### One fluid may not rule them all - "the negative sign issue"



#### You Zhou **Niels Bohr Institute**

Based on:



UNIVERSITY OF COPENHAGEN

• W. Zhao, YZ, H. Xu, W. Deng and H. Song, Phys. Lett. B 780, 495 • W. Zhao, YZ, K. Murase and H. Song, in preparation













- comparable with Pb-Pb at low N<sub>ch</sub>, weak multiplicity dependence
- ordering  $v_2 > v_3 > v_4$
- Multi-particle cumulants
  - $v_2{4} \approx v_2{6} \approx v_2{8}$
  - Long-range multi-particle correlations in all systems -> anisotropic flow!



#### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

### Flow in small systems



### Flow vector correlations in small systems







- Correlations between flow harmonics, via symmetric cumulants



#### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan



Similar correlations between different order flow vectors observed

• Correlations between flow symmetry planes, via asymmetric cumulants



### Everything flows, everywhere flows (?)



![](_page_3_Picture_3.jpeg)

Nov 6<sup>th</sup>, 2019

### You Zhou (NBI) @ QM2019, Wuhan

#### • And many more in this QM, excellent reviews, see J. Nagle Mon. 14:30, K.K. Gajdosova Thur. 9:30

![](_page_3_Picture_7.jpeg)

#### **C.** Loizides NPA956 (2016) 200

Observable or effect	Pb–Pb	p–Pb (high mult.)	pp (high mult.)
Low $p_{\rm T}$ spectra ("radial flow")	yes	yes	yes
Intermediate $p_{\rm T}$ ("recombination")	yes	yes	yes
Particle ratios	GC level	GC level except $\Omega$	GC level except $\Omega$
Statistical model	$\gamma_s^{\rm GC} = 1, 10-30\%$	$\gamma_s^{ m GC} \approx 1,20-40\%$	MB: $\gamma_s^{\rm C} < 1, 20-40\%$
HBT radii $(R(k_{\rm T}), R(\sqrt[3]{N_{\rm ch}}))$	$R_{ m out}/R_{ m side}pprox 1$	$R_{ m out}/R_{ m side}\lesssim 1$	$R_{ m out}/R_{ m side}\lesssim 1$
Azimuthal anisotropy $(v_n)$	$v_1 - v_7$	$v_1 - v_5$	$v_2 - v_4$
(from two particle correlations)			
Characteristic mass dependence	$v_2 - v_5$	$v_2, v_3$	$v_2$
Directed flow (from spectators)	yes	no	no
Charge-dependent correlations	yes	yes	yes
Higher-order cumulants	" $4 \approx 6 \approx 8 \approx LYZ$ "	" $4 \approx 6 \approx 8 \approx LYZ$ "	" $4 \approx 6$ "
(mainly $v_2\{n\}, n \ge 4$ )	+higher harmonics	+higher harmonics	
Symmetric cumulants	up to $SC(5,3)$	only $SC(4,2), SC(3,2)$	only $SC(4,2), SC(3,2)$
Non-linear flow modes	up to $v_6$	not measured	not measured
Weak $\eta$ dependence	yes	yes	not measured
Factorization breaking	yes $(n = 2, 3)$	yes $(n = 2, 3)$	not measured
Event-by-event $v_n$ distributions	n = 2 - 4	not measured	not measured
Direct photons at low $p_{\rm T}$	yes	not measured	not observed
Jet quenching through dijet asymmetry	yes	not observed	not observed
Jet quenching through $R_{AA}$	yes	not observed	not observed
Jet quenching through correlations	yes (Z-jet, $\gamma$ -jet, h-jet)	not observed (h-jet)	not measured
Heavy flavor anisotropy	yes	yes	not measured
Quarkonia production	suppressed <sup>†</sup>	suppressed	not measured

![](_page_4_Picture_3.jpeg)

Nov 6<sup>th</sup>, 2019

### You Zhou (NBI) @ QM2019, Wuhan

### Summary Table

#### **CERN Yellow Report: CERN-LPCC-2018-07**

![](_page_4_Picture_8.jpeg)

event-by-event central p+p, p+Pb and Pb+Pb collisions at  $\sqrt{s} = 5.02 \text{ TeV}$ 

Ryan D. Weller<sup>a</sup>, Paul Romatschke<sup>a,b,\*</sup>

![](_page_5_Figure_3.jpeg)

superSONIC describes  $v_2$  and  $v_3$  data in pp, p-Pb and Pb-Pb using a single choice for the fluid parameter \* Suggests common hydrodynamic origin including pp collisions

![](_page_5_Picture_6.jpeg)

![](_page_5_Picture_7.jpeg)

![](_page_5_Picture_8.jpeg)

### "One fluid to rule them all"

One fluid to rule them all: Viscous hydrodynamic description of

Physics Letters B 774 (2017) 351–356

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_5_Picture_13.jpeg)

### Headlines in newspapers

### **A Tiny Droplet of the Early Universe?**

Tiny droplets of early universe matter created

Proton-Size Droplets of Primordial Soup May Be the Tiniest in the Universe

**Researchers Create The Tiniest Droplets Of Early Universe Matter Yet** 

### Q: Tiniest size of the hydrodynamic fluid?

![](_page_6_Picture_6.jpeg)

Task: Search for hydrodynamic flow in pp **Tool: iEBE-VISHNU** 

![](_page_6_Picture_9.jpeg)

Nov 6<sup>th</sup>, 2019

![](_page_6_Picture_12.jpeg)

## Preparations for hydro calculations

![](_page_7_Figure_1.jpeg)

- Tune parameters in hydrodynamic framework
  - fit particle spectra and integrated  $v_n$
  - good agreements between data and iEBE-VISHNU with HIJING-IC
  - not worse than superSONIC calculations

![](_page_7_Picture_6.jpeg)

![](_page_7_Picture_7.jpeg)

#### Nov 6<sup>th</sup>, 2019

#### W. Zhao, YZ, H. Xu, W. Deng, H. Song, PLB 780 (2018) 495

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_7_Picture_13.jpeg)

# Validation of hydro framework

0.1

0.05

#### R. Weller, P. Romatschke, PLB 774 (2017) 351

 $v_2(p_T)$ superSONIC for p+p, √s=5.02 TeV, 0-1% \*data for √s=13 TeV 0.12 \*\*v<sub>2</sub>, subtracted VA/2 0.1 ATLAS, N<sub>ch</sub>=60+ ATLAS\*, Nch=60+ CMS\*\*, Ntrk=110-150 0.08 <sup>5</sup> 0.06 0.04 0.02 0 1.5 0.5 0 p<sub>T</sub> (GeV)

Examinations of tuned hydrodynamic framework

![](_page_8_Picture_5.jpeg)

Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

W. Zhao, YZ, H. Xu, W. Deng, H. Song, PLB 780 (2018) 495

![](_page_8_Figure_9.jpeg)

• Describe quantitatively  $p_T$  differential  $v_2$  of both charged and identified particles -> so far so good

![](_page_8_Picture_11.jpeg)

![](_page_9_Picture_0.jpeg)

![](_page_9_Picture_1.jpeg)

W. Zhao, YZ, H. Xu, W. Deng, H. Song, PLB 780 (2018) 495

![](_page_9_Figure_3.jpeg)

 $c_2{4} = -v_2^4$ 

![](_page_9_Picture_5.jpeg)

Nov 6<sup>th</sup>, 2019

# Negative c<sub>2</sub>{4}

![](_page_9_Picture_10.jpeg)

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_9_Picture_13.jpeg)

![](_page_10_Picture_0.jpeg)

W. Zhao, YZ, H. Xu, W. Deng, H. Song, PLB 780 (2018) 495

![](_page_10_Figure_3.jpeg)

![](_page_10_Picture_4.jpeg)

Nov 6<sup>th</sup>, 2019

### You Zhou (NBI) @ QM2019, Wuhan

## No trivial bias

- Not non-flow (resonance decays)
  - $c_2{4} = c_2{4}_{2sub} = c_2{4}_{3sub}$
- Not multiplicity fluctuations
  - Using unit  $N_{ch}$  bin and then rebin into wider bin, same as experiments
- Not statistical stability
  - Huge hydro data sample has been produced

![](_page_10_Picture_14.jpeg)

# Combine $v_2\{2\}$ and $v_2\{4\}$

**v**<sub>2</sub>{**2**}

 $\checkmark$ 

![](_page_11_Figure_3.jpeg)

- $v_2{2}$ : flow + flow fluctuations  $v_2{4}$ : flow – flow fluctuations
- \* iEBE-VISHNU (HIJING-IC) works for  $v_2{2}$  but not  $v_2{4}$

![](_page_11_Picture_7.jpeg)

#### Nov 6<sup>th</sup>, 2019

### You Zhou (NBI) @ QM2019, Wuhan

![](_page_11_Figure_10.jpeg)

![](_page_11_Figure_11.jpeg)

This hydro calculation does not describe neither flow nor flow fluctuations in pp

×10<sup>-6</sup>

15

10

5

 $c_2^{4}$ 

12

**v**<sub>2</sub>{4} X

pp  $\sqrt{s} = 13 \text{ TeV}$ 

 $\blacksquare \text{ATLAS, } c_2 \{4\}_{3-\text{sub}}$ 

 $0.3 < p_{\tau} < 3.0 \text{ GeV/}c$ 

• CMS,  $c_{2}\{4\}$ 

60

40

80

ÍEBE-ÝISHNÚ

(HIJING)

Para-l

Para-II

Para-III

Para-IV

120

100

140

 $N_{\rm ch}$ 

![](_page_11_Picture_14.jpeg)

### Preparations with 3 initial conditions

![](_page_12_Figure_1.jpeg)

Testing two other different initial conditions

better than TRENTo)

![](_page_12_Picture_4.jpeg)

### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

• Hydrodynamic calculations could fit the data with selected set of parameters (HIJING and super-MC work

![](_page_12_Picture_9.jpeg)

# V<sub>2</sub>(PT) with three IC

![](_page_13_Figure_1.jpeg)

**\clubsuit** Describe qualitatively  $p_T$  differential  $v_2$  of both charged and strange particles

![](_page_13_Picture_3.jpeg)

![](_page_13_Picture_4.jpeg)

![](_page_13_Picture_5.jpeg)

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_13_Picture_7.jpeg)

|4

## All hydro give positive $c_2{4}$

![](_page_14_Figure_1.jpeg)

Hydrodynamic calculations using super-MC and TRENTo initial conditions gives even larger positive  $c_2{4}$ , and far away from data

![](_page_14_Picture_3.jpeg)

![](_page_14_Picture_4.jpeg)

![](_page_14_Picture_5.jpeg)

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_14_Picture_8.jpeg)

## Not only for iEBE-VISHNU

B. Schenke, C. Shen, and P. Tribedy, <u>arXiv:1908.06212</u>

![](_page_15_Figure_2.jpeg)

Not only for iEBE-VISHNU but maybe a current difficulty

![](_page_15_Picture_4.jpeg)

Nov 6<sup>th</sup>, 2019

#### Details, see: B. Schenke, Wed. 14.00 C. Shen, Thu.9.00

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_15_Picture_8.jpeg)

### Initial eccentricity distributions

![](_page_16_Figure_1.jpeg)

**Positive -\epsilon\_2{4**}<sup>4</sup> results in positive c<sub>2</sub>{4}

- Expected if  $v_2 \propto \varepsilon_2$ lacksquare
- $\clubsuit$  Negative  $-\epsilon_2{4}^4$  also leads to positive  $c_2{4}$ 
  - Unexpected if  $v_2 \propto \varepsilon_2$

![](_page_16_Picture_7.jpeg)

#### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

• Corresponds to wider  $p(\varepsilon_2)$  distribution (larger  $\varepsilon_2$  fluctuations, larger  $<\varepsilon_2>$ )

![](_page_16_Picture_13.jpeg)

## Non-linear hydrodynamic response

![](_page_17_Figure_1.jpeg)

![](_page_17_Picture_2.jpeg)

 $\clubsuit$  Hydro modifies the p(v<sub>2</sub>) distributions, especially the larger  $\varepsilon_2$  region

![](_page_17_Picture_5.jpeg)

#### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_17_Figure_8.jpeg)

\* For the same  $\varepsilon_2$  region, a significant non-linear (cubic) hydrodynamic response of  $v_2$  to  $\varepsilon_2$ 

non-linear response  $\longrightarrow$  additional fluctuations  $\implies$  positive c<sub>2</sub>{4}

![](_page_17_Picture_12.jpeg)

### Flow vector correlations

![](_page_18_Figure_1.jpeg)

Hydrodynamic calculations could qualitatively describe the asymmetric cumulants ac{3}, and symmetric cumulants SC(4,2)

![](_page_18_Picture_3.jpeg)

Nov 6<sup>th</sup>, 2019

![](_page_18_Picture_7.jpeg)

## Flow harmonic correlations

![](_page_19_Figure_1.jpeg)

Negative SC(3,2) observed in data, while all hydrodynamic calculations give positive SC(3,2)! It seems that hydrodynamic calculations have the difficulty to generate multi-particle (single/

mixed harmonic) cumulants correctly

Nov 6<sup>th</sup>, 2019

![](_page_19_Picture_5.jpeg)

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_19_Picture_7.jpeg)

![](_page_20_Picture_0.jpeg)

![](_page_20_Picture_1.jpeg)

- $N_{ch}$
- Other sensitive observables on flow and flow correlations
  - cumulants in hydro)
  - How about small systems?

![](_page_20_Figure_6.jpeg)

![](_page_20_Picture_7.jpeg)

![](_page_20_Picture_8.jpeg)

![](_page_20_Picture_9.jpeg)

### What next (EXP)

Recheck statistical stability: what is the limit of 3-sub-event of 4-particle cumulants in large

• Multi-particle mixed harmonic correlations (shows -, +, -, + signs of 4-, 6-, 8- and 10-particle

You Zhou (NBI) @ QM2019, Wuhan

![](_page_20_Picture_14.jpeg)

**2** |

## What next (TH)

### Other initial conditions?

- dipole formulation of BFKL evolution (arXiv:1907.12871)
- Using global bayesian analysis to constrain the parameters in IC?

### Improvements in hydro framework?

- With other 2+1D hydro framework
- 3+1D hydro?

### None of the above works?

- Other mechanisms, AMPT-escape, Kinematic, string shoving etc. One fluid might not (yet) rule pp collision?

![](_page_21_Picture_10.jpeg)

### Nov 6<sup>th</sup>, 2019

![](_page_21_Picture_15.jpeg)

### Conclusions

- Probe hydrodynamic flow in pp collisions using iEBE-VISHNU works well for all 2- and 3-particle correlations,

  - can not reproduce negative signs of  $c_2{4}$  and NSC(3,2)
- Further testings on new IC as well as hydro developments must be performed, to confirm if the fluid may rule pp collisions

![](_page_22_Picture_5.jpeg)

![](_page_22_Picture_6.jpeg)

![](_page_22_Picture_7.jpeg)

![](_page_22_Picture_10.jpeg)

### The " research investment

![](_page_23_Figure_1.jpeg)

- The negative signs have been headache for a while ...
- Whoever helps to solve the puzzle first, she/he is invited to give a seminar at NBI in Copenhagen

![](_page_23_Picture_4.jpeg)

#### Nov 6<sup>th</sup>, 2019

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_23_Picture_8.jpeg)

![](_page_23_Picture_10.jpeg)

![](_page_23_Picture_12.jpeg)

![](_page_24_Picture_0.jpeg)

![](_page_24_Picture_2.jpeg)

![](_page_24_Picture_3.jpeg)

![](_page_24_Picture_4.jpeg)

### Backup

#### You Zhou (NBI) @ QM2019, Wuhan

![](_page_24_Picture_8.jpeg)

### Symmetric Cumulants in small systems

![](_page_25_Figure_1.jpeg)

Nov 6<sup>th</sup>, 2019

![](_page_25_Picture_2.jpeg)

![](_page_25_Figure_4.jpeg)

#### Symmetric cumulants

**Correlation** between  $v_2^2$  and  $v_4^2$  in all systems

Anti-correlation between  $v_2^2$  and  $v_3^2$  at high multiplicities, a **transition** to positive correlation followed by both small and large systems

Not described by non-flow only models, but qualitatively predicted by model with initial stage correlations

#### What's next:

lacksquare

ullet

SC(m<sup>i</sup>,n<sup>j</sup>), SC(m,n,k)

challenges: statistics

![](_page_25_Picture_14.jpeg)

## Normalized Symmetric Cumulants

![](_page_26_Figure_1.jpeg)

![](_page_26_Picture_2.jpeg)

- Use NSC<sup>v</sup>(3,2) to constrain initial state model
- $v_n$  might not be linearly correlated with  $\varepsilon_n$  in small systems (e.g. pp)
  - it generates additional fluctuations which changes sign of  $c_2{4}$
  - one should not compare initial NSC<sup> $\epsilon$ </sup>(3,2) in model calculations to NSC<sup>v</sup>(3,2) data
  - It also make less sense to compare  $v_2\{6\}\{8\}/v_2\{4\}$  and  $\varepsilon_2\{6\}\{8\}/\varepsilon_2\{4\}$  in SS.

#### Nov 6<sup>th</sup>, 2019

![](_page_26_Picture_11.jpeg)

![](_page_27_Figure_0.jpeg)

![](_page_27_Picture_1.jpeg)

![](_page_27_Picture_2.jpeg)

![](_page_27_Picture_3.jpeg)

![](_page_27_Figure_4.jpeg)

You Zhou (NBI) @ QM2019, Wuhan

![](_page_27_Picture_7.jpeg)

![](_page_28_Picture_0.jpeg)

### **\*** iEBE-VISHNU

![](_page_28_Picture_2.jpeg)

![](_page_28_Picture_3.jpeg)

![](_page_28_Picture_4.jpeg)

C. Shen, Z. Qiu, H. Song, J. Bernhard, S. Bass and U. Heinz. Comput. Phys. Commun. 199, 61 (2016)

![](_page_28_Picture_6.jpeg)

![](_page_28_Picture_7.jpeg)

![](_page_28_Picture_8.jpeg)

![](_page_28_Picture_12.jpeg)

![](_page_29_Picture_0.jpeg)

![](_page_29_Picture_1.jpeg)

In HIJING initial model, the produced jets pairs and excited nucleus are treated as independent strings, and these strings break into partons and quickly form hot spots for succeeding hydrodynamics. The center positions of strings  $(x_c, y_c)$  are sampled by Saxon-Woods distribution, and positions of partons within the strings are sampled by, exp  $\left(-\frac{(x-x_c)^2+(y-y_c)^2}{2\sigma_c^2}\right)$ 

HIJING constructs energy density by energy decompositions of individual partons via a Gaussian smearing:

$$\epsilon = K \sum_{i} \frac{E_i^*}{2\pi\sigma^2\tau_0 \Delta\eta_s} \exp\left(-\frac{(x-x_i)^2 + (y-y_i)^2}{2\sigma^2}\right)$$

![](_page_29_Picture_5.jpeg)

![](_page_29_Picture_6.jpeg)

# **HIINGIC**

	1					
	$\sigma_{R}$	$\sigma_0$	$ au_0$	$\eta/s$	$T_{ m sw}({ m MeV})$	
Para-I	1.0	0.4	0.1	0.07	147	1
Para-II	0.8	0.4	0.2	0.08	148	1
Para-III	0.4	0.2	0.6	0.20	148	1
Para-IV	0.6	0.4	0.4	0.05	147	1

![](_page_29_Figure_10.jpeg)

![](_page_29_Figure_13.jpeg)

![](_page_29_Picture_14.jpeg)

![](_page_30_Picture_0.jpeg)

In super-MC the entropy density is:

$$s(\mathbf{r}) = \frac{\kappa_s}{\tau_0} \sum_{k=1}^3 \gamma_k^{(i)} \, \frac{e^{-(\mathbf{r} - \mathbf{r}_k^{(i)})^2 / (2\sigma_g^2)}}{2\pi \sigma_g^2},\tag{5}$$

where  $\gamma_k$  is sampled from  $\Gamma$  distribution,  $\mathbf{r}_k^{(i)}$  is quark's positions,  $\sigma_g$  is width of gluons.

	$\sigma_{R}$	$\sigma_0$	$ au_0$	$\eta/s$	$T_{\rm sw}({ m MeV})$
Para-I	1.0	0.4	0.1	0.07	147
Para-II	0.8	0.4	0.2	80.0	148
Para-III	0.4	0.2	0.6	0.20	148
Para-IV	0.6	0.4	0.4	0.05	147

Nov 6<sup>th</sup>, 2019

**``** 

![](_page_30_Picture_6.jpeg)

### super-MC & TRENTo

In TRENTO the initial entropy density is:

$$s = s_0 \left(\frac{\tilde{T}_A^p + \tilde{T}_B^p}{2}\right)^{1/p}, \qquad (6)$$

where  $\tilde{T}(x,y) \equiv \int dz \frac{1}{n_c} \sum_{i=1}^{n_c} \gamma_i \rho_c (\mathbf{x} - \mathbf{x_i} \pm \mathbf{b}/2)$ ,  $n_c$  is the number of

the independent constituents and  $\rho_c(\mathbf{x}) = \frac{1}{(2\pi v^2)^{3/2}} \exp\left(-\frac{\mathbf{x}^2}{2v^2}\right)$ ,

	p	v	k	$n_c$	$ au_0$	$\eta/s$	$T_{\rm sw}({ m MeV})$
Para-I	0.5	0.3	1.5	4	0.2	0.08	149
Para-II	0.0	0.2	0.81	6	0.6	0.28	149
Para-III	0.5	0.2	1.0	4	0.8	0.28	149

![](_page_30_Picture_15.jpeg)