## NNLO QCD + NLO EW for diboson processes

Jonas M. Lindert

## 図 <br> University

Multi-Boson Interactions 2019
Thessaloniki, 26. August 2019

## Status



Remarkable agreement of inclusive diboson cross sections with NNLO QCD

Allows for stringent SM tests
Dibosons important background for Higgs and BSM searches

Tails, tails, tails,....



## NNLO QCD corrections vor VV

All VV processes known through NNLO QCD:
[Talk by M. Wiesemann]
$\rightarrow$ inclusive/on-shell Z,W \& differential/off-shell Z,W (Ieptonic)
YY - inclusive and differential [Catani, Cieri, de Florian, Ferrera, Grazzini 'I2], [Campbell, Ellis, Li, Williams 'I6], [Grazzini, Kallweit, MW 'I7]
$\mathbf{Z}_{\boldsymbol{\gamma}}$ - inclusive/on-shell and differential/off-shell [Grazzini, Kallweit, Rathlev, Torre 'I3], [Grazzini, Kallweit, Rathlev 'I5]; see also: [Campbell et al. 'I7]
$\mathbf{W} \mathbf{\gamma}$ - inclusive/on-shell and differential/off-shell [Grazzini, Kallweit, Rathlev, Torre 'I3], [Grazzini, Kallweit, Rathlev 'I5]
$\mathbf{Z Z}$ - inclusive/On-shell [Cascioli, Gehrmann, Grazzini, Kallweit, Maierhöfer, von Manteuffel, Pozzorini, Rathlev, Tancredi,Weihs 'I4]; see also: [Heinrich et al. 'I7]

- differential/off-shell [Grazzini, Kallweit, Rathlev 'I5], [Kallweit, MW 'I8]

WW - inclusive/On-shell [Gehrmann, Grazzini, Kallweit, Maierhöfer, von Manteuffel, et al. 'I4]

- differential/off-shell [Grazzini, Kallweit, Pozzorini, Rathlev, MW 'I5]

WZ - inclusive/on-shell [Grazzini, Kallweit, Rathlev, MW 'I6]

- differential/off-shell [Grazzini, Kallweit, Rathlev, MW 'I7]


## Perturbative expansion

$$
\begin{aligned}
& \mathrm{d} \sigma=\mathrm{d} \sigma_{\mathrm{LO}}+\alpha_{S} \mathrm{~d} \sigma_{\mathrm{NLO}}+\alpha_{\mathrm{EW}} \mathrm{~d} \sigma_{\text {NLO EW }} \\
& \quad \text { NLO QCD } \quad \text { NLO EW } \\
& \\
& +\alpha_{S}^{2} \mathrm{~d} \sigma_{\mathrm{NNLO}}+\alpha_{\mathrm{EW}}^{2} \mathrm{~d} \sigma_{\text {NNLOEW }}+\alpha_{S} \alpha_{\mathrm{EW}} \mathrm{~d} \sigma_{\text {NNLO QCDxEW }}+\ldots \\
& \text { NNLO QCD } \quad \text { NNLO EW } \quad \text { NNLO QCD-EW }
\end{aligned}
$$

Numerically $\mathcal{O}(\alpha) \sim \mathcal{O}\left(\alpha_{s}^{2}\right) \Rightarrow$ NLO EW $\sim$ NNLO QCD

## NLO EW

-4I-DF-ZZ Biedermann, Denner, Dittmaier, Hofer, Jäger; 'I6,'I6
-2l-DF-WW Biedermann, Billoni, Denner, Dittmaier, Hofer, Jäger, Salfelder; 'I6
-2I-SF-ZZ \& 2l-SF-ZZWW \& 2l-DF-WW Kallweit, JML, Pozzorini, Schönherr, 'I7
-3I-DF-WZ \& 3I-DF-WZ Biedermann, Denner, Hofer, 'I7

## Relevance of EW higher-order corrections: Sudakov logs in the tails

I. Possible large (negative) enhancement due to soft/collinear logs from virtual EW gauge bosons:


$\rightarrow$ overall large (negative) effect in the tails of distributions: PT, $m_{\text {inv }}, H_{T}, \ldots$ (relevant for BSM searches!)

## Relevance of EW higher-order corrections: collinear QED radiation

II. Possible large enhancement due to soft/collinear logs from photon radiation $\sim \alpha \log \left(\frac{m_{f}^{2}}{Q^{2}}\right)$ in sufficiently exclusive observables.


$\rightarrow$ important for radiative tails, Higgs backgrounds etc.
$\rightarrow$ typically considered via QED PS (PHOTOS / YFS)

## Relevance of EW higher-order corrections: photon-induced channels

III. QED factorisation and thus photon luminosities needed to absorb IS photon singularities.
$\rightarrow$ Possible large enhancement due to photon-induced channels in the tails of kinematic distributions,

$\rightarrow$ large differences between different photon descriptions. Now settled: LUXqed superior
$\rightarrow \mathrm{O}(10 \%)$ contributions from photon-induced channels

## Nontrivial features in NLO QCD $\rightarrow$ NLO EW

I. QCD-EW interplay

3. virtual EW corrections more involved than QCD (many internal masses)


Automation of fixed-order NLO EW well advanced:
MadGraph_aMC@NLO, Sherpa+OpenLoops/Recola, MUNICH+OpenLoops, ...

## Validation between tools

- There are subtle differences in implementation of these schemes in particular in the context of CMS (complex mass scheme).
$\rightarrow$ Have been studied for ZZ in the context of [LHI7, I 803.07977]
a) $\operatorname{PSP} 1$
$B / 10^{-15}$
$V_{\text {finite }} / 10^{-16}$
$V_{1} / 10^{-17}$
$V_{2} / 10^{-17}$

Madoop 5.26592465401088 Recola 5.26592465401090 OpenLoops 5.26592465401100 GoSAm $\quad 5.26592465401086$ NLOX 5.26592465401084
c) $\operatorname{PSP} 1$

MadLoop 4.63762790127829
Recola 4.63762790127830 OpenLoops 4.63762790127838
GoSam 4.63762790127830
6.60297993618509 6.60088670209820 6.60088670210145 .60088670209788 6.60088670211436
$2.63915540074976-3.09566543908773$ $2.63915540075328-3.09566543908732$ $2.63915540078563-3.09566543905505$ $.63915540076095-3.09566543909091$ $2.63915540076702-3.09566543908783$

$$
V_{1} / 10^{-15} \quad V_{2} / 10^{-15}
$$

$4.07216839247769-2.23061748556626$ $4.07216839245629-2.23061748556050$ $4.07216839246097-2.23061748560388$ $4.07216839247955-2.23061748556541$
inclusive cross sections:

$\rightarrow$ very convincing agreement between automated tools

# Diboson production at NLO QCD+EW: Collinear QED radiation 

[Kallweit, JML, Pozzorini, Schönherr; ' I 7]



YFS (Multi-Photon-Resummation) preserves resonance structure
$\rightarrow$ EW effects agree at the few percent level.

Source of differences:

- Multi-poton effects in YFS
-Resonance-assignment in YFS

CSS (Catani-Seymour-Shower) unaware of resonance structure $\rightarrow$ QED effects largely overestimated

- Fully consistent PS matching at NLO EW under development
- Naive NLO EW+PS matching available in Sherpa+OpenLoops (applicable at particle level)
$\Rightarrow$ CSS dipole shower (not resonaonce aware) $\Rightarrow$ significant mismodelling
$\Rightarrow$ YFS resummation (resonaonce aware) $\Rightarrow$ valid approximation


## The need for off-shell calculations

[Biedermann, M. Billoni, A. Denner, S. Dittmaier, L. Hofer, B. Jäger, L. Salfelder ;'I 6]


$\vec{p}_{\mathrm{T}, \mathrm{e}^{-}}$

$\Rightarrow$ sizeable differences in fully off-shell vs. double-pole approximation in tails

## Combination of NNLO QCD and NLO EW

- In Matrix+OpenLoops all (massive) diboson processes are available at

NNLO QCD + NLO EW (parton-level) [M. Grazzini, S. Kallweit, JML, S. Pozzorini, M. Wiesemann; very soon]

| 41-SF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{+} \ell^{-}$ | (ZZ) |  |
| :---: | :---: | :---: | :---: |
| 41-DF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{\prime+} \ell^{\prime-}$ | (ZZ) |  |
| 2l-SF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \nu_{\ell^{\prime}}{\bar{\nu} \ell^{\prime}}$ | (ZZ) |  |
| 2l-SF-ZZWW | $p p \rightarrow \ell^{+} \ell^{-} \nu_{\ell} \bar{\nu}_{\ell}$ | (ZZ/WW) | (soon to be made public) |
| 2l-DF-WW | $p p \rightarrow \ell^{+} \ell^{\prime-} \nu_{\ell} \bar{\nu}_{\ell^{\prime}}$ | (WW) |  |
| 31-SF-WZ | $p p \rightarrow \ell^{+} \ell^{-} \ell \nu_{\ell}$ | (WZ) |  |
| 31-DF-WZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{\prime} \nu_{\ell^{\prime}}$ | (WZ) |  |

- Combination of QCD and EW
additive: $\quad \mathrm{d} \sigma_{\mathrm{QCD}+\mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}}=\mathrm{d} \sigma^{\mathrm{LO}}\left(1+\delta_{\mathrm{QCD}}^{(\mathrm{N}) \mathrm{NLO}}+\delta_{\mathrm{EW}}\right)+\mathbf{d} \sigma^{\mathrm{ggLO}}$
multiplicative: $\quad \mathbf{d} \sigma_{\mathbf{Q C D} \times \mathbf{E W}}^{(\mathrm{N}) \mathbf{N L O}}=\mathrm{d} \sigma^{\mathrm{LO}}\left(1+\delta_{\mathbf{Q C D}}^{(\mathrm{N}) \mathrm{NLO}}\right)\left(1+\delta_{\mathrm{EW}}\right)+\mathbf{d} \sigma^{\mathrm{ggLO}}$


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| 41-SF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{+} \ell^{-}$ | (ZZ) |  |
| :---: | :---: | :---: | :---: |
| 41-DF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{\prime+} \ell^{\prime-}$ | (ZZ) |  |
| 2l-SF-ZZ | $p p \rightarrow \ell^{+} \ell^{-} \nu_{\ell^{\prime}} \bar{\nu}_{\ell^{\prime}}$ | (ZZ) |  |
| 2l-SF-ZZWW | $p p \rightarrow \ell^{+} \ell^{-} \nu_{\ell} \bar{\nu}_{\ell}$ | (ZZ/WW) | (soon to be made public) |
| 2l-DF-WW | $p p \rightarrow \ell^{+} \ell^{\prime-} \nu_{\ell} \bar{\nu}_{\ell^{\prime}}$ | (WW) |  |
| 31-SF-WZ | $p p \rightarrow \ell^{+} \ell^{-} \ell \nu_{\ell}$ | (WZ) |  |
| 3l-DF-WZ | $p p \rightarrow \ell^{+} \ell^{-} \ell^{\prime} \nu_{\ell^{\prime}}$ | (WZ) |  |

- Combination of QCD and EW
additive: $\quad \mathrm{d} \sigma_{\mathrm{QCD}+\mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}}=\mathrm{d} \sigma^{\mathrm{LO}}\left(1+\delta_{\mathrm{QCD}}^{(\mathrm{N}) \mathrm{NLO}}+\delta_{\mathrm{EW}}\right)+\mathbf{d} \sigma^{\mathrm{ggLO}}$
multiplicative: $\quad \mathbf{d} \sigma_{\mathbf{Q C D} \times \mathbf{E W}}^{(\mathrm{N}) \mathbf{N L O}}=\mathrm{d} \sigma^{\mathrm{LO}}\left(1+\delta_{\mathbf{Q C D}}^{(\mathrm{N}) \mathrm{NLO}}\right)\left(1+\delta_{\mathrm{EW}}\right)+\mathbf{d} \sigma^{\mathrm{ggLO}}$


## Giant K-factors and EW corrections



# Giant K-factors and EW corrections 



- NLO QCD/LO=~<1.5
("normal K-factor")
- Reliable estimate of O (as a) from

$$
\mathrm{d} \sigma_{\mathrm{QCD}+\mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}} \text { vs. } \mathbf{d} \sigma_{\mathrm{QCD} \times \mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}}
$$

$p p \rightarrow \mathrm{e}^{-} \mathrm{e}^{+} \nu_{\mu} \bar{\nu}_{\mu}$
$\xrightarrow[\text { jet veto }]{\longrightarrow}$
$H_{\mathrm{T}}^{\text {jet }}<0.2 H_{\mathrm{T}}^{\text {lep }}$
-However:
additional uncertainty due to efficiency of jet veto

## Photon-induced contributions



## Photon-induced contributions



# Giant K-factors and EW corrections: pTII 




inclusive

- Same features as pTVI


# Giant K-factors and EW corrections: pTII 





## Without giant K-factors: stable predictions for pTV2



## inclusive

Reliable estimate of $O$ (as a) from


## Without giant K-factors: stable predictions for mVV





## inclusive

Reliable estimate of $O$ (as a) from

$$
\begin{gathered}
\mathrm{d} \sigma_{\mathrm{QCD}+\mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}} \\
\mathrm{VS} . \\
\mathbf{d} \sigma_{\mathrm{QCD} \times \mathrm{EW}}^{(\mathrm{N}) \mathrm{NLO}}
\end{gathered}
$$

## Without giant K-factors: stable predictions for mVV



# Giant K-factors and EW corrections: MET 



## inclusive

- in WW at large MET>MW: W's are forced off-shell

- very large NLO QCD corrections (back-to-back opens up)
arge gamma-induced
(also here Bremsstrahlung opens up back-to-back)


# Giant K-factors and EW corrections: MET 




## Conclusions

- NNLO QCD + NLO EW available in MATRIX+OpenLoops for all (massive) diboson processes
- soon public
- $V$ +gamma in the making
- QCD uncertainties at NNLO often reach few percent level.
- EW corrections enhanced at high energies reaching several tens of percent.
- In observables subject to 'giant K-factors': QCD+EW vs. QCDxEW introduces O(I) uncert.
- Can be cured via jet-veto.
- Relevant contribution of photon-induced processes
- Open issues:
- When measuring diboson processes at large pTVI/MET/mVV should always a jet veto be considered? Increased sensitivity to aTGCs?
- How to obtain reliable inclusive predictions? In particular relevant for background simulations.
- MEPS@NLO multi-jet merging including EW corrections (see $\mathrm{V}+$ jets, I 5 I I .08692)
$\Rightarrow$ how to retain NNLO QCD precision?
- How to estimate NNLO EW - O( $\left.\boldsymbol{\alpha}^{2}\right)$ ?


## BACKUP

## Relevance of EW higher-order corrections I

Numerically $\mathcal{O}(\alpha) \sim \mathcal{O}\left(\alpha_{s}^{2}\right) \Rightarrow$ NLO EW ~NNLO QCD
Possible large (negative) enhancement due to soft/collinear logs from virtual EW gauge bosons:



Universality and factorisation: [Denner, Pozzorini; '0 I ]

$$
\delta \mathcal{M}_{\mathrm{LL}+\mathrm{NLL}}^{1-\mathrm{loop}}=\frac{\alpha}{4 \pi} \sum_{k=1}^{n}\left\{\frac{1}{2} \sum_{l \neq k} \sum_{a=\gamma, Z, W^{ \pm}} I^{a}(k) I^{\bar{a}}(l) \ln ^{2} \frac{\hat{s}_{k l}}{M^{2}}+\gamma^{\mathrm{ew}}(k) \ln \frac{\hat{s}}{M^{2}}\right\} \mathcal{M}_{0}
$$

$\rightarrow$ overall large effect in the tails of distributions: $\mathrm{PT}, \mathrm{m}_{\mathrm{inv}}, \mathrm{H}_{\mathrm{T}}, \ldots$

## Relevance of EW higher-order corrections II

## Real photon radiation

- soft/coll. photon unresolved
- needed to cancel QED singularities



## Photon initial states

- QED factorisation needed to absorb IS photon singularities
- possible strong enhancement, e.g. for V



## Real W,Z,h radiation (HBR)

- partial cancellation with virtual Sudakov logs (KLN theorem not applicable)
(strongly dependent on experimental selection)
- free from singularities $\Rightarrow$ separate processes
- themselves receive large virtual EW corrections
\& inclusion requires care (double-counting issues)


