



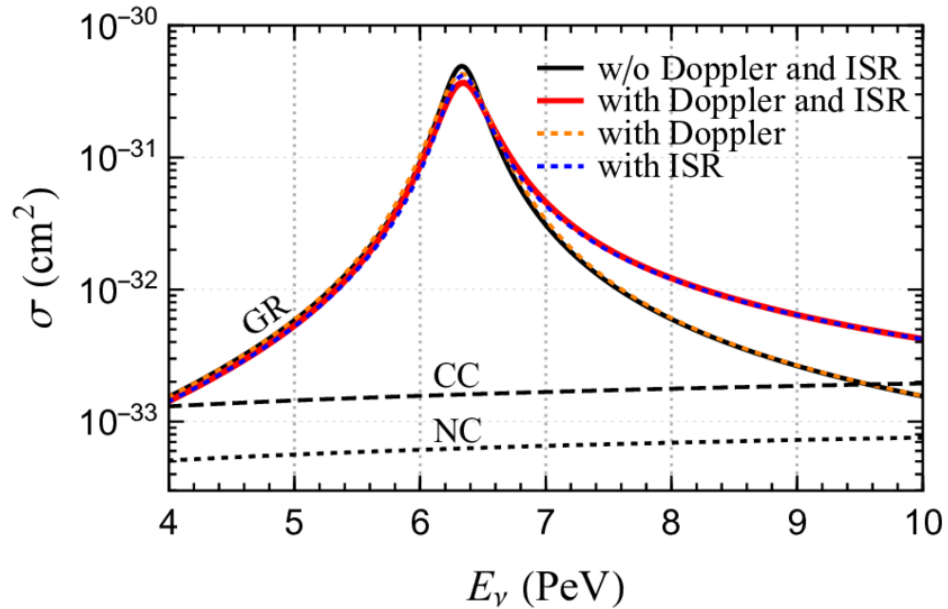
IceCube – Neutrinos from Active Galaxies

Sam Leadley

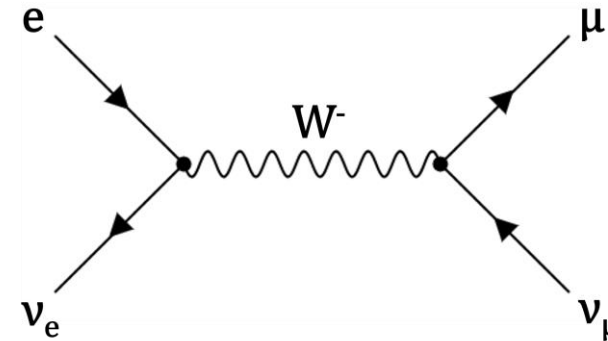
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Glashow Resonance



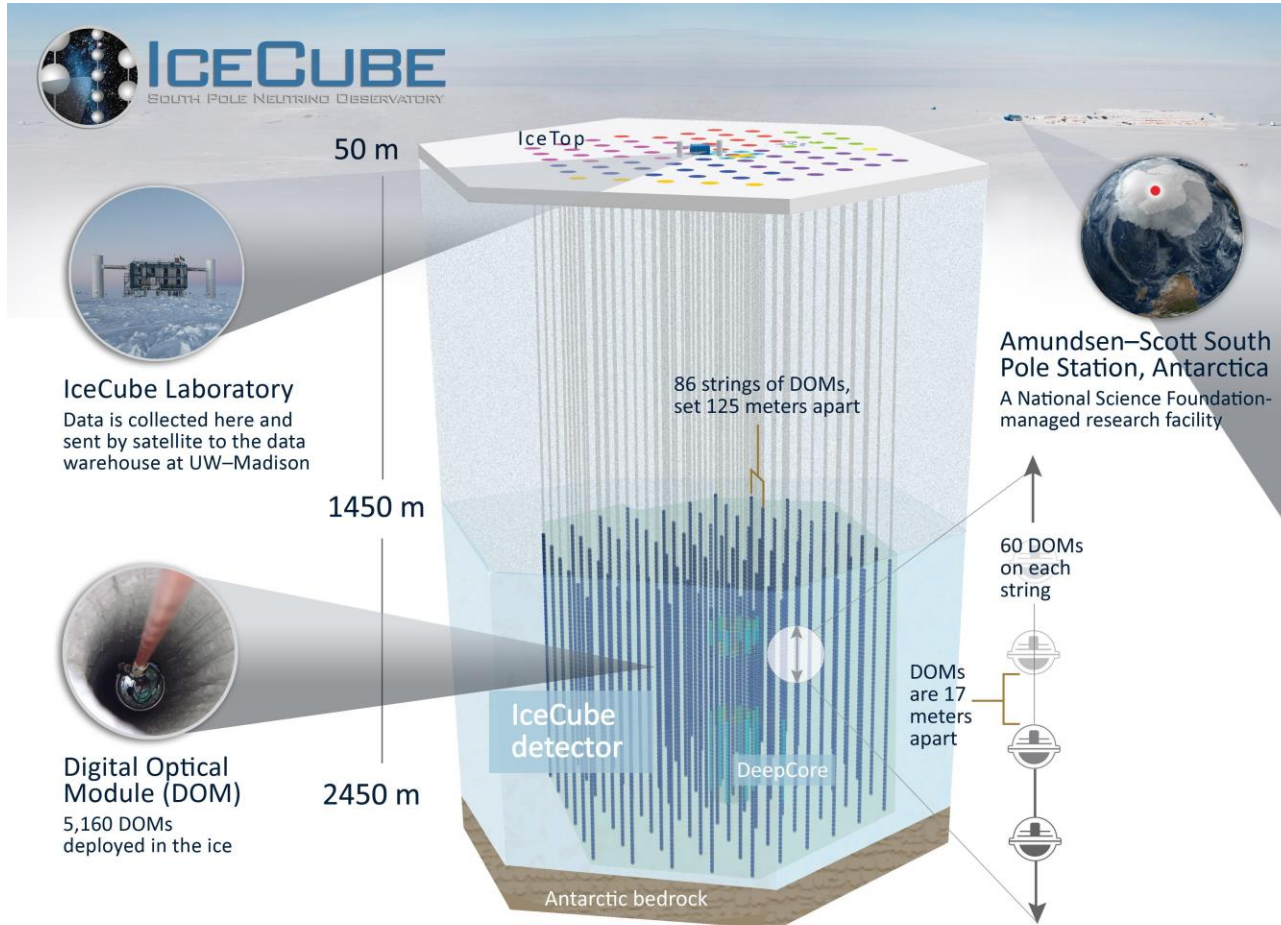
Cross section for s-channel W boson production via CC and NC DIS, and the Glashow Resonance.



Feynman diagram for the Glashow resonance process observed at IceCube.

- Possible neutrino **origins**:
 - **Atmospheric**
 - Expect $2 \cdot 10^{-7}$ events over 4.6 yrs
 - 1 event in **40 million years** (unlikely)
 - **Cosmic**
 - This event: **5.2σ**
 - More events: **$>8\sigma$**

IceCube – The Detector



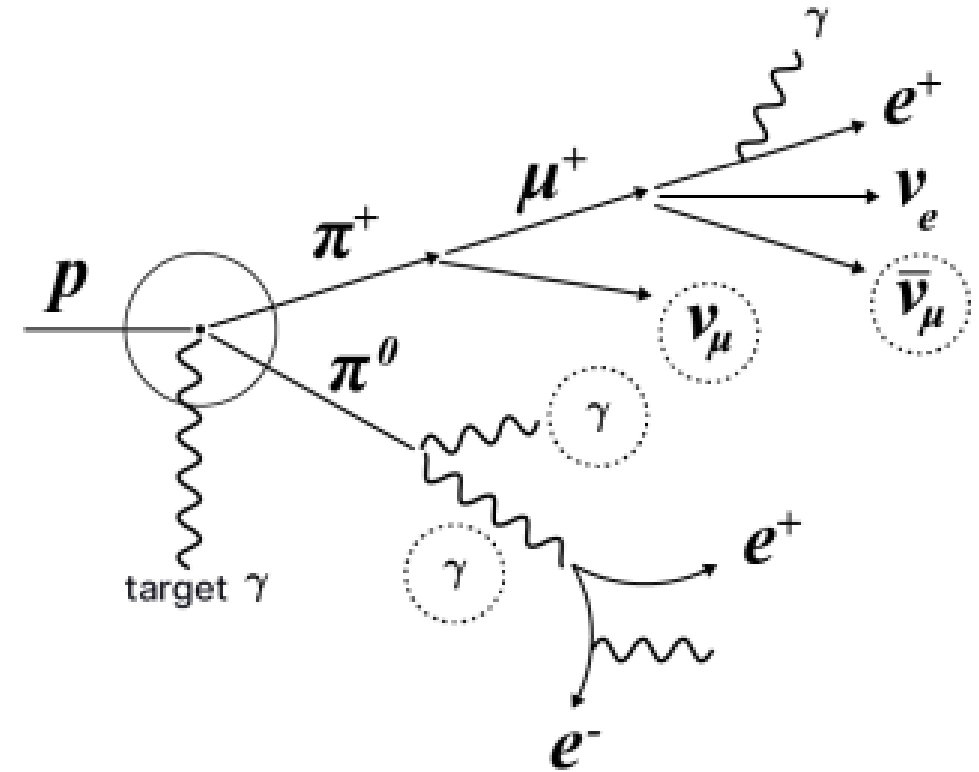
- 1km³ Čerenkov detector
 - 5160 DOMs (optical sensors)
 - 86 strings
- 1500m below the South Pole
- Most sensitive to **muons** and **muon neutrinos**
 - Penetrating
 - No scattering

Where are the neutrinos from?

- **Signatures** of cosmic neutrinos
 - High energy (**TeV/PeV**)
 - Neutrinos + lower energy **EM shower**
 - **No muon** accompanying
 - Energy spectrum of **neutrino flux**:

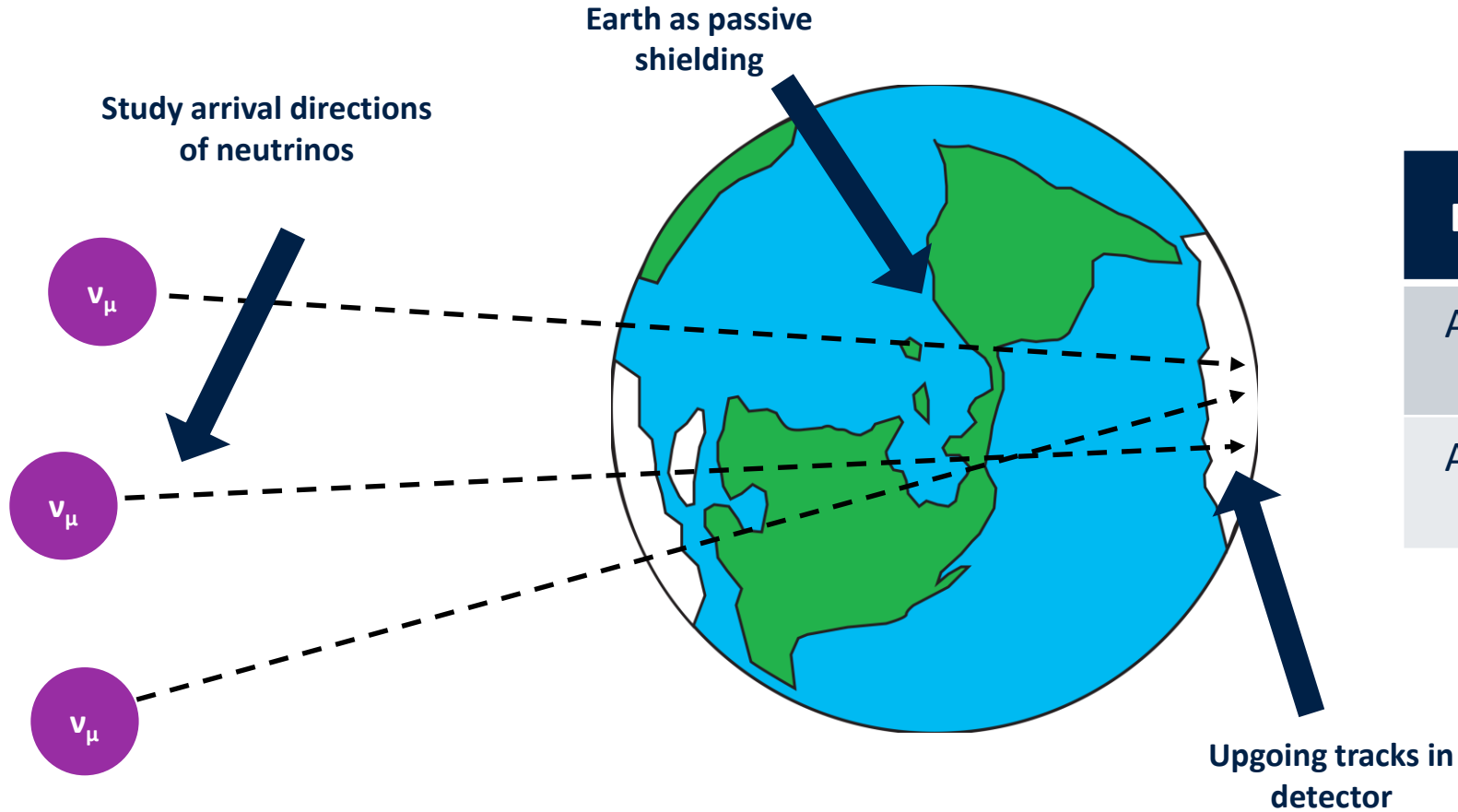
$$\frac{dN}{dE} \sim E^{-\gamma} \rightarrow \text{Calculate } \text{spectral index } \gamma$$

- **Where** are the neutrinos from?
 - GRB?
 - Supernova?
 - Elsewhere?



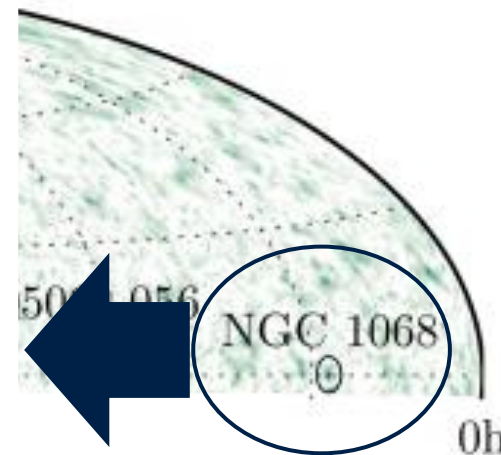
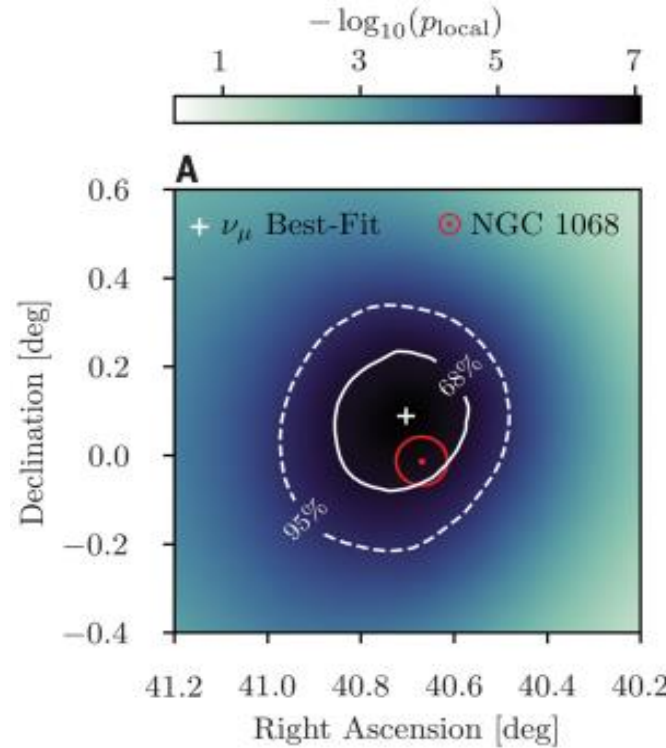
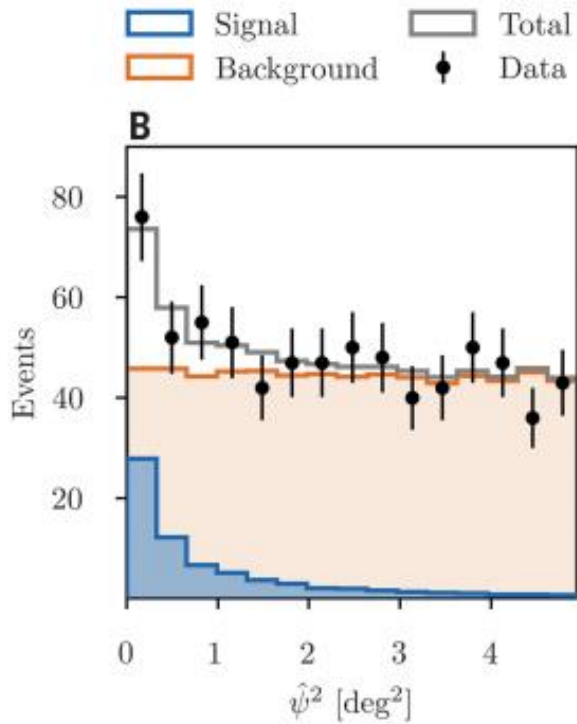
A typical interaction resulting in cosmic ray neutrinos.

Methodology



Background	Mitigation
Atmospheric muons	Upgoing tracks – use Earth as shield
Atmospheric neutrinos	Constant background, spectral index (γ)

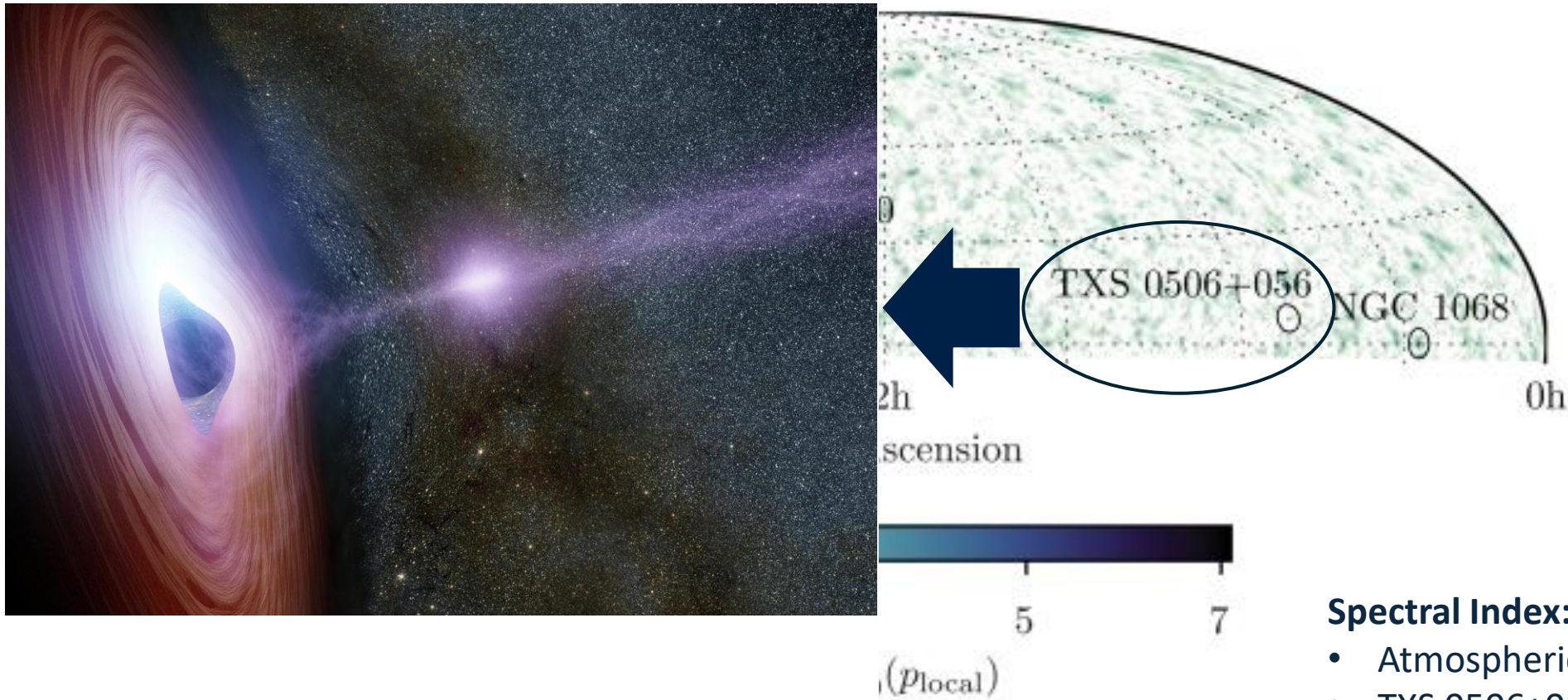
Results



Spectral Index:

- Atmospheric γ : 3.7
- M77 γ : 3.2 ± 0.2

Results



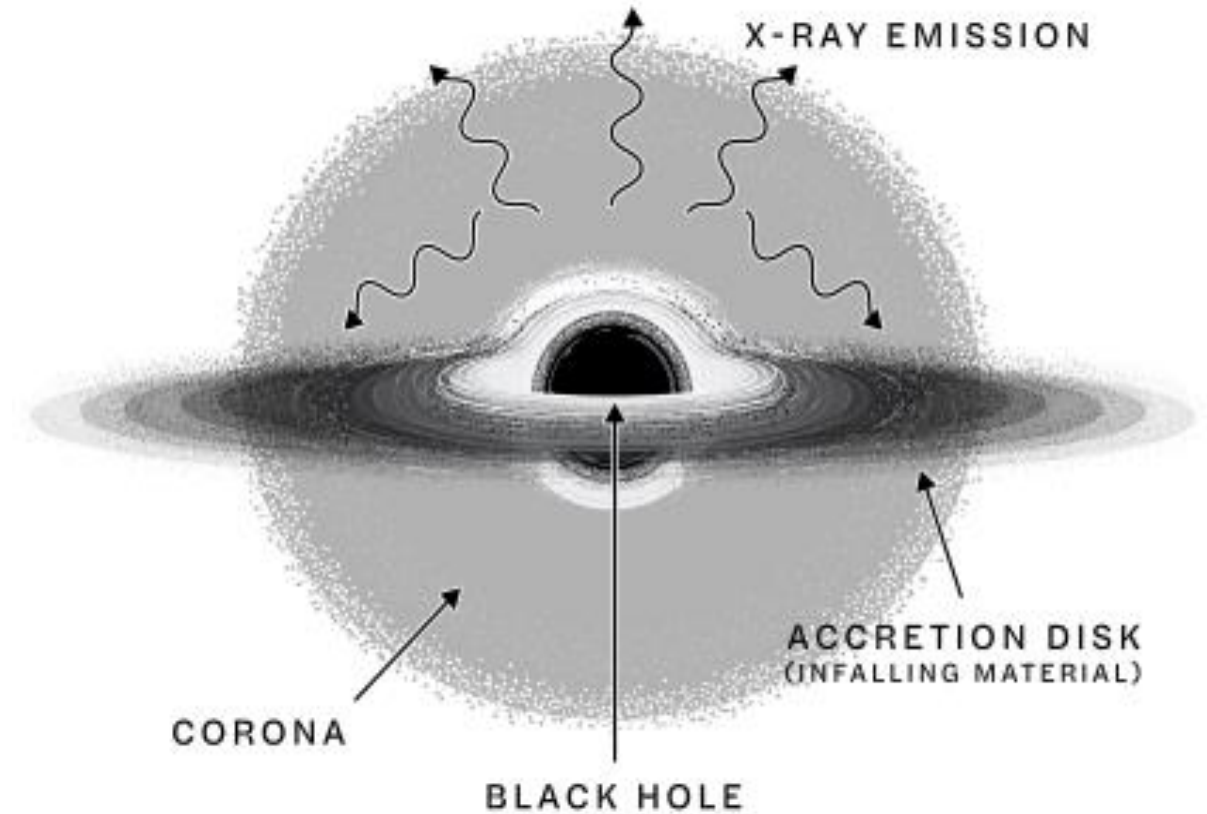
Spectral Index:

- Atmospheric γ : 3.7
- TXS 0506+056 γ : 2.37 ± 0.09

Are neutrinos from galaxy cores?

- Protons **accelerated** in accretion disk
- Interactions with **corona** produce showers
 - Corona has **high density** of X-rays and matter
 - **Different opacity** to protons and photons:

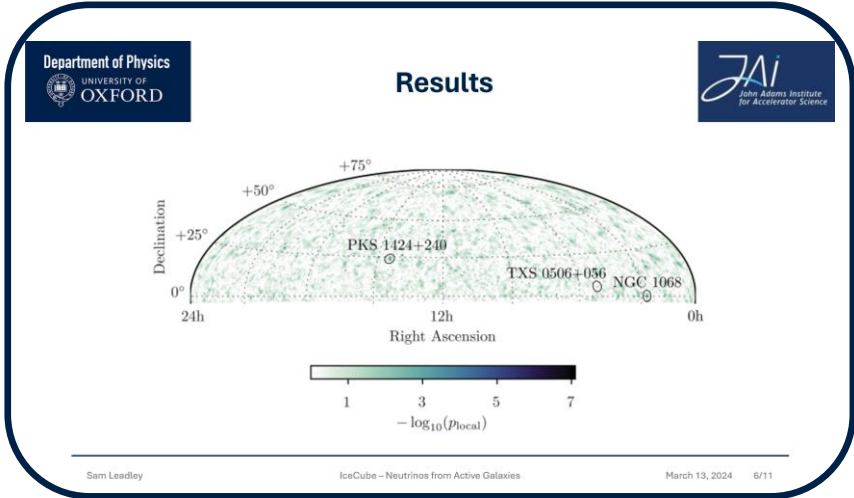
$$\sigma_\gamma \approx 1000\sigma_p$$
- Large opacity required for **neutrino production**
 - **Suppression** of photon flux guaranteed
- **Luminosity** of M77 suggests $R \approx 10R_s$



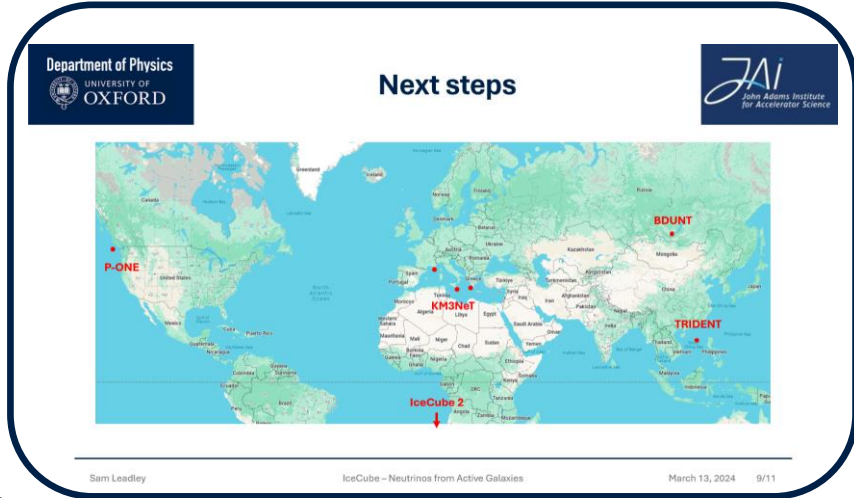
Next steps



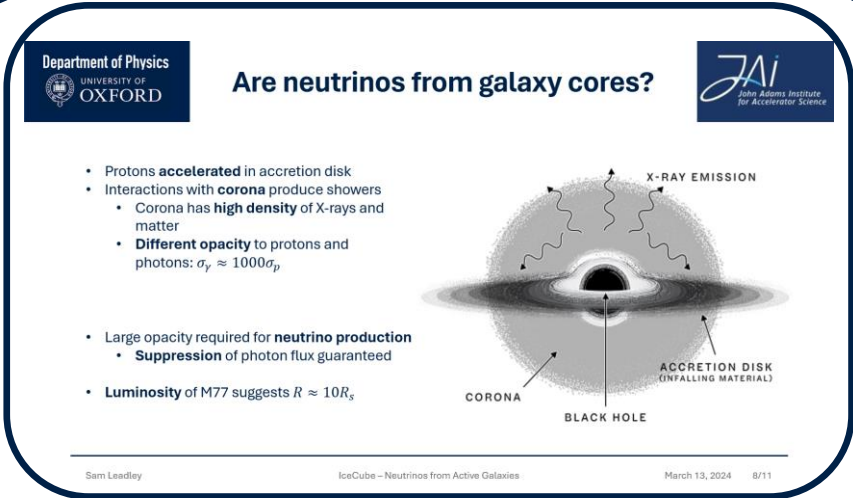
Summary and Conclusion



Proposed mechanisms explain missing high energy neutrino flux



First high energy source of cosmic neutrinos discovered



Progress requires better sky coverage and angular resolution



References

Slide 2

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Slide 3

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- R. Abbasi et. al., An Absence of Neutrinos Associated with Cosmic Ray Acceleration in Gamma-Ray Bursts (18/04/2012)
- IceCube Collaboration, Evidence for neutrino emission from the nearby active galaxy NGC 1068 (03/11/2022) (image)

Slide 5

- IceCube Collaboration, Neutrinos from Active Galaxies (11/03/2023)

Slide 6

- IceCube Collaboration, Neutrinos from Active Galaxies (11/03/2023) (images)

Slide 7

- NASA/JPL-Caltech, Shifting Coronas Around Black Holes - <https://www.jpl.nasa.gov/images/pia20051-shifting-coronas-around-black-holes-artist-concept> (27/10/2015) (image)
- IceCube Collaboration, Neutrinos from Active Galaxies (11/03/2023) (image)

Slide 8

- IceCube Collaboration, Neutrinos from Active Galaxies (11/03/2023) (image)

Slide 9

- Wikipedia, List of Neutrino Experiments (04/03/2024)

Extra Material

- IceCube Collaboration, Evidence for High-Energy Extraterrestrial Neutrinos at the IceCube detector (16/12/2013)
- B. Zhou & J. Beacom, W-boson and trident production in TeV-PeV neutrino observatories (18/02/2020)