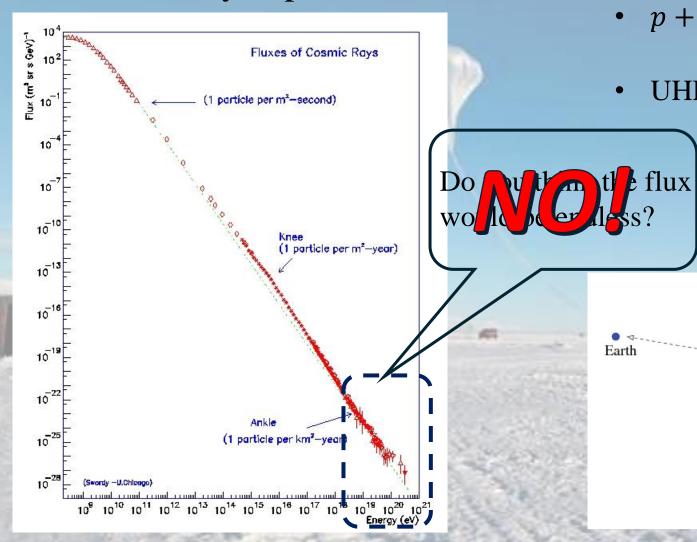
A Journey of ANITA Experiment

Po-Wei Huang
Graduate Symposium

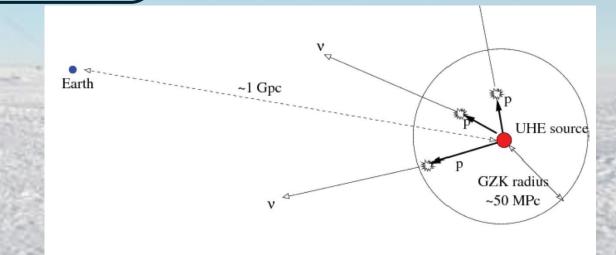
Motivation - UHE Neutrino Search

Cosmic Rays Spectrum [4]

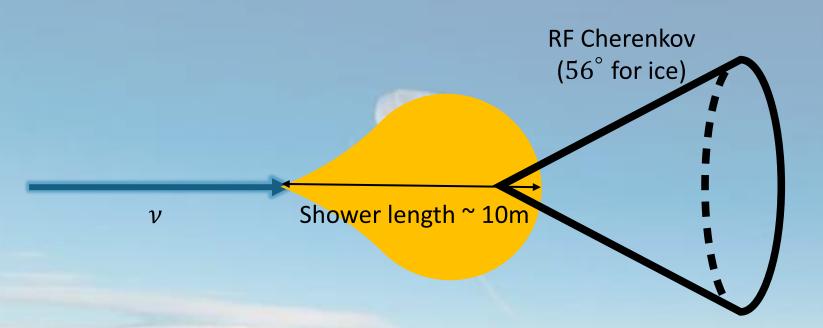


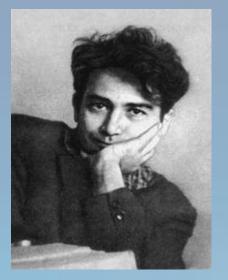
GZK Cutoff

- $p + \gamma_{CMB} \to \Delta^+ \to n + \pi^+$ [4] when $E_p > 10^{19.5} eV$
- UHE ν from π^+ decay



Detection Method – Askaryan Effect



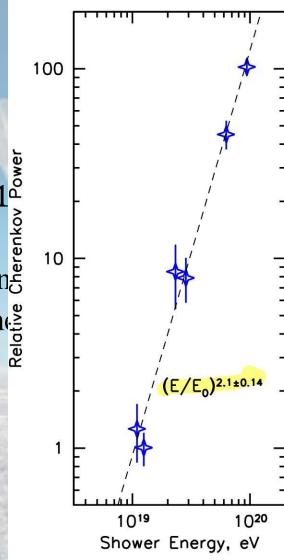


Gurgen Ashotovich Askaryan a.k.a Гурген Аскарьян (1928 -1997)

- ~20% Negative Excess Charge produce Cherenkov Radio Signals [4]
- $P_{Cherenkov} \propto N_e^2 \propto E^2$ [2]
- Size of shower depends on frequency

Test of Askaryan Effect

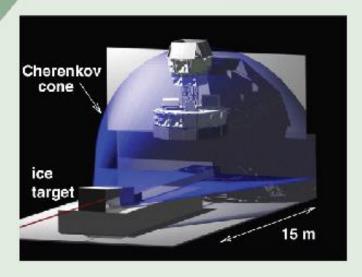
- Silica Sand (2000)
- Rock Salt (2002)
- Ice Test in SLAC (2006)
 - 28.5 *GeV e*⁻ beams with 1^a
 - 10 tons ices covering 90%
 - Angular, Frequency respons
 - Shower Energy Measurem



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3

ANITA Detector



• $N_{\nu} = flux \otimes Area \otimes \sigma_{\nu} \otimes \varepsilon_{Det} \otimes T$

○ Area matters! → Where to put detector?

GPS Antennas

Seavey Quad-ridge Horns [6] [7]

- 1.Bandpass: 200 1200MHZ
- 2. Separating Vertical & Horizontal Polarisation
- 3. $\sim 60^{\circ}$ Ø, θ field of view

Photovoltaic Array

ANITA Flights

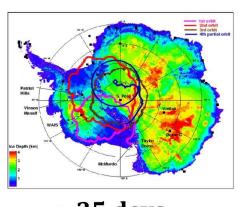
ANITA-I (2006-2007) ANITA-II (2008-2009) ANITA-III (2014-2015) ANITA-IV (2016)

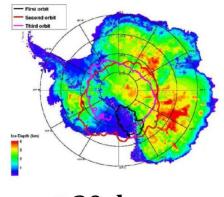


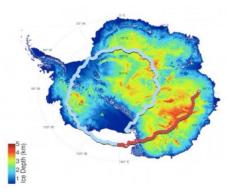


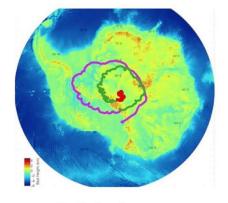










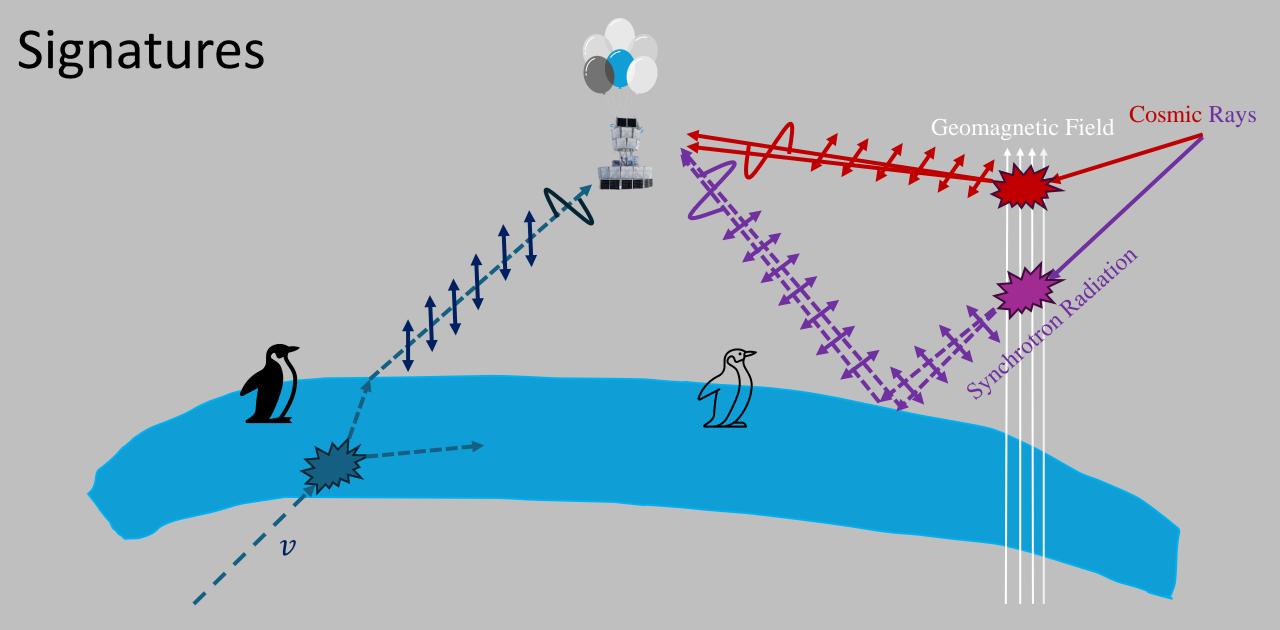


• 35 days

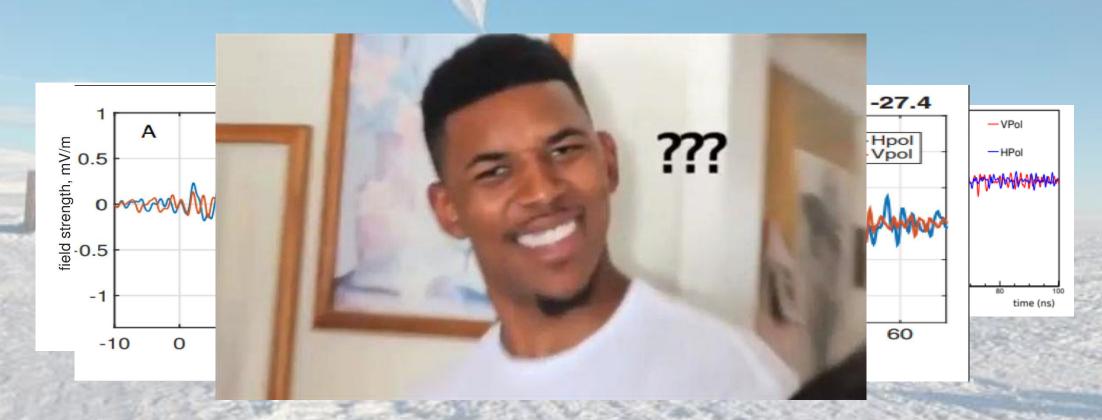
30 days

22 days

28 days

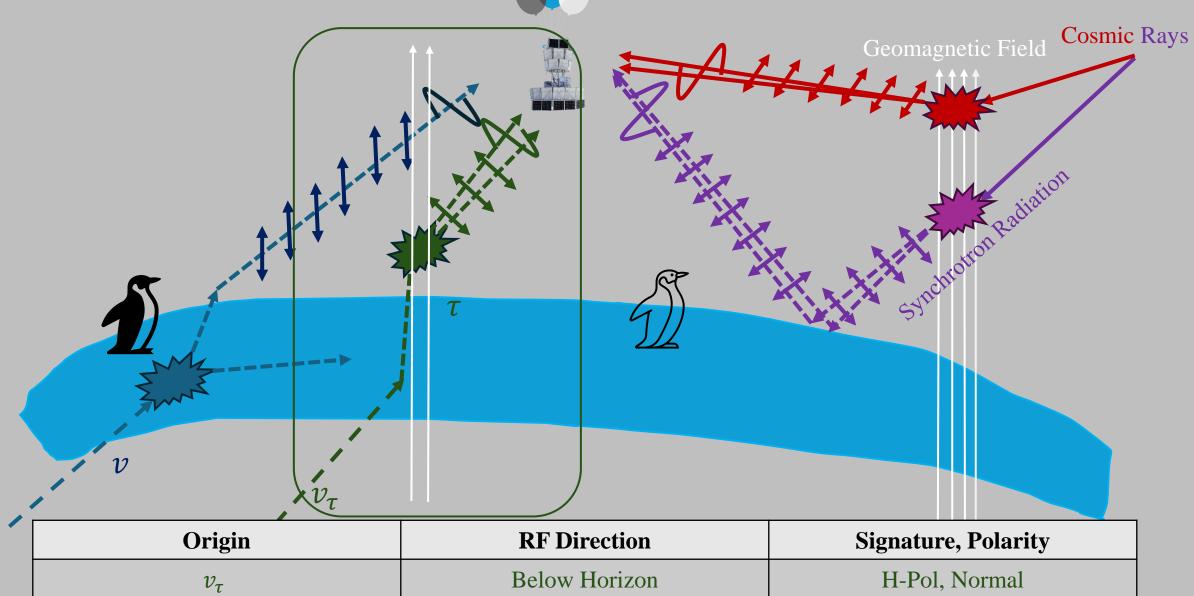


Origin	RF Direction	Signature, Polarity
CR-reflected	Below Horizon	H-Pol, Inverted
CR-Direct	Above Horizon	H-Pol, Normal
Neutrino	Below Horizon	V-Pol, Normal



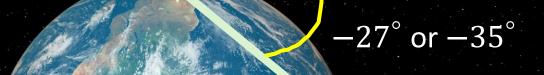
Potential Explanation





Are they really τ neutrino?

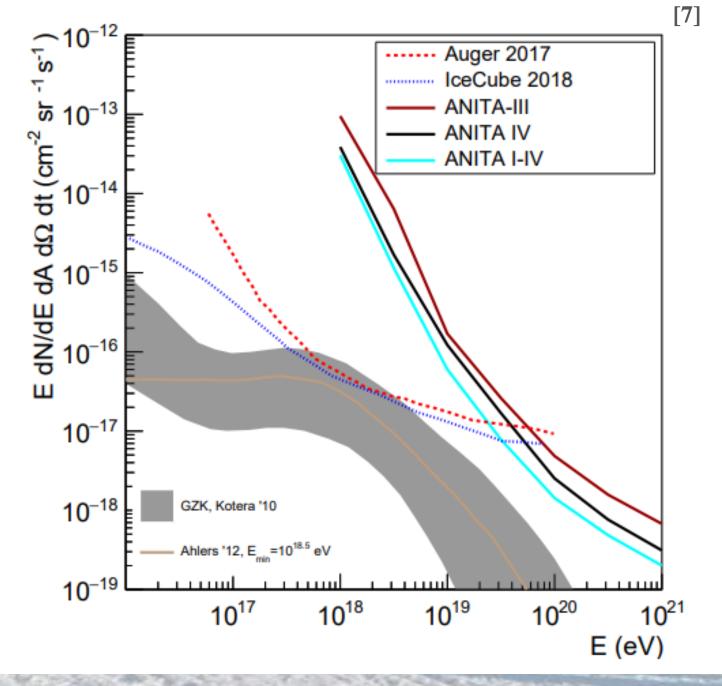
Travel 100 times more than Standard Model Prediction



	Elevation	Cord Length
Event 3985267	-27°	5500km
Event 15717147	-35°	7000km

Flux Result

• Flux limit was set under the treatment of neutrino candidate as background



Conclusion

- Motivation: UHE Neutrino Search
- Signatures to exclude direct and reflected CR
- Two anonymous events in ANITA-I and ANITA-III



Thanks for joining the ANITA journey!!!

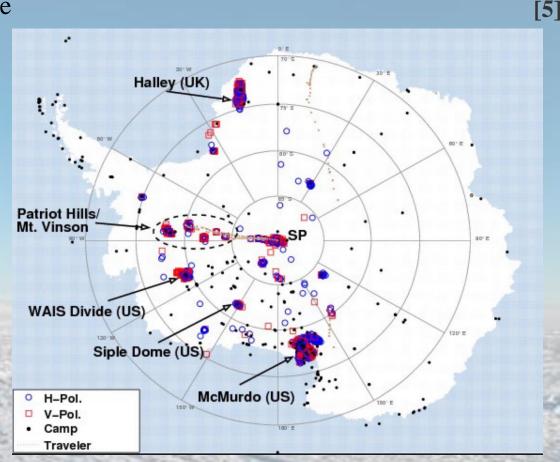
Reference

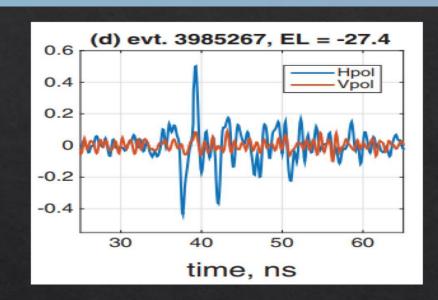
- [1] P.W. Gorham, J.Nam, Characteristics of Four Upward-Pointing Cosmic-Ray-like Events Observed with ANITA, Phys. Rev. Lett. 117, 071101 (2016)
- [2] P. W. Gorham *et al.* (ANITA Collaboration), Observations of the Askaryan Effect in Ice, Phys. Rev. Lett. 99, 171101 (2007)
- [3] Cosmin Deaconu, Recent Results from ANITA, arXiv:1810.00820 (astro-ph) (2018)
- [4] JiWoo Nam, Cosmic Rays Experiments, Lecture Slides in Instrumental Methods in Particle Physics (2023)
- [5] JiWoo Nam, Introduction Sides in Instrumental Methods in Particle Physics (2023)
- [6] M. Detrixhe *et al.* (ANITA Collaboration), Ultrarelativistic magnetic monopole search with the ANITA-II balloon-borne radio interferometer, Phys. Rev. D 83, 023513 (2011)
- [7] Cosmin Deaconu, Searches for Ultra-High Energy Neutrinos with ANITA, 36th International Cosmic Ray Conference –ICRC (2019)

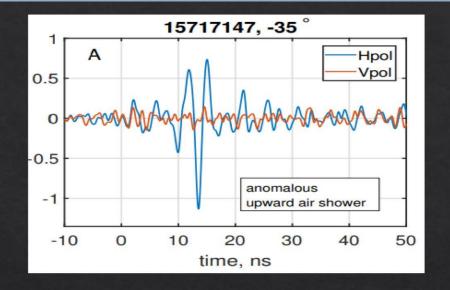


Background

- CW (Continuous Wave) from geostationary military communication satellites
 - Masking out azimuthal sectors with high trigger rate
- Thermal Noise
 - Incoherent waveform
 - Set triggering threshold, Coincidence tagging
- Anthropogenic Signals
 - Tend to cluster spatially
 - Only events that are outliers survive for final cuts
- Digitizer Glitch
- Payload Blast

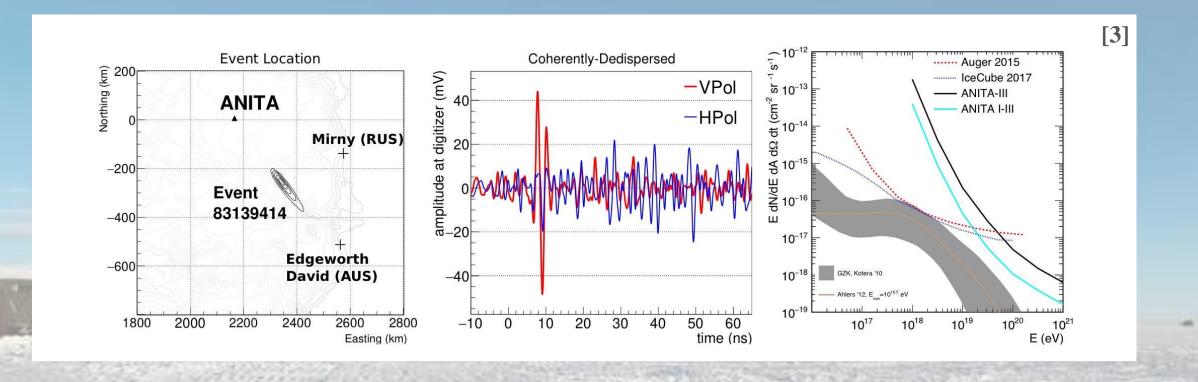






Event, flight	3985267, ANITA-I	15717147, ANITA-III
Date, time	2006-12-28	2014-12-20
El, Az	-27.4±0.3°, 159.6±0.7°	$-35.4\pm0.3^{\circ}$, $61.41\pm0.7^{\circ}$
RA, Dec	282.14,+20.33	50.78, +38.65
E _{shower}	0.6±0.4 EeV	$0.56^{+0.3}$ -0.2 EeV
Cord length trough Earth	~5500 km	~7000 km

Potential Neutrino Events In ANITA-III



Setup in SLAC Ice Experiment

