

# Direct Detection of sub-GeV Dark Matter with strong interactions through a light mediator

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**Light Dark World 2019**

ESI, Vienna, 12.08.2019

In collaboration with

- Rouven Essig
- Chris Kouvaris
- Mukul Sholapurkar

Based on

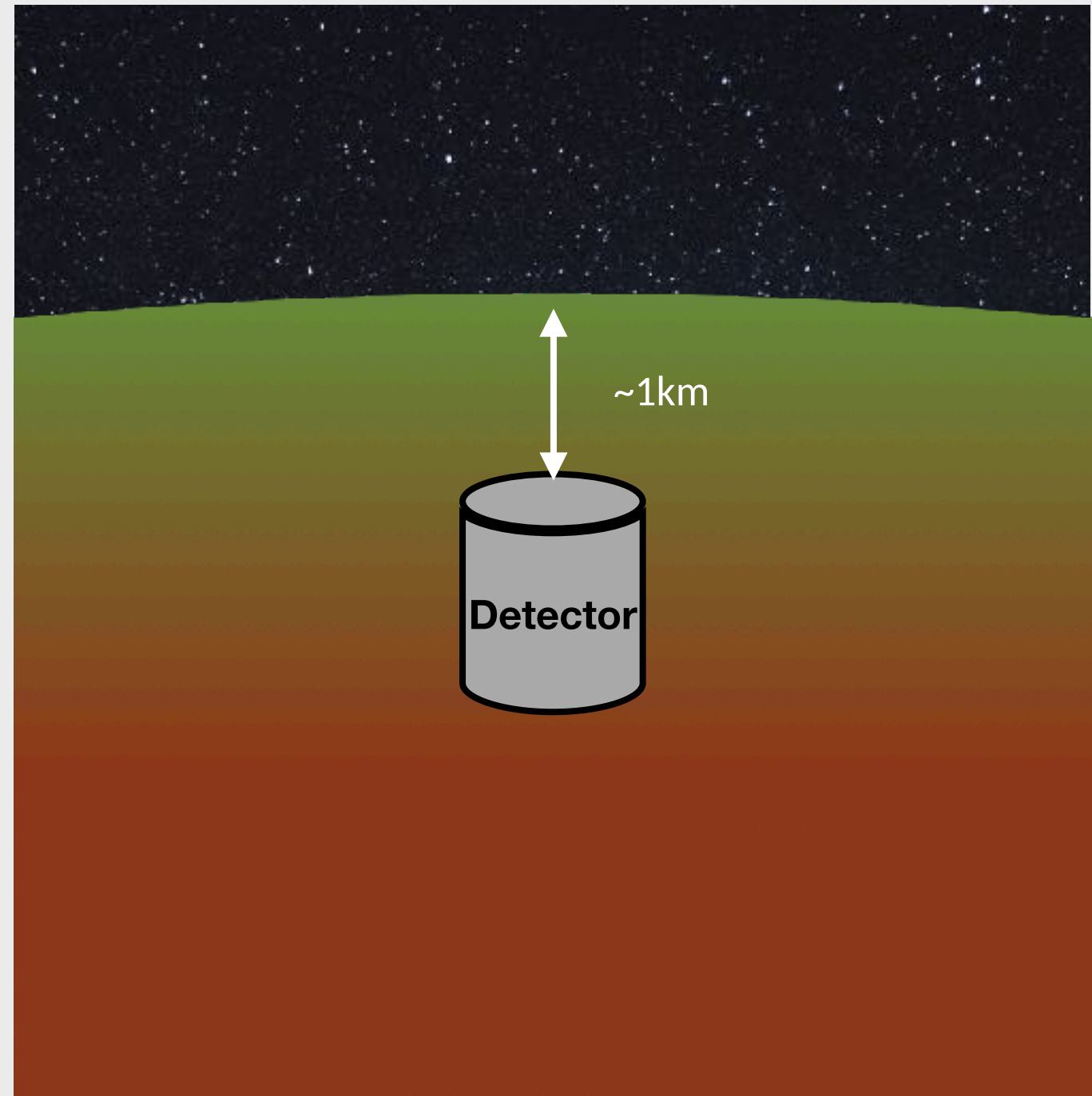
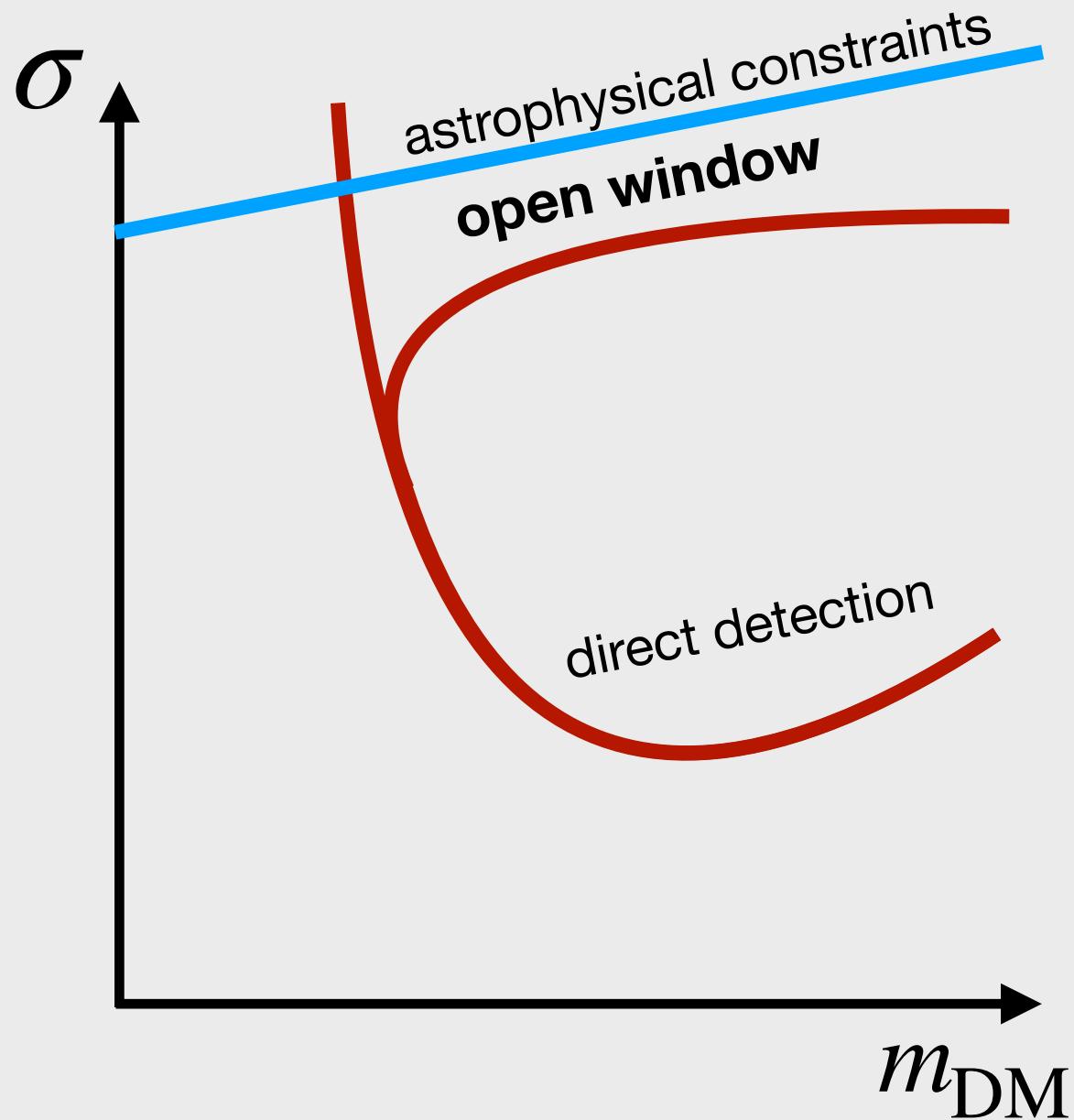
- [[arXiv:1905.06348](https://arxiv.org/abs/1905.06348)]
- [[arXiv:1802.04764](https://arxiv.org/abs/1802.04764)]



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UNIVERSITY OF TECHNOLOGY

# Direct detection of strongly interacting DM

Goodman and Witten, Phys.Rev. D31 (1985) 3059  
Starkman et al, Phys.Rev. D41 (1990) 3594



## Opening the window on strongly interacting dark matter

Glenn D. Starkman and Andrew Gould  
*Institute for Advanced Study, Princeton, New Jersey 08540*

Rahim Esmailzadeh  
*Center for Particle Astrophysics, University of California, Berkeley, California 94720*

Savas Dimopoulos\*  
*CERN TH-Division, 1211 Geneva 23, Switzerland*

Cracking Open the Window for Strongly Interacting Massive Particles as the Halo Dark Matter

Paul J. Steinhardt

A Window

What about **sub-GeV** DM searches via electron recoils?

What about **light mediators**?

Can we find a low-mass window?

Opening the Window on Strongly Interacting Dark Matter with IceCube

Ivone F. M. Albuquerque (1 and 2), Carlos Pérez de los Heros (3) ((1) Center for Particle Astrophysics FERMIAR, (2) Universidade de São Paulo, São Paulo, Brazil, (3) Department of Physics and Astronomy, University of Illinois at Urbana-Champaign, Urbana, IL 61801, USA)  
 (Submitted on 8 Jan 2010 (v1), last revised 16 Feb 2010 (this version, v2))

Closing the Window on Strongly Interacting Dark Matter

Closing the window on  $\sim$ GeV Dark Matter with moderate ( $\sim \mu$ b) interaction with nucleons

M. Shafi Mahdawi, Glennys R. Farrar

(Submitted on 1 Sep 2017 (v1), last revised 20 Dec 2017 (this version, v3))

# Outline

- I. Simulating terrestrial effects on direct detection experiments**
- II. Nuclear recoil experiments**
- III. Sub-GeV DM and light mediators**

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## Part I

# Simulating terrestrial effects on direct detection experiments

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# How to include the effect of Earth scatterings?

J. Kopp et al., Phys. Rev. D80 (2009) 083502  
 R. Essig et al., Phys. Rev. D85 (2012) 076007  
 R. Essig et al., Phys. Rev. Lett. 109 (2012) 021301  
 R. Essig et al., JHEP 1605 (2016) 046

- Nuclear recoil spectrum:

$$\frac{dR}{dE_R} = \frac{1}{m_N} \frac{\rho_\chi}{m_\chi} \int_{v > v_{\min}(E_R)} dv v f_\chi(v) \frac{d\sigma_N}{dE_R}$$

- Energy spectrum for DM induced atomic ionizations:

$$\frac{dR_{\text{ion}}}{dE_e} = \frac{1}{m_N} \frac{\rho_\chi}{m_\chi} \sum_{nl} \frac{\sigma_e}{8\mu_{\chi e}^2 E_e} \int dq q \int_{v > v_{\min}(\Delta E_e, q)} dv \frac{f_\chi(v)}{v} |F_{\text{DM}}(q)|^2 |f_{\text{ion}}^{nl}(k', q)|^2$$

- Energy spectrum for DM induced excitations in semiconductors:

$$\frac{dR_{\text{crystal}}}{dE_e} = \frac{\rho_\chi}{m_\chi} \frac{1}{M_{\text{cell}}} \frac{\sigma_e \alpha m_e^2}{\mu_{\chi e}^2} \int dq \frac{1}{q^2} \int_{v > v_{\min}(E_e, q)} dv \frac{f_\chi(v)}{v} |F_{\text{DM}}(q)|^2 |f_{\text{crystal}}(E_e, q)|^2$$

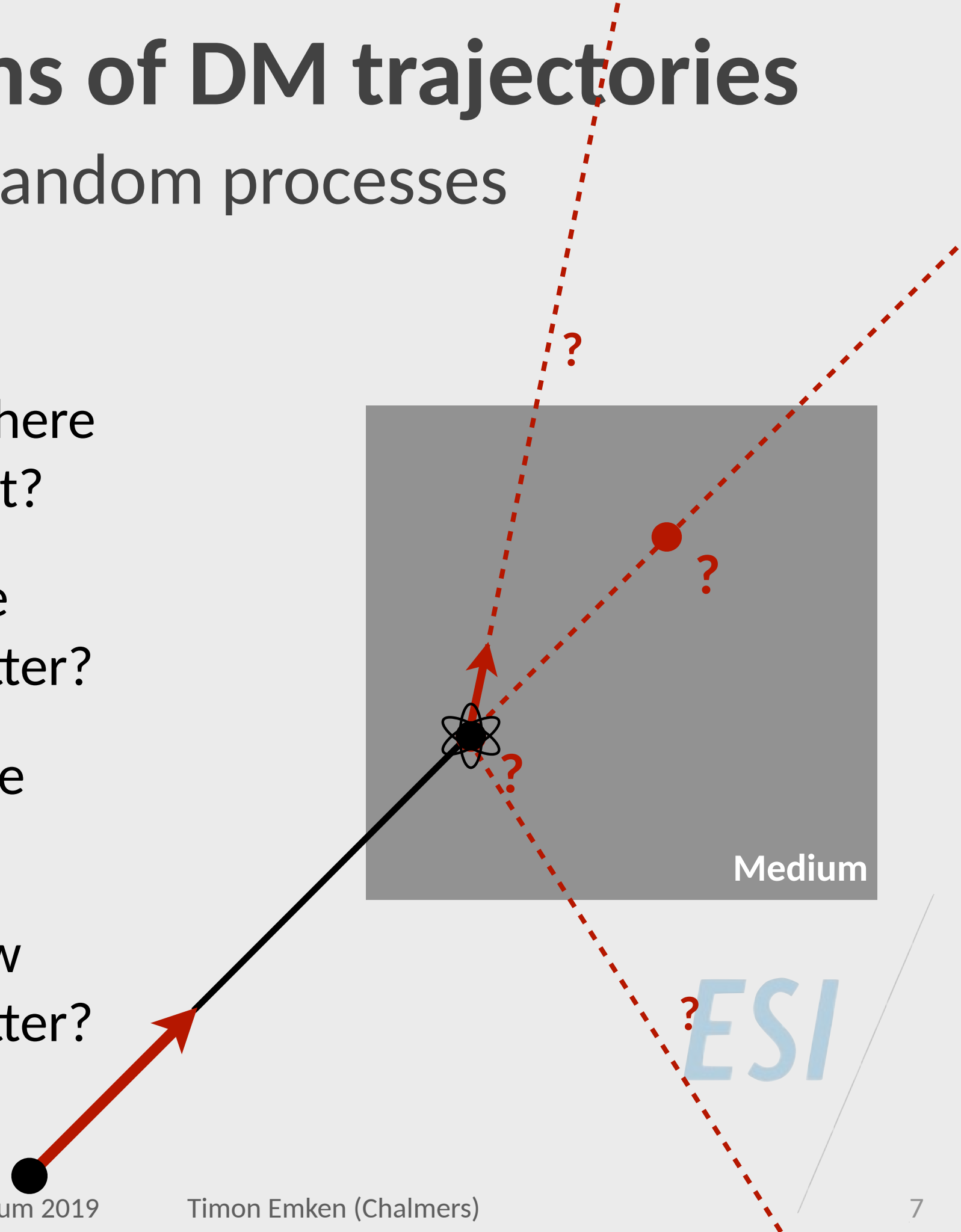
Earth scatterings modify the underground DM **density** and **velocity distribution**.

# MC simulations of DM trajectories

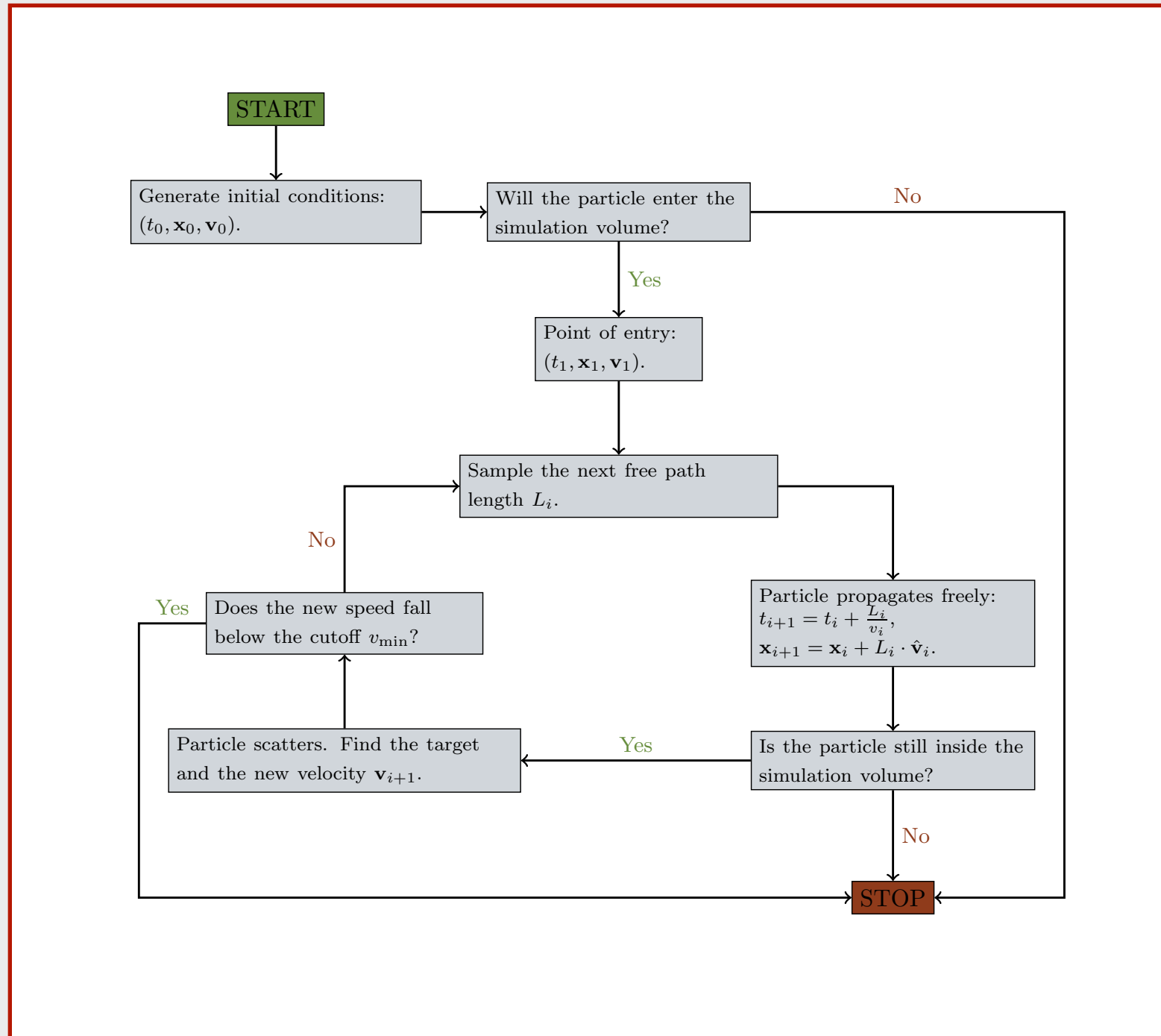
The fundamental random processes

1. **Initial Conditions:** Where does the particle start?
2. **Free distance:** Where does the particle scatter?
3. **Target:** What does the particle scatter on?
4. **Scattering angle:** How does the particle scatter?

**Repeat steps 2.-4.**



# MC simulation algorithm



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$t=0$  s



Emken 2016

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# MC simulation of the overburden of detectors

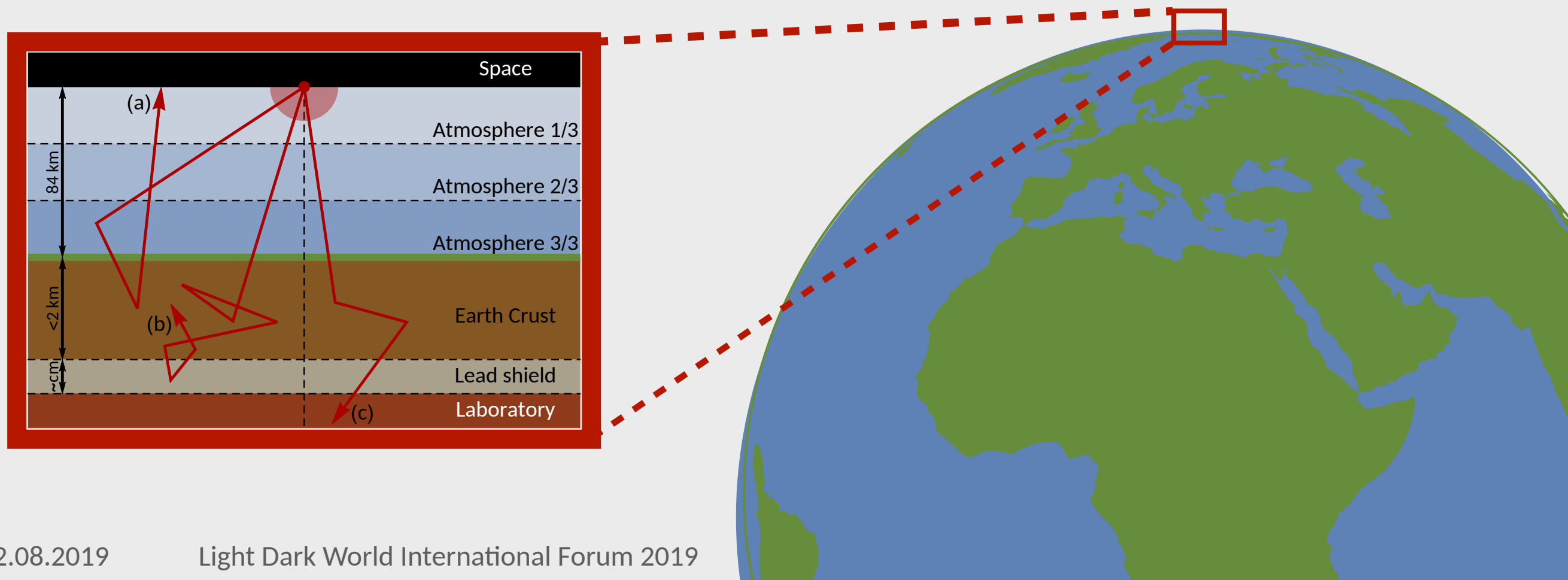
TE, C. Kouvaris, I.M. Shoemaker, Phys.Rev. D96 (2017) no.1, 015018

M.S. Mahdawi, G.R. Farrar, JCAP 1712 (2017) 004

TE, C. Kouvaris, Phys.Rev. D97 (2018) no.11, 115047

M.S. Mahdawi, G.R. Farrar, JCAP 1810 (2018) no.10, 007

To find the **critical** cross-section, where a given experiment loses sensitivity to strongly interacting DM, we only simulate the overburden, not the entire Earth.



## Part II

# Nuclear recoil experiments

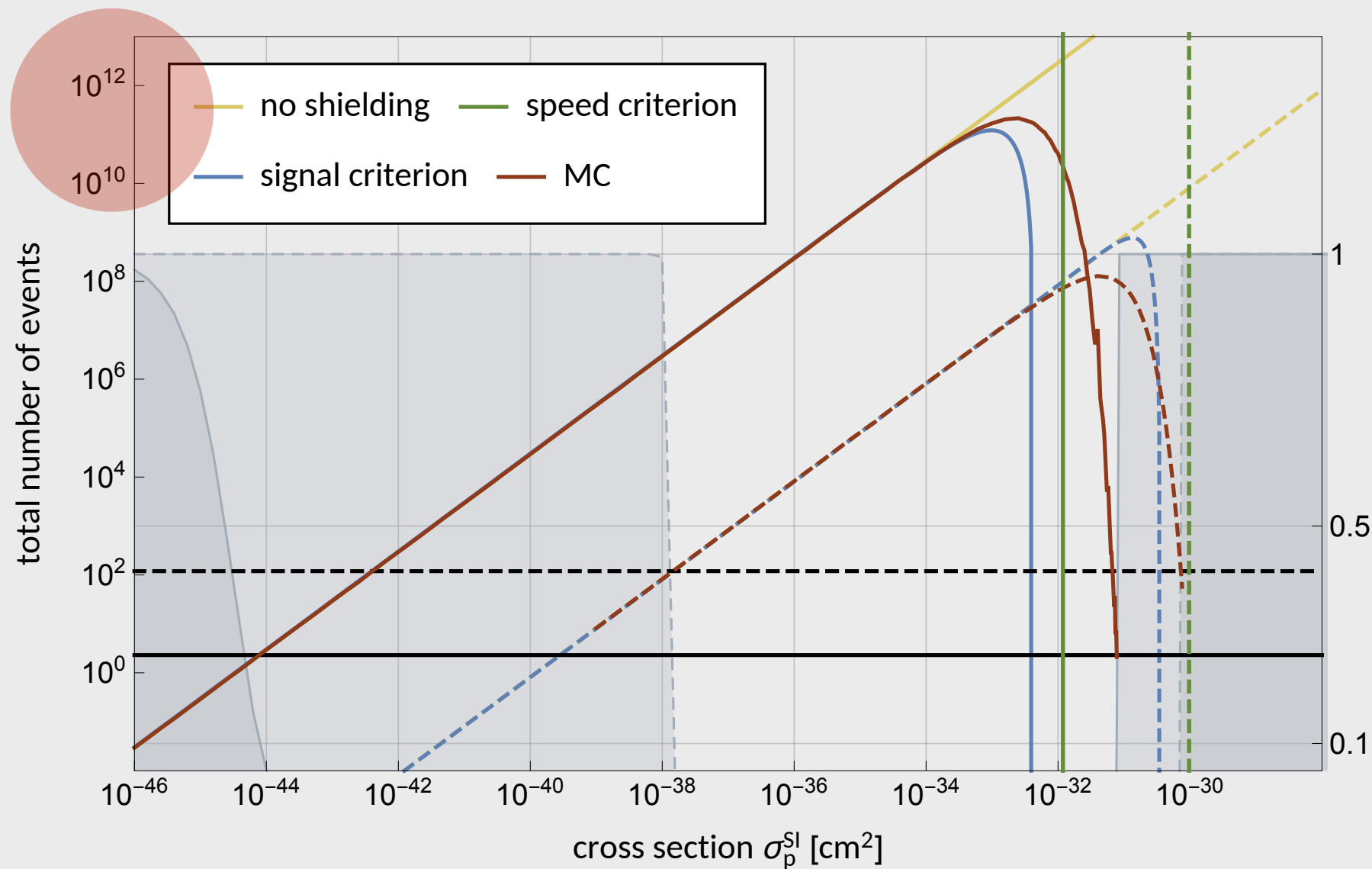
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# Shielding *vs.* Detection



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# Shielding *vs.* Detection



Rare-event techniques are absolutely crucial!

- Importance Sampling
- Importance Splitting

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## Rare event simulation I

# Importance Sampling

M.S. Mahdawi, G.R. Farrar, JCAP 1712 (2017) 004

- Rare event technique, which modifies the PDFs of the simulation.

$$\langle Y \rangle_I = \int_I dx Y(x) f(x) = \int_I dx Y(x) \frac{f(x)}{\hat{g}(x)} \hat{g}(x)$$

- Try to “mimic” the successful runs by introducing a bias into the simulations.
- Compensate by a weight factor.

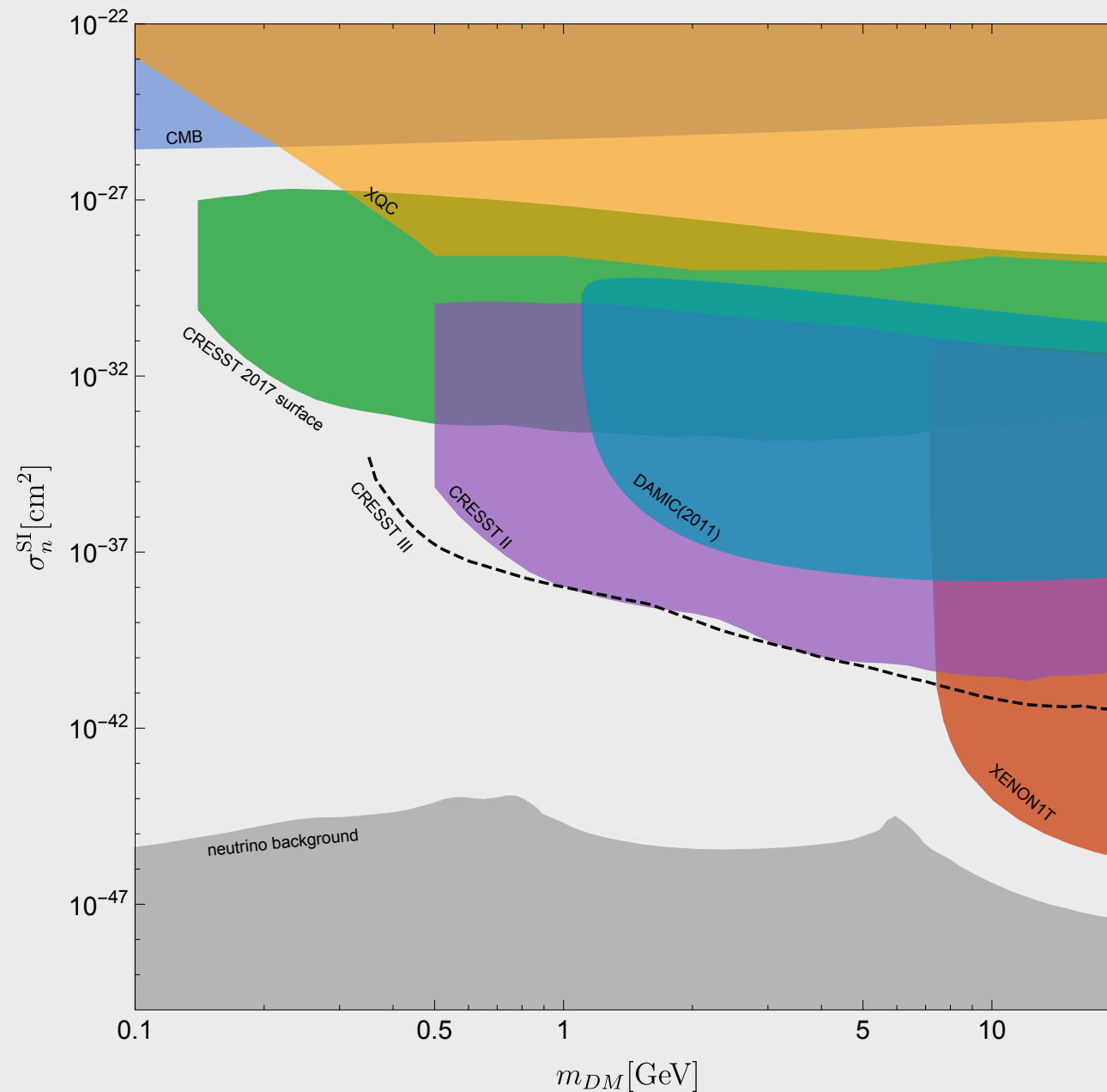
$$f_\lambda(x) = \frac{1}{\lambda} \exp\left(-\frac{x}{\lambda}\right)$$
$$g_\lambda(x) = \frac{1}{(1 + \delta_\lambda)\lambda} \exp\left(-\frac{x}{(1 + \delta_\lambda)\lambda}\right)$$

$$f_\theta(\cos \theta) = \frac{1}{2}$$
$$g_\theta(\cos \theta) = \frac{1 + \delta_\theta \cos \theta}{2}$$

Statistical weight:  $w_{\lambda,i} = \frac{f_\lambda(l_i)}{g_\lambda(l_i)}$

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# Constraints on the DM-nucleon scattering cross-section



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## Part III

# Sub-GeV DM and light mediators

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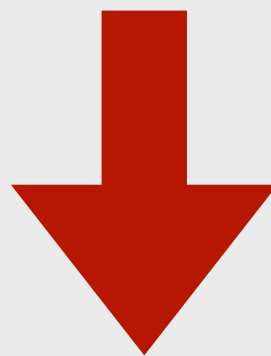


# Including DM-electron scatterings

The incoming DM flux gets attenuated by

1. Elastic nuclear scatterings.
2. Elastic DM-electron scatterings.
3. Inelastic DM-electron scatterings (ionizations/excitations).

detection process  $\neq$  attenuation/stopping process



We need a model.

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# Including DM-electron scatterings

## The Dark Photon Model

- Extend the SM by a DM particle and a U(1) gauge group with kinetic mixing.

$$\mathcal{L}_D = \bar{\chi}(i\gamma^\mu D_\mu - m_\chi)\chi + \frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} + m_{A'}^2 A'_\mu A'^\mu + \varepsilon F_{\mu\nu}F'^{\mu\nu}$$

- For kinetic mixing with the photon, the DM couples to electric charge.

$$\frac{d\sigma_N}{dq^2} = \frac{\sigma_p}{4\mu_{\chi p}^2 v_\chi^2} F_{\text{DM}}(q)^2 F_N(q)^2 Z^2$$

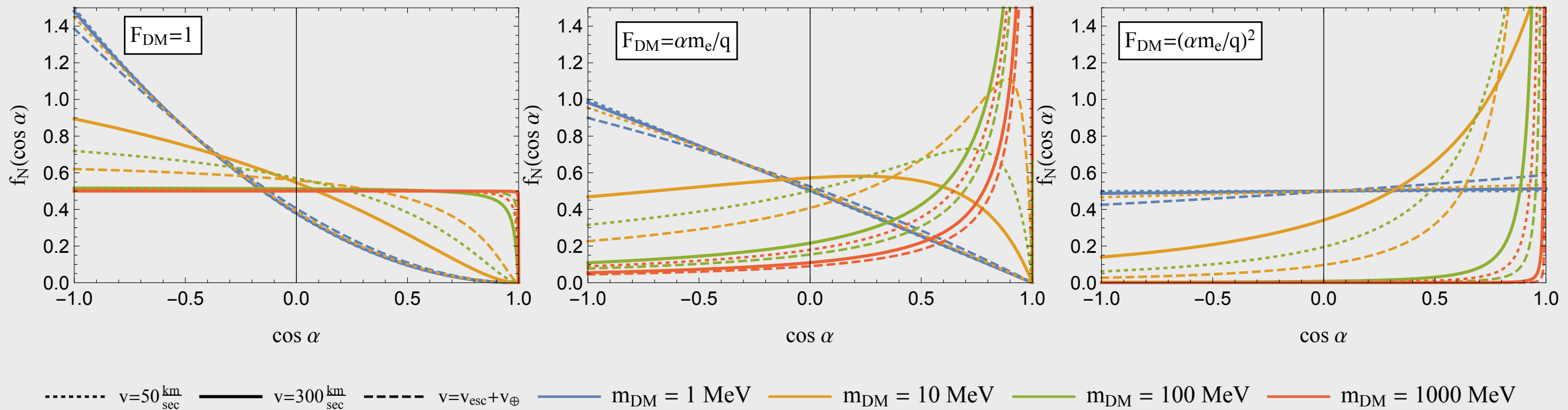
- Hierarchy between the DM-proton and DM-electron cross section:

$$\frac{\sigma_p}{\sigma_e} = \left( \frac{\mu_{\chi p}}{\mu_{\chi e}} \right)^2$$

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S.K. Lee et al, PRD92 (2015) 083517

# New scattering kinematics



DM form factor

VS

Charge screening

$$F_{\text{DM}}(q) = \begin{cases} 1, & \text{for heavy mediator,} \\ \frac{q_{\text{ref}}}{q}, & \text{for ED interactions,} \\ \left(\frac{q_{\text{ref}}}{q}\right)^2, & \text{for light mediator.} \end{cases}$$

$$F_A(q) = \frac{a^2 q^2}{1 + a^2 q^2}$$

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# Rare event simulation II

## Geometric Importance Splitting

- “More interesting” particles get split into copies.
- Requires the definition of an importance function,

$$I : \mathbb{R}^3 \rightarrow \mathbb{R}$$

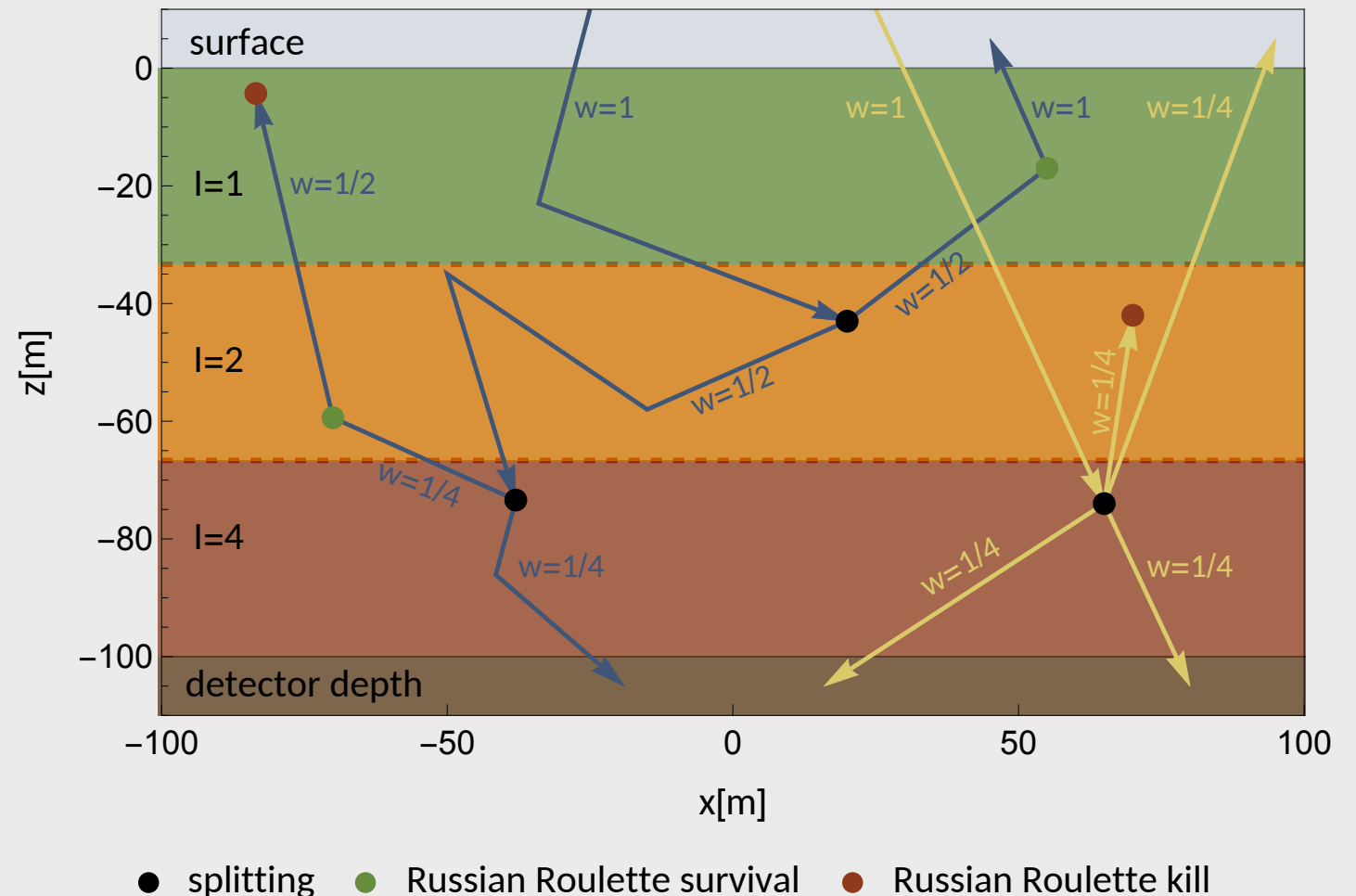
- If the importance increases,

$$\nu \equiv \frac{I_{i+1}}{I_i} > 1$$

- the particle gets split into

$$n = \begin{cases} \nu, & \text{if } \nu \in \mathbb{N}, \\ \lfloor \nu \rfloor, & \text{if } \nu \notin \mathbb{N} \wedge \xi \geq \Delta, \\ \lfloor \nu \rfloor + 1, & \text{if } \nu \notin \mathbb{N} \wedge \xi < \Delta, \end{cases}$$

copies.



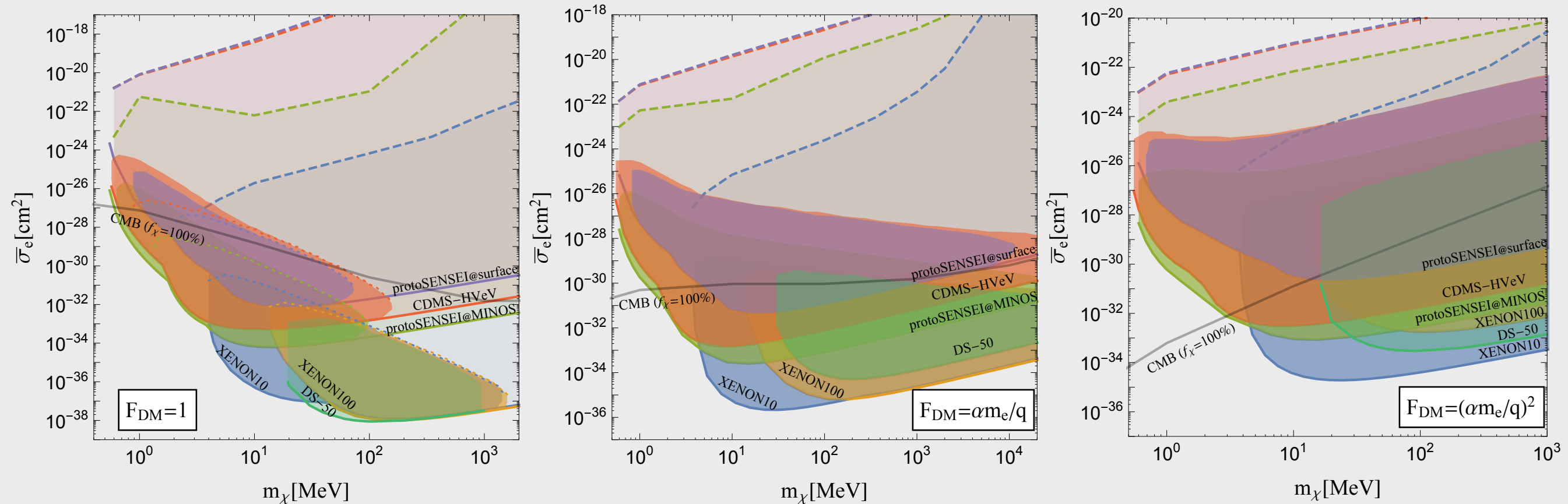
New statistical weight

$$w_{i+1} \equiv \frac{w_i}{n}$$

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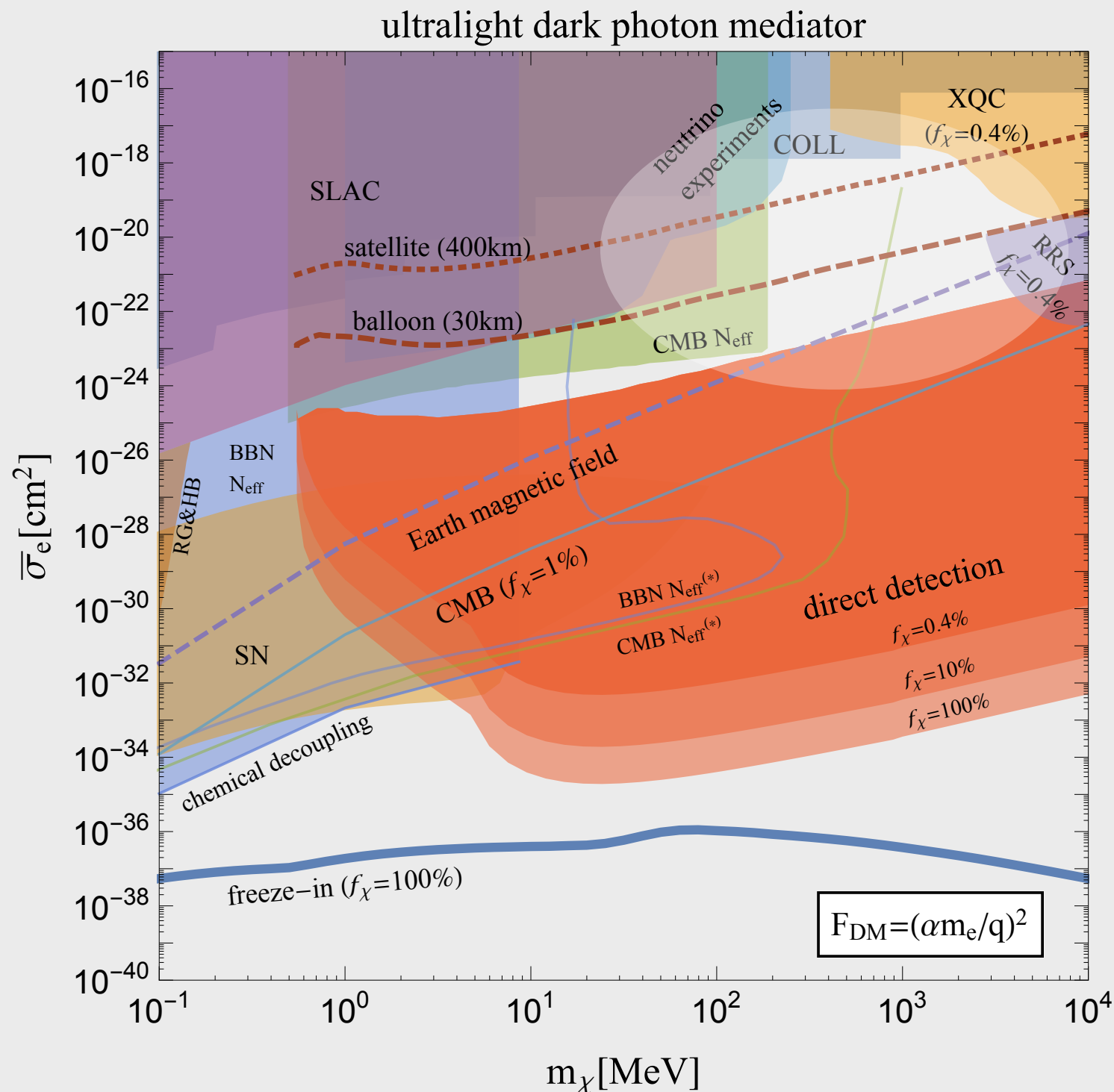
- Otherwise: Russian Roulette

# DM-Electron Scattering constraints



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# Constraints on a sub-dominant component of strongly interacting DM



- So is there an open window in parameters space?
- Probably not for milli-charged DM.
- Definitely not for

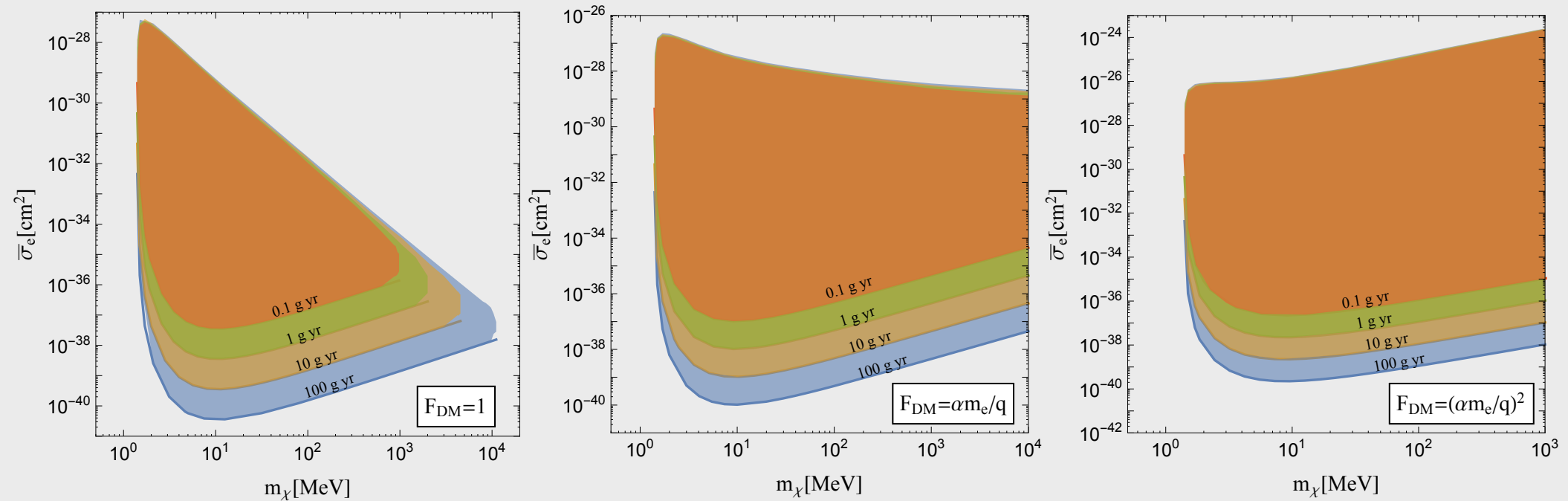
$$f_\chi = 100\%$$

- Yes, under certain conditions:
  - Sub-dominant component.
    - $f_\chi < 0.4\%$
  - Ultralight, but not massless mediator.
  - Small dark gauge coupling.

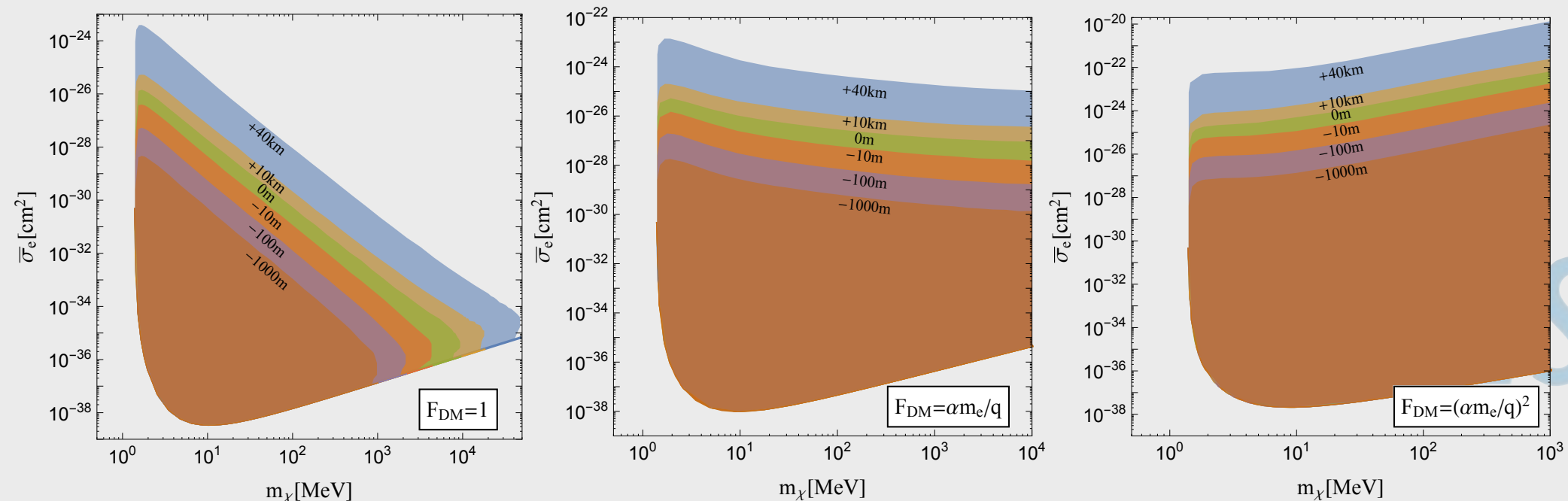
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# How to push towards **stronger** interactions

- Higher Exposures

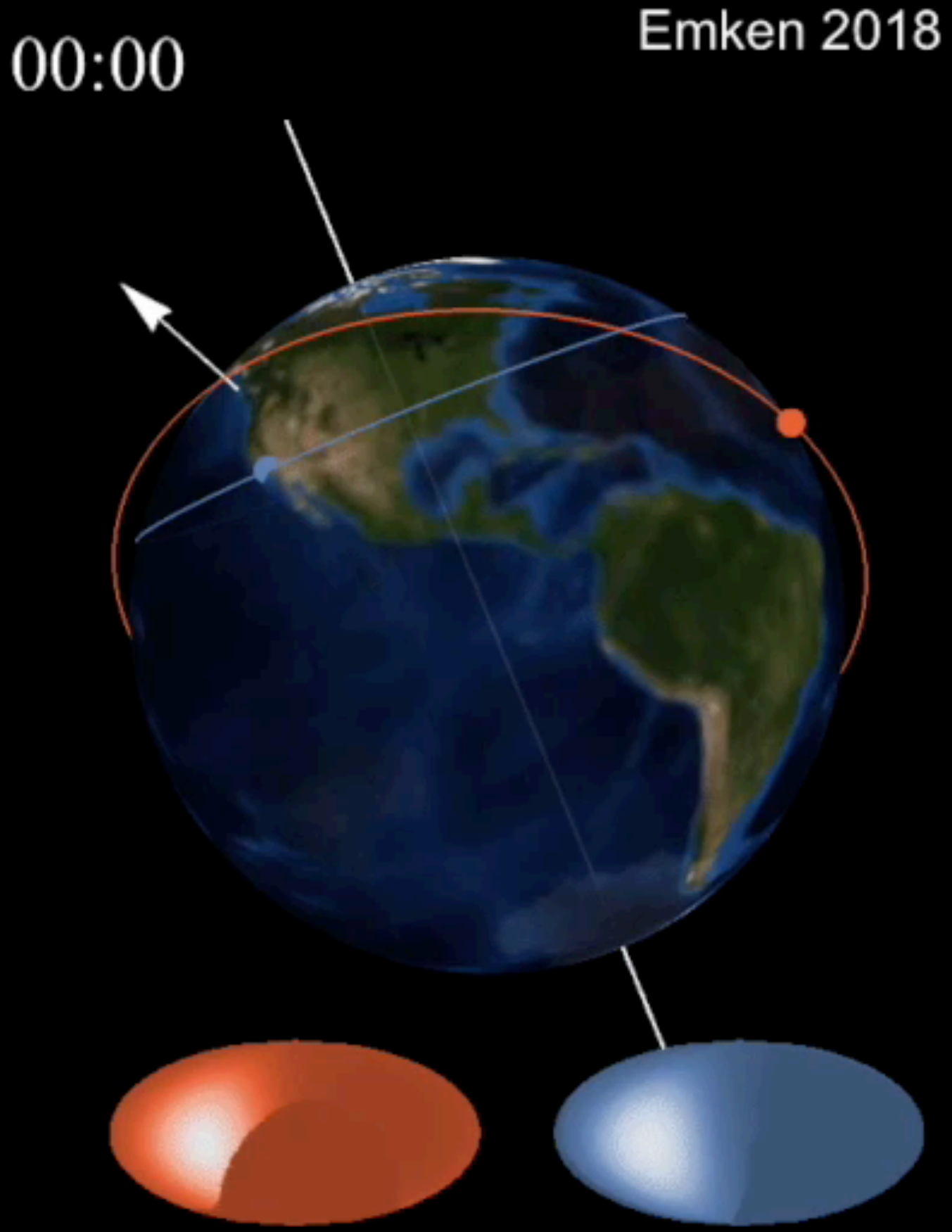
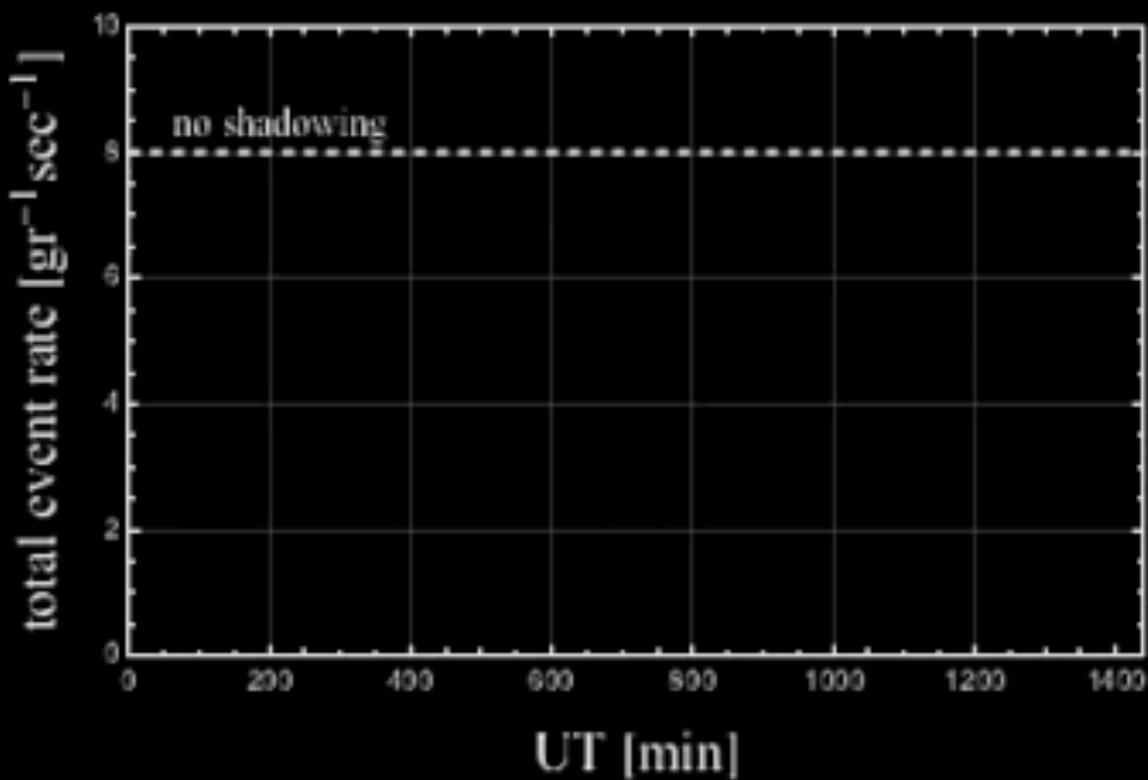
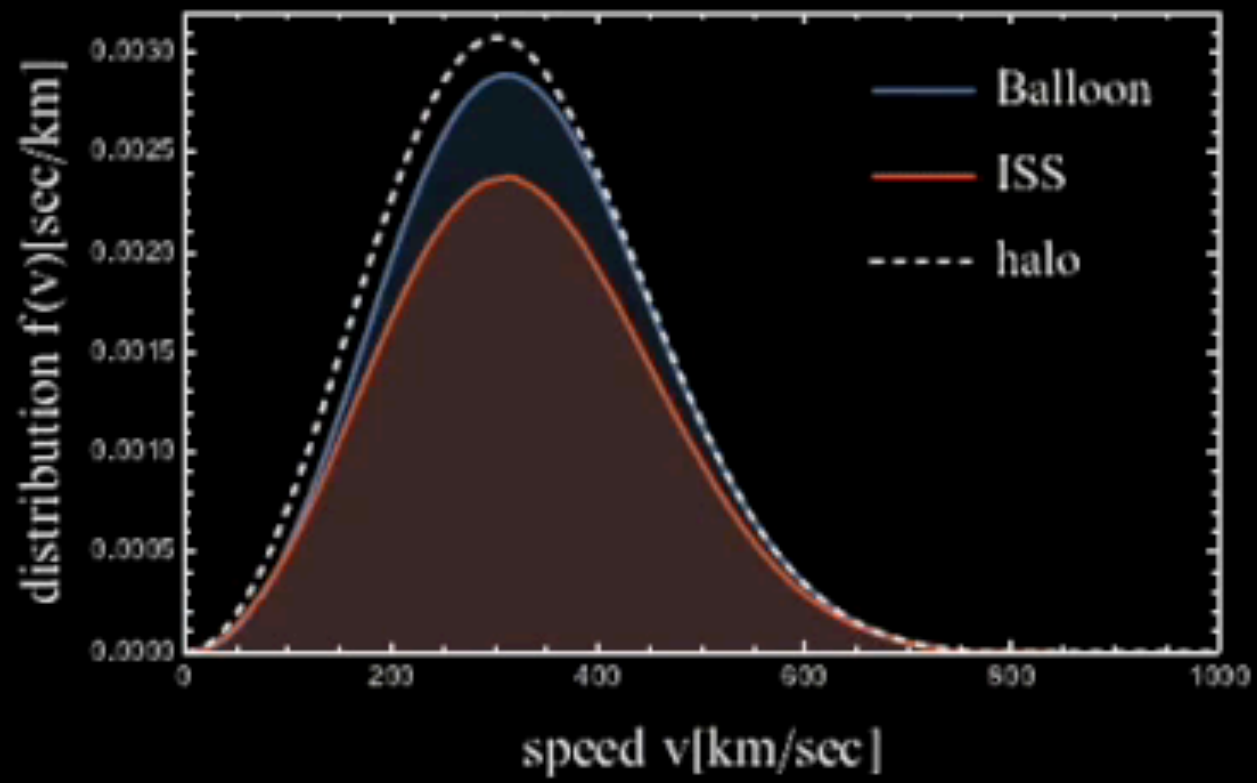


- Shallower Laboratories





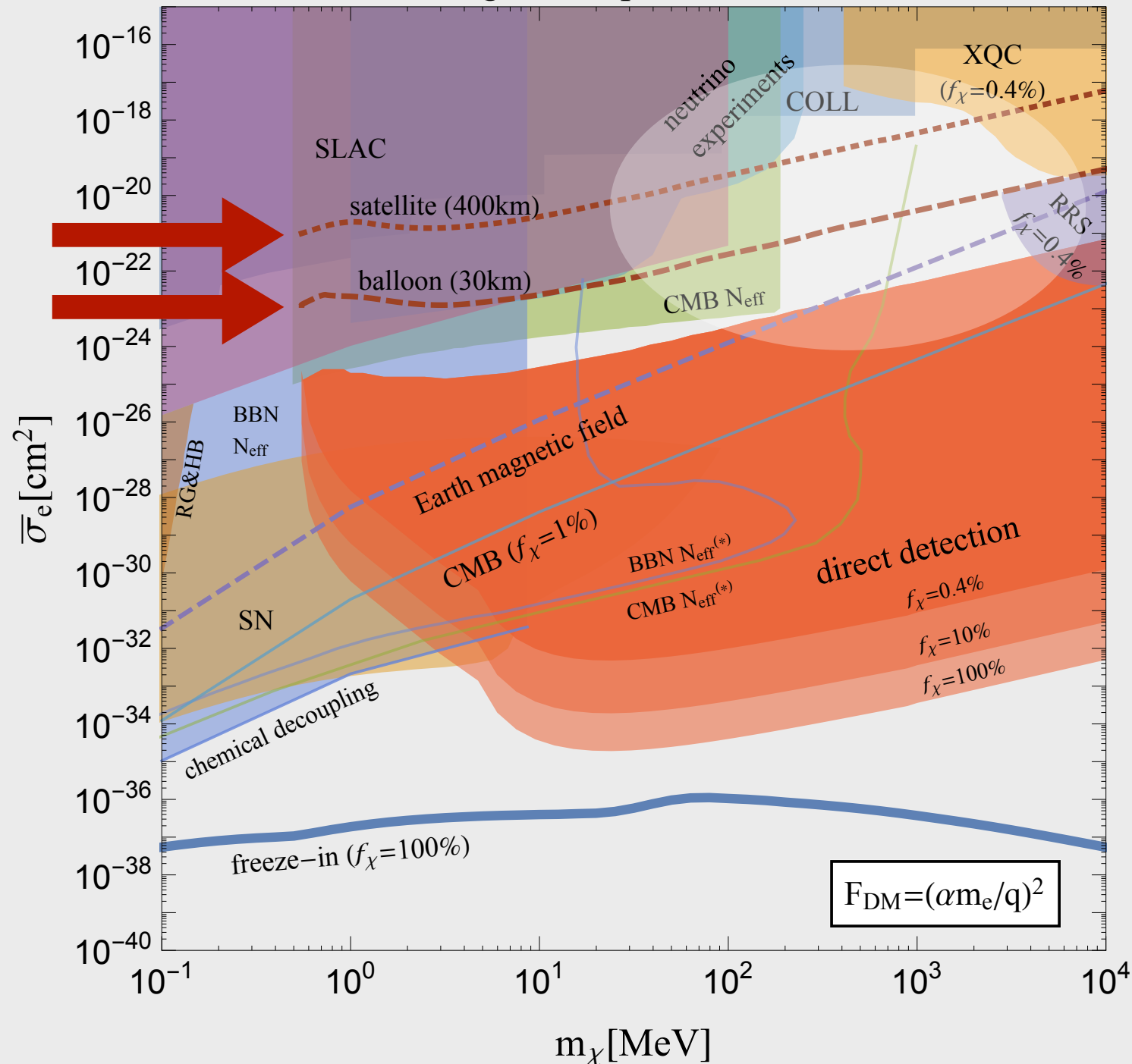
# Balloon/satellite experiments





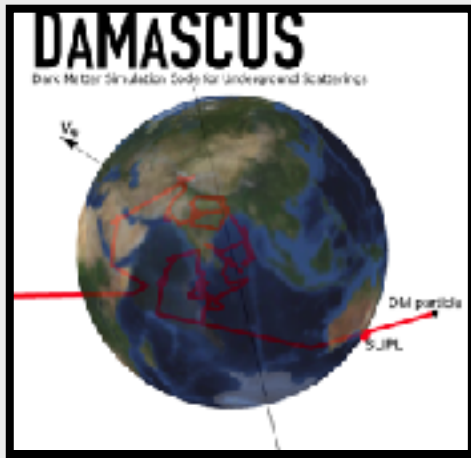
# Constraints on a sub-dominant component of strongly interacting DM

ultralight dark photon mediator



- So is there an open window in parameters space?
- Probably not for milli-charged DM.
- Definitely not for
  - $f_\chi = 100\%$
- Yes, under certain conditions:
  - Sub-dominant component.
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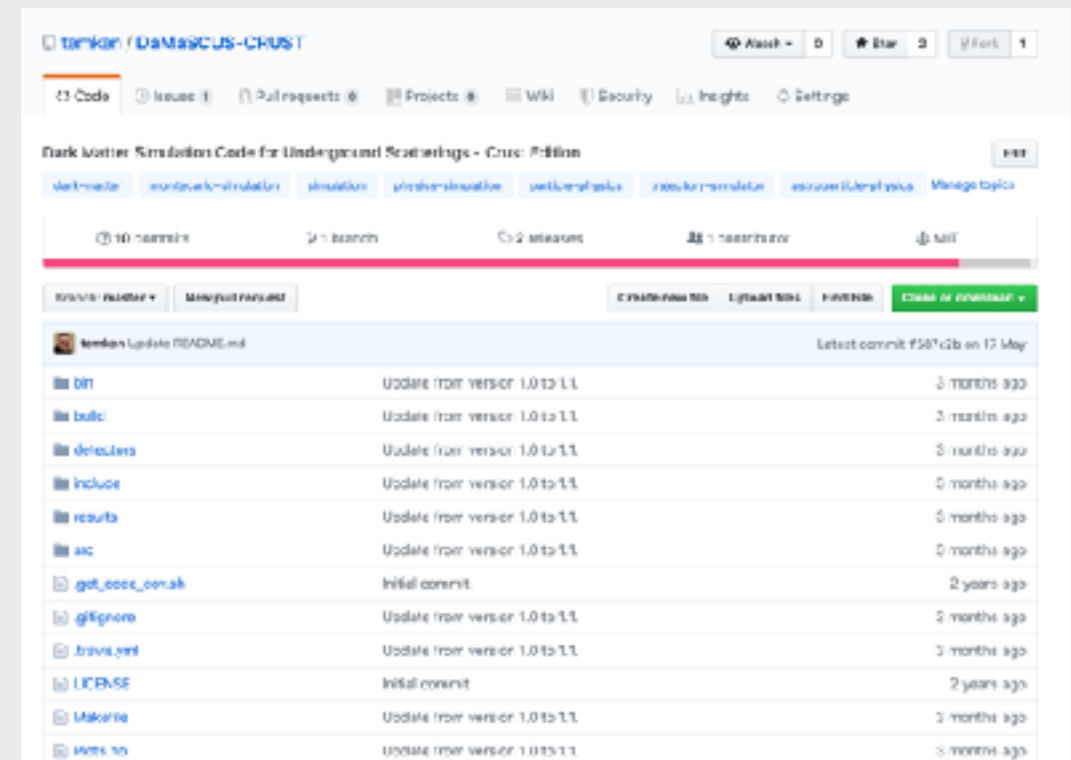
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# The DaMaSCUS code

Dark Matter Simulation Code for Underground Scatterings

- Written in C++.
- Fully parallelized with *MPI*.
- Results were generated on the ABACUS2.0 supercomputer.
- The code is public.



<http://github.com/temken/>

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# Light Dark World International Forum 2019

**Thank you!**

# Backup Slides

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# Kernel Density Estimation (KDE)

A non-parametric method to estimate an unknown PDF based on data.

For a data set  $\{x_1, x_2, \dots, x_N\}$  we can estimate the PDF via

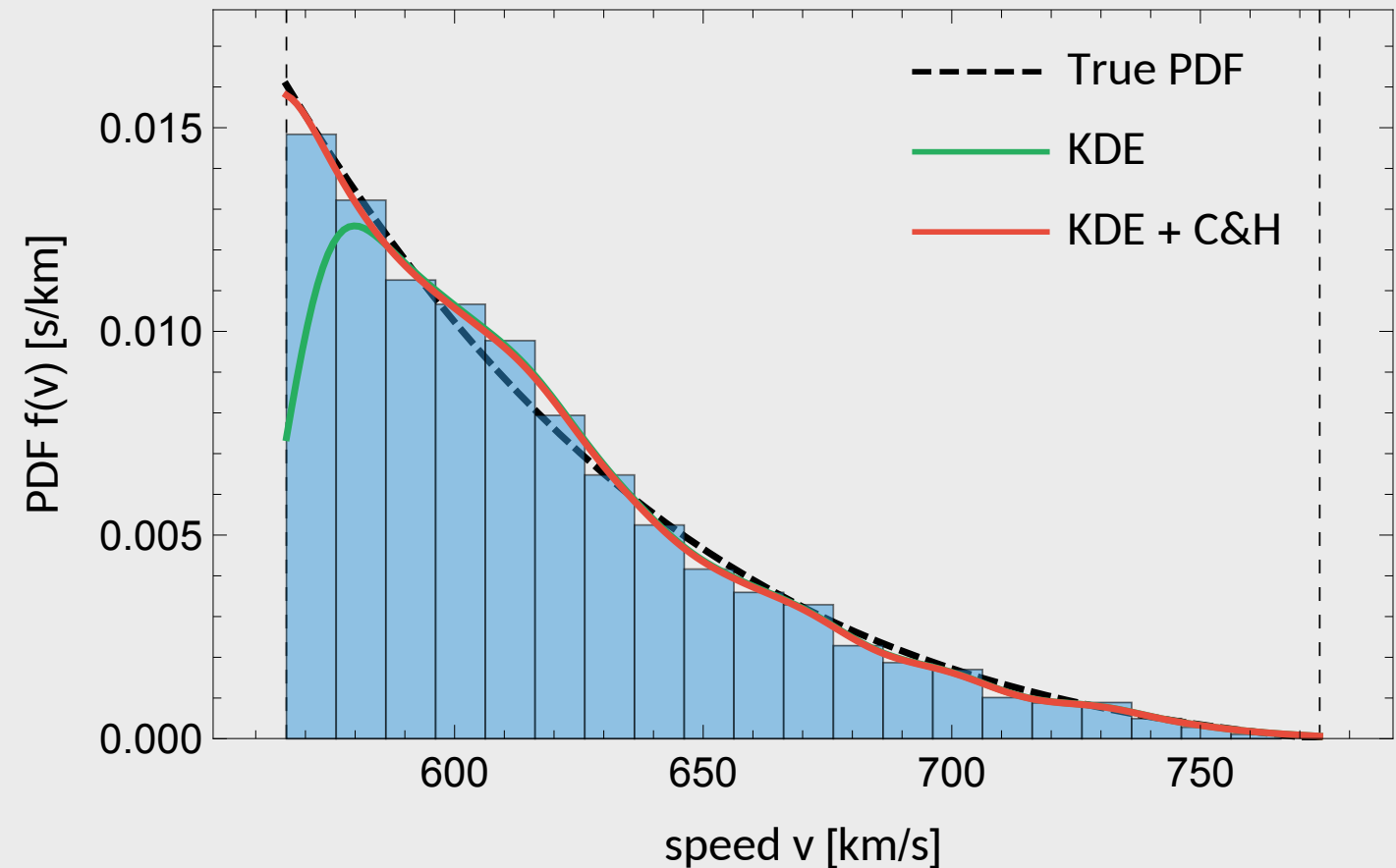
$$\hat{f}_h(x) = \frac{1}{h} \sum_{i=1}^N K\left(\frac{x - x_i}{h}\right).$$

E.g. with a Gaussian Kernel,

$$K(x) = \frac{1}{\sqrt{2\pi}} \exp\left(-\frac{x^2}{2}\right).$$

We set the bandwidth  $h$  using Silverman's rule of thumb,

$$h = \left(\frac{4}{3N}\right)^{1/5} \hat{\sigma}.$$



The bias at the domain's boundary has to be compensated, e.g. by a pseudo-data method by Cowling and Hall

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R. Karunamuni, T. Alberts, *Statistical Methodology* 2 (2005), 191

A. Cowling, P. Hall, *Journal of the Royal Statistical Society*, B58 (1996), 551