

# Disappearing tracks at FCC

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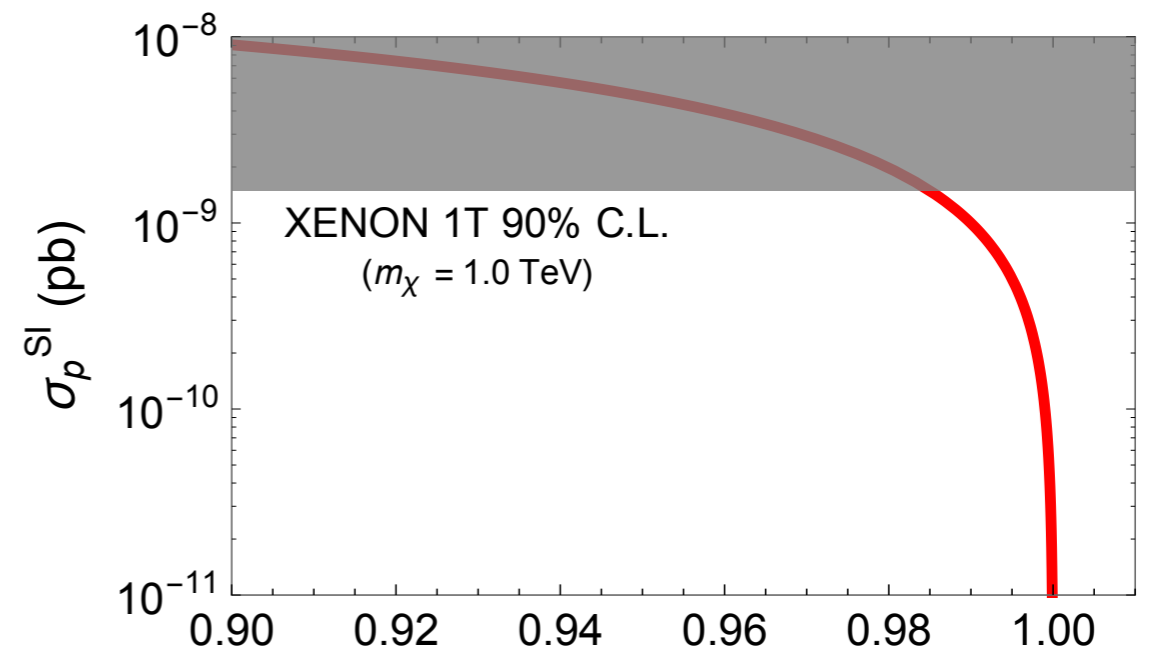
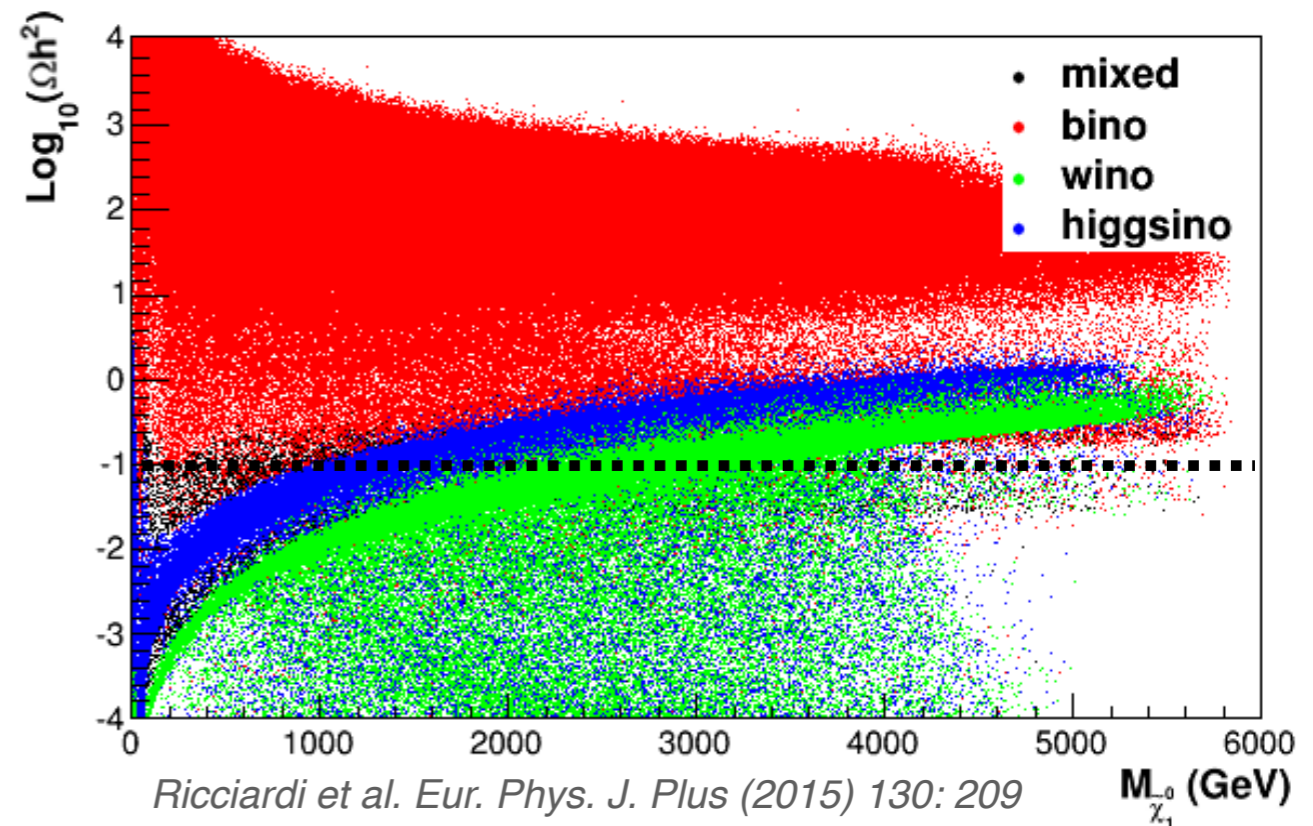


[arXiv:1901.02987](https://arxiv.org/abs/1901.02987)

[arXiv:1901.10389](https://arxiv.org/abs/1901.10389)

# Neutralino dark matter

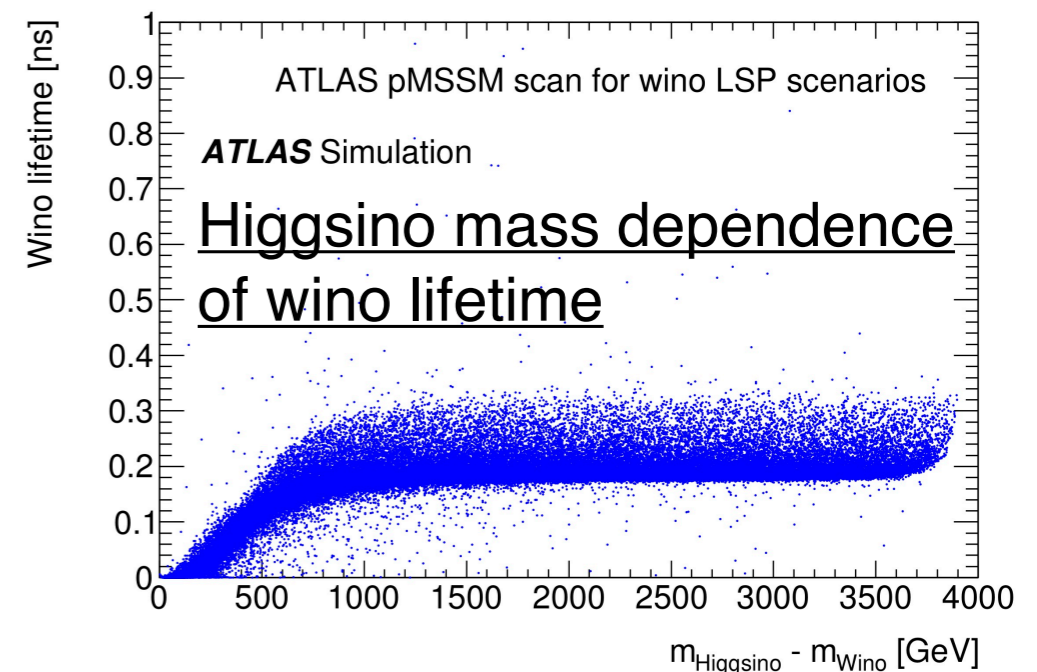
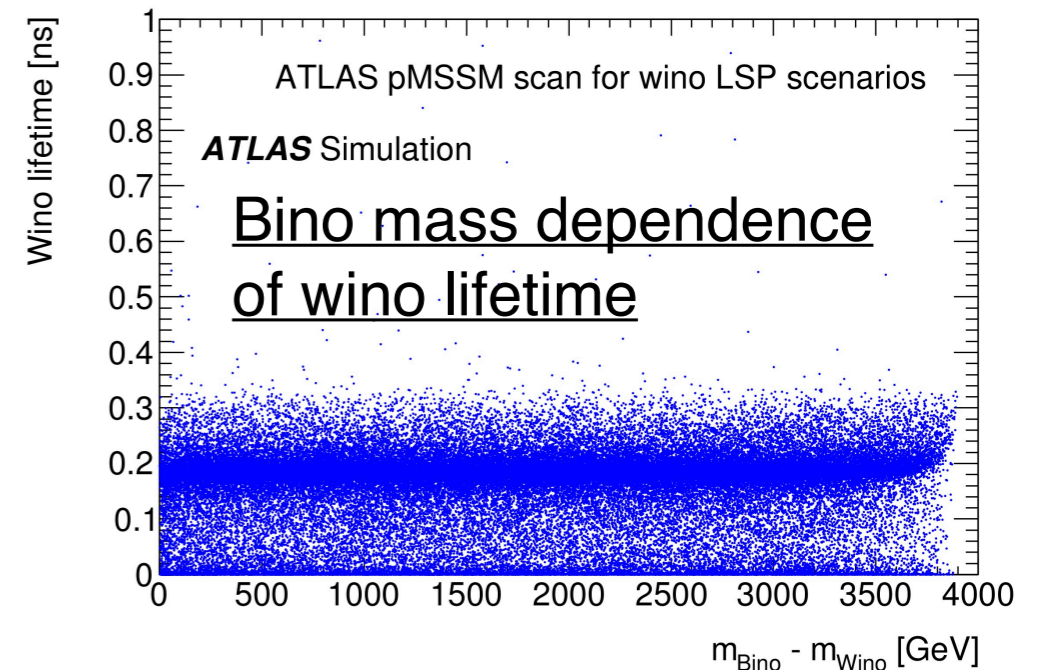
- The lightest neutralino is a good candidate of dark matter.
- Bino DM would be overproduced, while **Wino** and **Higgsino** would give the correct relic density.
- If thermally produced, the dark matter mass should be  **$\sim 1$  TeV for Higgsino** and  **$3$  TeV for Wino** scenarios.
- “Mixed” LSP models are strongly constrained by underground direct-searches.



Adv.High Energy Phys. 2018 (2018) 6828560  $f_{\text{higgsino}}$

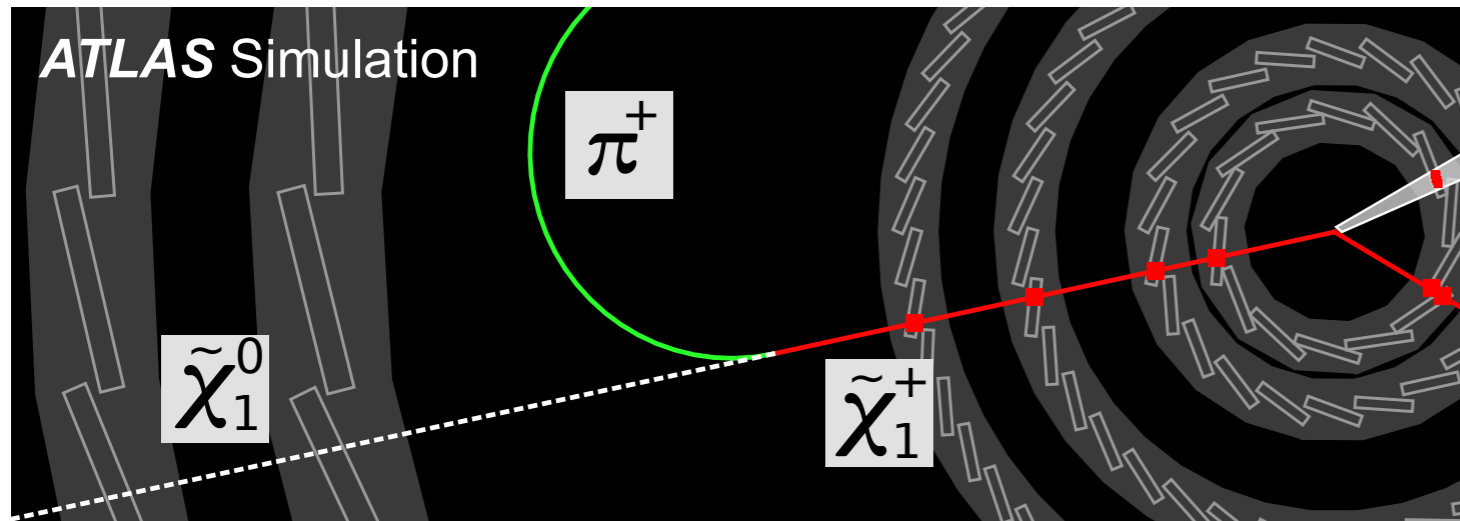
# Wino and higgsino at collider

- If the LSP is almost pure higgsino or wino, the **mass difference** between the neutral and the charged states is tiny ( $\mathcal{O}(100 \text{ MeV})$ ) as a function of the LSP mass (nearly independent of other SUSY particle masses)
- Due to the small mass difference, the lifetime of chargino is macroscopic,
  - **Wino** :  $c\tau \sim 6 \text{ cm}$
  - **Higgsino** :  $c\tau \sim 1\text{--}2 \text{ cm}$



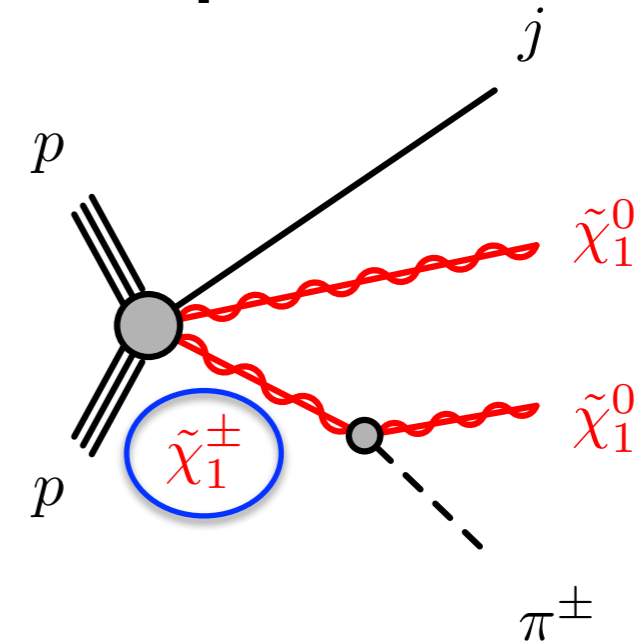
ATLAS SUSY-2016-06

# Disappearing track search @ LHC

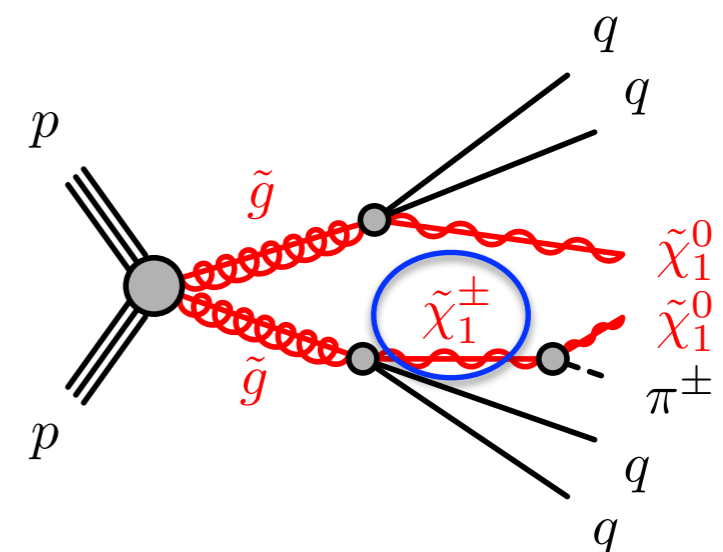


- Long lived chargino searches using **disappearing track**
  - Short tracks which do not have associated hits in the outer part of the tracker and calorimeters.
- ATLAS (36 fb<sup>-1</sup> results)
  - **Wino** excluded below **460 GeV** ( $c\tau = 0.2$  ns) JHEP 06 (2018) 022
  - **Higgsino** excluded below **152 GeV**. ATL-PHYS-PUB-2017-019
- CMS
  - **Direct production** (38 fb<sup>-1</sup>) JHEP 08 (2018) 016
  - **Full Run 2 result** for gluino-decay channel. CMS-PAS-SUS-19-005

## Direct production

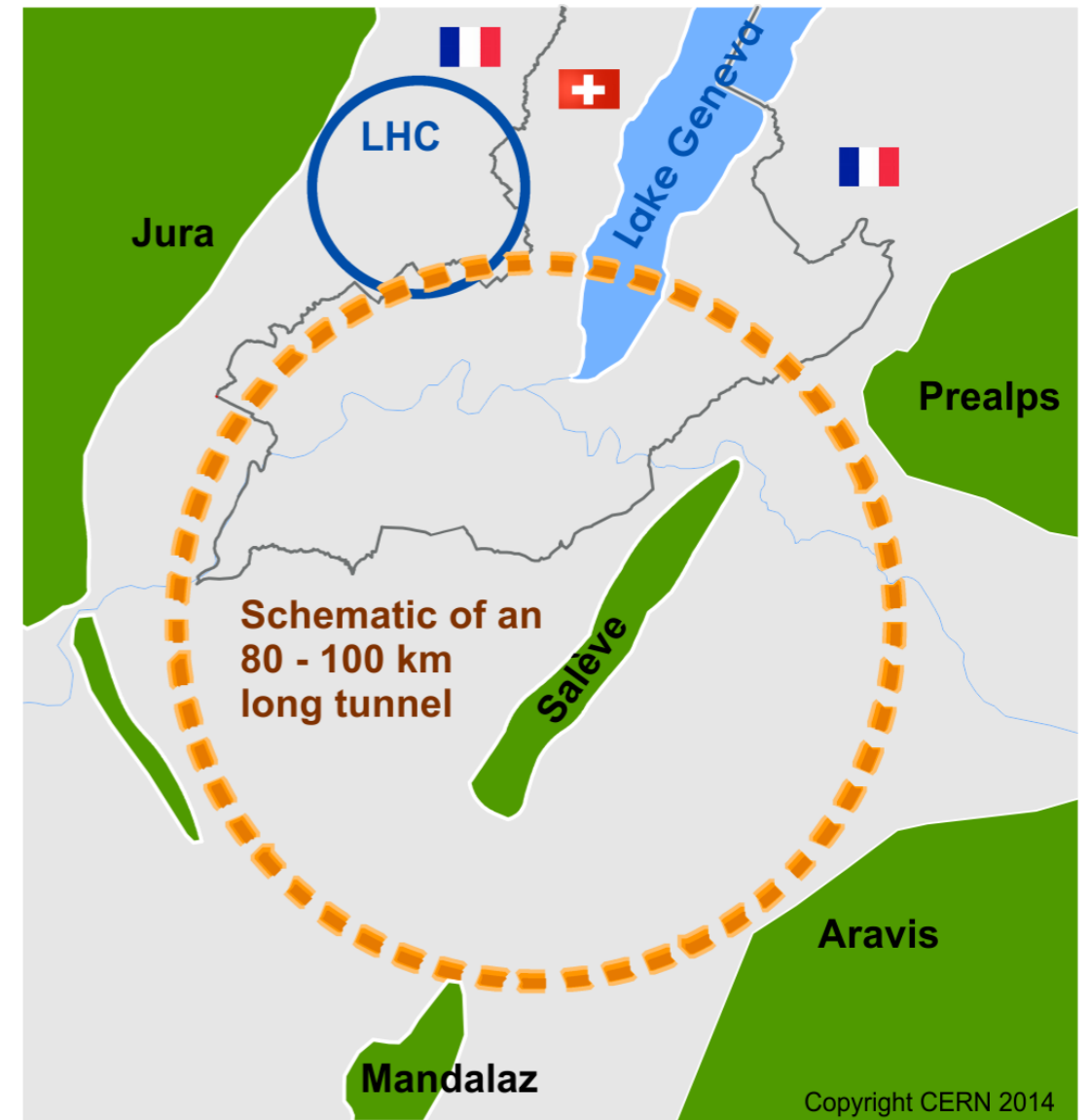
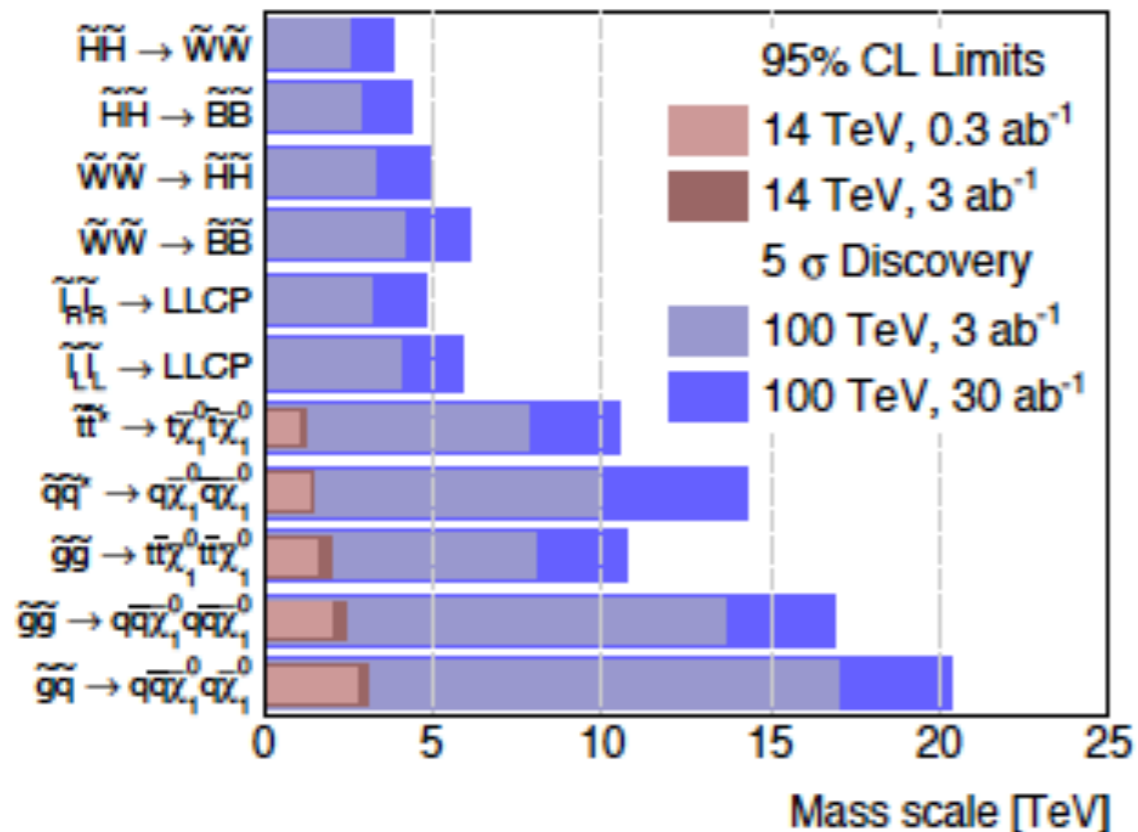


## From gluino decay



# Future Circular Collider (FCC-hh)

- 100 km tunnel in Geneva area
- $pp$  collider with  $\sqrt{s} = 100$  TeV
- 200—1000 collisions per bunch crossing
- Total integrated lumi.  $\sim 20$   $\text{ab}^{-1}$
- Much higher sensitivities to various new physics than LHC

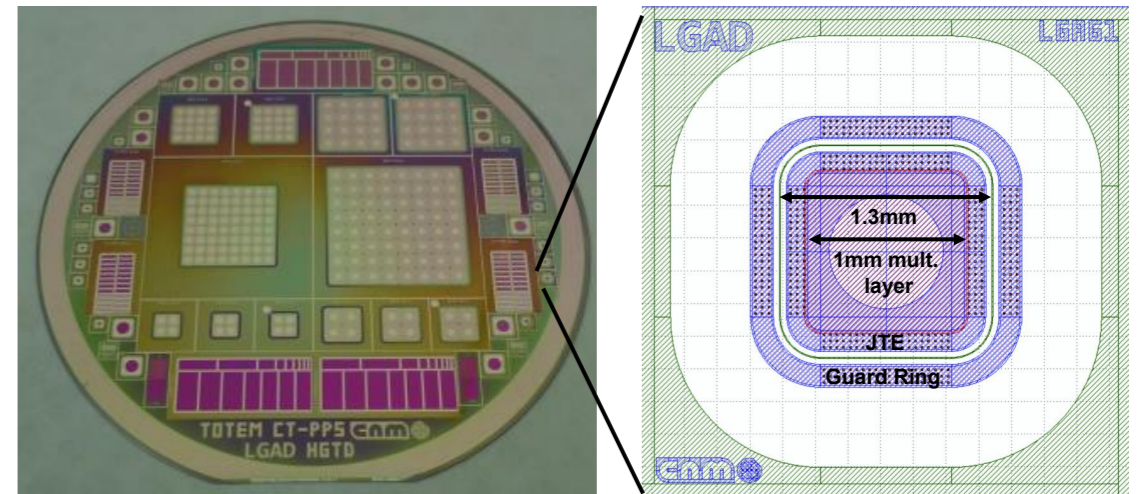


*Details in Filip's talk*

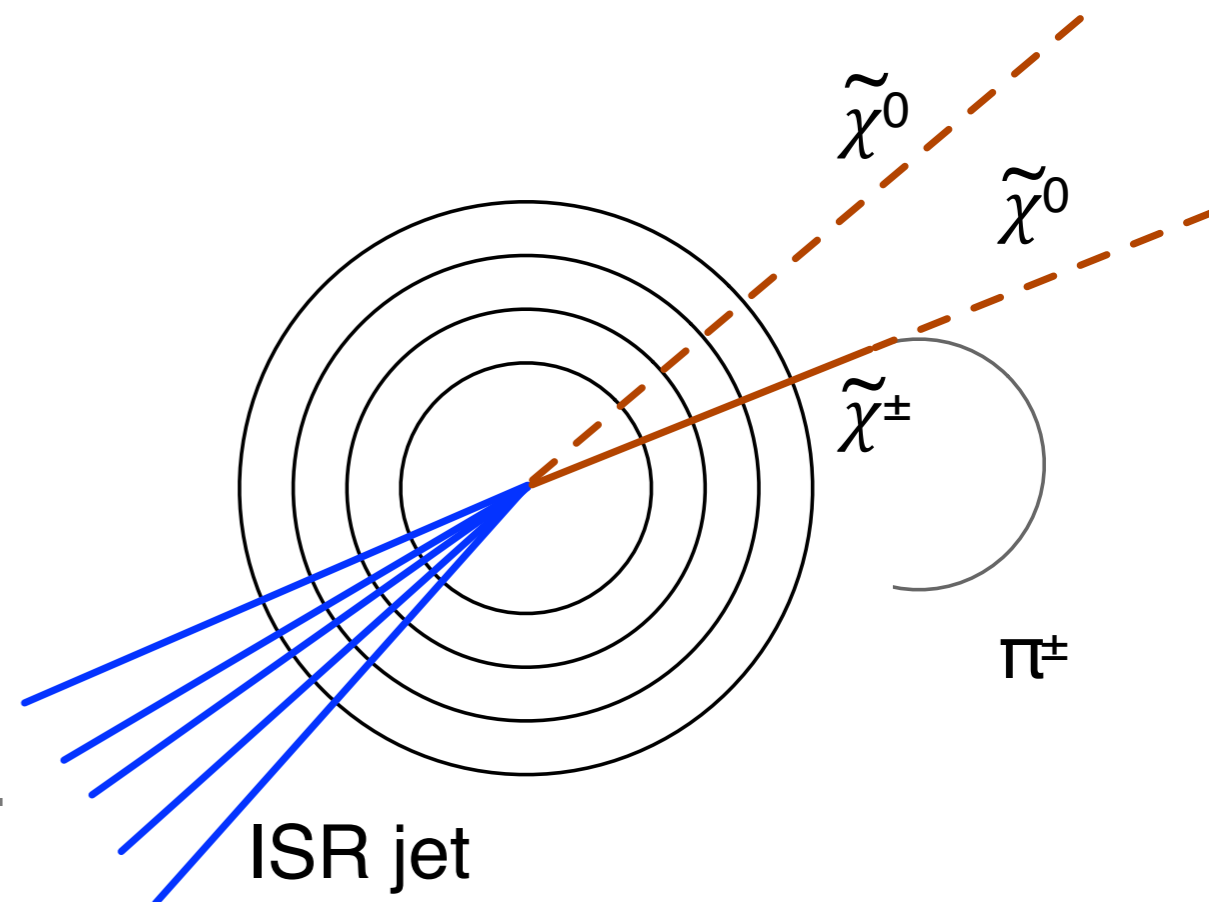
Future Circular Collider Study. Volume 3, CERN-ACC-2018-0058. Submitted to Eur. Phys. J. ST.

# DT + Hit-time information

- **Low Gain Avalanche Detectors (LGAD)** have time resolution of **10–30 ps**
- We assume here that the detector can be used in FCC as the inner pixel-detector (not at an additional timing-layer).
- We can use the hit-time for two purpose
  1. **BG fake tracks** (random-combination) decrease by requiring consistent time of pixel-hits on track.
  2. Measure **the velocity of a particle.**
    - Vertex time can be determined by ISR jets
    - If hit-time resolution is 20 ps, velocity resolution for charginos could be  $\sim 6\%$ .



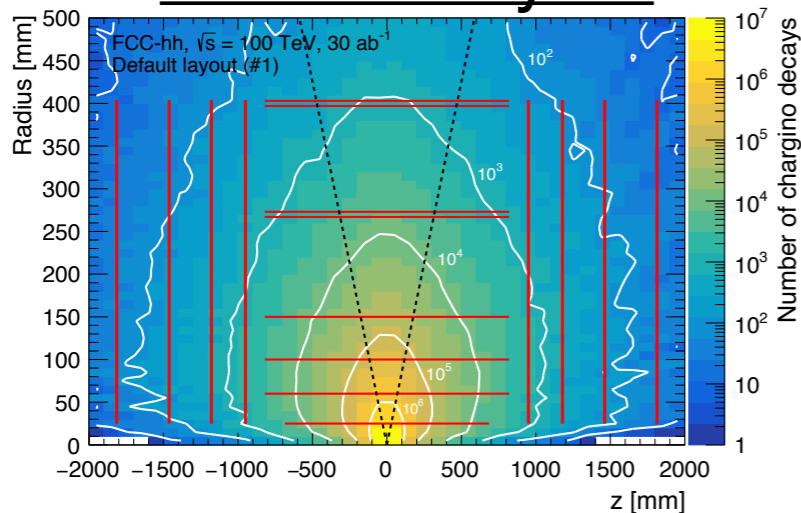
2017 JINST 12 P05003



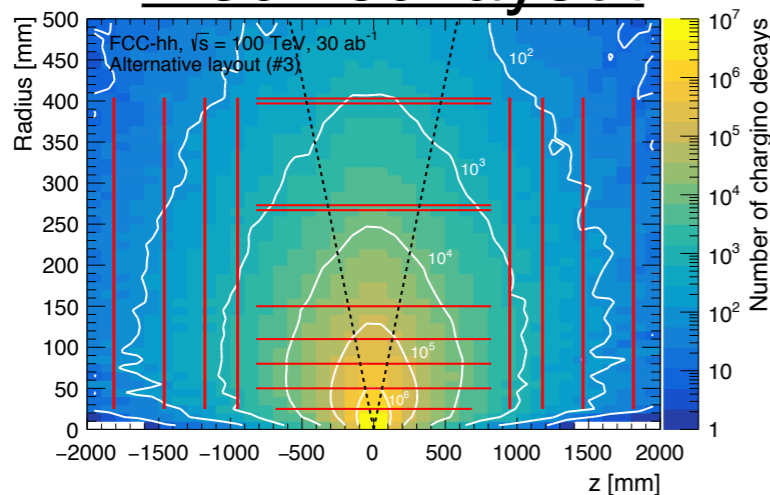
# Expected reach at FCC

- Signal and signature : direct production, ISR jet + MET + a disappearing track
- Modified pixel-detector layout (5 layers within 15cm from the beamline)
- BG rejection using time information ( $\chi^2/\text{ndf}$  in the fit)
- Fake track BG estimated using Geant4 simulation

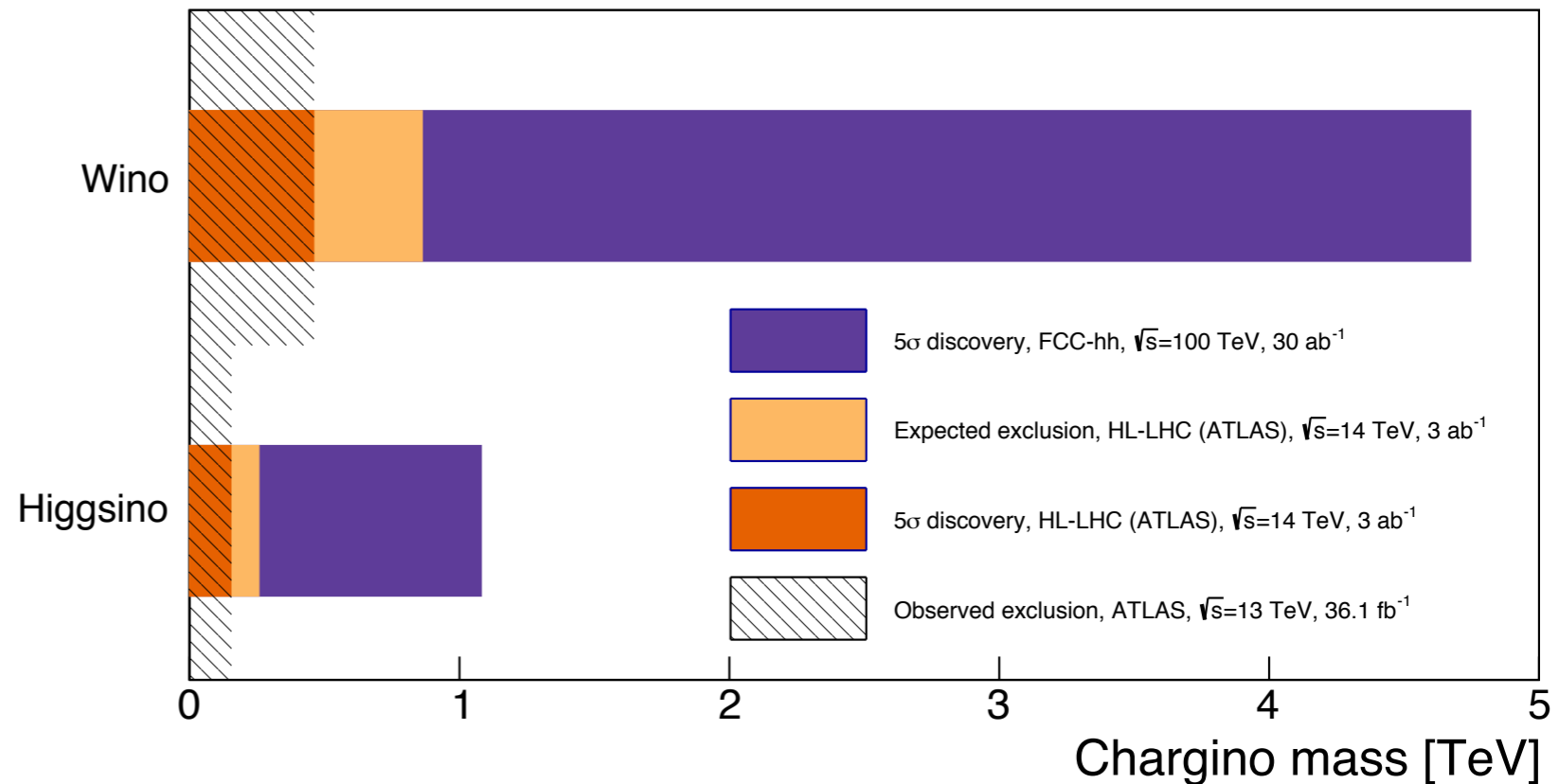
## Nominal layout



## Modified layout



## Disappearing Track



**We can discover 3 TeV wino or 1 TeV higgsino**

# Strong production of wino

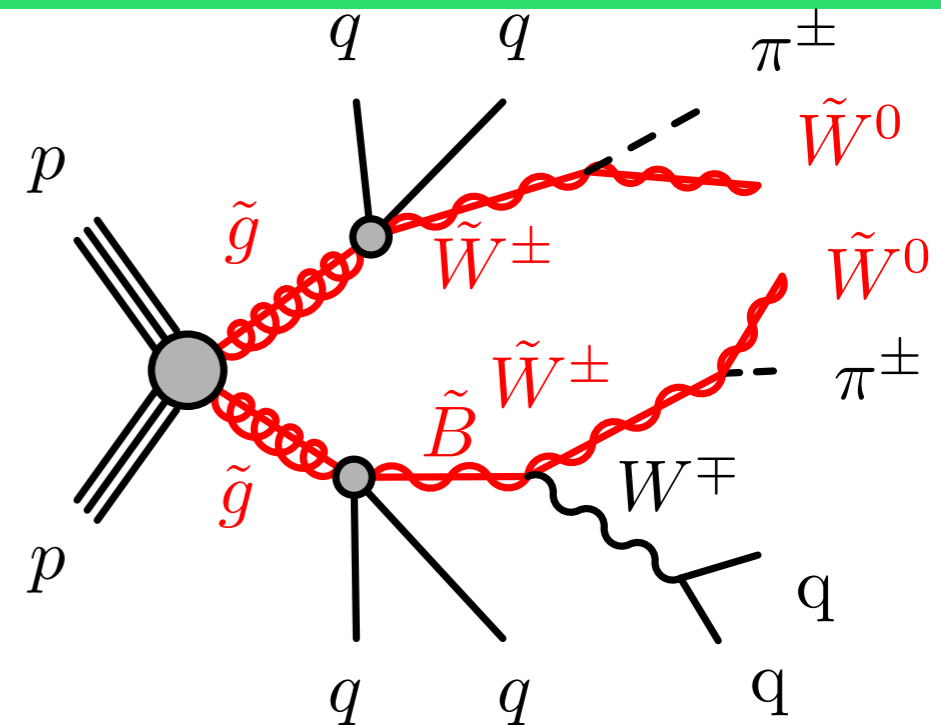
$$\tilde{g} \rightarrow \bar{q}q^{(\prime)}\tilde{W}$$

$$\tilde{g} \rightarrow \bar{q}q\tilde{B}, \quad \tilde{B} \rightarrow \tilde{W}^\pm W^\mp, \quad \tilde{W}^0 h$$

$$\text{Br}(\tilde{g} \rightarrow \bar{q}q\tilde{B}) = \text{Br}(\tilde{g} \rightarrow \bar{q}q^{(\prime)}\tilde{W}) = 0.5$$

## Benchmark points in PGM model of SUSY

	Point 1	Point 2 *	Point 3*
$m_{3/2}$ [TeV]	250	302	350
$L$ [TeV]	800	756	709
$m_{\tilde{B}}$ [GeV]	3660	4060	4470
$m_{\tilde{W}}$ [GeV]	2900	2900	2900
$m_{\tilde{g}}$ [GeV]	6000	7000	8000
$\sigma(pp \rightarrow \tilde{g}\tilde{g})$ [fb]	7.9	2.7	1.0



## Detector assumption

- 5 pixel layers in 10 cm from beamline

## Analysis pre-selection

- $E_{\text{T}^{\text{miss}}} > 1 \text{ TeV}$
- Two disappearing tracks
  - 1st :  $r > 10 \text{ cm}$
  - 2nd :  $r > (5^* \text{ or}) 10 \text{ cm}$

\* these results are in the paper

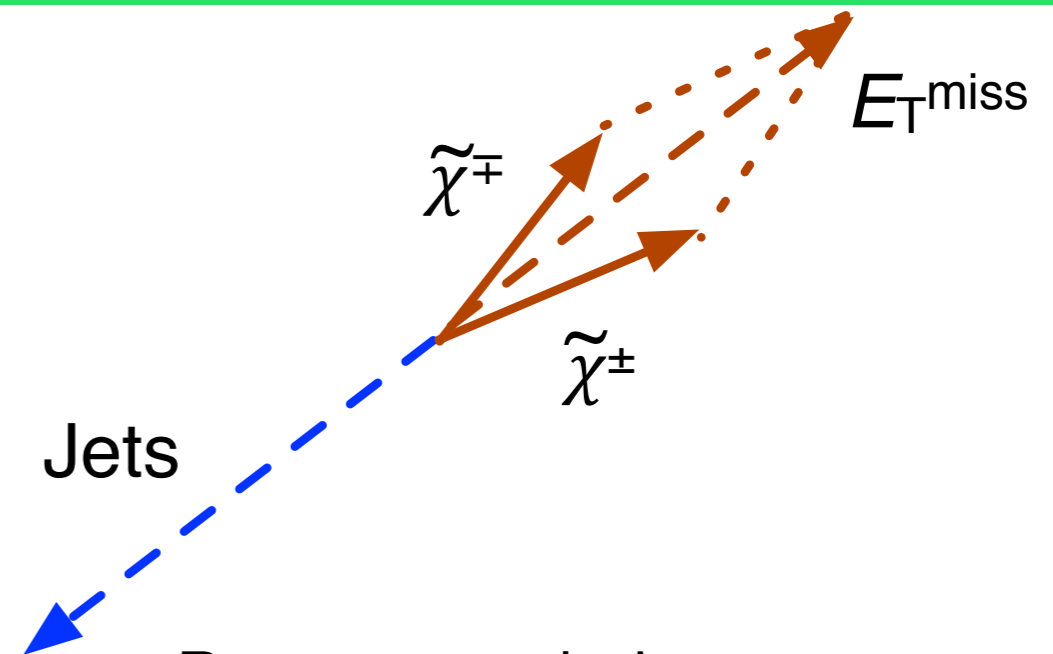


# Wino mass measurement

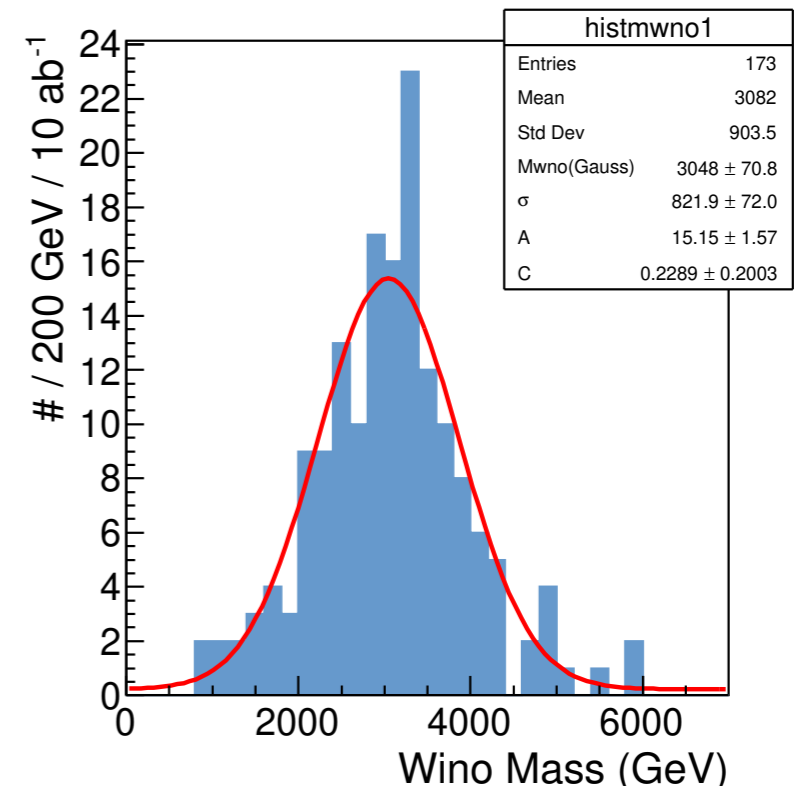
$$\text{mass} = \text{momentum} / \beta \cdot \sqrt{(1 - \beta^2)},$$

$$\beta = v/c$$

- **Velocity** is measured by the tracker using **time information**
- How to measure the **momentum** ?
  - We can not use the momentum of charged-wino tracks because of the **too poor resolution**; the track length is too short ( $< 10\text{cm}$ ) !
  - Instead, we can reconstruct from  **$E_T^{\text{miss}}$  and direction of charged winos**, because pion carry little momentum ( $O(100 \text{ MeV})$ )



Reconstructed wino mass

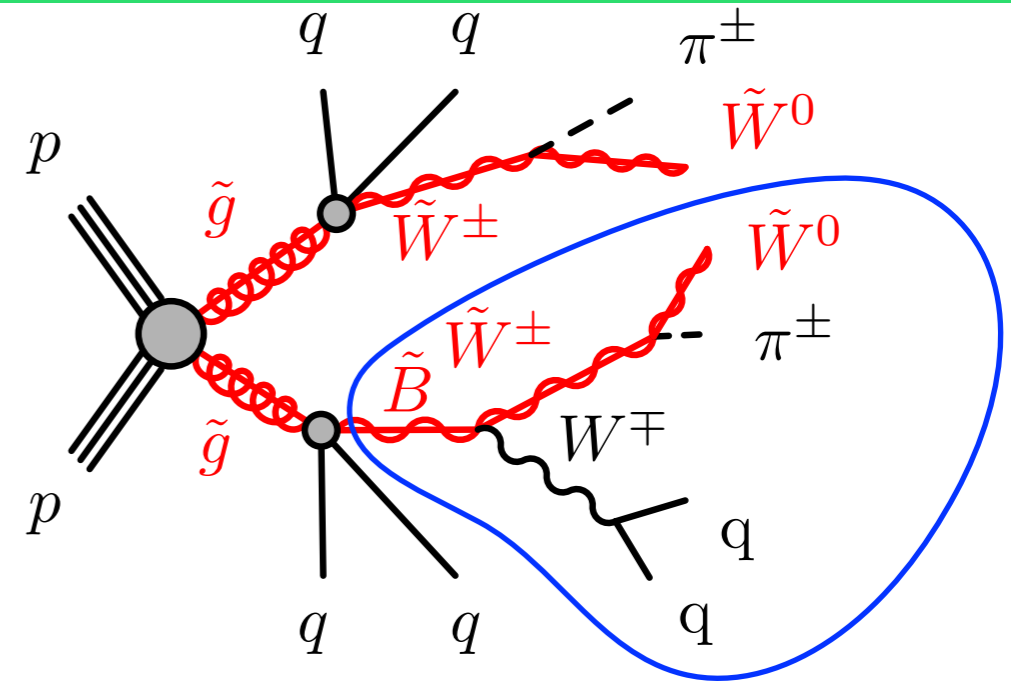


2% precision @ 2.9 TeV

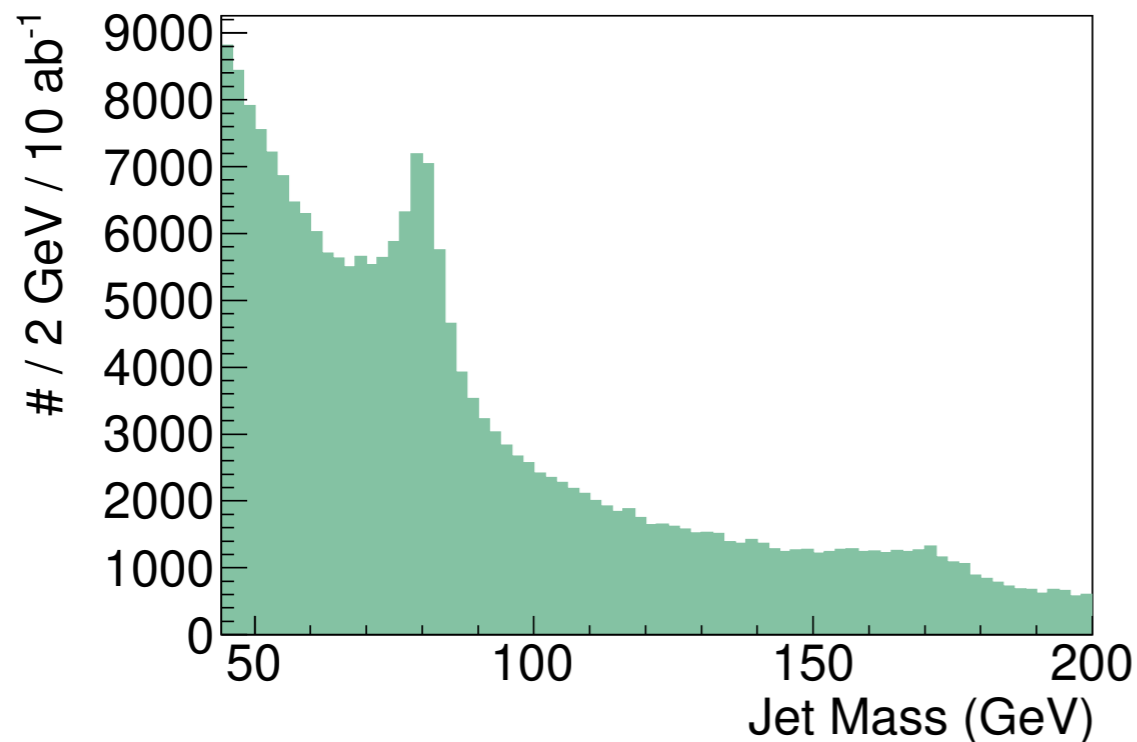
*Isolated lepton-veto is also applied*

# Bino mass measurement

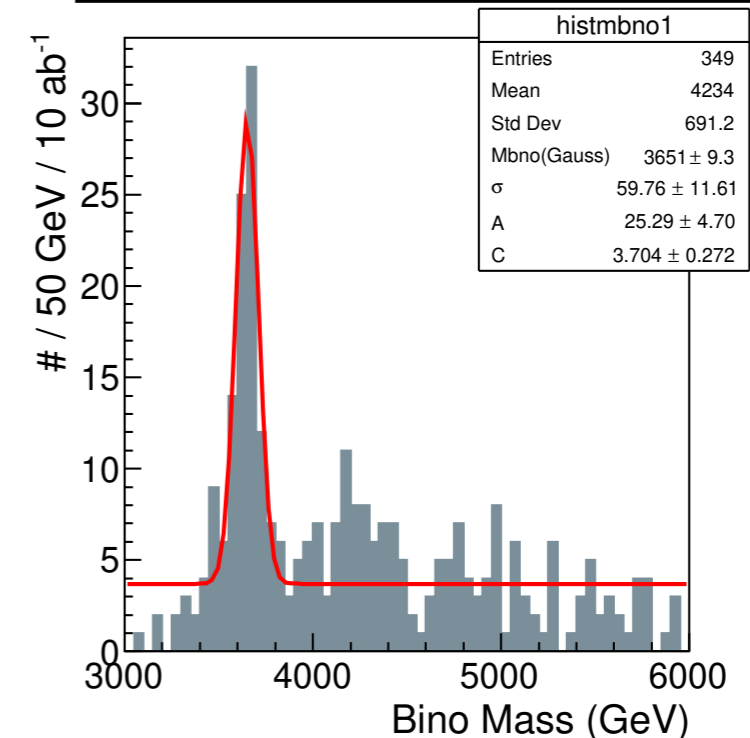
- Reconstruct **Bino mass** from **Wino and W momentum**
  - **Wino momentum** : reconstruct from the measured **velocity** and Wino **mass**.
  - **W momentum** : reconstruct using **fat jets**.



Fat jet mass



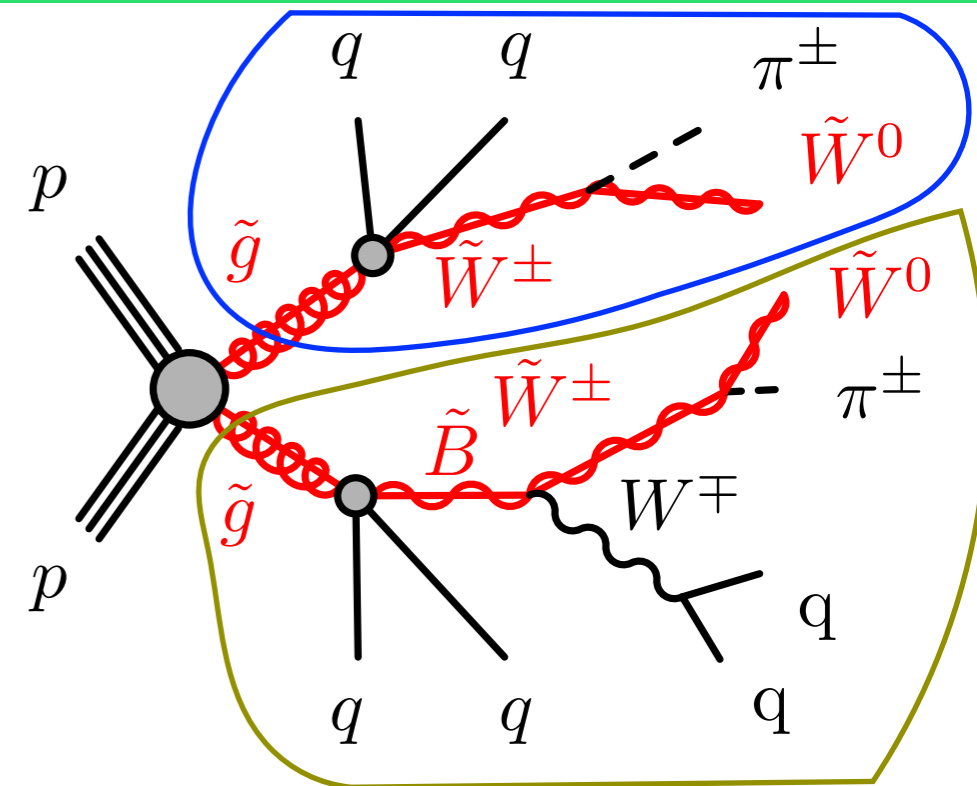
Reconstructed Bino mass



<2% precision @ 3.7 TeV

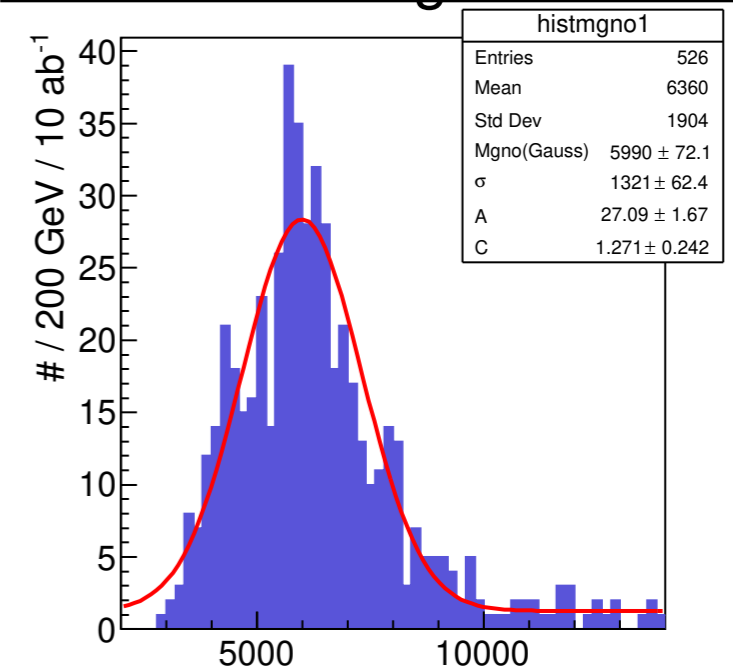
# Glauino mass measurement

- **Glauino mass** is reconstructed by “**hemi-sphere**” analysis
  1. Define two hemi-spheres using two disappearing-track directions
  2. Iteratively assign jets to each hemi-sphere and update the directions
  3. Reconstruct the glauino mass from jets and Winos.



- Glauino mass can be estimated from the **cross-section**.
  - Comparing the two estimates would be **good test of SUSY hypothesis**.

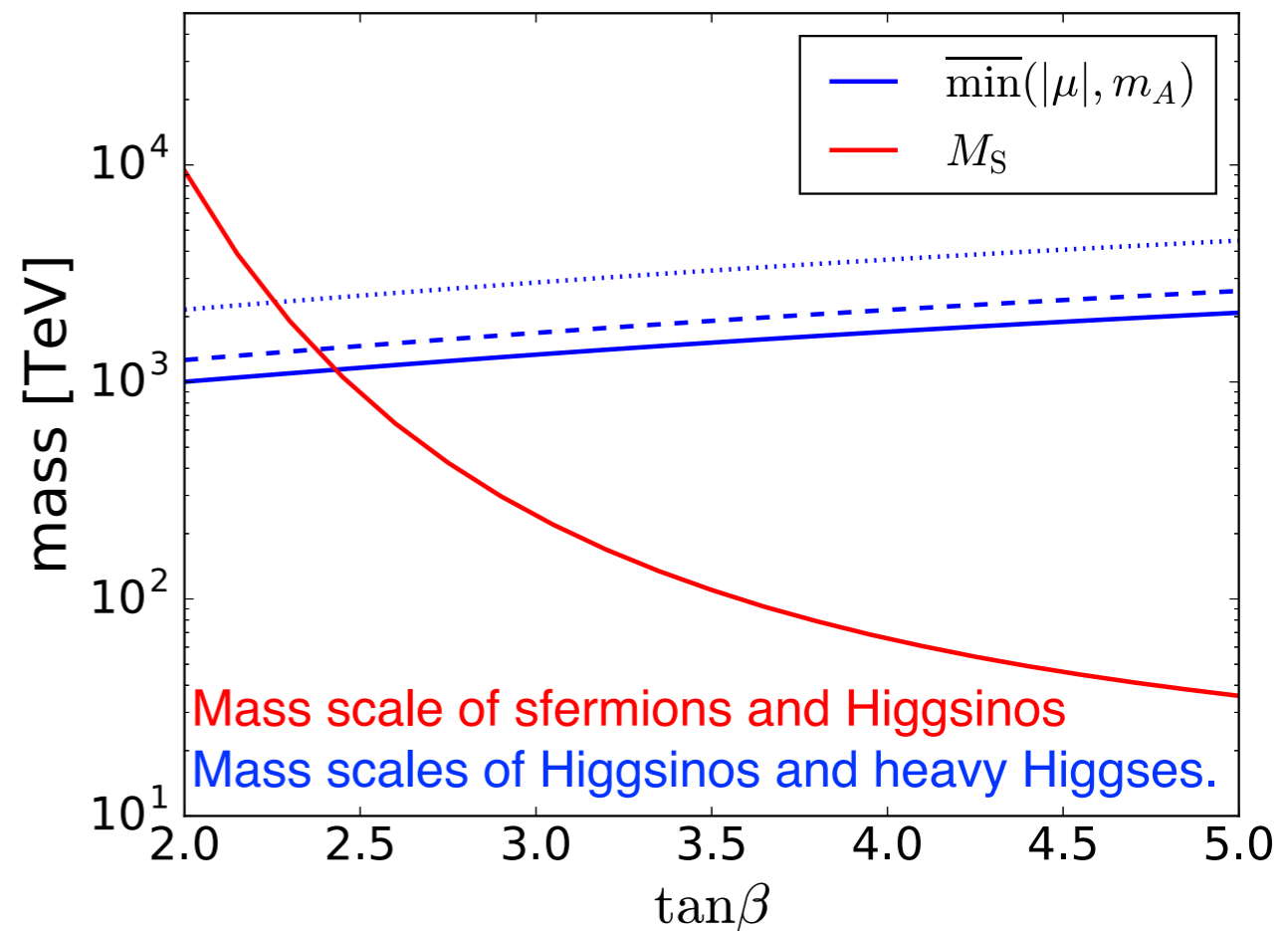
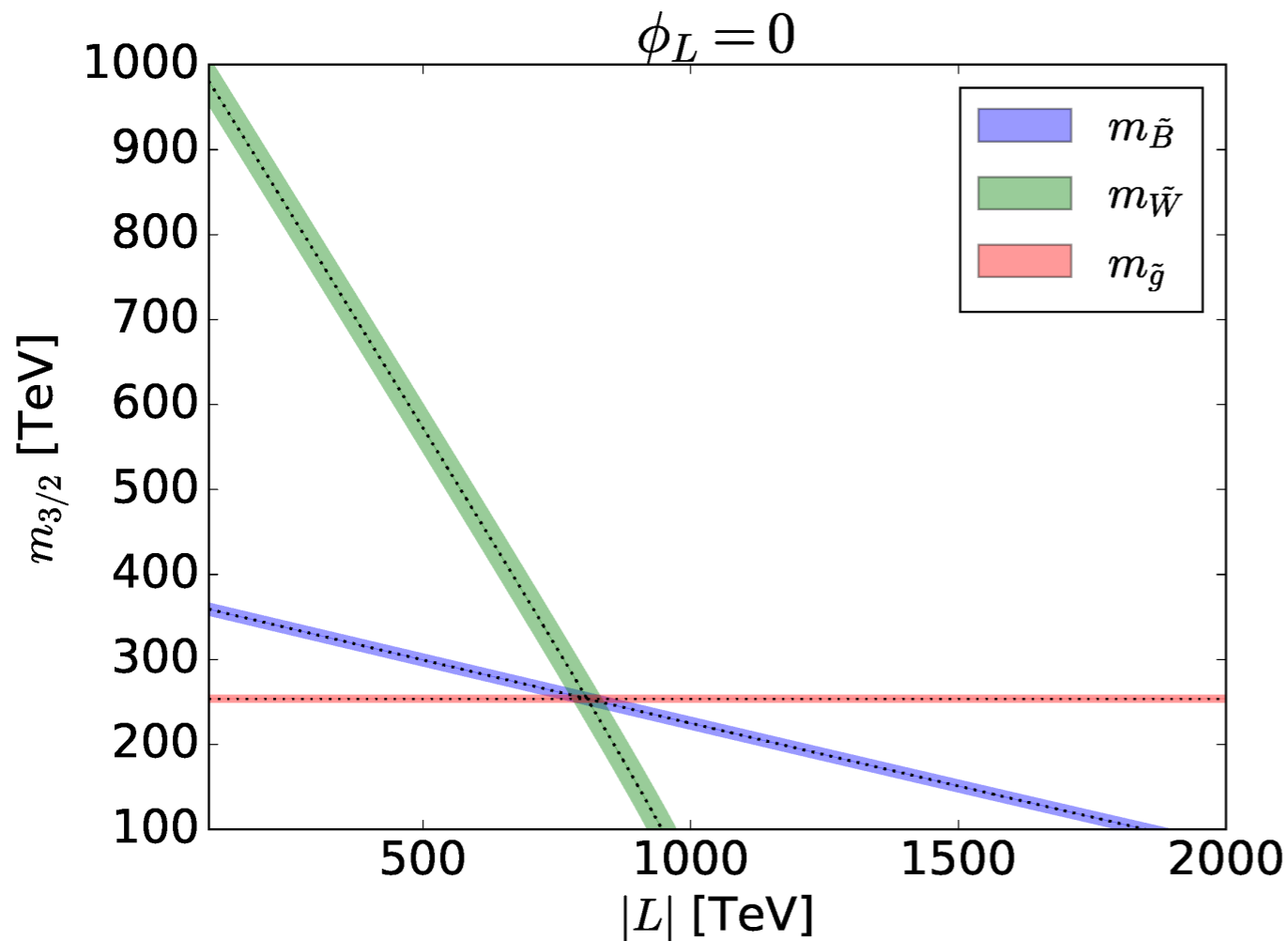
Reconstructed glauino mass



<2% precision @ 6 TeV

*Isolated lepton-veto is also applied*

# Implication



**Three model parameters** can be constrained by the gaugino mass measurements.

The gaugino mass measurements would imply also **the next particle-mass scale** (Higgsino, Higgses and sferimions)

# Conclusion

- **Timing-capable inner-most detector** in collider experiments is very effective to reduce background.
  - This will enable to **discover higgsino or wino DM**.
- Using the velocity measurement of charginos, it becomes possible to **measure all gaugino masses**.
  - This will give understanding of **the next energy scale** (beyond FCC).
  - Can we do a similar search in **LHC or HL-LHC** using pixel  $dE/dx$ , instead of timing, to measure the velocity ?