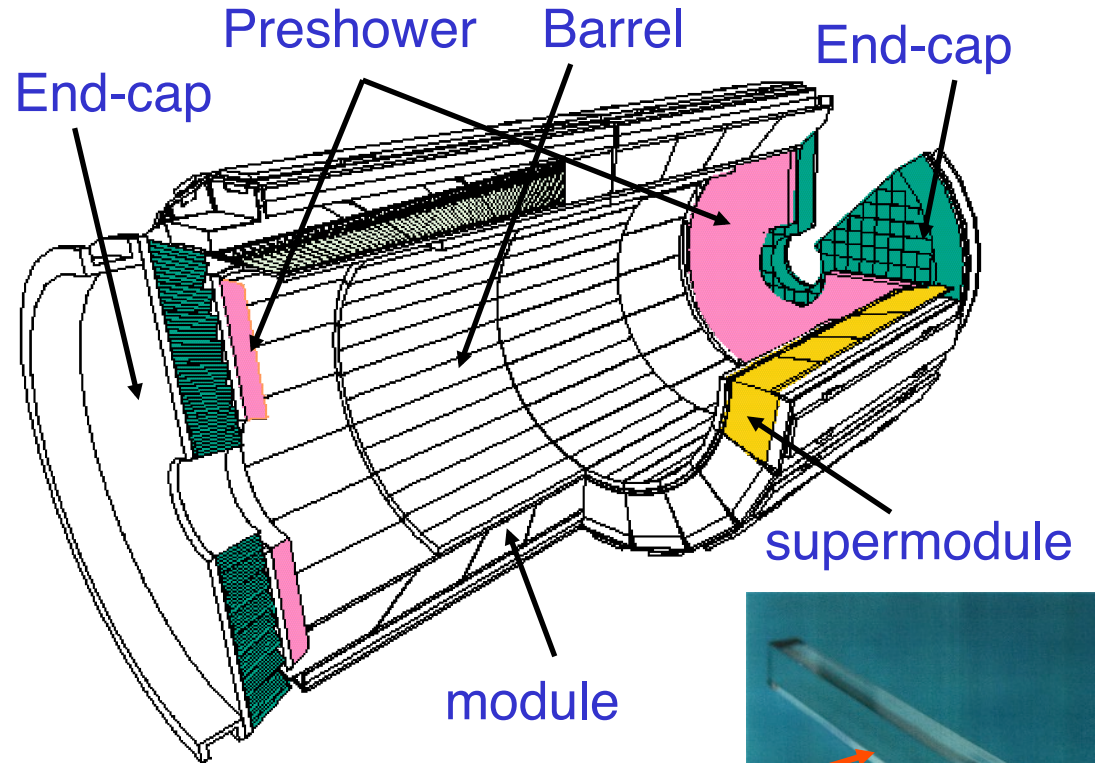
From the UA1 electromagnetic calorimeter to the crystal (PbWO_4) calorimeter of CMS



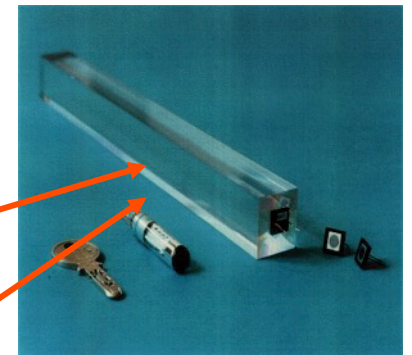
754 PM's
in barrel

UA1 ECAL (2x24 gondolas)
 Scint.-Pb sandwich, 1.2mmPb/1.5mmSci
 $\Delta\phi\Delta\eta = 180^\circ \times 0.14$; $27X_0$ deep,
 four segments in depth
 ECAL acceptance: $|\eta| < 3.0$

It is here that the W and Z were found!



a TeV em shower contained in a crystal of $24 \times 2 \times 2 \text{cm}^3$



74.000 crystals

$L_{\text{rad}} = 9 \text{mm}$, $R_{\text{mol}} = 2 \text{cm}$ ($\Delta\eta = 0.014$)
 $\sim 10 \text{nsec}$ response read out with avalanche photo-diodes in barrel

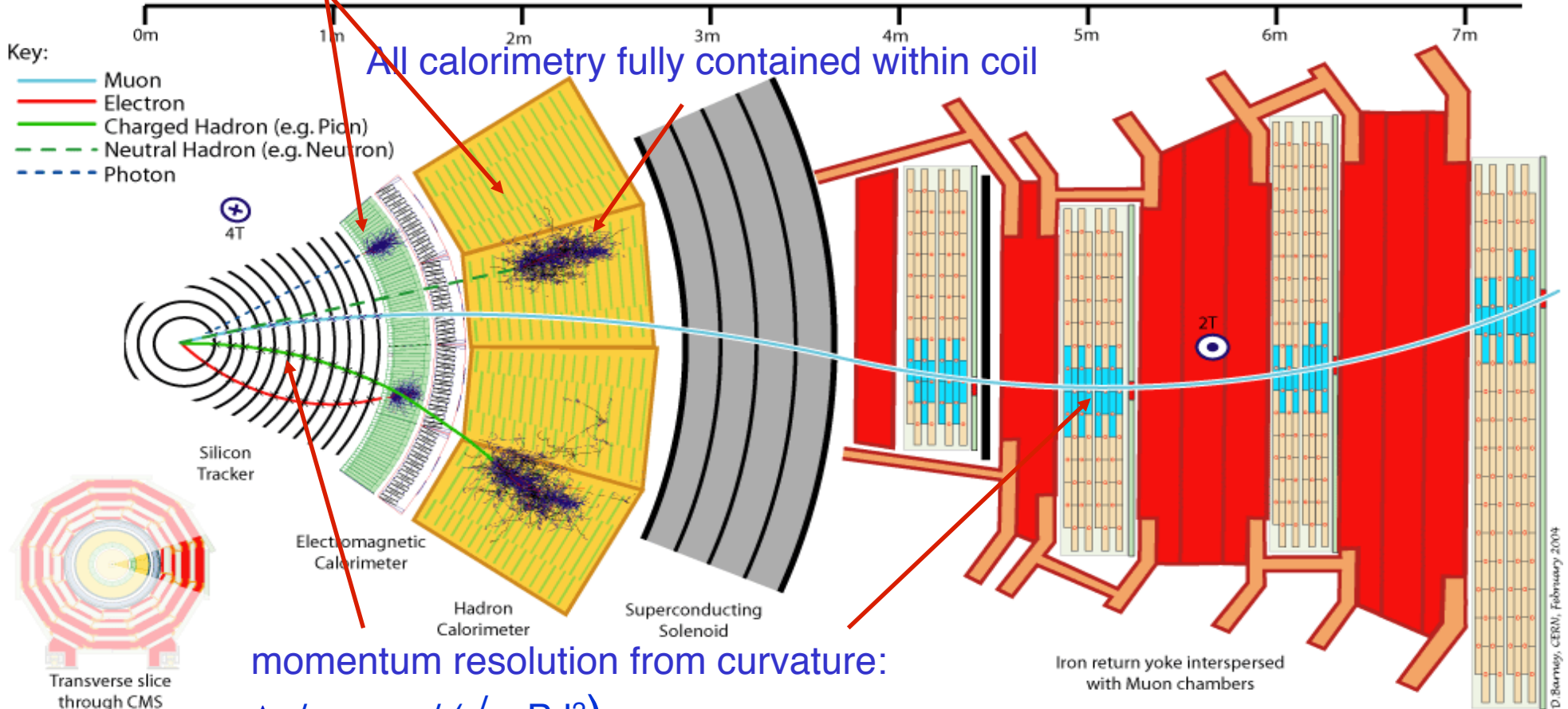
**Factor $\sim 10^2$
in granularity!**



Central region of CMS; detector functions

Tracking + Ecal + Hcal + Muons for $|\eta| < 2.4$

$\Delta E/E \sim \sqrt{((a/\sqrt{E})^2 + (b/E)^2 + c^2)}$ generic resolution in calorimeters



All calorimetry fully contained within coil

momentum resolution from curvature:

$\Delta p/p \sim \epsilon p / (\sqrt{n} B l^2)$

Redundancy and robustness in muon system

Si TRACKER

Silicon Microstrips and Pixels

CALORIMETERS

ECAL
Scintillating PbWO₄ crystals

HCAL
Plastic scintillator/brass sandwich

MUON BARREL

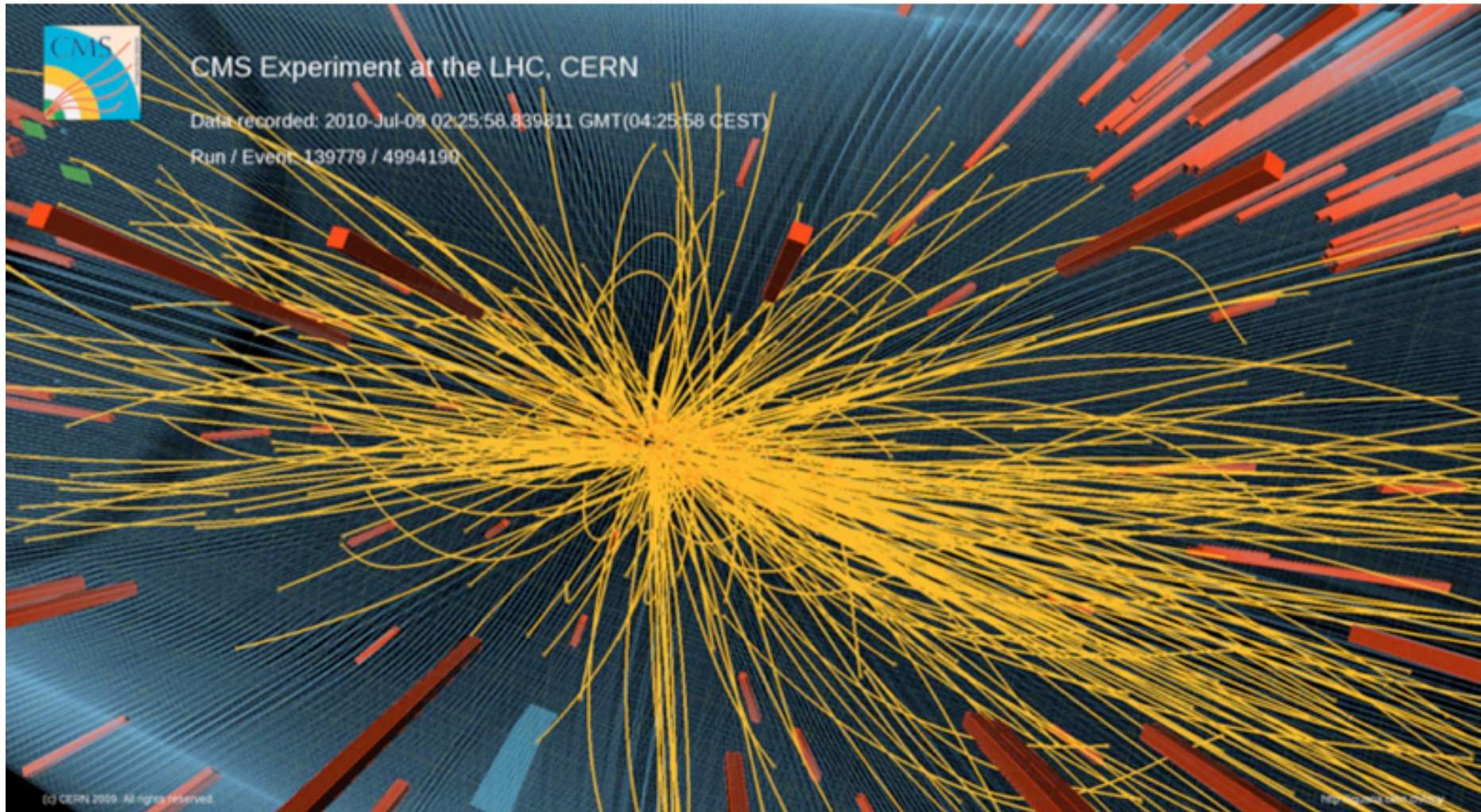
Drift Tube Chambers (DT) Resistive Plate Chambers (RPC)



**A typical proton-proton collision at the LHC,
there are \sim one billion such collisions per second**

(in reality they are clustered by ~ 25 every 25 nanoseconds – pile-up)

Expected Higgs boson production rate is less than one in a billion pp collisions!





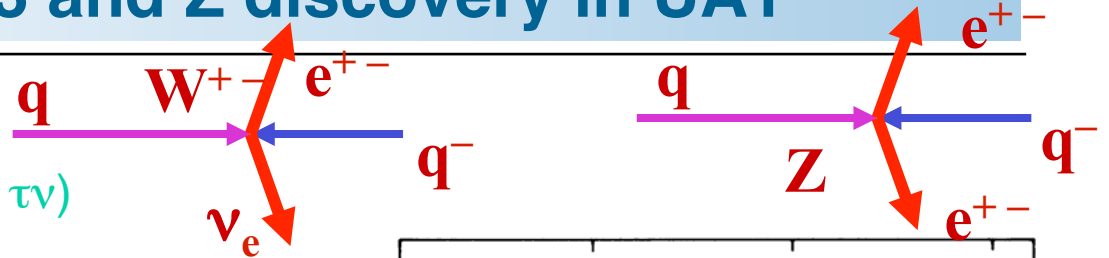
Physics

jets, W, Z, top, Higgs boson



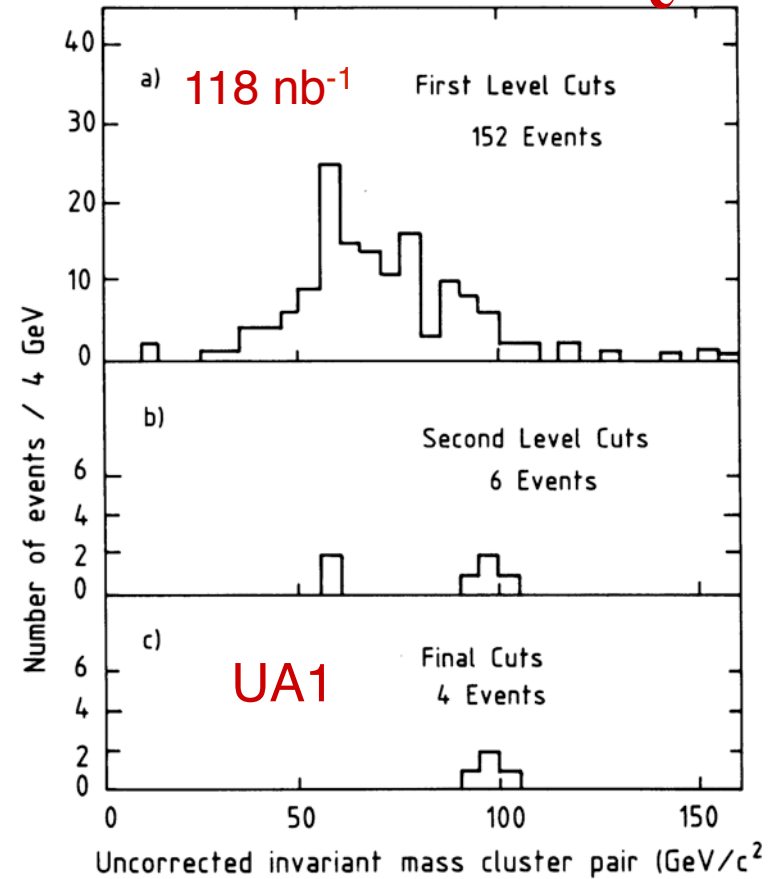
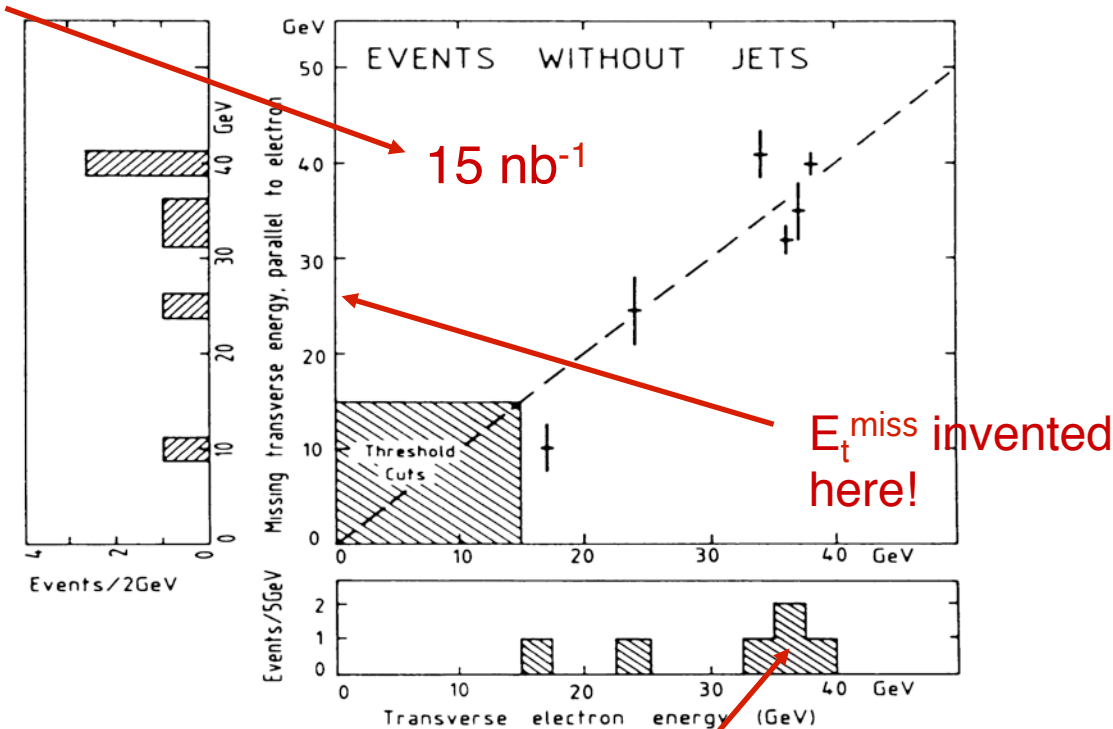
Run of winter 1982, W discovery, followed by run of spring 1983 and Z discovery in UA1

Search for leptonic decays:



6 events selected (5 $W \rightarrow e\nu$ + 1 $W \rightarrow \tau\nu$)

Correlation between missing transverse energy and e^+ transverse energy for the first W events



$m_W = 81 \pm 5 \text{ GeV}$ (UA1)
from first "Jacobian peak"

$m_Z = 95.5 \pm 2.5 \pm (3.0) \text{ GeV}$ (UA1)
 $\sigma_Z \text{BR}(Z \rightarrow ll) = 41 \pm 21 (\pm 7) \text{ pb}$

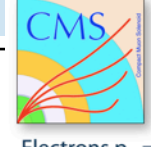
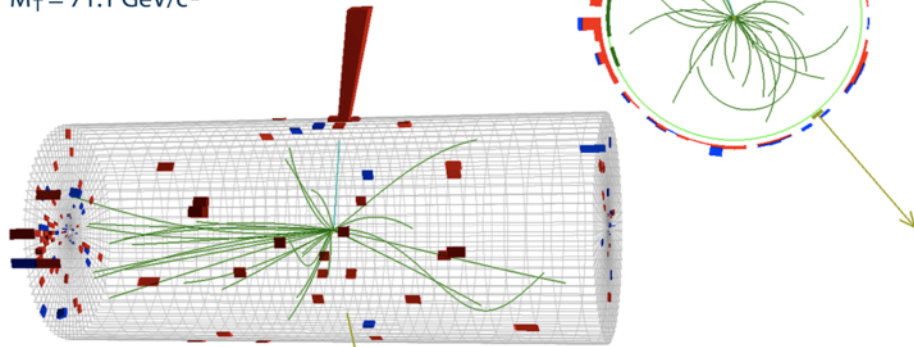


LHC : First $W \rightarrow e\nu$ and $Z \rightarrow e^+e^-$ events, April 2010



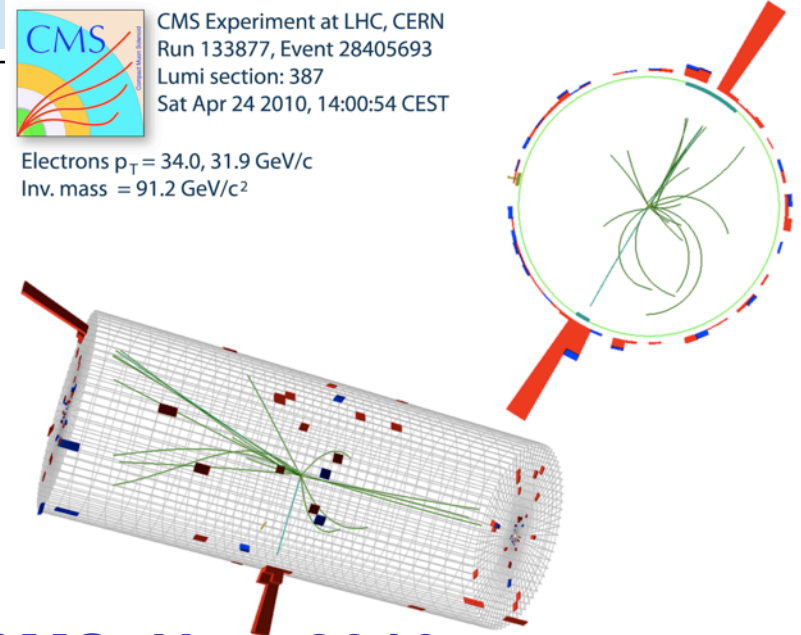
CMS Experiment at LHC, CERN
Run 133874, Event 21466935
Lumi section: 301
Sat Apr 24 2010, 05:19:21 CEST

Electron $p_T = 35.6$ GeV/c
 $ME_T = 36.9$ GeV
 $M_T = 71.1$ GeV/c²

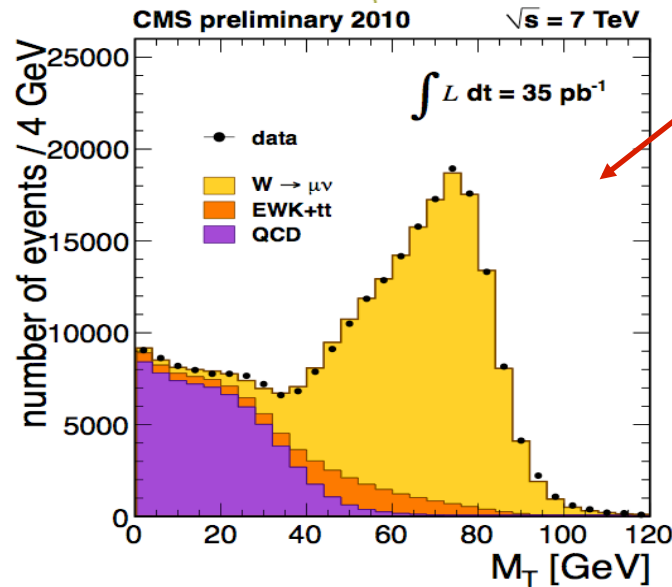


CMS Experiment at LHC, CERN
Run 133877, Event 28405693
Lumi section: 387
Sat Apr 24 2010, 14:00:54 CEST

Electrons $p_T = 34.0, 31.9$ GeV/c
Inv. mass = 91.2 GeV/c²

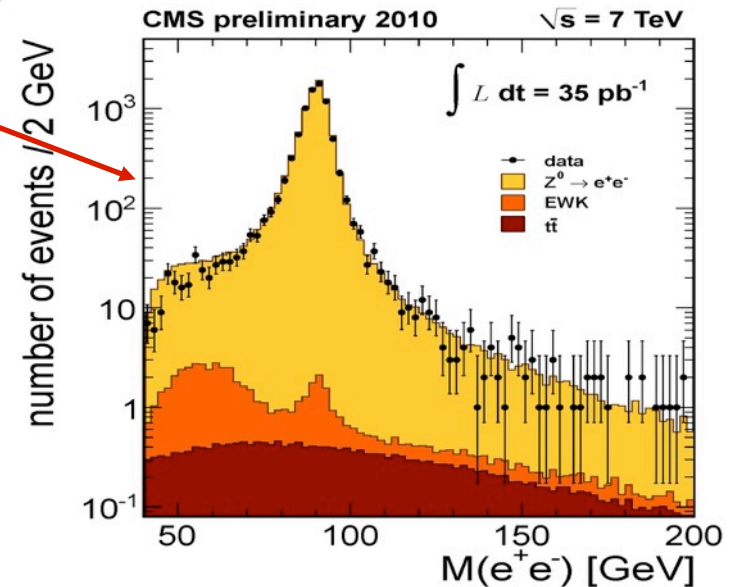


W and Z spectra in CMS, Nov. 2010



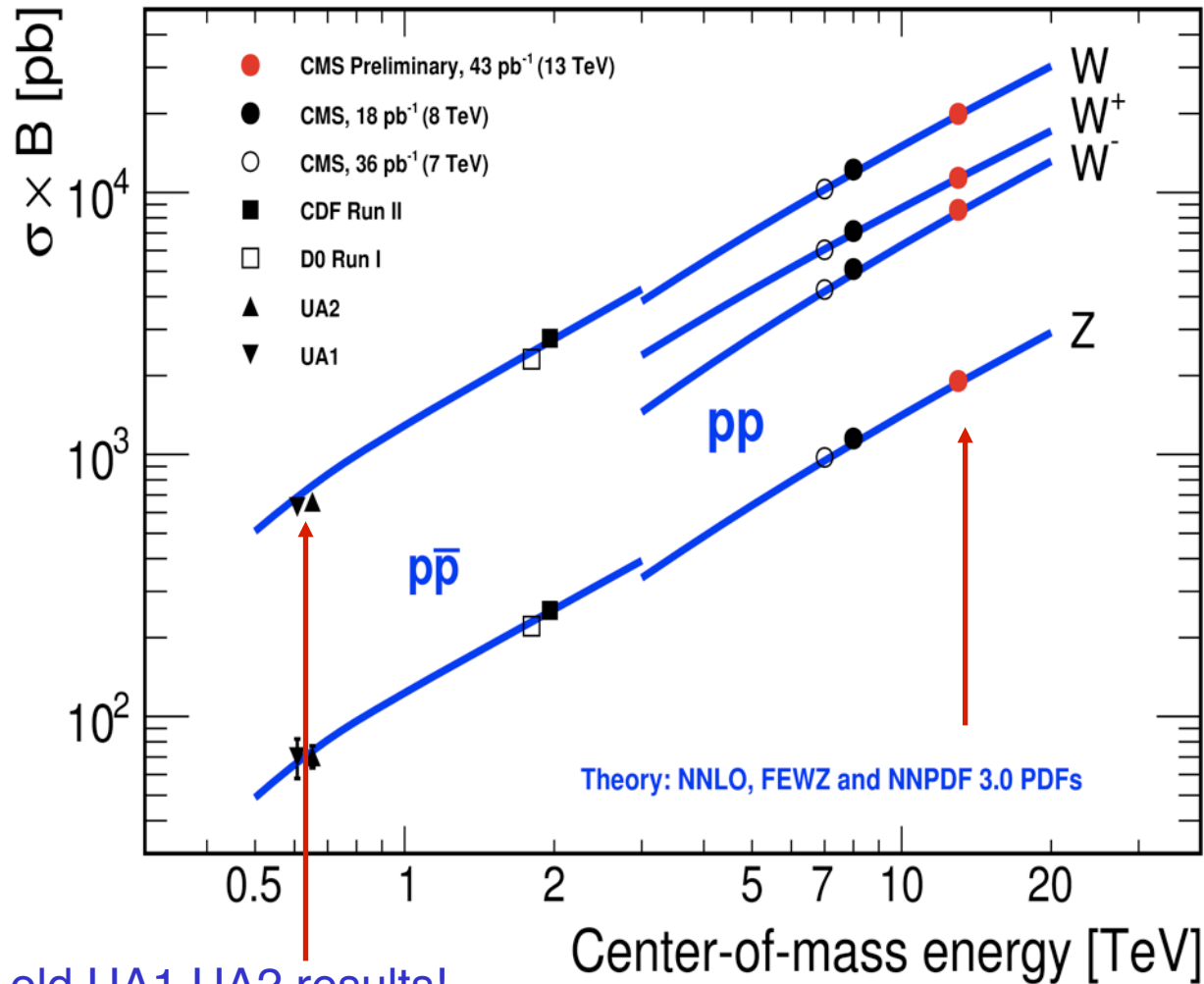
By end-2012 we had
~150.000.000 W and
~15.000.000 Z
decaying leptonically!!

Note progress since
first UA1 W, Z!

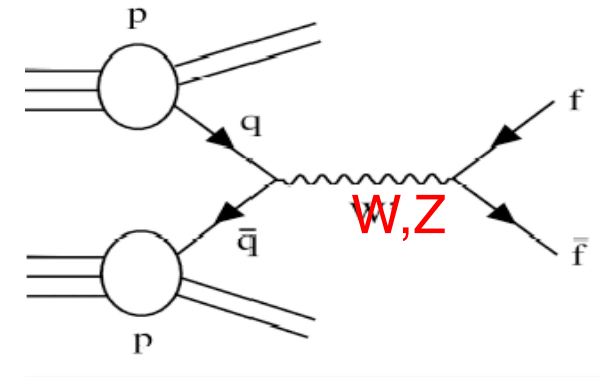




W, Z ATLAS and CMS studies, 2018, total cross sections from the CERN pp-bar collider to the Tevatron and the LHC



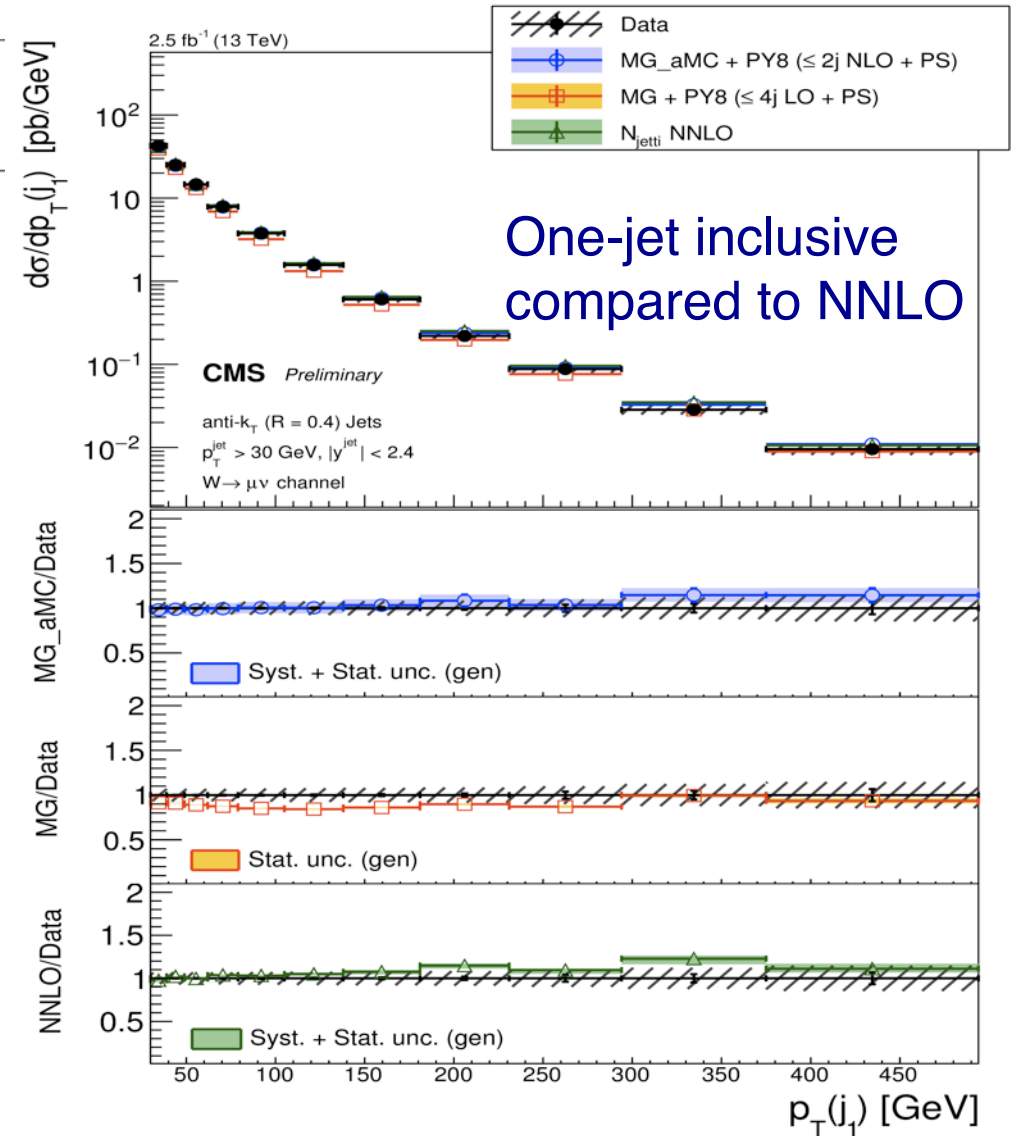
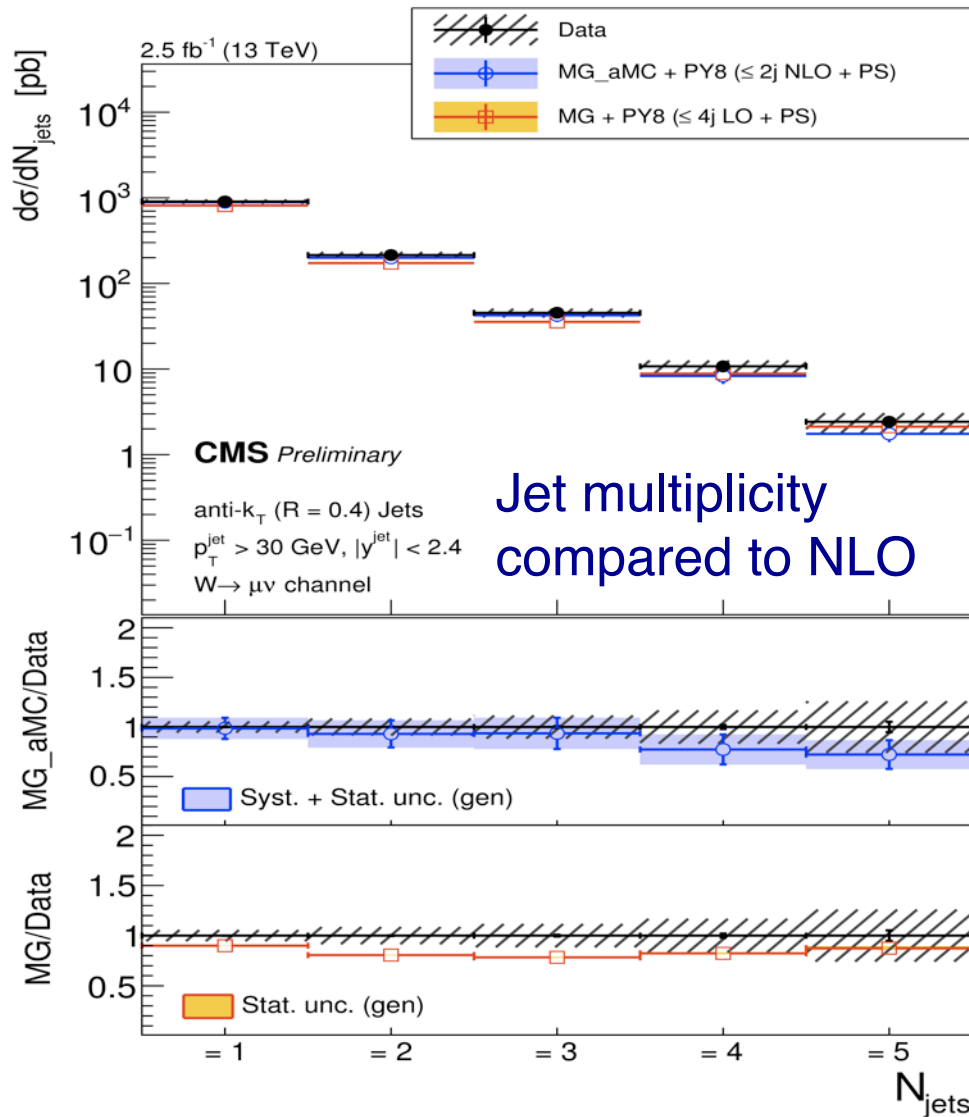
old UA1,UA2 results!



Comprehensive W and Z
results from ATLAS and CMS
– New results available at 7, 8
and 13 TeV
– Allow to test the SM with
sub-percent uncertainties
– Measured cross sections
agree with NNLO QCD
and NLO EW



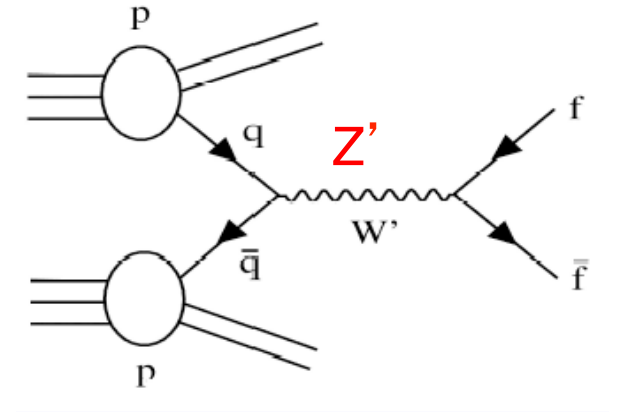
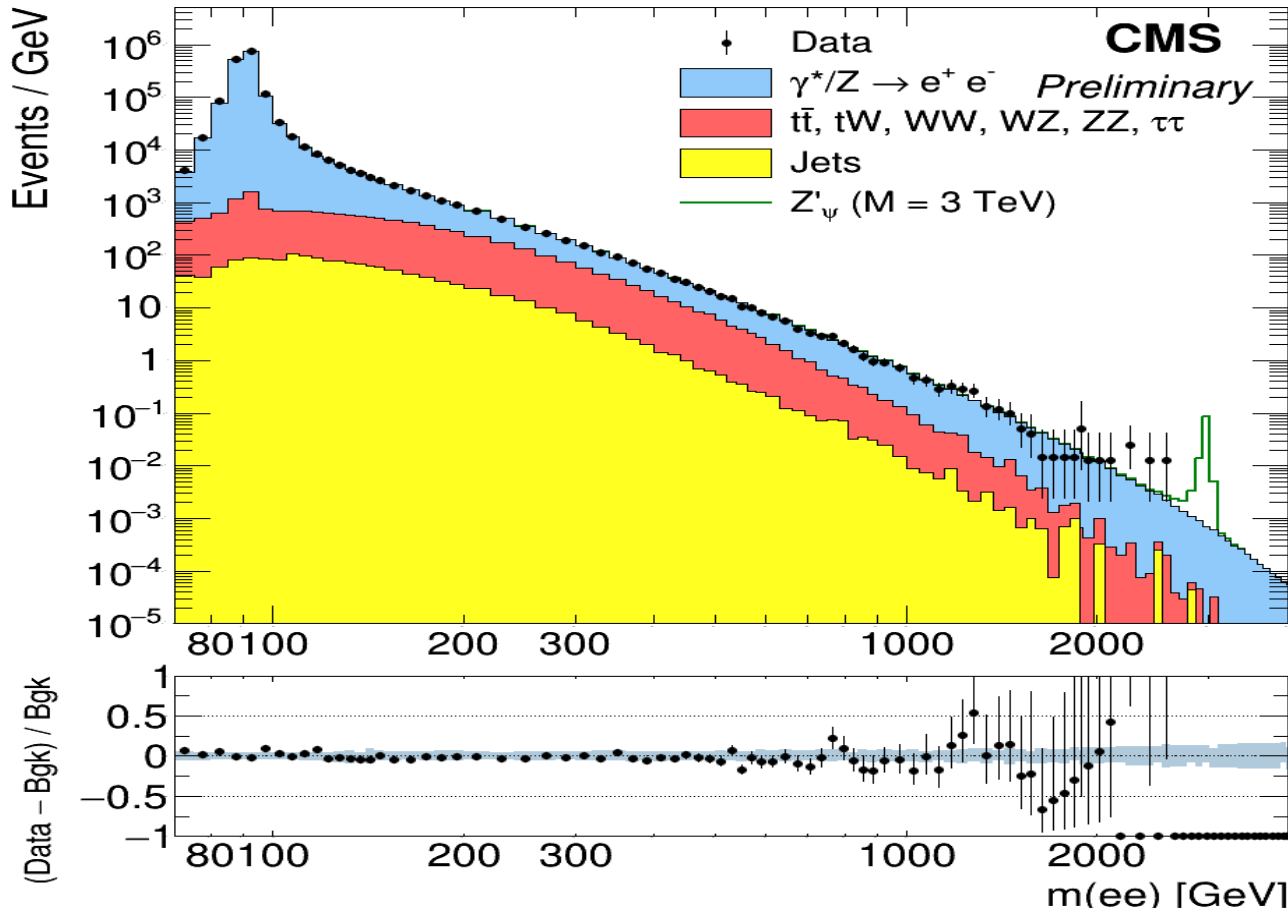
W + jets differential cross sections compared to QCD expectations, NLO, NNLO





Z' searches with dileptons, CMS, status in 2018

35.9 fb⁻¹ (13 TeV)



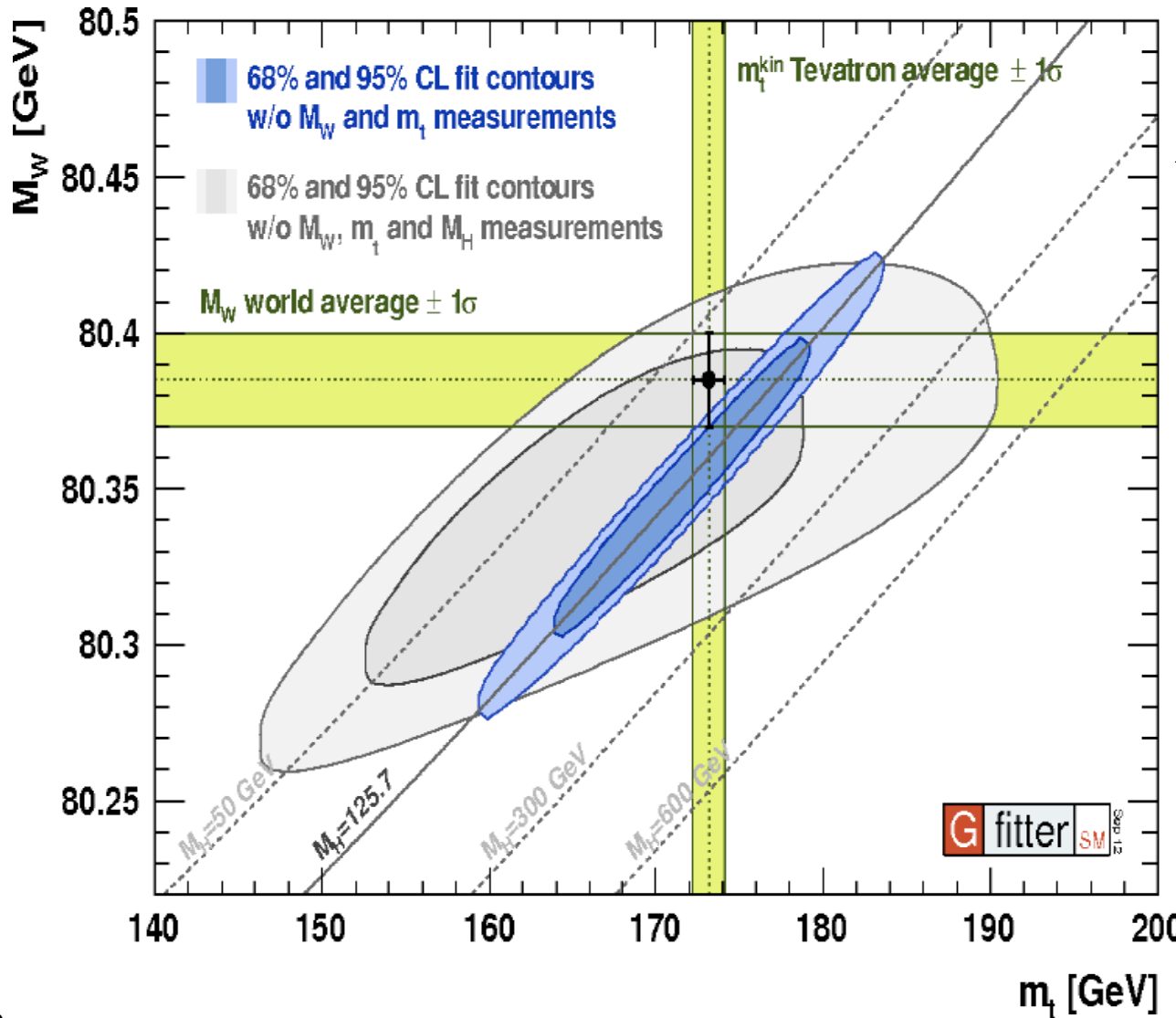
Limits extending beyond 4 TeV!

Channel	Model	Obs. limit (TeV)	Exp. limit (TeV)
ee (2017)	Z'_{SSM}	4.10	4.15
	Z'_{ψ}	3.35	3.55
ee (2016 and 2017) + $\mu\mu$ (2016)	Z'_{SSM}	4.7	4.7
	Z'_{ψ}	4.1	4.1



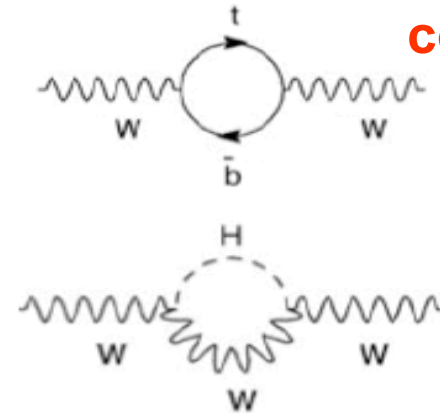
Testing the coherence of the Standard Model - importance of precision measurements of W, top and Higgs masses - all correlated through radiative corrections!

Higgs masses - all correlated through radiative corrections!



$$M_W^2 = \frac{\pi\alpha}{\sqrt{2}G_F} \frac{1}{\sin^2 \theta_W (1+\Delta r)}$$

← radiative corrections



$$(\Delta r)_{top} \approx \frac{3G_F m_t^2}{8\sqrt{2}\pi^2} \frac{1}{\tan^2 \theta_W}$$

$$(\Delta r)_{Higgs} \approx \frac{11G_F M_Z^2 \cos^2 \theta_W}{24\sqrt{2}\pi^2} \ln \frac{m_h^2}{M_Z^2}$$



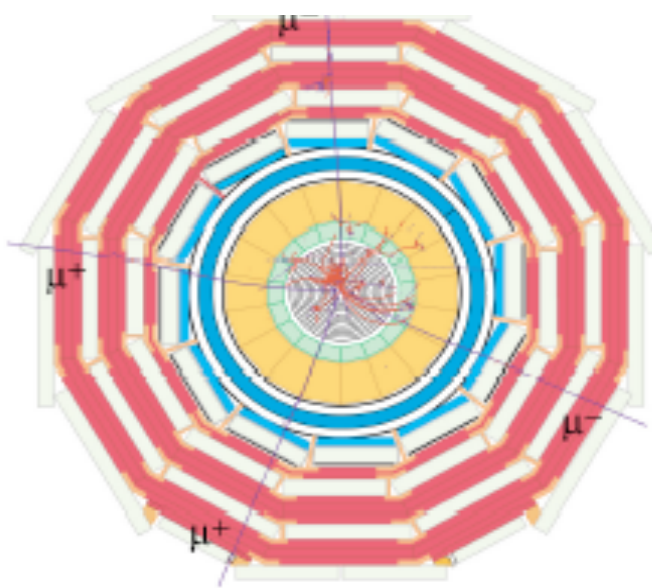
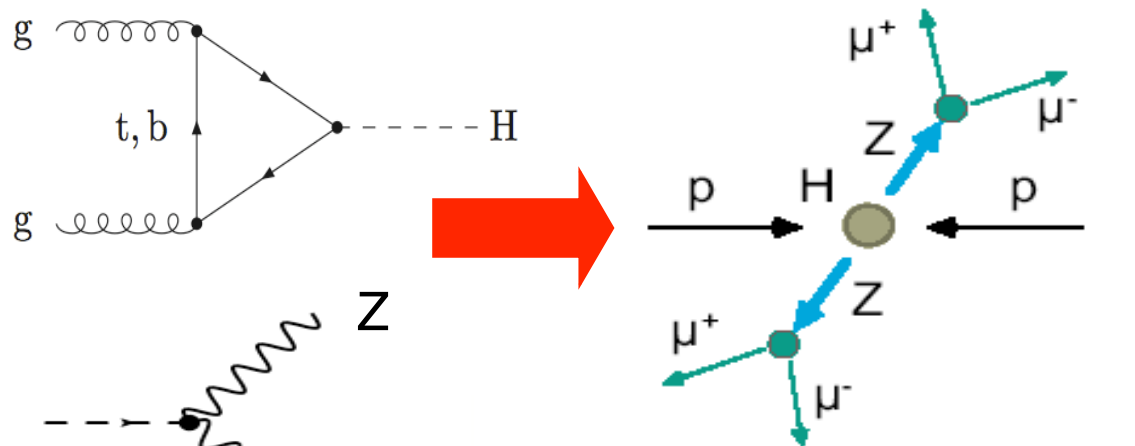
the Higgs.....



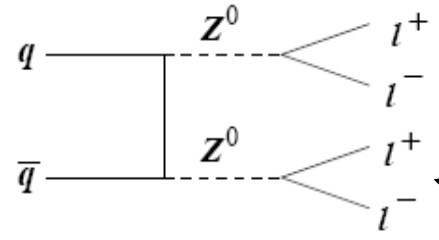
Production and detection of the Higgs in CMS – if $m_H \sim 150$ GeV ($H \rightarrow ZZ/ZZ^* \rightarrow 4$ leptons)

expectations from 1992!

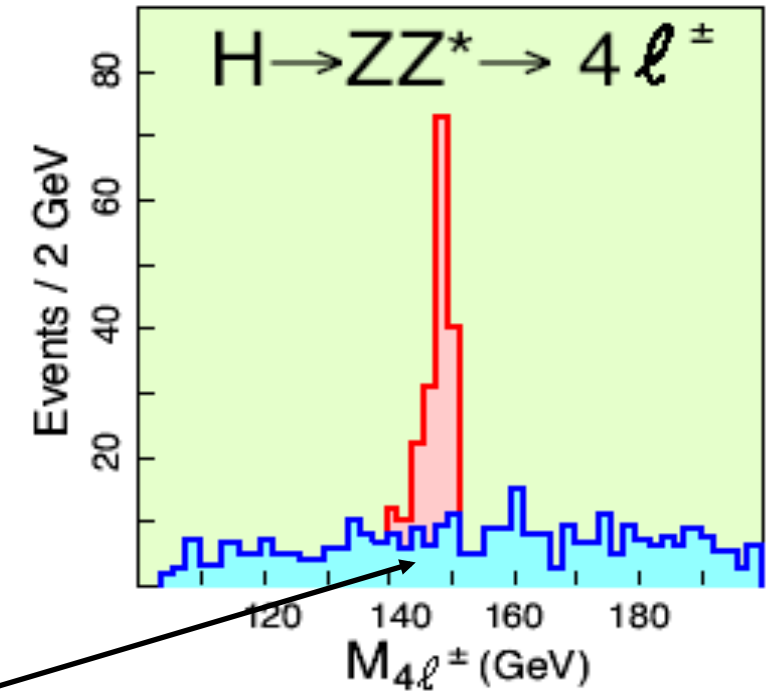
Expectations for signal and background if the Higgs has a mass ~ 150 GeV
Similar situation in the mass range $\sim 130 - \sim 400$ GeV



Electroweak ZZ background!!



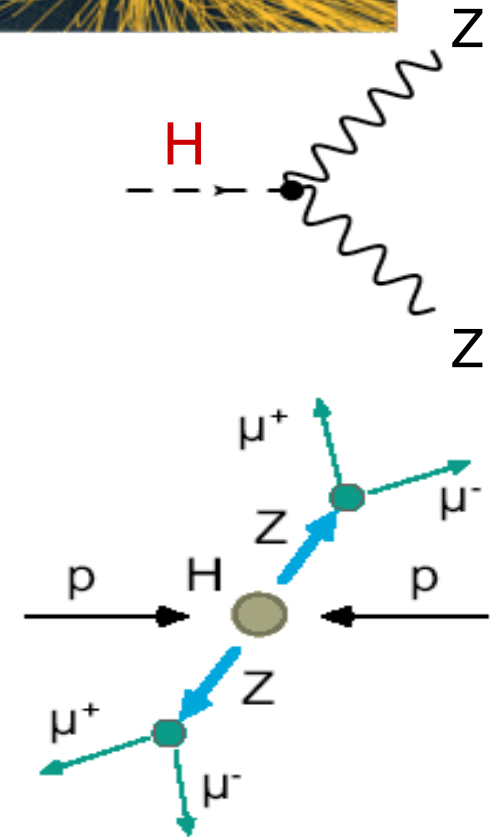
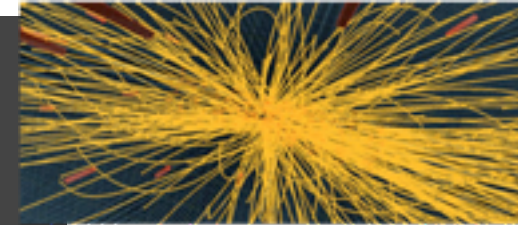
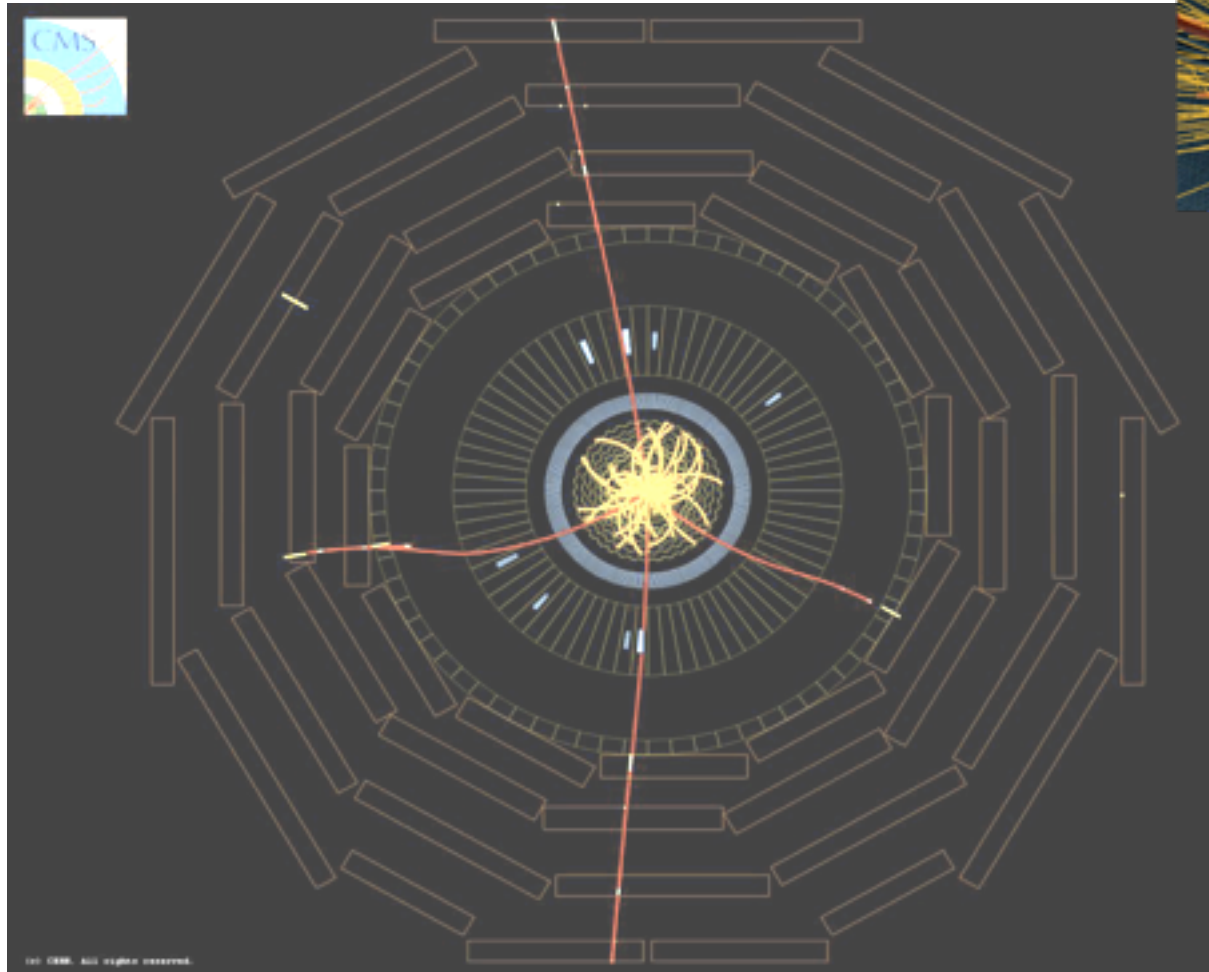
$M_{Higgs} = 150$ GeV



Electroweak ZZ background



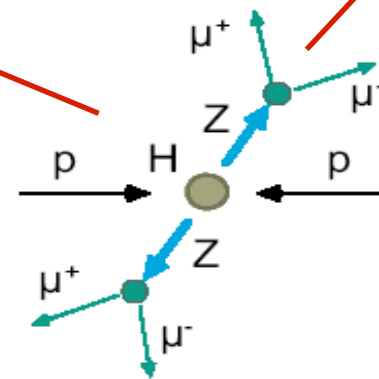
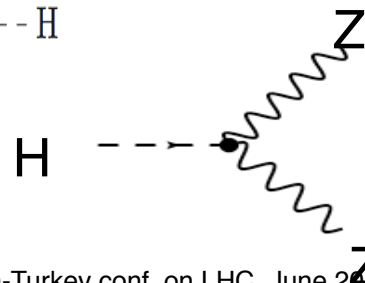
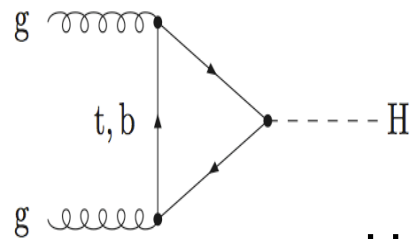
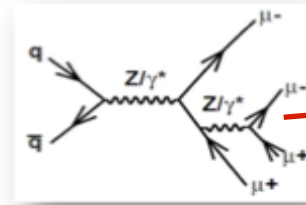
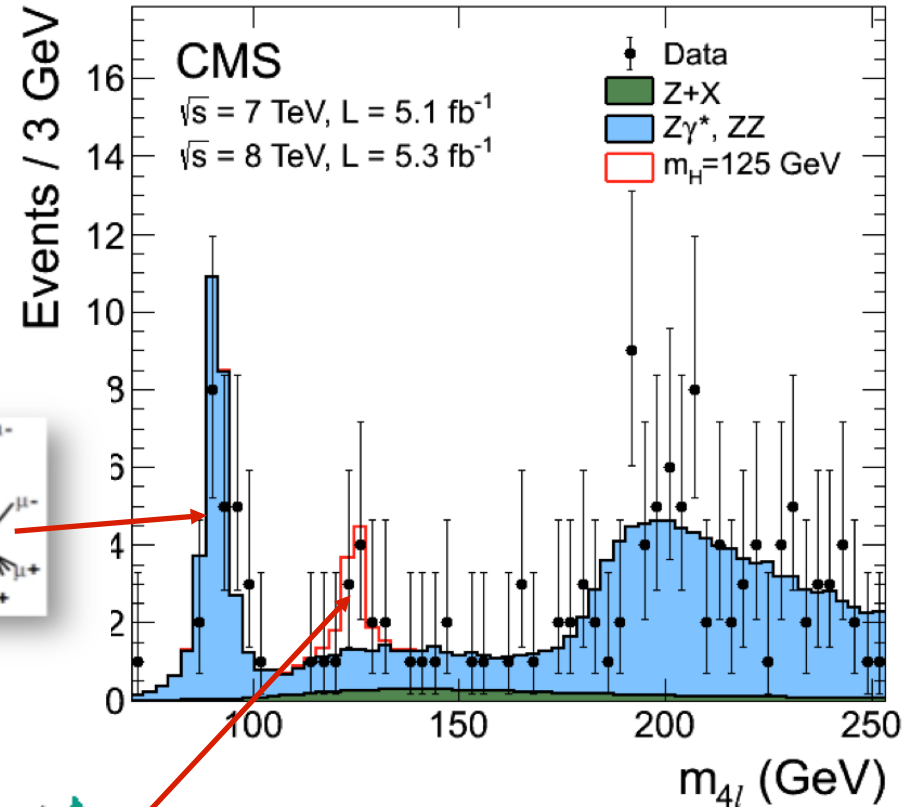
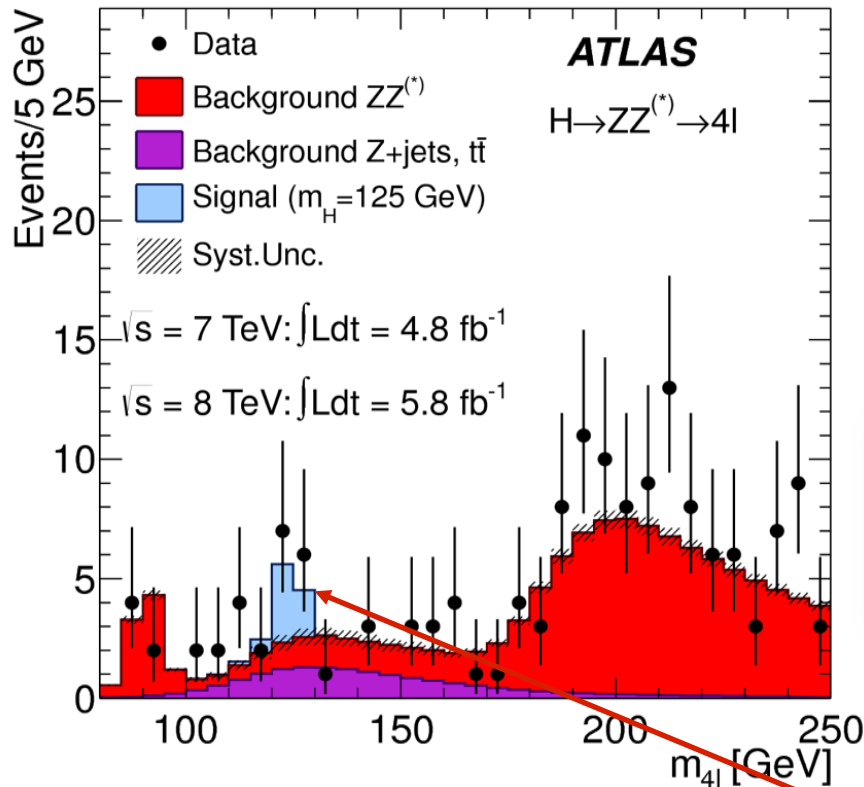
First events: $H \rightarrow ZZ \rightarrow \mu\mu\mu\mu$ candidate event in CMS, $\sqrt{s} = 8$ TeV data, 2012



Such an event with a Higgs boson decaying into 4 muons appears about once in 10^{15} proton-proton collisions in the LHC



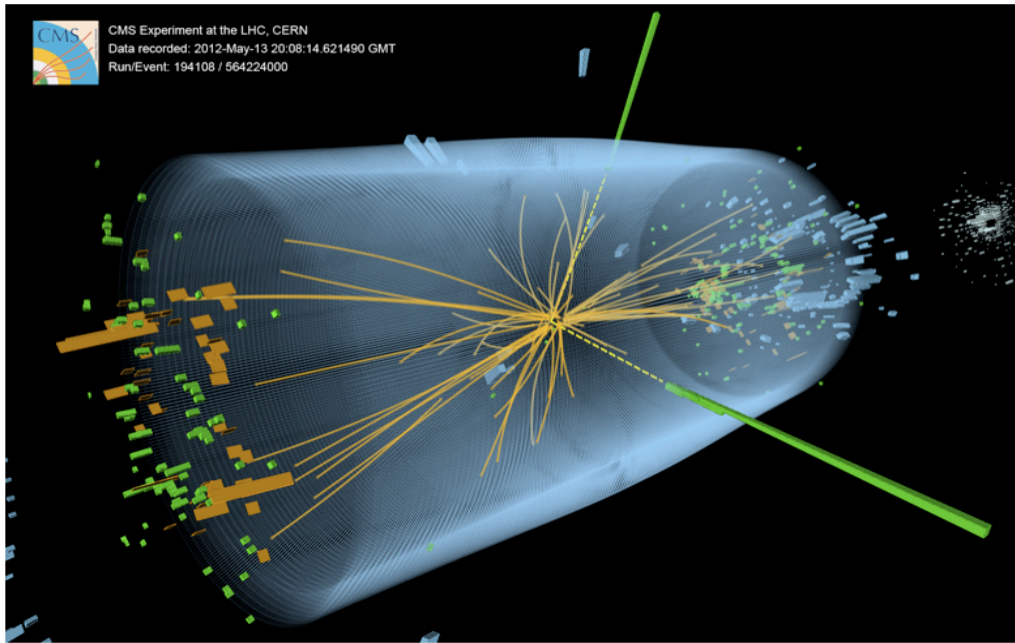
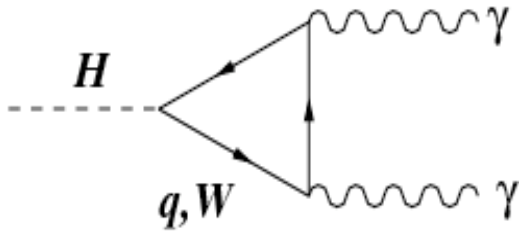
Higgs decay to 4 leptons, ATLAS and CMS, July 4th 2012



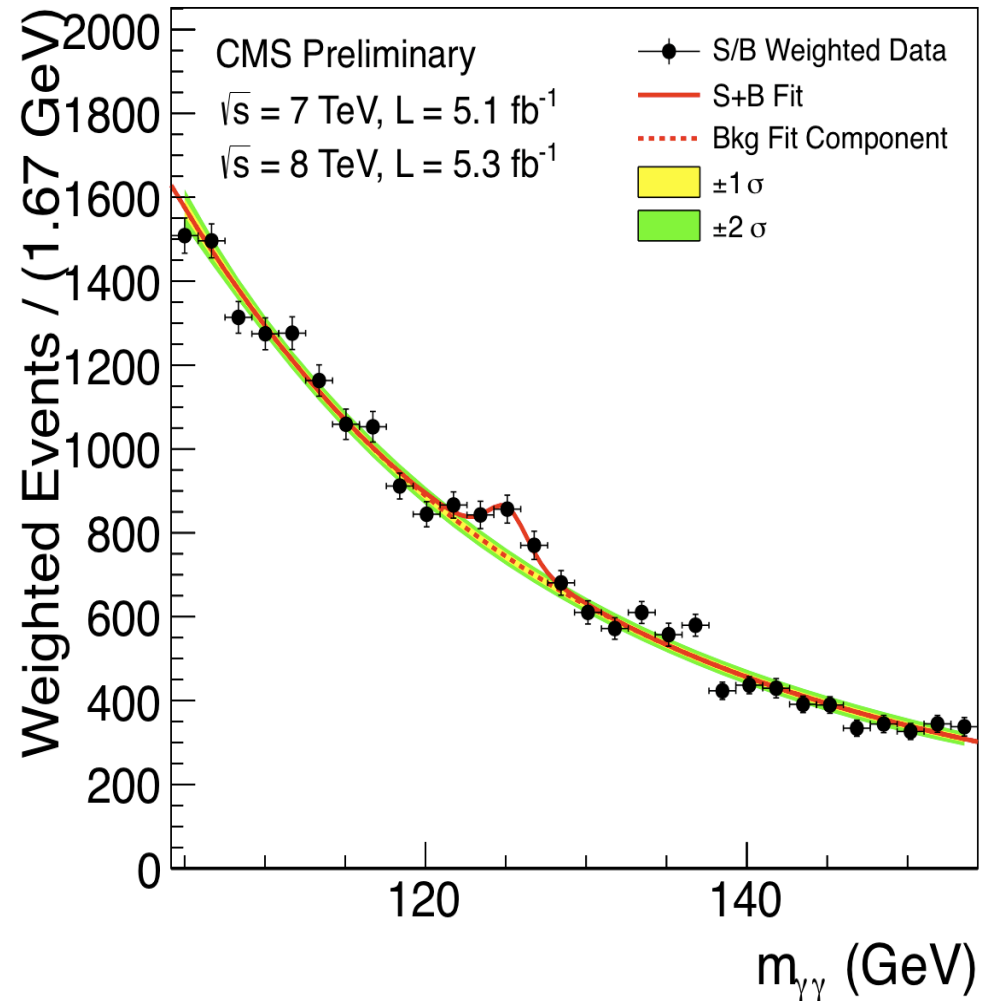
notice these modest beginnings! As for W, Z!



H \rightarrow $\gamma\gamma$ candidate in CMS, situation July 2012

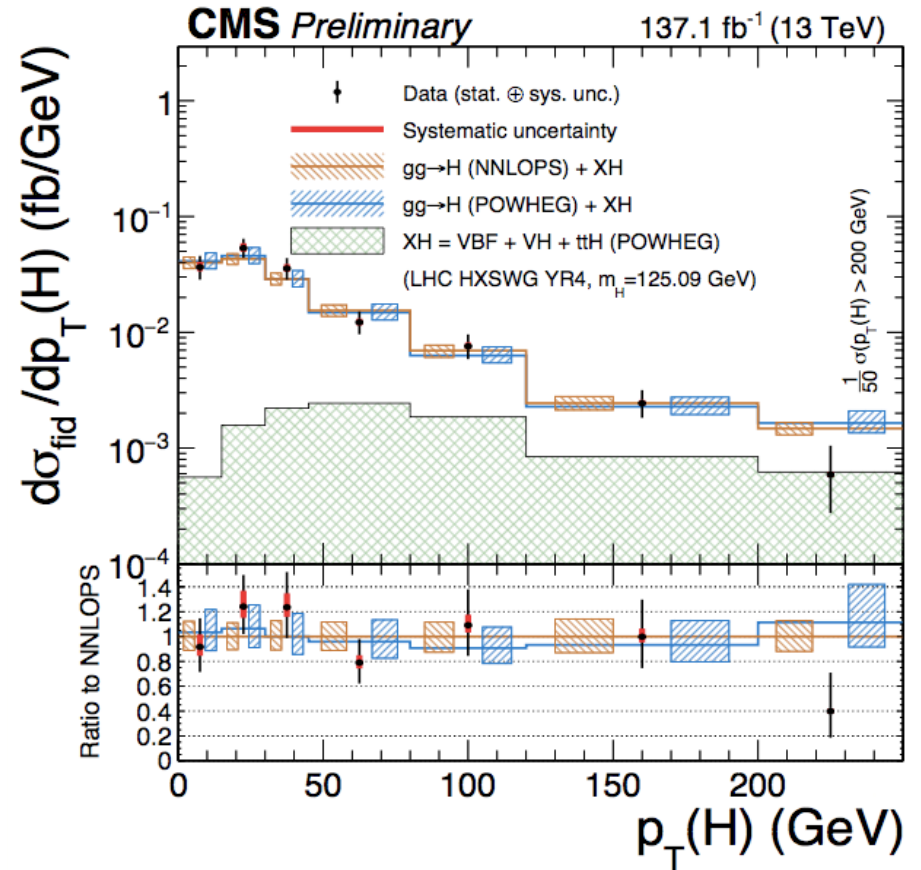
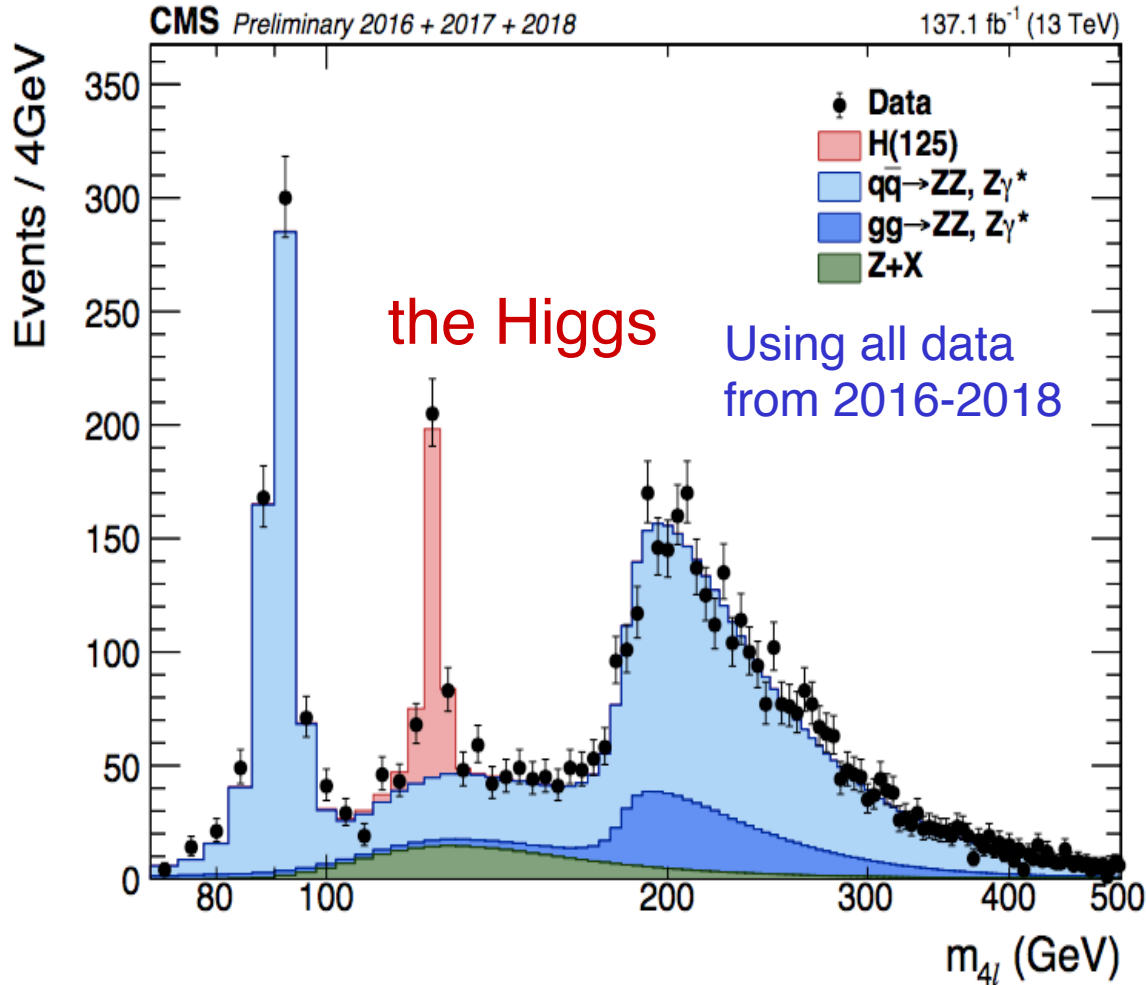


$$m_{\gamma\gamma}^2 = 2 E_1 E_2 (1 - \cos\theta_{12})$$





H → ZZ → 4 leptons, full 13TeV stat. April 2019

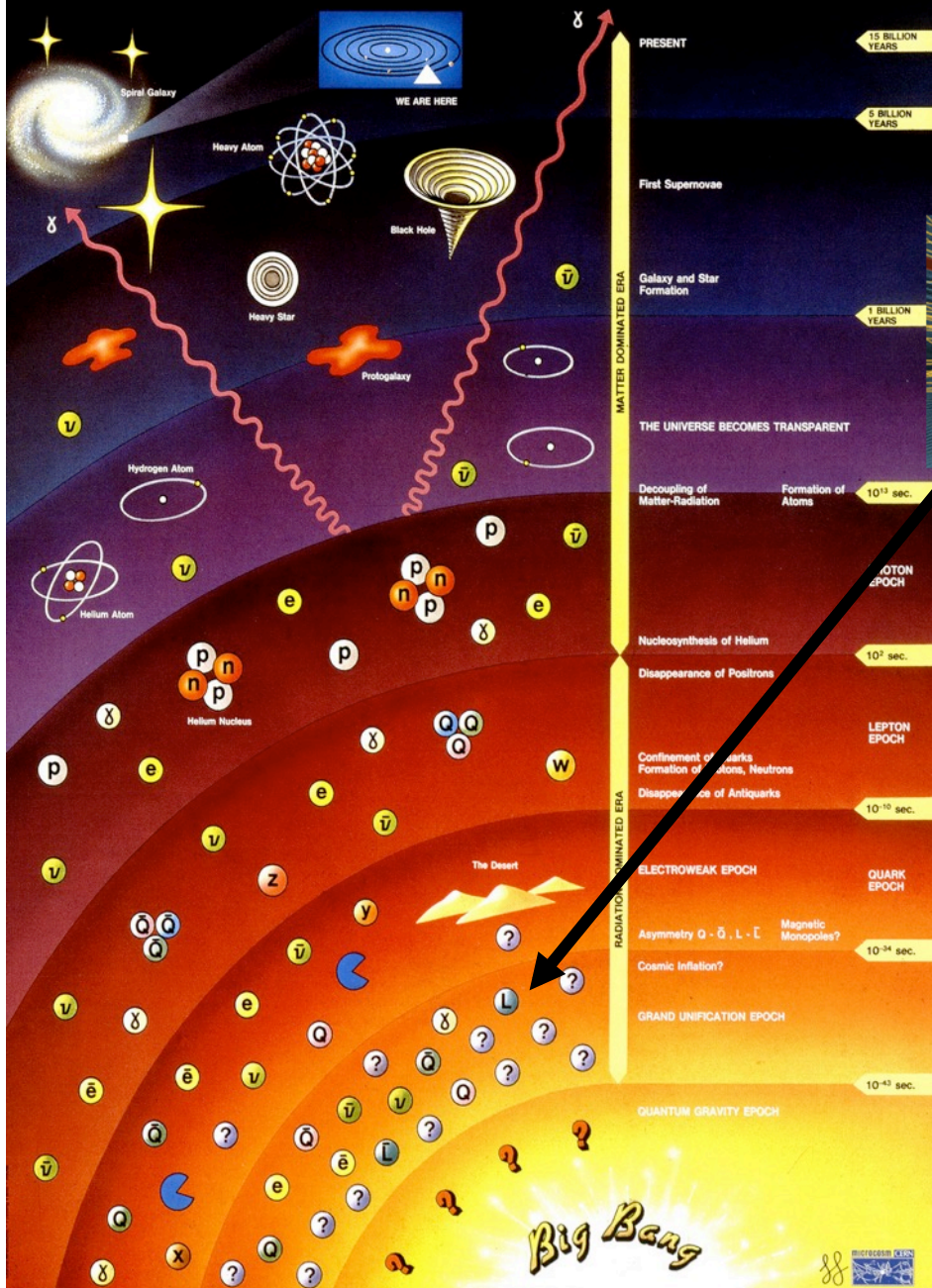


$$\sigma_{ZZ} = 17.1 \pm 0.3(\text{stat}) \pm 0.4(\text{syst}) \pm 0.4(\text{theo.}) \pm 0.3(\text{lumi}) \text{ pb}$$

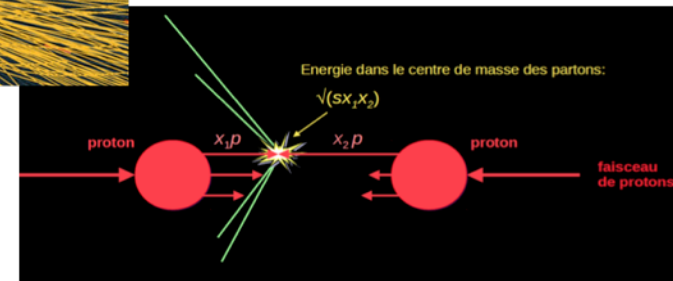
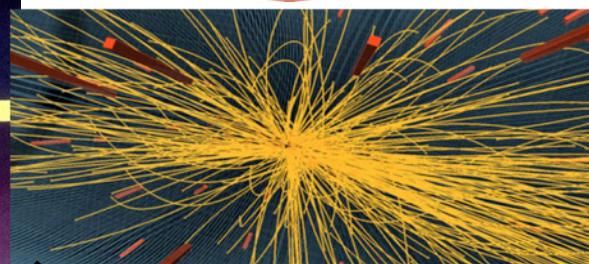
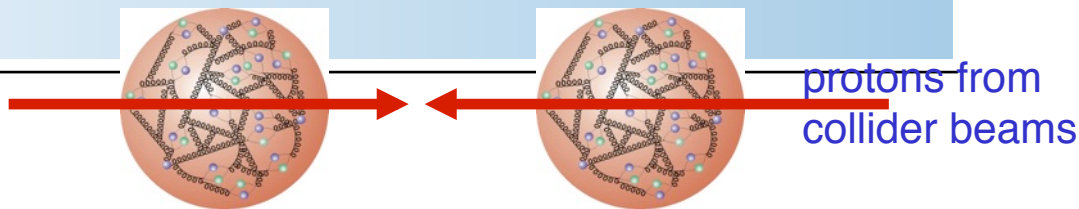
$$\mu = \sigma / \sigma_{SM} = 0.99^{+0.33}_{-0.26} \quad m_H = 124.50^{+0.48}_{-0.46} \text{ GeV} = 124.50^{+0.47}_{-0.45}(\text{stat.})^{+0.13}_{-0.11}(\text{sys.}) \text{ GeV}$$

June 2019

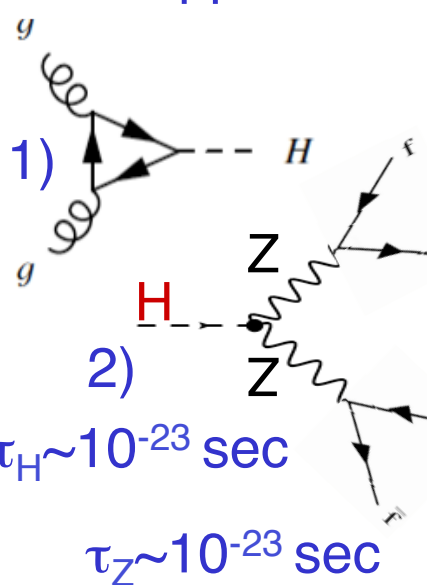
History of the Universe



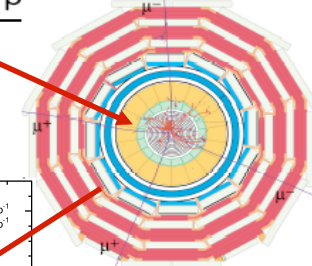
Mini Big-Bangs and the Higgs



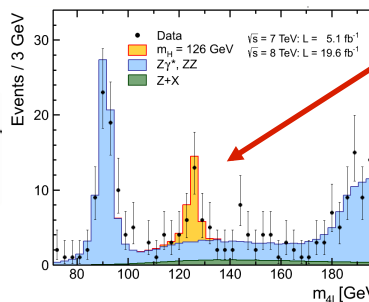
once in $\sim 10^{15}$ collisions this will happen:



after 20 years of construction



and after 3 years of collisions

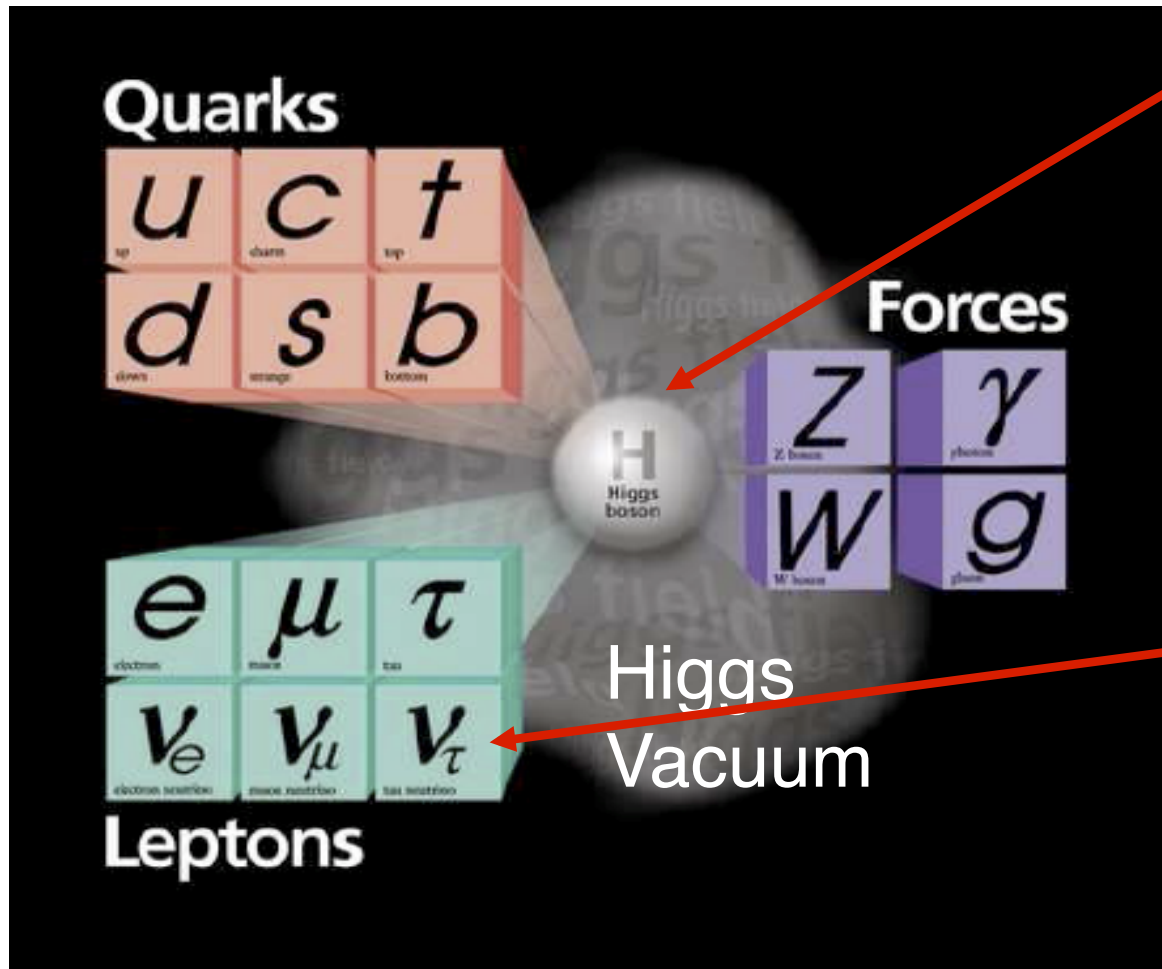




Particle content of the Universe - status studies and questions for the future

Matter

Intermediate bosons



- New Higgs sector
- Properties
 - Couplings
 - Symmetry breaking
 - Vacuum stability

- Neutrino sector
- Majorana/Dirac
 - Masses & mixings
 - CP violation
 - Lepton flavor violation

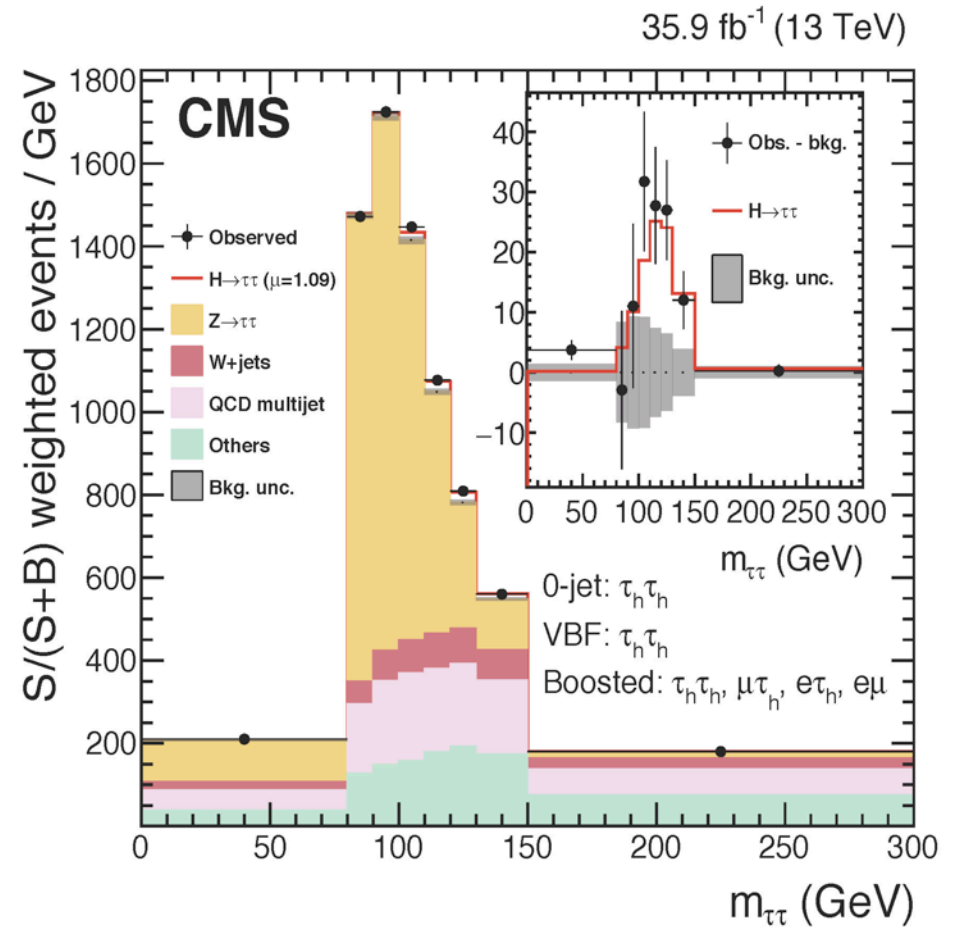
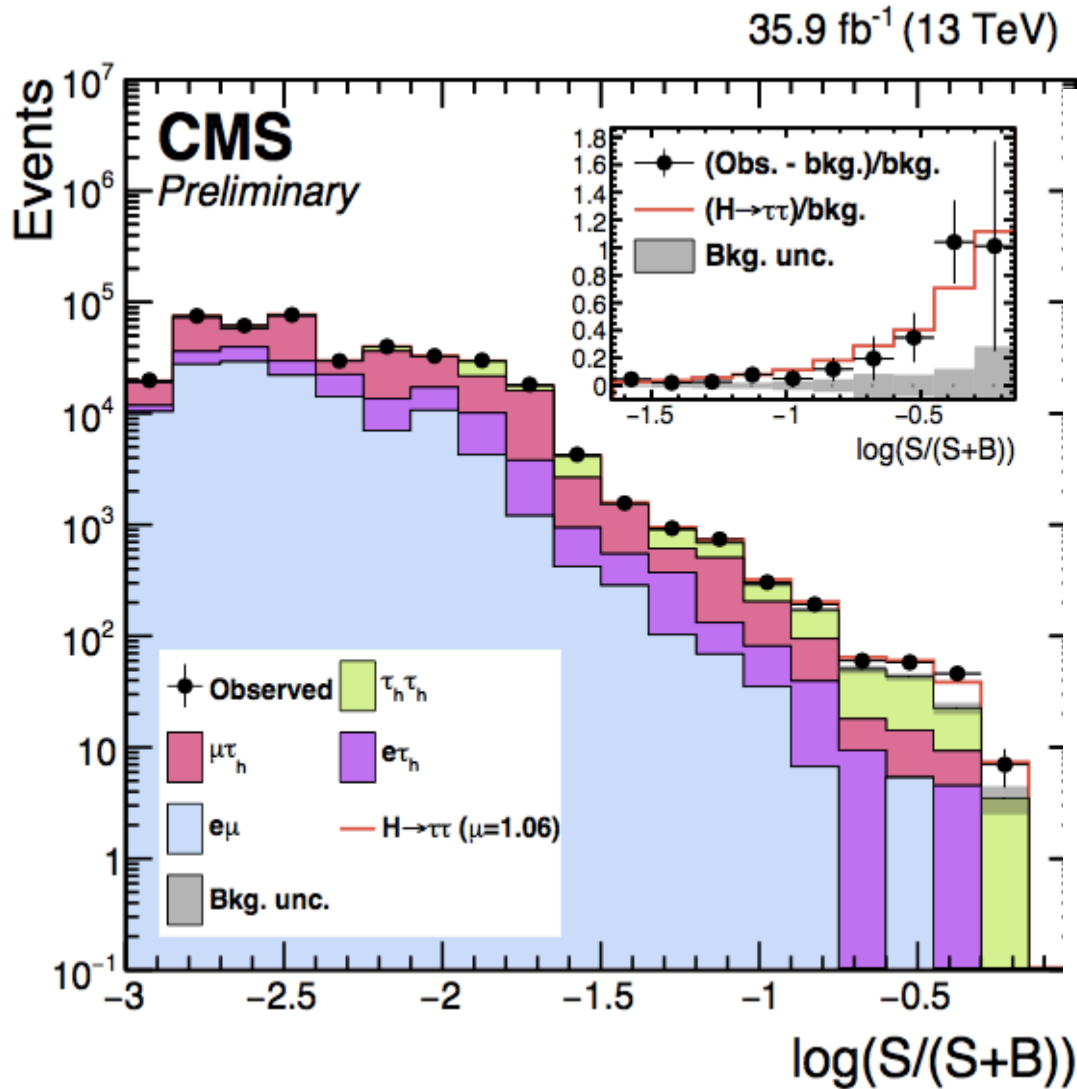


The World Lagrangian our entire knowledge of the material world in this concise formulation, except for gravity - requiring a different formulation (General Relativity)

$$\begin{aligned} \mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ & + i\bar{\psi} \not{D} \psi + h.c. \\ & + \chi_i Y_{ij} \chi_j \phi + h.c. \\ & + |D_m \phi|^2 - V(\phi) \end{aligned}$$



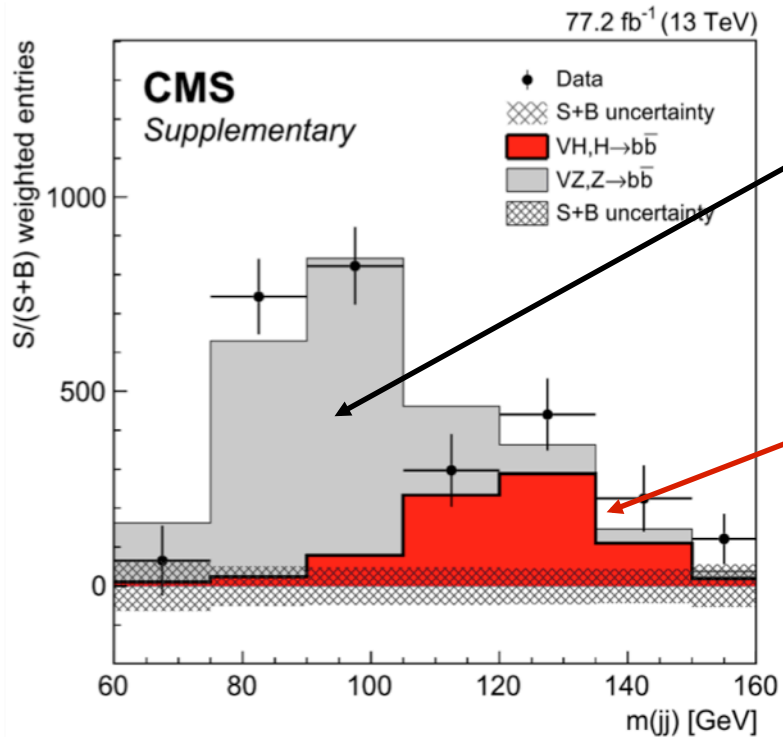
Higgs $\rightarrow \tau^+\tau^-$, CMS





Observation of $H \rightarrow bb$ decay

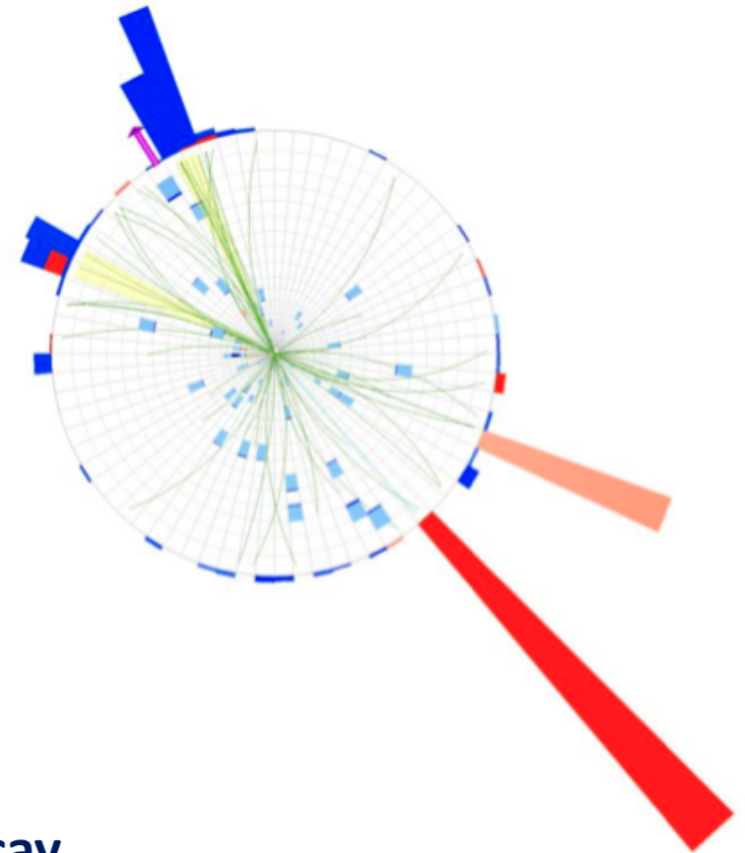
Candidate for HZ production followed by $Z \rightarrow ee$ and $H \rightarrow bb$



$ZZ \rightarrow ee + bb$

$ZH \rightarrow ee + bb$

$WH \rightarrow l\nu + bb$

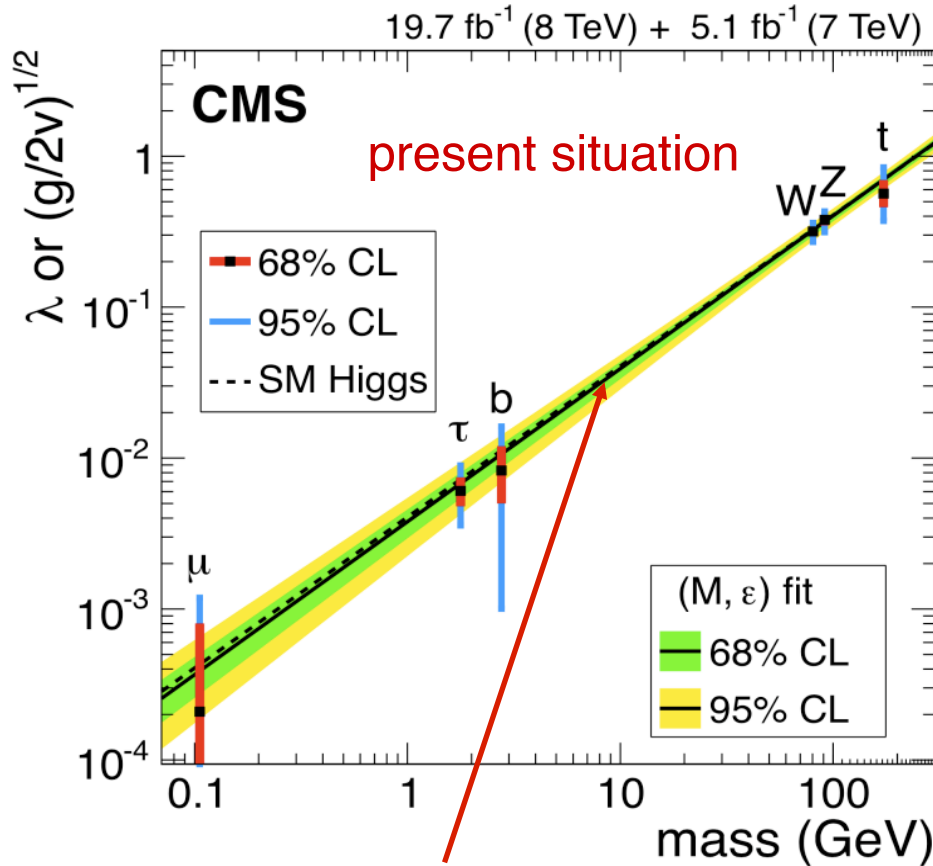


CMS has reached a 5.6σ observation of the $H \rightarrow bb$ decay, with signal strength $\mu = 1.04 \pm 0.20$

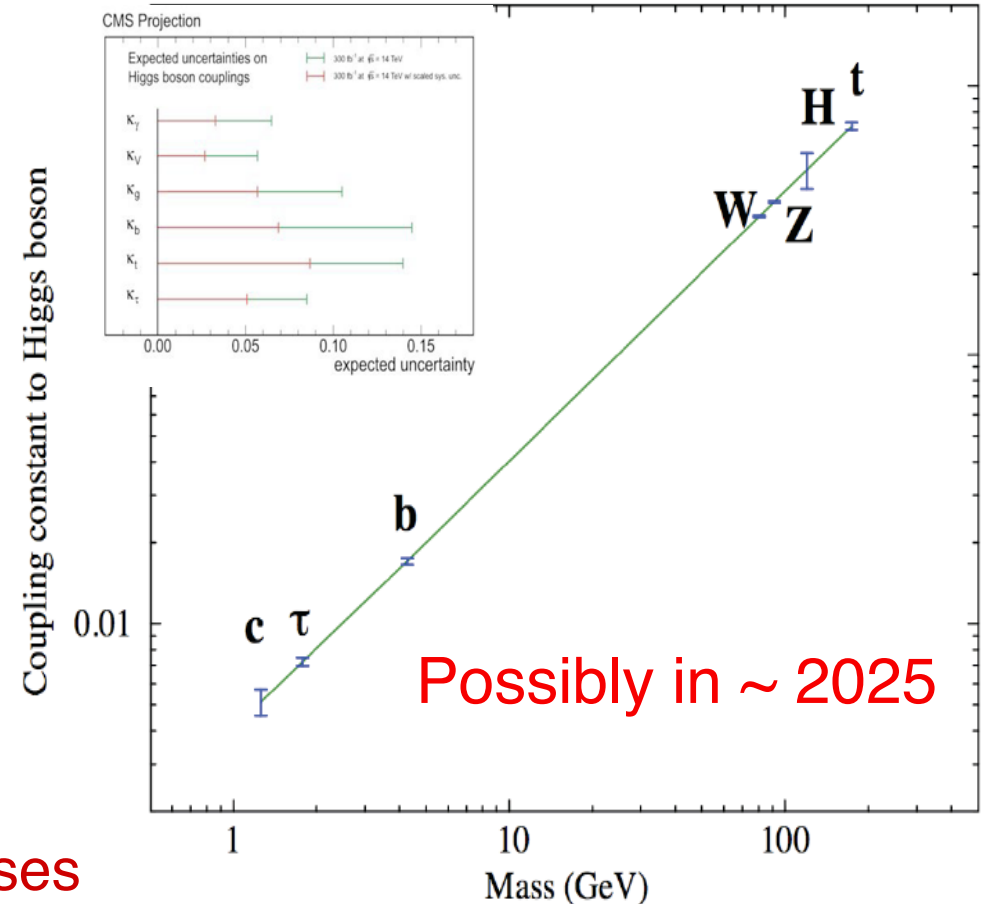
- Combination of several production channels, dominated by $VH(bb)$
- Result contained in HIG-18-016, arXiv:XXXX.XXXXX submitted to PRL



Higgs couplings, the key prediction of the theory, present situation and expectations for LHC by about 2025



Coupling Mass Relation



➔ Higgs couplings \propto particle masses

We need precise measurements of properties looking for small deviations to clarify/understand the exact nature of the observed Higgs boson; for ex. in the decoupling limit of the MSSM the h is very much SM-like! **Is it really elementary?**



Higgs couplings - why is precision needed....I

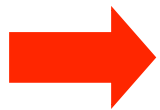
New physics affects the Higgs couplings:

SUSY $\frac{g_{hbb}}{g_{h_{SM}bb}} = \frac{g_{h\tau\tau}}{g_{h_{SM}\tau\tau}} \simeq 1 + 1.7\% \left(\frac{1 \text{ TeV}}{m_A}\right)^2$ for $\tan\beta = 5$

Composite Higgs $\frac{g_{hff}}{g_{h_{SM}ff}} \simeq \frac{g_{hVV}}{g_{h_{SM}VV}} \simeq 1 - 3\% \left(\frac{1 \text{ TeV}}{f}\right)^2$

Top partners $\frac{g_{hgg}}{g_{h_{SM}gg}} \simeq 1 + 2.9\% \left(\frac{1 \text{ TeV}}{m_T}\right)^2$, $\frac{g_{h\gamma\gamma}}{g_{h_{SM}\gamma\gamma}} \simeq 1 - 0.8\% \left(\frac{1 \text{ TeV}}{m_T}\right)^2$

Other models may give up to 5% deviations with respect to the Standard Model

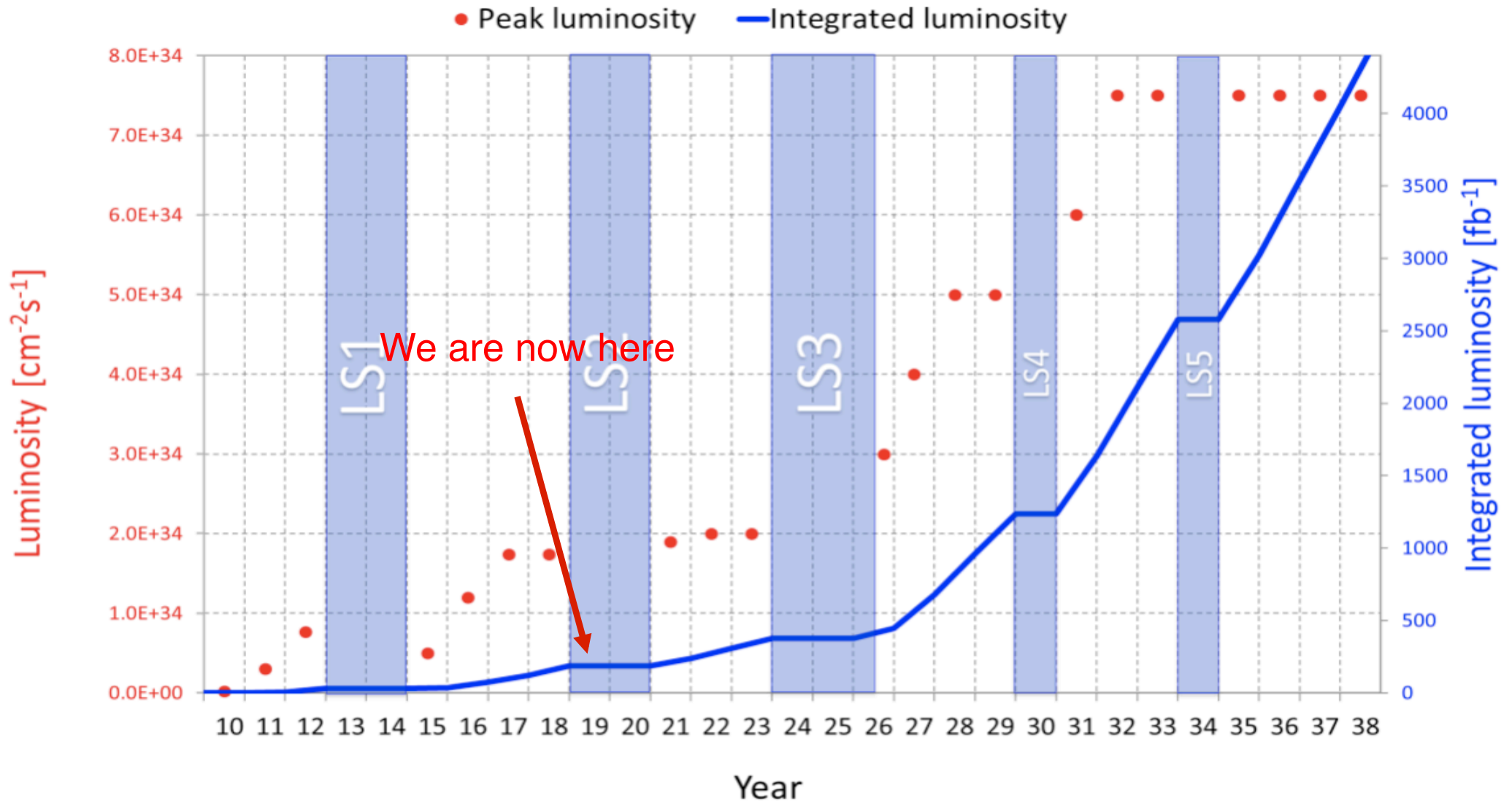


Sensitivity to “TeV” new physics needs per-cent to sub-per-cent accuracy on couplings for 5 sigma evidence/discovery



Evolution of LHC luminosity till 2038

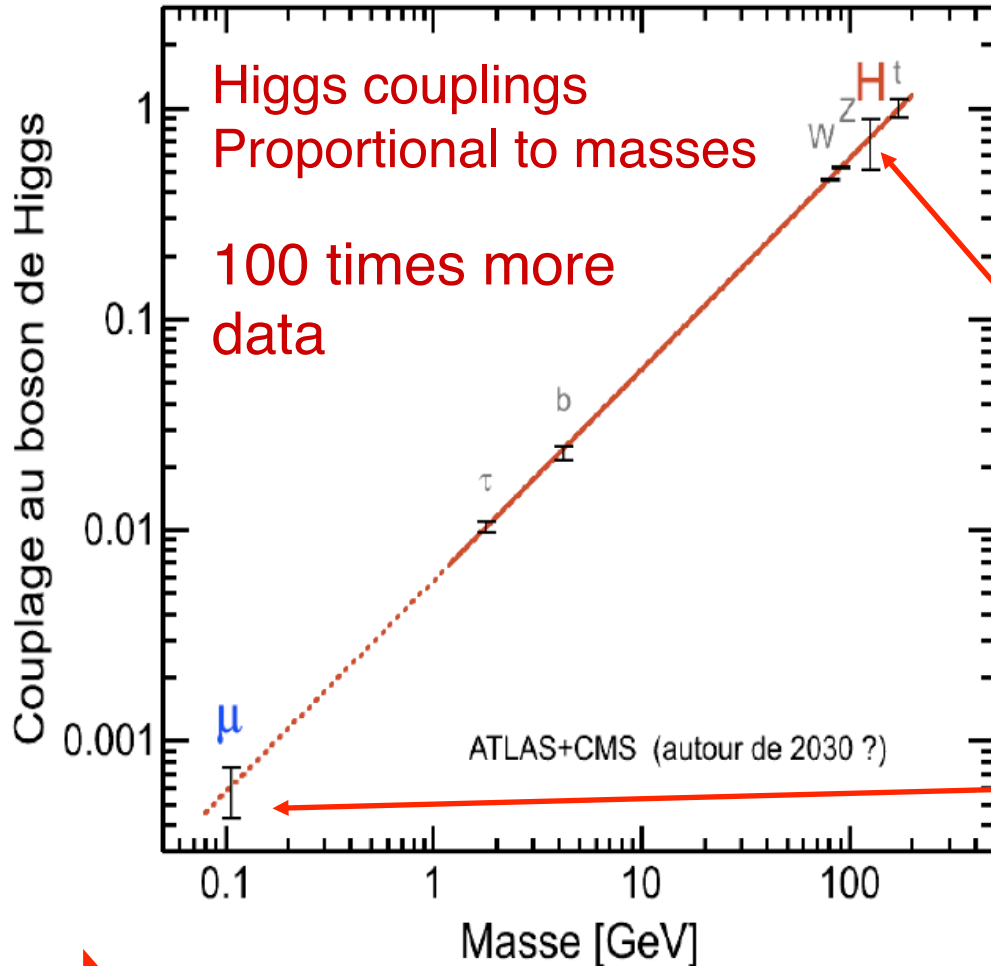
- october 2018 HL-LHC workshop -





Higgs couplings - future up to HL-LHC and 3000fb⁻¹

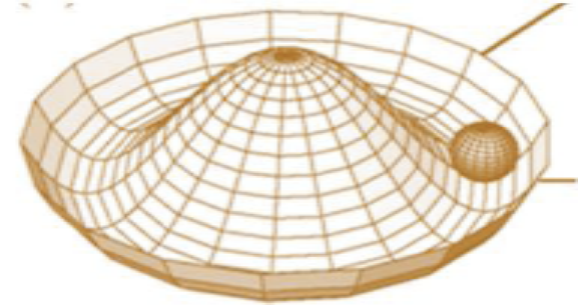
Relation Couplage-Masse



$$g_{Hff} = m_f/v$$

$$g_{HHH} = 3m_H^2/v$$

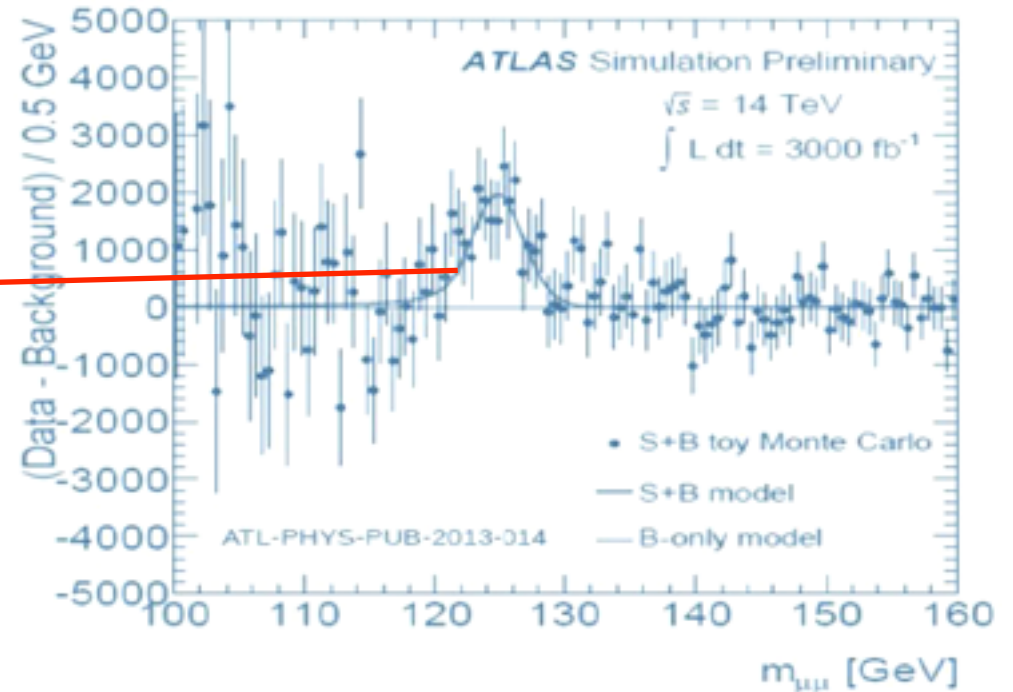
$$g_{HVV} = 2m_V^2/v$$



$$v = (-\mu^2/\lambda)^{1/2} = 246$$

Higgs self-coupling and Higgs potential

➔ With 3000fb⁻¹ data at 13 TeV (by ~ 2025) couplings to 8 particles, including coupling to muons

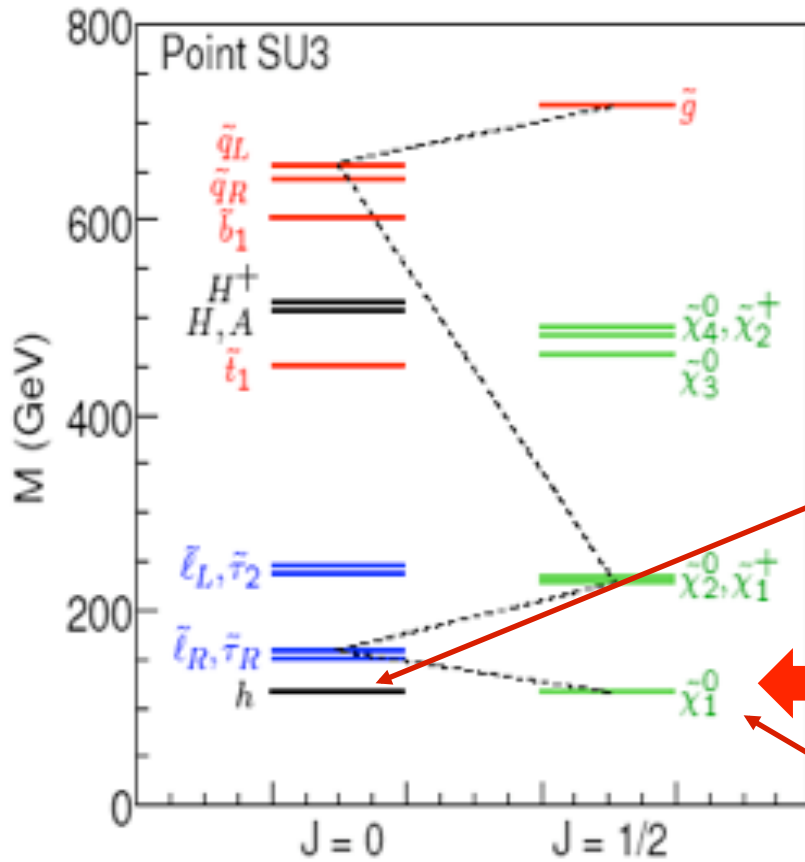




Looking further for physics beyond the Standard Model

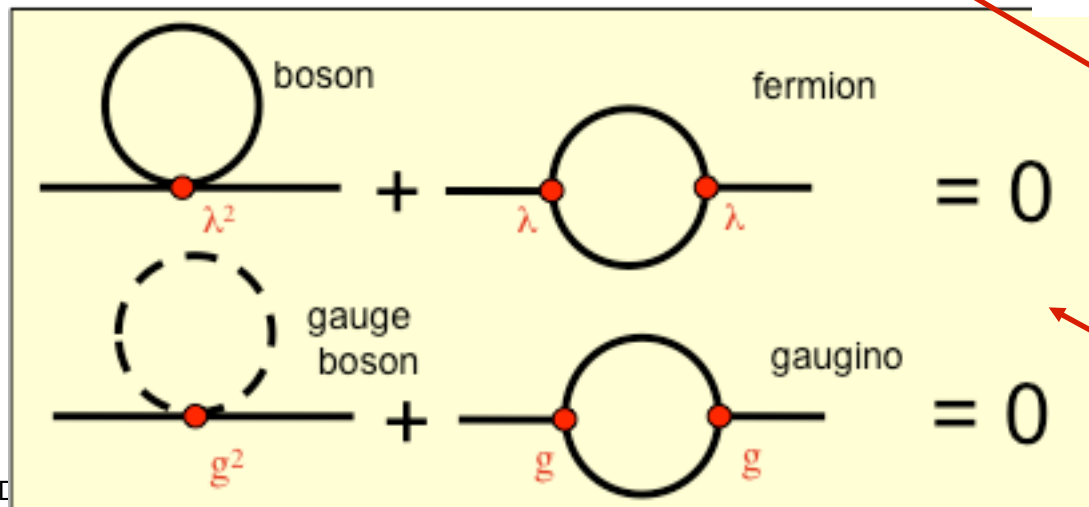
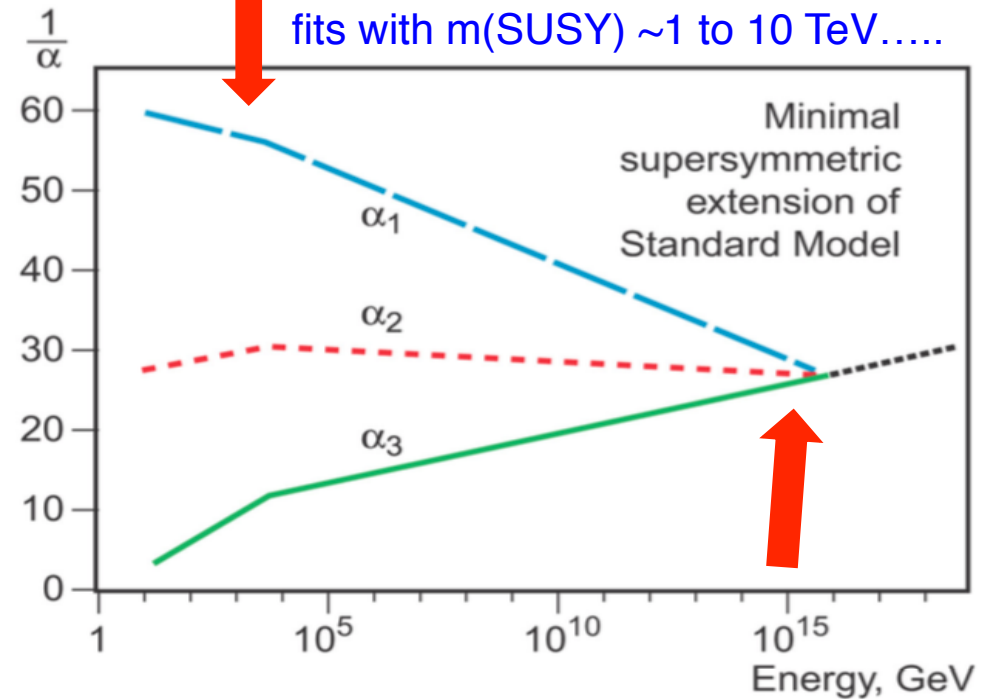
- SUSY
- DM.....

SUSY - could solve a number of problems at ew scale.....



Present Higgs?

LSP



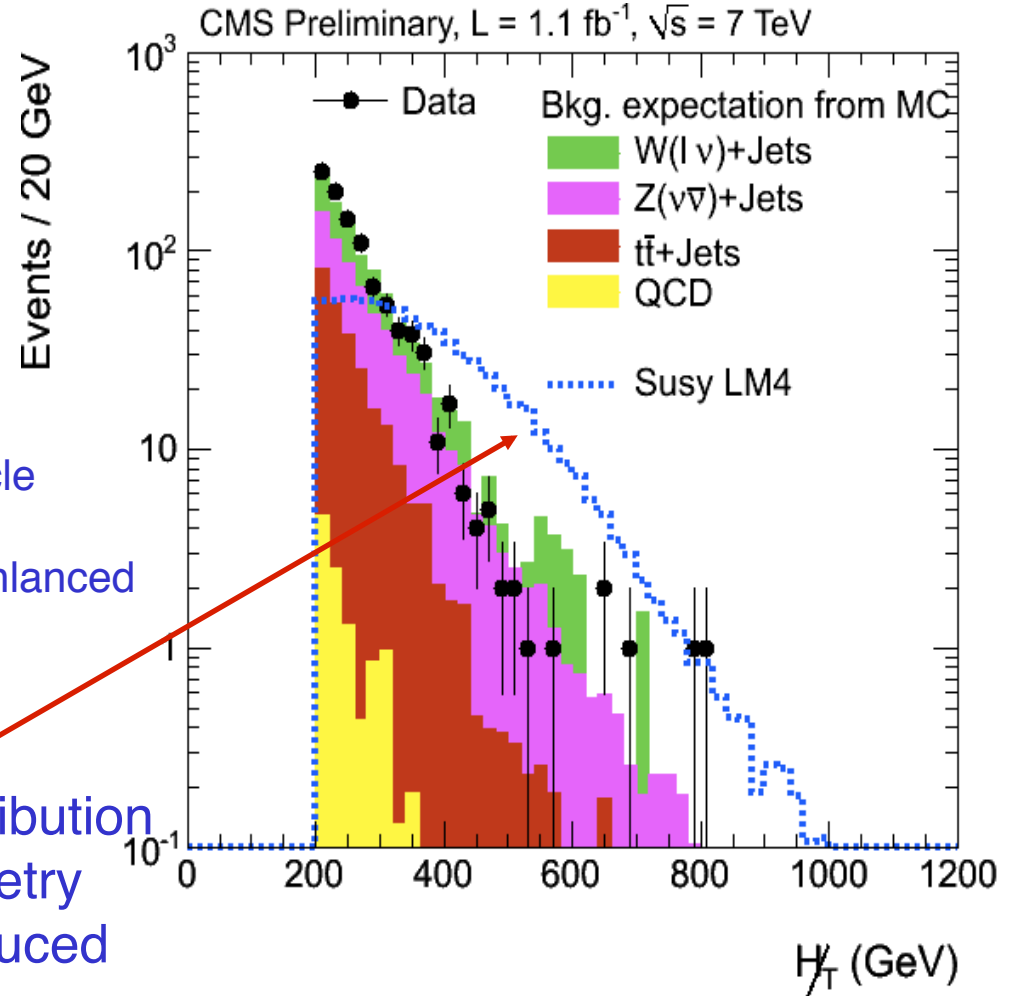
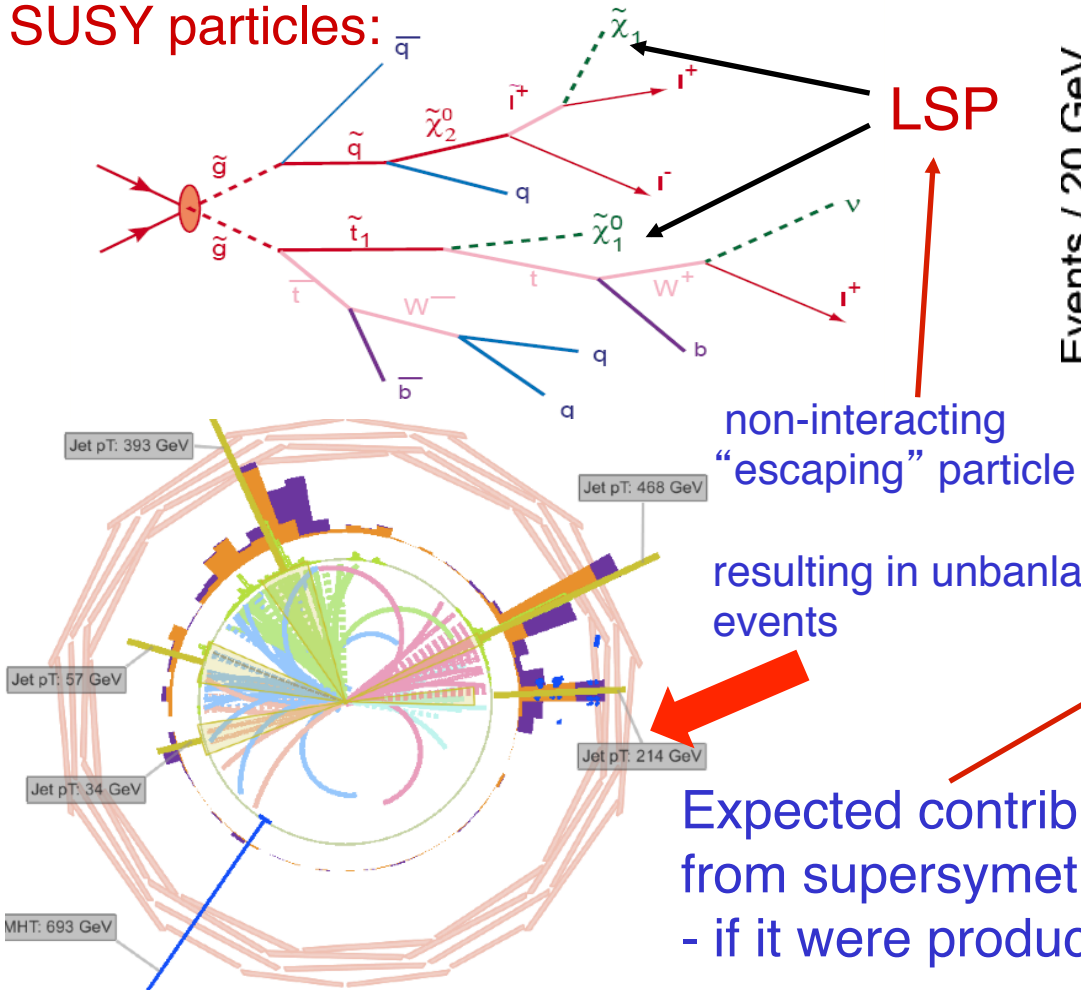
Possibly the DM particle....

Stabilizes Higgs boson mass



DM – could SUSY and the LHC bring the solution? SUSY search in jets + missing- E_T channel

Proton-proton collision at the LHC producing a pair of SUSY particles:

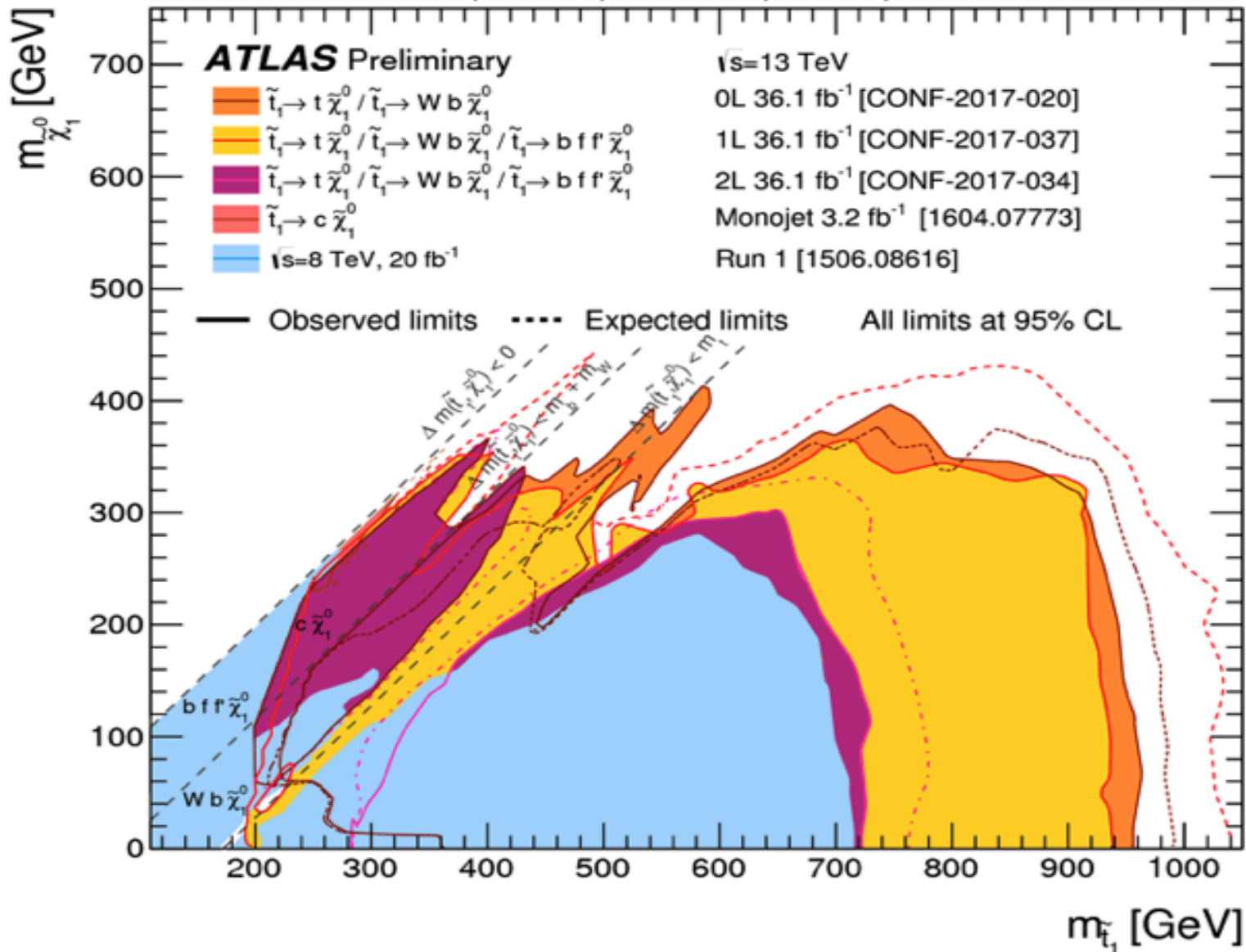


Nothing has yet been observed at the LHC....
insufficient statistics or energy?



Stop searches, ATLAS

$\tilde{t}_1\tilde{t}_1$ production, $\tilde{t}_1 \rightarrow b f \tilde{\chi}_1^0$ / $\tilde{t}_1 \rightarrow c \tilde{\chi}_1^0$ / $\tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0$ / $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$ Status: May 2017



$m(\tilde{t}_1) > 950$ GeV

ATLAS SUSY Searches* - 95% CL Lower Limits

May 2017

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [fb^{-1}]$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 e, μ /1-2 τ	2-10 jets/3 b	Yes	20.3	\tilde{g}, \tilde{g}	1.85 TeV	$m(\tilde{g})=m(\tilde{g})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{q}	1.57 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(1^{st} \text{ gen. } \tilde{q})=m(2^{nd} \text{ gen. } \tilde{q})$	ATLAS-CONF-2017-022
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	3.2	\tilde{q}	608 GeV	$m(\tilde{g})-m(\tilde{\chi}_1^0) < 5$ GeV	1604.07773
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.02 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.01 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}_2^0)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0$	3 e, μ	4 jets	-	36.1	\tilde{g}	1.825 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2017-030
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0$	0	7-11 jets	Yes	36.1	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2017-033
	GMSB ($\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	\tilde{g}	2.0 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	1607.05979
	GGM (bino NLSP)	2 γ	-	Yes	3.2	\tilde{g}	1.65 TeV	$c\tau(\text{NLSP}) < 0.1$ mm	1606.09150
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 950$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu < 0$	1507.05493
GGM (higgsino-bino NLSP)	γ	2 jets	Yes	13.3	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) > 680$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu > 0$	ATLAS-CONF-2016-066	
GGM (higgsino NLSP)	2 e, μ (Z)	2 jets	Yes	20.3	\tilde{g}	900 GeV	$m(\text{NLSP}) > 430$ GeV	1503.03290	
Gravitino LSP	0	mono-jet	Yes	20.3	$F^{1/2}$ scale	865 GeV	$m(\tilde{G}) > 1.8 \times 10^{-4}$ eV, $m(\tilde{g})=m(\tilde{g})=1.5$ TeV	1502.01518	
3 rd gen. \tilde{g} med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	36.1	\tilde{g}	1.92 TeV	$m(\tilde{\chi}_1^0) < 600$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	36.1	\tilde{g}	1.97 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 300$ GeV	1407.0600
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	36.1	\tilde{b}_1	950 GeV	$m(\tilde{\chi}_1^0) < 420$ GeV	ATLAS-CONF-2017-038
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^0$	2 e, μ (SS)	1 b	Yes	36.1	\tilde{b}_1	275-700 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^\pm) + 100$ GeV	ATLAS-CONF-2017-030
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^0$	0-2 e, μ	1-2 b	Yes	4.7/13.3	\tilde{t}_1	117-170 GeV	$m(\tilde{\chi}_1^\pm) = 2m(\tilde{\chi}_1^0)$, $m(\tilde{\chi}_1^0) = 55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $\tilde{\chi}_1^0$	0-2 e, μ	0-2 jets/1-2 b	Yes	20.3/36.1	\tilde{t}_1	90-198 GeV	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, ATLAS-CONF-2017-020
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	3.2	\tilde{t}_1	90-323 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0) = 5$ GeV	1604.07773
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-600 GeV	$m(\tilde{\chi}_1^0) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	36.1	\tilde{t}_2	290-790 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	ATLAS-CONF-2017-019
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 e, μ	4 b	Yes	36.1	\tilde{t}_2	320-880 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	ATLAS-CONF-2017-019
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	36.1	$\tilde{\ell}$	90-440 GeV	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\tilde{\ell}\bar{\nu})$	2 e, μ	0	Yes	36.1	$\tilde{\chi}_1^\pm$	710 GeV	$m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}\nu(\tilde{\tau}\bar{\nu})$	2 τ	-	Yes	36.1	$\tilde{\chi}_1^\pm$	760 GeV	$m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_1, \nu\tilde{\ell}_1, \ell(\tilde{\nu}\nu), \ell\tilde{\nu}\tilde{\ell}_1, \ell(\tilde{\nu}\nu)$	3 e, μ	0	Yes	36.1	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	1.16 TeV	$m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^\pm Z\tilde{\chi}_1^0$	2-3 e, μ	0-2 jets	Yes	36.1	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	580 GeV	$m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^\pm h, \tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	270 GeV	$m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_R\ell$	4 e, μ	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	635 GeV	$m(\tilde{\chi}_2^0) = m(\tilde{\chi}_3^0)$, $m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	1 $e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	115-370 GeV	$c\tau < 1$ mm	1507.05493
	GGM (bino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	2 γ	-	Yes	20.3	\tilde{W}	590 GeV	$c\tau < 1$ mm	1507.05493
	Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	36.1	$\tilde{\chi}_1^\pm$	430 GeV	$m(\tilde{\chi}_1^\pm) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm) = 0.2$ ns
Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$		dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^\pm$	495 GeV	$m(\tilde{\chi}_1^\pm) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm) < 15$ ns	1506.05332
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9	\tilde{g}	850 GeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $10 \mu\text{s} < \tau(\tilde{g}) < 1000$ s	1310.6584
Stable \tilde{g} R-hadron		trk	-	-	3.2	\tilde{g}	1.58 TeV	-	1606.05129
Metastable \tilde{g} R-hadron		dE/dx trk	-	-	3.2	\tilde{g}	1.57 TeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $\tau > 10$ ns	1604.04520
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{\ell}, \tilde{\mu}) + \tau(e, \mu)$		1-2 μ	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$	1411.6795
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3	$\tilde{\chi}_1^0$	440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ee/\mu\mu/\mu\mu\nu$		displ. $ee/\mu\mu/\mu\mu\nu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g}) = 1.3$ TeV	1504.05162
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$		displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480$ mm, $m(\tilde{g}) = 1.1$ TeV	1504.05162
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau/\mu\tau$	$e\mu, e\tau, \mu\tau$	-	-	3.2	$\tilde{\nu}_\tau$	1.9 TeV	$\lambda_{311} = 0.11, \lambda_{132/133/233} = 0.07$	1607.08079
	Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{g}) = m(\tilde{g}), c\tau_{LSP} < 1$ mm	1404.2500
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^\pm\tilde{\chi}_1^0 \rightarrow ee\nu, e\mu\nu, \mu\mu\nu$	4 e, μ	-	Yes	13.3	$\tilde{\chi}_1^\pm$	1.14 TeV	$m(\tilde{\chi}_1^0) > 400$ GeV, $\lambda_{12k} \neq 0$ ($k = 1, 2$)	ATLAS-CONF-2016-075
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^\pm\tilde{\chi}_1^0 \rightarrow \tau\tau\nu_e, e\tau\nu_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.08 TeV	$BR(\tilde{t}) = BR(\tilde{b}) = BR(c) = 0\%$	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.55 TeV	$m(\tilde{\chi}_1^0) = 800$ GeV	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	2.1 TeV	$m(\tilde{\chi}_1^0) = 1$ TeV, $\lambda_{112} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	1.65 TeV	$m(\tilde{t}_1) = 1$ TeV, $\lambda_{323} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	15.4	\tilde{t}_1	410 GeV	-	ATLAS-CONF-2016-022, ATLAS-CONF-2016-084
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\ell$	2 e, μ	2 b	-	36.1	\tilde{t}_1	0.4-1.45 TeV	$BR(\tilde{t}_1 \rightarrow b\ell/\mu) > 20\%$	ATLAS-CONF-2017-036
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	510 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10^{-1}

1

Mass scale [TeV]



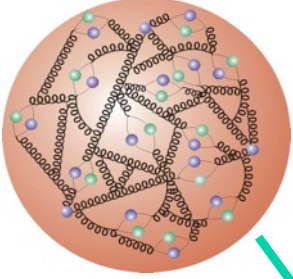
CERN - the longer term plans.... (~ 2035 till ~ 2060)

ILC - CLIC - FCC

and the international competition....

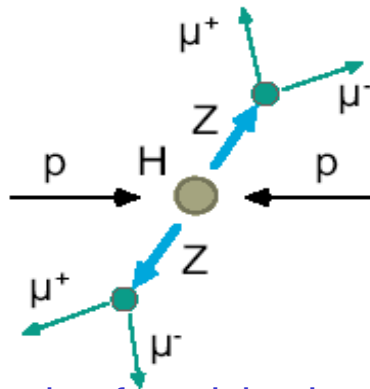
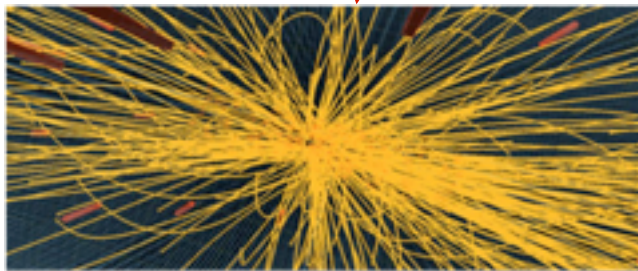
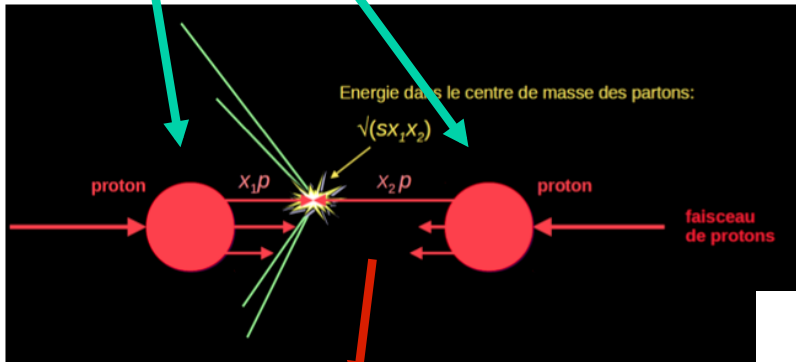


Higgs production at e^+e^- colliders, simple, clean final states in contrast to pp collisions



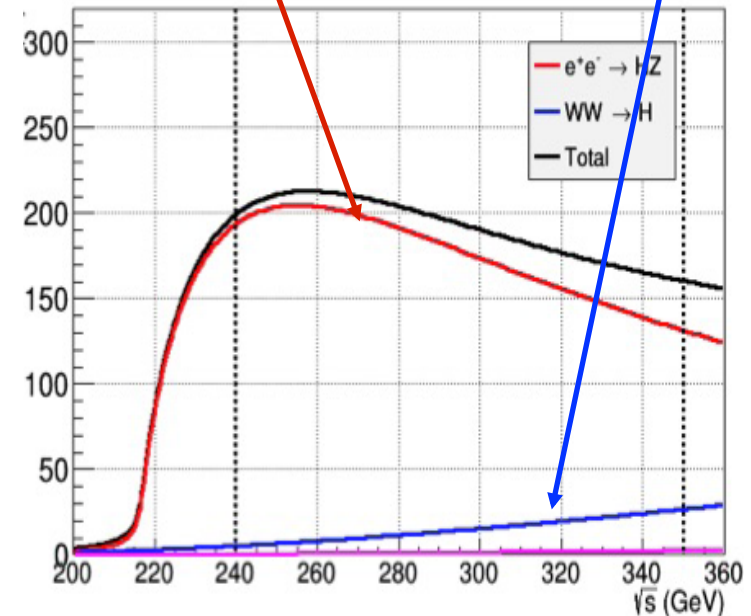
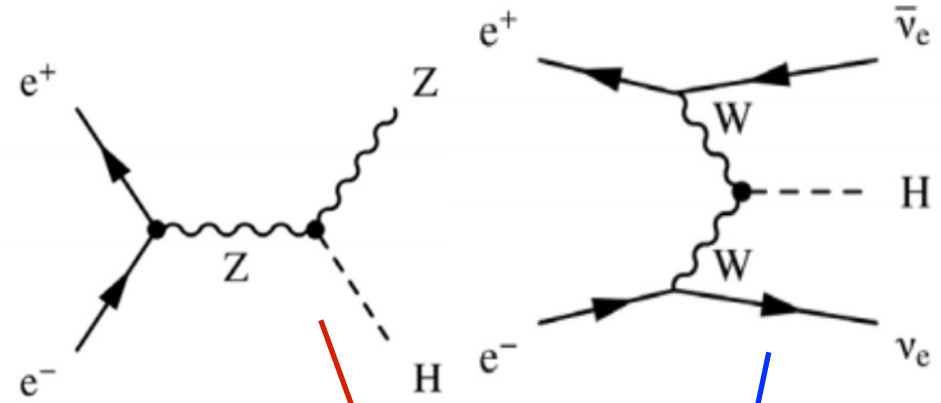
The proton is a complex system, with three 'valence' quarks, a 'sea' of quark-antiquark pairs and a multitude of gluons

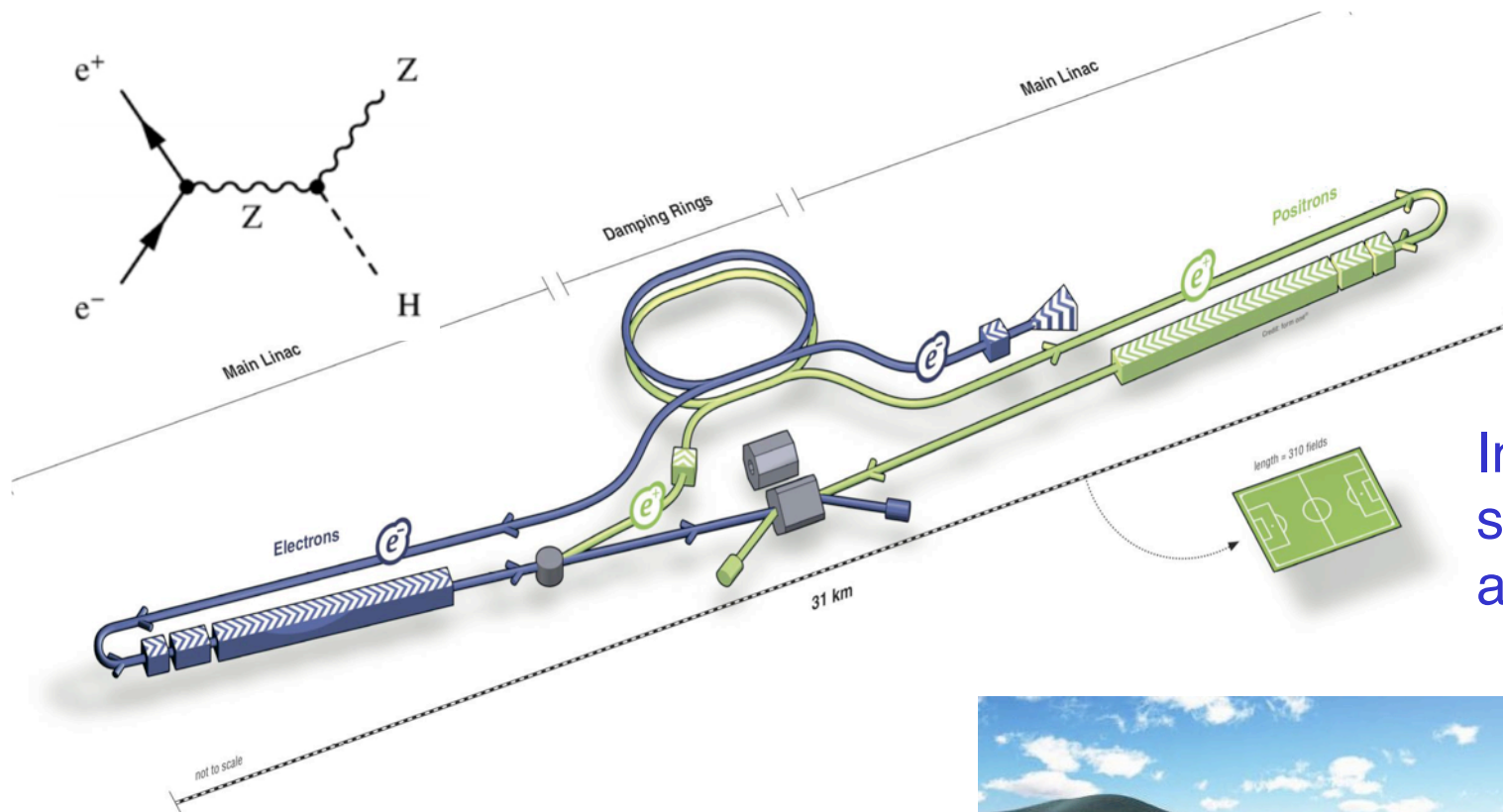
Typical pp collisions:



In proton-proton collisions a multitude of particles in the final state, thus Higgs was searched initially in particularly simple final states!

e^+e^- collisions leading to Higgs production: Simple and clean final states



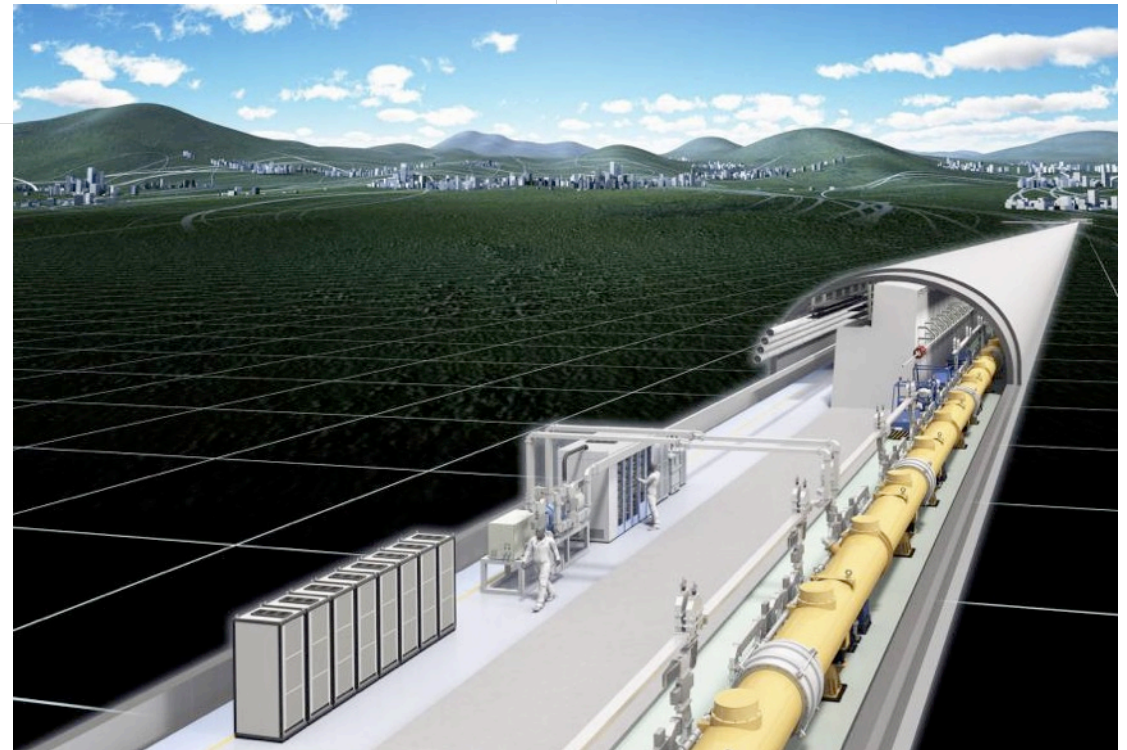


The ILC project

Initially at 250 GeV
subsequently 500 GeV
and ultimately 800 GeV

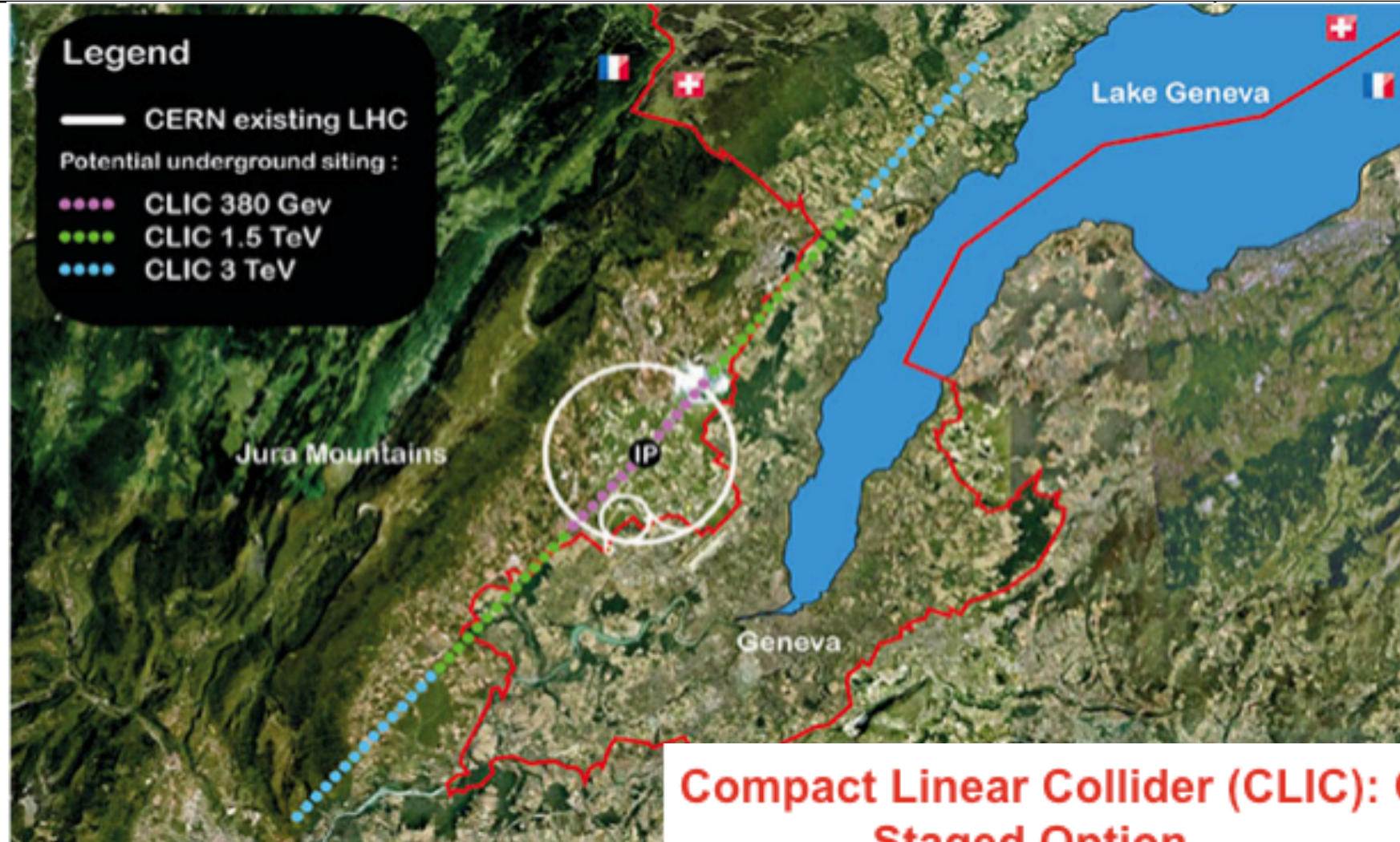
ILC project technically almost ready to go, machine for detailed Higgs production, but new direct discovery potential somewhat limited....

➔ Japan willing to put up half the price.....price about 7 GCHF





Compact Linear Collider CLIC - CERN



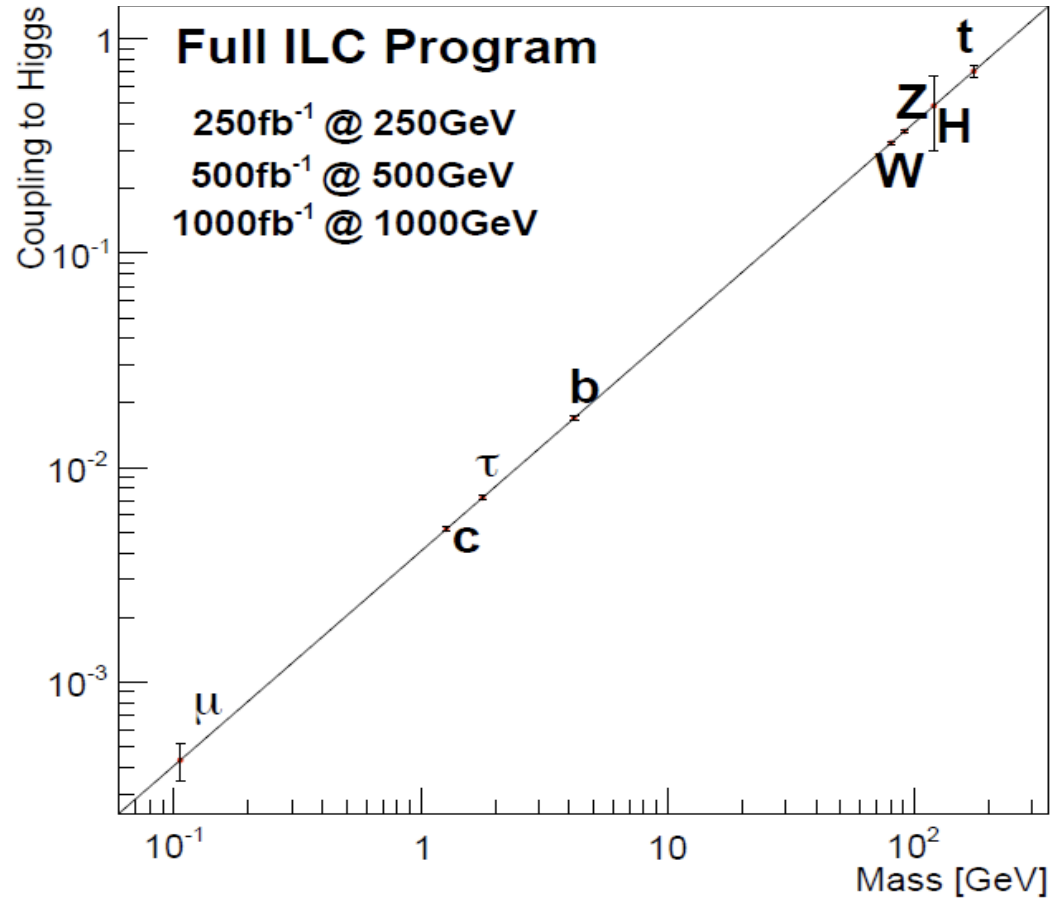
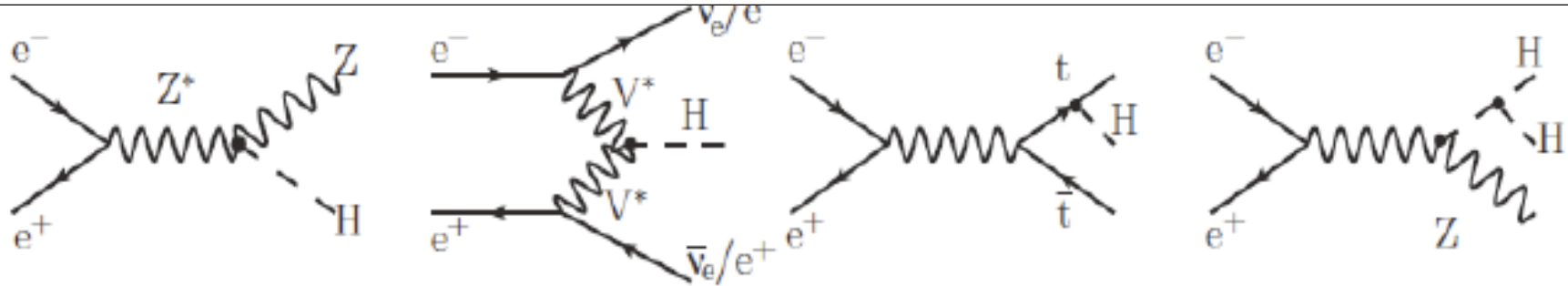
Compact Linear Collider (CLIC): CERN Staged Option

e^-e^+ , \sqrt{s} : 380 GeV, 1.5 TeV, 3 TeV
Length: 11 km, 29 km, 50 km

price of order 10 GCHF



Higgs at ILC.....expected precision on couplings





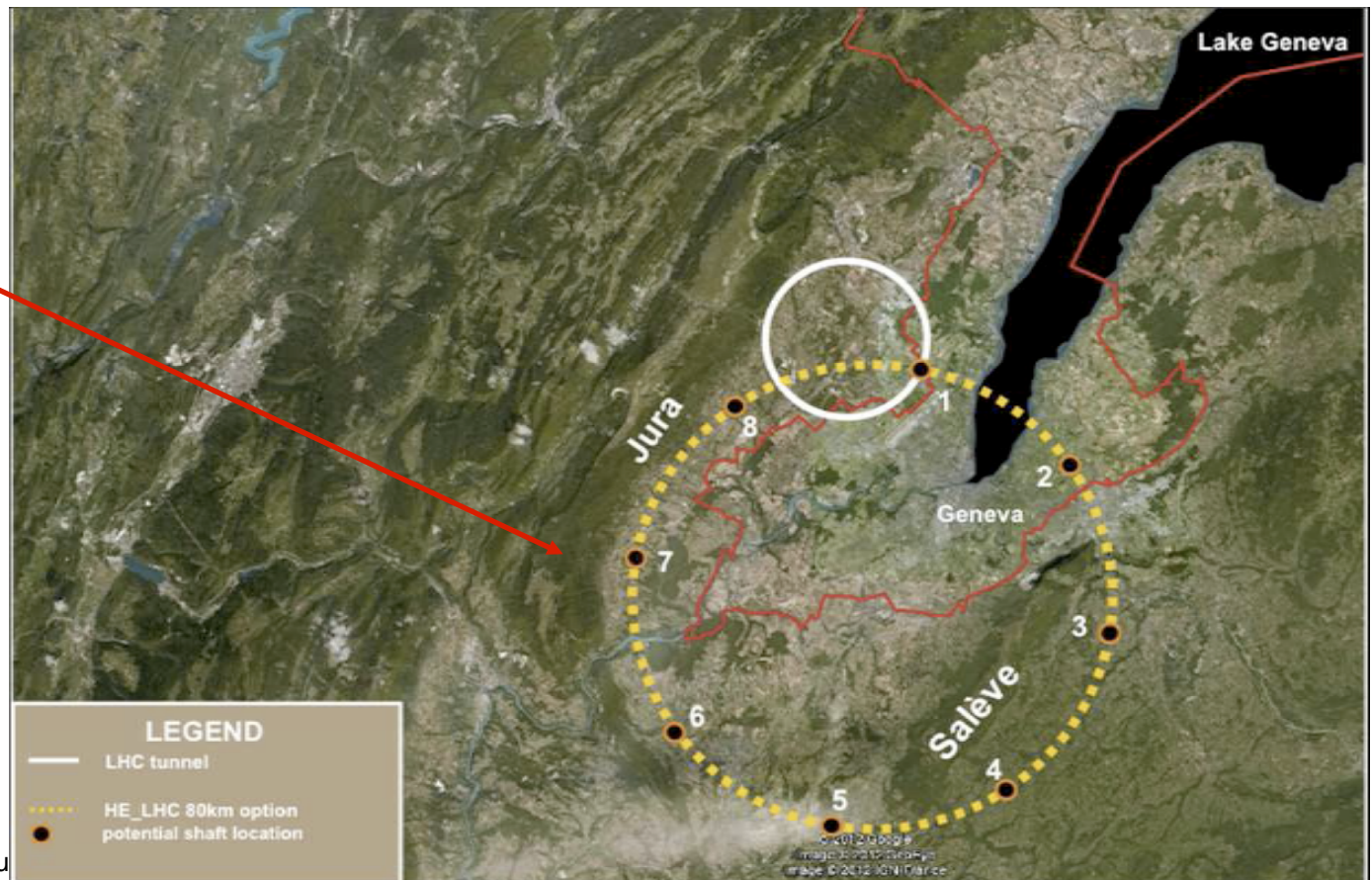
Possible long-term future (beyond ~2035/40) - the FCC project (ee, pp, ep, Pb-Pb)

An Ultra Large Collider to reach 100 TeV in pp mode, an order of magnitude larger than LHC, with an e^+e^- initial phase at ~ 350 GeV, and potential for e-p and Pb-Pb. Projects at the technology frontier at level of design studies and generating requiring/motivating ambitious R&D efforts

An 80-100 km tunnel encompassing all of the Geneva area....

➔ China is proposing to built it in China, first the ee phase, but help needed for the proton-proton phase.

price of order 15 GCHF

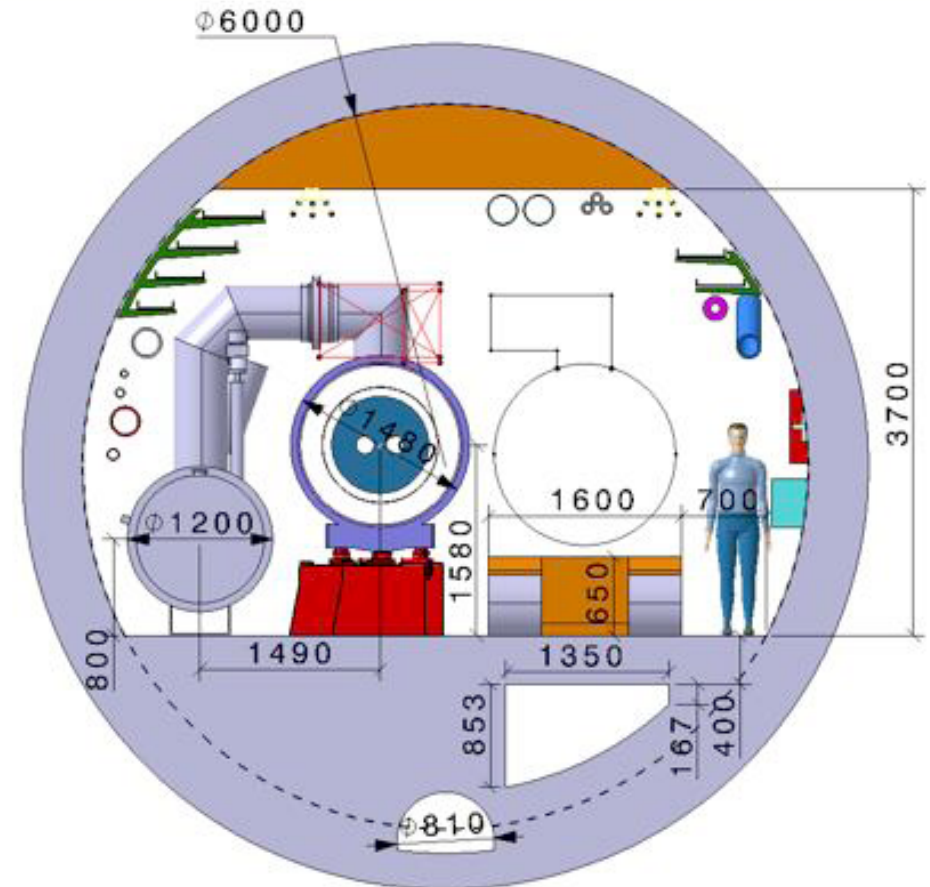
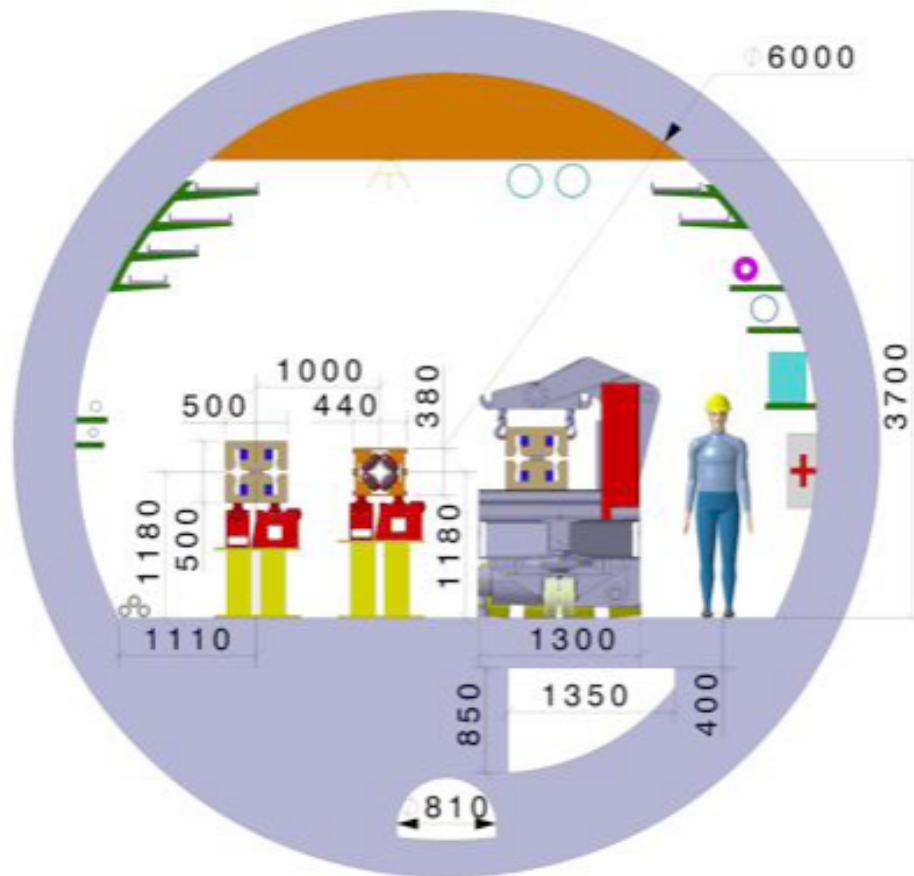




FCC ee and hh tunnel instalations

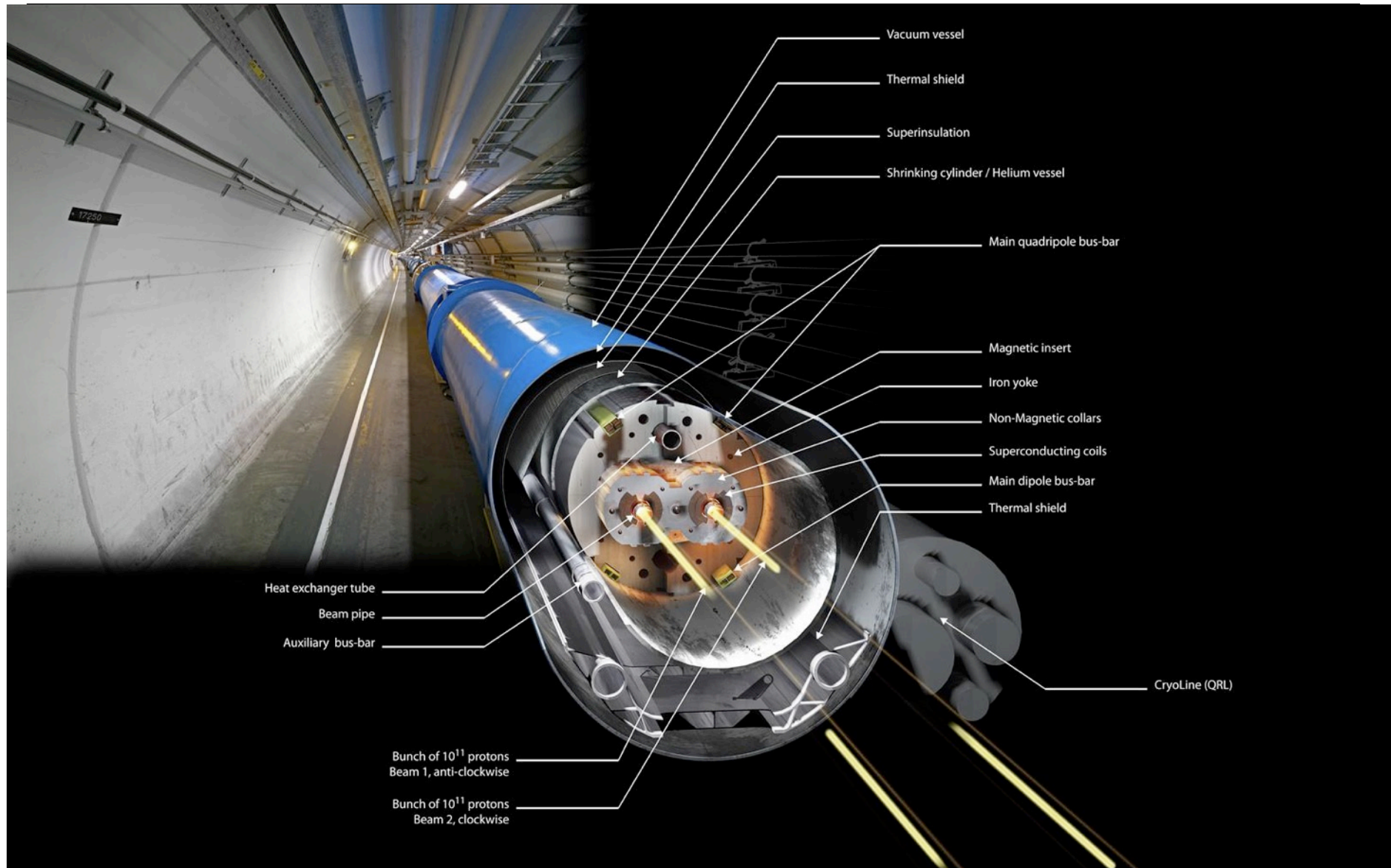
FCC-ee
new 5.5 m inner diameter

FCC-hh



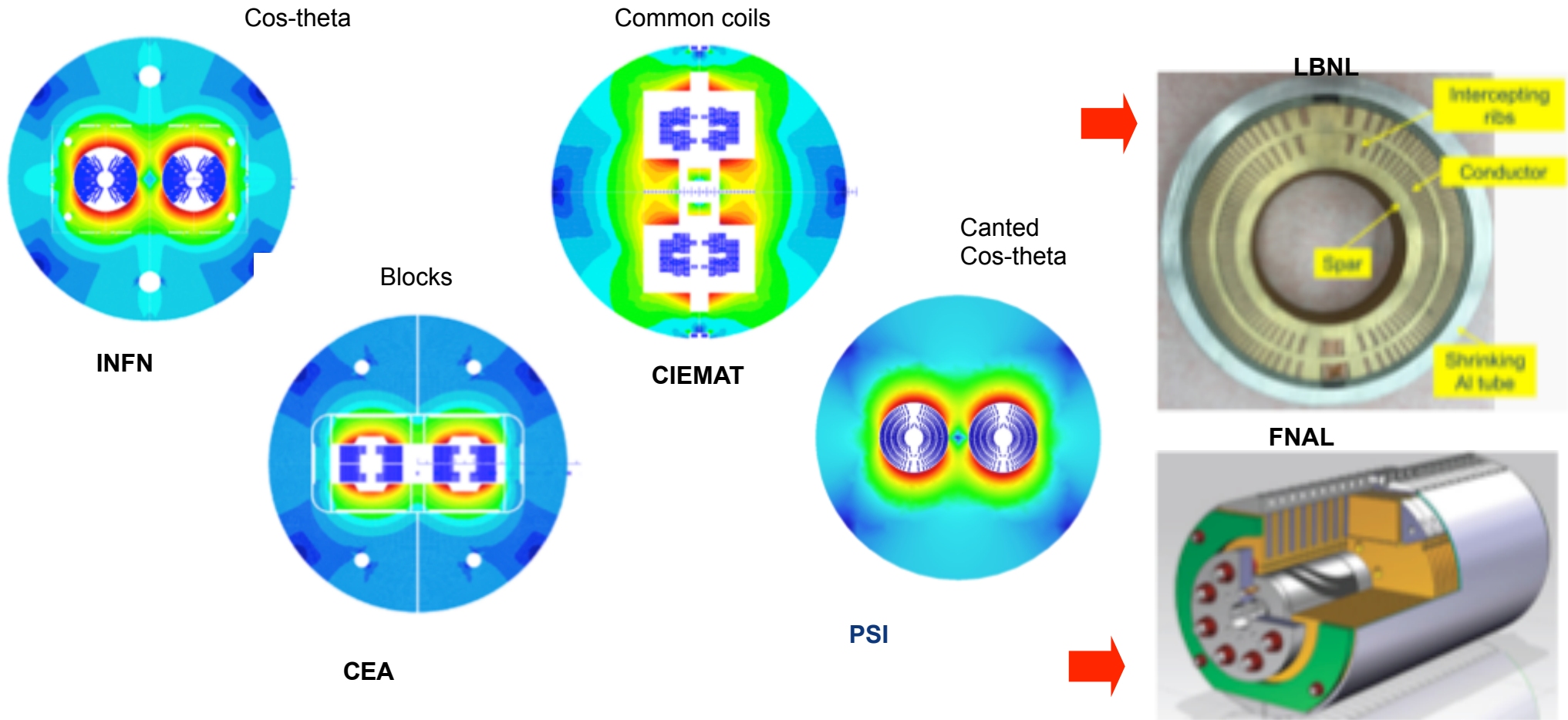


LHC magnets - key of the LHC project.....





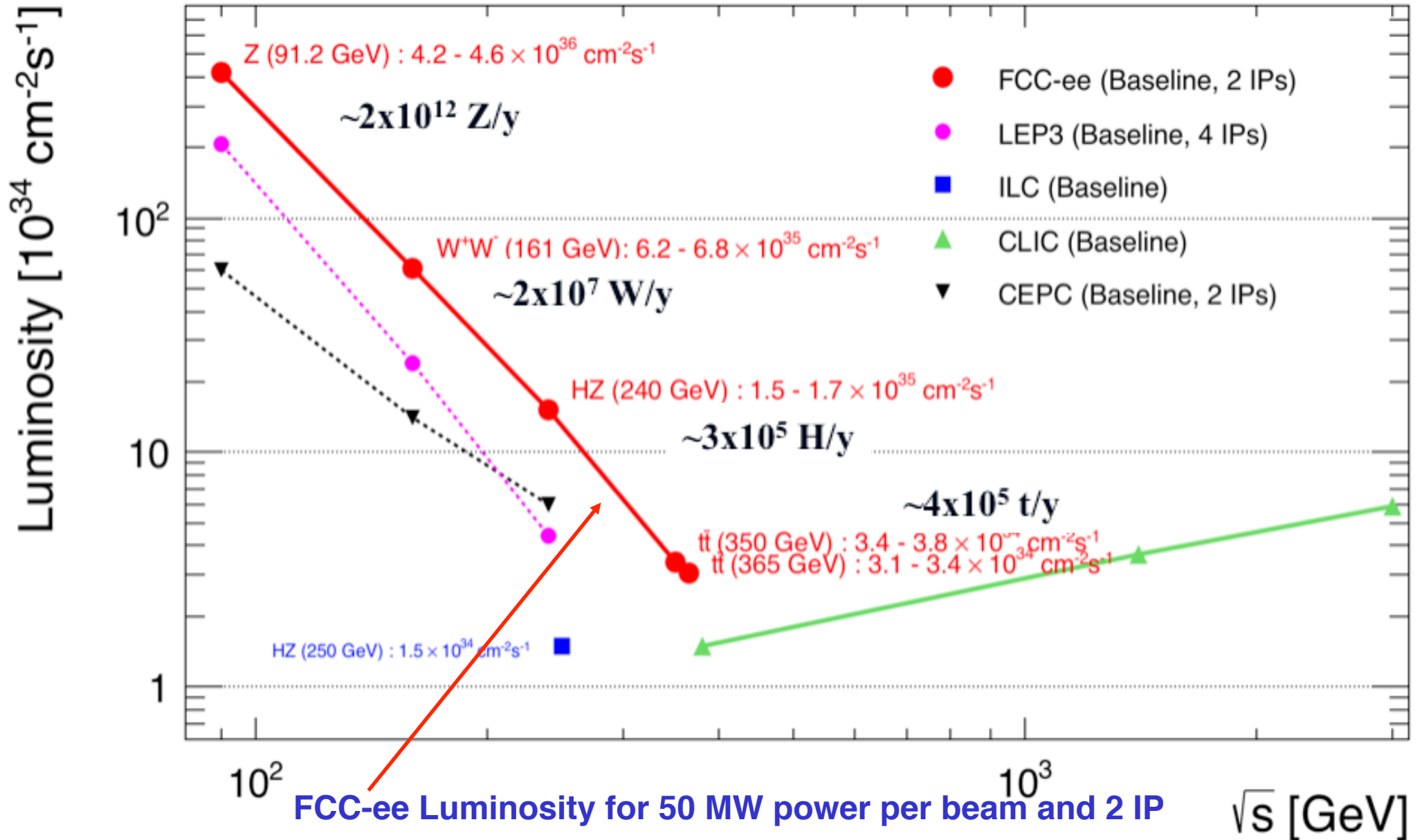
World-wide 16 Tesla dipole design activities and options - for FCC-pp



➔ Short model magnets (1.5 m lengths) will be built from 2018 – 2022
Russian 16 T magnet program launched by BINP recently.



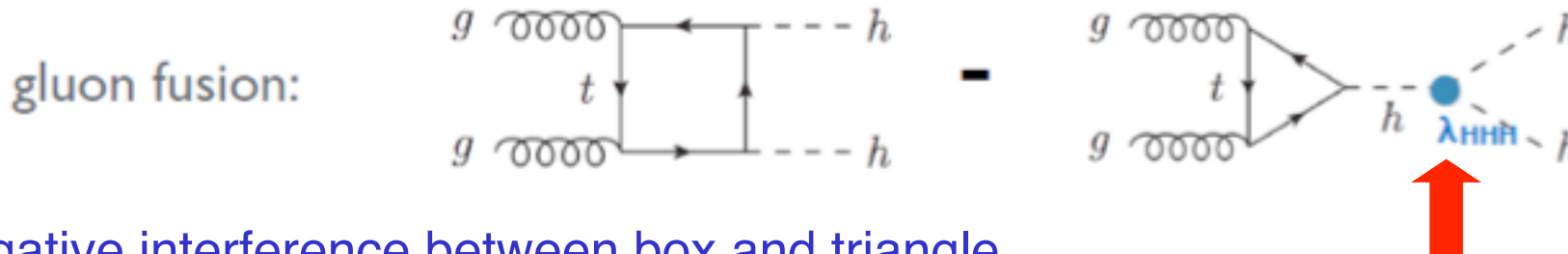
Electron-positron colliders, ILC, CLIC, FCC-ee perspectives



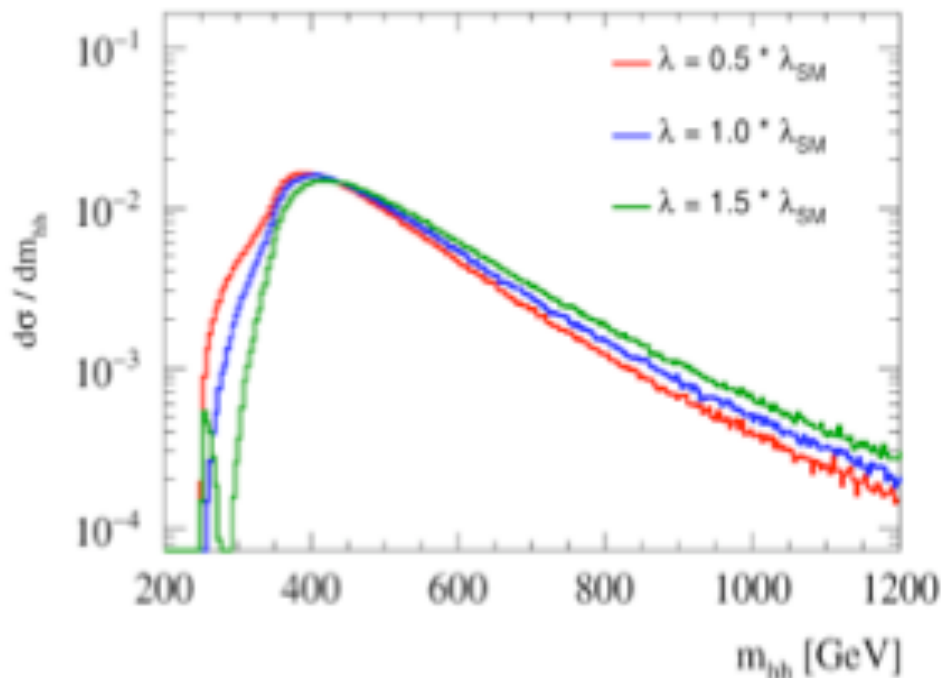


FCC-hh: Higgs self coupling at 100 TeV

$\sigma(100 \text{ TeV})/\sigma(14 \text{ TeV}) \approx 40 \times$ factor 10 for luminosity



Negative interference between box and triangle
Sensitivity to λ from low M_{HH} region



Mode	Sensitivity : $\delta K_\lambda(\text{Stat})$ with 30 ab^{-1}
HH \rightarrow bb $\gamma\gamma$	3.5%
HHj \rightarrow bbbbj + bb $\tau\tau$ j	10-30%
HH \rightarrow bb4l	20-30%
HH \rightarrow bbWW \rightarrow bblvjj	40%

FCC--hh is the machine for the H
self coupling

$$\delta k_\lambda / k_\lambda < 3.5\%$$



Unification of fundamental interactions, past and future possibilities, perspectives...

10^{-35} m
 10^{19} GeV

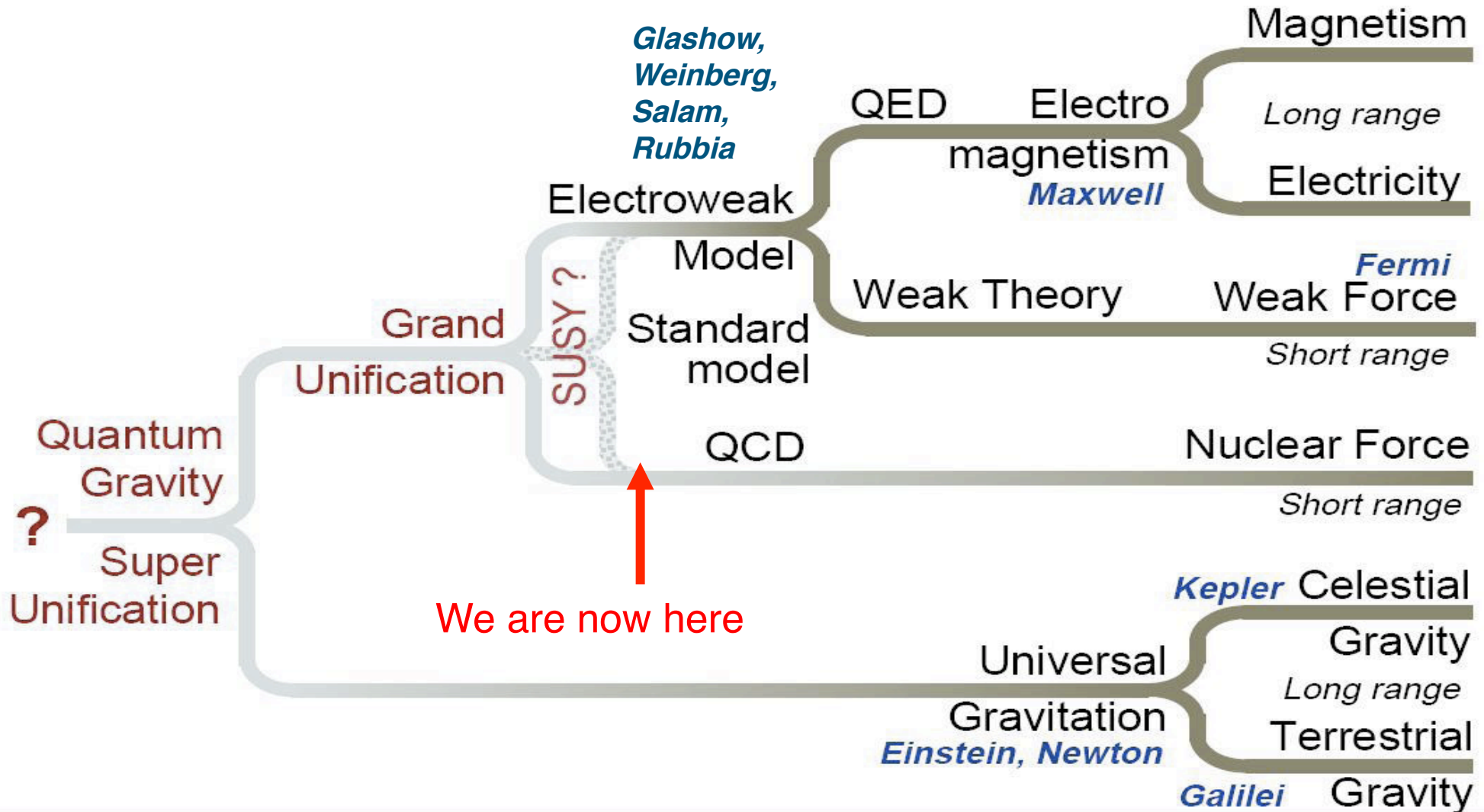
10^{-32} m
 10^{16} GeV

10^{-18} m
 10^2 GeV

10^{-16} m
1 GeV

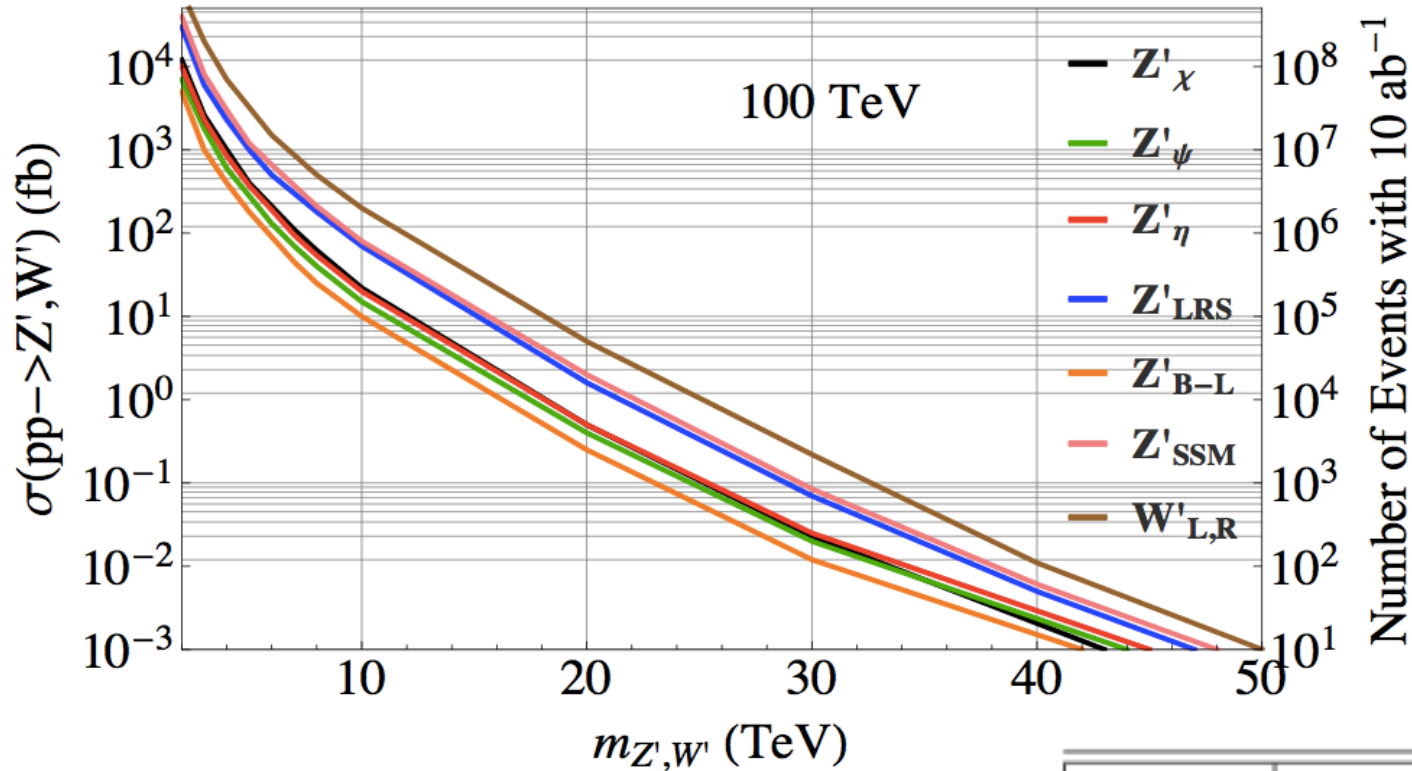
10^{-15} m
1 MeV

10^{-10} m
10 eV





New weak gauge interactions, Fcc-hh reach with 10 atob^{-1}



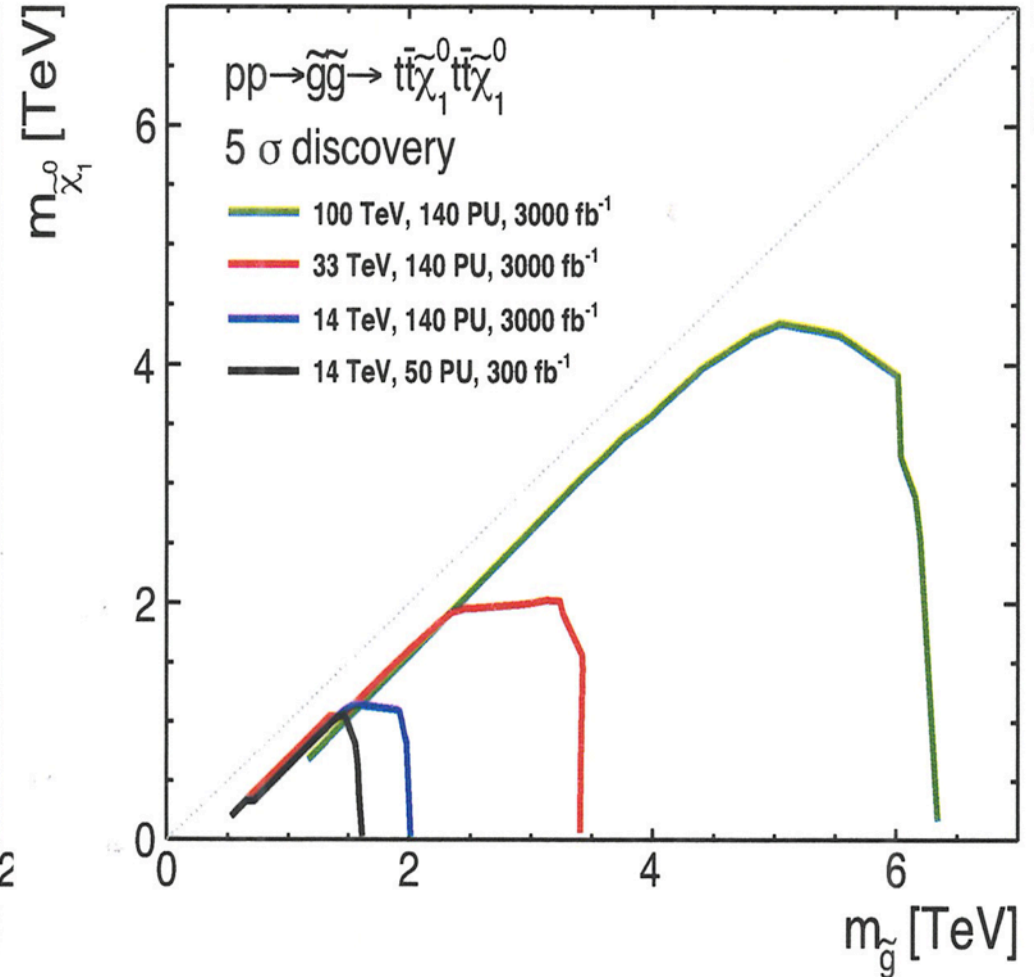
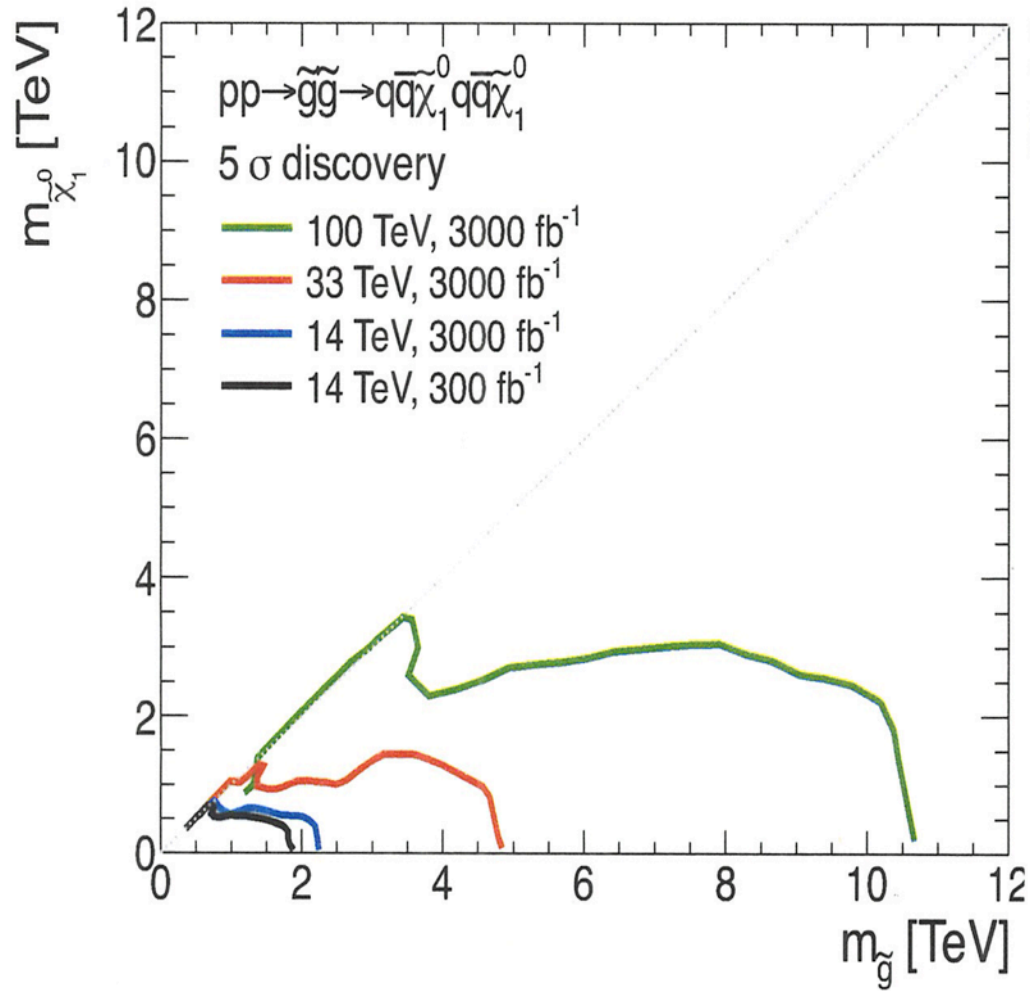
Depending on the Z' model
reach up to 40-50 TeV

Discovery reach
T.Rizzo, arXiv:1403.5465

Model	1 ab^{-1}	10 ab^{-1}	100 ab^{-1}
SSM	23.8	33.3	41.3
LRM	22.6	31.5	39.5
ψ	20.1	29.1	37.2
χ	22.7	30.6	38.2
η	20.3	29.8	38.0
I	22.4	29.2	36.2



Expected gluino reaches at 33 and 100 TeV in FCC-hh machine with 3000 fb⁻¹





Conclusions

The LHC is an incredible technological and scientific endeavor
- on a world-wide scale

The four major experiments ATLAS, CMS, ALICE and LHCb are all operating very successfully. The physics harvest up to now is extraordinary, the Higgs, beautiful and detailed studies in EWK and QCD physics, rare decay modes, QGP studies etc

In 2015 LHC started operating at ~ 13 TeV; many technical challenges. Aims: clarification of the Higgs (i.e. is it THE Higgs boson or a A Higgs from an extension of the SM), looking for SUSY etc

The particle physics community is already working on long-term options/projects to take over in ~ 20 -30 years from now



SPARES

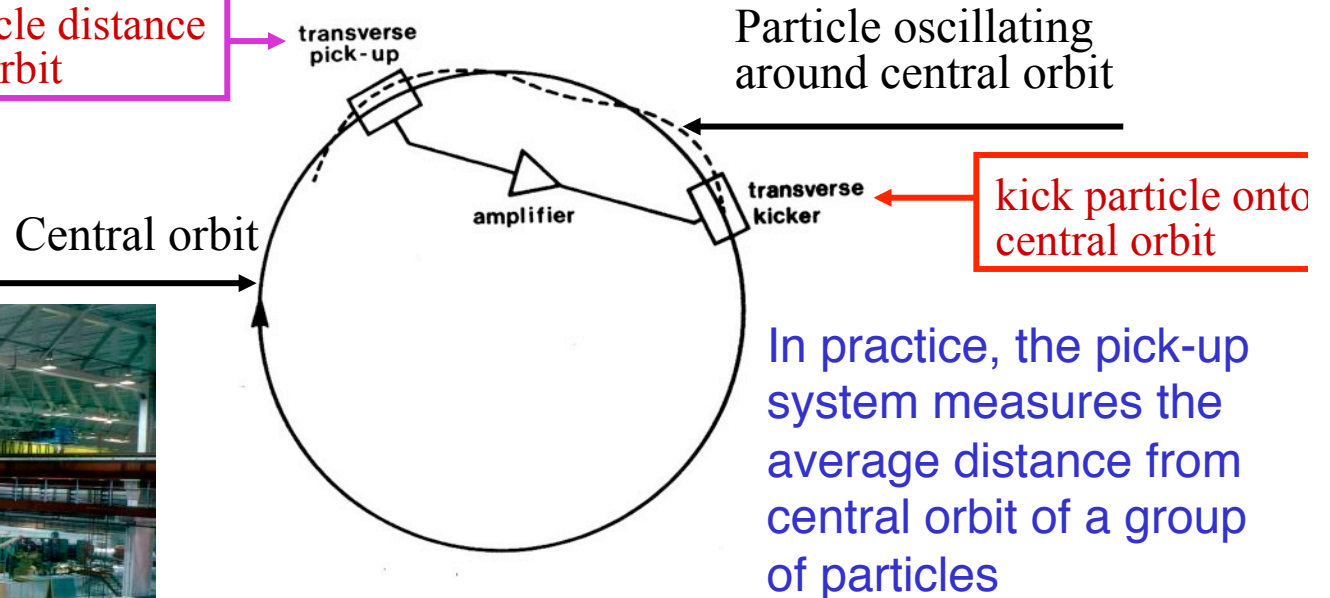


Stochastic cooling (S. van der Meer)

Principle of operation: for ex. cooling in the horizontal plane:

Measure particle distance from central orbit

3.5 GeV/c large-aperture ring for antiproton storage and cooling

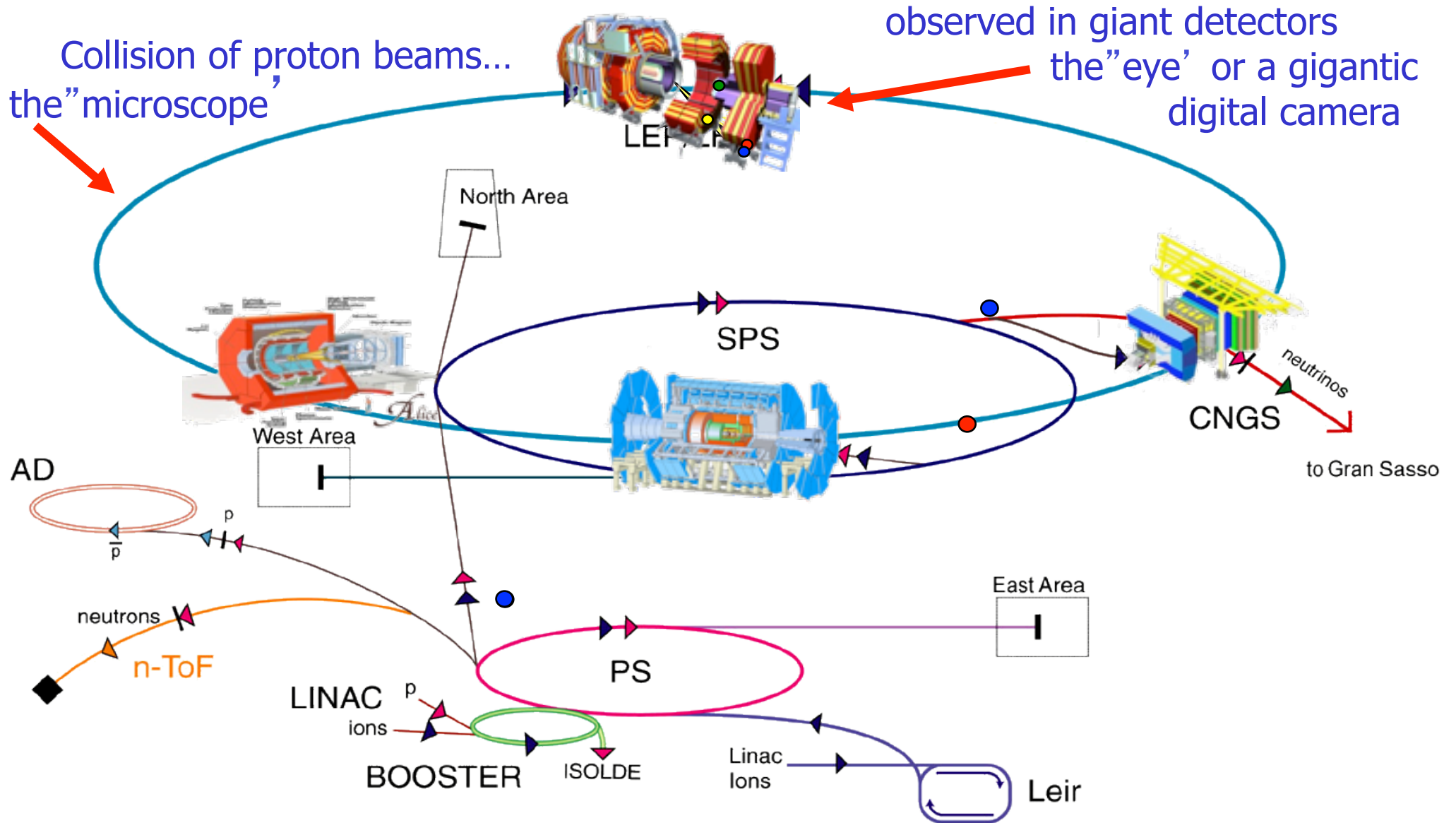


- Independent pick-up – kicker systems to cool:
- horizontal motion
 - vertical motion
 - longitudinal motion (decrease of $\Delta p/p$)

- Antiproton production rate : 1 antiproton (at 3.5 GeV) per 10^6 incident protons of 26 GeV on target. About 10^{11} cooled antiprotons were accumulated per day; luminosity lifetime \sim one day



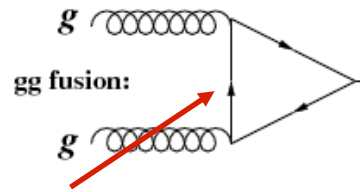
The Large Hadron Collider and experiments; the sequence of accelerators - injectors





Higgs production mechanisms and decay modes

87.4 % @120 GeV

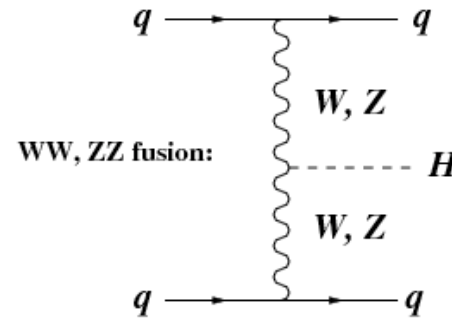


indirect coupling to gg,
top in the loop

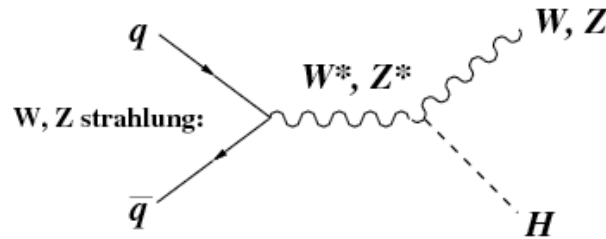
$$g_{HVV} = 2m_V^2/v$$

$$g_{HWW} = g m_W$$

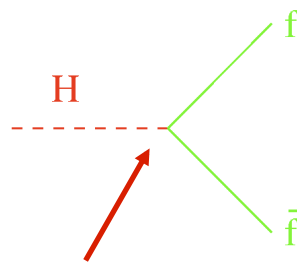
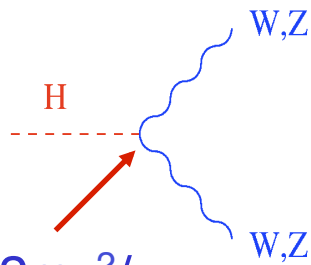
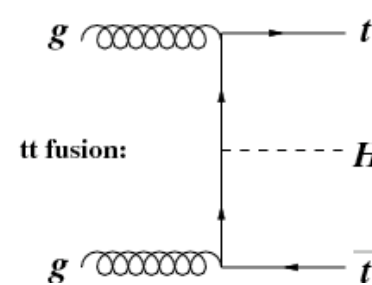
6.7 %



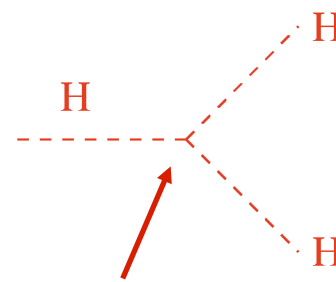
5.4 %



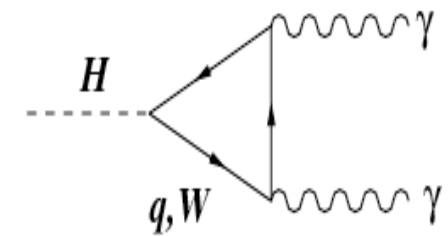
0.5 %



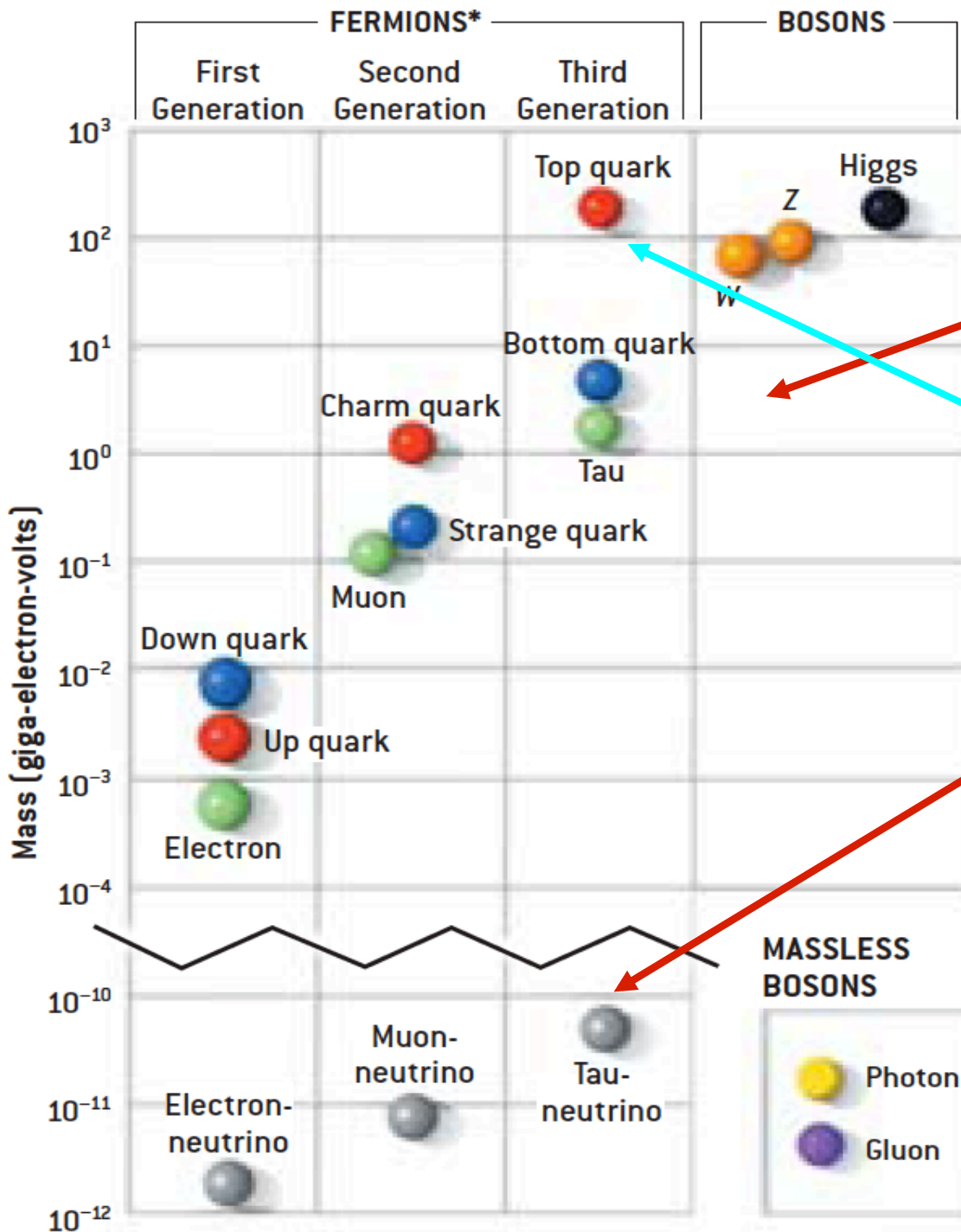
$$g_{Hff} = m_f/v$$



$$g_{HHH} = 3m_H^2/v$$



indirect coupling to $\gamma\gamma$



Particle masses and the Higgs boson

Higgs boson responsible for the masses of W, Z, quarks, leptons

Top quark

the Higgs boson is not responsible for the neutrino masses in the SM!
...a problem for the future....

Mass-less particles/fields, not coupled to the Higgs boson field



CMS phase-2 upgrades to face the challenges of HL-LHC

Muon system

- GEM Glass RPCs
- Extended η coverage
- New DT minicrates

Tracker

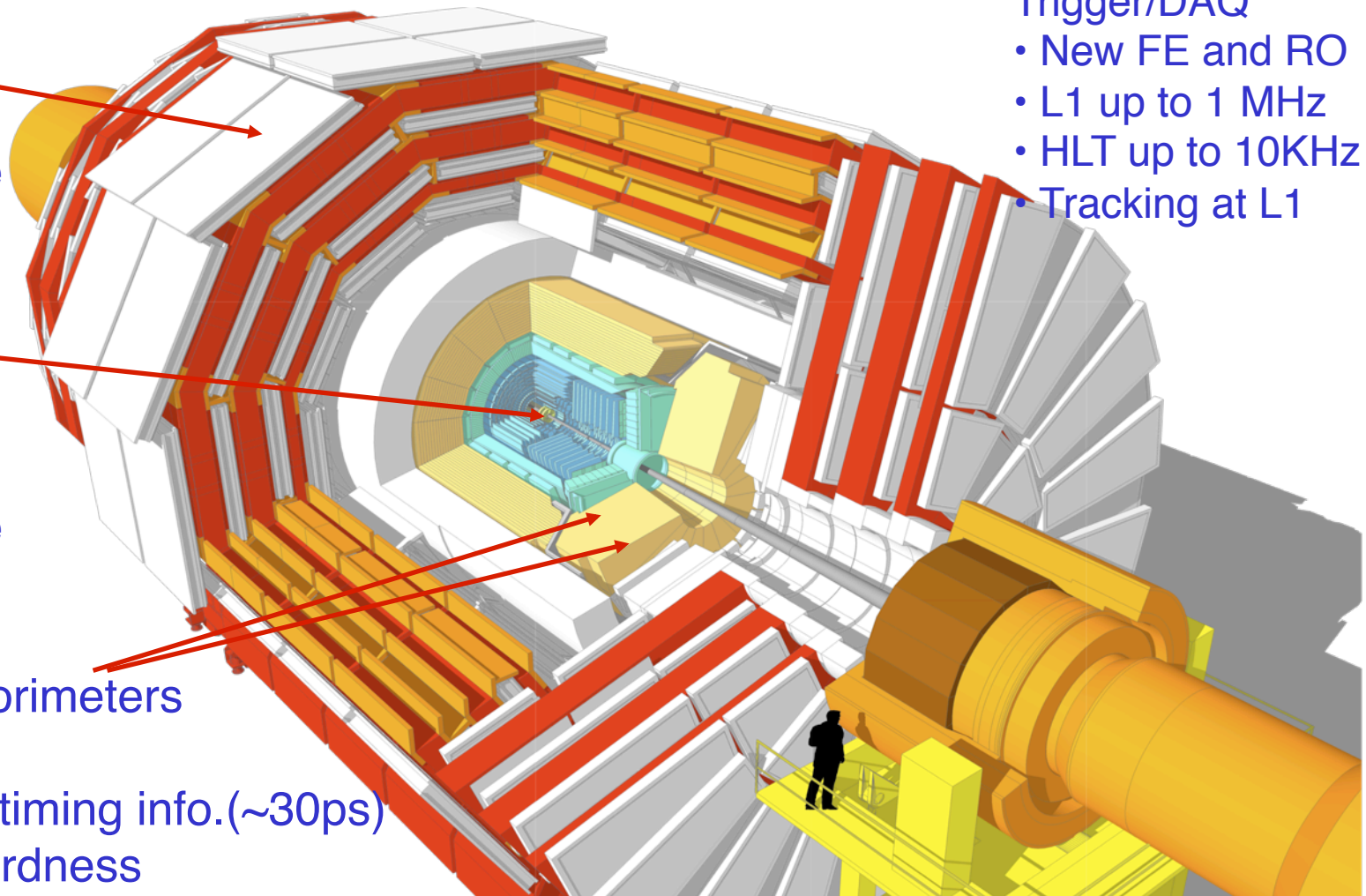
- Higher granularity
- Less material
- Better p_T resolution
- Extended η coverage
- Track trigger at L1

Replace endcap calorimeters

- Higher granularity
- 3D reconstruction, timing info. ($\sim 30\text{ps}$)
- Higher radiation hardness

Trigger/DAQ

- New FE and RO
- L1 up to 1 MHz
- HLT up to 10KHz
- Tracking at L1



The CMS-Phase 2 (HL-LHC) upgrade is a major construction effort
financial level is at 50% of the original CMS detector cost, ~ 300 MCHF



Linear colliders in competition

Competition of two concepts.

- Length determined by efficiency (gradient) of RF cavities.

- ILC – International Linear Collider, 0.5 TeV, based on Superconducting RF cavities (**gradient 31.5 MV/m**)

1st option $E_{cm} = 250 \text{ GeV}$

- CLIC – Compact Linear Collider, developed by CLIC Collaboration (CERN), 0.5 -3 TeV , based on warm RF cavities at 12 GHz with very high el. **gradient $\sim 100 \text{ MV/m}$**

1st option $E_{cm} = 380 \text{ GeV}$

Competition but also cooperation:

CLIC + ILC >> Linear Collider Collaboration



ILC - International Linear Collider

QuickTime™ and a
decompressor
are needed to see this picture.

Collisions: Between electrons - positrons, in bunches of 5 nm
Energy: Up to 0.5 TeV with an option to upgrade to 1 TeV
Acceleration Technology: 16,000 superconducting accelerating
cavities made of pure niobium
Length: Approximately 31 kilometres
Accelerating Gradient: 31.5 megavolts per metre



1.3 GHz superconducting cavities for ILC

Gradient 31.5 MV/m, restricted by quality of forming and welding techniques (surface roughness, impurities...) of bulk niobium





FCC-ee challenges

Deal with small beam life time

⇒ Top up scheme needed

QuickTime™ and a
decompressor
are needed to see this picture.

QuickTime™ and a
decompressor
are needed to see this picture.

QuickTime™ and a
decompressor
are needed to see this picture.



FCC-hh parameters

QuickTime™ and a
decompressor
are needed to see this picture.



FCC-hh main cross sections

